

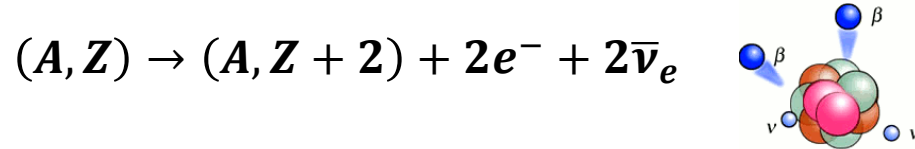


# BINGO active veto

Hawraa Khalife (Postdoc at IRFU/CEA)  
GDR DI2I  
10–12 July 2023 @ SUBATECH (Nantes)

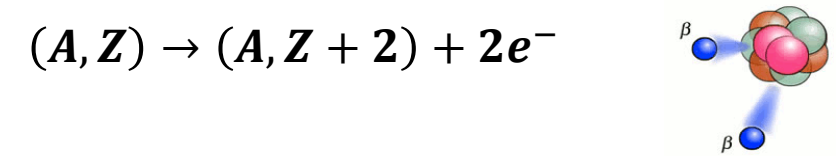
# double beta decay

## Two-neutrino mode: $2\nu\beta\beta$

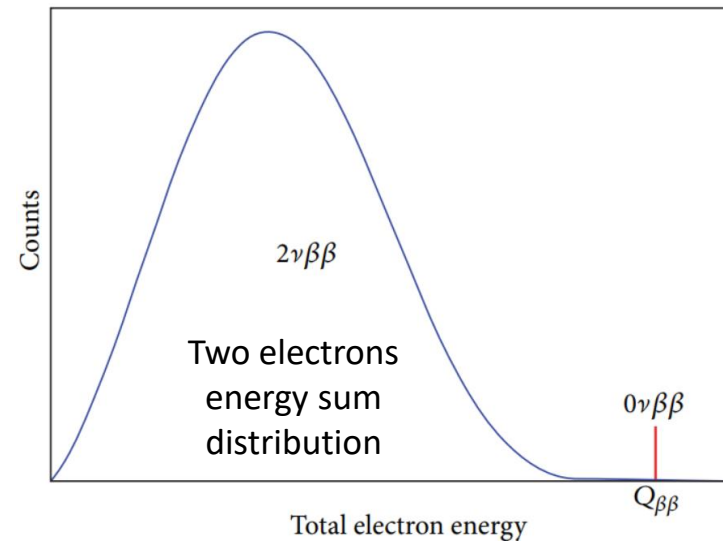


- It can occur for even-even nuclei
- Allowed in the SM
- It is energetically allowed for 35 nuclides and has been observed in 1/3 of them ( $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{96}\text{Zr}$ ,  $^{100}\text{Mo}$ ,  $^{116}\text{Cd}$ ,  $^{128}\text{Te}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$ ,  $^{238}\text{U}$ )
- Extremely long half-life  $T_{1/2} \sim 10^{18} - 10^{24}$  yr

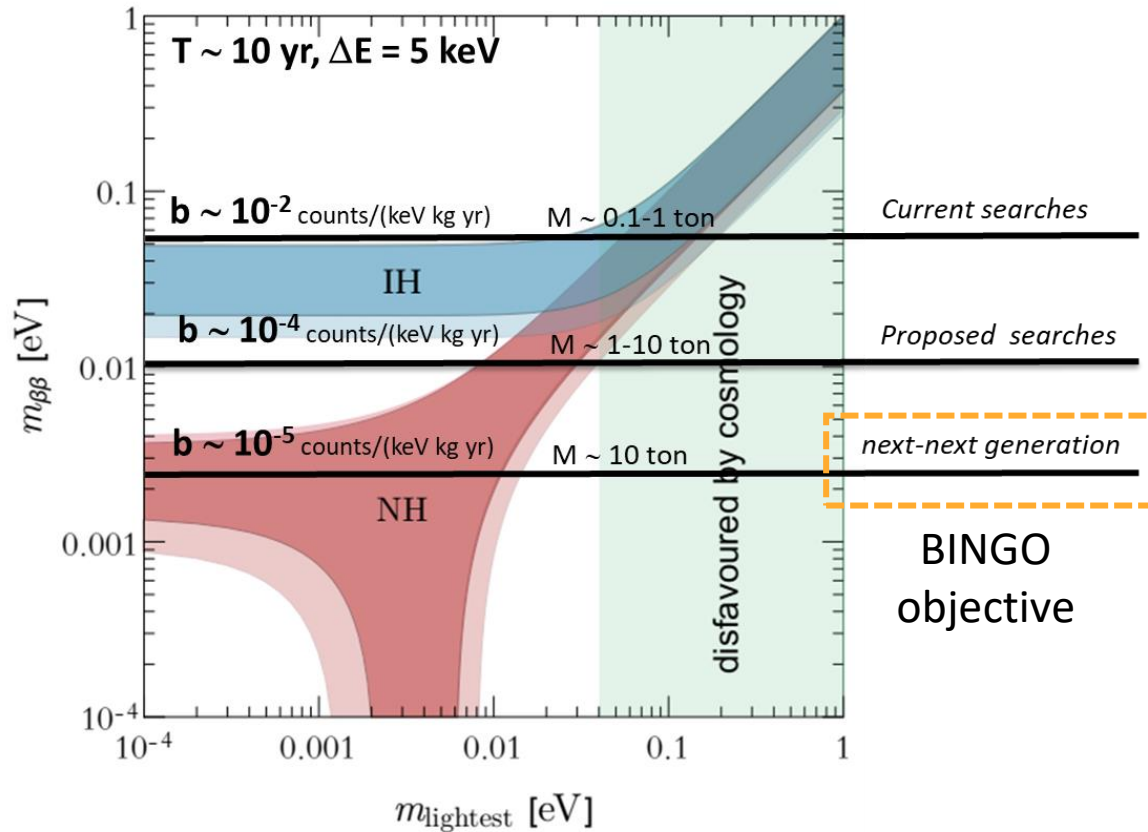
## Neutrinoless mode: $0\nu\beta\beta$



- Lepton number violation
- Neutrinos are massive Majorana particles ( $\bar{\nu} = \nu$ )
- Measure  $T_{1/2}^{0\nu}$  that will lead to effective Majorana neutrino mass  $m_{\beta\beta}$  measurement



