

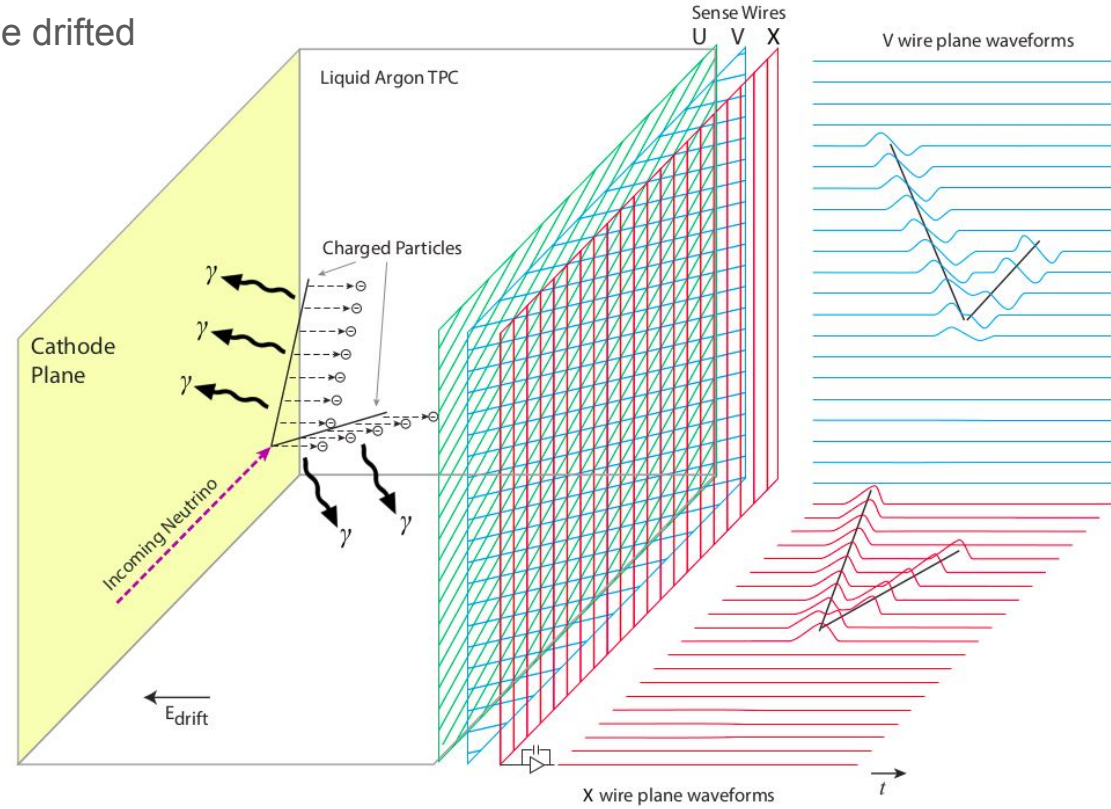
Light simulation and Analysis

Henrique Souza

APC - 19/04/2023

Motivation

- The LArTPC is characterized by the free drifted electrons signal and light emission
- The detection of scintillation light can provide the absolute time (T_0) of events and internal triggering for non beam events
- **Besides**, light signals can improve position, time and energy resolution. Improve particle identification (PID) and improve background rejection by the proper fiducialization of the detector.



- Liquid argon scintillation
 - Mechanism
 - Composition and time response
 - Propagation
- Detection
 - Description of DUNE's photon detectors: X-Arapuca
- Simulation
 - How is the light simulation implemented
- Analysis
 - Where to find data real data and its structure

