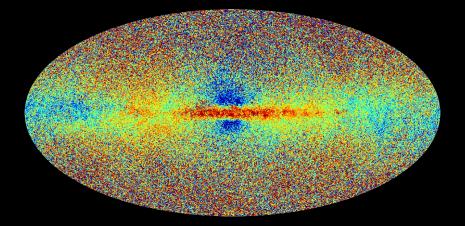
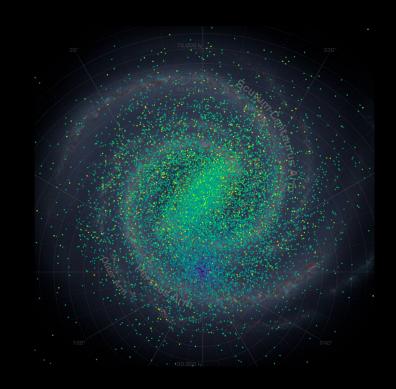
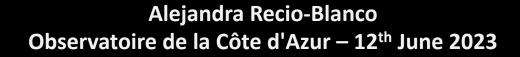
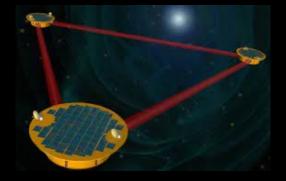
Gaia mission: lessons learned (personal view)





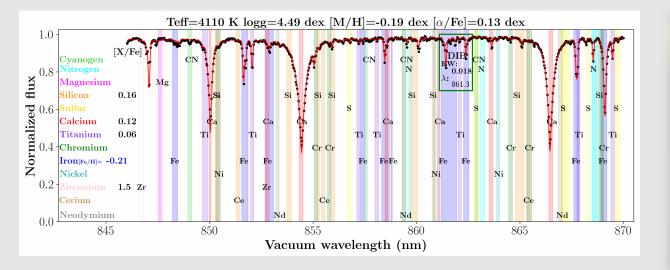




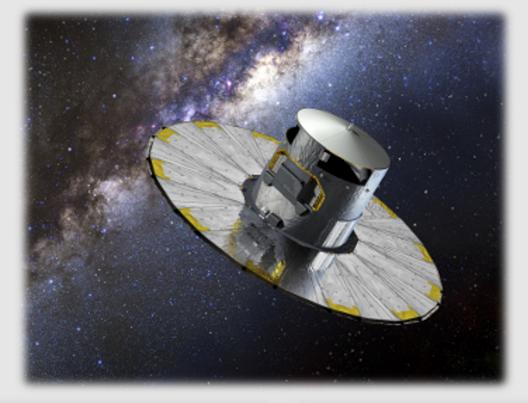


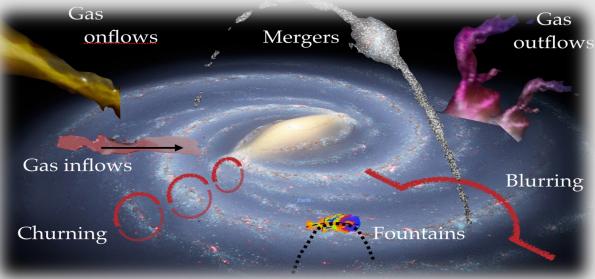
Outline of the talk

- 1. The expected and the unexpected
- 2. Keys of the Gaia revolution
- 3. Synergies Gaia-LISA



Credits:ESA/GAIA/DPAC-CU8-CU6 Recio-Blanco and the GSPspec team

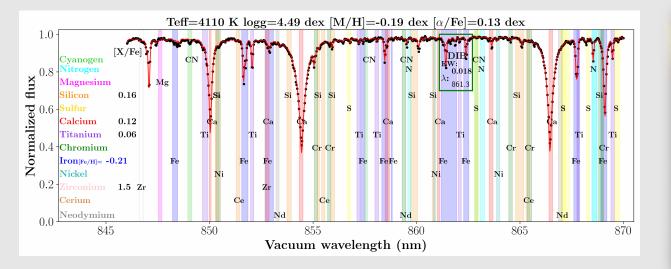




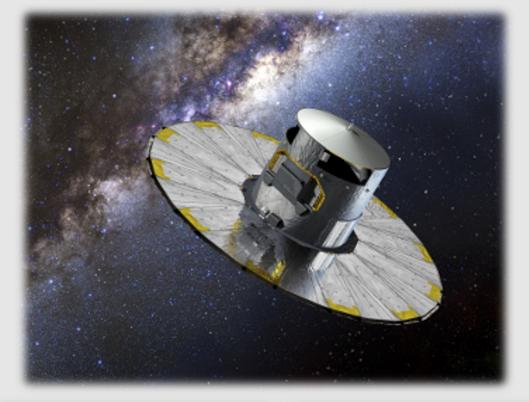
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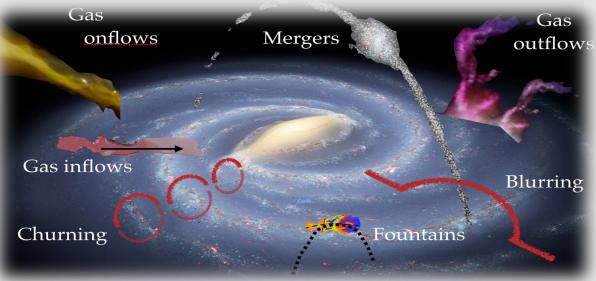
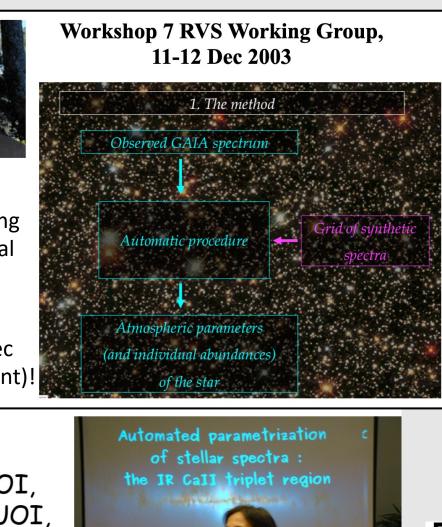




Photo during ESA external fellowship

copy-right Eric Lagadec (PhD student)!