# Sensitivity to effective Majorana mass



$$b = \frac{\text{number of background counts}}{M \times T \times \Delta E}$$

$b$ : background index

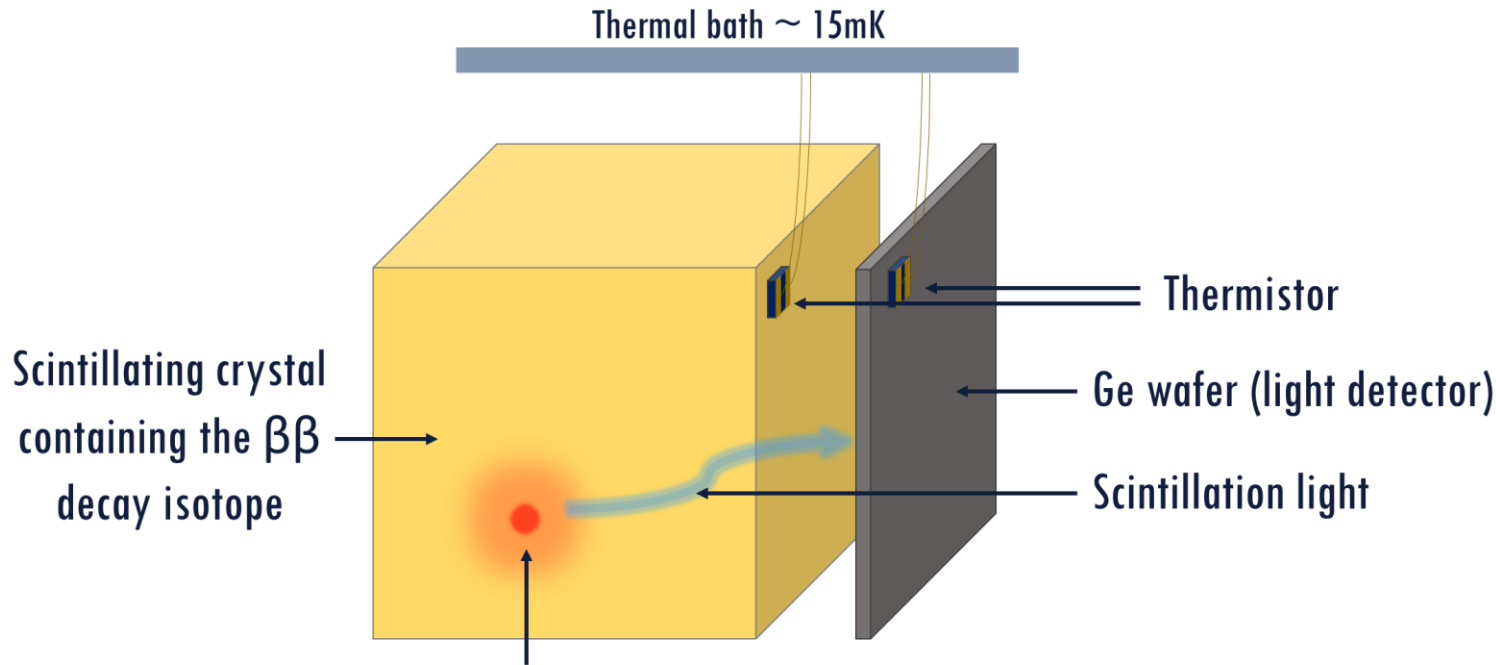
$M$ : detector mass

$T$ : time of measurement

$\Delta E$ : energy resolution of the detector (FWHM)

BINGO aims to reduce the background in the region of interest, which is one of the most limiting factors for current and future  $0\nu 2\beta$  bolometric experiments

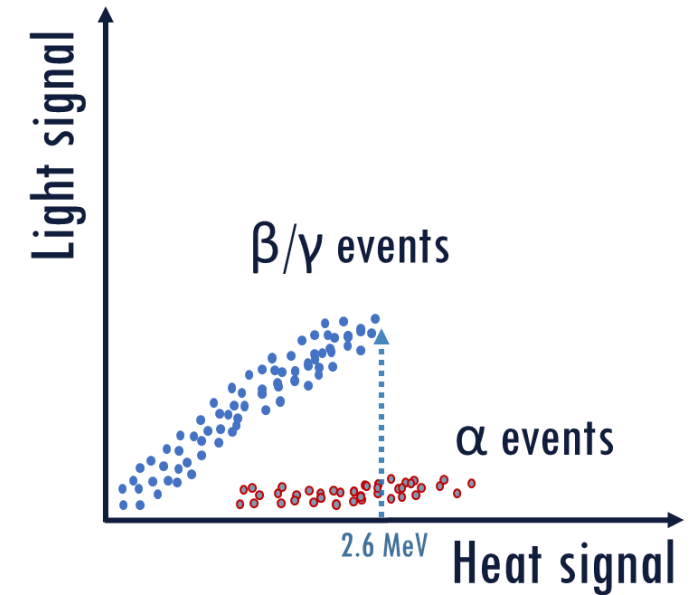
# Bolometric technique (Dual read-out)



Particle interaction leads to:

- Heat signal
- Light signal

The dual heat-light readout allows to reject the alpha contamination thanks to the lower light output of  $\alpha$ s



## Features

- high energy resolution
- full active volume (no dead layer)
- flexible material choice (**BINGO** will use  $\text{Li}_2\text{MoO}_4$  ( $^{100}\text{Mo}$ ) and  $\text{TeO}_2$  ( $^{130}\text{Te}$ ) with  $Q_{\beta\beta}$  at 3034 keV and 2527 keV respectively)

# Mitigating the background

## $\alpha$ surface contamination

- $\text{Li}_2\text{MoO}_4$  (scintillator): dual heat-light readout
- $\text{TeO}_2$  (bad scintillator): Cherenkov light detection using Neganov-Luke (NL) light detector (proposed by BINGO)

## $\beta$ Surface contamination

- New technology: surface sensitive detectors (proposed by CROSS)
- **Use an active shield (the crystal should be facing the veto)**

## $2\nu 2\beta$ pile up

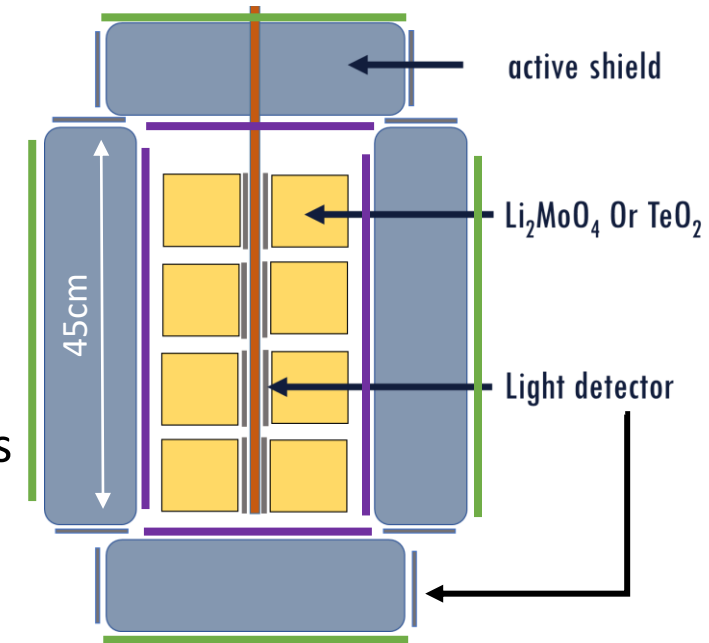
- $\text{Li}_2\text{MoO}_4$ : increase light detector speed and signal to noise ratio (NL LDs)
- $\text{TeO}_2$ : 100 times lower rate than  $^{100}\text{Mo}$

## $\gamma$ radioactivity

- Detector holders: fine selection and reduction of materials (new assembly design proposed by BINGO)
- **Shields and external materials: use an active shield**

