



Light dark matter search with DarkSide-50

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on behalf of the DarkSide-50 Collaboration



DarkSide-50

The Dual-Phase TPC

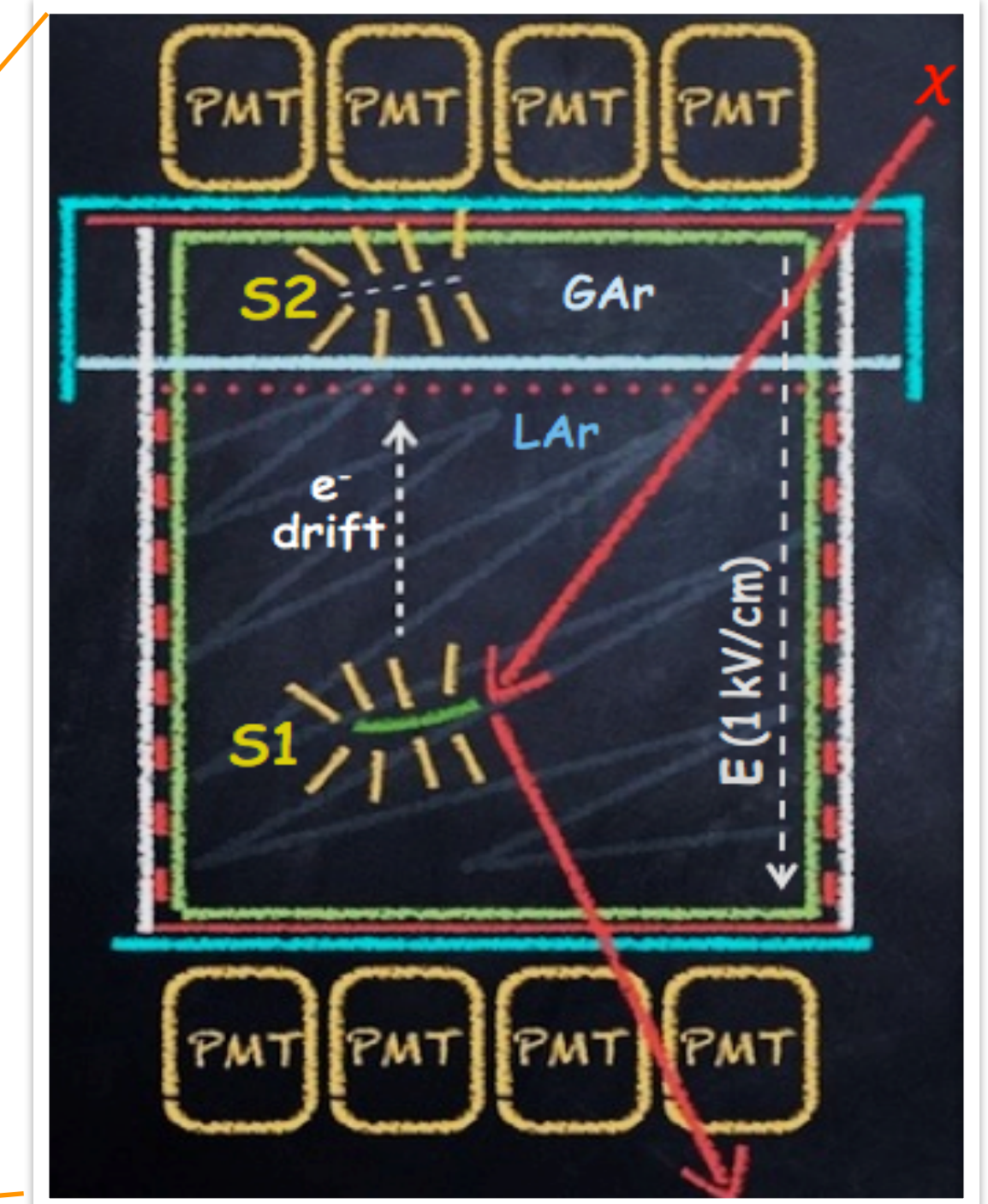
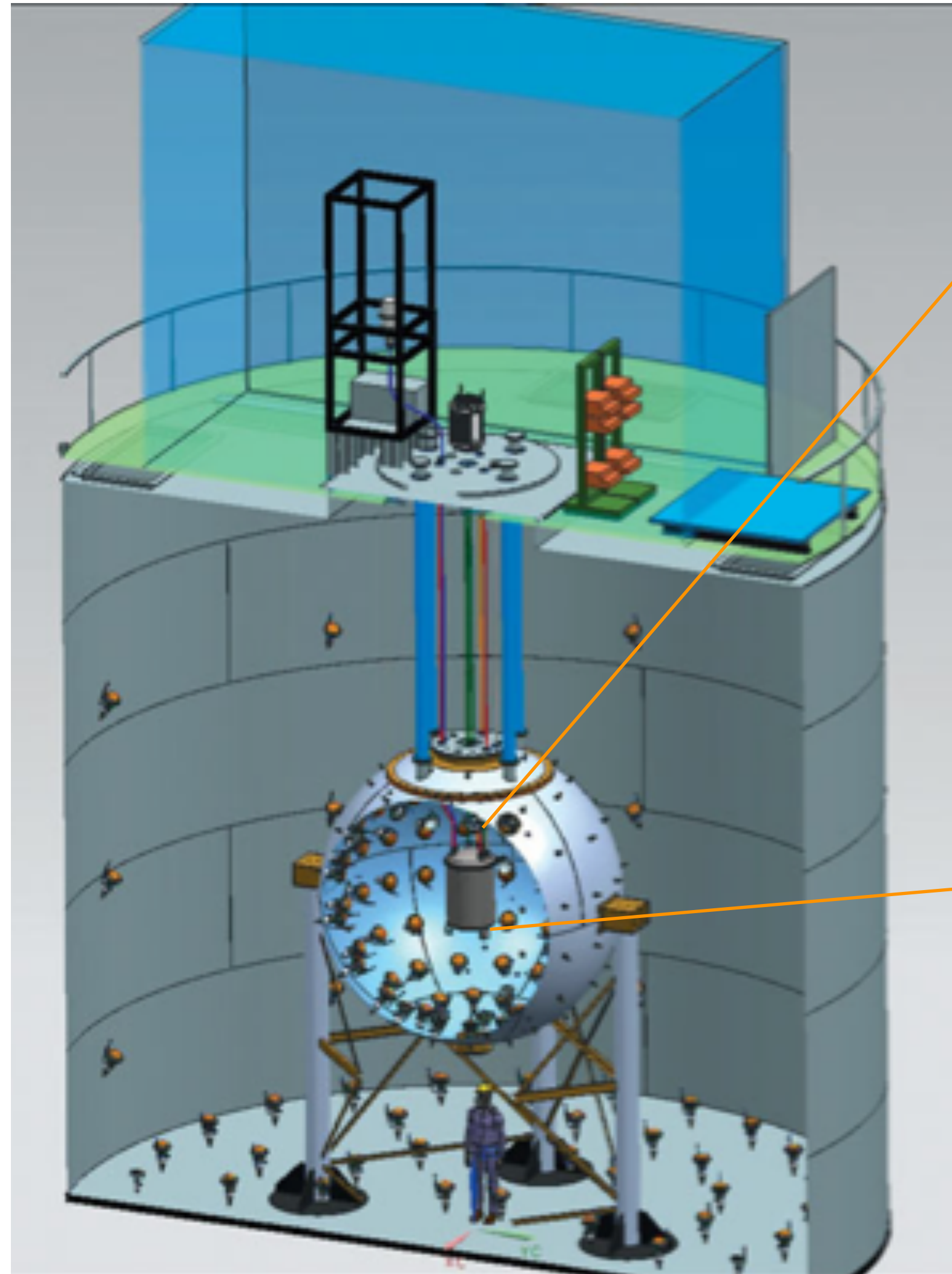
- 50 kg active mass of UAr
- 19 top + 19 bottom R11065 HQE 3" PMTs
- 36 cm height, 36 cm diameter
- Low field of 0.2 kV/cm drift

Liquid Scintillator Veto against neutrons

- 4 m diameter sphere
- Boron-loaded: 1:1 PC and TMB
- 110 8" PMTs
- LY ~ 500 pe/MeV

Cherenkov Water Detector

- 11 m diam. x 10 m
- 80 PMTs



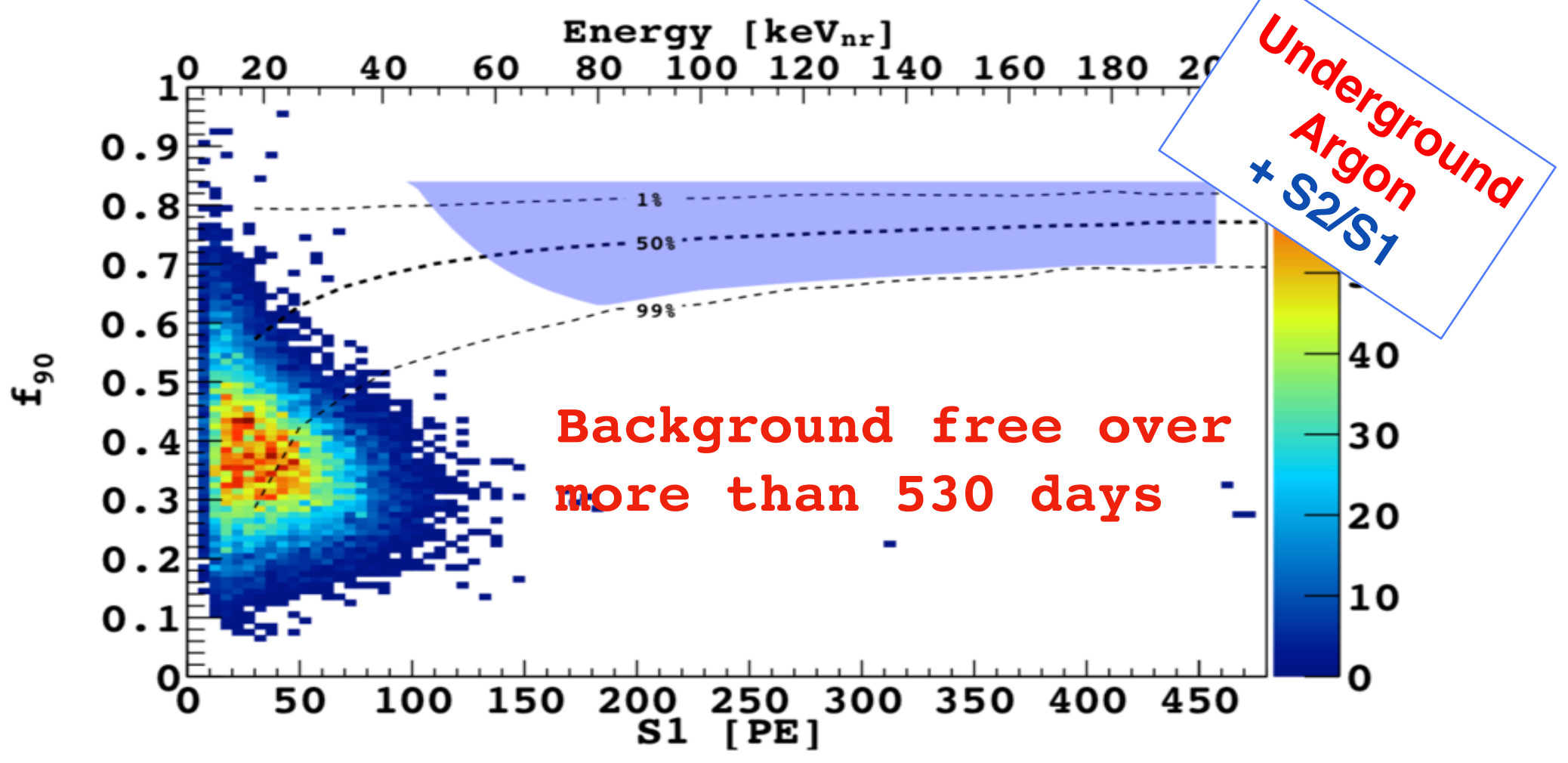
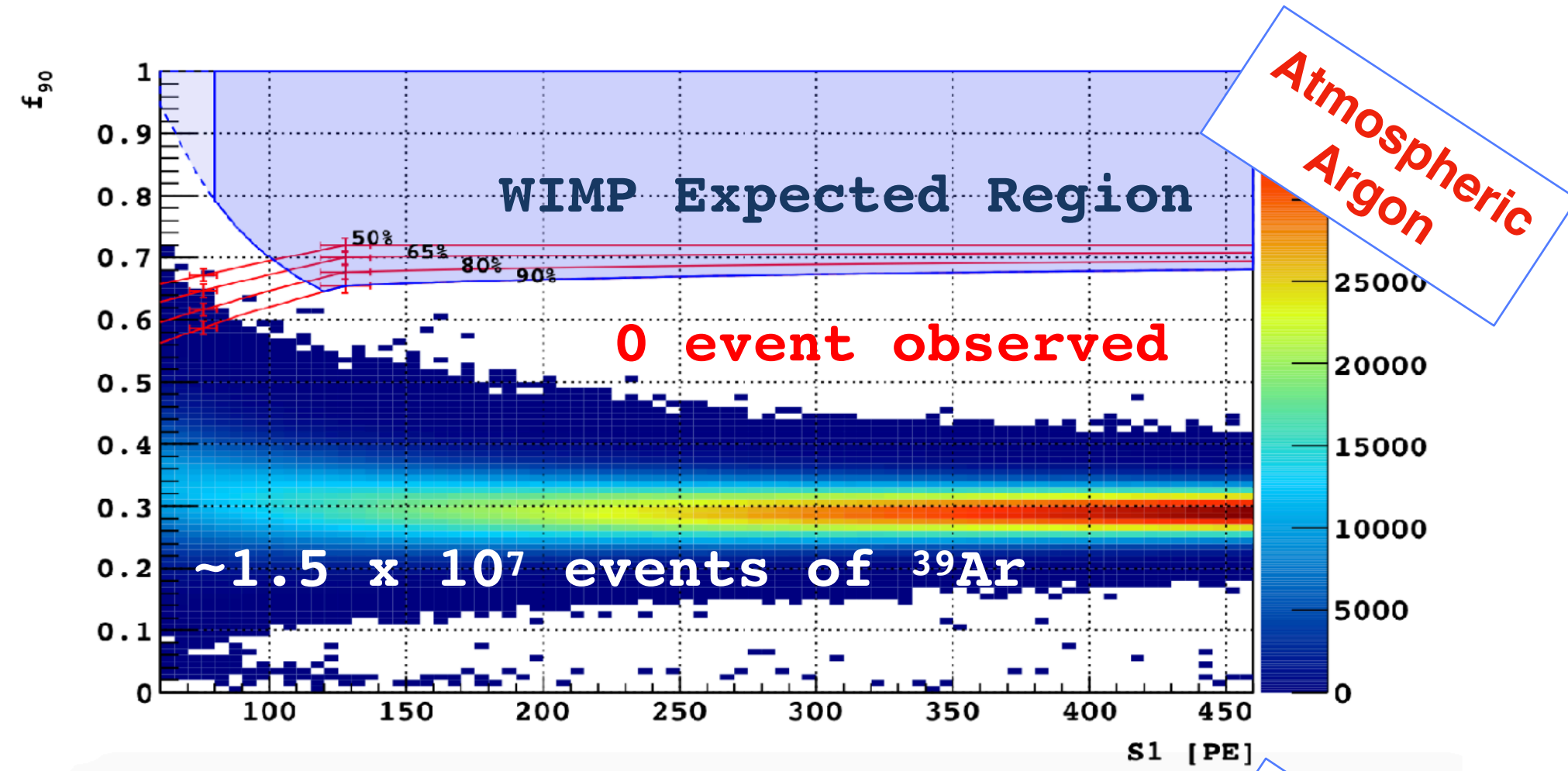
Efficient electronic recoil (ER) background rejection thanks to LAr scintillation **Pulse Shape Discrimination** through the “f90” observable (fraction of light detected in the first 90 ns)



