

# CRITICAL POINT FLUCTUATIONS IN HEAVY-ION COLLISIONS WITHIN MOLECULAR DYNAMICS WITH EXPANSION

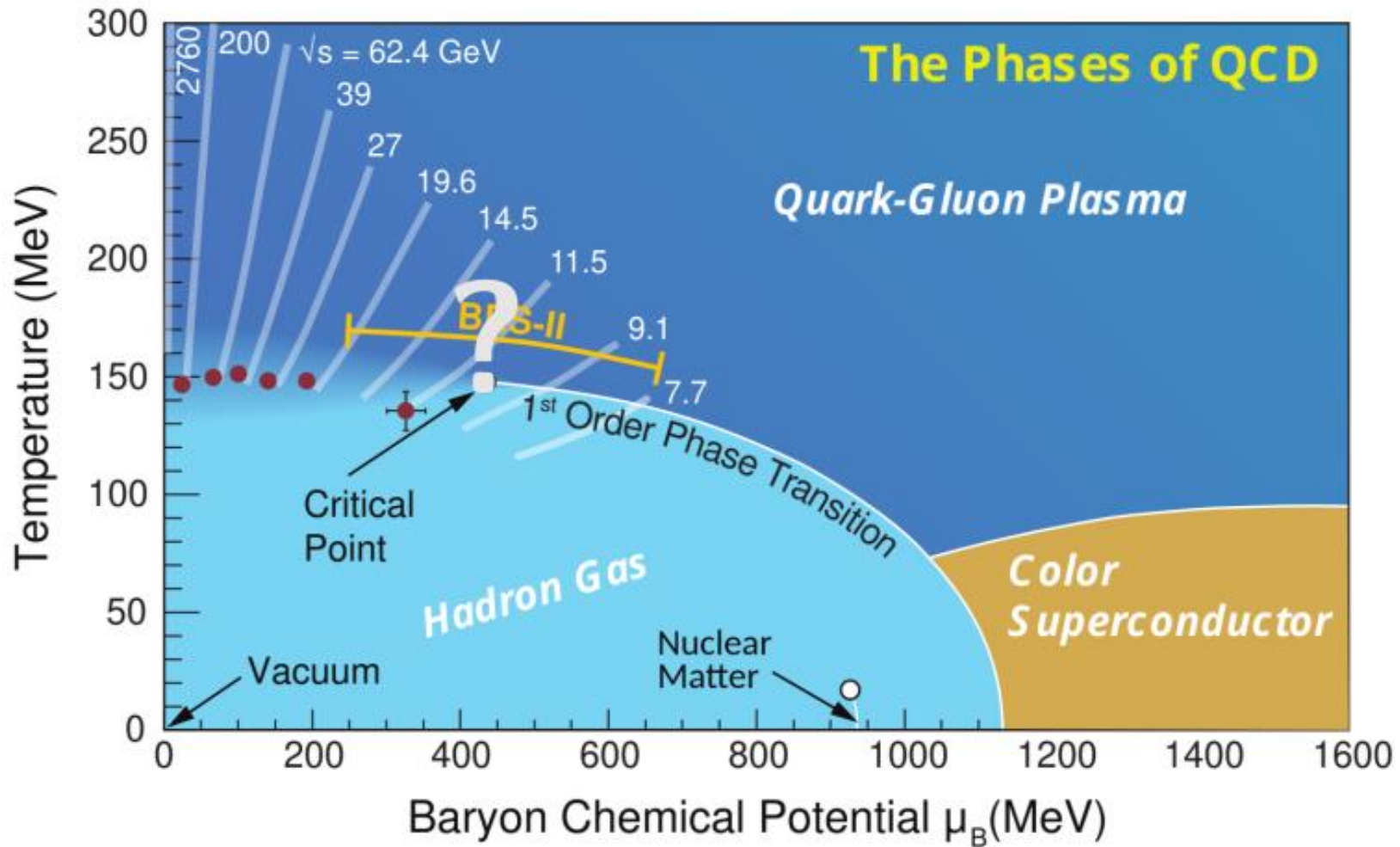


Volodymyr Kuznietsov

Based on: V.K., M.I. Gorenstein, V. Koch, V. Vovchenko, arXiv: 2404.00476 [nucl-th]

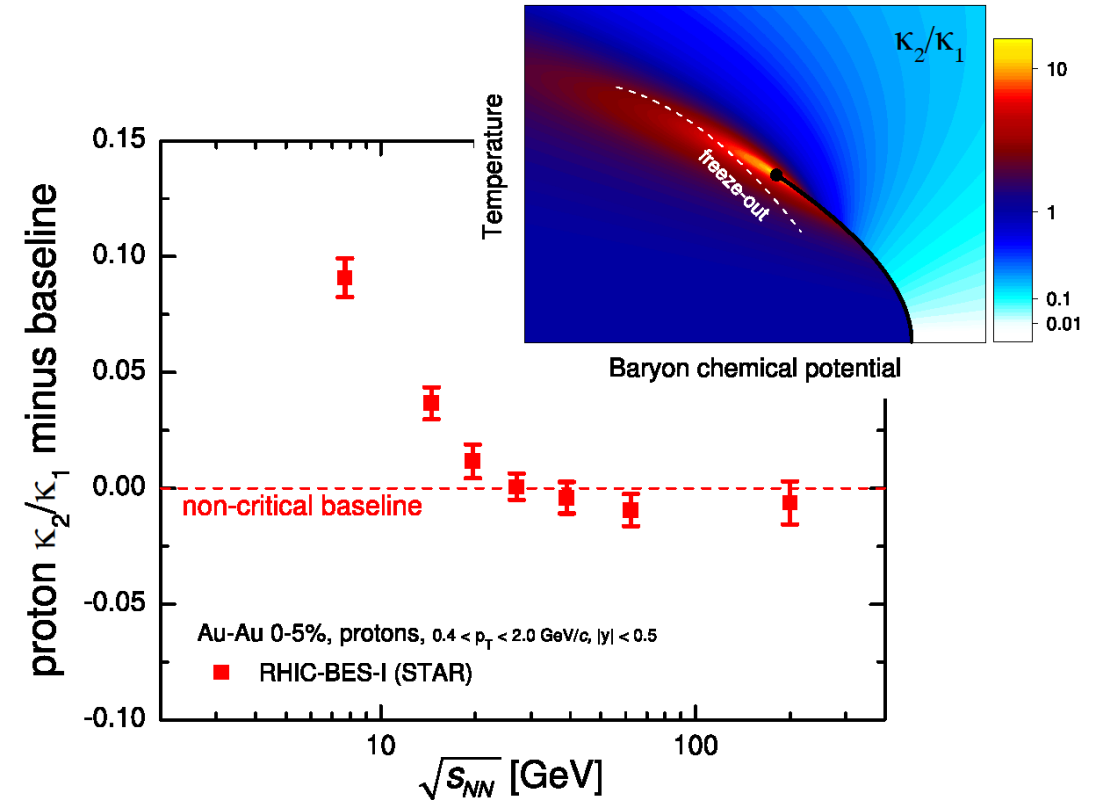
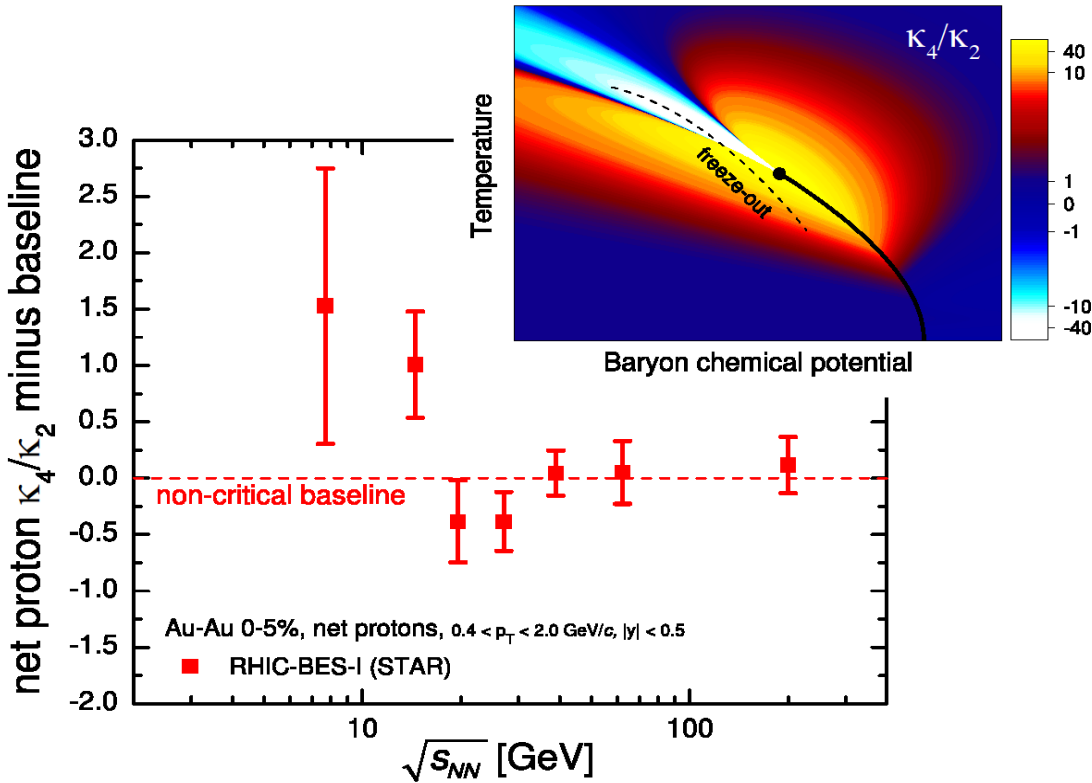