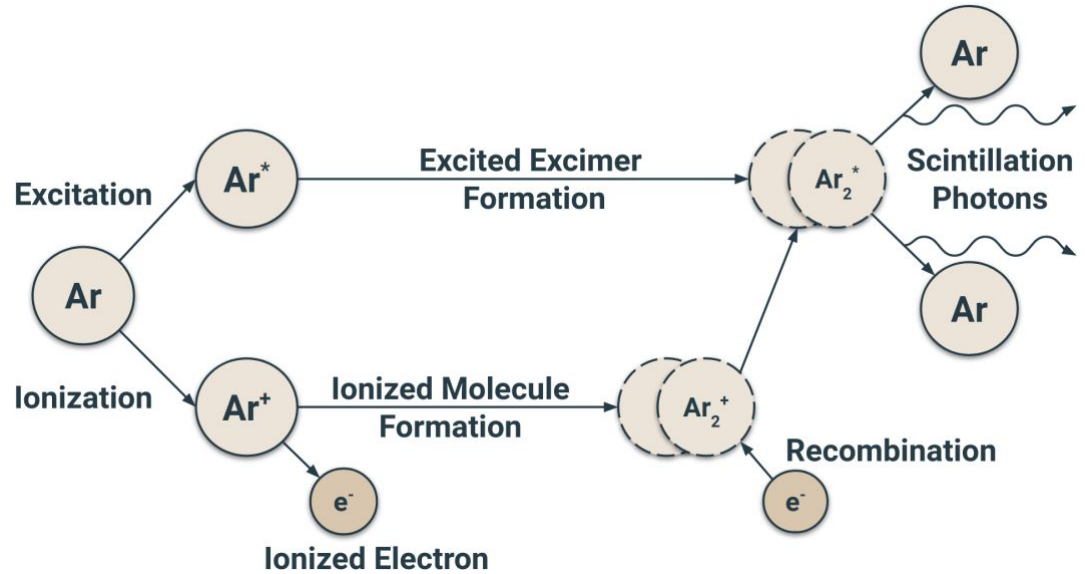
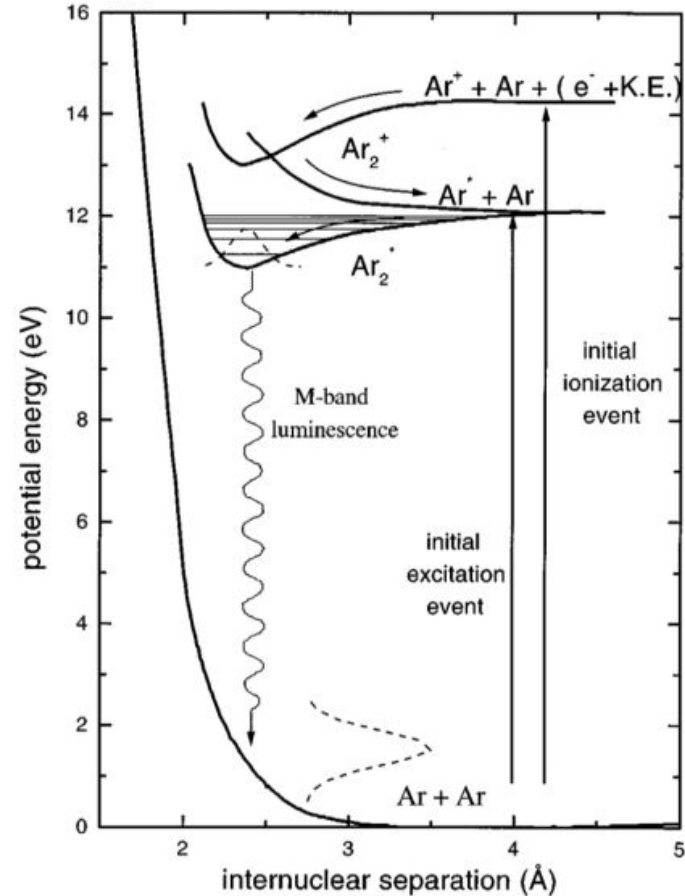
Disclaimer: many of these slides were possible from past slides, specially [this](#) presentation from Andrzej Szec and thanks to Laura Paulucci for sending support material.