# The BINGO veto

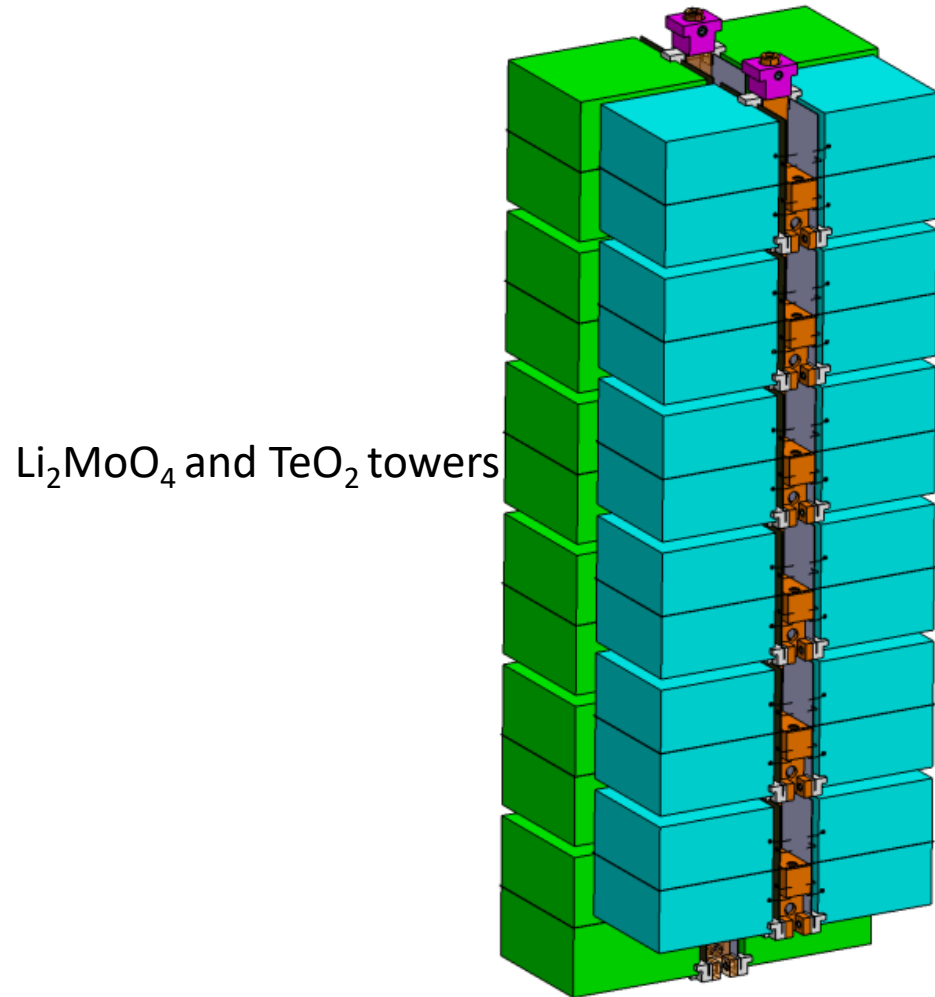
- An active inner shield will be used to surround the  $\text{Li}_2\text{MoO}_4$  and  $\text{TeO}_2$  towers
  - Suppress the external gamma background and reject surface radioactivity from the crystals that face the active shield
  - The background reduction will be achieved through anti-coincidence between the veto and  $\text{Li}_2\text{MoO}_4/\text{TeO}_2$
- The shield will be composed of BGO scintillator (initially  $\text{ZnWO}_4$  was chosen):
- Each bar (in fact two bars on top of each other) will be read by two light detectors (with Neganov-Luke effect for signal amplification)
- A reflecting material will be added on the lateral side of the veto to increase light collection in LDs
- On the internal side of veto, facing the crystals, a material will be added that should **not** be an  $\alpha$  stopper and that should prevent scintillation light from BGO to reach  $\text{Li}_2\text{MoO}_4$  and  $\text{TeO}_2$  LDs (Al, Au , ...?)



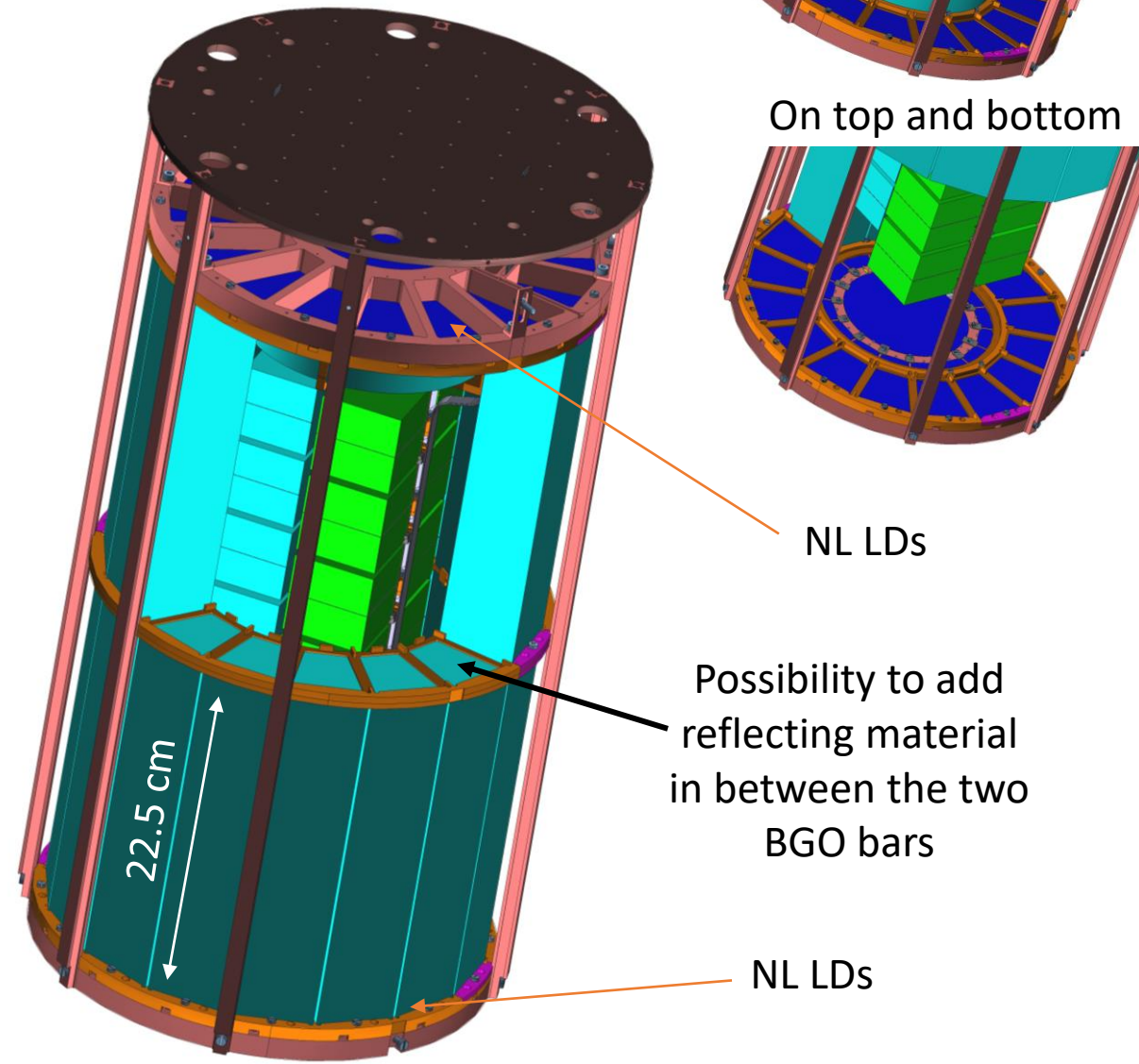
# BGO internal $^{207}\text{Bi}$ contamination

- BGO can show a significant internal  $^{207}\text{Bi}$  contamination up to 300mBq/kg
- A careful selection of the raw materials for growing BGO is ongoing from different providers worldwide. Different samples are being screened to select the ones with lower  $^{207}\text{Bi}$  contamination (some samples have shown good radiopurity)
- **MC simulation is ongoing to see the effect of  $^{207}\text{Bi}$  contamination on the experiment**

# The veto design



$\text{Li}_2\text{MoO}_4$  and  $\text{TeO}_2$  towers



On top and bottom

NL LDs

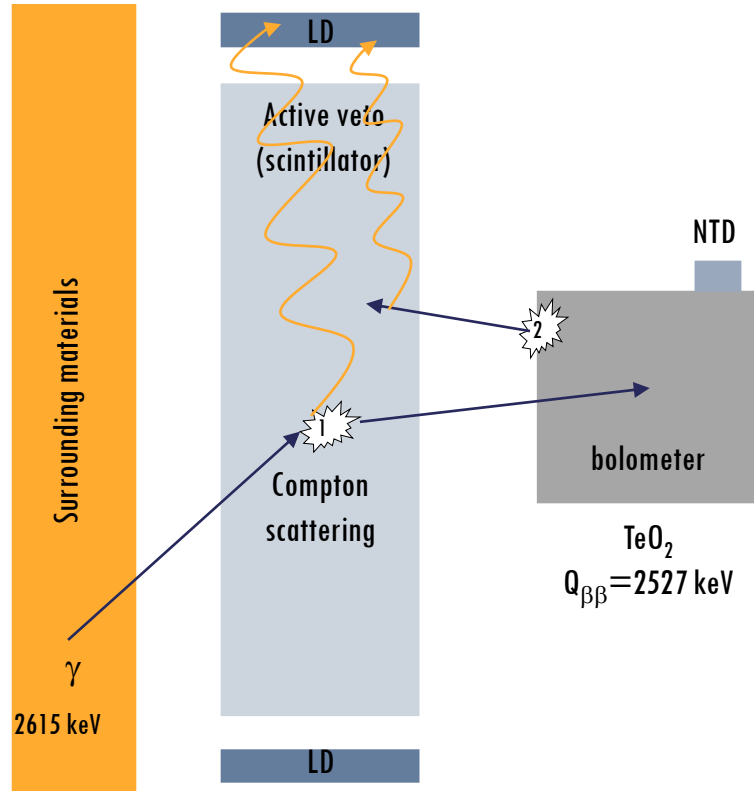
Possibility to add reflecting material in between the two BGO bars