POURQUOI, POURQUOI, POURQUOI?



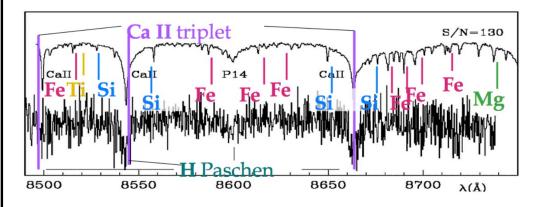
Alejandra Recic Collaborators: Patrick de 1 6. Kordoy Observatoire di

Brussels, 2011

DPAC creation in 2006 First CU8 DPAC meeting

esa

Generalized Stellar Parametrizer – Spectroscopy (**GSP-spec**) GWP-S-823-0000



Alejandra Recio-Blanco on behalf of the GSP-spec group Nice, 16-17 March 2006

When we started, automated spectra parameterization practically did not exist... and almost nobody cared!

Success in the expected

• Confirmed success of the astrometric observations Ground breaking results with a strong impact in many fields of astrophysics

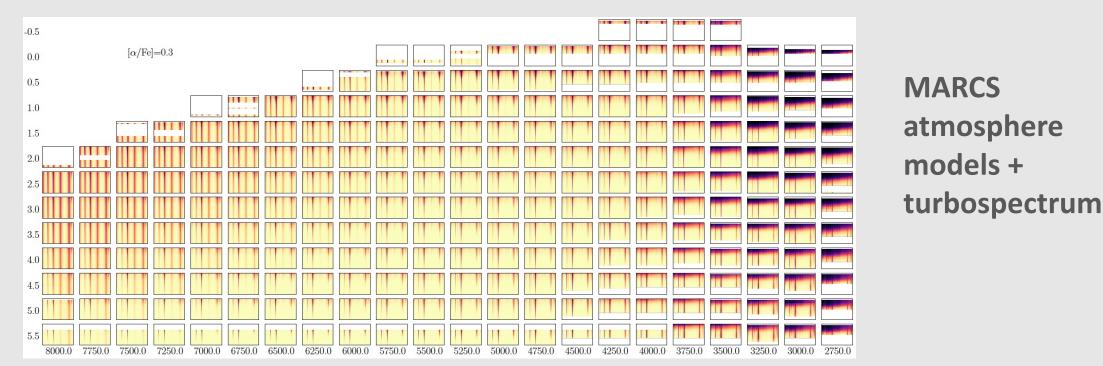


Success in the expected

• Confirmed importance of building up on existing expertise (e.g. data analysis, Hipparcos mission,...)

It allows to be more efficient, even pioneers in some aspects

• Confirmed importance of investment in modeling/simulations Crucial to prepare the pipelines and to understand the results!



Success in the expected

- Support from CNES and its data processing centre is a key of the success
 - CNES teams always supportive, helping to implement as much as possible all our (sometimes crazy) ideas. They are brilliant people !
 - CNES support in PhD thesis, post-docs and engineers is crucial before and during operations.



Expected difficulties

- VERY strong pressure on the teams.
- VERY long term engagements.
- Problems detected in the payload when operations start (e.g. stray-light)
- Hard dead-lines can be imposed...
- Misbehaviour (including harassment) happens. It is important to quicky have a code of conduct and reference people for these aspects.

Success in the unexpected

• Pessimistic predictions about spectroscopic data quality and derived products



1.0

0.8

0.6

0.4

0.035

0.030

0.025

8450

8500

8550

Biased by ground-based experience with similar instrumental configurations

Ground-based RAVE data

20070911_2229m20_137

8600

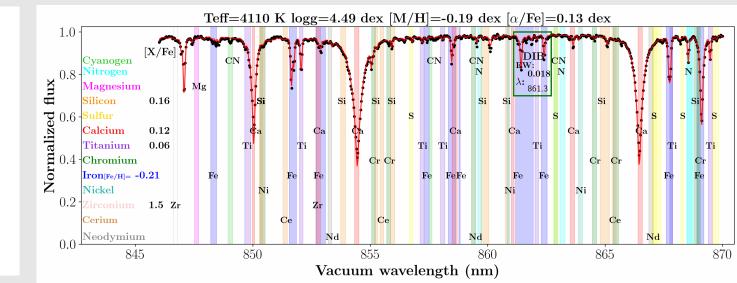
Wavelength

8650

8700

8750

8800



Gaia RVS data

Real data are much better than expected -> high precision in RVs, astrophysical parameters, chemical abundances, ...