Liquid argon scintillation

- Mechanism of light production
- Time components
- Scintillation yield
- Electric field
- Light propagation

Mechanism

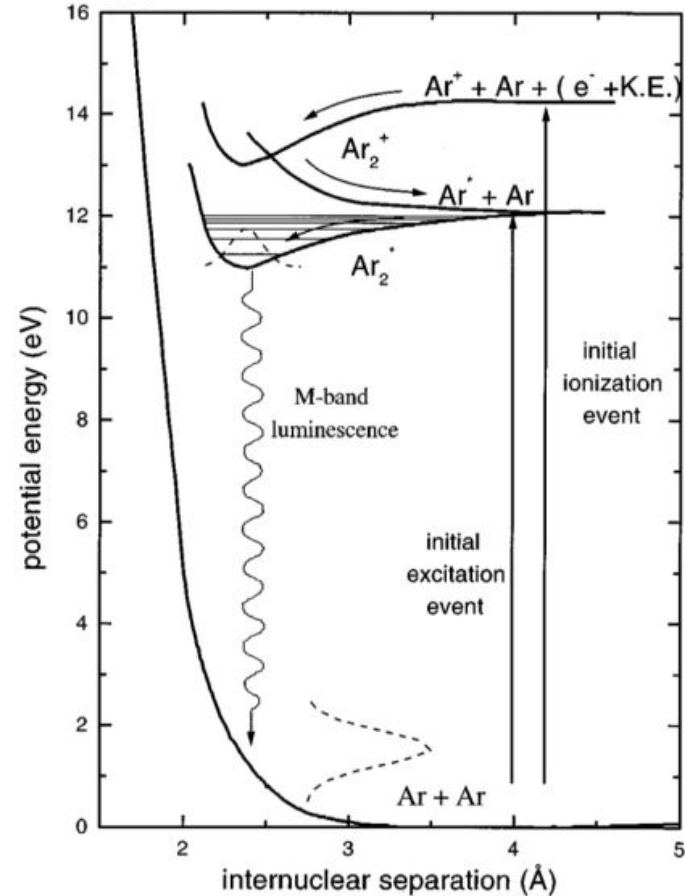
Ph. Rev. B 56 (1997), 6975



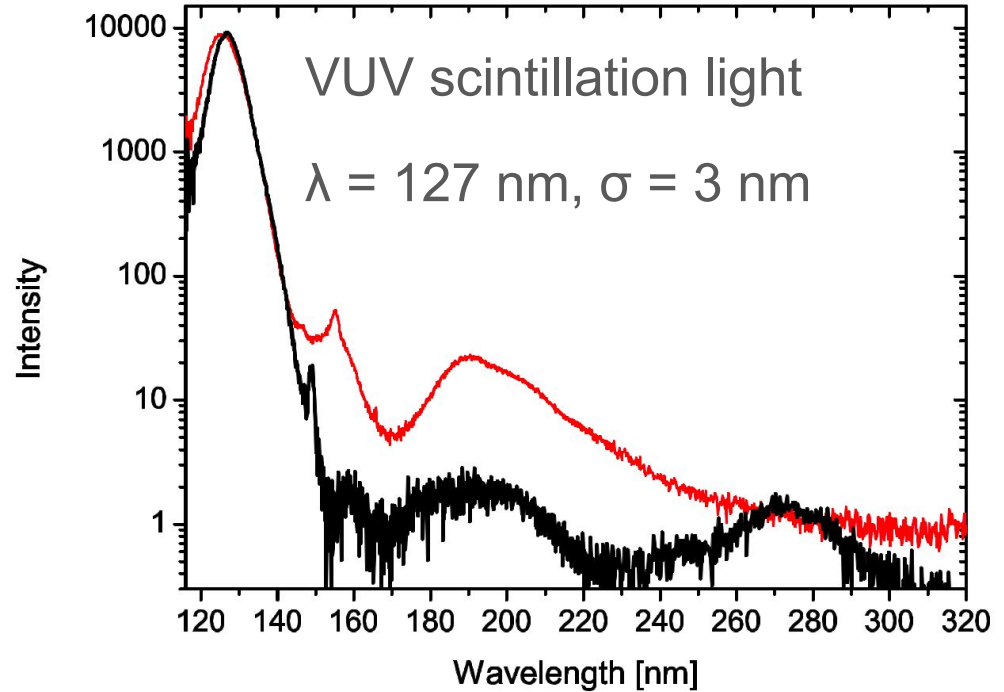
Ar₂^{*} should be de-excited to the ground state by emitting a Vacuum ultraviolet (VUV) photon