NL LDs



# How does the veto work

The crystals on the periphery will be exposed directly to the veto



1

- If a 2615 keV  $\gamma$  deposits a small amount of energy in the surrounding material ( $\sim 80 \text{ keV}$ ) and the rest in  $\text{TeO}_2 \rightarrow$  background in ROI
- Thanks to the active veto and the LDs, these events can be rejected:
  - The energy deposition in the active veto will lead to scintillation light detected by the LD
  - Using anti-coincidence these events can be rejected from  $\text{TeO}_2$

2

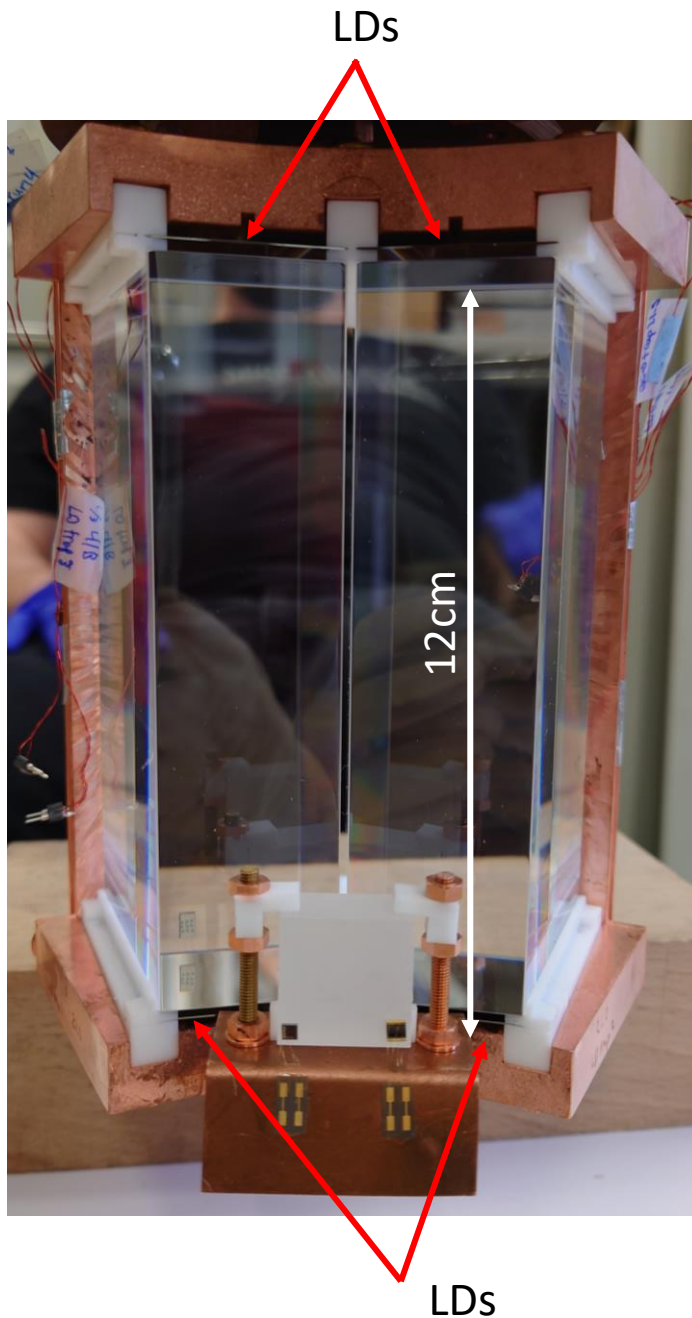
- Some surface contamination on the crystal can be dangerous if part of the energy escapes. This can also be rejected by anti-coincidence with the veto

# Requirements

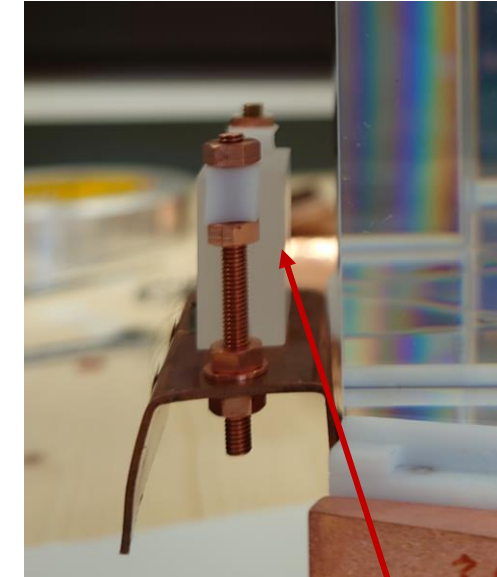
For this to work, a Neganov-Luke LD is needed:

- Higher signal to noise ratio in order to achieve a low enough energy threshold in LD
- 50 keV energy threshold in BGO is required
  - It corresponds to the dangerous small energy deposition of the 2615 keV line in BGO
- 50 keV in the BGO scintillator corresponds to a few keV in LD
  - Few keV in LD is achieved by taking into account the expected gain from NL effect (10-20)

