

Standard and Non-Standard Neutrino oscillations with NOvA

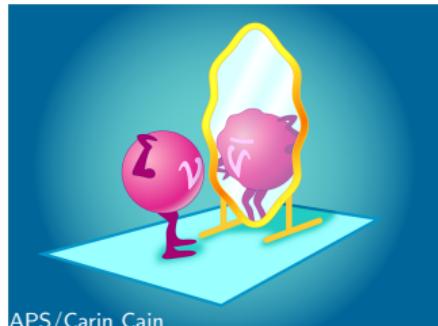
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on behalf of NOvA
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19 March 2023



1. Neutrino Oscillations
2. NOvA experiment
3. Bayesian Inference into the PMNS model
4. Sterile neutrinos

Open Questions



Do neutrinos violate
Charge-Parity?



Do we understand ν
propagation in
matter?



Is there a light, **sterile**
neutrino?

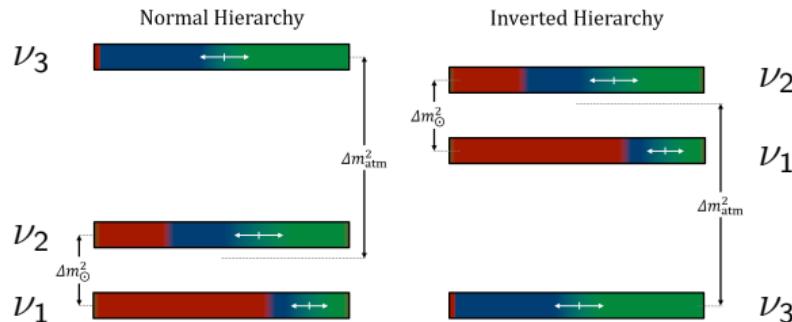


How are the **neutrino mass states ordered?**



How ν_μ/ν_τ mix into the
mass states?

Neutrino oscillation physics



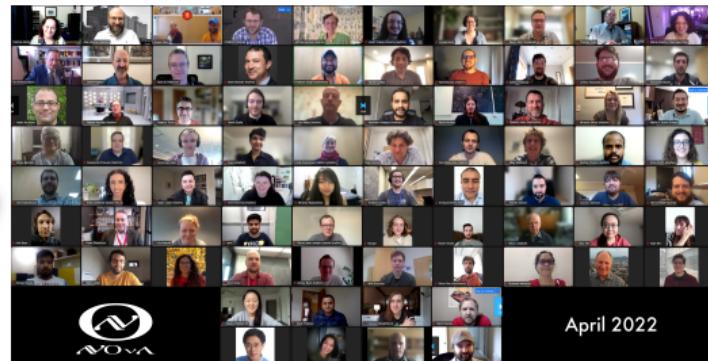
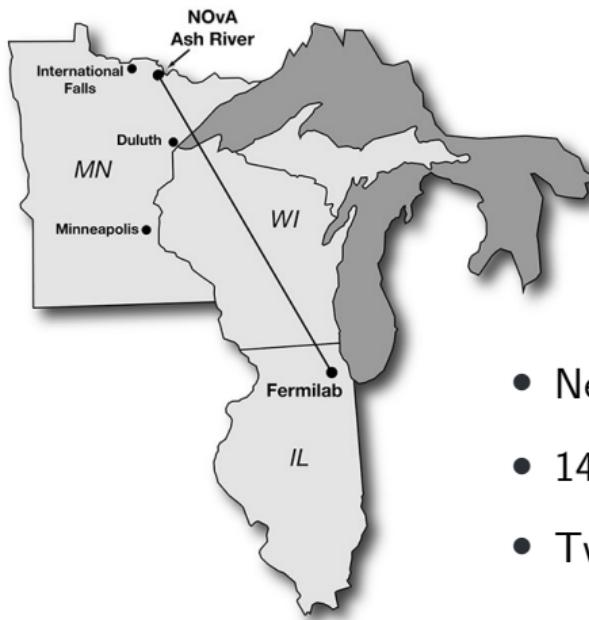
- Flavour eigenstates; ν_e , ν_μ and ν_τ (interact)
- Mass eigenstates; ν_1 , ν_2 and ν_3 (propagate)

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\substack{\text{atmospheric, beam} \\ \text{Super-K, IceCube,} \\ \text{Opera, NOvA, T2K}}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix}}_{\substack{\text{reactor, beam} \\ \text{Double Chooz, Daya Bay,} \\ \text{RENO, NOvA, T2K}}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\substack{\text{solar, reactor} \\ \text{Super-K, SNO,} \\ \text{KamLAND}}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

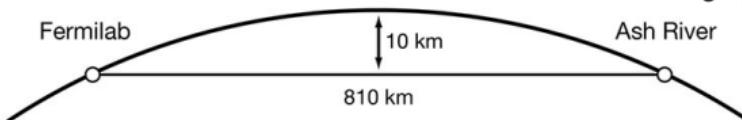
$s_{ij} = \sin \theta_{ij}$
 $c_{ij} = \cos \theta_{ij}$

- θ_{23} : Larger or smaller than 45? Important for $\nu_\tau - \nu_\mu$ symmetries.
- δ_{CP} : Potential contribution to matter-antimatter asymmetry in the universe.
- $\pm \Delta m_{32}^2$: Symmetries in neutrino physics, consequences for 2β decay search.

NOvA Experiment

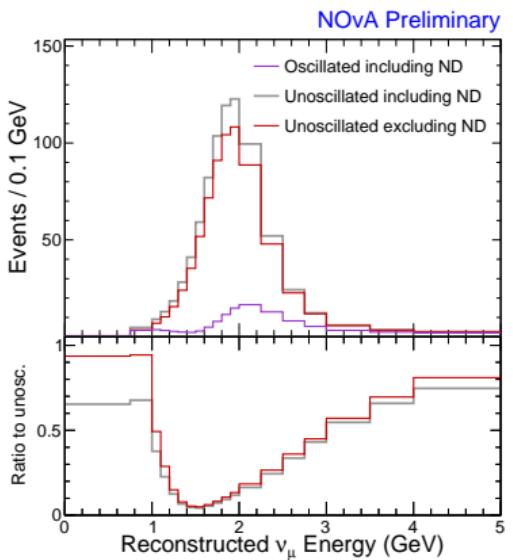


- Neutrino beam from Fermilab's NuMI Beamlne.
- 14 mrad off-axis beam narrowly peaked at ~ 2 GeV.
- Two functionally identical detectors:
 - Near Detector (ND), 0.3 kton, 1 km baseline.
 - Far detector (FD), 14 kton, 810 km baseline.



Neutrino oscillations with accelerators

ν_μ signal



Location of the dip: $|\Delta m_{32}^2|$

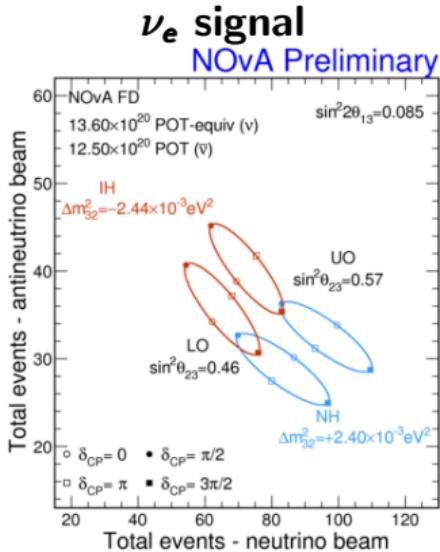
(does not depend on the sign)

Depth of the dip: $\sin^2(2\theta_{23})$

Difficult to separate $\theta_{23} > 45^\circ$ and $\theta_{23} < 45^\circ$

Is $\nu_\mu = \nu_\tau$ in ν_3 mass state?

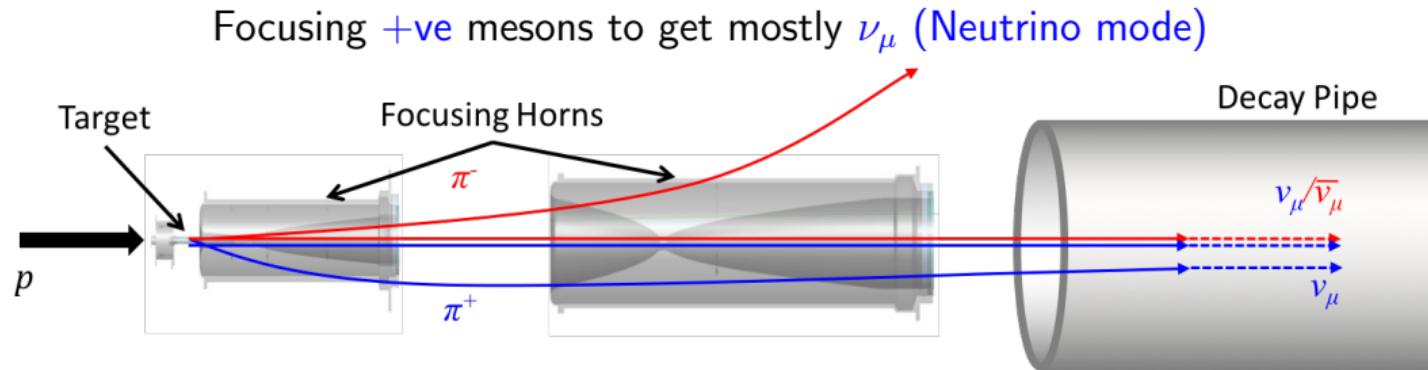
ν_e signal



Combination of ν_e and $\bar{\nu}_e$ excess;
 $\sin^2(\theta_{23})$, $\sin^2(\theta_{13})$, δ_{CP}

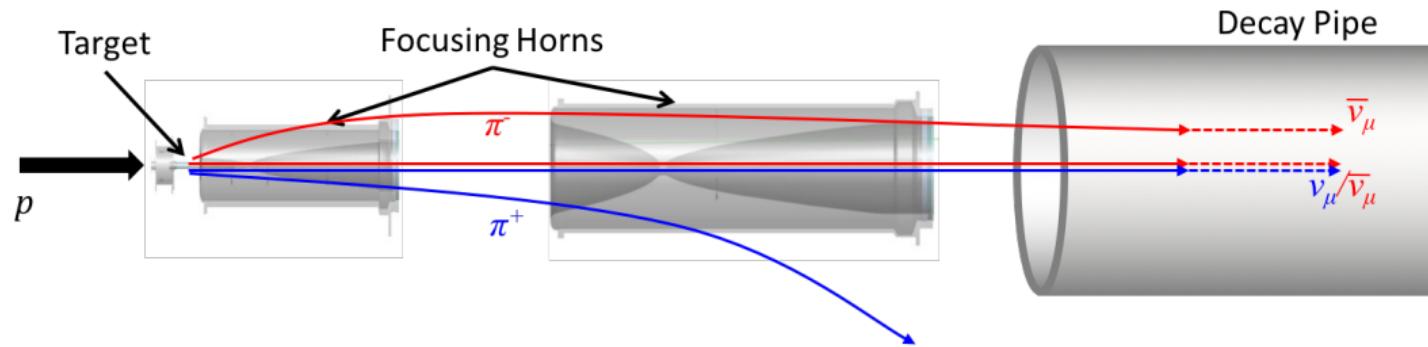
Good dependence on the sign of Δm_{32}^2