# QCD PHASE DIAGRAM



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# FLUCTUATIONS AS CP SIGNATURE



[1] Non-critical baseline: Vovchenko, Koch, Shen, PRC 105, 014904 (2022)

[2] STAR data (BES-I): J. Adam et al., PRL 126, 092301 (2021)

# CONNECTION TO THE EXPERIMENT

## Experiment

- Momentum space
- Expanding in vacuum
- Non-conserved particle numbers
- Inhomogenous
- Fluctuating volume

## Theory

- Coordinate and/or momentum space
- In contact with the heat bath
- Conserved charges
- Uniform
- Fixed volume

Here we study critical fluctuations in a microscopic approach  
**(MD)**

# 5 LENNARD-JONES POTENTIAL

The Lennard-Jones potential reads

$$V_{LJ} = 4\epsilon[(\sigma/r)^6 - (\sigma/r)^{12}]$$

In reduced dimensionless variables it can be rewritten as

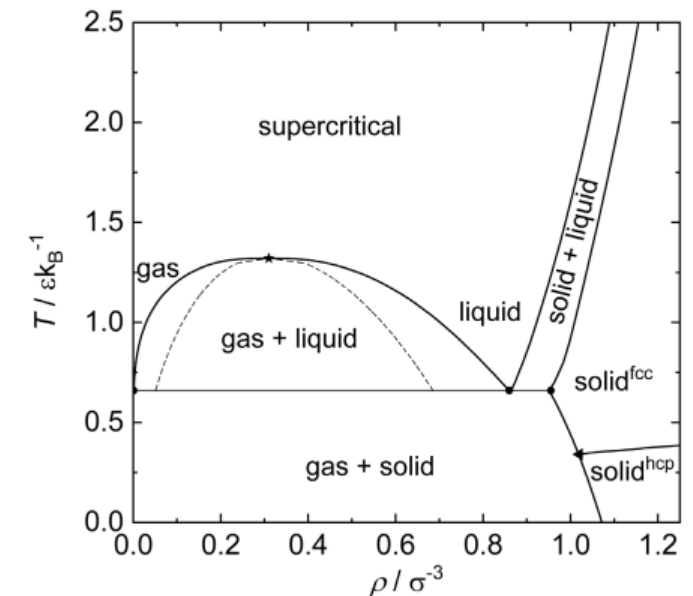
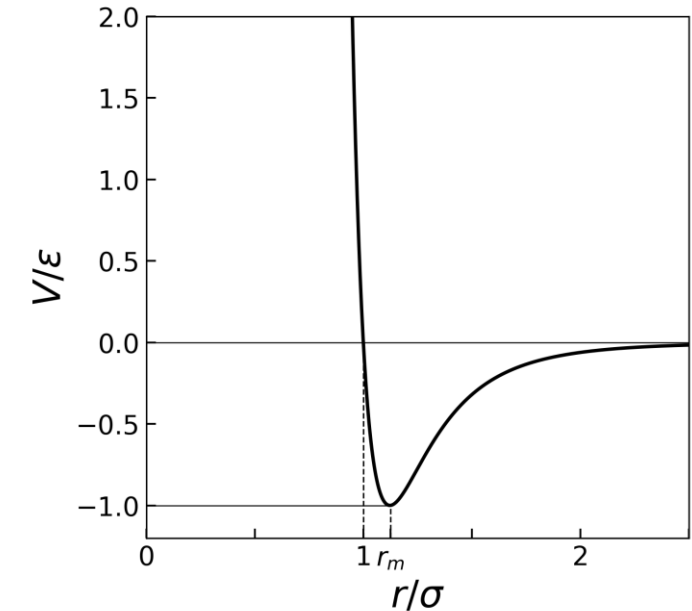
$$\tilde{V}_{LJ} = 4(\tilde{r}^{-6} - \tilde{r}^{-12})$$

where the reduced variables are use:

$$\tilde{r} = r/\sigma, \quad \tilde{t} = t\sqrt{\epsilon/(m\sigma^2)}, \quad \tilde{V}_{LJ} = V_{LJ}/\epsilon$$

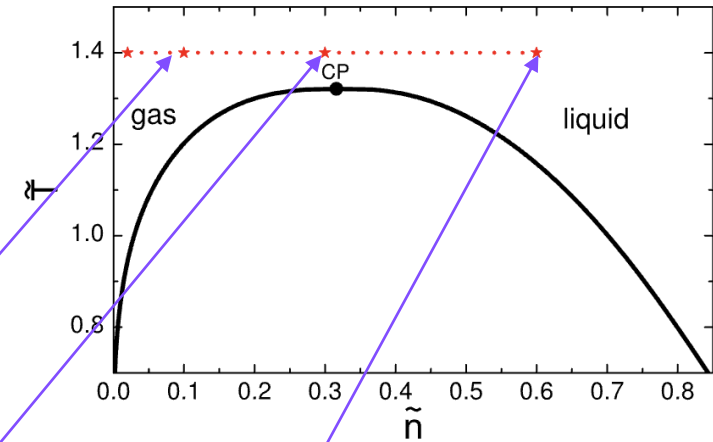
Advantages:

- Well-studied system
- The LJ fluid contains a critical point in the **3D-Ising universality class**, same as QCD critical point.
- Critical fluctuations are automatically projected on finite particle number statistics.