# Prototype test



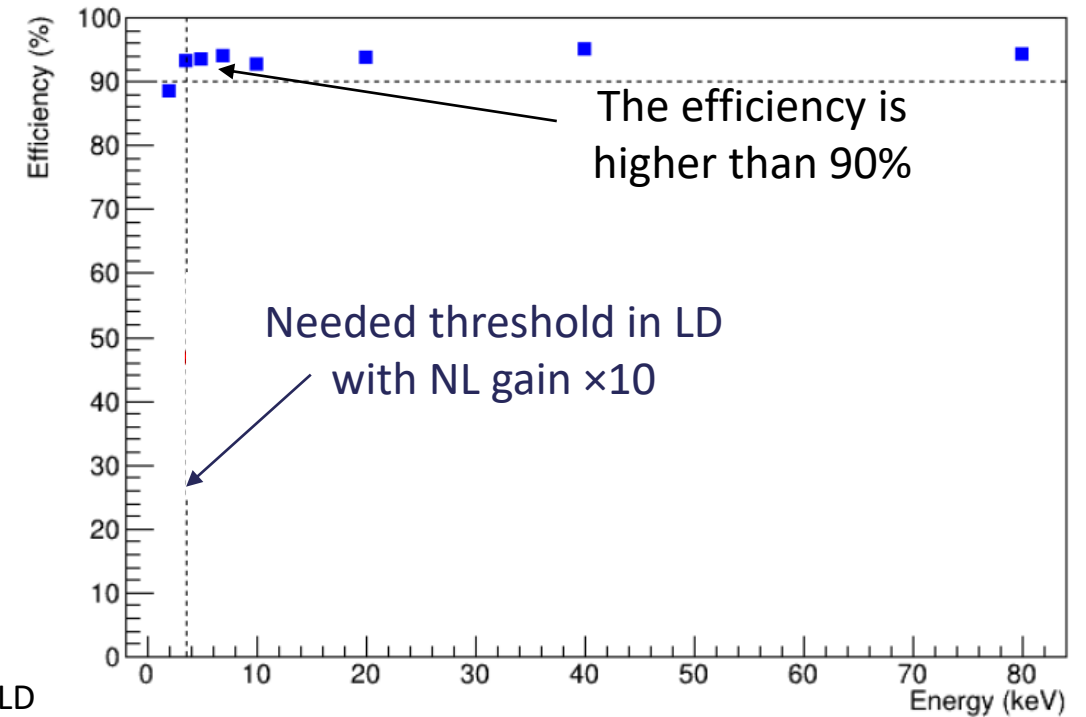
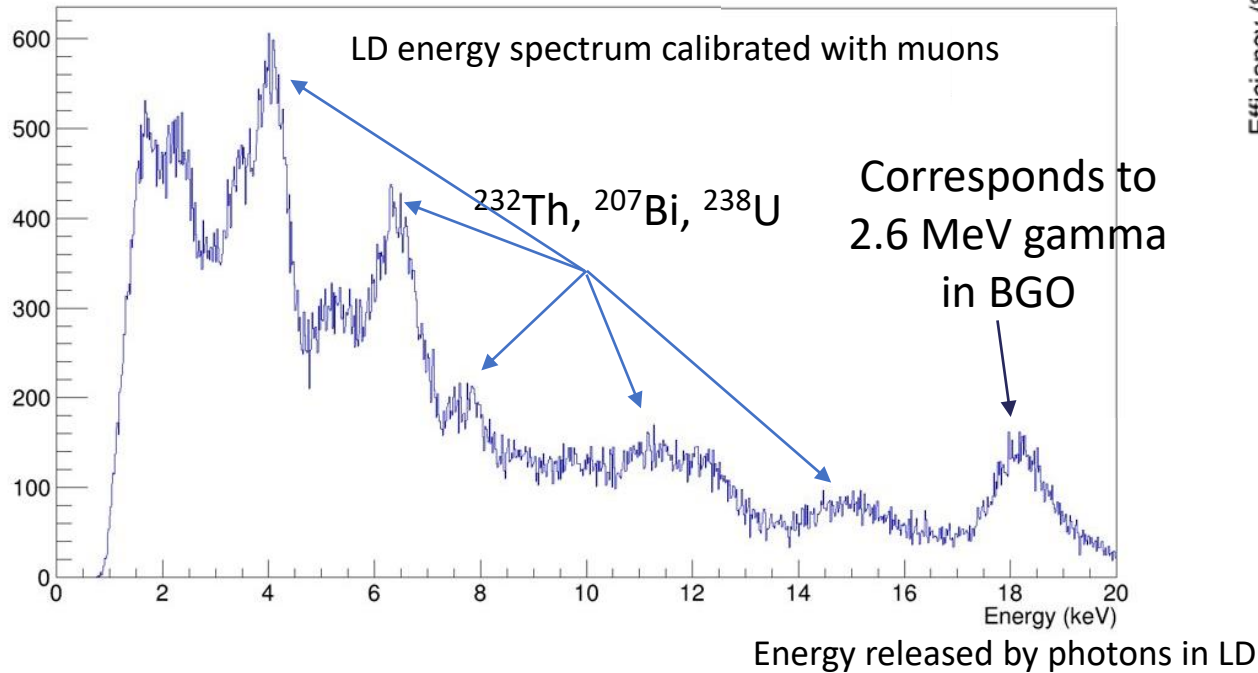
- 2 BGO crystals (~1.6kg each)
- 2 normal LDs facing each BGO
- $\text{TeO}_2$  crystal facing both BGOs
- Reflecting foil on the back
- The test was performed above-ground in a pulse-tube cryostat at IJCLab (Orsay)



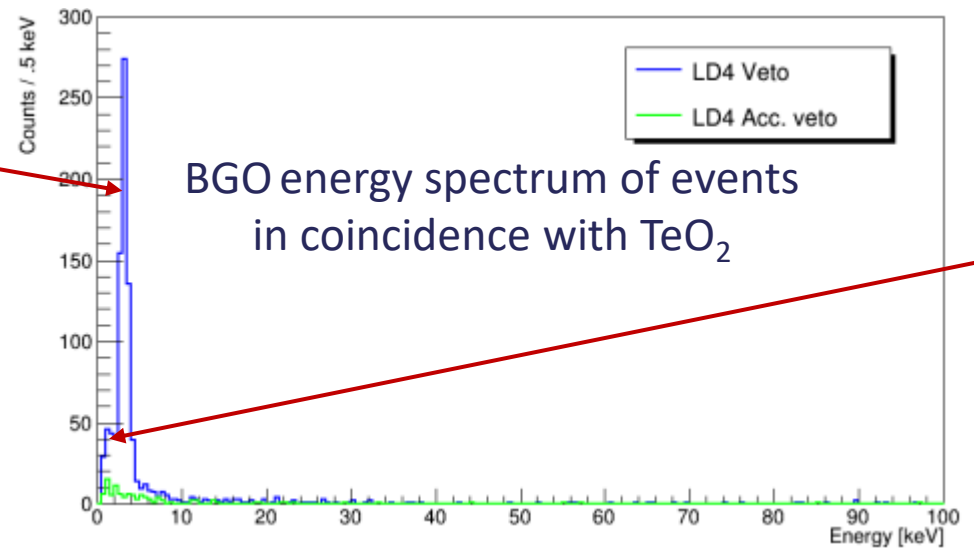
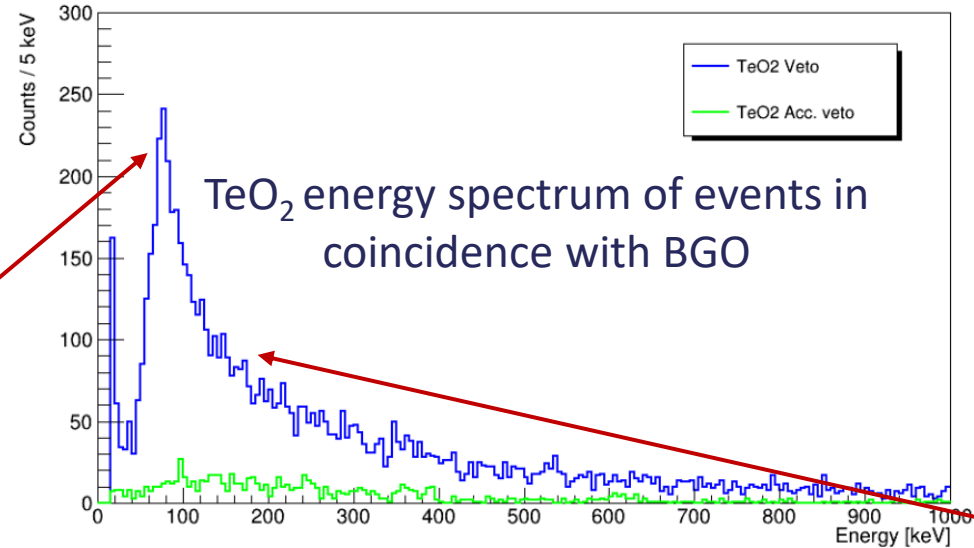
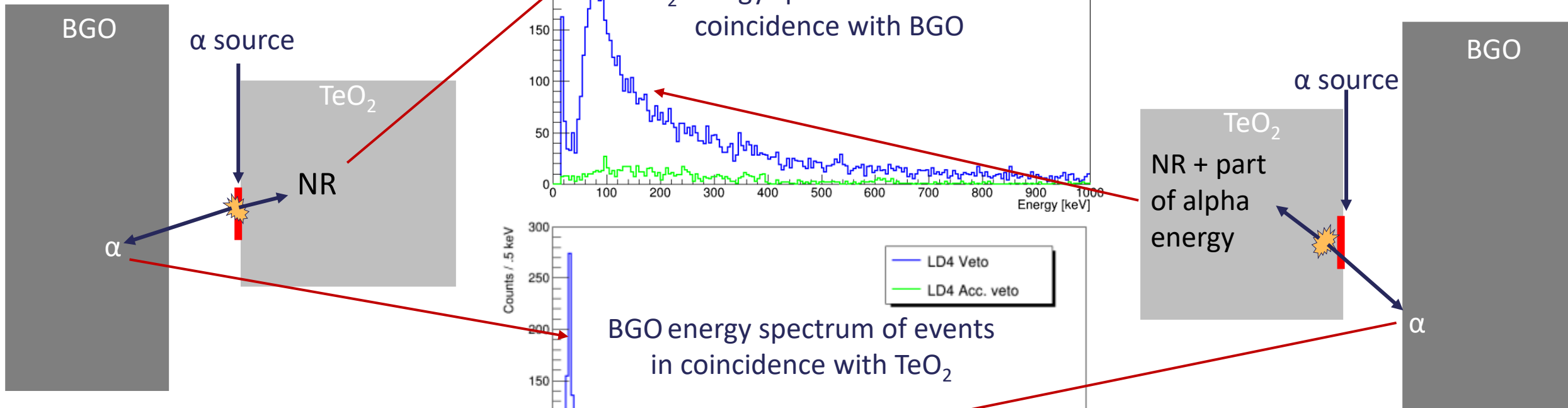
Uranium  $\alpha$  source deposited on  $\text{TeO}_2$  to produce surface contamination (at 4.2 and 4.8 MeV)