Channel for CP violation detection



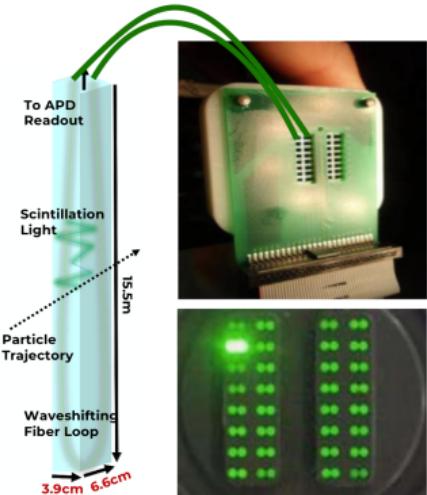
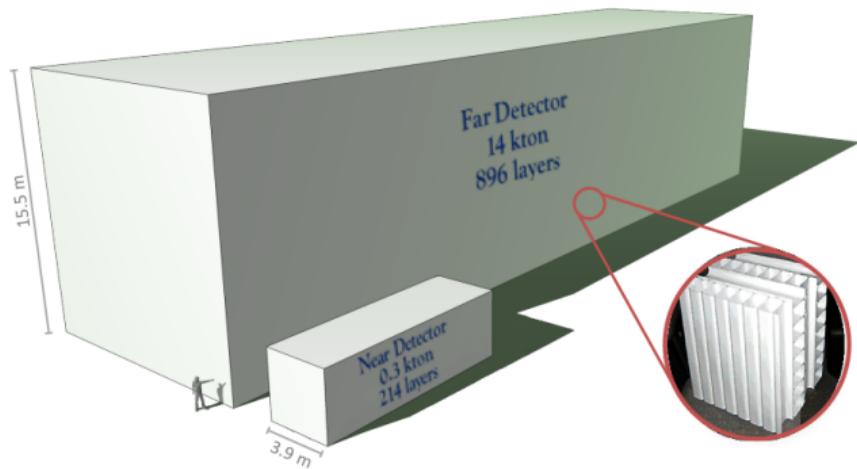
- Beam of 120 GeV protons incident on carbon target.
- Focusing +ve or -ve mesons to obtain mostly ν_μ or $\bar{\nu}_\mu$.
 - Achieved by reversing the polarity of the magnetic horns.
- Neutrinos appear from the decaying mesons. 675 m decay pipe.

Focusing -ve mesons to get mostly $\bar{\nu}_\mu$ (Antineutrino mode)



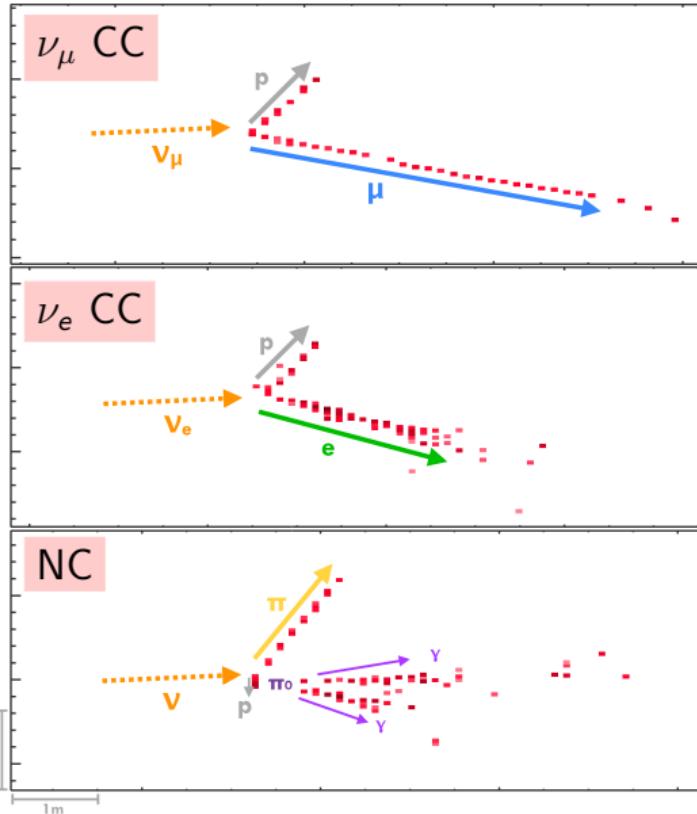
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NOvA Detectors



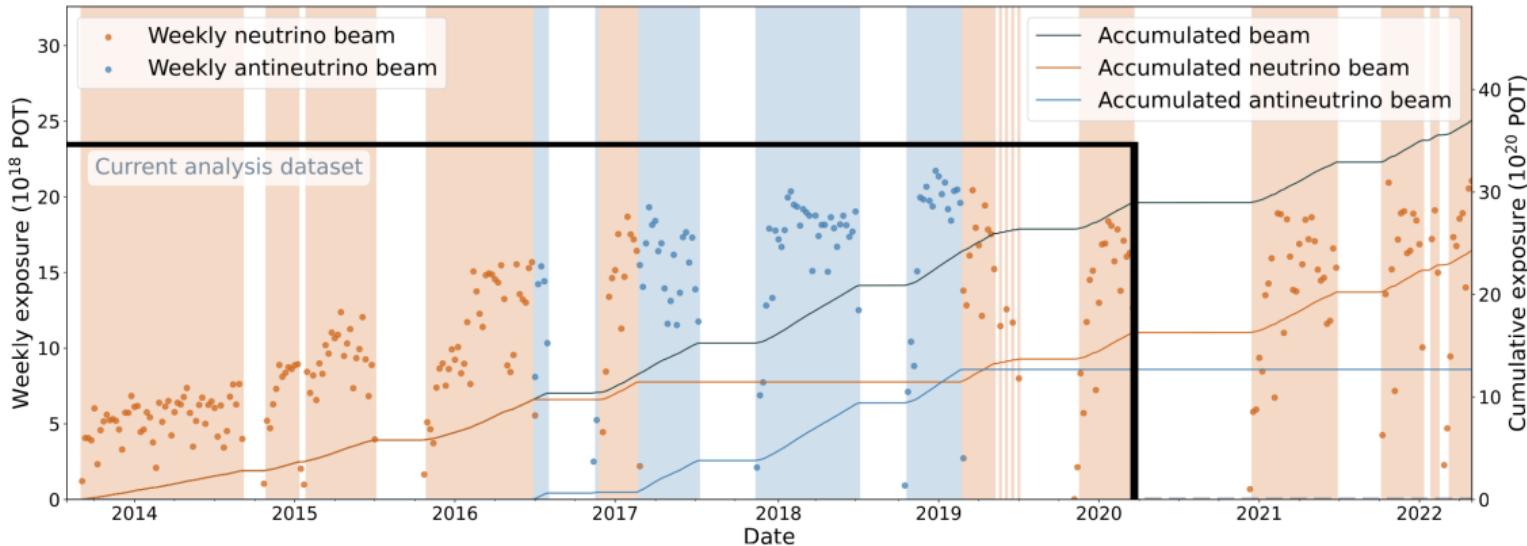
- Extruded cells filled with liquid scintillator, with 62% active volume.
- Wavelength-shifting fibre collects and transports light to Avalanche photodiode.
 - Each APD sees 32 NOvA cells.
- Cells with alternating horizontal & vertical planes for 3D reconstruction.
- Optimized for electron showers.

Event topologies



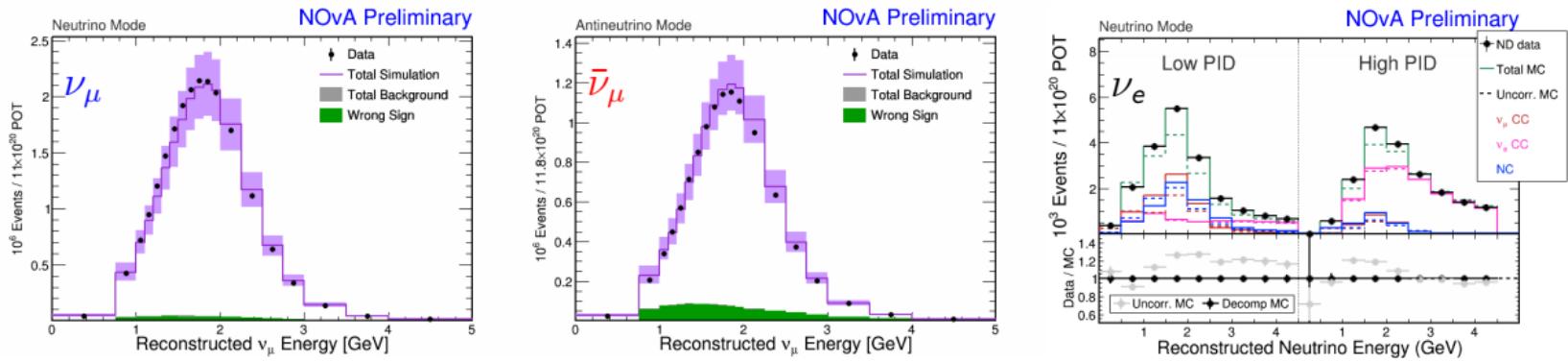
- Modern **convolutional neural network (CNN)** techniques used to identify neutrino flavour.
- Learns features of different event topologies.
- Data-driven validations based on ND and FD control samples.
- New test beam data will help with future validations/improvements.

Collected beam data



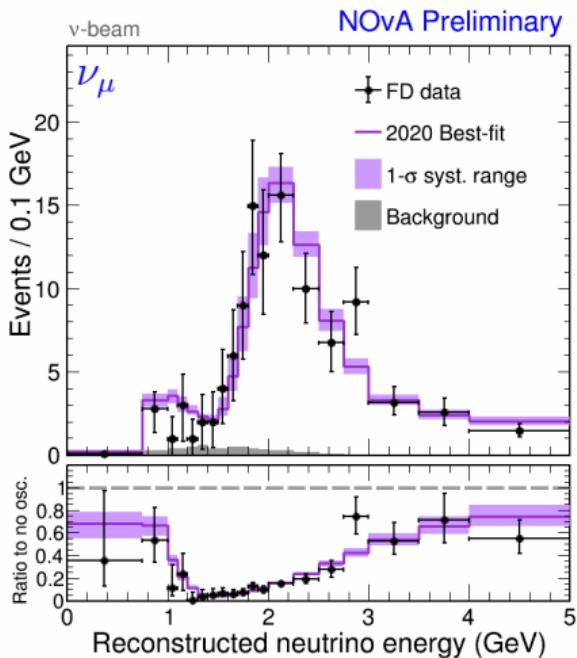
- Collected 37×10^{20} protons-on-target up to date.
- Data up to early 2020 included in the analysis shown here.
 - 13.6×10^{20} in ν -beam mode.
 - 12.5×10^{20} in $\bar{\nu}$ -beam mode.