Mechanism

Ph. Rev. B 56 (1997), 6975

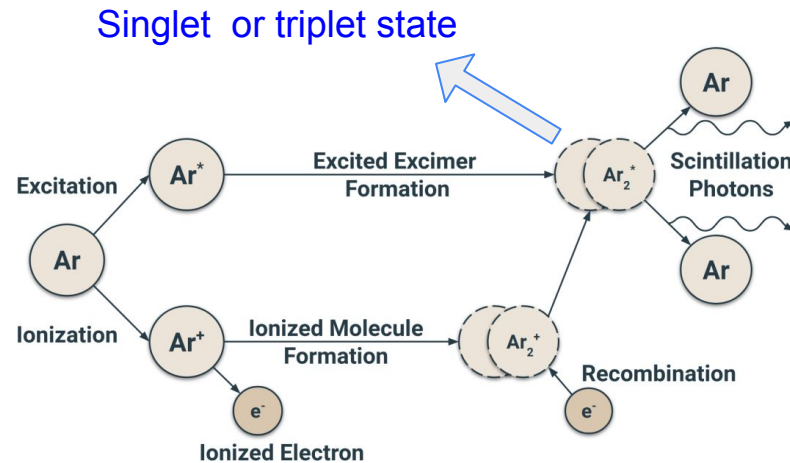
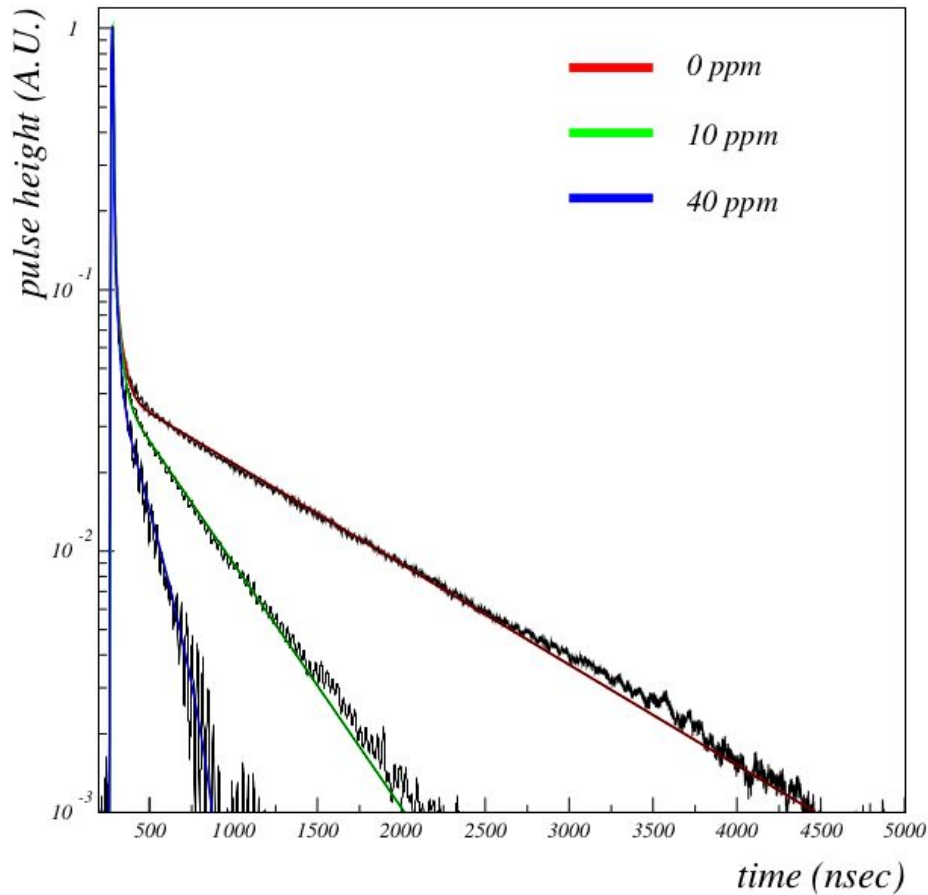