The DS-50 WIMP Search

Pulse Shape Discrimination

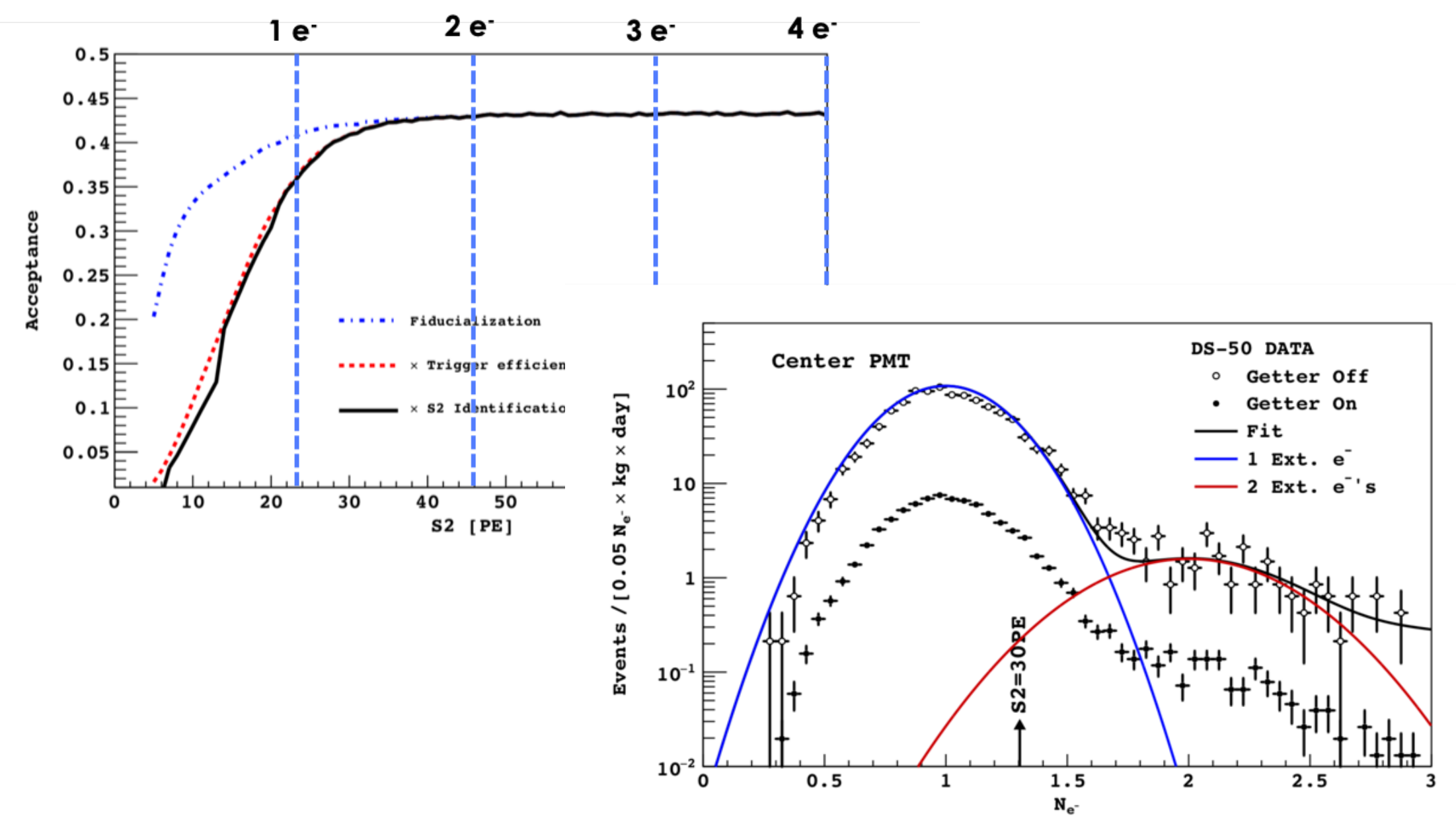
“high-mass” WIMP search



“low-mass” WIMP search

- ~~Scintillation (S1)~~
 - Detection efficiency (g1) ~ 16%
- Ionization (S2) ✓
 - Efficiency to extract 1 e- in the gas pocket ~ 100%
 - Amplification factor (g2) = ~23 pe / e-

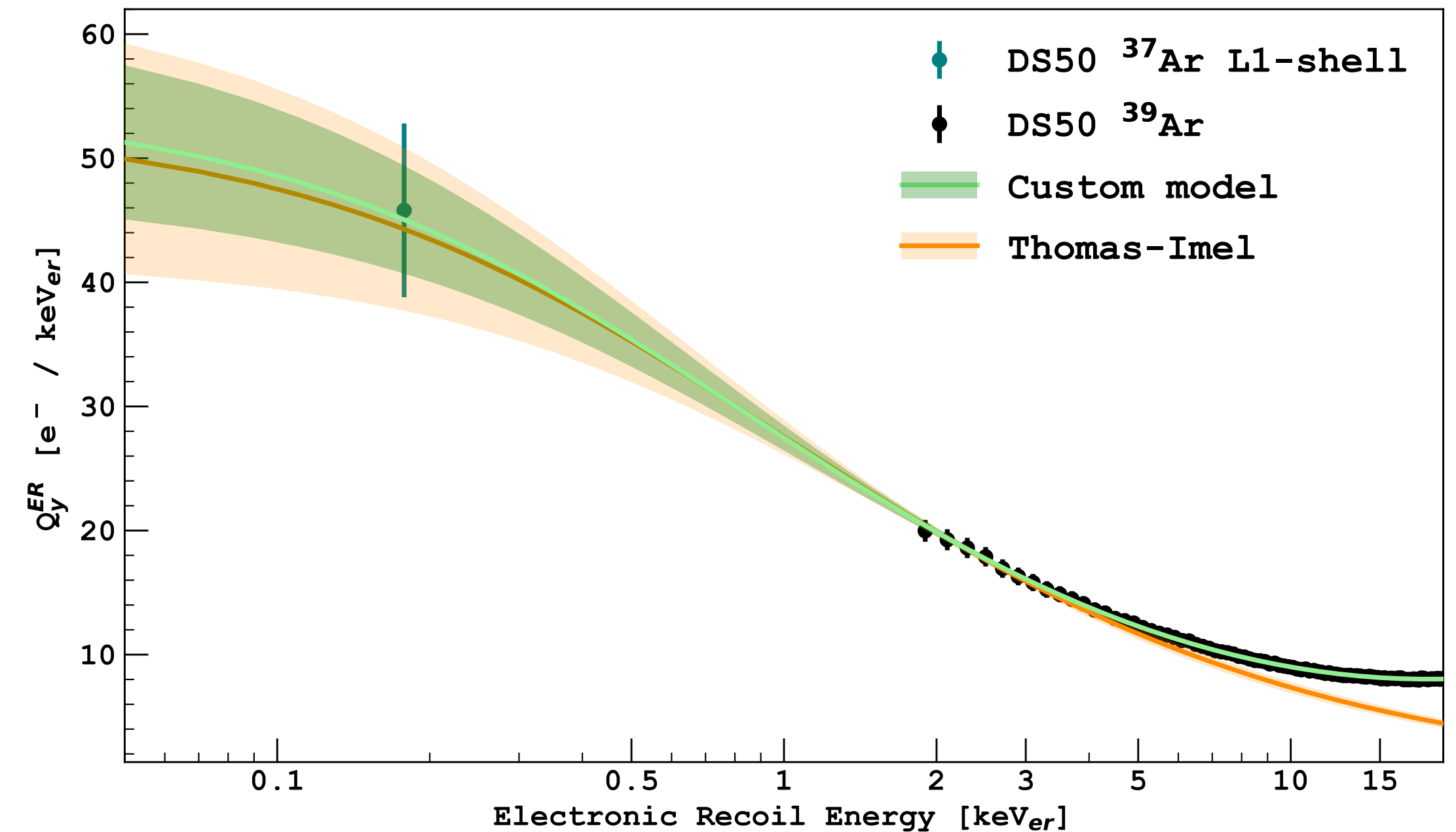
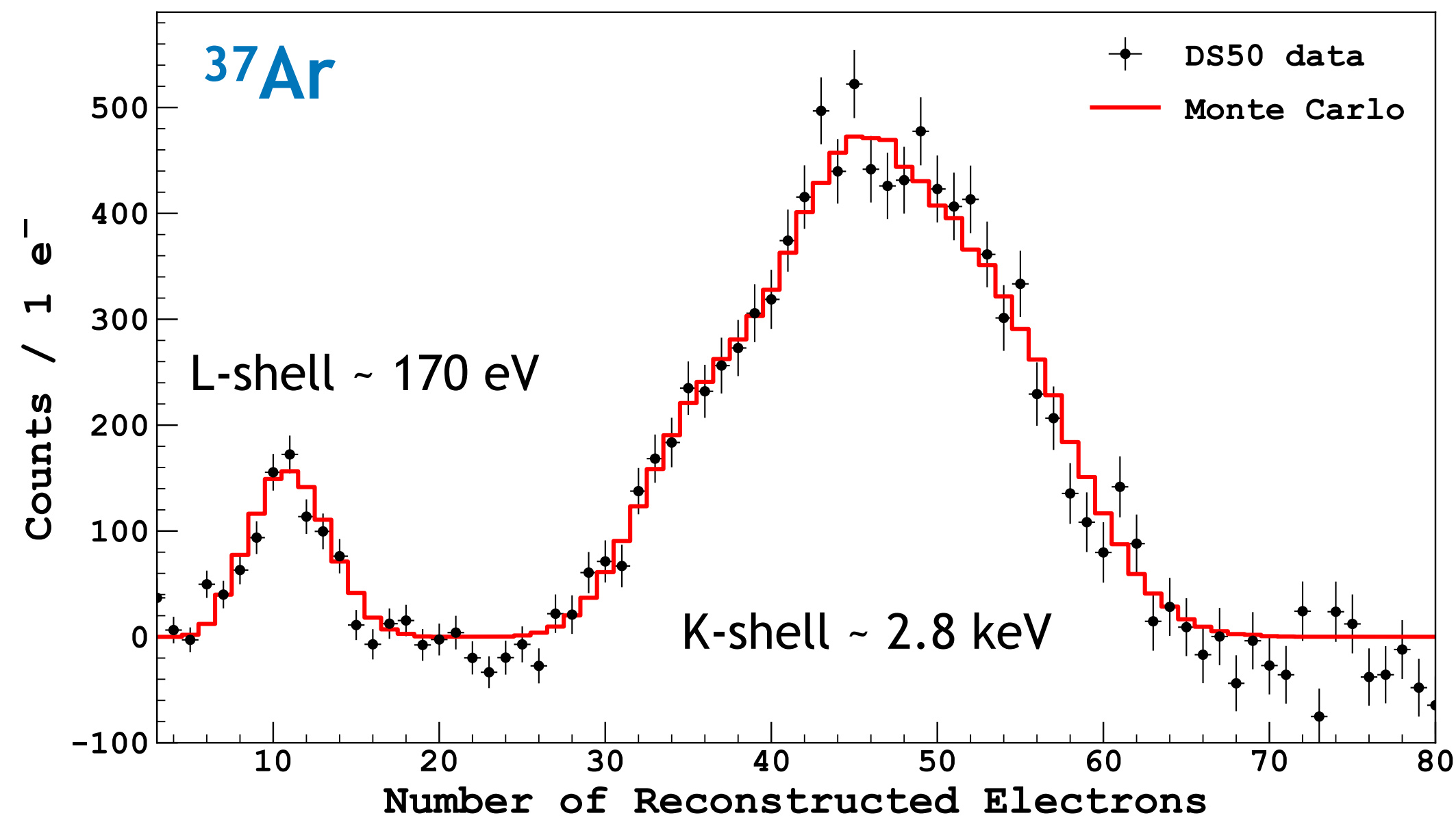
Ionization Only





Thomas-Imel + extended custom model

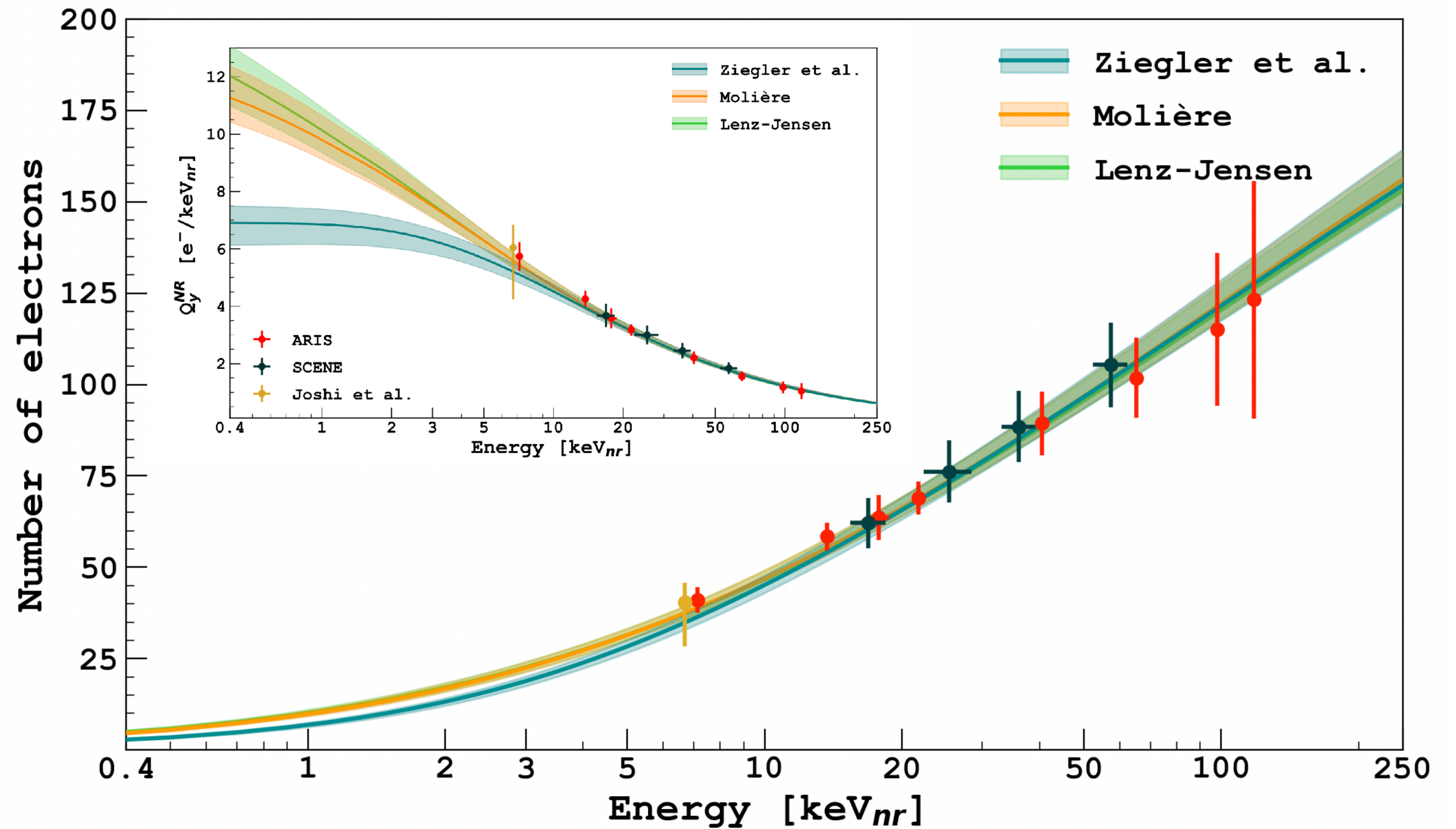
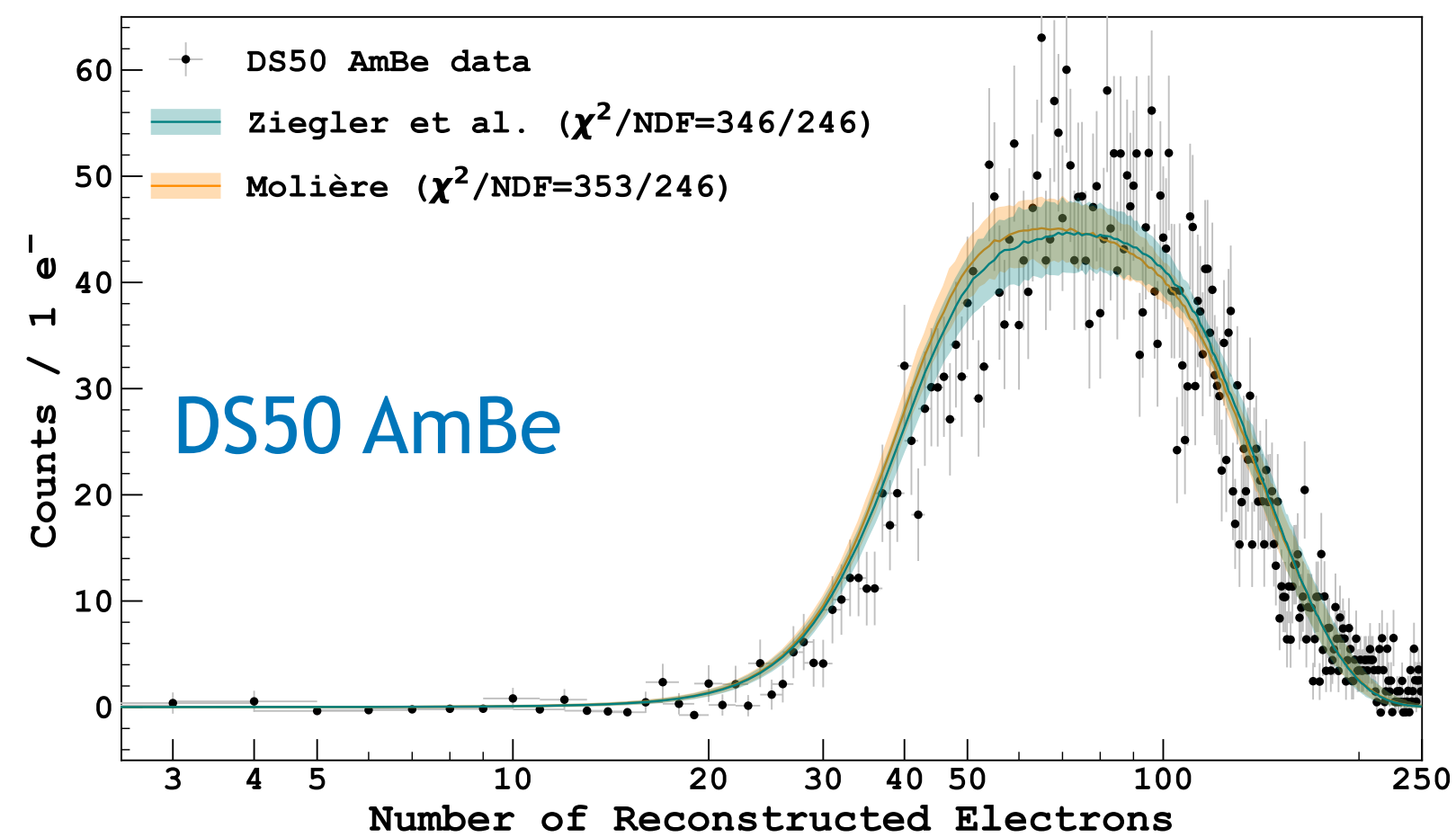
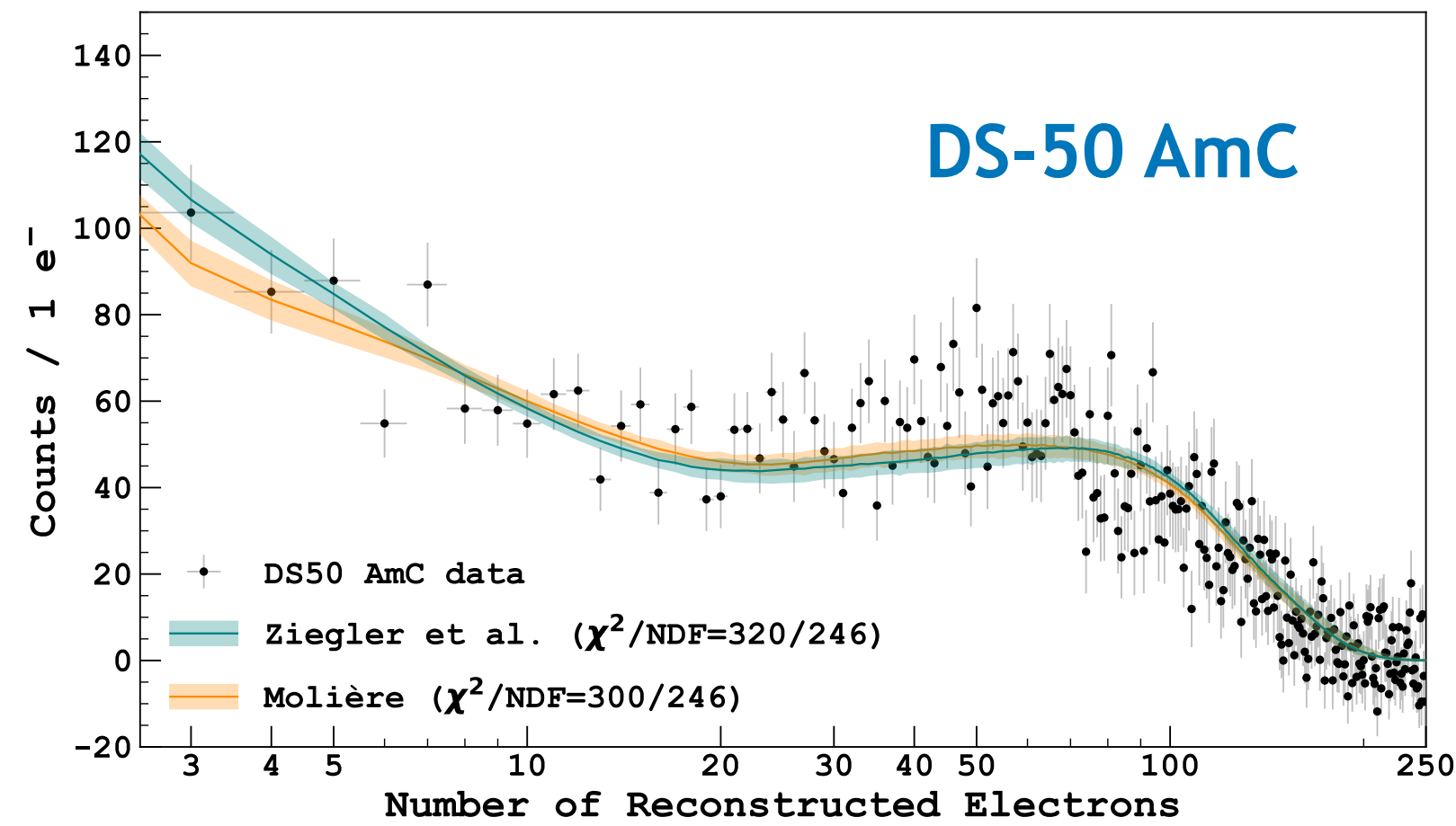
$$Q_y^{ER} = \left(\frac{1}{\gamma} + p_0 \left(\frac{E_{er}}{\text{keV}_{er}} \right)^{p_1} \right) \frac{\ln(1 + \gamma \rho E_{er})}{E_{er}}$$