Important

SPACE data

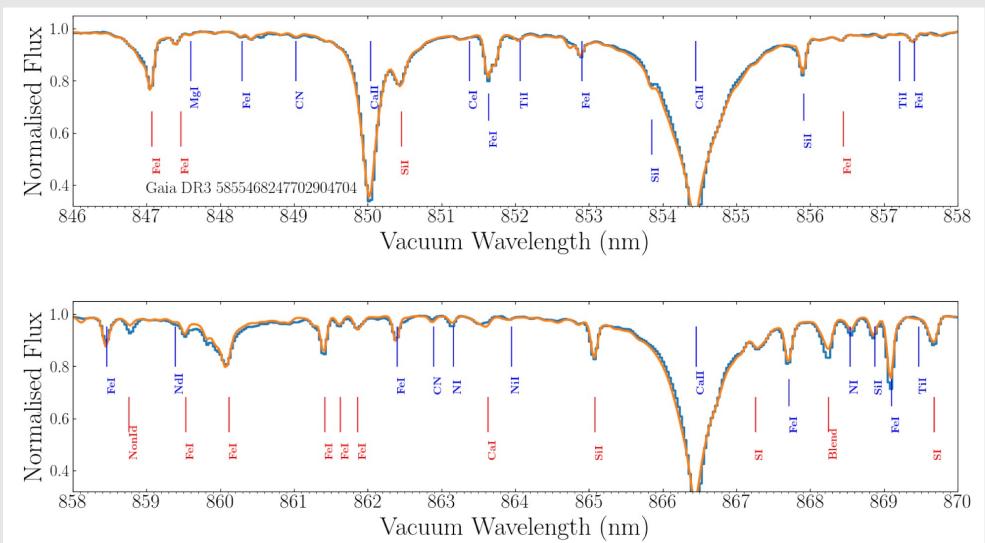
ground based data

- Continuous observations for years (34 months for DR3, ~25 000 h of continuous observations)
- Stable conditions (no atmosphere)
- Very good control and modeling of systematics
- Extremely homogeneous treatment
- High number statistics providing hundreds of thousands of high SNR (>150) data Parametrization quality comparable to ground-based surveys of higher spectral resolution and wavelength coverage.



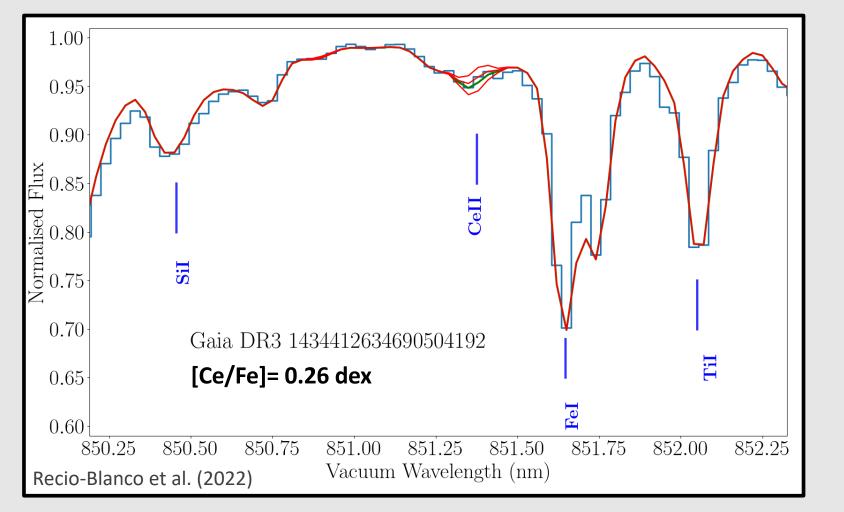


High quality spectra: continuous observations for 3 years, no atmosphere, control of systematics, ... Gaia is not a ground-based survey!





High quality spectra: continuous observations for 3 years, no atmosphere, control of systematics, ... Gaia is not a ground-based survey!



Heavy elements ex. Cerium

Unexpected difficulties

• Too optimistic predictions about some products (e.g photometric parameters)



Pre-launch simulations not realistic enough?

Important world-wide crisis (covid!) right before a crucial data release
Gaia was not locked down, but we were!! (with young children, or completely alone...)

 Analysis of key objects (e.g. standard candles, black holes) spread over CUs/WPs not interacting with each other in a standard analysis flow.
Creation of new specific Task Forces composed of cross-CU teams Slide from the Standard Candles Task Force kick off meeting (Resp. ARB)

Would these people have worked together if they had been Gaia DPAC members?





Ejnar Hertzsprung (CU5)

Antonia Maury (CU4)



Henry N. Russel (CU5)



Friedrich W. Bessel (CU3)



William and Margaret Huggins (CU6)

Angelo Secchi (CU6) The Harvard "computers" (DPCC)



Annie J. Cannon (CU8)



Cecilia Payne (CU8)



Henrietta Leavitt (CU7) Edwin Hubble (CU9)





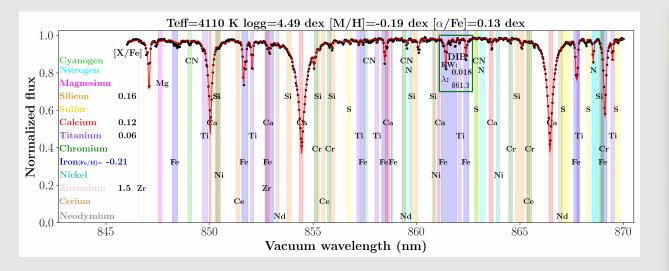
Harlow Shapley (CU7)