T. Heindl *et al* 2010 *EPL* **91** 62002



Time components

R Acciarri *et al* 2010 *JINST* **5** P06003



$$I(t) = A_S \exp\left(-\frac{t}{\tau_S}\right) + A_T \exp\left(-\frac{t}{\tau_T}\right)$$

$$\tau_S = 6 \text{ ns}$$

$$\tau_T = 1600 \text{ ns}$$

Scintillation time and yield

Composition of fast/slow (singlet/triplet) depends on particle Linear Energy Transfer (LET).

Muon/e⁻: 23 % (fast) and 77% (slow)

Alphas: 77 % (fast) and 23 % (slow)

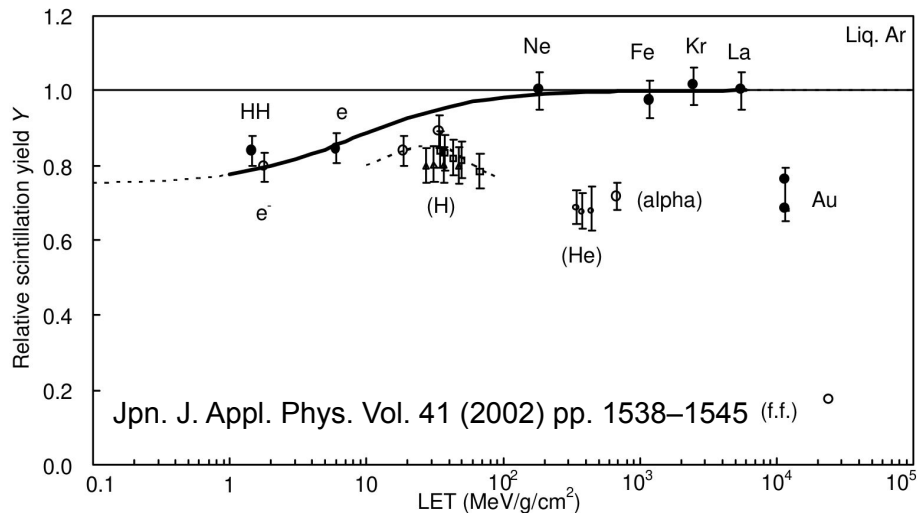


TABLE I. Decay times for the fast τ_S and the slow τ_T components of luminescence from liquid argon. The intensity ratios I_S/I_T of the fast component to the slow component are also shown. F.F. stands for fission fragments. All decay times are in nsec.

Particle	τ_S	τ_T	I_S/I_T	Reference
Electron	6.3 ± 0.2	1020 ± 60	0.083	Kubota <i>et al.</i> ^a
	(5.0 ± 0.2)	(860 ± 30)	(0.045)	$(E = 6 \text{ kV/cm})^a$
	4.6	1540	0.26	Carvalho and Klein ^b
	4.18 ± 0.2	1000 ± 95		Keto <i>et al.</i> ^c
		1110 ± 50		Suemoto and Kanzaki ^d
	6 ± 2	1590 ± 100	0.3	This work
α	~ 5	1200 ± 100		Kubota <i>et al.</i> ^c
	4.4	1100	3.3	Carvalho and Klein ^b
	7.1 ± 1.0	1660 ± 100	1.3	This work
F.F.	6.8 ± 1.0	1550 ± 100	3	This work

Ph. Rev. B 27 (1983), 5279

Scintillation yield also depend on LET.