Near detector data

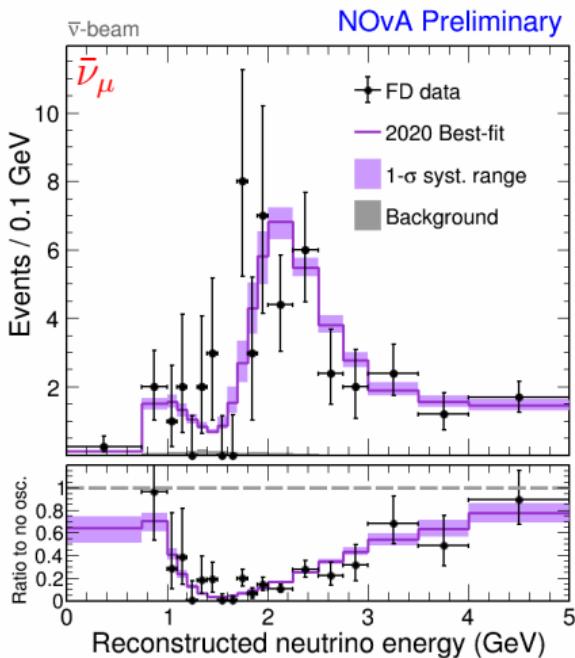


- ν_μ and $\bar{\nu}_\mu$ ND samples are used to correct the FD unoscillated predictions via extrapolation.
- We can then apply the $P(\nu_\mu \rightarrow \nu_e)$ curve to the corrected predictions.
- The ν_e samples are used to correct the irreducible ν_e background in the beam at the FD.
- The ν_e samples also used to estimate how the backgrounds to the ν_e appearance oscillate.

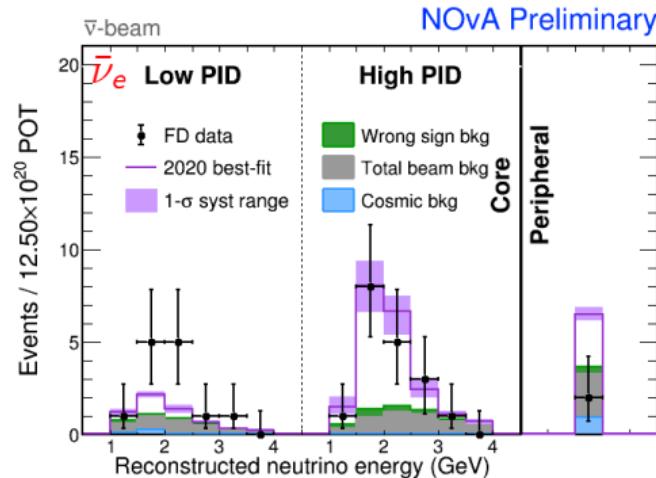
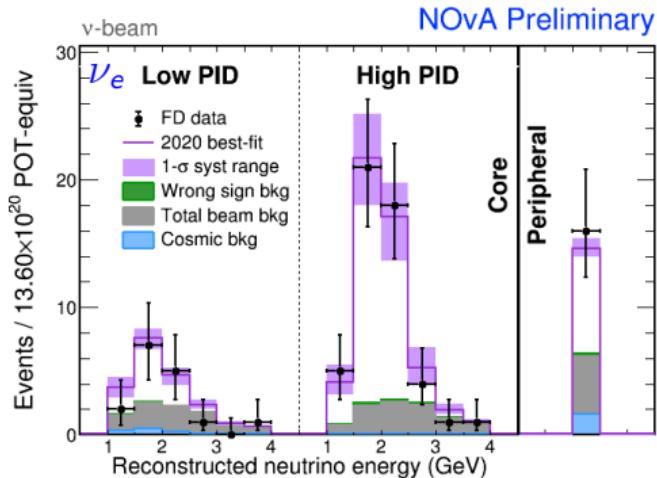
Far detector data Muon neutrinos



- Observed: 211
- Best Fit Prediction: 222.3
- Background: 8.2



- Observed: 105
- Best Fit Prediction: 105.4
- Background: 2.1



- Observed: 82
- Best Fit Prediction: 85.8
- Background: 26.8

- Observed: 33
- Best Fit Prediction: 33.2
- Background: 14

$> 4\sigma$ evidence of electron antineutrino appearance

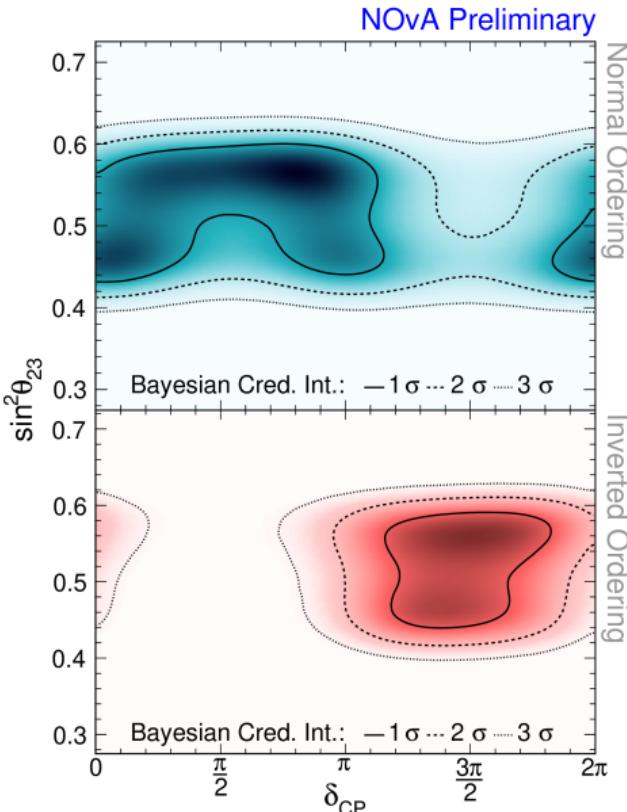
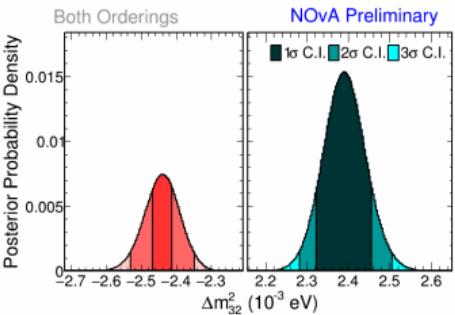
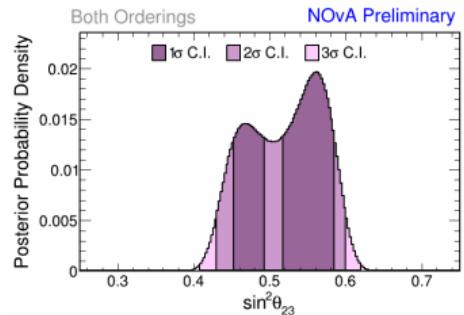
Results

Standard PMNS results

Both Orderings

Normal Ordering

Inverted Ordering



- General conclusions compatible with the 2020 Frequentist analysis.
- Prefer upper octant and normal mass ordering.
- Neither preference is significant, below $\sim 1\sigma$.
- “Not worth more than a bare mention” *

	N. Ordering	I. Ordering	
U. Octant	41.7%	20.9%	62.6%
L. Octant	25.8%	11.5%	37.4%
	67.5%	32.5%	

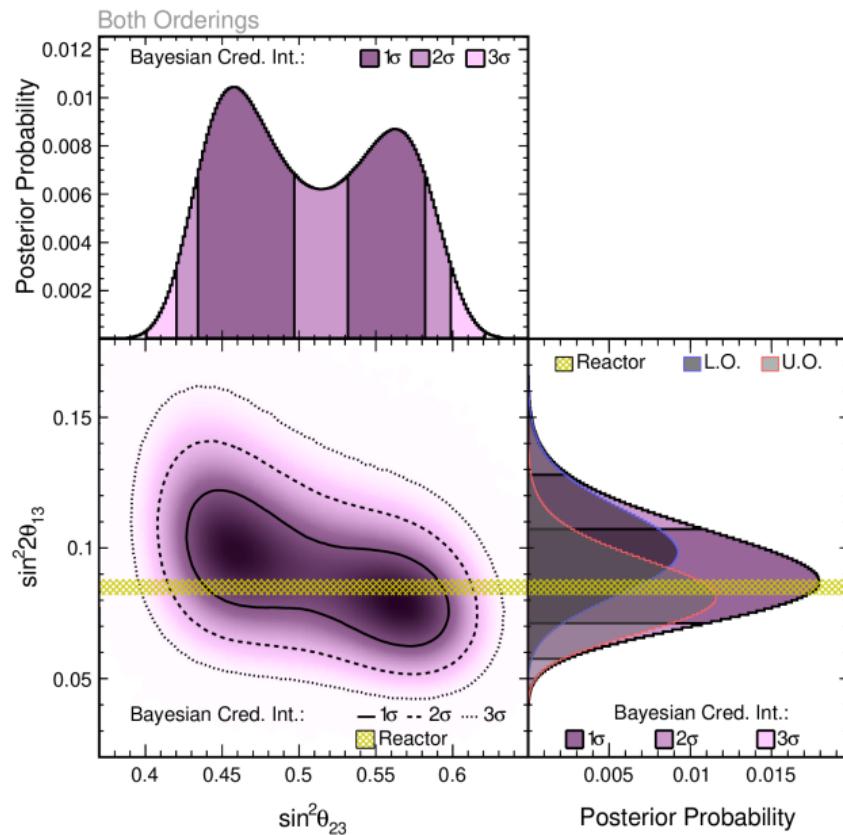
* Jeffreys ISBN:9780191589676, Raftery & Kass [doi:10.2307/2291091](https://doi.org/10.2307/2291091)

First NOvA-only θ_{13} measurement

Both Orderings

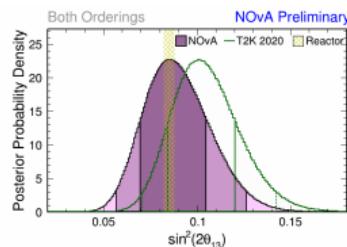
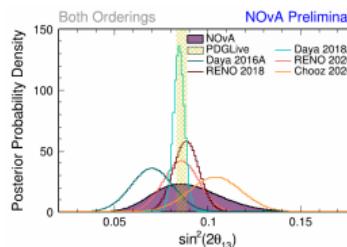
Normal Ordering

Inverted Ordering



- $\sin^2(2\theta_{13}) = 0.085^{+0.020}_{-0.016}$

- NOvA in a good agreement with the reactor experiments.
- No tensions between short-distance $P(\bar{\nu}_e \rightarrow \bar{\nu}_e)$ and long-distance $P(\nu_\mu \rightarrow \nu_e)$ & $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$.