# Efficiency study at the expected threshold

- 1000 fake pulses at different energies were injected into the data to estimate the efficiency after data processing
- The required energy threshold for the veto scintillator should be around 50 keV, which corresponds to around 0.3-0.4 keV in LD when taking into account the light yield (LY) which is about 7 keV/MeV

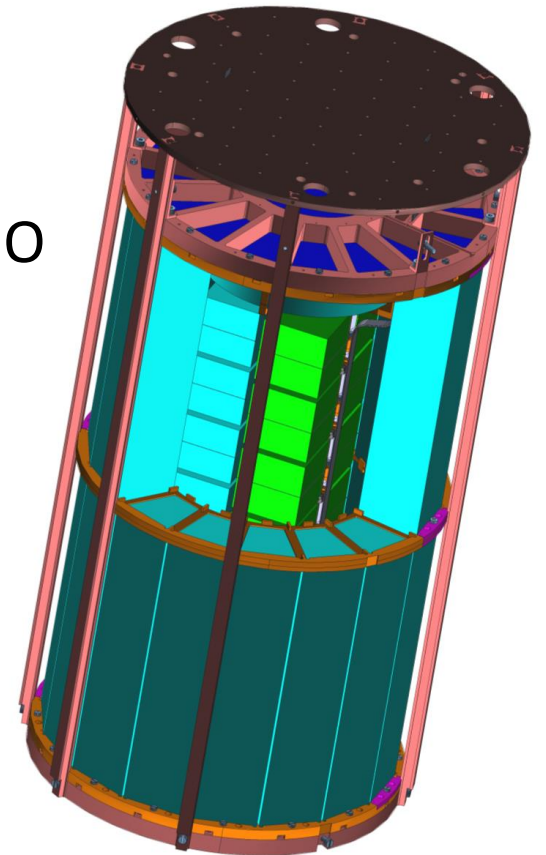


# Hunting the $\alpha$ source with coincidence



# Status of the veto

- A new veto design is being produced
- Studying the best reflection materials is still ongoing
  - A reflecting material will be added on the lateral side of the veto
  - A reflecting material will be added between the two bars (possibly)
  - On the internal side of veto, facing the crystals
- We have now hints of good providers for the raw materials to make BGO
- Four new trapezoidal bars to be produced by autumn 2023
- A section of the final veto design (two trapezoidal BGO bars) will be measured above-ground in a cryostat to test the mechanics/NL LD/reflecting materials in September 2023



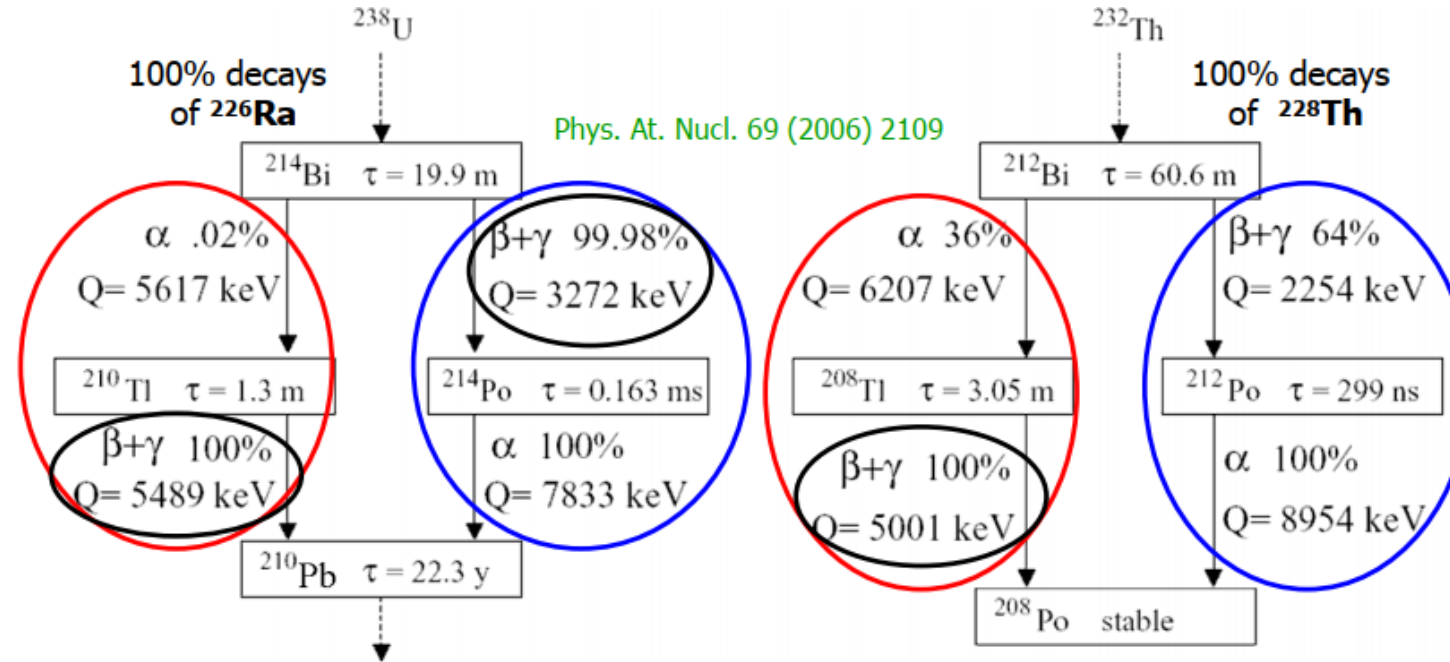
# Conclusions

- BINGO introduces innovative technology for background rejection
- For the first time, an active veto will be developed to surround the double-beta bolometers
- It will be used to reject external gamma radioactivity, in addition to surface radioactivity from the crystals facing the veto
- Prototype tests and MC simulation are still ongoing before moving to the final full geometry

backups

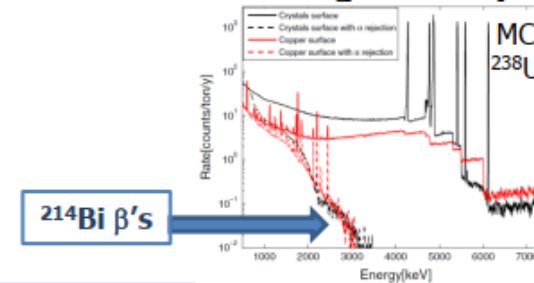
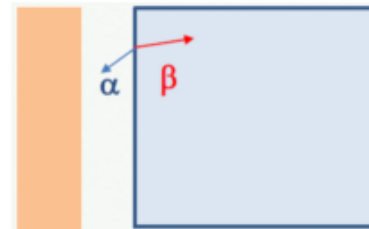
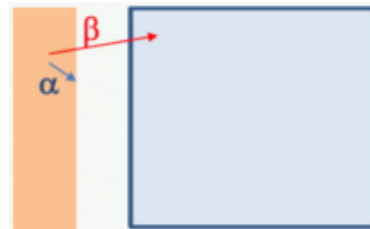