Outline of the talk

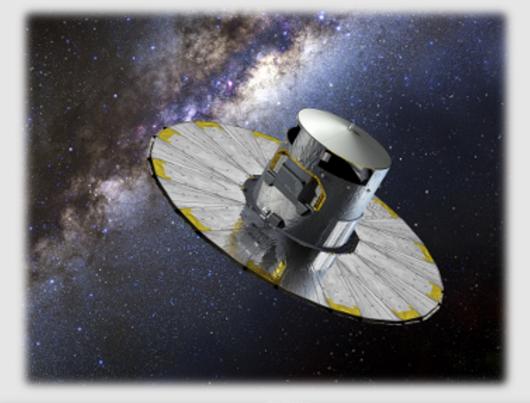
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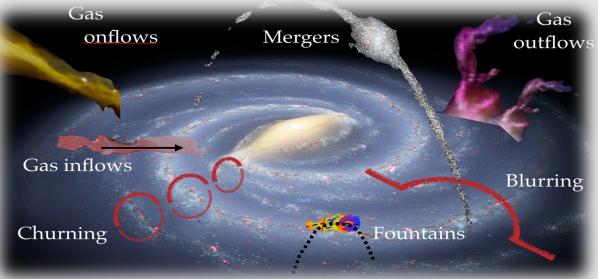
2. Keys of the Gaia revolution

3. Synergies Gaia-LISA



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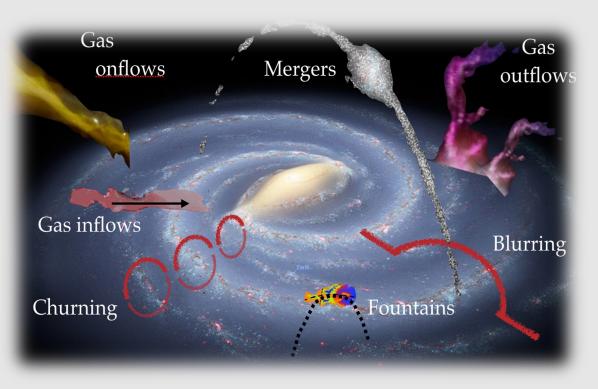


Gaia revolutions: roots and keys

Gaia combines the **astrometric** approach of **classical astronomy** with the **physical** approach of **modern astrophysics**. This is enhanced by:

- High number statistics
- High precision
- Time series observations

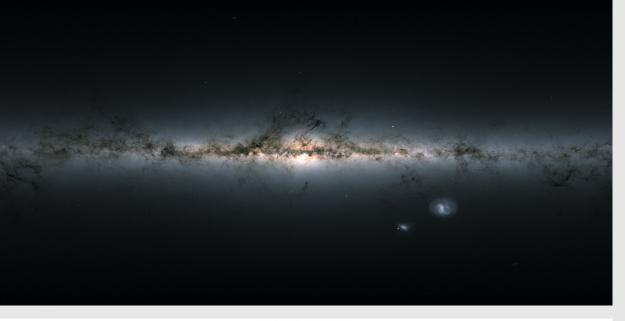
Detailed evolution of the Galaxy in its environment

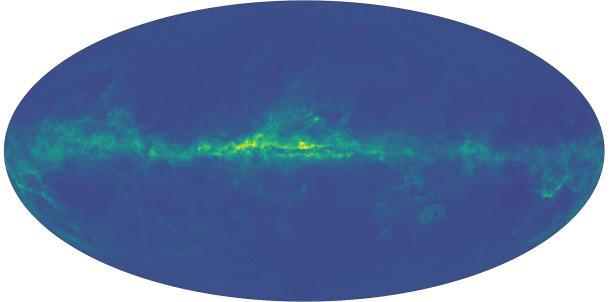


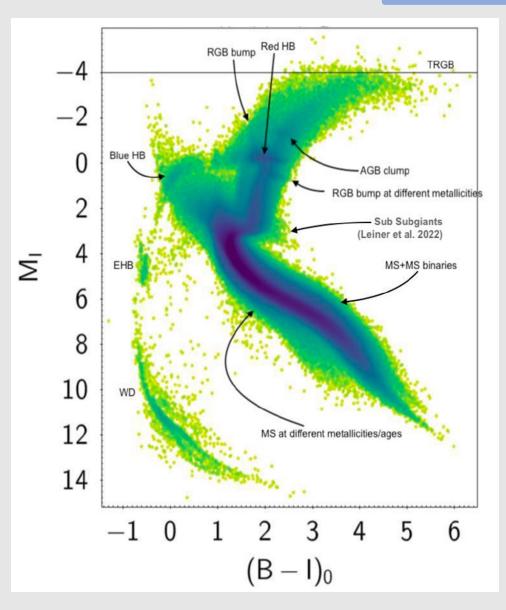
The keys of the Gaia revolution

- Parallaxes: the depth of the sky...
- Number statistics: 1.8 billion stars (astrometry+photometry) 33 million stars (spectroscopy) Nb increasing!