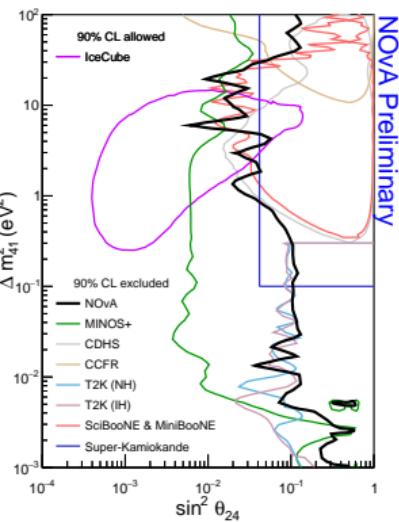


Sterile Neutrinos: 3+1

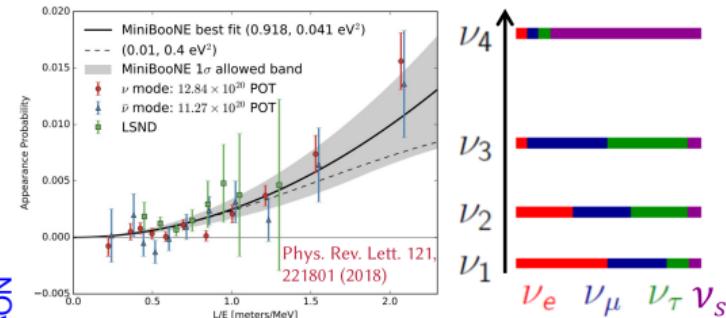
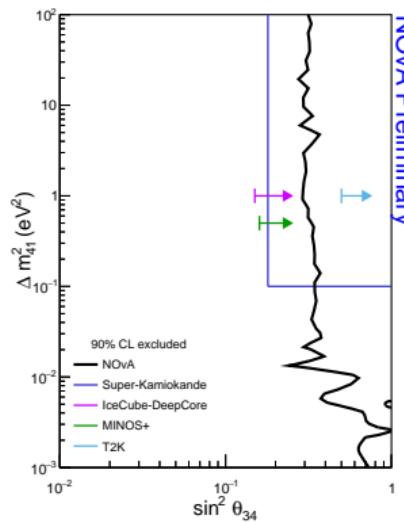
Anomalous neutrino event rates in the short-baseline experiments

- ν_μ disappearance could occur in the near detector at large Δm_{41}^2 .
- No evidence of sterile neutrinos found.
- Competitive results for Δm_{41}^2 , θ_{24} & θ_{34}

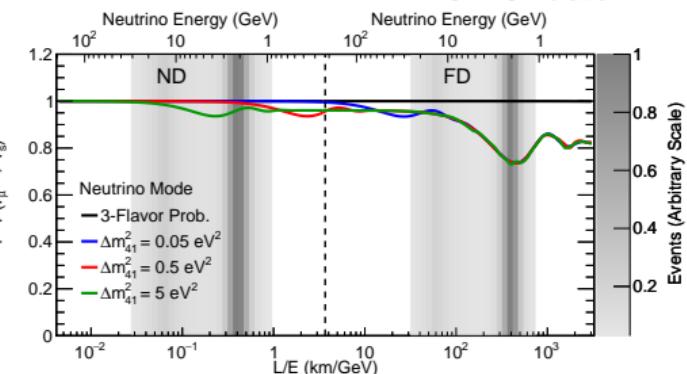
Neutrino Beam



Neutrino Beam



Neutrino Beam

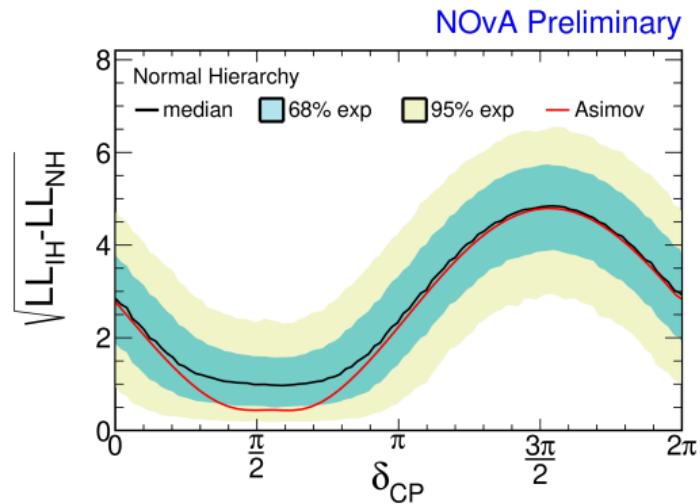


Bright Future

- NOvA Test-Beam to measure detector response.
- MW-capable horn and target installed.
 - New power record reached last year!
- Expect $> 2\times$ more in both ν and $\bar{\nu}$ data.
 - Analysed $26\text{e}20$ POT.
 - $11\text{e}20$ POT more collected since.
 - Goal by 2027: $67\text{--}72\text{e}9$ POT.



- NOvA-T2K effort to produce joint result.



Conclusions

- **Competitive constraints on 3+1 sterile neutrinos.**

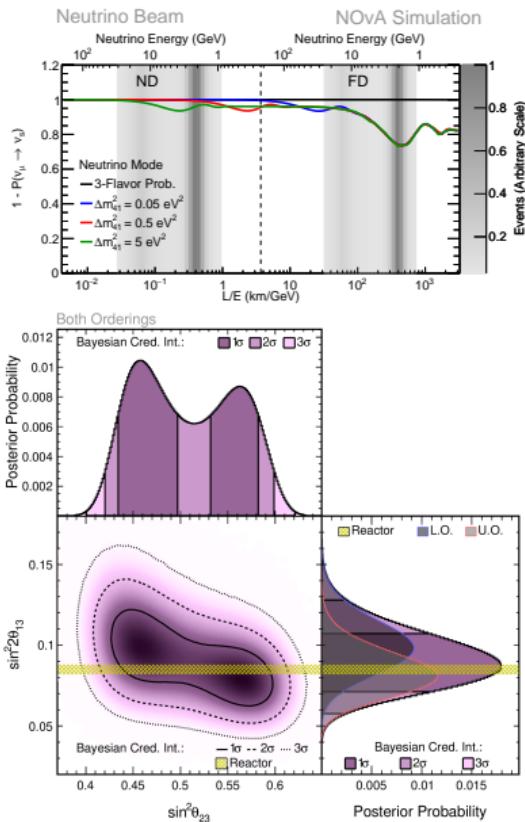
- NOvA data consistent with no sterile neutrinos.

- First NOvA-only measurement of θ_{13}

$$\sin^2(2\theta_{13}) = 0.085^{+0.020}_{-0.016}$$

- **PMNS formalism explains NOvA data very well:**

- No tension between Accelerator and Reactor neutrinos.
- No high preference for either CP-Violation or CP-Conservation.
- Good model-data agreement.



Thank you!



BACKUPS

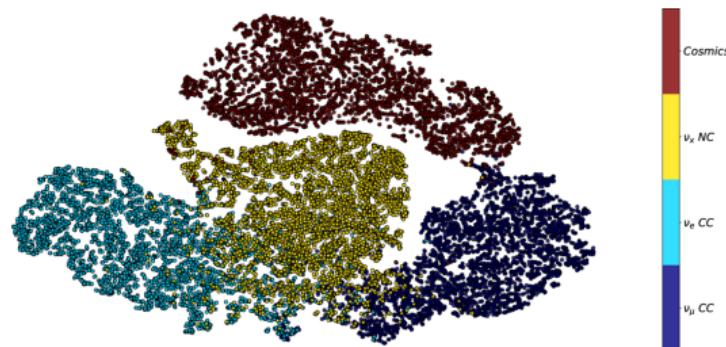
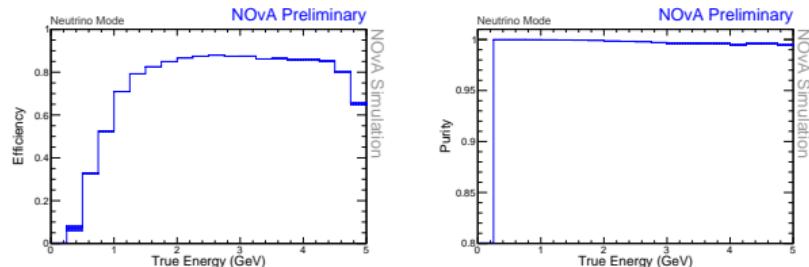
Event identification

Pre-selections:

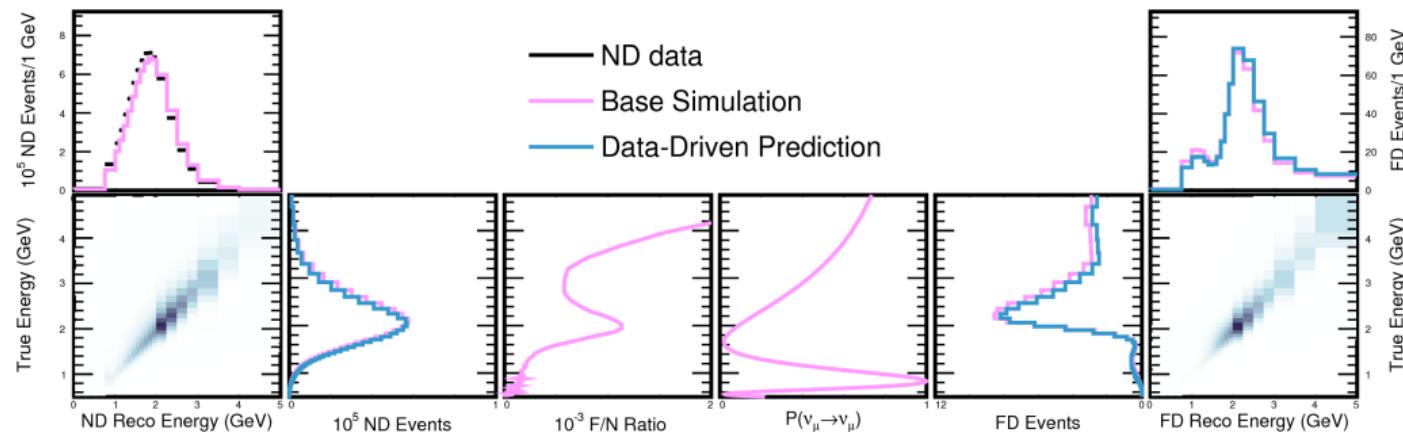
- Contained inside the detector.
- Inside of the beam spill-window.
- Cosmic particles rejection via BDT.

Event Identification:

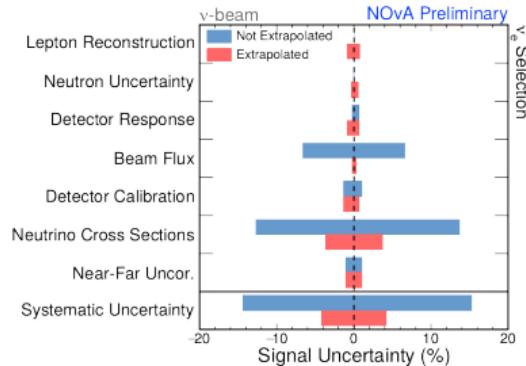
- Modern CNN techniques used to identify neutrino flavour.
- Learns features of different event topologies.
- Data-driven validations based on ND and FD control samples.
- Results in high purity samples.