Muon/e⁻ ~ 0.8

Alphas ~ 0.7

Scintillation time and yield

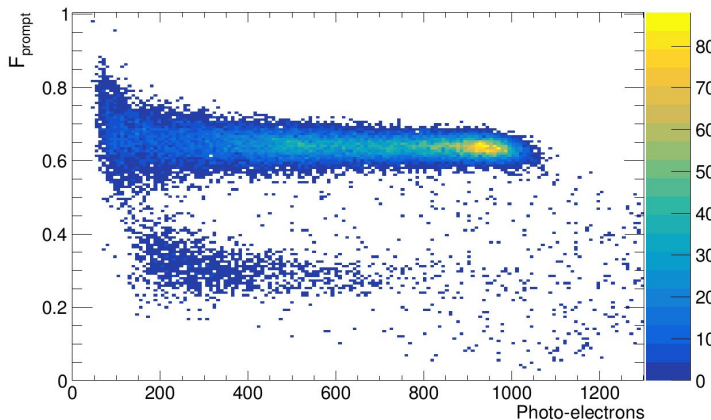
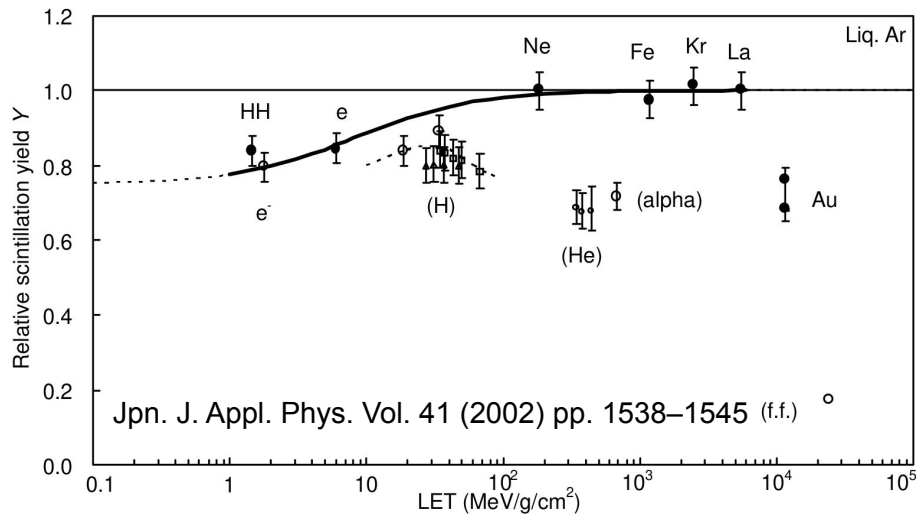


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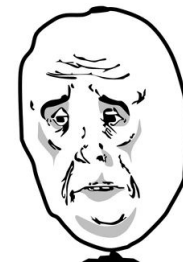
Ph. Rev. B 27 (1983), 5279



Scintillation yield also depend on LET.