Ionization Response to Nuclear Recoils

Global fit to DS-50 calibration data with neutrons sources + external datasets (ARIS and SCENE)



Ziegler's screening function is the baseline for this analysis



Dataset and Data Selection

First results on light dark matter candidates with LAr in 2018

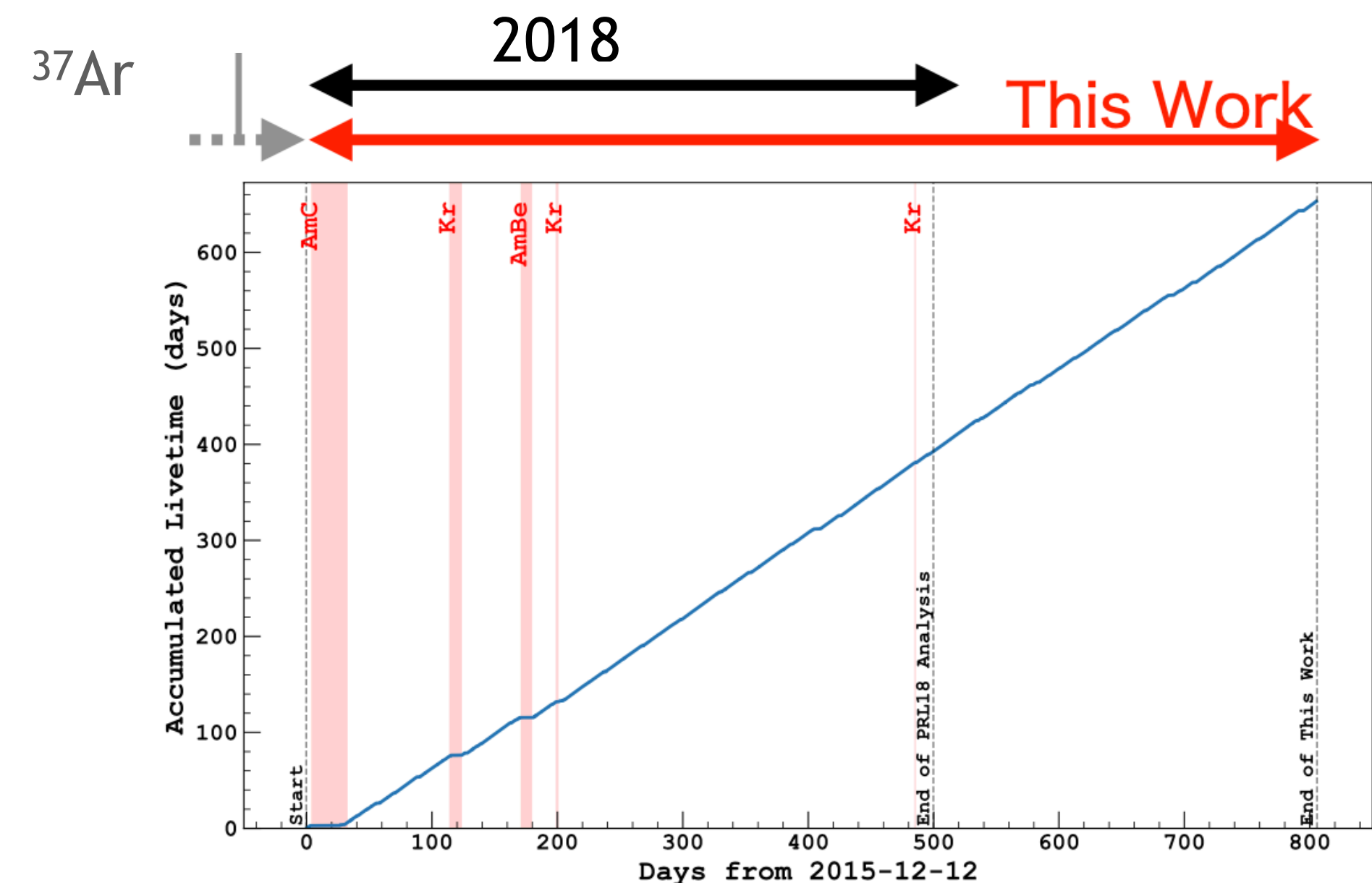
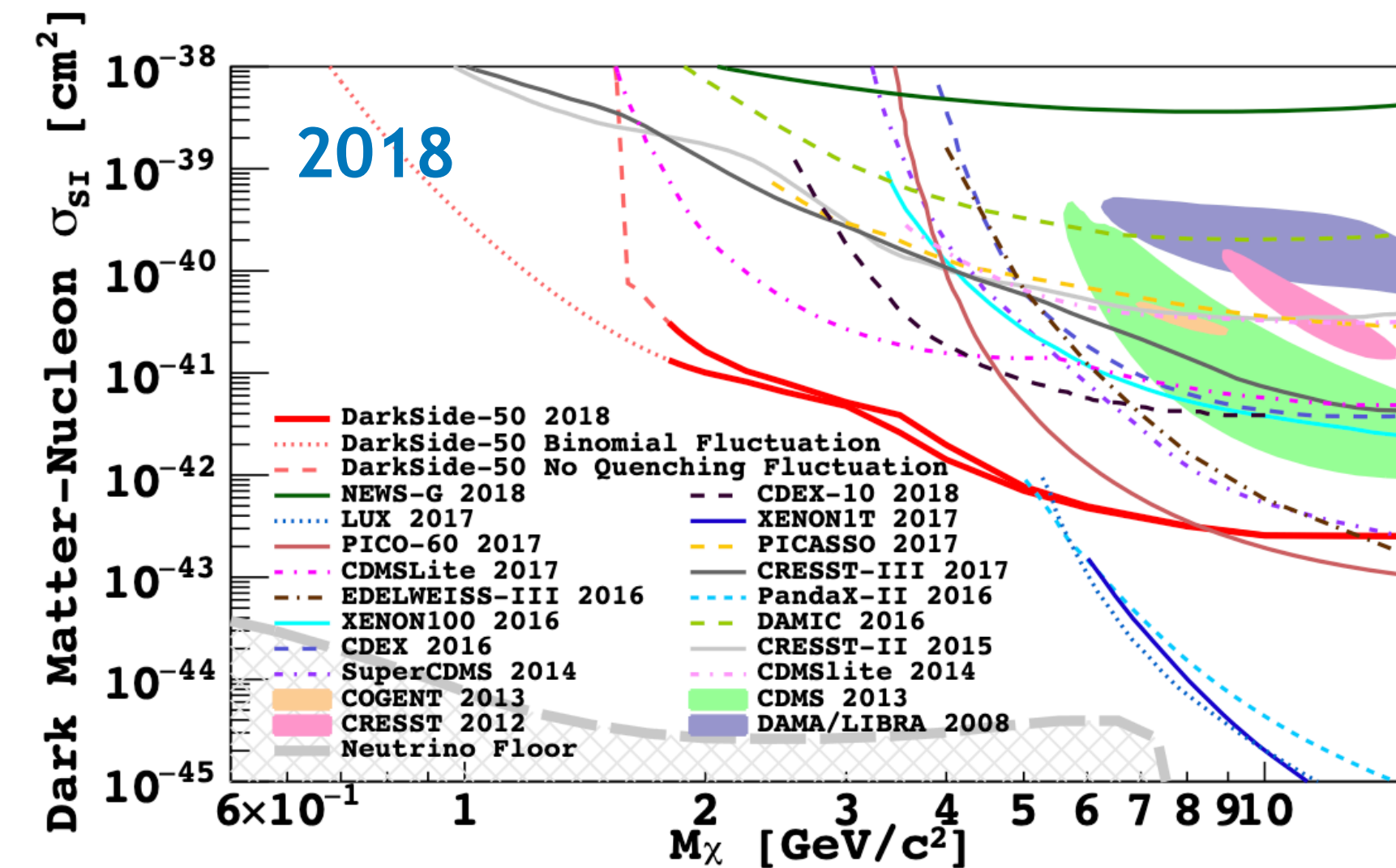
- DarkSide-50, [Phys. Rev. Lett. 121 \(2018\) 081307](#)
- DarkSide-50, [Phys. Rev. Lett. 121 \(2018\) 111303](#)

New exposure

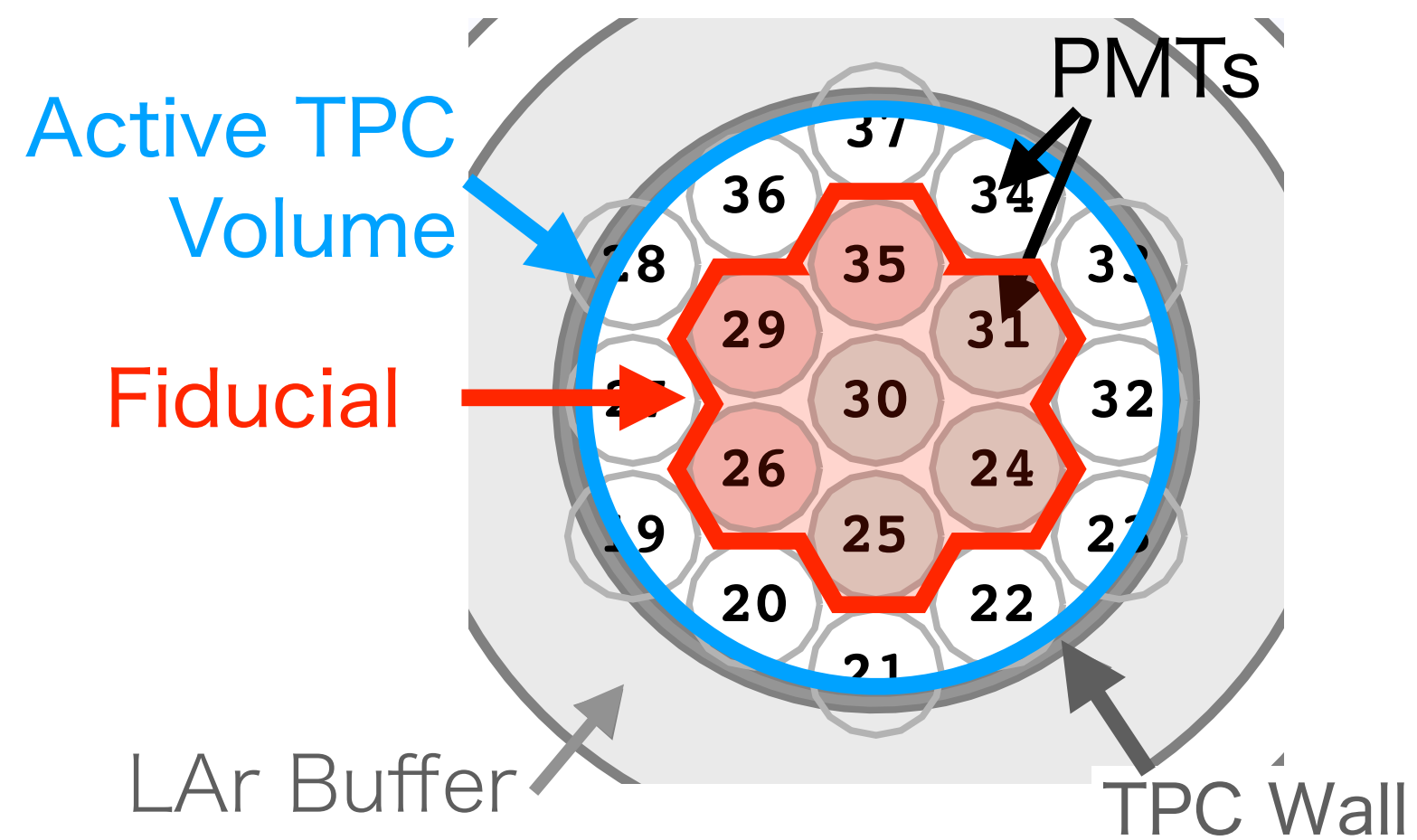
- 650 live-days / 12 ton-day
- x 1.8 exposure used in 2018

Data Selection

- **Quality cuts**
 - Pulse-shape: remove anomalous pulses due to the pile-up of multiple S2's or S1+S2
 - Acceptance: 95% at 4 Ne and 99% at >15 Ne
- **Selection cuts**
 - Fiducialization against external bg
 - S2/S1 against S2's from alphas on the walls
 - Time veto against spurious (or "single") electrons