# SIMULATION SETUP

$$m \frac{d^2 \vec{r}_{i,j}}{d\tilde{t}^2} = -\vec{\nabla} V_{LJ}(\vec{r}_{i,j})$$



- Three points on the phase diagram,  $\tilde{n} \approx 0.3n_c$ ,  $\tilde{n} \approx 0.95n_c$ ,  $\tilde{n} \approx 1.9n_c$  ( $\tilde{T} \approx 1.9T_c$ )
- $N_{ev} = 32\,000$  events at each density
- Initialize each event with random initial coordinates and momenta
- Run each event for long time ( $\tilde{t} = 100$ ) write snapshots to file at regular time intervals
- Calculate observables as event-by-event (ensemble) or time average

The simulations are performed on PhysGPU cluster at UH. Code is available at:

<https://github.com/vlvovch/lennard-jones-cuda>

# 7 OBSERVABLES AND ERGODICITY

Time average

$$\langle A \rangle_\tau = \frac{1}{\tau} \int_{\tilde{t}_{eq}}^{\tilde{t}_{eq} + \tau} A(t) dt$$

versus ensemble average:

$$\langle A \rangle_M = \frac{1}{M} \sum_{i=0}^M A_i$$

Ergodic hypothesis:

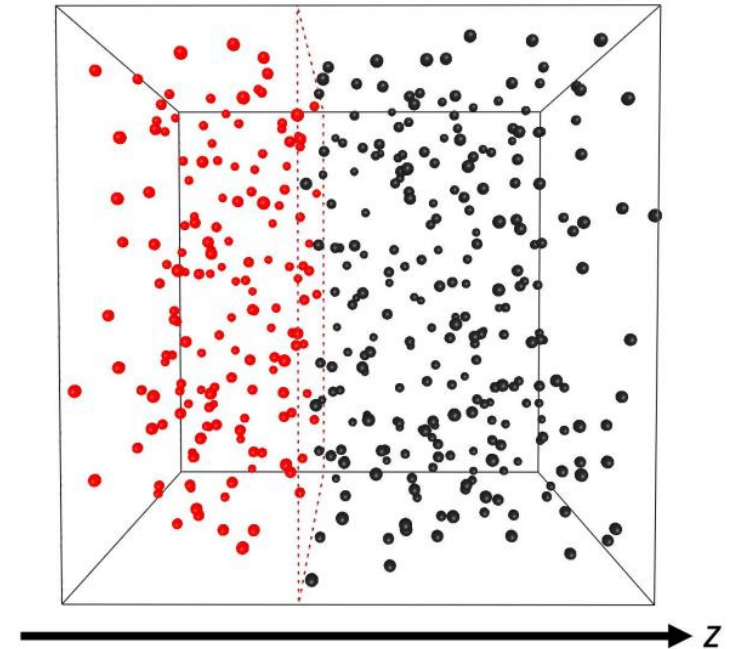
$$\lim_{M \rightarrow \infty} \langle A \rangle_M = \lim_{\tau \rightarrow \infty} \langle A \rangle_\tau$$

Namely:

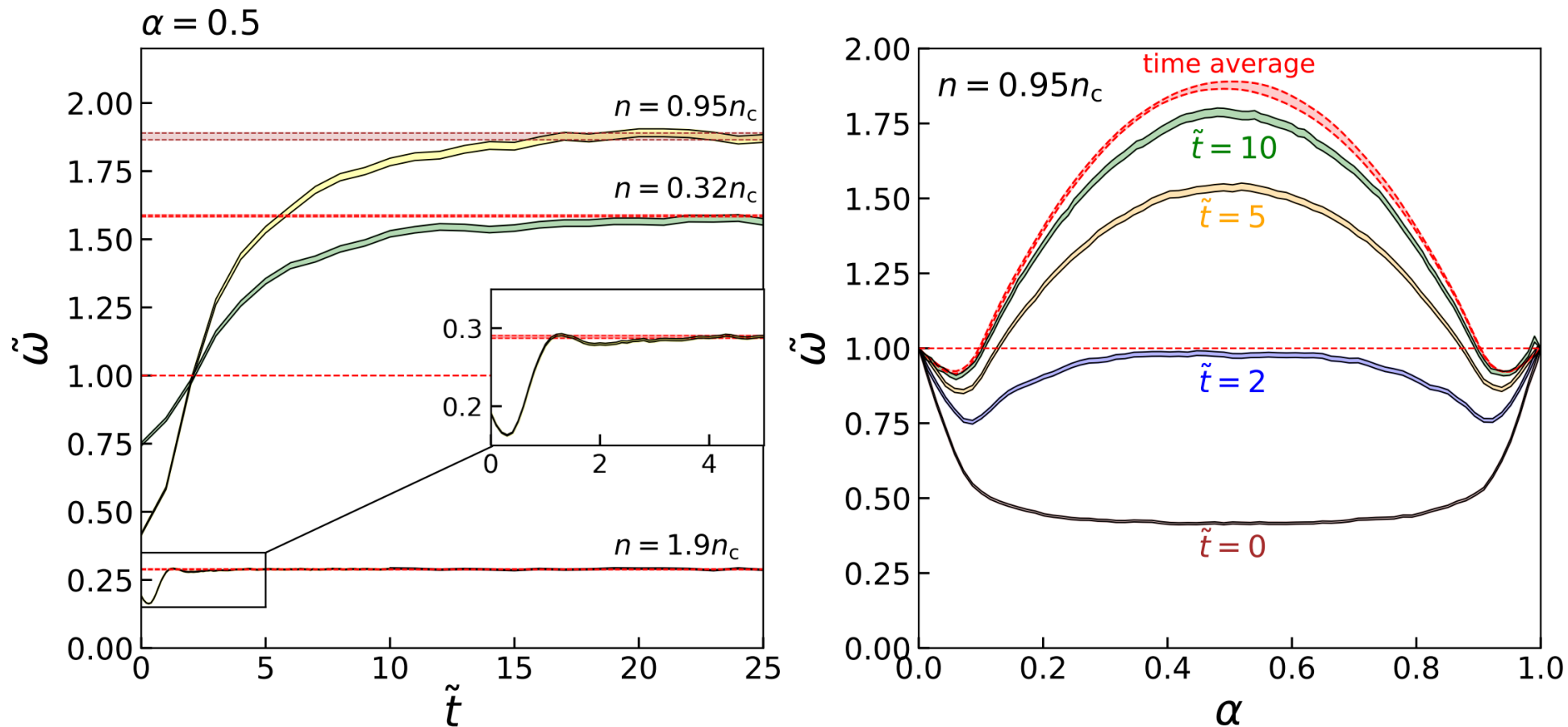
$$\langle N \rangle = \frac{1}{M} \sum_{i=0}^M N_i(\alpha), \quad \tilde{\omega} = \frac{1}{1 - \alpha} \frac{\langle N \rangle^2 - \langle N^2 \rangle}{\langle N \rangle}, \quad \alpha = \frac{\langle N \rangle}{N_{tot}}$$

V. A. Kuznietsov et al., PRC 105, 044903 (2022)

\* $1/(1 - \alpha)$  correction is related to the V. Vovchenko et al., Phys. Let. B, 2020