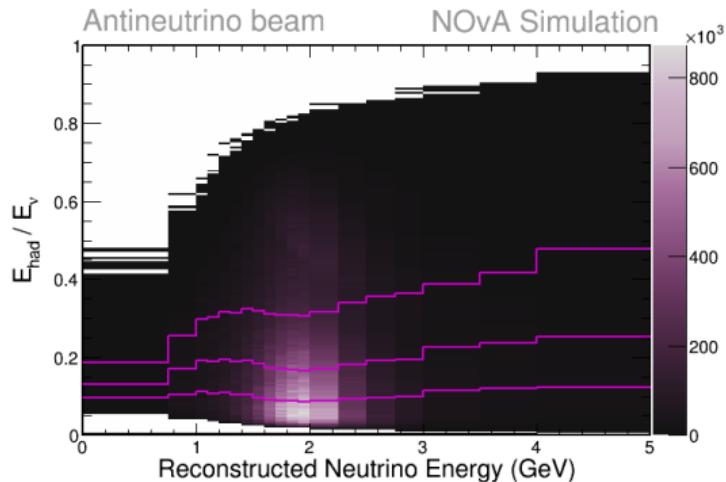
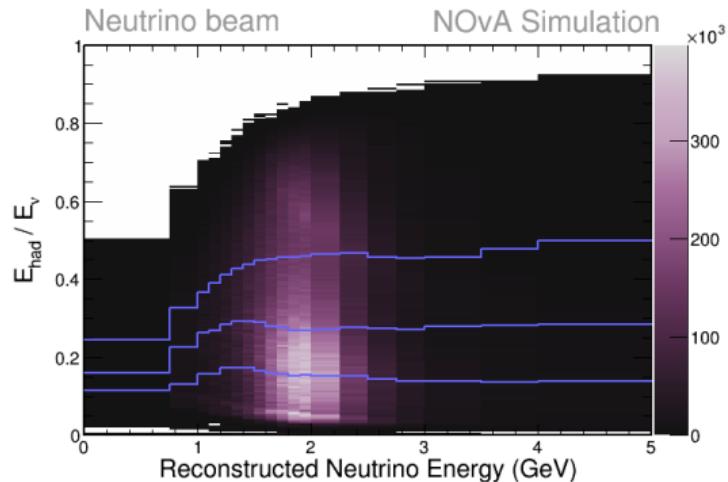
ND \rightarrow FD Extrapolation



- Take advantage of detector similarity to extrapolate ND predictions to FD.
- Many systematic effects e.g. cross-sections, flux and efficiency are shared.
- Helps dealing with the “unknown unknowns” .
- Extrapolate different kinematic samples separately to deal with Near/Far acceptance differences.

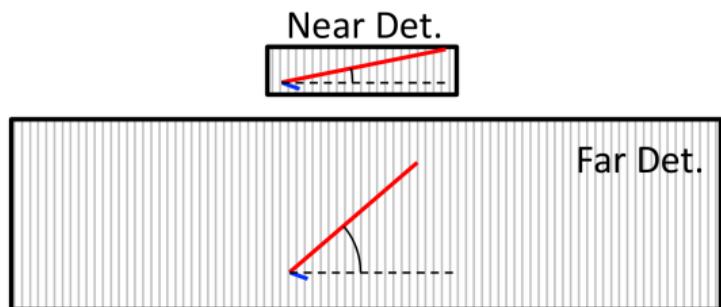
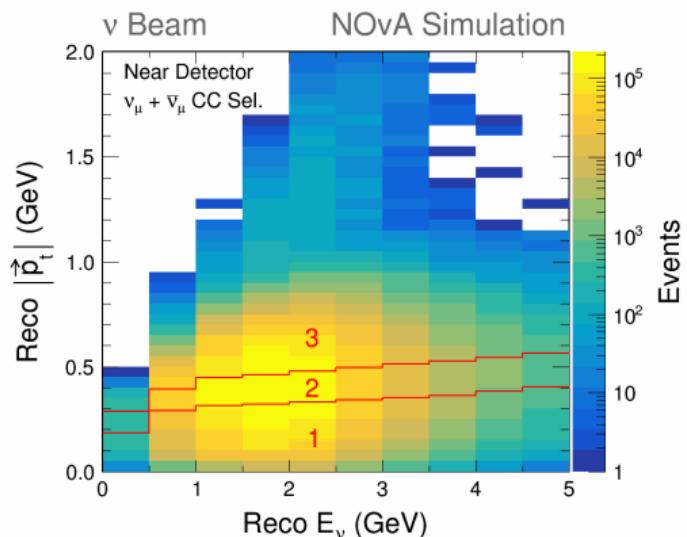


ND \rightarrow FD Extrapolation: E_{had}

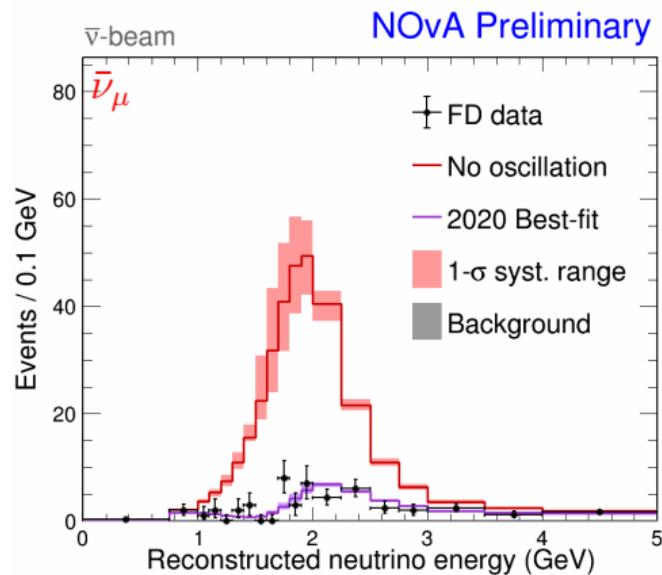
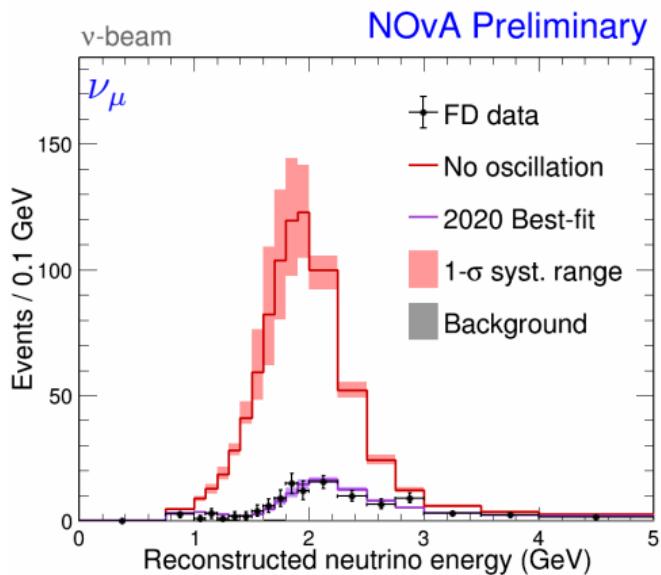


- The energy resolution varies between the detectors.
- Extrapolation split in four E_{had}/E_{ν} quartiles.
- Matches the Hadronic energy resolution between ND and FD.

ND \rightarrow FD Extrapolation: Lepton $|p_T|$



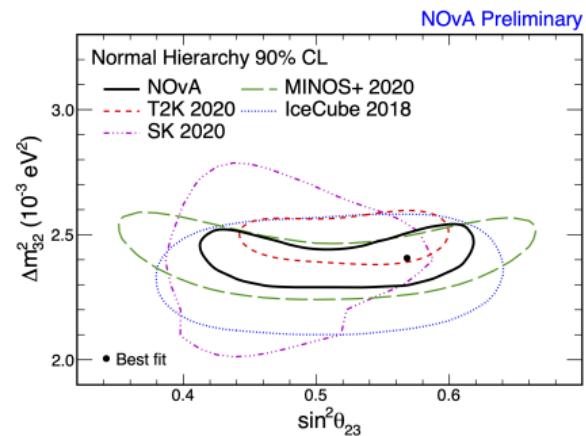
- Different lepton angle distributions due to the difference in detectors' size.
- Extrapolation split in three ranges of lepton transverse momenta.
- Done separately for each E_{had} quartile.
- Matches the detector acceptances between ND and FD.



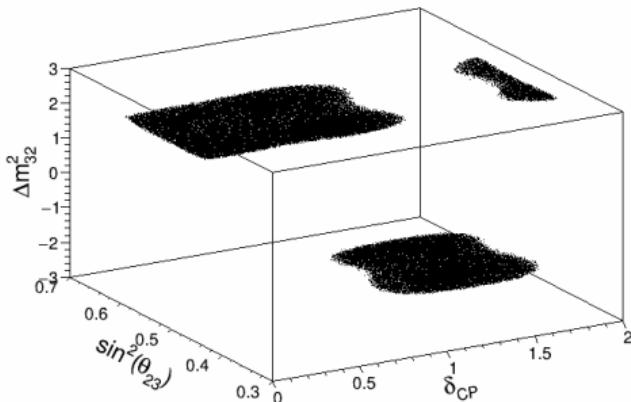
- Observed: 211
- Best Fit Prediction: 222.3
- Background: 8.2

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Fitting/Sampling techniques



- NOvA fits 10 data samples: $4\nu_\mu$, $4\bar{\nu}_\mu$, $1\nu_e$ and $1\bar{\nu}_e$.
- All the previous NOvA results were Frequentist with use of profiling.
- New Bayesian frameworks implemented in NOvA.
- New studies now easier: Jarlskog-Invariant, NOvA-only θ_{13} , Bayes factors and possibly more!
- Other experiments often provide Marginalized and/or Bayesian results.



Bayes Theorem:

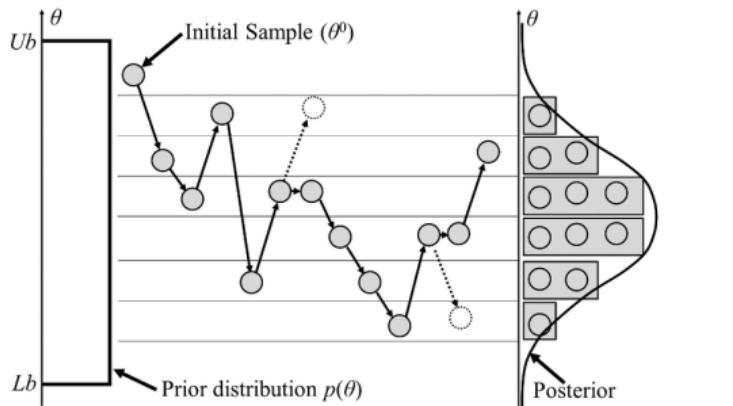
$$\mathcal{P}(\vec{\theta}|D) \approx \mathcal{P}(D|\vec{\theta})\mathcal{P}(\vec{a})$$

Posterior \approx Likelihood \times Prior

$$\text{Posterior} \approx e^{-\frac{\chi^2}{2}} \times \text{Prior}$$

- Bayesian results given in terms of posterior probability distributions.
- Need to produce N-dimensional probability distribution for marginalized results.
- MCMC generates samples on N-dimensional.
 - Sample density corresponds to posterior probability density.

Markov Chain Monte Carlo for NOvA



L. Jaewook et. al. (2015). Energies. 8. 5538-5554. 10.3390/en8065538.

Bayes Theorem:

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Posterior \approx Likelihood \times Prior

$$\text{Posterior} \approx e^{-\frac{\chi^2}{2}} \times \text{Prior}$$

- MCMC generates samples by iteratively deviating parameters from their previous values.
- At each iteration we can either accept, or reject the step.
 - Accept: new step added to the end of the chain.
 - Reject: previous values repeated at the end of the chain.
- Over time, this “chain” ensemble starts resembling posterior probability.