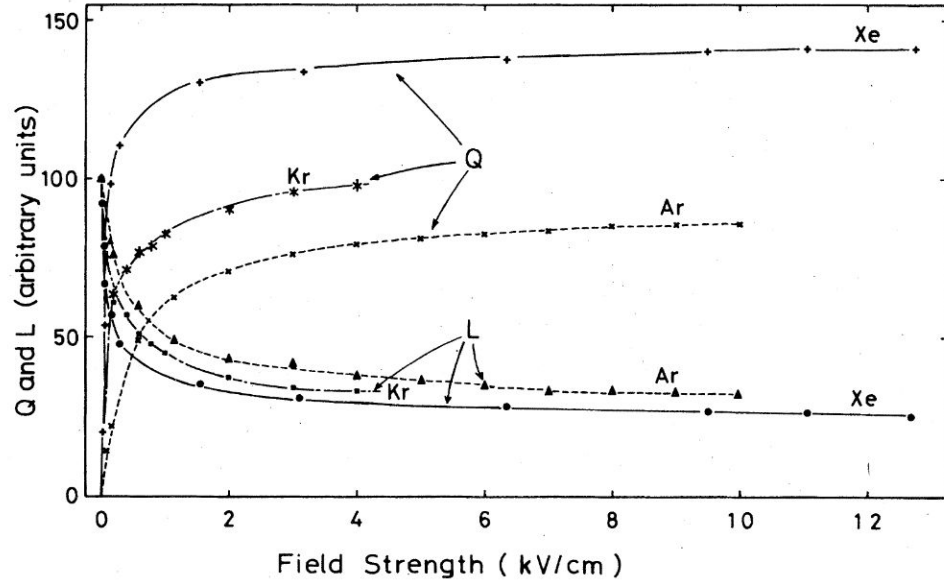
Muon/ e^- ~ 0.8

Alphas ~ 0.7



Electric field (simplified model)

Phys. Rev. B 20, 3486

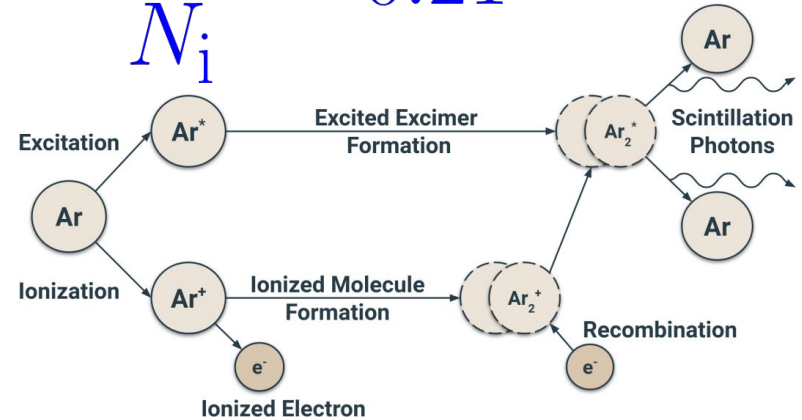


$$Q = N_e = N_i R_C$$

$$L = N_\gamma = N_{\text{ex}} + N_i(1 - R_C)$$

$$Q + L = N_{\text{ex}} + N_i = \Delta E / W_\gamma \quad 19.5 \pm 1 \text{ eV}$$