# 8 TIME VS ENSEMBLE AVERAGE, EQUILIBRATION



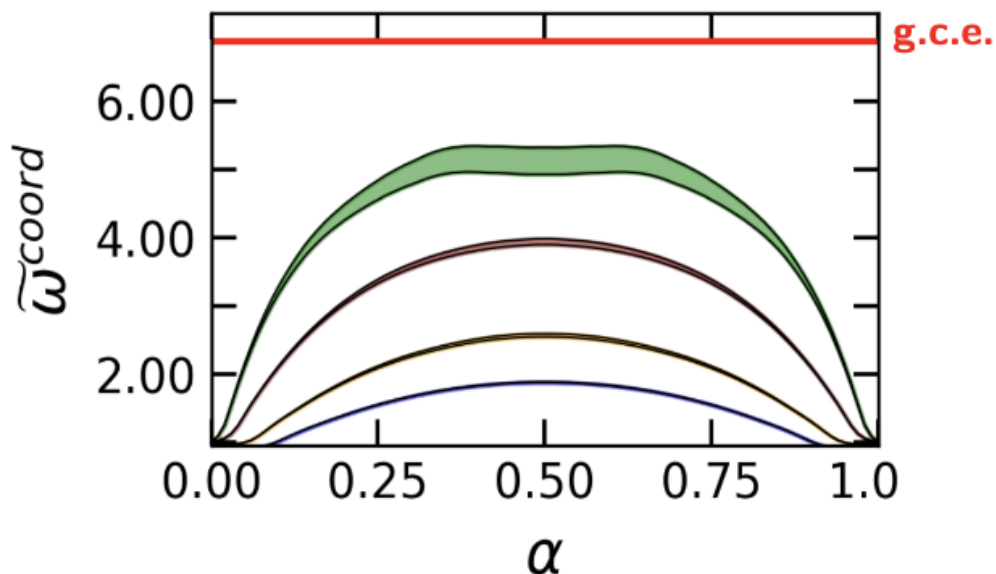
Ergodic hypothesis works



# PREVIOUS RESULTS: TIME AVERAGE IN BOX<sup>+</sup>

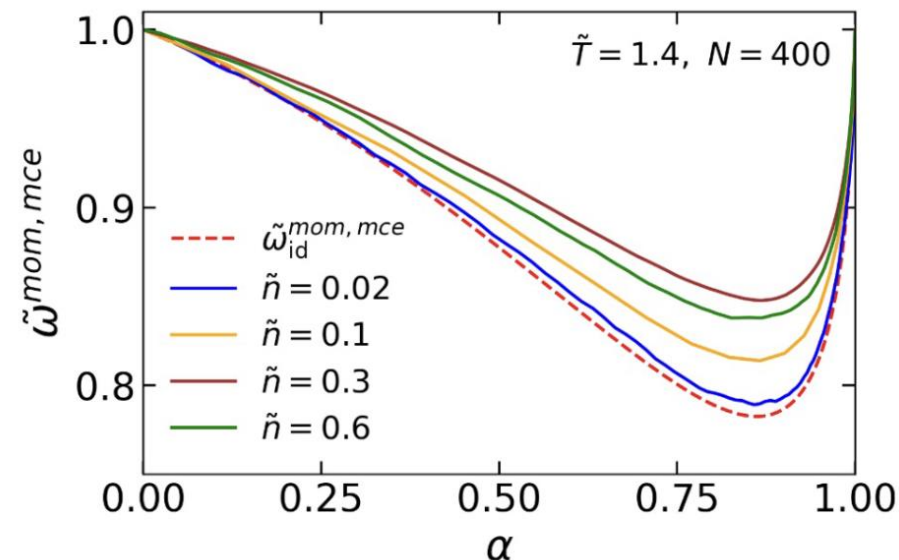
$$\tilde{\omega} = \frac{1}{1-\alpha} \frac{\langle N \rangle^2 - \langle N^2 \rangle}{\langle N \rangle}$$