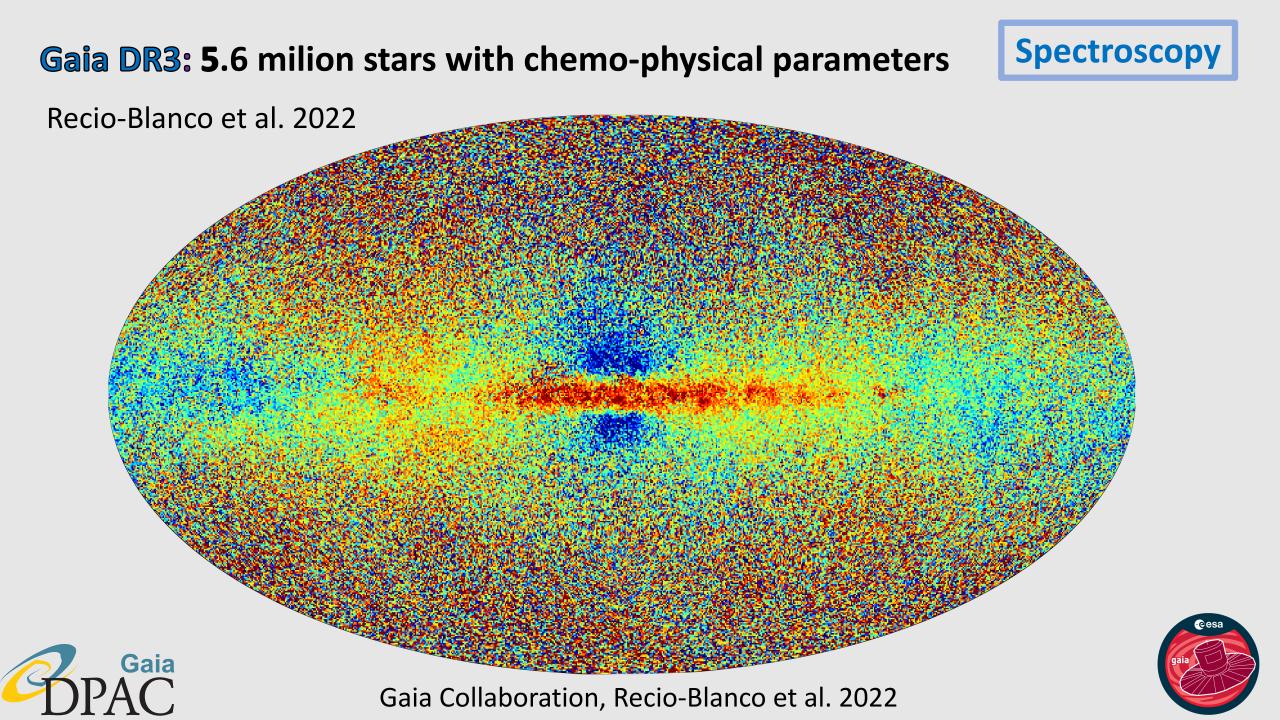


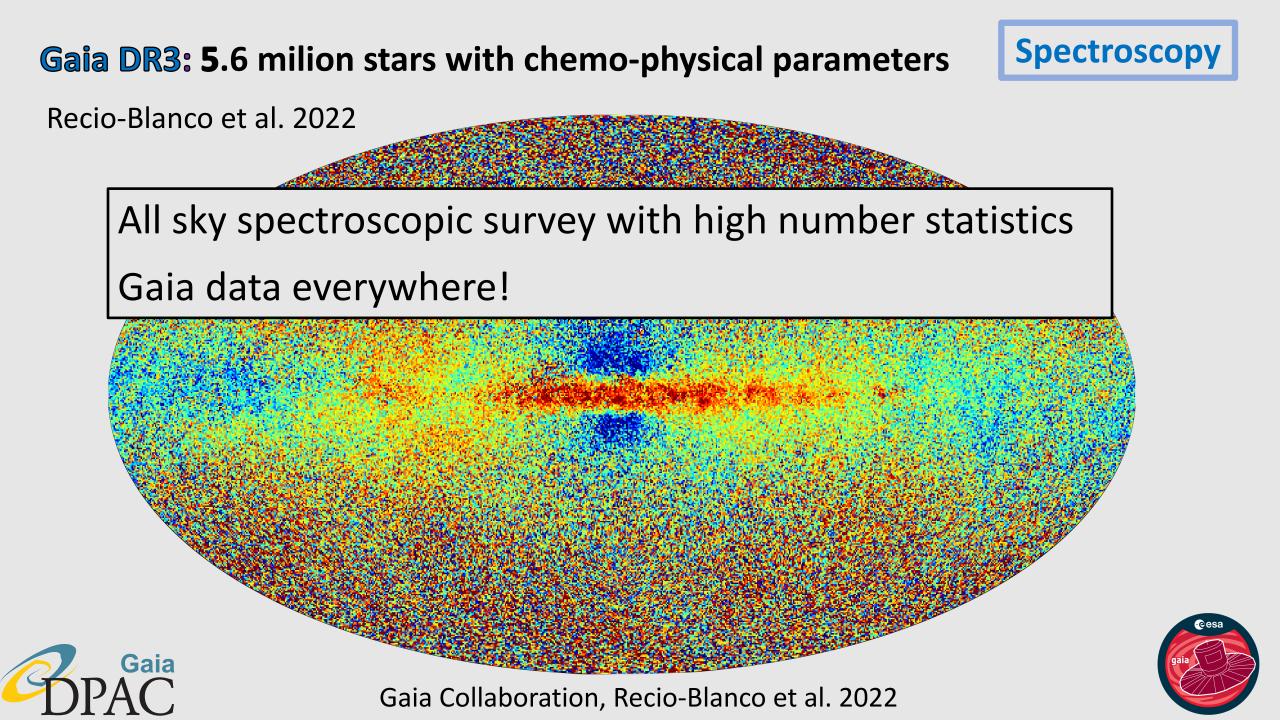




Gaia Collaboration, Montegriffo et al. 2022





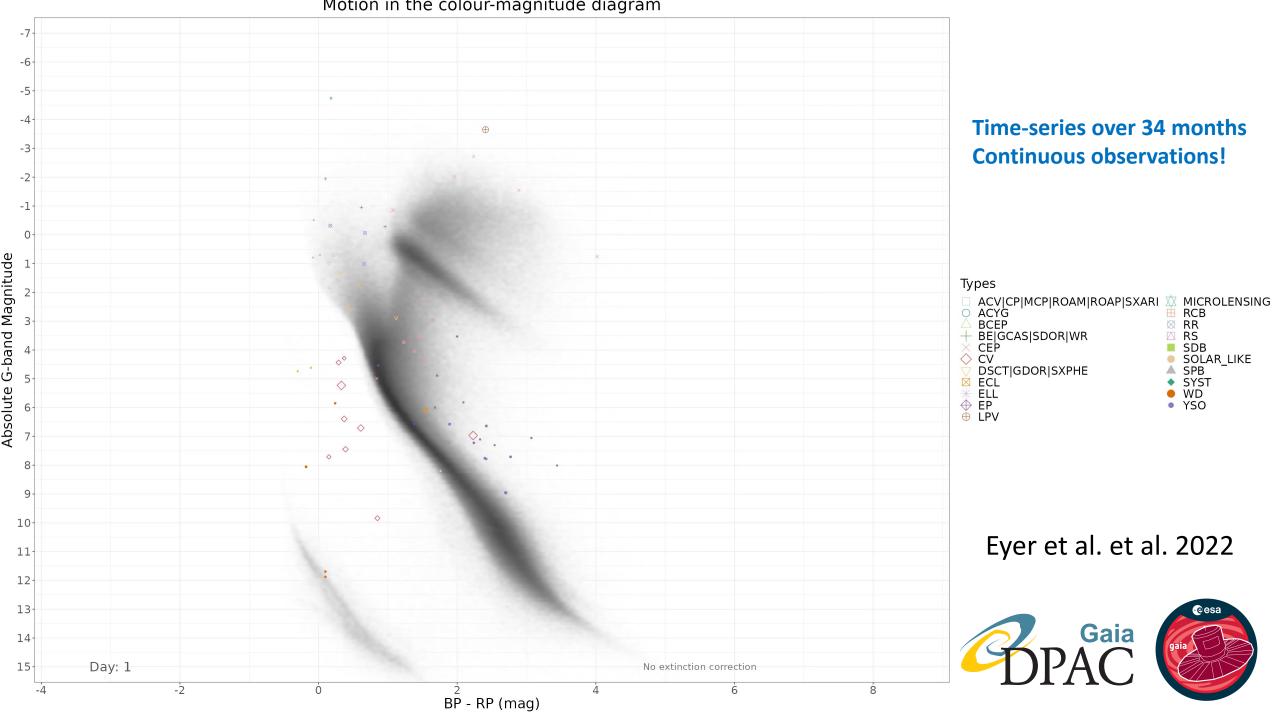


The keys of the Gaia revolution

- Parallaxes: the depth of the sky...
- Number statistics: 1.8 billion stars (astrometry+photometry) 33 million stars (spectroscopy) Nb increasing!
- Stability and precision: space observations (no Earth's atmosphere) extremely good control of systematics

The keys of the Gaia revolution

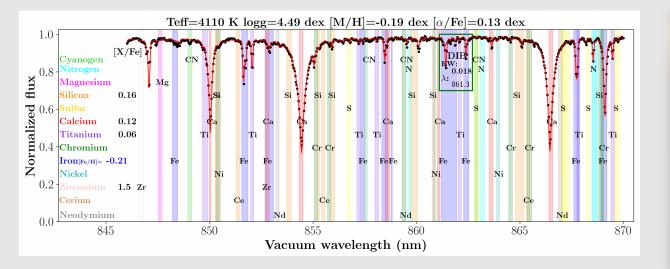
- Parallaxes: the depth of the sky...
- Number statistics: 1.8 billion stars (astrometry+photometry) 33 million stars (spectroscopy) Nb increasing!
- Stability and precision: space observations (no Earth's atmosphere) extremely good control of systematics
- Time-series (continuous observations for years): evolution!
 - Proper motions
 - Solar System acceleration
 - Stellar variability
 - Binaries and their orbital solutions