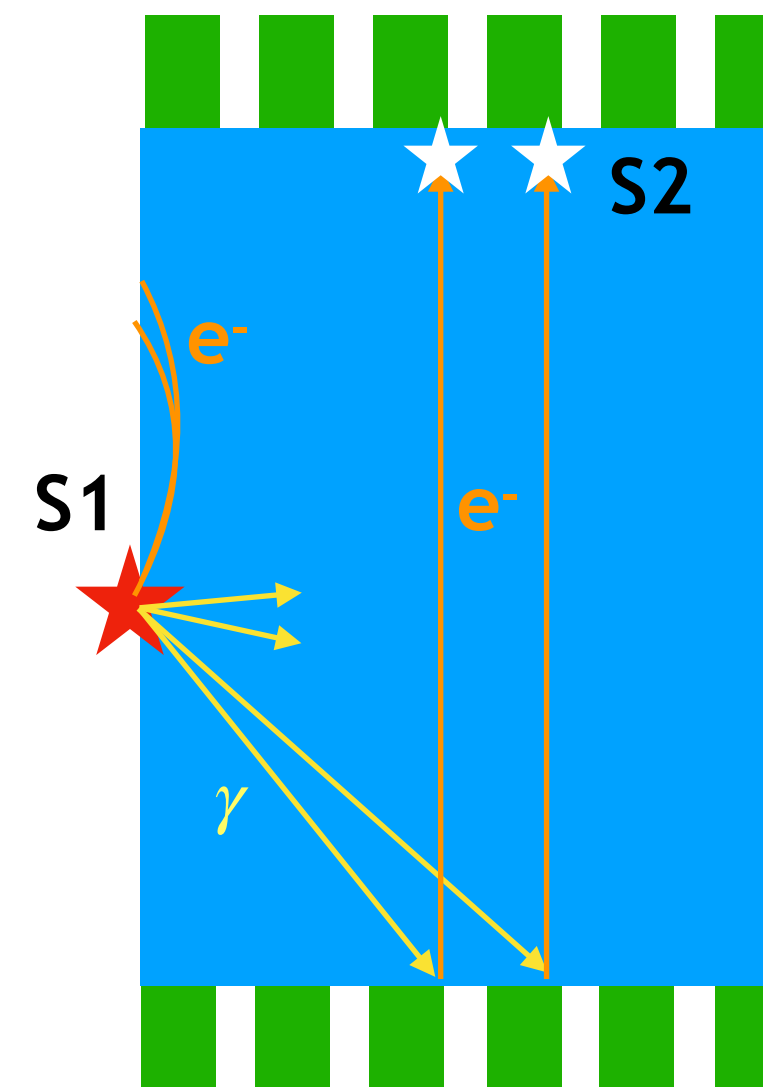


Fiducialization



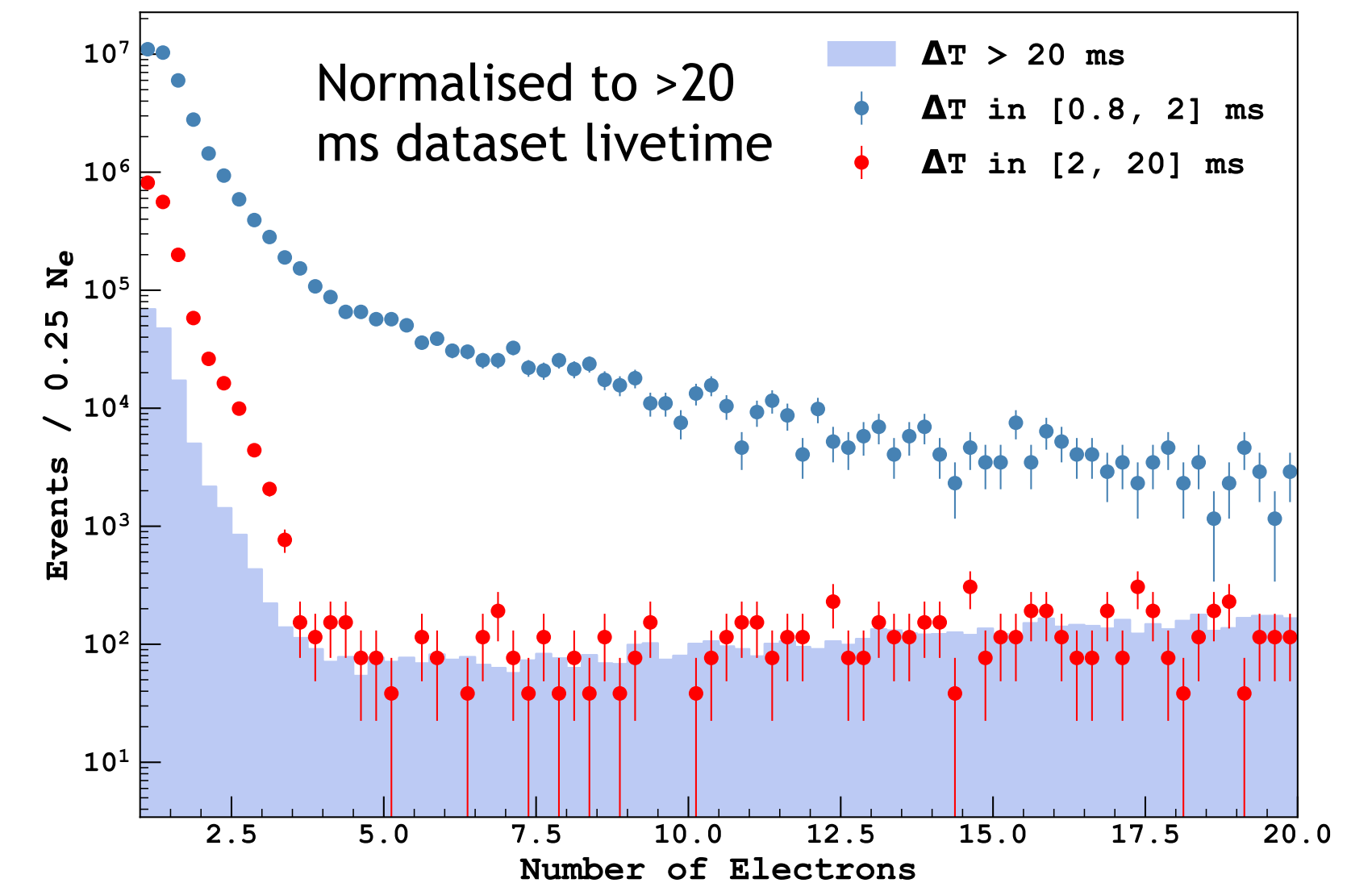
- Select events with max fraction of pes in one of the 7 central top PMTs
- Acceptance ~ 41%

Alpha-induced S2 pulses



- Reject events with “anomalous” S2/S1
- Cut tuned on calibration data
- Acceptance ~ 99%

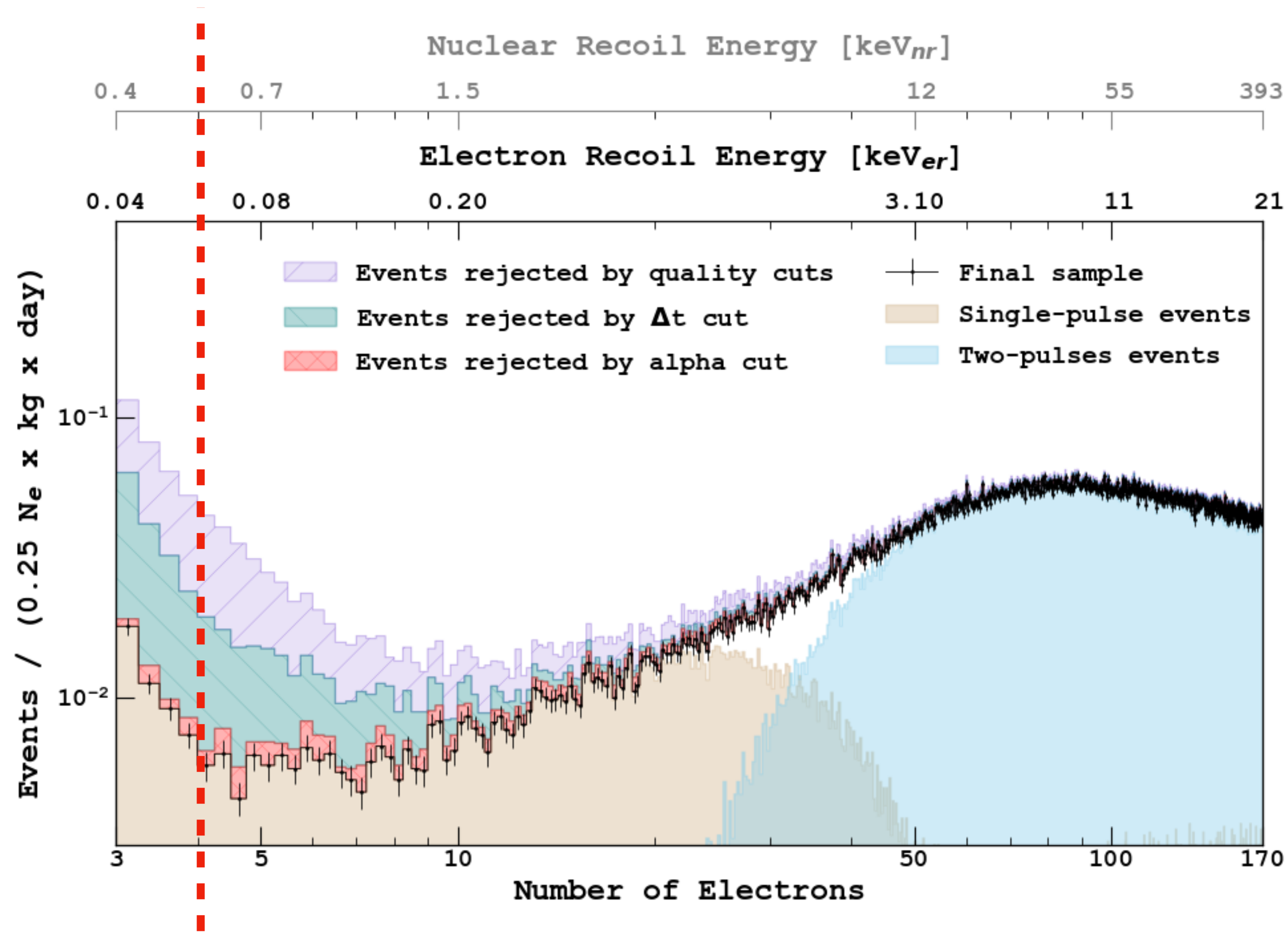
Spurious Electrons



- Reject correlated events (if within 20 ms from the previous one)
- Acceptance ~ 97%

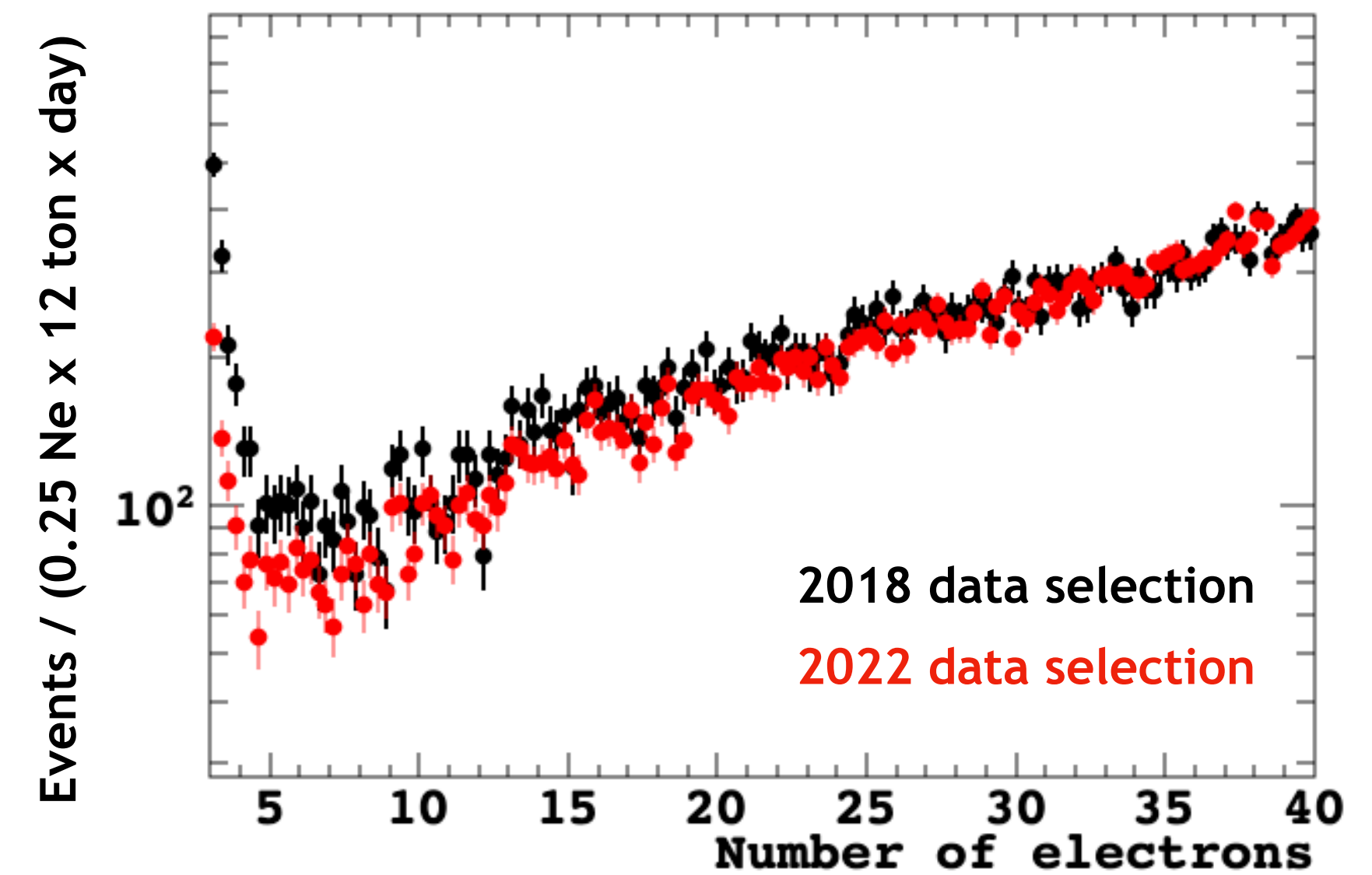


Data Selection



Analysis threshold

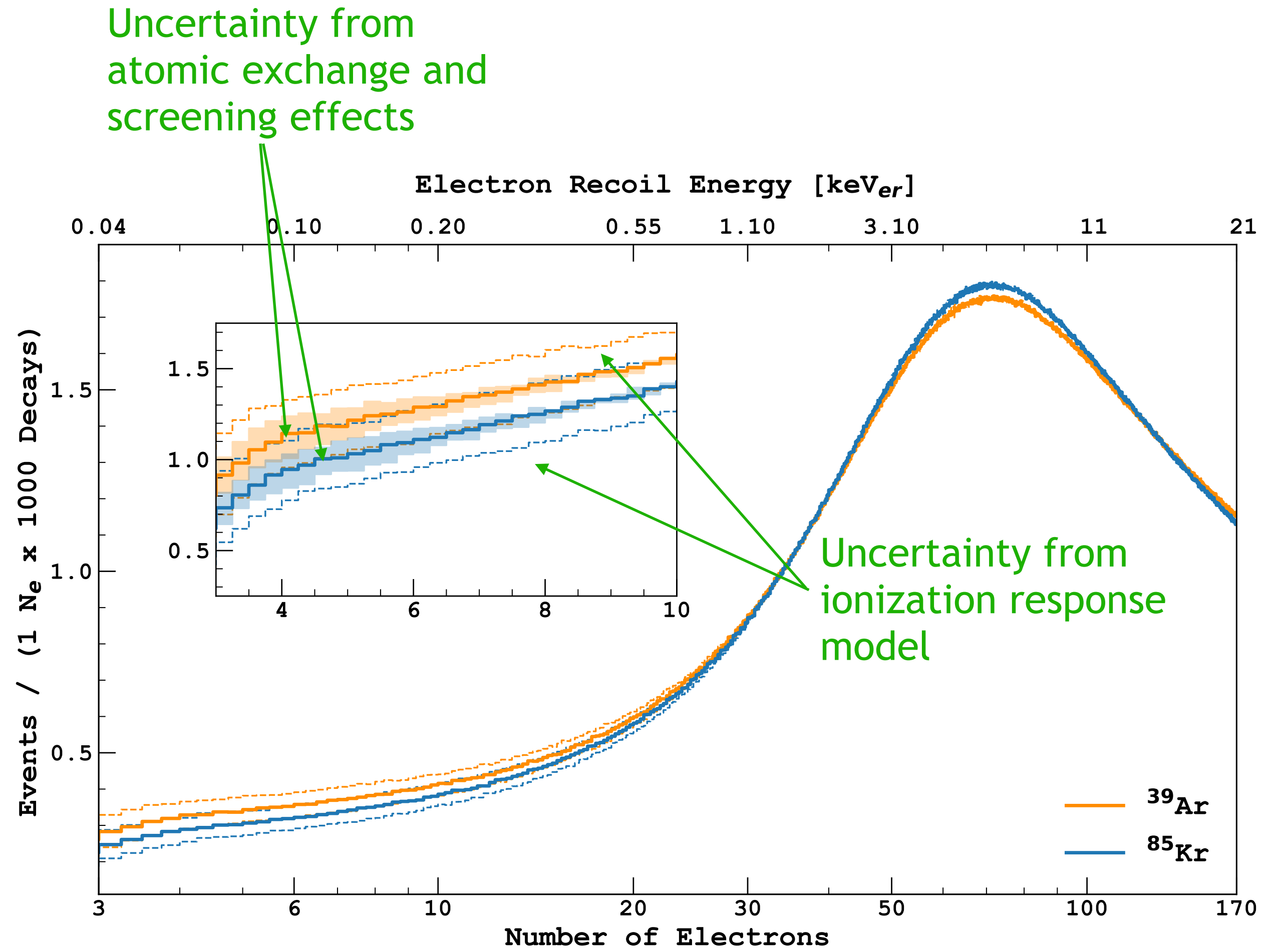
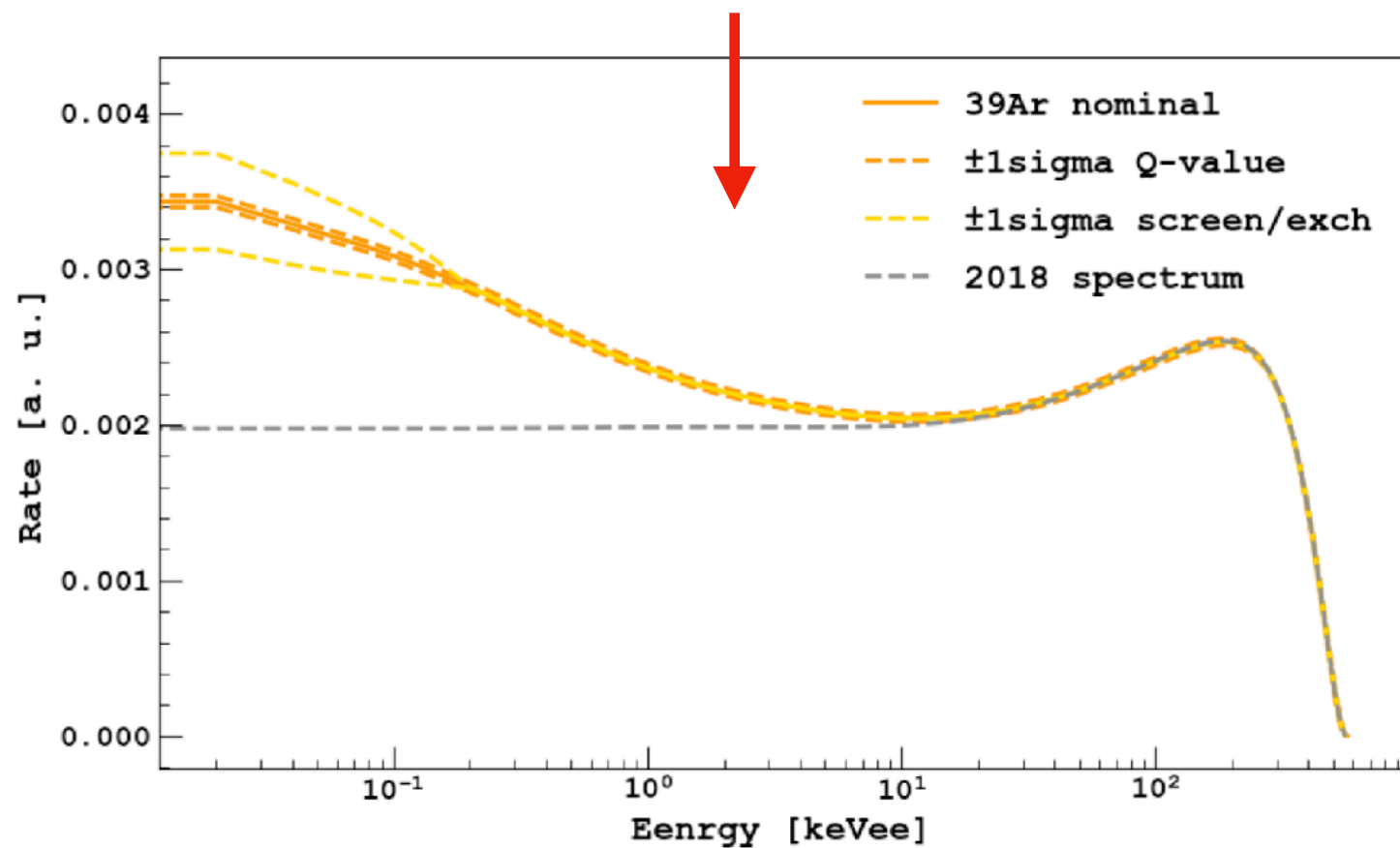
- Overall acceptance almost flat: 38.2% at 4 e⁻ and 40.2% at ≥15 e⁻
- Low-Ne region more depleted than in 2018