Coordinate space



Large correlations in coordinate space

Momentum space

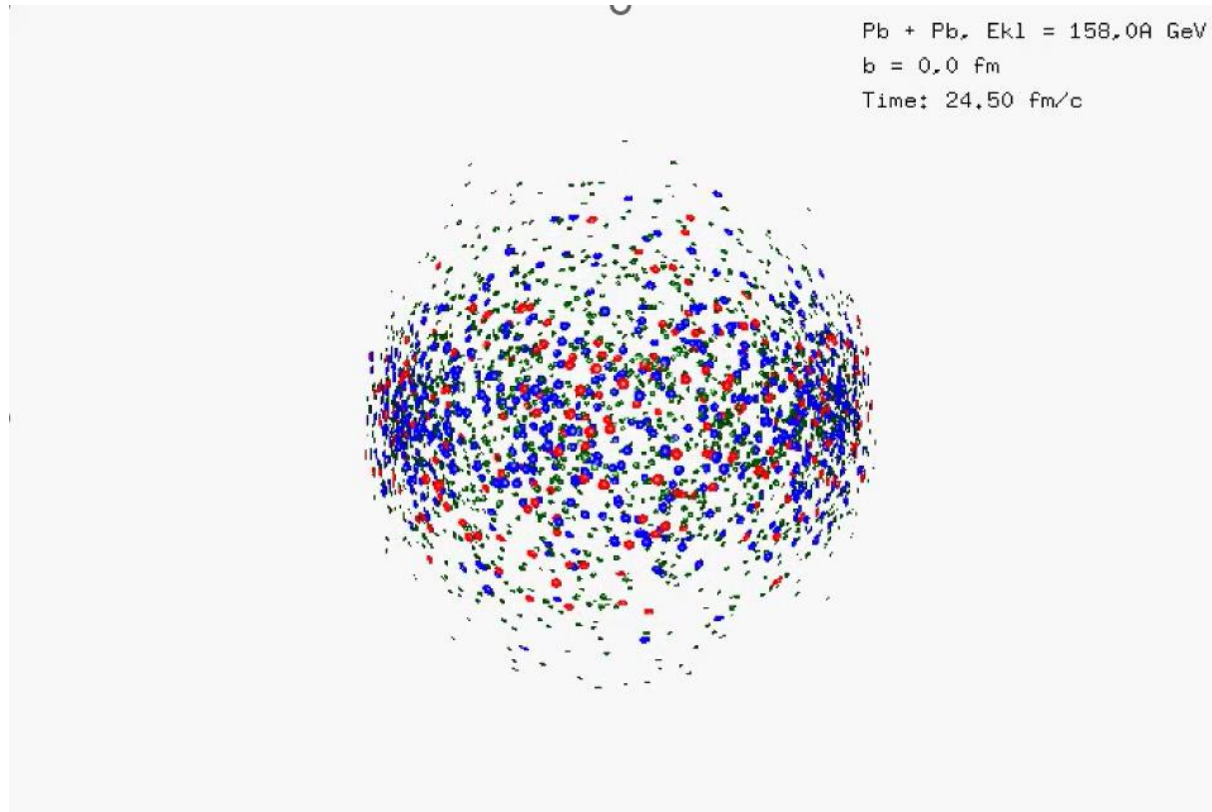


Signal disappears in momentum space

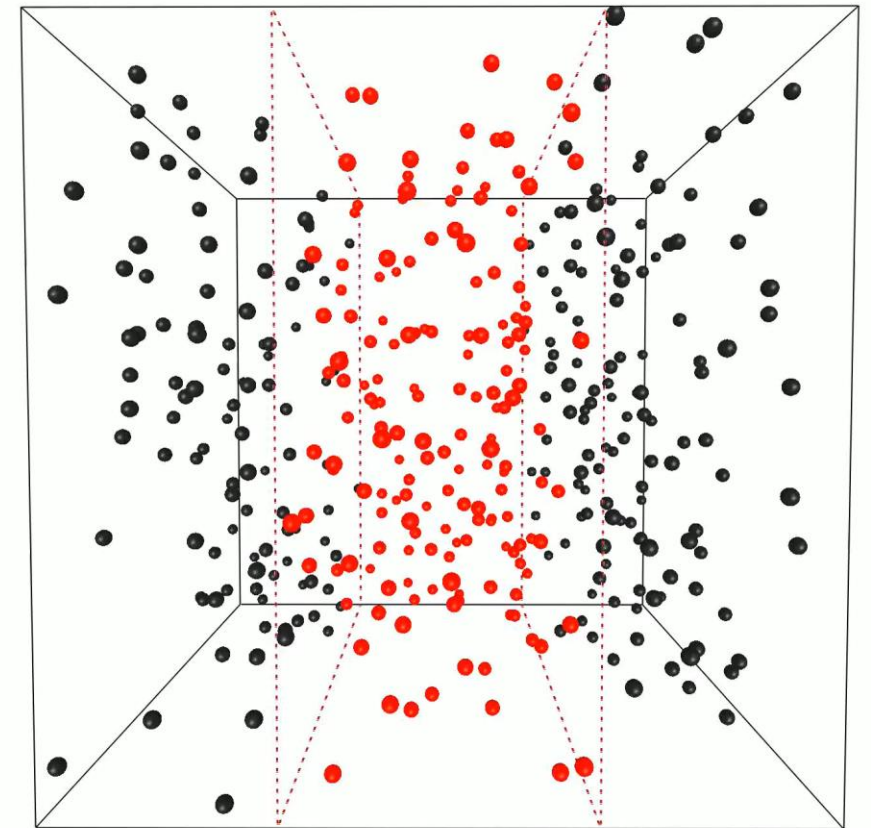
# EXPERIMENT VS BOX



## Heavy-ion simulation (UrQMD)



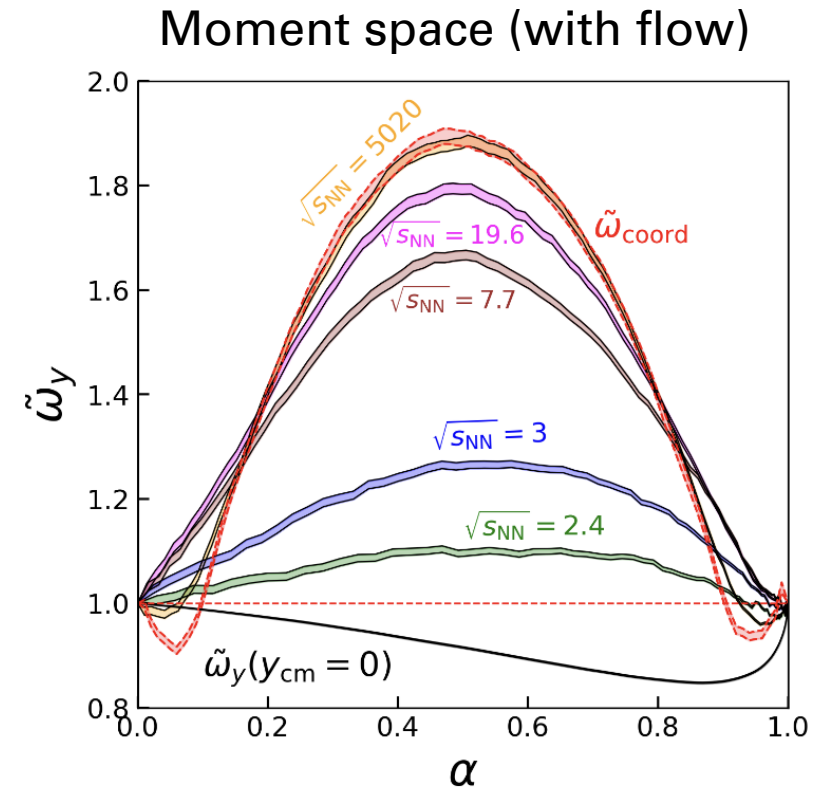
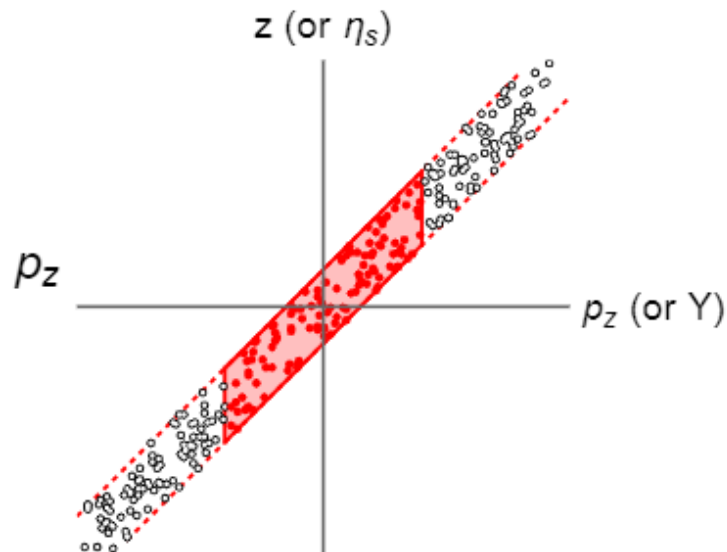
## MD box simulation (LJ)



# ADDING COLLECTIVE FLOW

We now add longitudinal flow (Bjorken-like) by boosting particles according to the  $z$ -coordinates

$$\tilde{y} = \tilde{y}^{LJ} + \sqrt{\frac{T_{frz}}{m\tilde{T}}} \kappa \left( \frac{\tilde{z}}{\tilde{L}} - \frac{1}{2} \right), \quad \kappa = 2y_{cm} \sqrt{\frac{m\tilde{T}}{T_{frz}}}$$



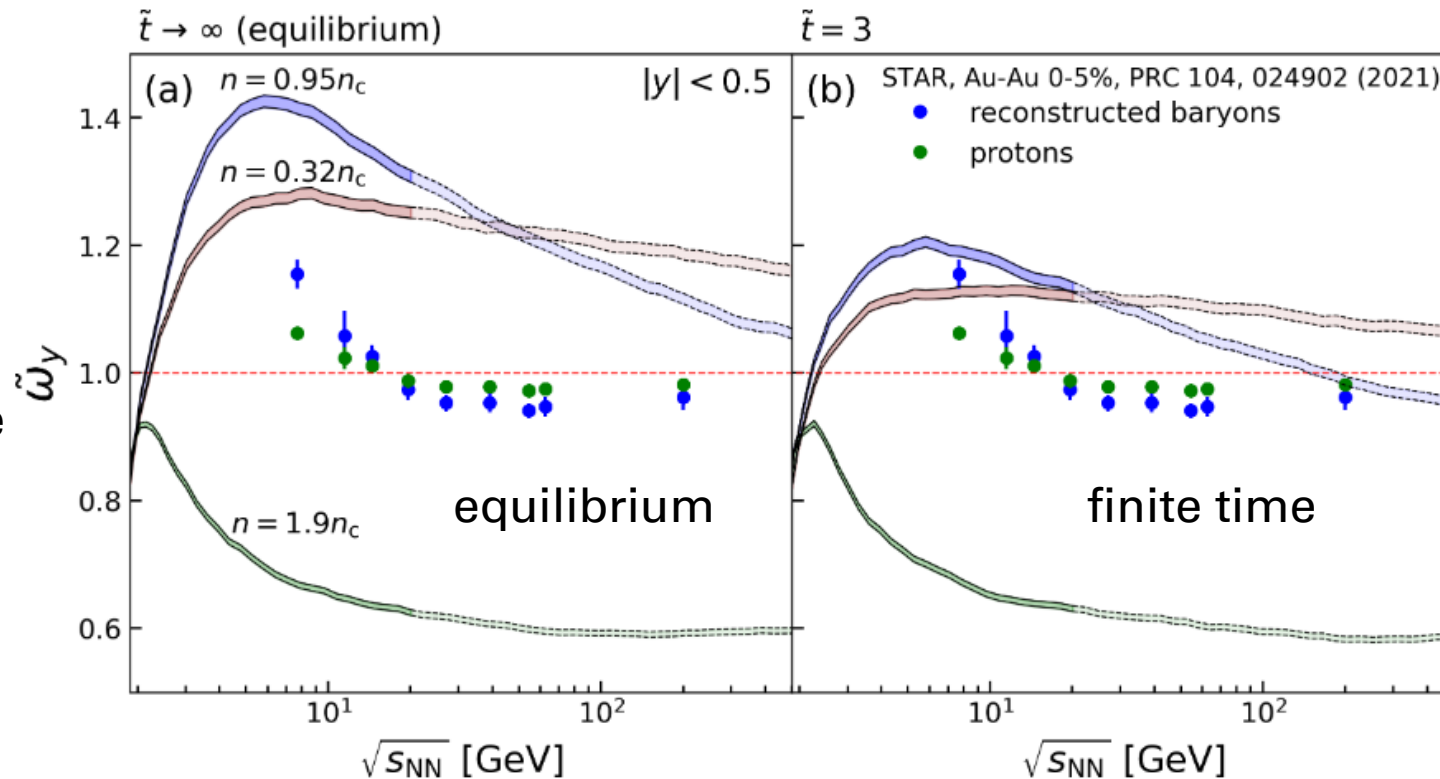
Heavy-ion collision inspired parameters:  $T_{frz} = 150$  MeV,  $m = 938$  MeV

Collective flow correlates momenta to coordinates and recovers correlations

# 12 FLUCTUATIONS FOR CONSTANT RAPIDITY CUT

$$\alpha = \alpha(\sqrt{s_{NN}}) = \frac{\langle N \rangle}{N_{tot}}, \quad y_{cut} = 0.5 \text{ (experiment)}$$