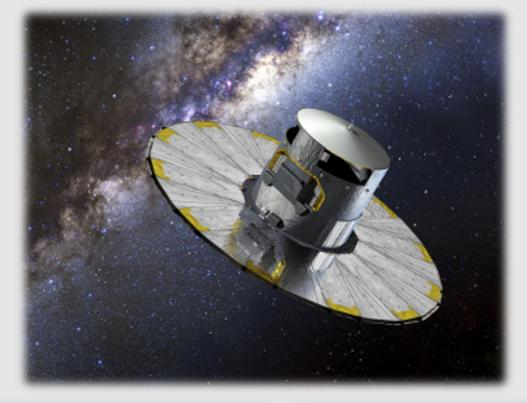
Motion in the colour-magnitude diagram

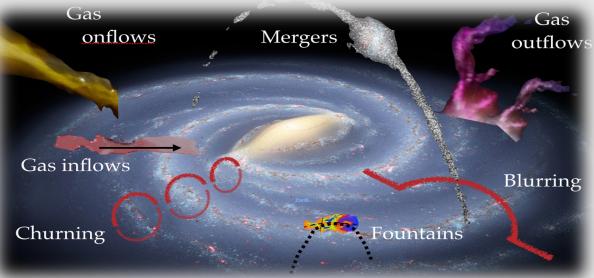
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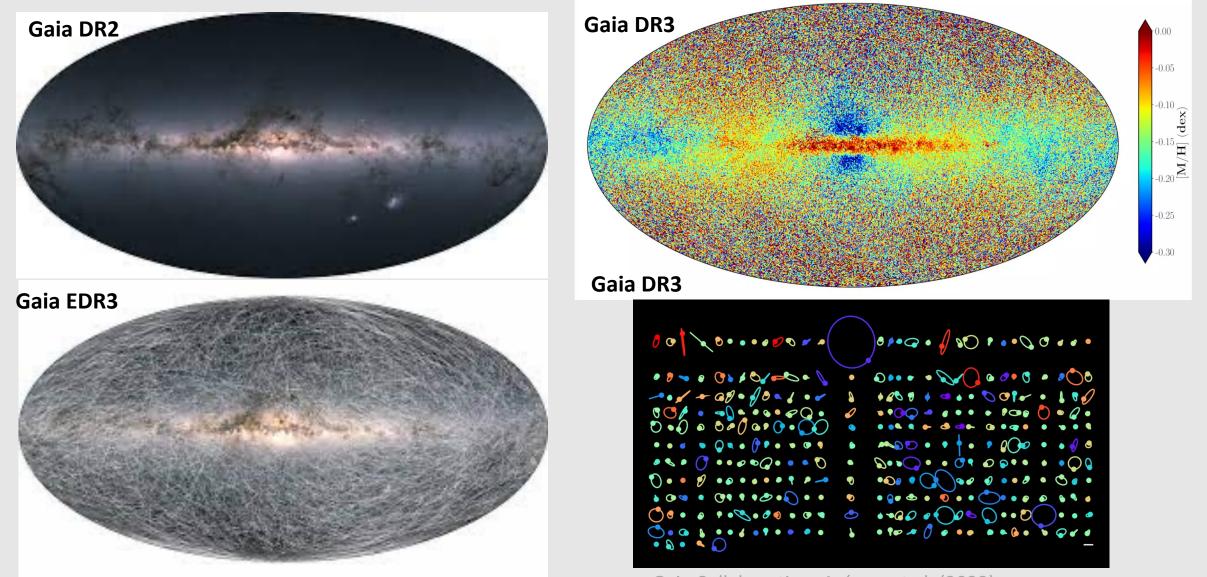


Credits:ESA/GAIA/DPAC-CU8-CU6 Recio-Blanco and the GSPspec team





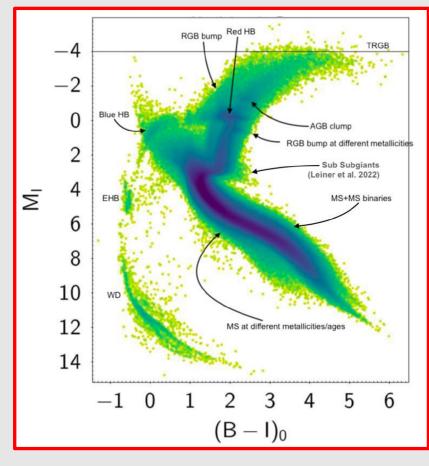
Gaia Collaboration, Recio-Blanco et al. (2022)



Gaia Collaboration, Arénou et al. (2022)

Binaries and compact objects

Gaia Collaboration, Montegrifo et al. (2022)



Gaia Collaboration, Arénou et al. (2022)

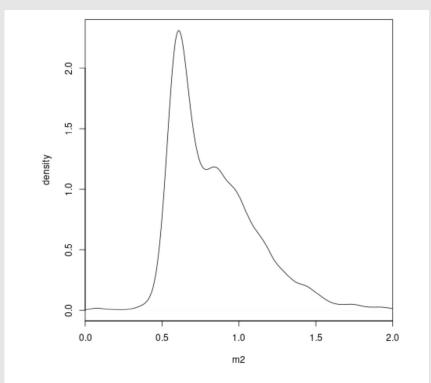


Fig. 17. Distribution of the secondary mass of astrometric solutions with fluxratio_upper = 0 in Table 3

Binaries and compact objects

Gaia Collaboration, Arénou et al. (2022) M. Barstow et al. Big improvement expected in DR4!

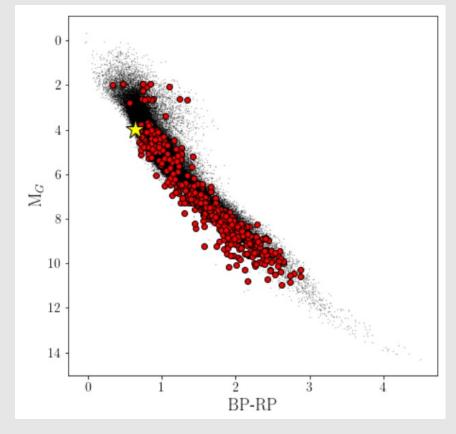
NSS catalogues

| Table | nss_solution_type | Solutions | Description |
|------------------------|------------------------|-----------|--|
| nss_acceleration_astro | Acceleration7 | 246947 | Second derivatives of position (acceleration) |
| | Acceleration9 | 91 268 | Third derivatives of position (jerk) |
| nss_two_body_orbit | Orbital | 134 598 | Orbital astrometric solutions |
| | OrbitalAlternative* | 629 | Orbital astrometric, alternative solutions |
| | OrbitalTargetedSearch* | 533 | Orbital astrometric, supplementary external input list |
| | AstroSpectroSB1 | 33 467 | Combined orbital astrometric + spectroscopic solutions |
| | SB1 or SB2 | 186 905 | Orbital spectroscopic solutions |
| | EclipsingSpectro | 155 | Combined orbital spectroscopic + eclipsing solutions |
| | EclipsingBinary | 86918 | Orbits of eclipsing binaries |
| nss_non_linear_spectro | FirstDegreeTrendSB1 | 24 083 | First order derivatives of the radial velocity |
| | SecondDegreeTrendSB1 | 32725 | Second order derivatives of the radial velocity |
| nss_vim_fl | VIMF | 870 | Variable-induced movers fixed |

Binaries and compact objects

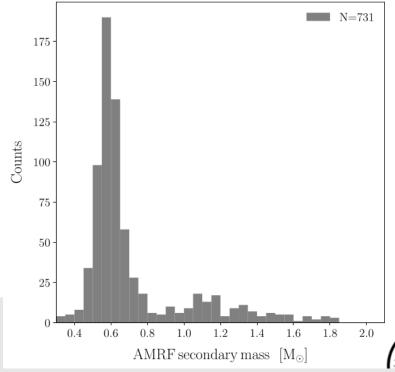
Gaia Collaboration, Arénou et al. (2022) M. Barstow et al.