Arianna Rosenbluth



Stanislaw Ulam

- Two algorithms in NOvA: Metropolis-Hastings and Hamiltonian MCMC.
- Hamiltonian MCMC is based on Stan library (<https://mc-stan.org>).  Stan

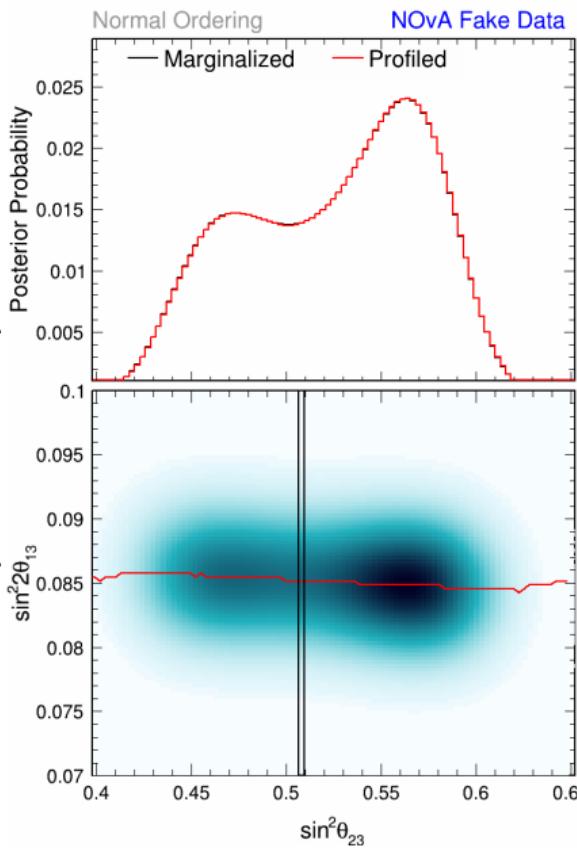
 - Stanislaw Ulam invented the methods of Monte-Carlo.

- Metropolis-Hastings was written from scratch in-house.
 - Named Aria after Arianna Rosenbluth, who first implemented the method.
- Importantly, both algorithms produce identical results.

References: Metropolis-Hastings [doi:10.1063/1.1699114](https://doi.org/10.1063/1.1699114), Hamiltonian [doi:10.1016/0370-2693\(87\)91197-X](https://doi.org/10.1016/0370-2693(87)91197-X)

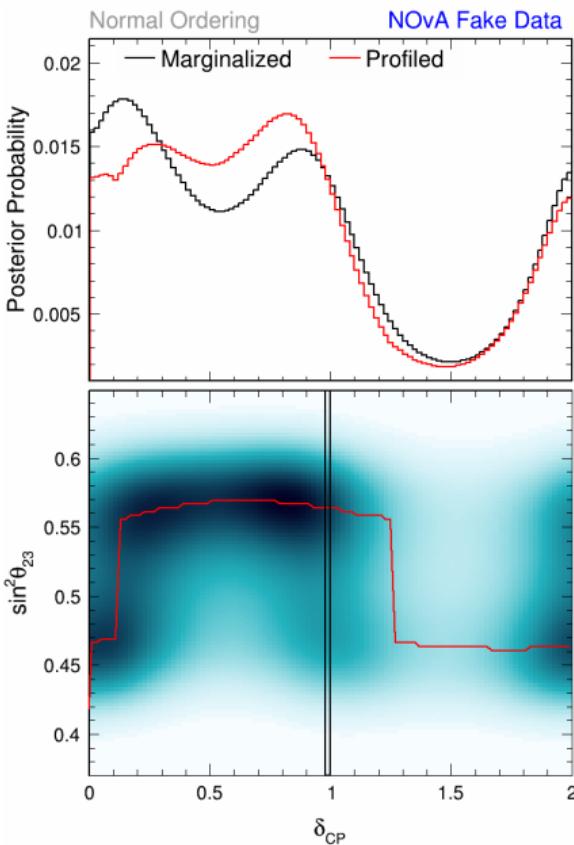
Bayesian vs Frequentist, Marginalization vs Profiling

- MCMC uses marginalization rather than profiling.
 - Not necessarily reserved to Bayesian methods!
 - Profiling: Maximize parameters not shown.
 - Marginalization: Integrate over parameters not shown.
- Example: marginalizing/profiling over $\sin^2 2\theta_{13}$.
 - Line of best fit to profile over $\sin^2 2\theta_{13}$.
 - Box with probabilities to sum over for marginalization.
- Use posterior probability densities, not χ^2 .
 - Bayes. Credible Intervals vs Freq. Confidence Levels.

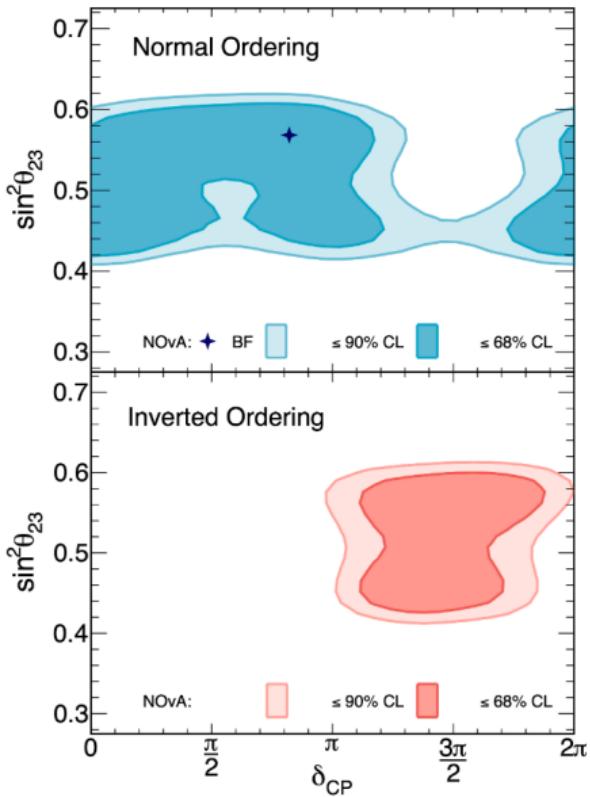


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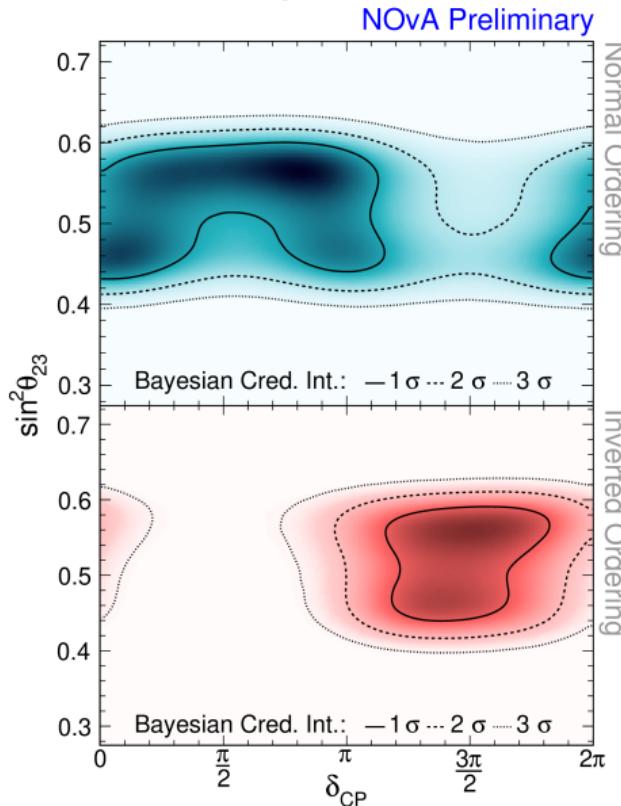


Frequentist result

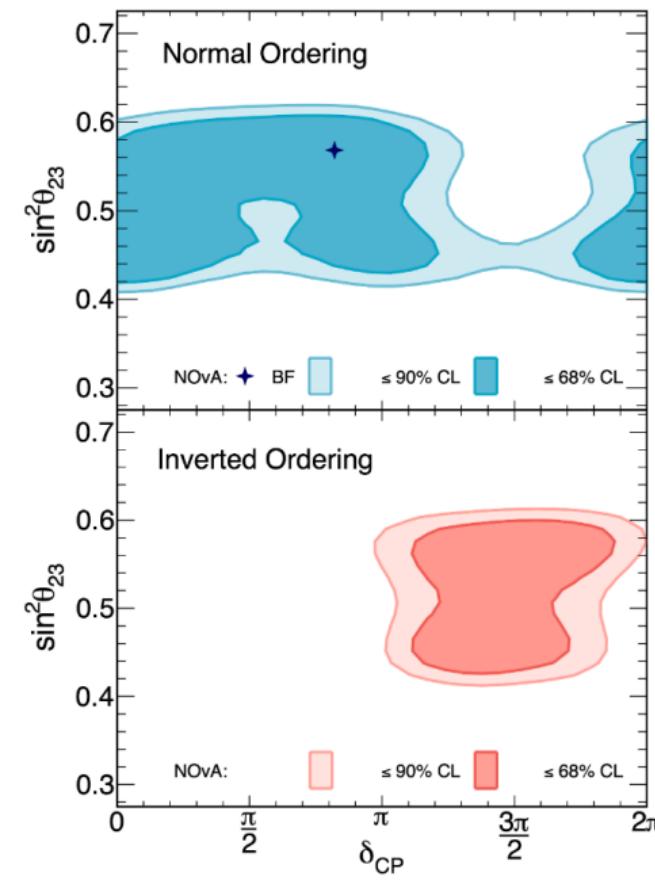


Frequentist results reference: [arXiv:2108.08219](https://arxiv.org/abs/2108.08219)

Bayesian result

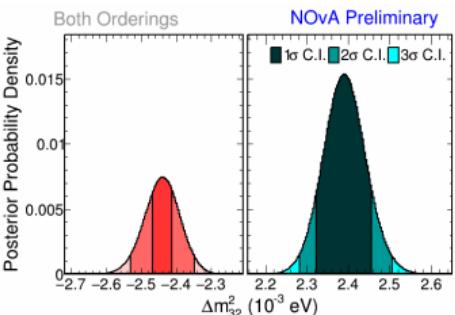
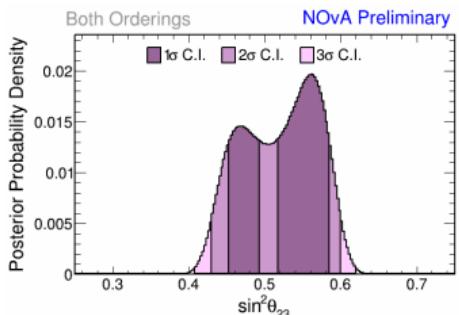


2020 Frequentist Results



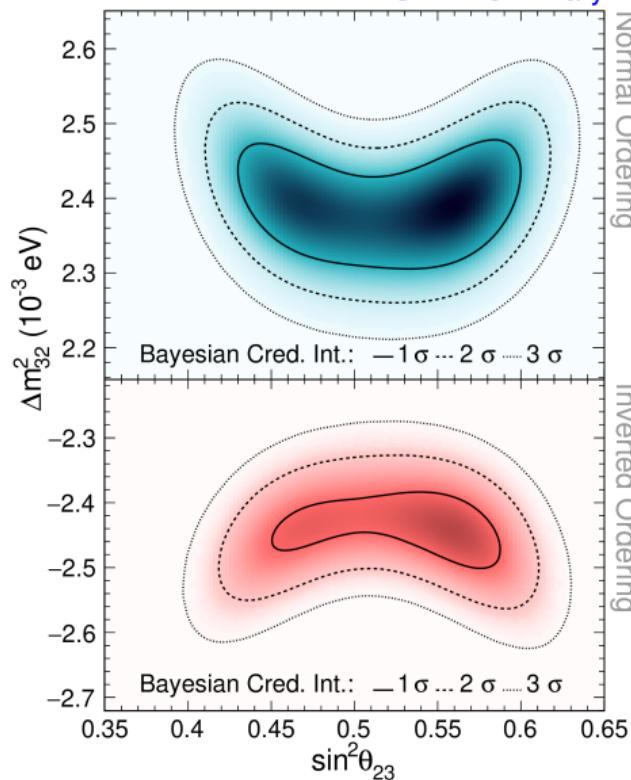
Best fit:

- Normal mass ordering.
- $\Delta m_{32}^2 = (2.41 \pm 0.07) \times 10^{-3} \text{ eV}^2$
- $\sin^2 \theta_{32} = 0.57^{+0.04}_{-0.03}$
- $\delta_{CP} = 0.82\pi$
 - Disfavour IO $\delta_{CP} = \pi/2$ at $> 3\sigma$.
 - Disfavour NO $\delta_{CP} = 3\pi/2$ at 2σ .