(Reconstructed STAR BES-I data was used, BES-II in future)

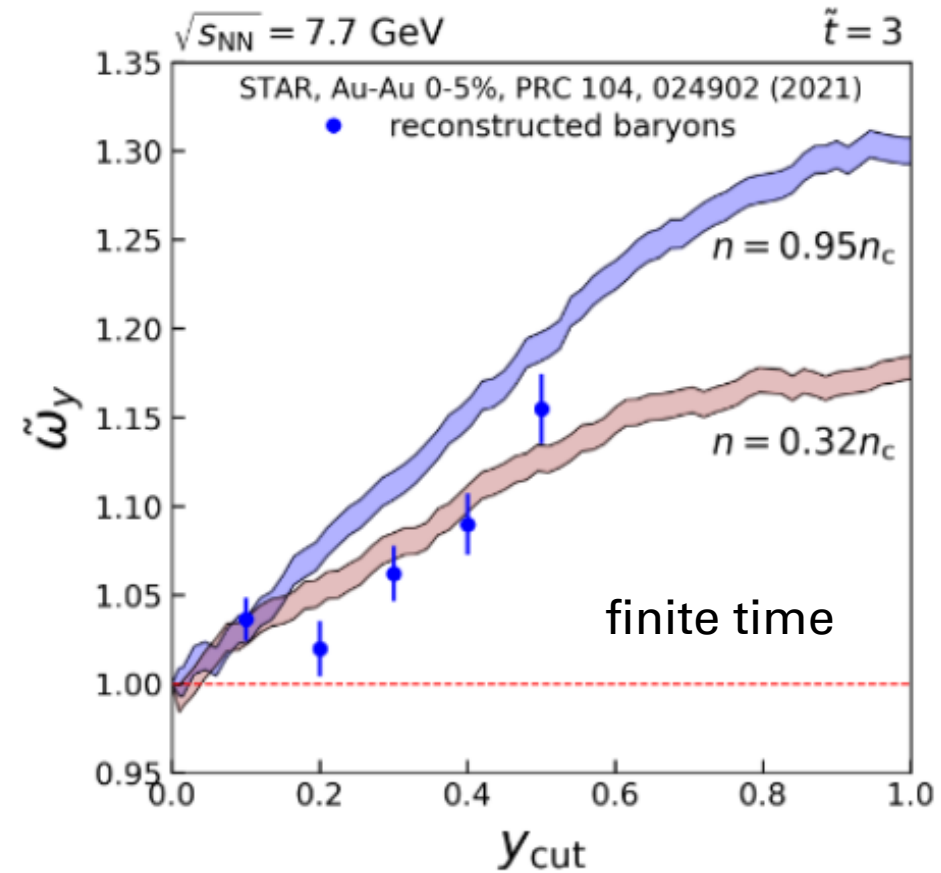
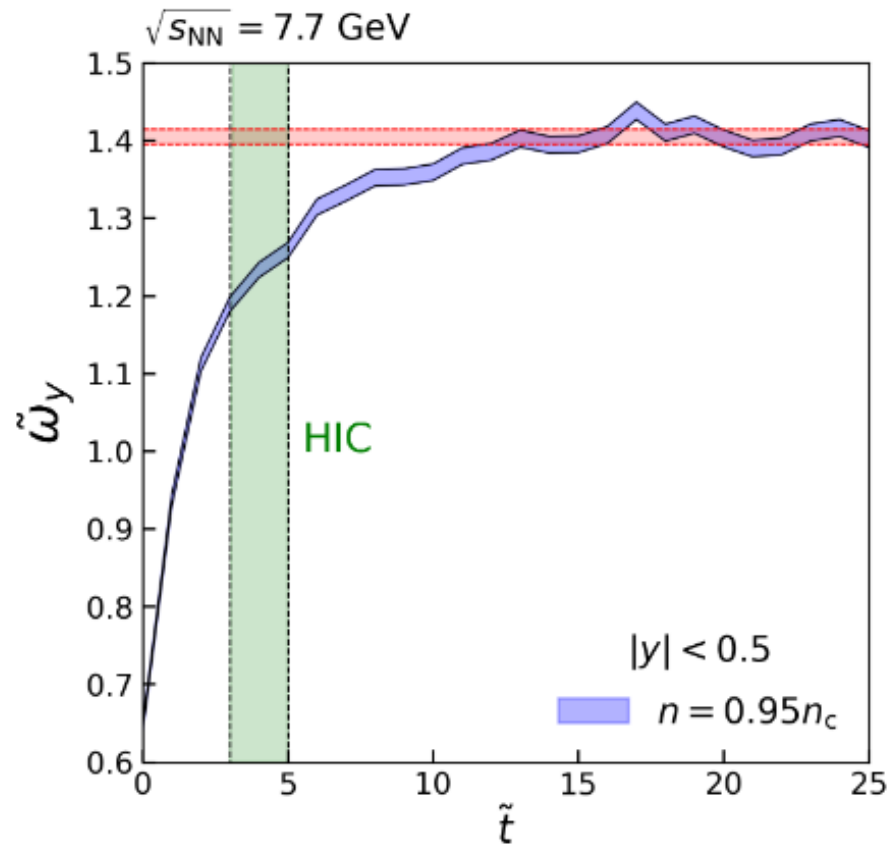


STAR data provided for the qualitative comparison

Data: M. S. Abdallah *et al.* (STAR Collaboration), PRC 104, 024902 (2021)  
 1- $\alpha$  correction based on hydro: Vovchenko, Koch, Shen, PRC105, 014904 (2022)

# 13 FLUCTUATIONS FOR CONSTANT RAPIDITY CUT

$\tau$ : 7 – 10 fm  $\rightarrow$  3 – 5 at units of  $\tilde{t}$



Data: M. S. Abdallah *et al.* (STAR Collaboration), PRC 104, 024902 (2021)

1- $\alpha$  correction based on hydro: Vovchenko, Koch, Shen, PRC105, 014904 (2022)

# SUMMARY

- Critical fluctuations are studied in a microscopic setup
- Ergodic hypothesis is shown to work for 2nd-order fluctuations along the  $\tilde{T} = 1.4 \sim 1.06T_c$  isotherm, including the vicinity of the critical point
- The collective flow effect allow us to see the enhancement of fluctuations in the momentum space
- Fluctuations in experimental rapidity acceptance  $|y| < 0.5$  are studied. If critical point is close to freeze-out, the largest signal is observed at  $\sqrt{s_{NN}} \sim 5$  GeV as interplay between longitudinal flow and number of protons in acceptance (system size).

# OUTLOOK

- Putting our study in the context of the BES II
- Higher-order cumulants (need bigger statistics)
- Study of the mixed phase
- Implementing CP dynamics into the transport theory (UrQMD/SMASH)

**Thank you for your time!**  
**Questions?**

# 16 FLUCTUATIONS AS CP SIGNATURE

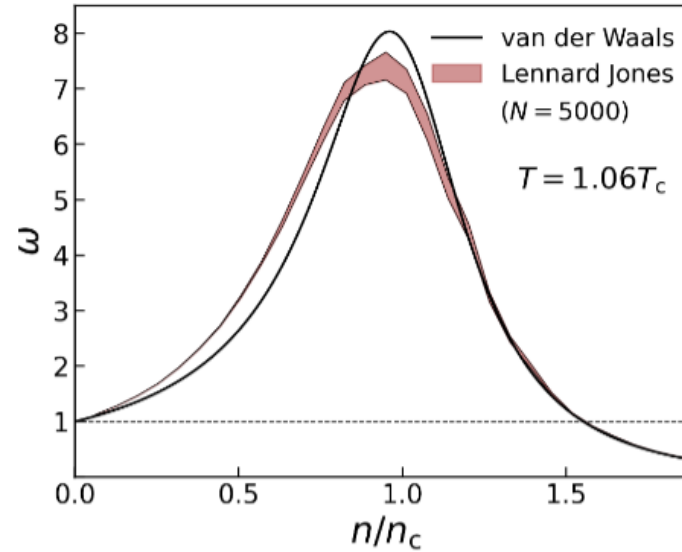


In GCE density cumulants shows singularity behavior in the critical point.