# Sources of background



## For DBD bolometers with particle ID:

- **BiPo events** (mixed  $\beta + \alpha$  decays with a total  $E > \sim 8$  MeV) **negligible contribute to ROI**
- **Delayed events of  $^{210,208}\text{Tl}$   $\beta$ 's can be rejected by an off-line gate after  $^{214,212}\text{Bi}$   $\alpha$ 's**
- **$\beta$ 's of  $^{212,214}\text{Bi}$  subchains ( $Q_\beta > Q_{\beta\beta}$ ) detected without  $\alpha$ 's contribute significantly to ROI**



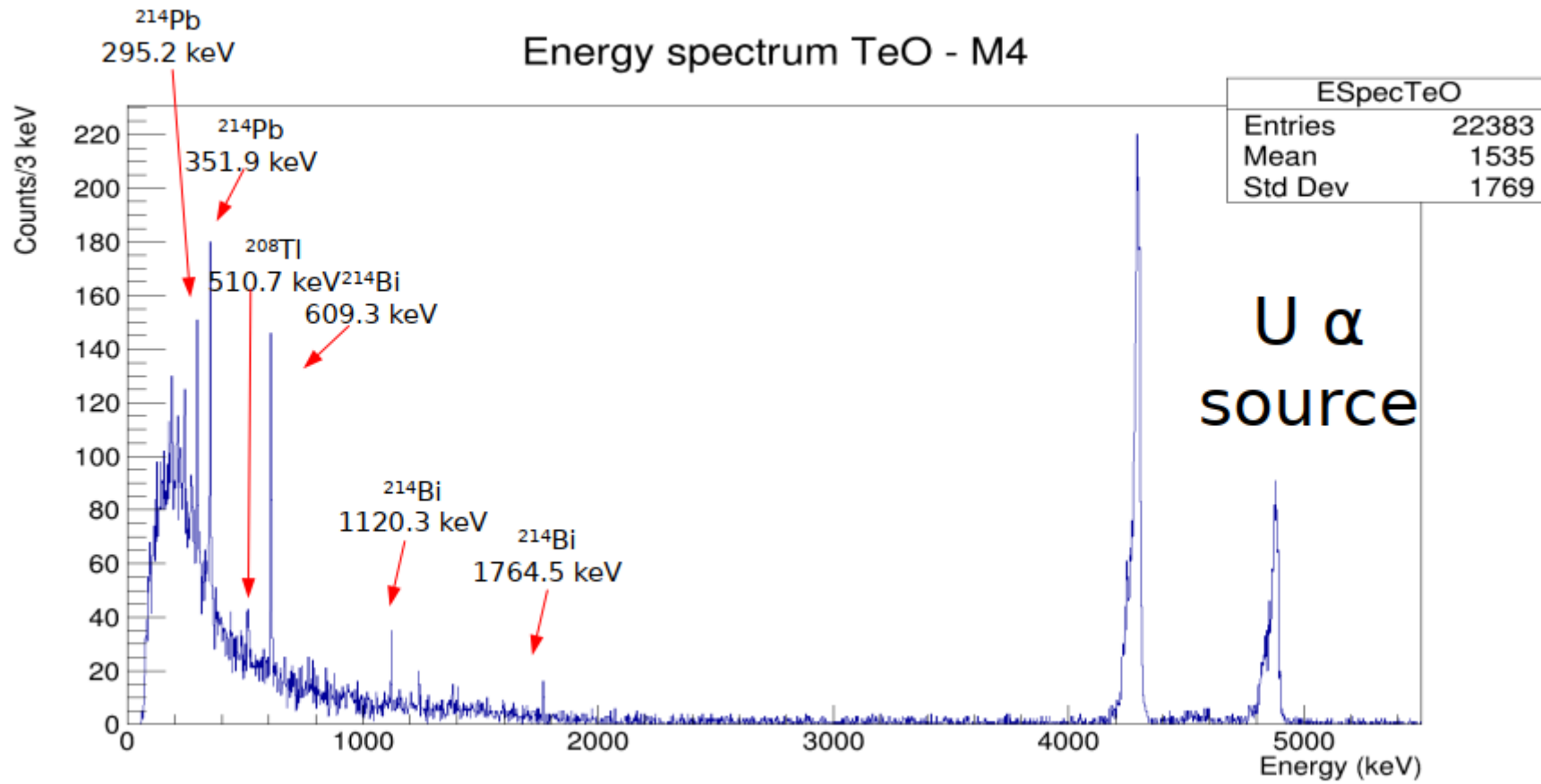
# $^{207}\text{Bi}$ contamination

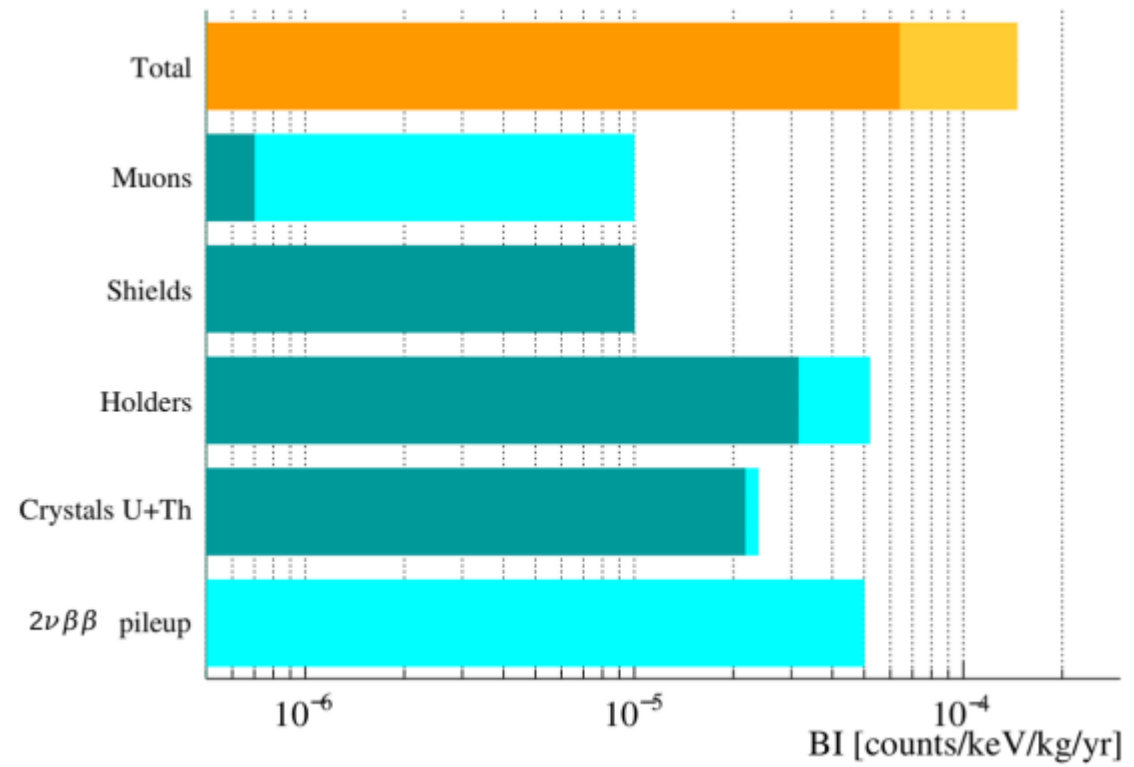
- With 328 mBq/kg,  $M_{\text{VETO}} = 115$  kg and assuming a 5 ms coincidence time window, dead time is  $\sim 17\%$
- Desirable to reach  $< 100$  mBq/kg  $\rightarrow$  deadtime is  $\sim 6\%$