Background Model: Internal ^{39}Ar and ^{85}Kr

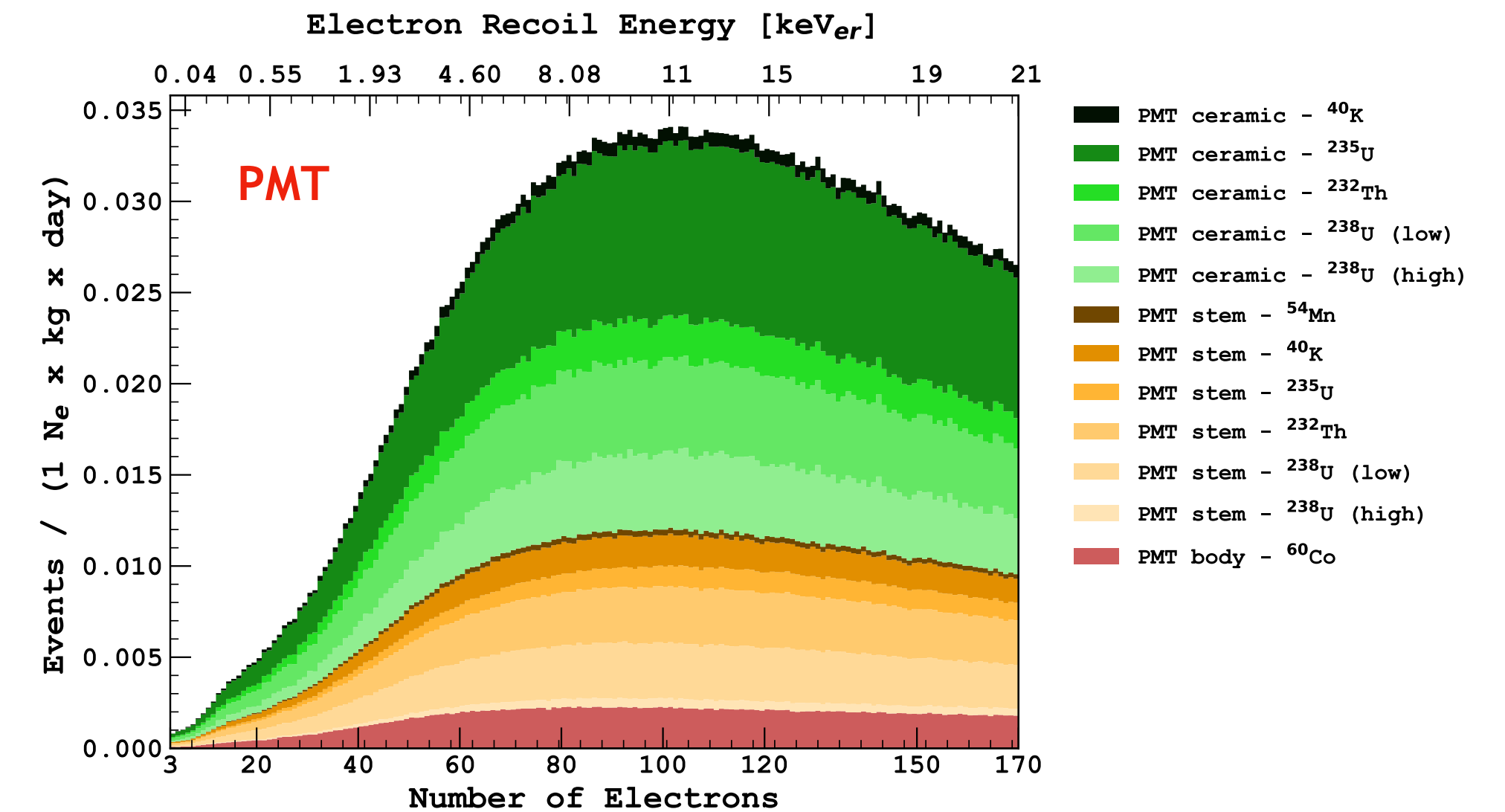
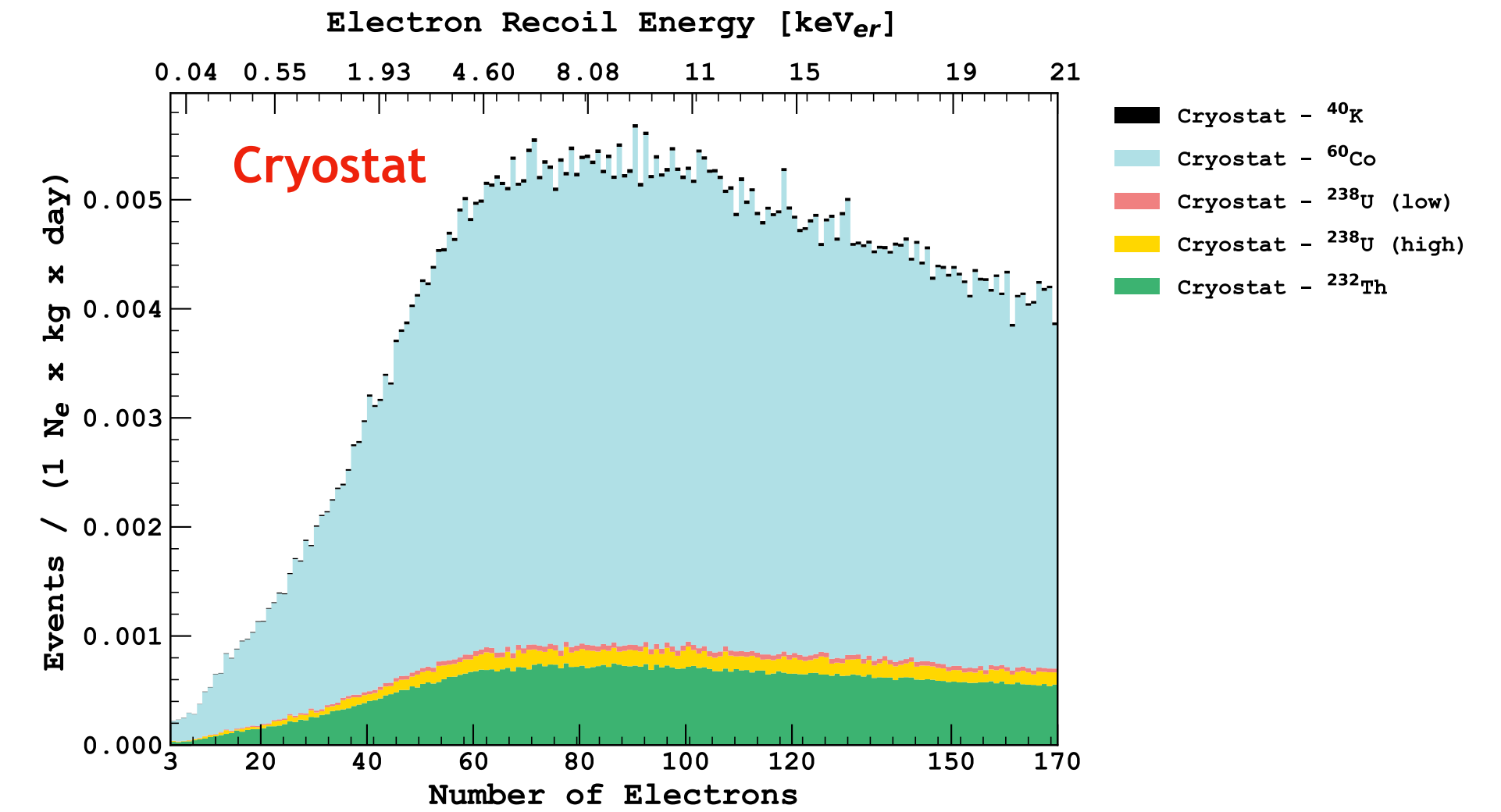
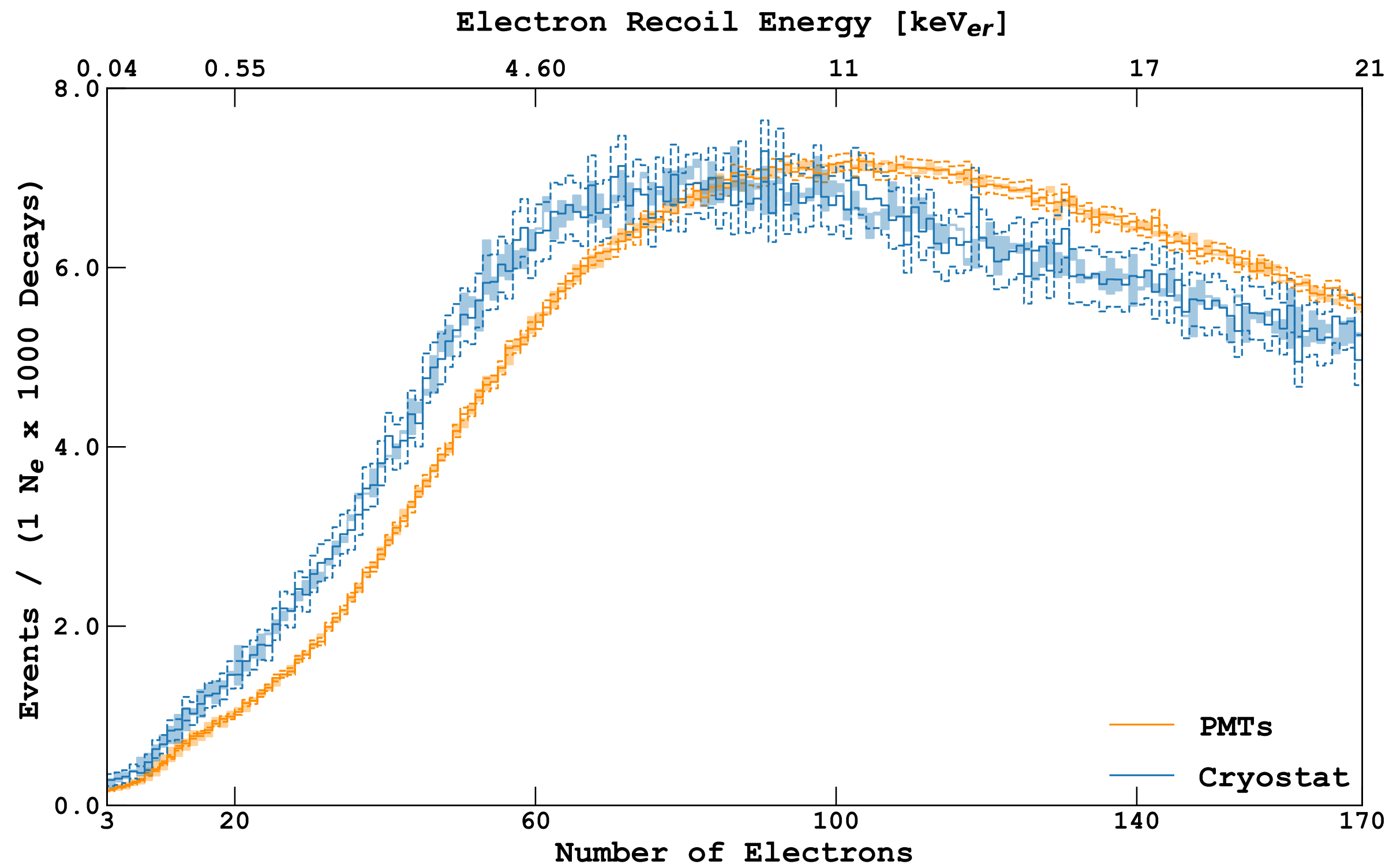
- Both ^{39}Ar and ^{85}Kr uniformly distributed in the LAr bulk
- ^{39}Ar activity: 0.7 ± 0.1 mBq/kg
 - from high energy spectral fit
- ^{85}Kr activity: 1.8 ± 0.1 mBq/kg
 - from high energy spectral fit
 - from fast coincidence through metastable state
 - from decay time fit
- Both unique first-forbidden beta decays: **additional atomic exchange and screening effects**





Background Model: External Gammas

- New background model from material screening campaign
- In 2018 it was extrapolated from the fit of the high energy spectrum





Profile Likelihood and Systematics

$$\mathcal{L} = \prod_{i \in \text{bins}} \mathcal{P}(n_i | m_i(\mu_s, \Theta)) \times \prod_{\theta_i \in \Theta} \mathcal{G}(\theta_i^0 | \theta_i, \Delta\theta_i) \times \prod_{i \in \text{bins}} \mathcal{G}(m_i^0 | m_i(\Theta), \delta m_i(\Theta))$$

Poisson probability of observing n_i events in the i^{th} -bin with respect to the expected ones, $m_i(\mu_s, \Theta)$, with μ_s the signal strength

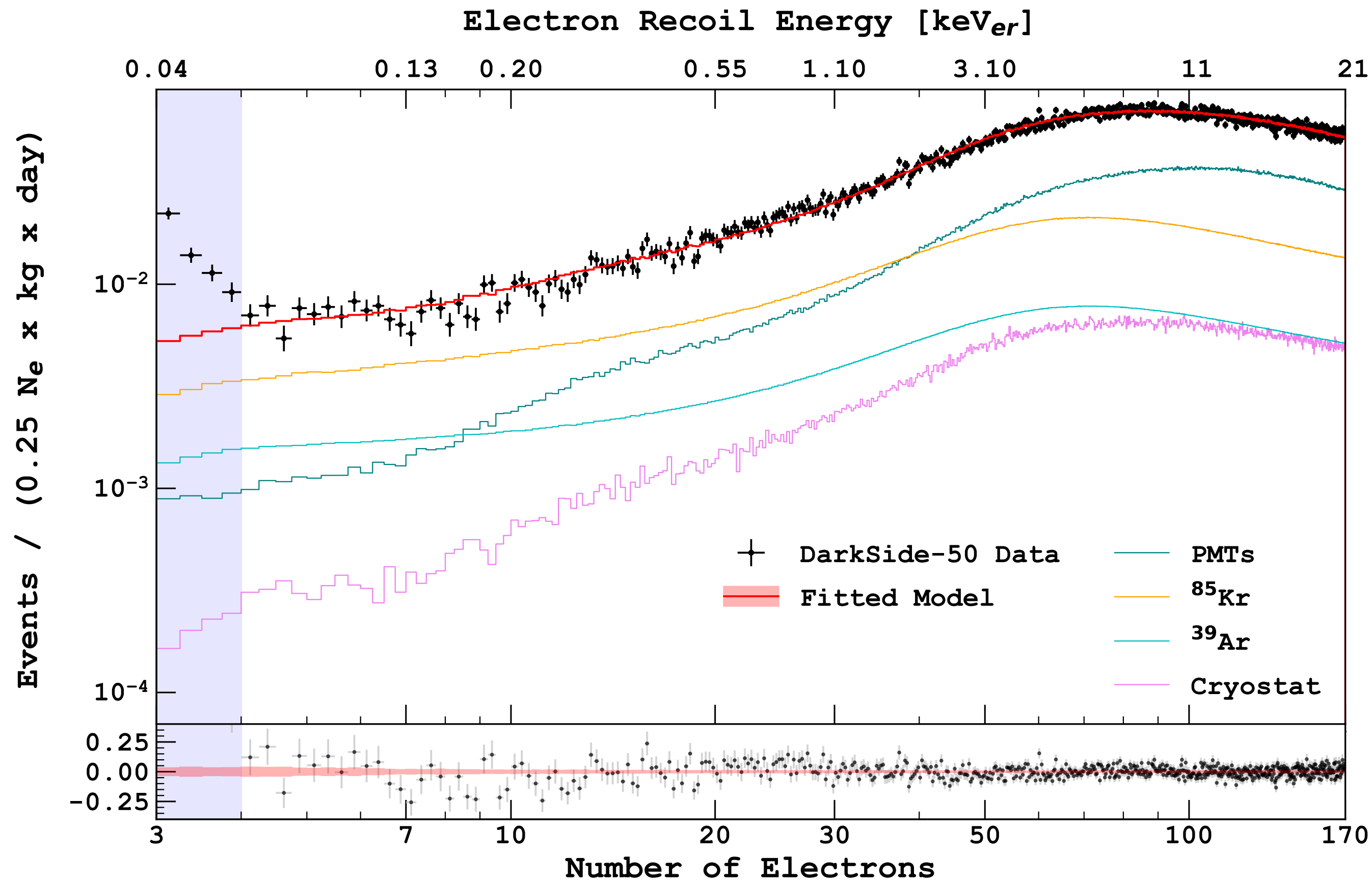
Gaussian penalties to account for the **nuisance parameters** (θ_0 and $\Delta\theta$ are the nominal central values and uncertainties)