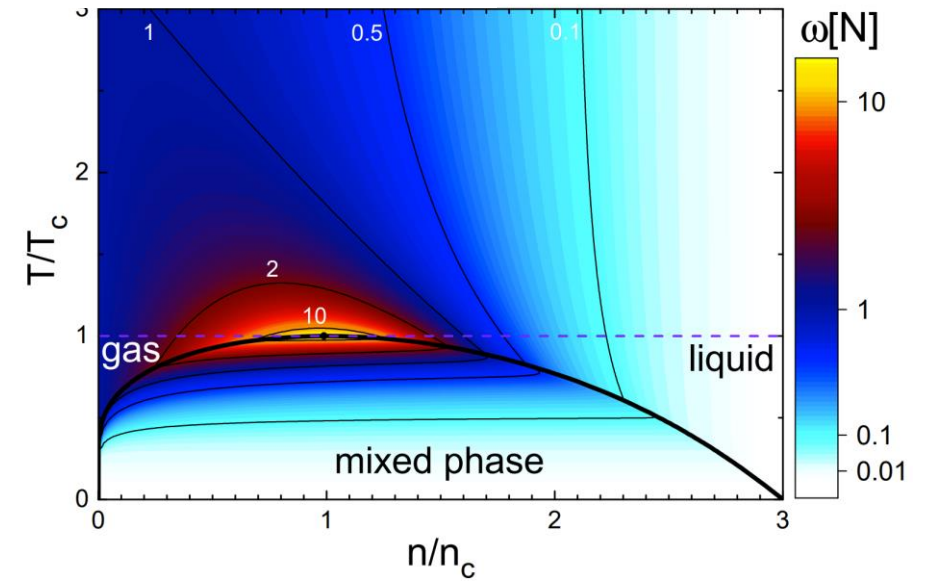
$$\ln(Z^{\text{gce}}) = \ln \left( \sum_{N=0}^{\infty} e^{\mu N} Z^{\text{ce}}(T, V, N) \right)$$

$$\kappa_n = \frac{\partial^n (\ln(Z^{\text{ce}}))}{\partial (\mu_N)^n}$$

The real expression for  $Z^{\text{gce}}$  is unknown in QCD matter.

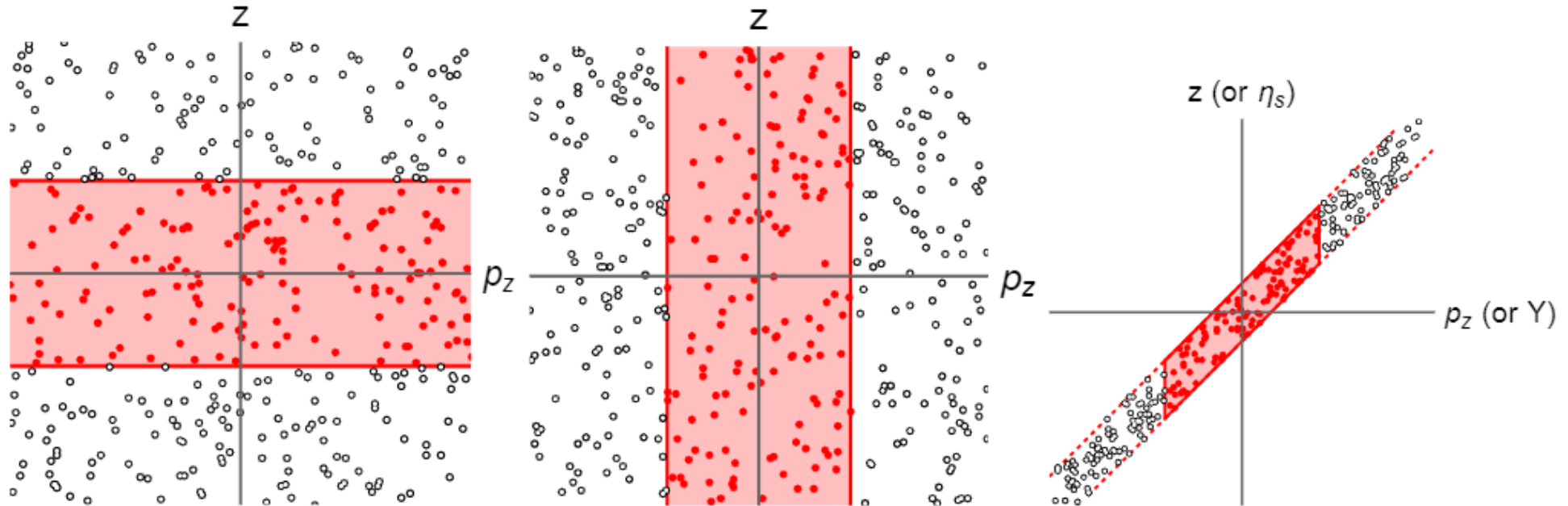


$$\kappa_2 \sim \varepsilon^2$$





# MOMENTUM VS COORDINATE CUTS



# SUB-ENSEMBLE METHOD: CORRECTION FOR GLOBAL CONSERVATION

In the case of interacting system one can find

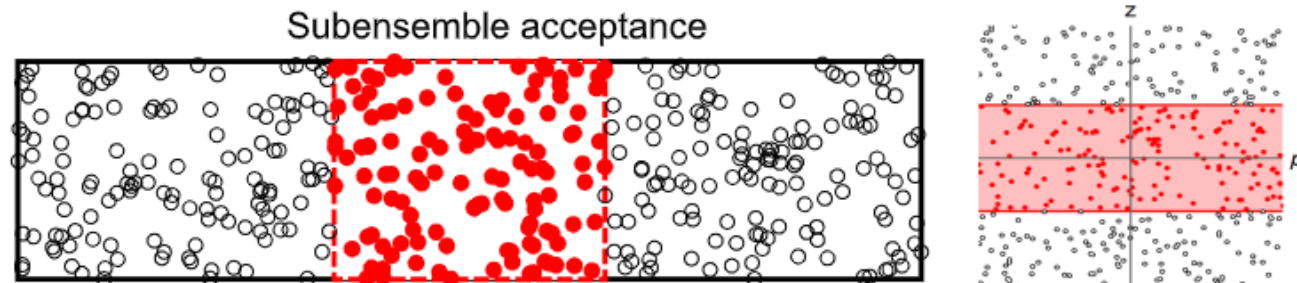
$$\kappa_1 = \alpha VT^3 \chi_1, \quad \kappa_2 = \alpha(1 - \alpha) VT^3 \chi_2$$

