

# Spinodal decomposition in a rapidly expanding fluid

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# QCD phase diagram



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# **Trajectories of heavy-ion collisions**



Chun Shen, Quark Matter 2018

- Realistic heavy-ion collisions do not follow a clean trajectory on the phase diagram
- The phase space volume passing near the critical point may not be large
- It makes sense to look for signatures of the first order phase transition as a larger system volume could pass through that

# **Spinodal separation**



Figure from Pathria and Beale, Statistical Mechanics

- The region between A and B is unstable (spinodal region). It is energetically favorable for phases to separate
- The region from 3 to A and B to 1 is metastable. This is where nucleation dominates
- If nucleation rate is small, spinodal separation is the dominant mechanism of phase transition

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• Consider the Helmholtz free energy functional at temperature T

$$F\{n(\mathbf{x},t)\} = \int d^3x \left[\frac{1}{2}K(\vec{\nabla}n)^2 + f(T,n)\right]$$
$$= \int d^3x \tilde{f}(T,n)$$

• For small changes in density

$$F\{n+\delta n\} - F\{n\} = \int d^3 x \left[ K \vec{\nabla} \delta n \cdot \vec{\nabla} n + \frac{\partial f}{\partial n} \delta n \right]$$
$$= \int d^3 x \left[ K \vec{\nabla} \cdot (\delta n \vec{\nabla} n) - K \delta n \nabla^2 n + \frac{\partial f}{\partial n} \delta n \right]$$

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• The chemical potential is

$$\tilde{\mu} = \frac{\delta F}{\delta n} = \frac{\partial f}{\partial n} - K \nabla^2 n = \mu - K \nabla^2 n$$

• The isotropic pressure is

$$\tilde{P} = n\mu - [f + \frac{1}{2}K(\vec{\nabla}n)^2]$$
  
=  $n\frac{\partial f}{\partial n} - f - Kn\nabla^2 n - \frac{1}{2}K(\vec{\nabla}n)^2$   
=  $P - Kn\nabla^2 n - \frac{1}{2}K(\vec{\nabla}n)^2$ 

• Assuming negligible dependence of *K* on temperature the entropy density is as usual

$$s = -\frac{\partial f}{\partial T}$$

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• Local energy density

$$\tilde{\epsilon} = f - T \frac{\partial f}{\partial T} + \frac{1}{2} \mathcal{K}(\vec{\nabla}n)^2 = \epsilon + \frac{1}{2} \mathcal{K}(\vec{\nabla}n)^2$$

• The local enthalpy is

$$\tilde{w} = \tilde{\epsilon} + \tilde{P} = n \frac{\partial f}{\partial n} - T \frac{\partial f}{\partial T} - Kn \nabla^2 n$$
$$= \tilde{\mu}n + Ts = w - Kn \nabla^2 n$$

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• Keeping the baryon number fixed, let's minimize the Helmholtz free energy

$$I = \int d^3x \left[ \tilde{f}(T, n) - \lambda n \right]$$

• The resulting Euler-Lagrange equation is

$$\partial_i \frac{\partial(\tilde{f} - \lambda n)}{\partial(\partial_i n)} - \frac{\partial(\tilde{f} - \lambda n)}{\partial n} = 0$$

• The Lagrange multiplier is

$$\lambda = \mu - \mathbf{K} \nabla^2 \mathbf{n} = \tilde{\mu} = ext{constant}$$

• If n(x) solves the above equation, then the equilibrium surface free energy is

$$\sigma = K \int_{-\infty}^{\infty} dx \left(\frac{dn}{dx}\right)^2$$

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### Metastable and unstable states

• For  $T < T_c$  and  $n_G \le n \le n_L$  we parameterize the pressure in terms of the density as

$$P_{in}(n) = P_X(T) + \sum_{i=1}^{4} c_i (n - n_G)^i$$

• The chemical potential is

$$\mu_{in}(n) = \mu_X(T) + d_0 \ln(n/n_G) + \sum_{i=1}^3 d_i (n - n_G)^i$$

- We get constraints from  $P_{in}(n_L) = P_X(T)$  and the continuity of first derivates of P at  $n_L$  and  $n_G$
- Further constraints are obtained by thermodynamics relations between P, n and  $\mu$  and by  $\mu_{in}(n_L) = \mu_X(T)$

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### Metastable and unstable states



### **Correlation length and surface energy**

• Near critical density  $n_c$ , we can reasonably assume  $n_L - n_c = n_c - n_G = \Delta n/2$ . For  $\delta n = n - n_c$  with  $-\Delta n/2 \le \delta n \le \Delta n/2$ . Then

$$\mu - \mu_X \approx d_3(\delta n + \Delta n/2)(\delta n + \Delta n/2)\delta n$$

• The planar surface equation is

$$\mathcal{K}rac{\partial^2 \delta \mathbf{n}}{\partial x^2} = \mu - \mu_X$$

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with solution

$$\delta n = \frac{\Delta n}{2} \tanh(x/2\zeta)$$

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### **Correlation length and surface energy**