- Prefer upper octant and normal mass ordering.
- Neither preference is significant, below $\sim 1\sigma$.

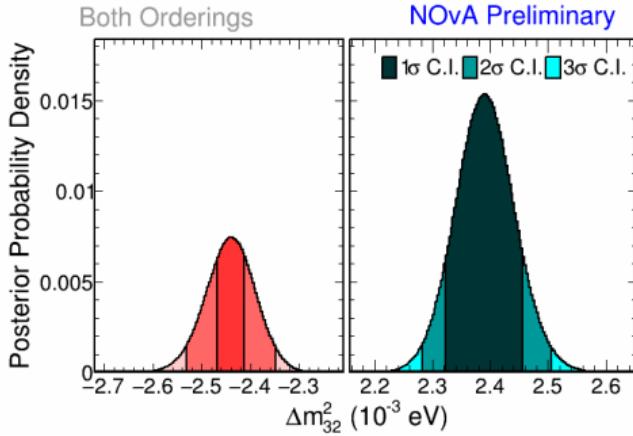
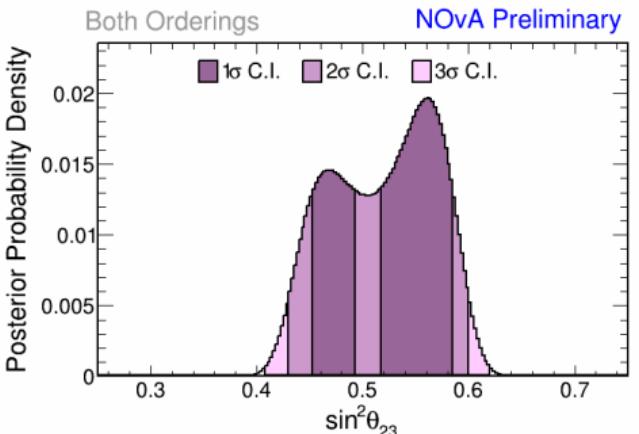
	N. Ordering	I. Ordering	
U. Octant	41.7%	20.9%	62.6%
L. Octant	25.8%	11.5%	37.4%
	67.5%	32.5%	

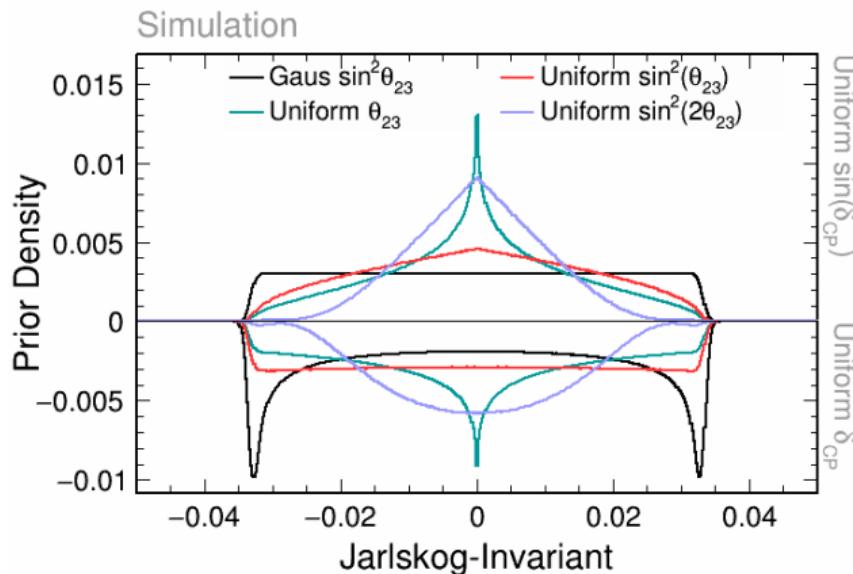


	N. Ordering	I. Ordering	
U. Octant	41.7%	20.9%	62.6%
L. Octant	25.8%	11.5%	37.4%
	67.5%	32.5%	

- Bayes Factors: odds ratio, how much more likely one model is than another.
- NO/IO: 2.1, UO/LO: 1.7
- Both can be interpreted as below 1σ or “not worth more than a bare mention” according to Jeffreys and Raftery & Kass scales.

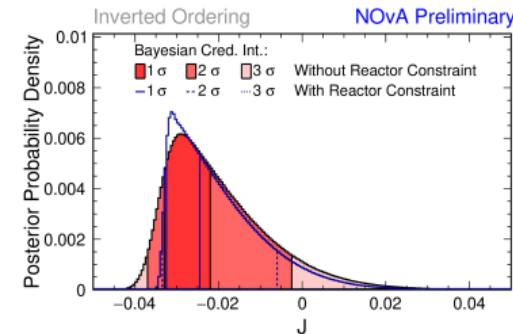
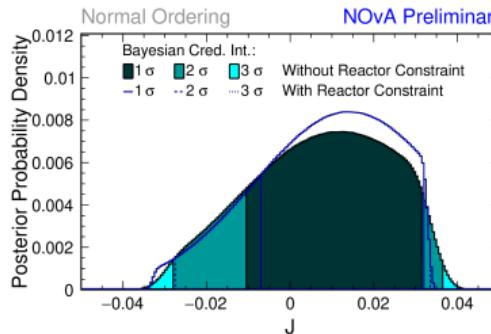
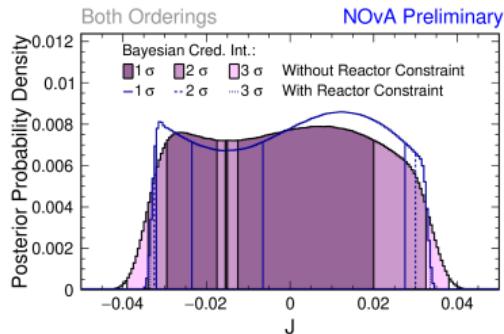
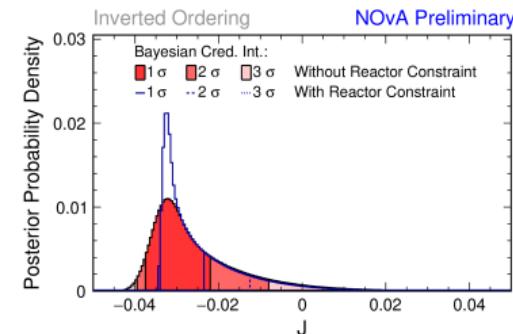
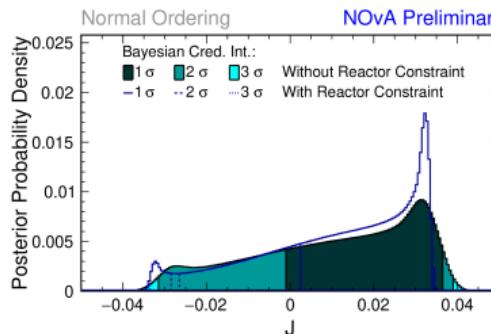
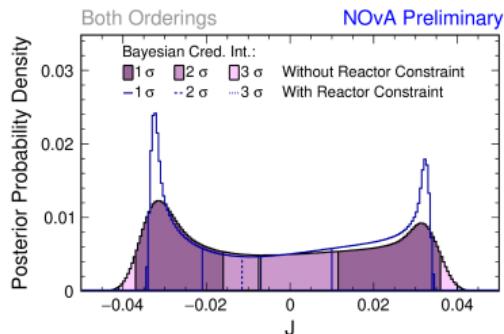
References: Jeffreys ISBN:9780191589676,
Raftery & Kass doi:[10.2307/2291091](https://doi.org/10.2307/2291091)

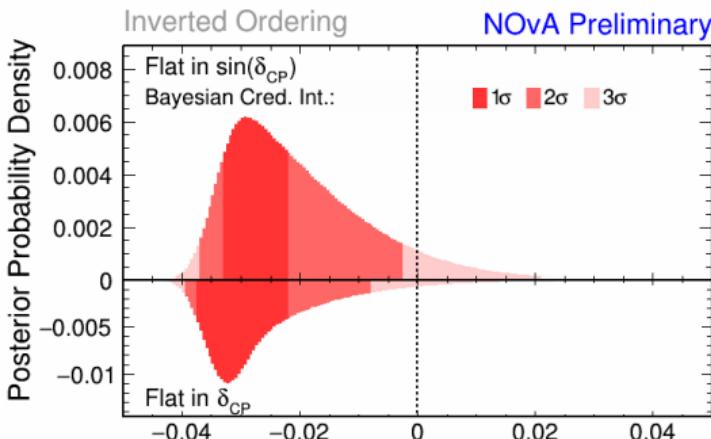
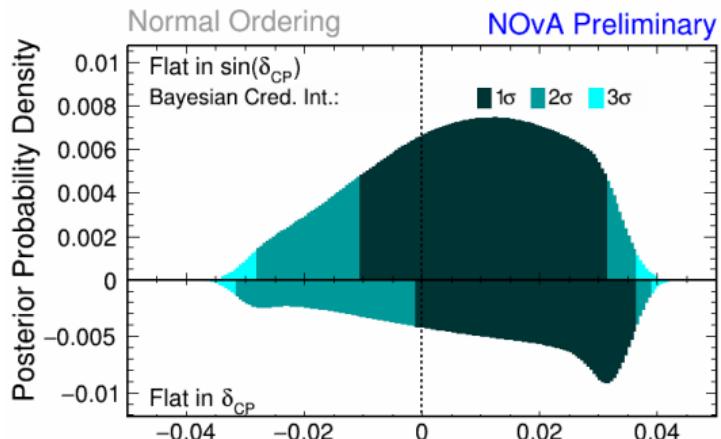




How about priors on other oscillation parameters for J?

- Changing θ priors to be uniform in θ , $\sin^2 \theta$ changes the prior contribution.
- It does not, however, change the posterior – our results.
- Likelihood is stronger than the prior for high-stats data, overwhelming it.
- This does not happen for δ_{CP} because we don't constrain it well.