Big improvement expected in DR4!



Compact objects

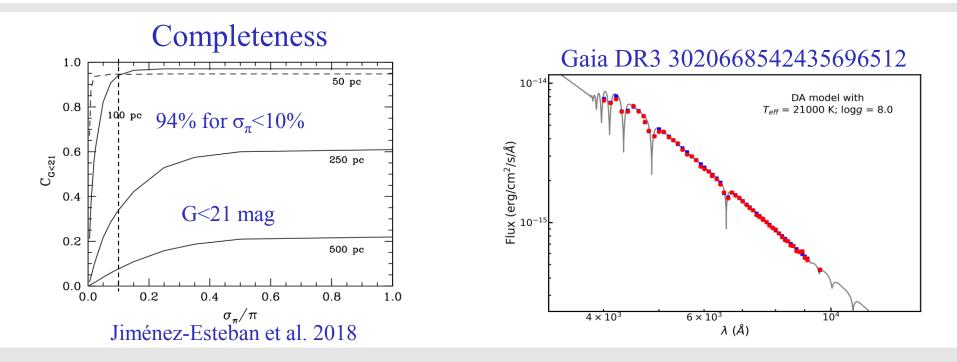
- Range of possible solutions which systems may contain WDs, NSs or BHs?
- SB1 hidden companion
- Eclipsing
- Orbital
- SB2 only as a 3rd component



Binaries and compact objects

Big improvement expected in DR4!

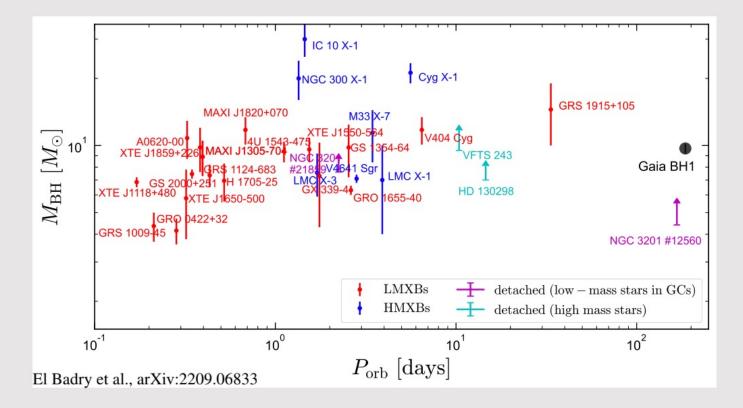
- Jiménez-Esteban et al. 2018 \rightarrow 100 pc WD sample is almost complete
- Gaia-DR3 has provided ~100,000 low-resolution spectra of WDs



Binaries and compact objects



Gaia BH1

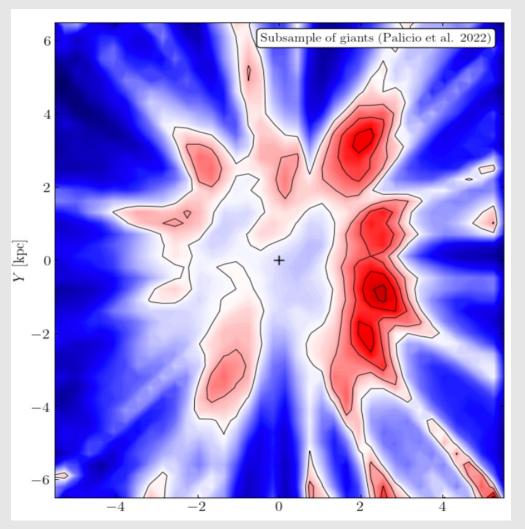


- Unique system among the known black holes: solar type star on wide orbit around dormant BH
- Evolutionary path to make such systems is unclear
- Gaia DR4 may uncover dozens more cases

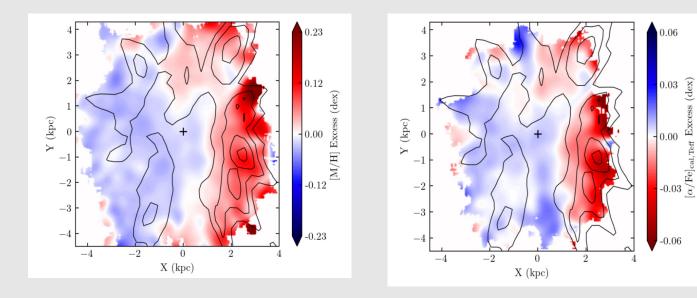
Other important information

- Binarity estimates on different Galactic populations
- Characterisation of WD progenitors through the Galaxy
- Large scale, high statistics (~100 million stars in DR5) metallicity catalogues to constrain models
- Detailed Galaxy characterisation to improve LISA predictions

Revising LISA white paper predictions with Gaia DR3 data



Adapted from Palicio et al. (2023a) by M. Barbillon Spiral arms detected in stellar density and chemical abundances for old stars (Age>1Gyr)



Barbillon et al., in prep.

Gaia-LISA PhD thesis starting in Sept. 2023.

A (personal) final comment

Research is not part of a ground-segment data analysis consortium, but...

- Research is crucial to <u>really</u> understand the data products
- Research is essential to <u>correctly prioritize the data analysis</u> (there are always choices to make, we can not do everything)

Solutions (compliant with publication policies) exist :

- Science verification work (e.g. Gaia Collaboration papers)
- Support to teams to do research with public data (post-docs, PhDs)