Statistical uncertainties of the **simulated** sample

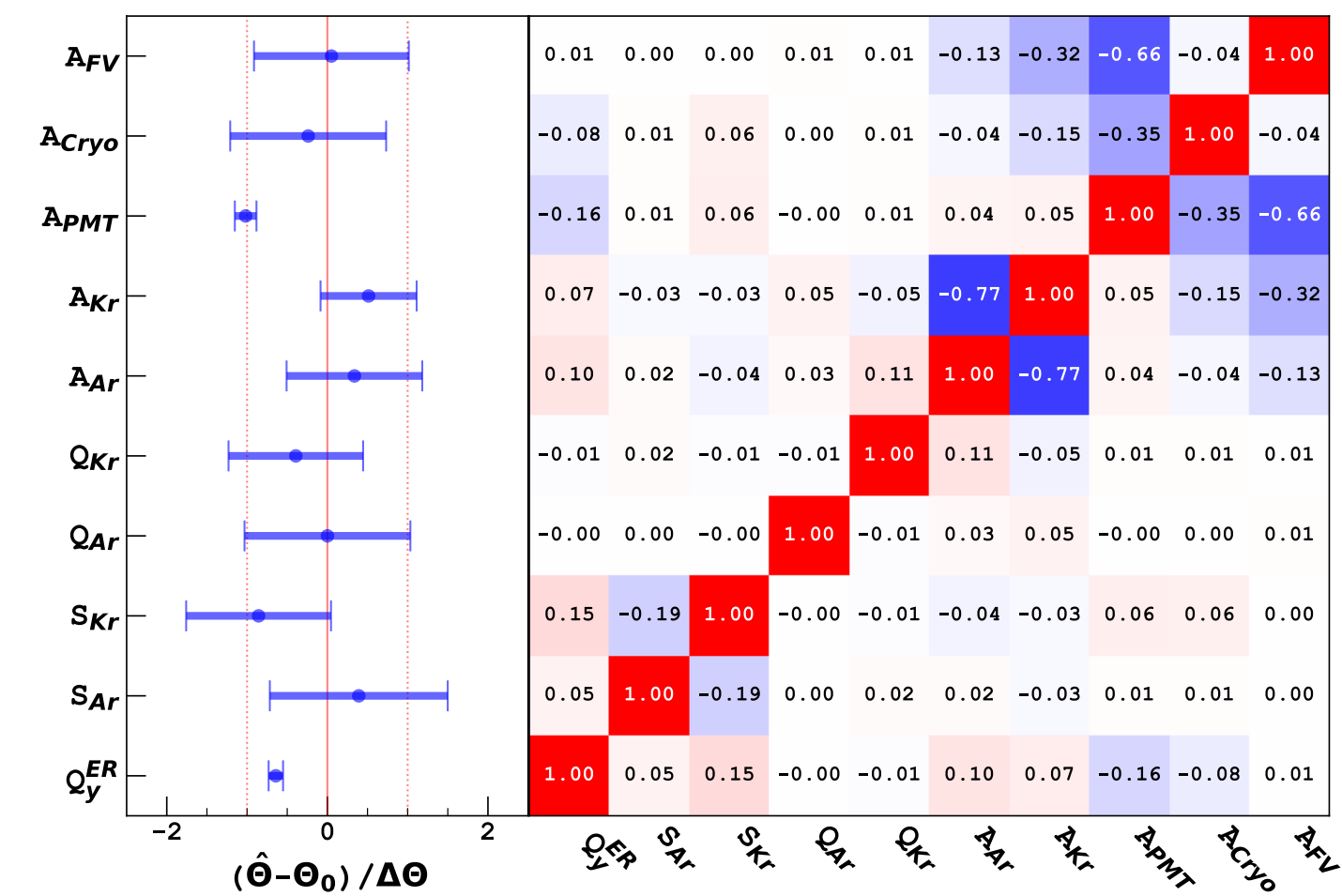
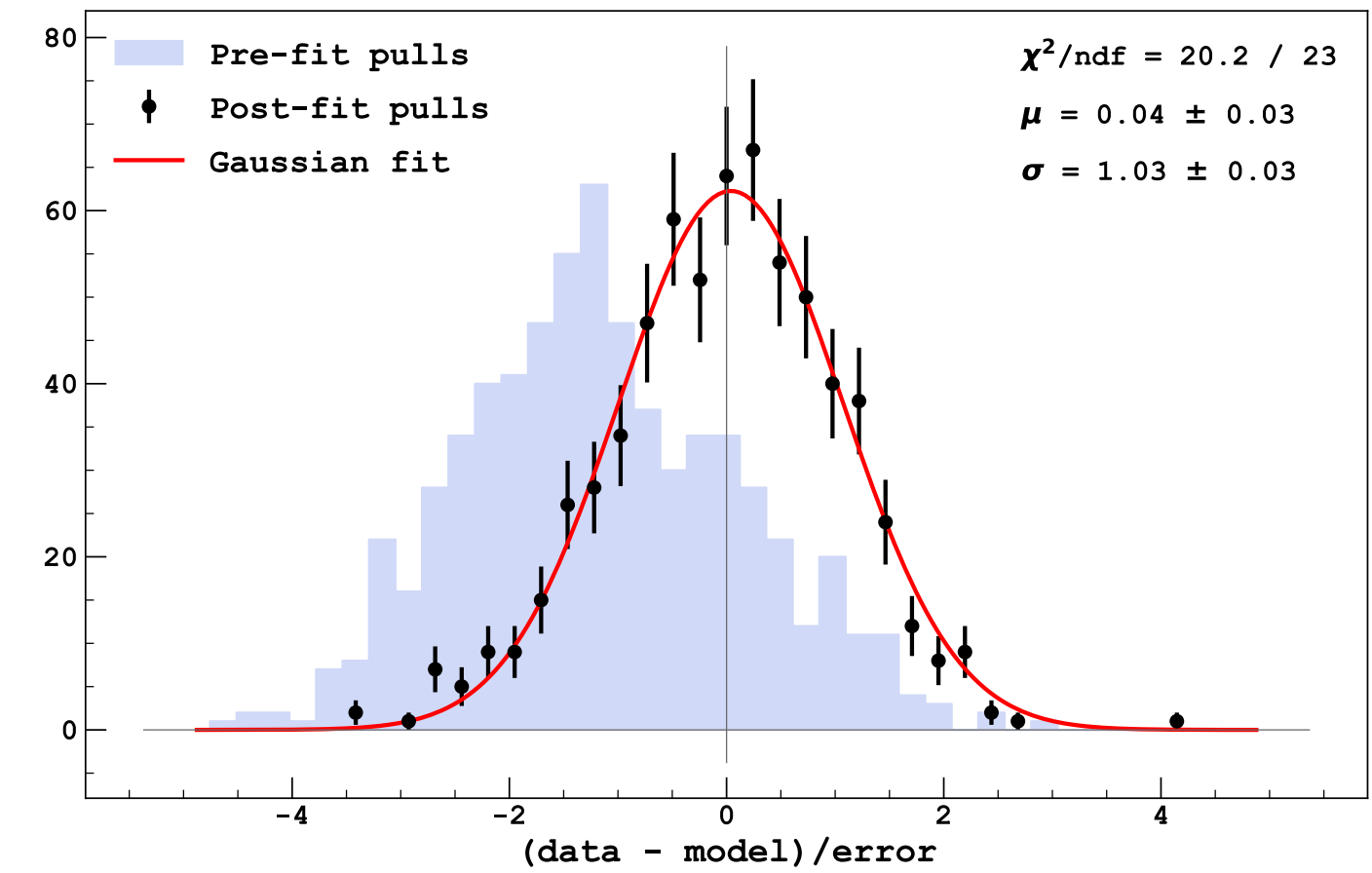
	Name	Source	Affected components
Amplitude	A_{FV}	uncertainty on the fiducial volume	WIMP, ^{39}Ar , ^{85}Kr , PMTs, Cryostat
	A_{Ar}	14.0% uncertainty on ^{39}Ar activity	^{39}Ar
	A_{Kr}	4.7% uncertainty on ^{85}Kr activity	^{85}Kr
	A_{pmt}	11.5% uncertainty on activity from PMTs	PMT
	A_{cryo}	6.6% uncertainty on activity from the cryostat	Cryostat
Shape	Q_{Kr}	0.4% uncertainty on the ^{85}Kr -decay Q-value	^{85}Kr
	Q_{Ar}	1% uncertainty on the ^{39}Ar -decay Q-value	^{39}Ar
	S_{kr}	spectral shape uncertainty on atomic exchange and screening effects	^{85}Kr
	S_{Ar}	spectral shape uncertainty on atomic exchange and screening effects	^{39}Ar
	Q_y^{er}	spectral shape systematics from ER ionization response uncertainty	^{39}Ar , ^{85}Kr , PMTs, Cryostat
	Q_y^{nr}	spectral shape systematics from NR ionization response uncertainty	WIMP

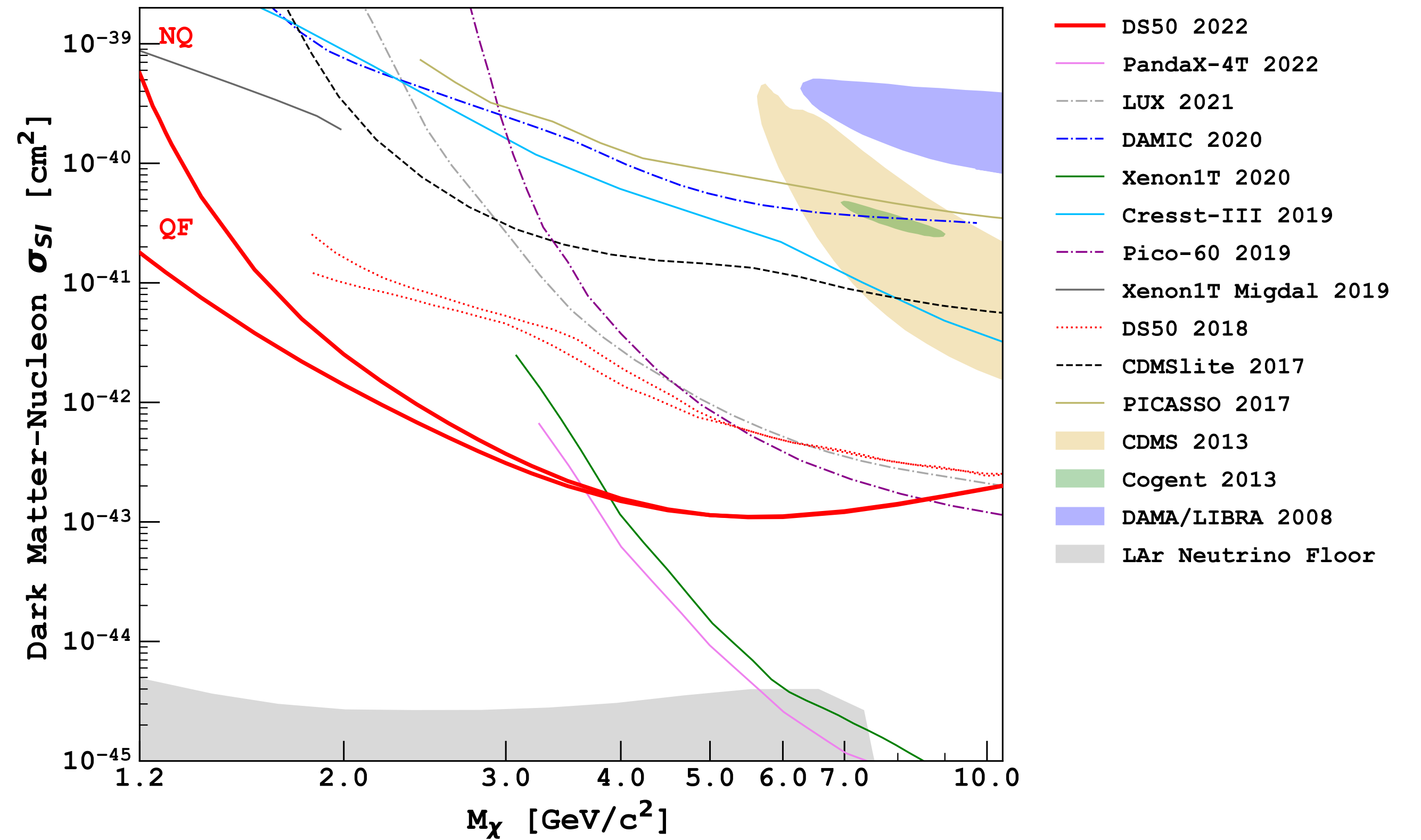
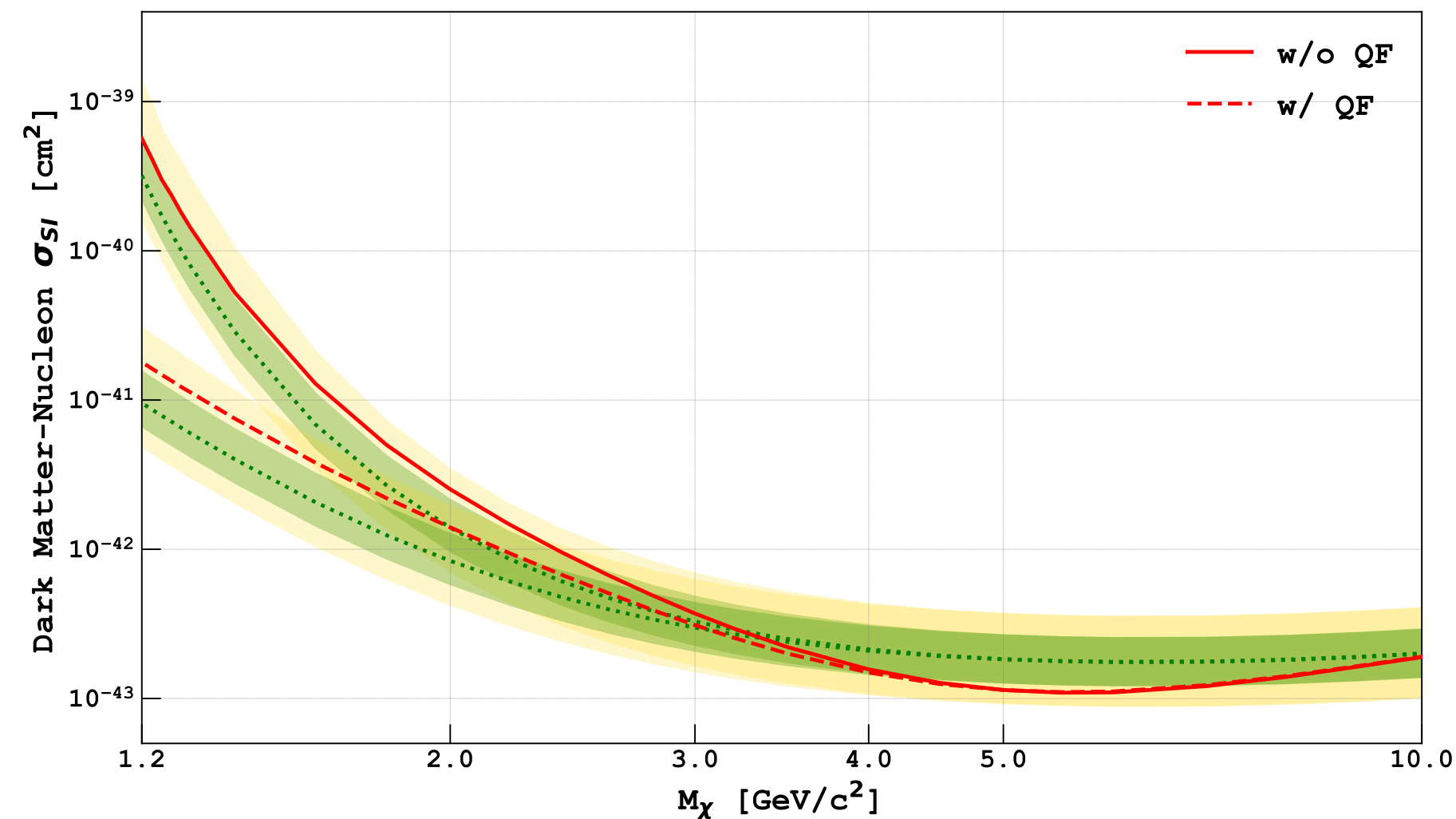
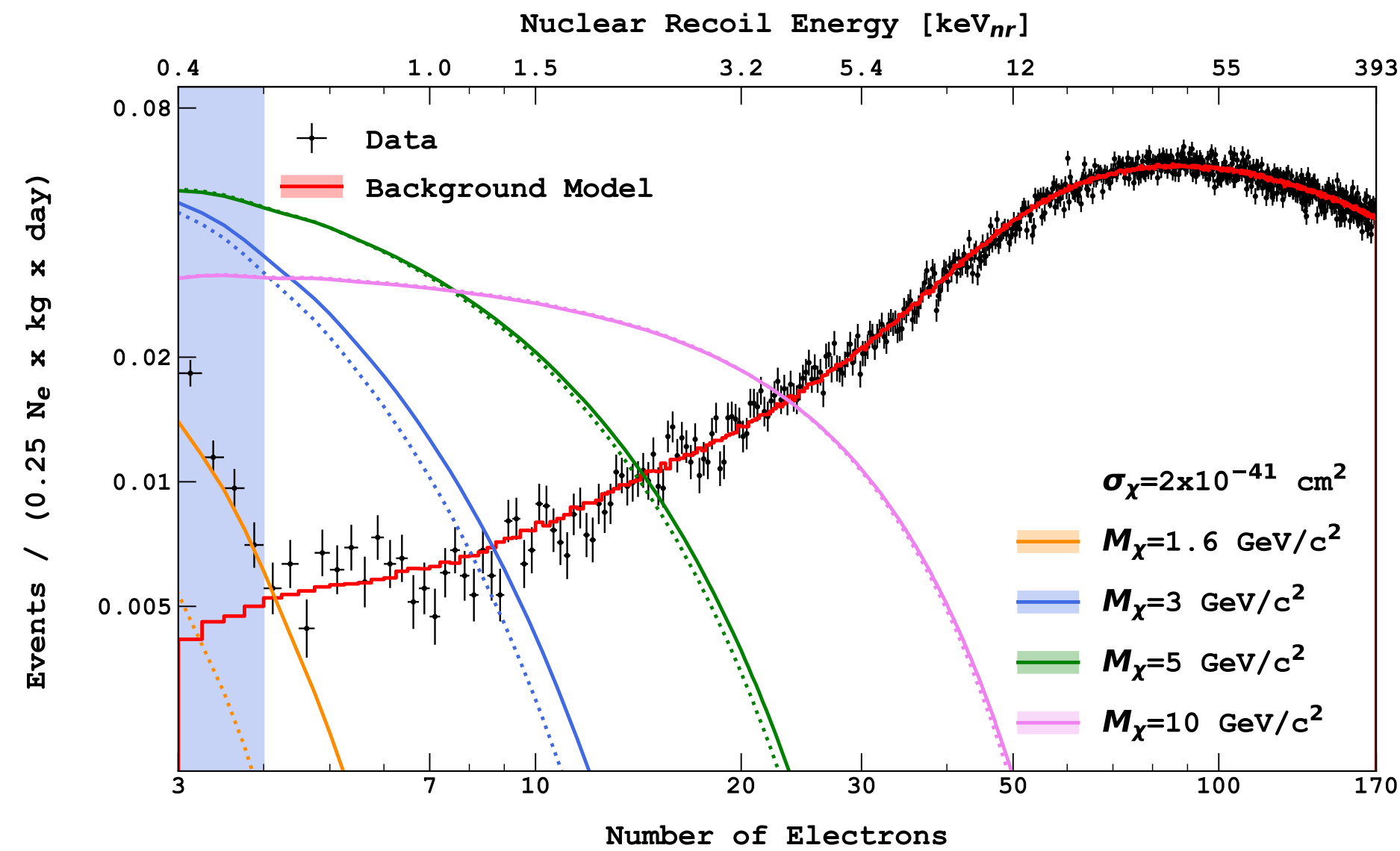