By definition in cumulants formalism

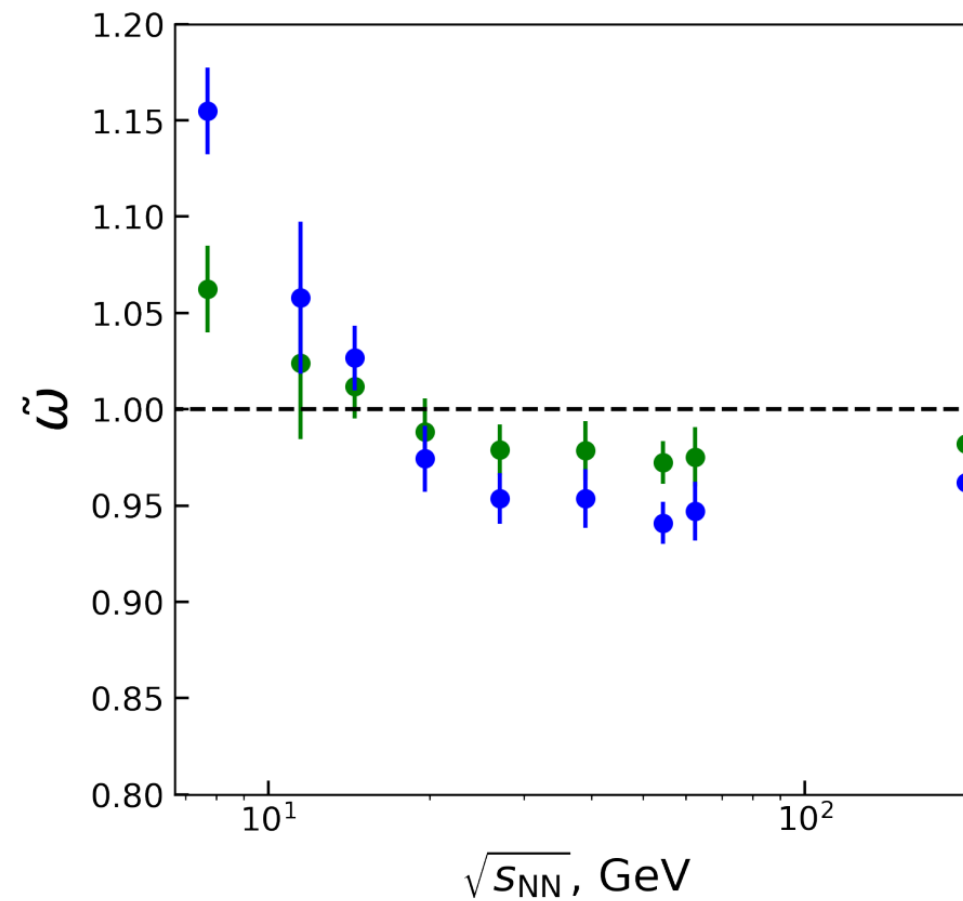
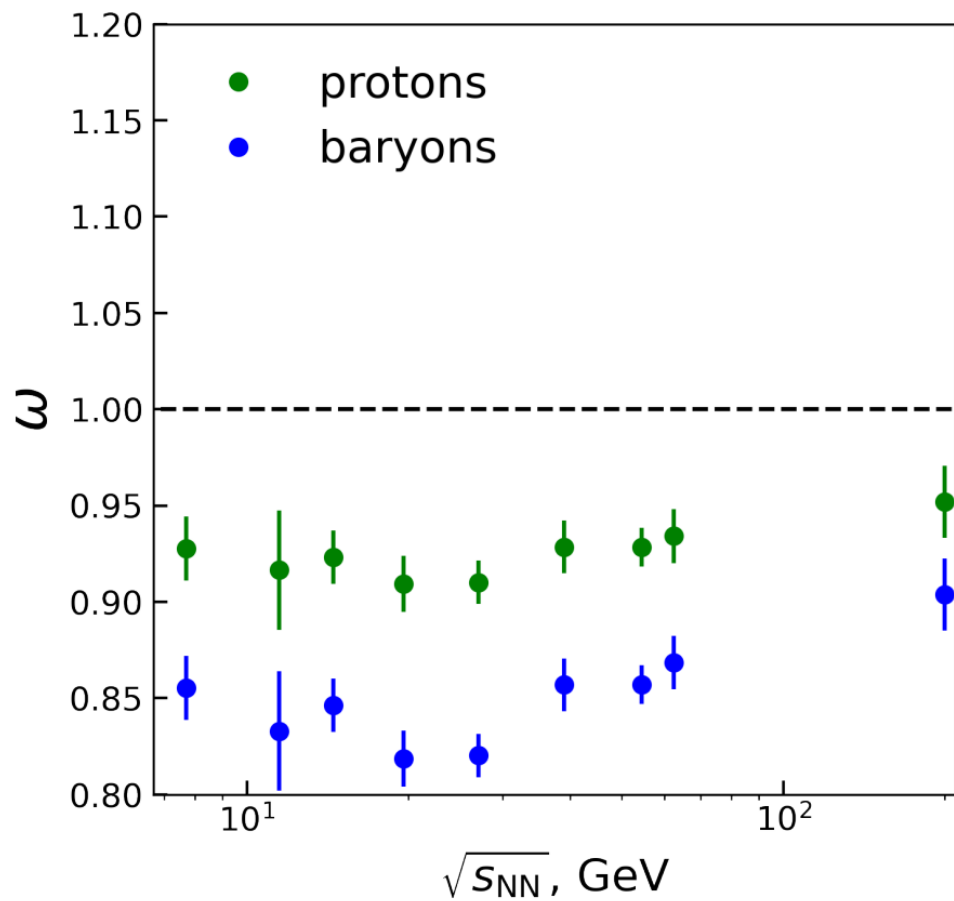
$$\omega^{coord} = \omega^{coord} / (1 - \alpha)$$

Finally, one can introduce the new measure

$$\tilde{\omega}^{coord} = \frac{\kappa_2}{\kappa_1} = (1 - \alpha) \frac{\chi_2}{\chi_1} = (1 - \alpha) \omega$$

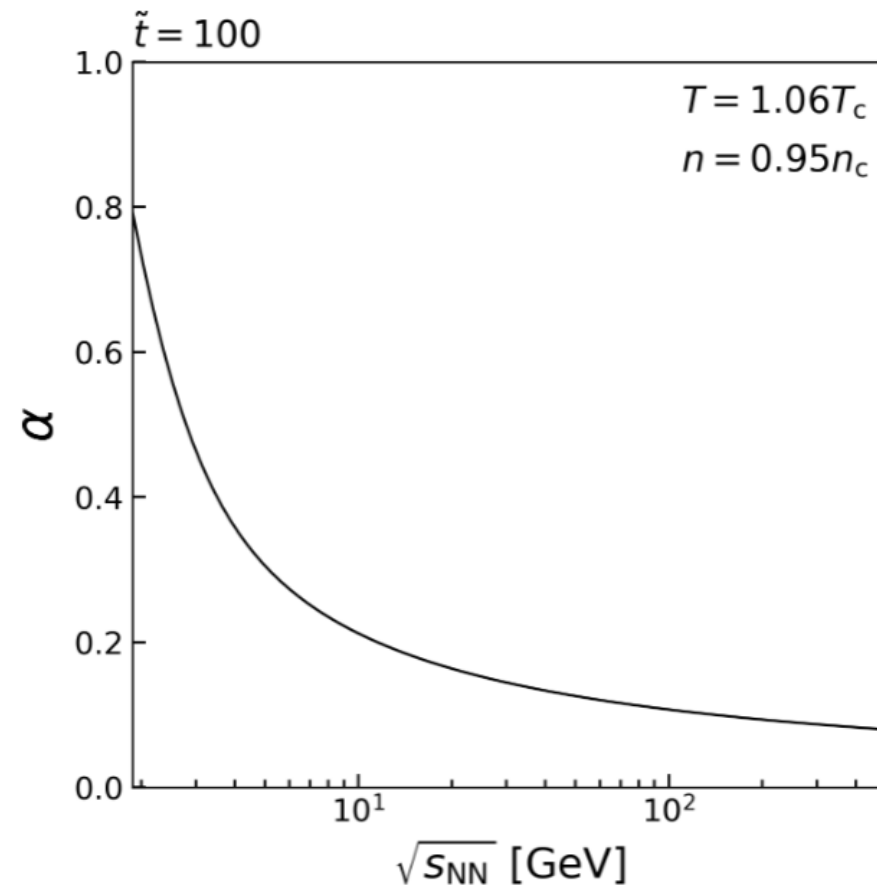


# RECONSTRUCTED STAR DATA



# ALPHA DEPENDENCE

$$\alpha = \alpha(\sqrt{s_{NN}}) = \frac{\langle N \rangle}{N}, \quad y_{cut} = 0.5 \text{ (experiment)}$$



# FLUCTUATIONS FOR CONSTANT ALPHA

$$\alpha = \text{const} = \frac{\langle N \rangle^{\text{acc}}}{N},$$

$$y_{\text{cut}} = y_{\text{cut}}(\sqrt{s_{\text{NN}}})$$