• The correlation length  $\zeta$  is

$$\zeta^2 = \frac{2\mathbf{K}}{\alpha\Delta\mathbf{n}^2}$$

• The surface free energy is

$$\sigma = K \int_{-\infty}^{\infty} dx \left(\frac{dn}{dx}\right)^2 = \frac{K \Delta n^2}{6\zeta}$$

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### **Correlation length and surface energy**



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# **Equation of state**

• Background EOS is obtained by matching pQCD EOS in QGP phase to HRG EOS using a smooth function. It is matched to Lattice calculations at  $\mu_B = 0$ 

[M. Albright, J. Kapusta and C. Young, Phys. Rev. C 90, (2014)]

• The critical point is embedded on it at  $T_c=130~{\rm MeV}$  and  $\mu_c=450~{\rm MeV}$  J. Kapusta, T. Welle and C. Plumberg, Phys. Rev. C 106, (2022)



# **Hydrodynamics**

• The relativistic energy momentum tensor is

$$T^{\mu\nu} = \tilde{P}(u^{\mu}u^{\nu} - g^{\mu\nu}) + \tilde{\epsilon}u^{\mu}u^{\nu} + K(D^{\mu}n)(D^{\nu}n)$$

where the gradient is orthogonal to velocity

$$D^{\mu}n \equiv \partial^{\mu}n - u^{\mu}u^{lpha}\partial_{lpha}n$$

• The local thermodynamic quantities are

$$\begin{split} \tilde{\mu} &= \mu + KD^2 n \\ \tilde{P} &= P + KnD^2 n + \frac{1}{2}K(D^{\mu}n)(D_{\mu}n) \\ \tilde{\epsilon} &= \epsilon - \frac{1}{2}K(D^{\mu}n)(D_{\mu}n) \\ \tilde{w} &= \tilde{P} + \tilde{\epsilon} = Ts + \tilde{\mu}n = w + KnD^2n \end{split}$$

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### 1+1 D flow

We use a simple model of 1 + 1 D inviscid fluid with baryon dimension in Bjorken coordinates. In Landau frame, to  $1^{\rm st}$  order

$$J^{\mu}= extsf{nu}^{\mu}+\sigma_{B} extsf{TD}^{\mu}\left(rac{ ilde{\mu}}{ extsf{T}}
ight)$$

The equations to solve are

$$\frac{\partial \epsilon(\mathbf{n}, T)}{\partial \tau} + \frac{\mathbf{w}(\mathbf{n}, T)}{\tau} + \frac{\mathbf{K}}{\tau^2} \frac{\partial \mathbf{n}}{\partial \xi} \frac{\partial^2 \mathbf{n}}{\partial \tau \partial \xi} - \frac{\mathbf{K}}{\tau^3} \mathbf{n} \frac{\partial^2 \mathbf{n}}{\partial \xi^2} = 0$$
$$\frac{\partial}{\partial \tau} (\tau \mathbf{n}) - \frac{\sigma_B T}{\tau} \frac{\partial^2}{\partial \xi^2} \left(\frac{\tilde{\mu}}{T}\right) - \frac{1}{\tau} \frac{\partial}{\partial \xi} (\sigma_B T) \frac{\partial}{\partial \xi} \left(\frac{\tilde{\mu}}{T}\right) = 0$$
$$\tilde{\mu} = \mu(\mathbf{n}, T) - \frac{\mathbf{K}}{\tau^2} \frac{\partial^2 \mathbf{n}}{\partial \xi^2}$$

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# Setting up the system

- Due to fourth order derivatives, it is difficult to do this numerically with any level of generality with finite difference methods
- Dedicated finite element routines exist for many systems at equilibrium
- Here, we set up boost invariant energy density and a sinusoidal baryon density such that the lowest point touches the phase transition curve from above



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# **Evolution across the phase boundary**

- We chose a relatively small *K* so that the change in *n* is negligible and the gradients in *n* are determined by the initial sinusoidal distribution
- We get the phase separation, but eventually all QCD matter is in hadronic phase



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# **Energy density**



- The two different curves correspond to K = 0 and  $K = 5 \times 10^{-5} \text{ MeV}^{-4}$ . The energy density increases or decreases in the coexistence phase for different K
- This energy change will be balanced by energy flow in the hadronic phase

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# Temperature



QGP Hadron Gas Coexistence Region 3 4

- Temperature evolution for different K
- The QGP temperature is lower that the hadron gas temperature as μ<sub>B</sub> has opposite sign

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# **Baryon chemical potential**





• Specifics of temperature and chemical potential changes are sensitive to the exact choice of the EOS

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- It has been postulated that the QCD phase diagram has a first order phase transition curve at high baryon chemical potential
- If such a curve exists, it is very likely that a large volume of QCD matter at some collision energy will undergo phase transition via spinodal decomposition
- We present the equations of relativistic hydrodynamics with the phase transition and solve it for a simple system
- Realistically simulating phase transitions for a rapidly expanding fluid in HIC is challenging and will likely require generalizations of finite element techniques used to study chemical and condensed matter systems