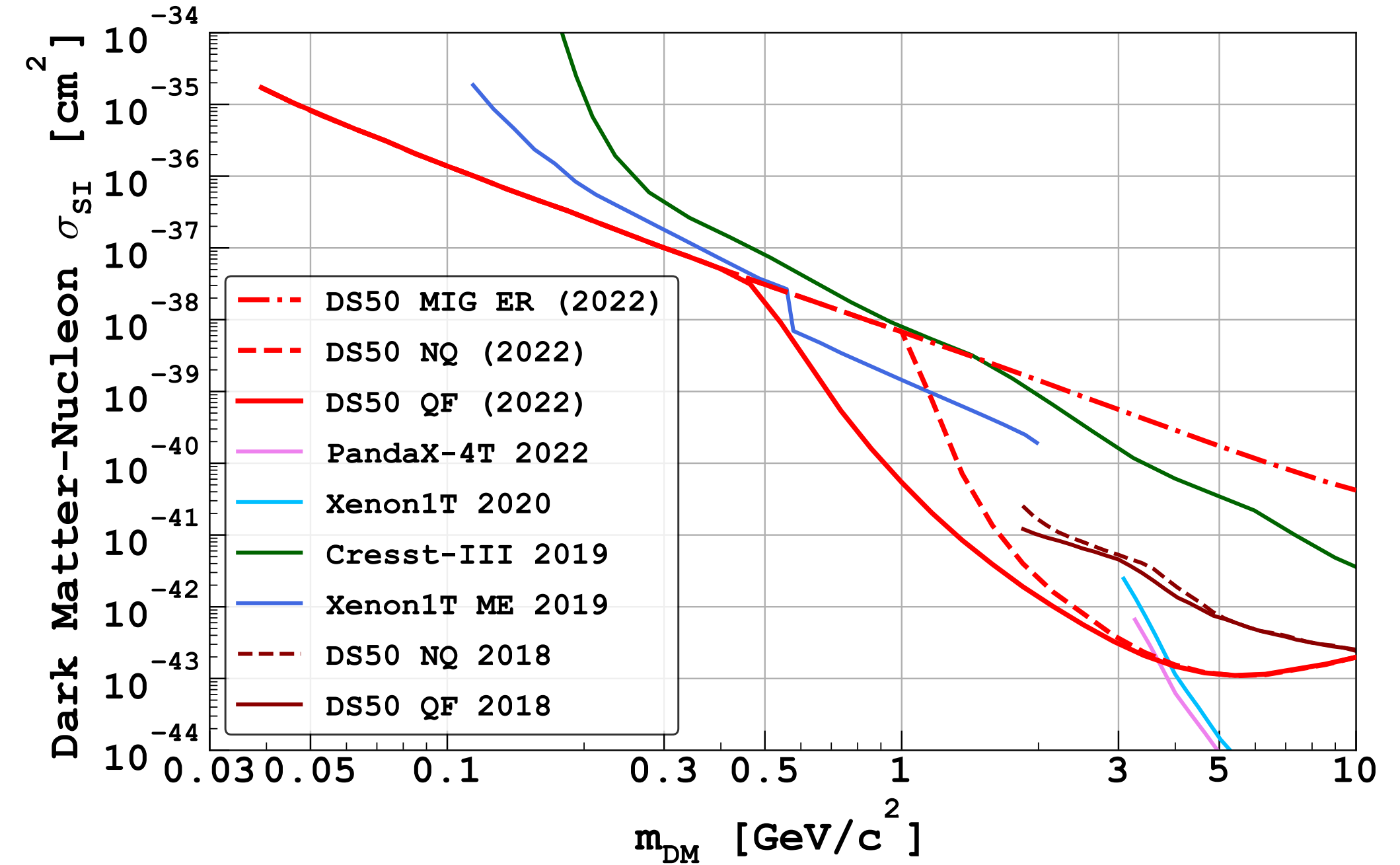
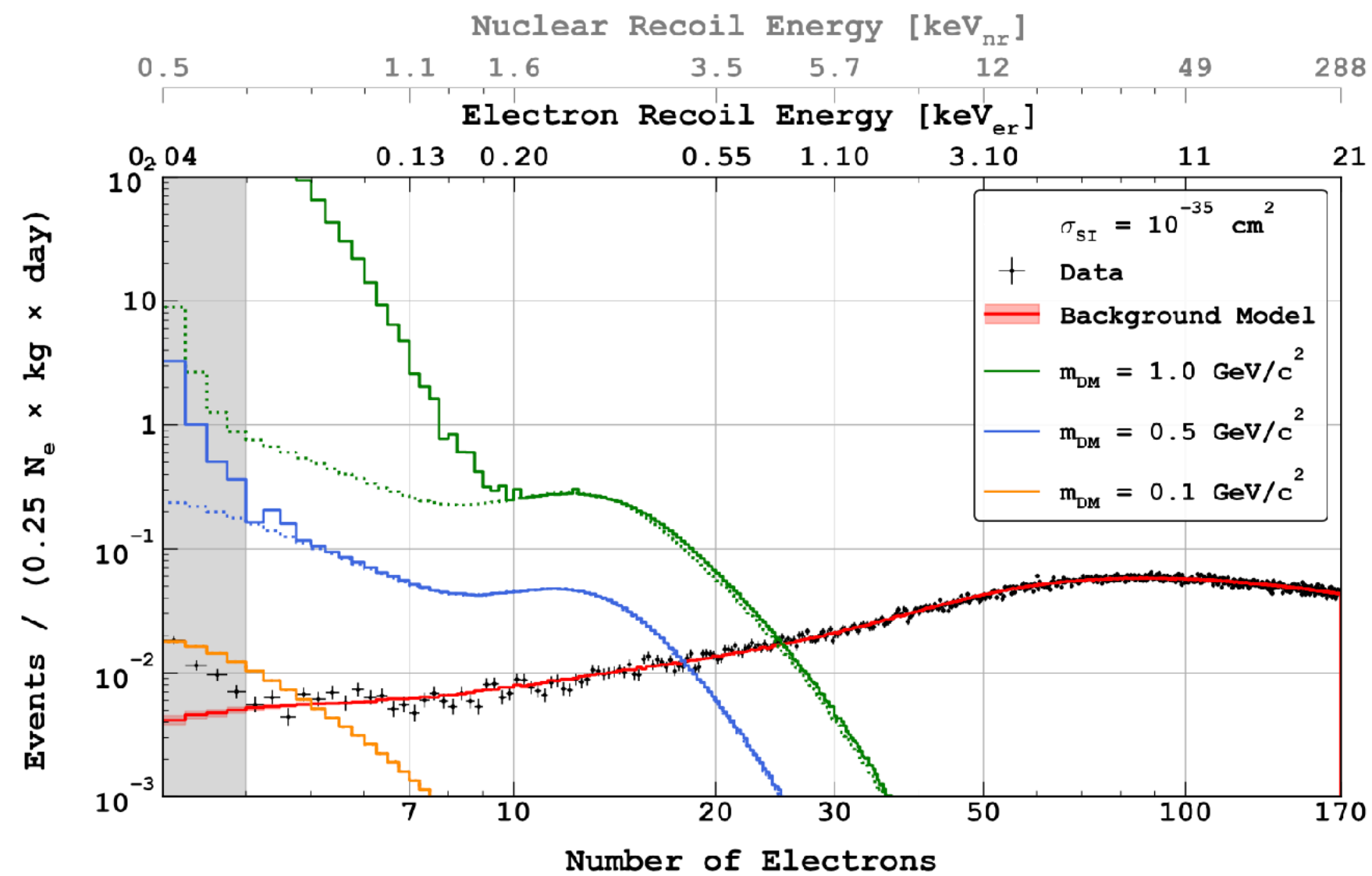
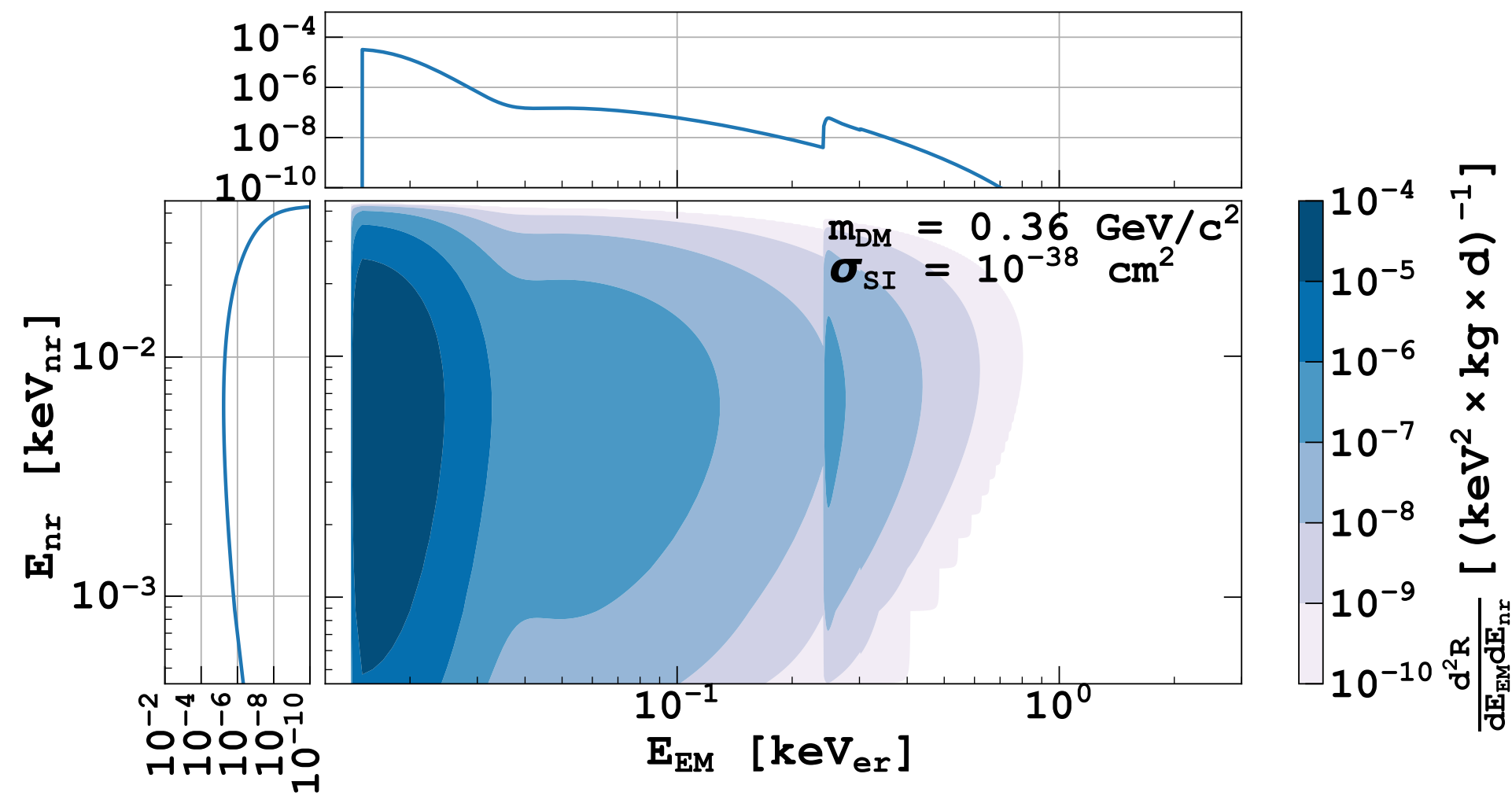
Background-Only Fit



Tritium activity < 1 $\mu\text{Bq/kg}$ (90% CL)