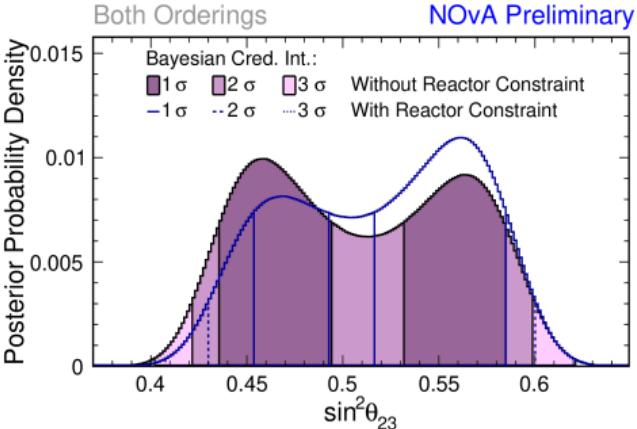
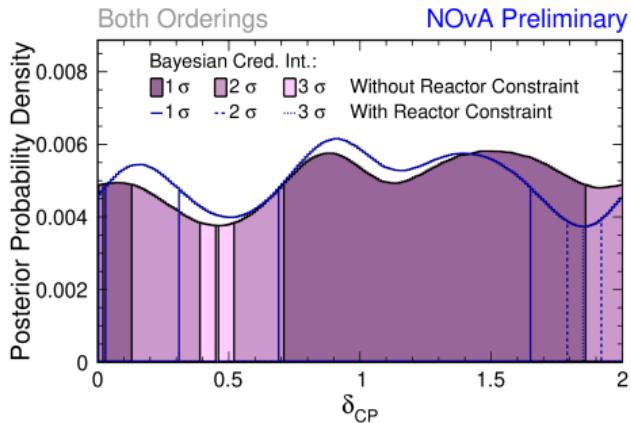
θ_{13} constraint comparisonsPrior uniform in $\sin(\delta_{CP})$:Prior uniform in δ_{CP} :



- The shape changes because θ_{13} is allowed to take more values.
- Nevertheless, the general conclusions about CP-conservation are similar.

Reactor constraint comparisons

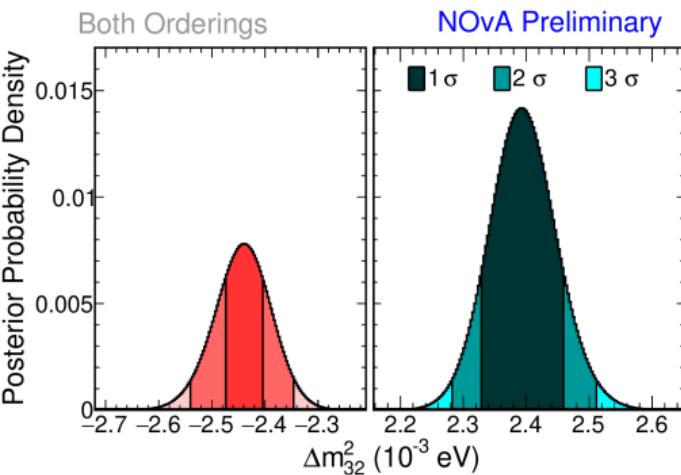
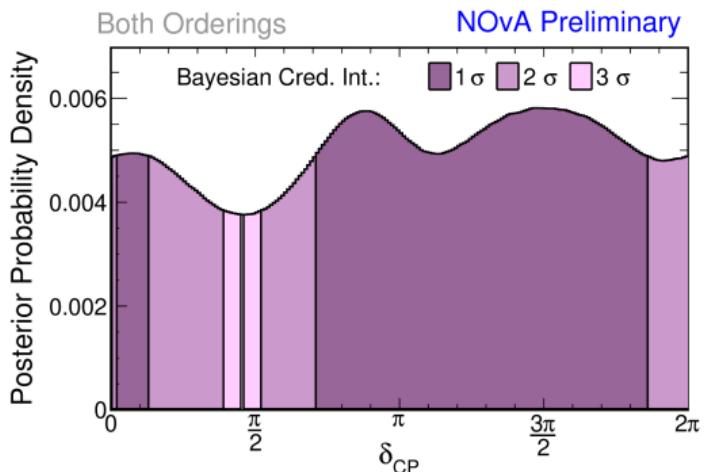
Both Orderings Normal Ordering Inverted Ordering



- Setting θ_{13} free does change our results slightly.
- Prefer lower octant with free θ_{13} , upper octant when constrained.
- These differences are low, however.
 - 1 σ intervals in both octants.
 - Low Bayes Factors.

Results without Reactor Constraint

Both Orderings Normal Ordering Inverted Ordering



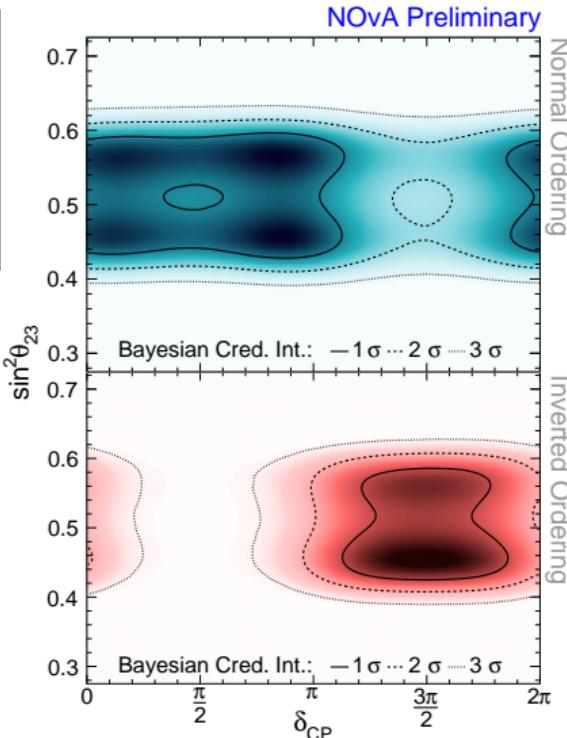
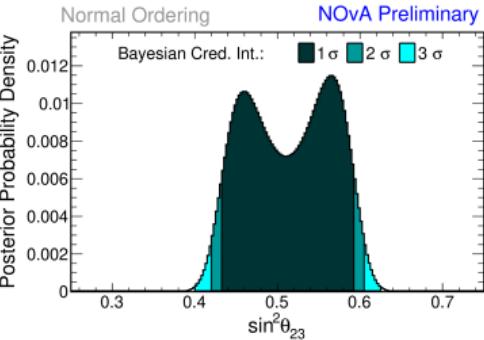
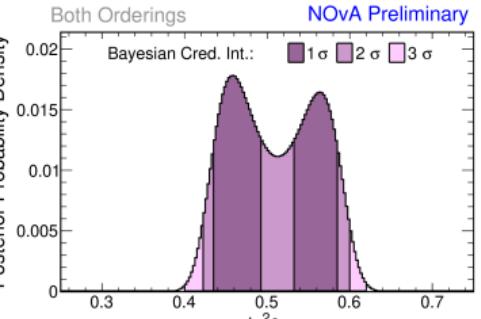
- All previous Frequentist results shown with external θ_{13} constraint.
- But NOvA has sensitivity to θ_{13} ! How does it affect our results?
 - Do we agree with the Reactors? Tensions in the PMNS model?
- Allowing unconstrained θ_{13} to give NOvA-only preferences:
 - δ_{CP} preferences don't change much.
 - Prefer normal mass ordering.
 - General conclusions similar to Reactor-constrained θ_{13} .

Results without Reactor Constraint

Both Orderings

Normal Ordering

Inverted Ordering

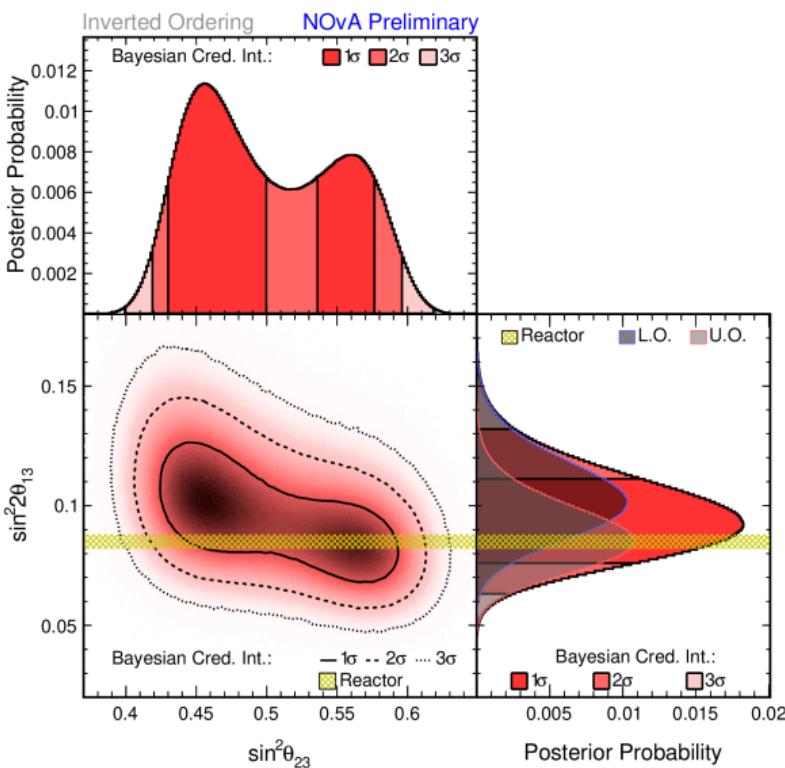
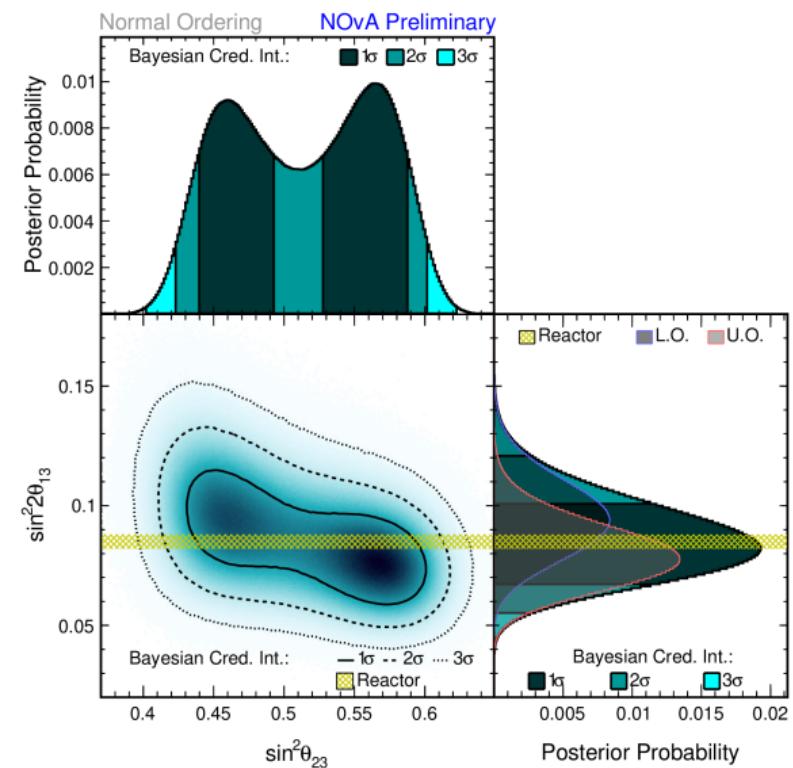


Lower octant: $\sin^2 \theta_{23} < 0.5$, Upper Octant: $\sin^2 \theta_{23} > 0.5$

- Prefer Lower Octant overall with NOvA-only θ_{13}
- Slight preference for Upper Octant in Normal Ordering.
- Higher preference for Lower Octant in Inverted Ordering.
- We need to look at θ_{13} to understand this.

NOvA-only θ_{13} measurements

Both Orderings Normal Ordering Inverted Ordering



Non-Standard Interactions: $\epsilon_{\mu\tau}$