Supplier6	Type	Purity [%]	Mass [g]	Activity [mBq/kg]	Activity [mBq/kg BGO]	803 keV
SICCAS (CN)	BGO		301	$173 \pm 16$	$173 \pm 16$	N
(RU)	BGO		301	$68 \pm 11$	$68 \pm 11$	N
Alfa Aesar (DE)*	$\text{Bi}_2\text{O}_3$	99.999	212	$< 21$ (95% CL)	$< 16$ (95% CL)	N
Santech (CN)	$\text{Bi}_2\text{O}_3$	99.990	206	$23 \pm 6$	$17 \pm 5$	Y
Zhuzhou (CN)	Bi	99.999	460	$37 \pm 5$	$25 \pm 3$	N

\*@ LSM for screening

# TeO<sub>2</sub> energy spectrum





# Bolometric compounds choices

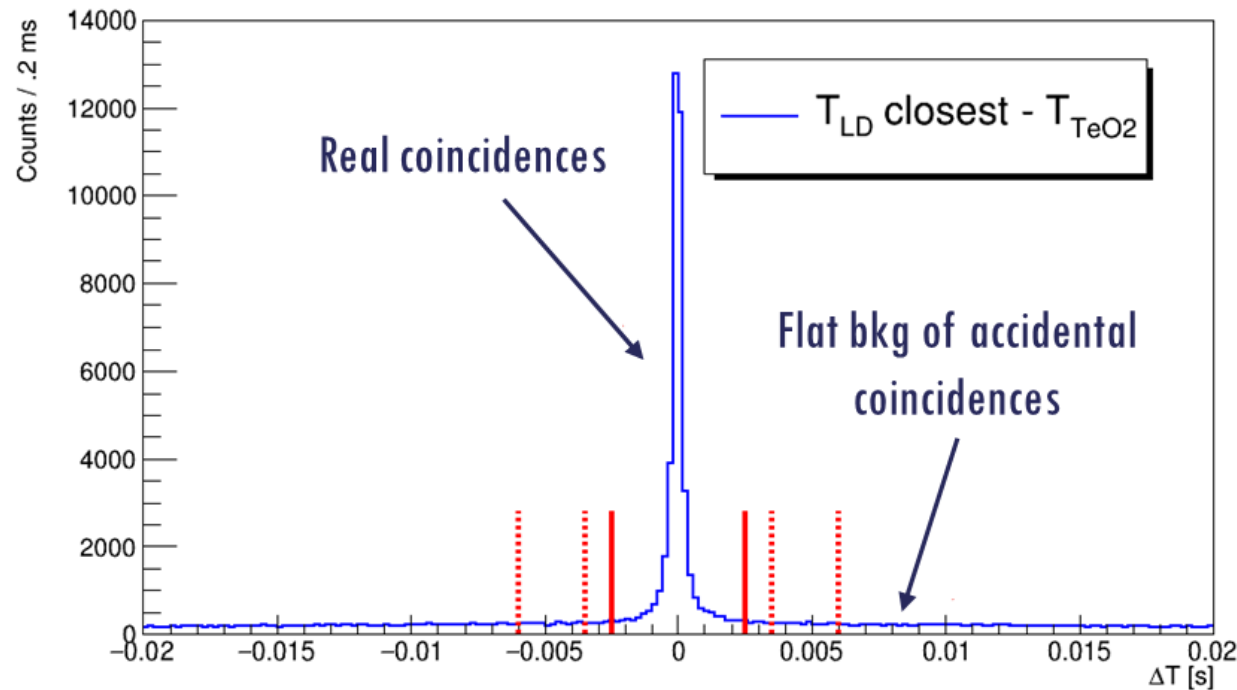


- Embeds  $^{100}\text{Mo}$  with a  $Q_{\beta\beta}$  at 3034 keV
- This crystal was validated by the CUPID-Mo demonstrator
  - Excellent energy resolution
  - High internal radio-purity
  - Easiness in crystallization
- High rate of  $2\nu 2\beta \rightarrow$  background in the region of interest (ROI) due to  $2\nu 2\beta$  random coincidences



- Embeds  $^{130}\text{Te}$  with a  $Q_{\beta\beta}$  at 2527 keV
- This crystal was validated by the CUORE experiment
  - Excellent energy resolution
  - High internal radio-purity
  - Easiness in crystallization
- $Q_{\beta\beta}$  below the end line (at 2615 keV line of  $^{208}\text{Tl}$ ) of natural gamma radioactivity
- Very poor scintillator  $\rightarrow$  no alpha background rejection

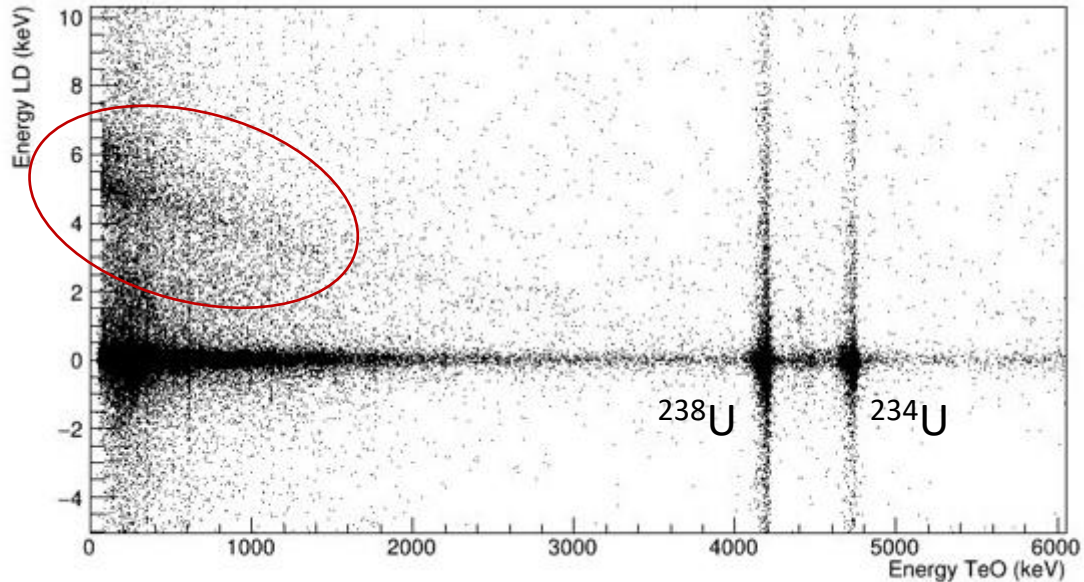
# Hunting the $\alpha$ source with coincidence



- Events in  $\text{TeO}_2$  are rejected if an event is found in a time window of 5ms in the light detectors.
- Accidental coincidences distribution is determined with the regions in red dashed lines since it should be the same under the peak

# Hunting the $\alpha$ source with coincidence

Coincidences between  $\text{TeO}_2$  and LD  
(meaning BGO)



- The marked events are alphas (with shared energy in  $\text{TeO}_2$ ) in LD
- At 0 keV in  $\text{TeO}_2$  we have a full alpha absorption in BGO (LD)
- At higher energies in  $\text{TeO}_2$  the alpha energy is shared between  $\text{TeO}_2$  and the LD
- The extrapolation of this population to higher energies in the  $\text{TeO}_2$  leads to full alpha absorption in  $\text{TeO}_2$