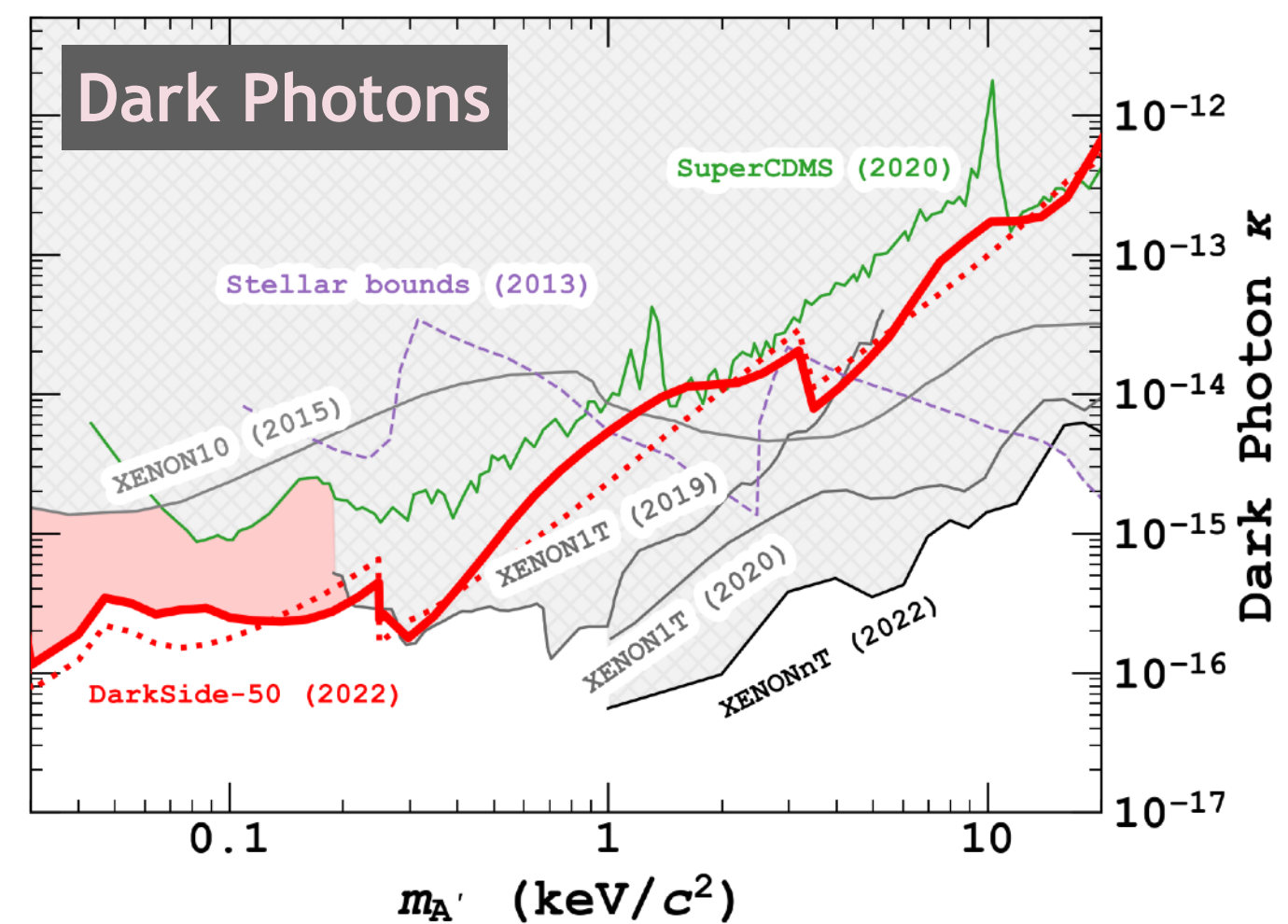
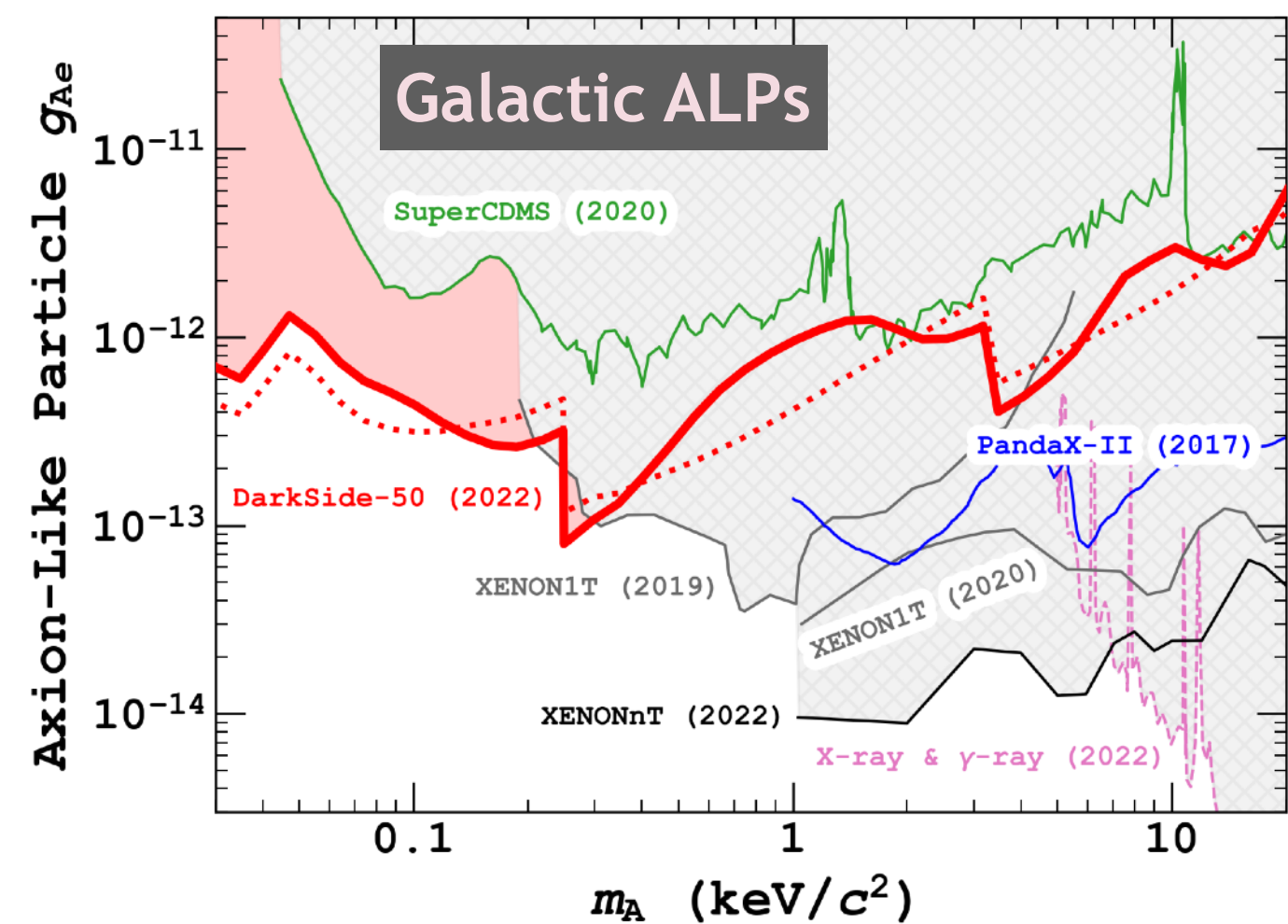
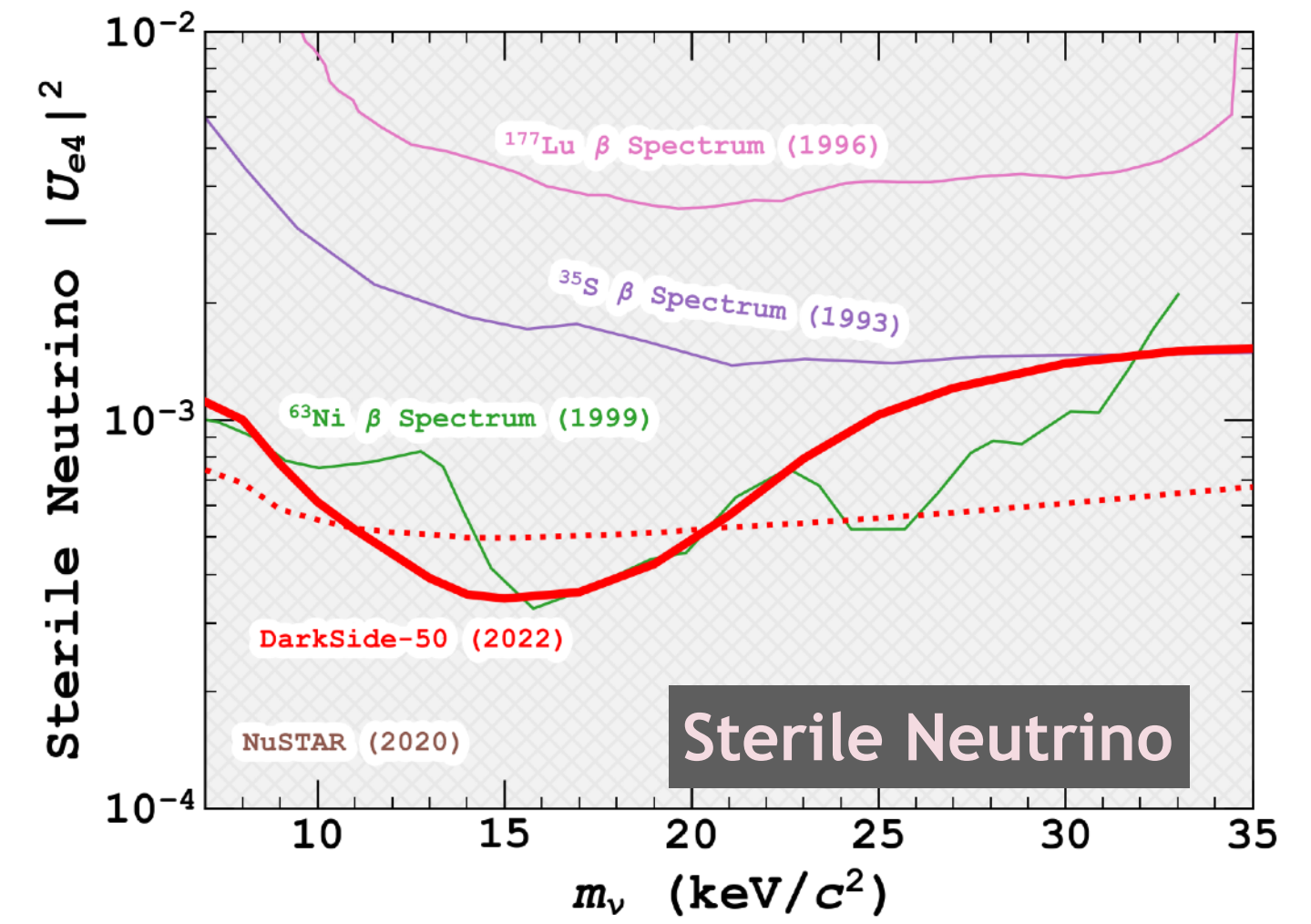
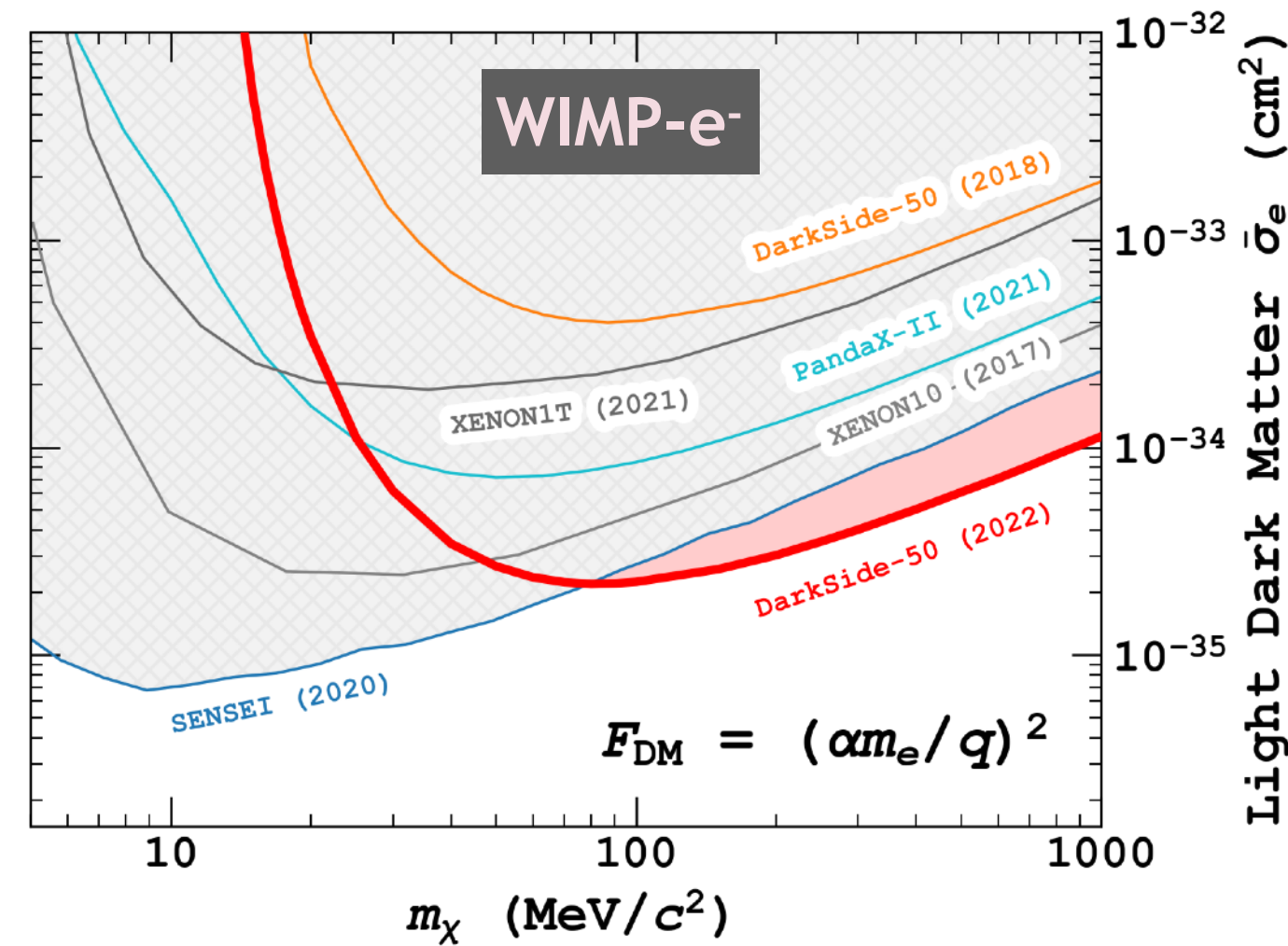
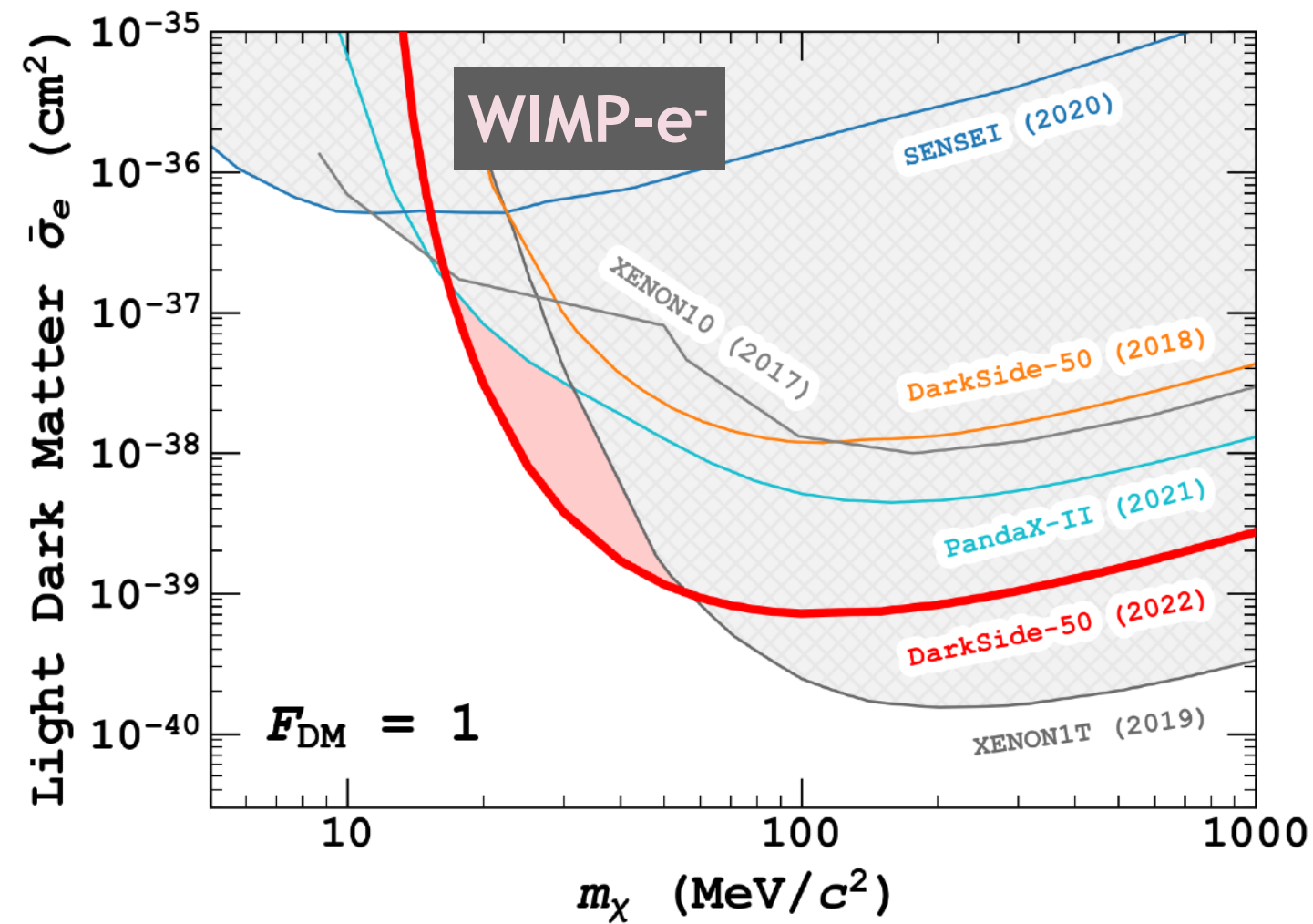






Results confirmed by using **Bayesian Networks** ([arXiv:2302.01830](https://arxiv.org/abs/2302.01830))

- detector response model included in the likelihood function
- Markov Chain Monte Carlo for posterior probability





Conclusions

The DarkSide-50 low-mass search

- Improved light dark matter limits from 2018 analysis thanks to:
 - **Calibration** of ionization response to ERs and NRs down to <1 keV
 - Extended **exposure**
 - Better **data selection**
- Best SI WIMP-nucleon limits down to $1.2 \text{ GeV}/c^2$ ($40 \text{ MeV}/c^2$) WIMP mass without (with) Migdal effect
- Improved limits on **WIMP-electron** interactions, galactic **ALPs**, **dark photons**, and **sterile neutrinos**
- More results on **annual modulations** and **non-standard operators** in progress

Follow up of the DarkSide-50 search

- Sensitivity projection for a **1 ton-year exposure** dual-phase LAr TPC optimized for light dark matter searches through the ionization channel ([arXiv:2209.01177](https://arxiv.org/abs/2209.01177))
- Sensitivity projections for **DarkSide-20k** in progress (see Marie Van Uffelen's talk for more details on DarkSide-20k)