$$\frac{N_{\text{ex}}}{N_i} \sim 0.21$$

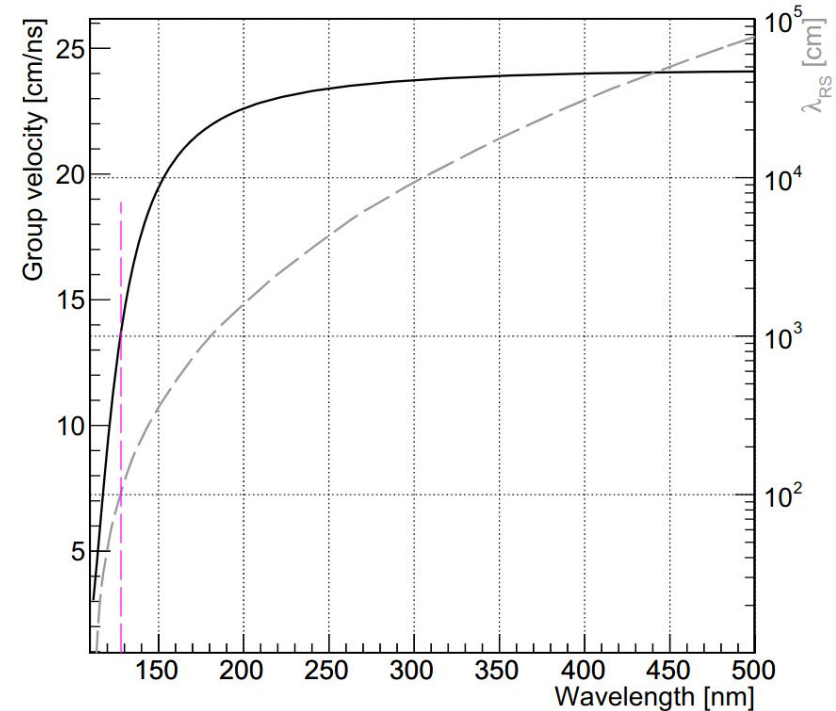
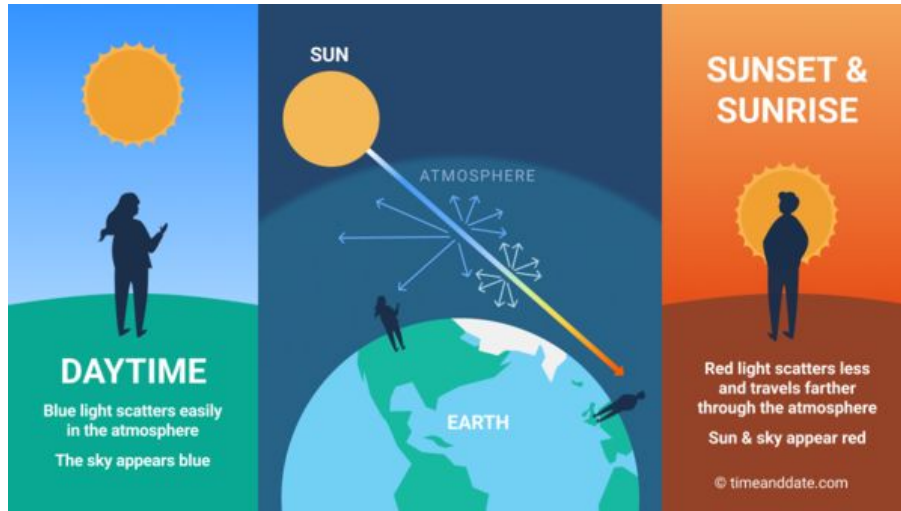


At 500 V/cm we have about 60% of light.
For muons, this corresponds to 24,000 photons/MeV

(See backup for estimating number of photons)

Light propagation

- Pure LAr is transparent to its own scintillation radiation
 - Attenuation is given by an exponential with decay length of ~ 20 m (3 ppm N_2)
- During propagation through LAr VUV photons may undergo elastic interactions on Ar atoms \Rightarrow **Rayleigh scattering**

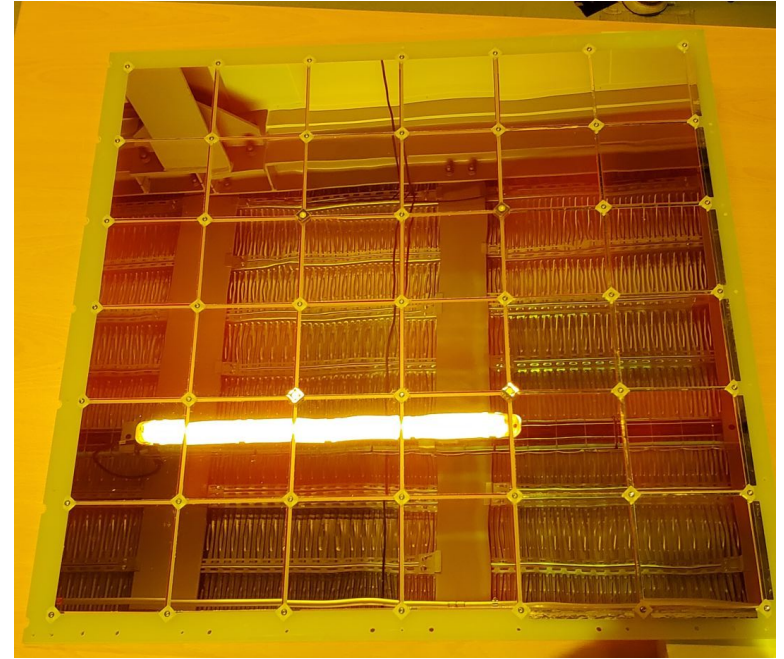
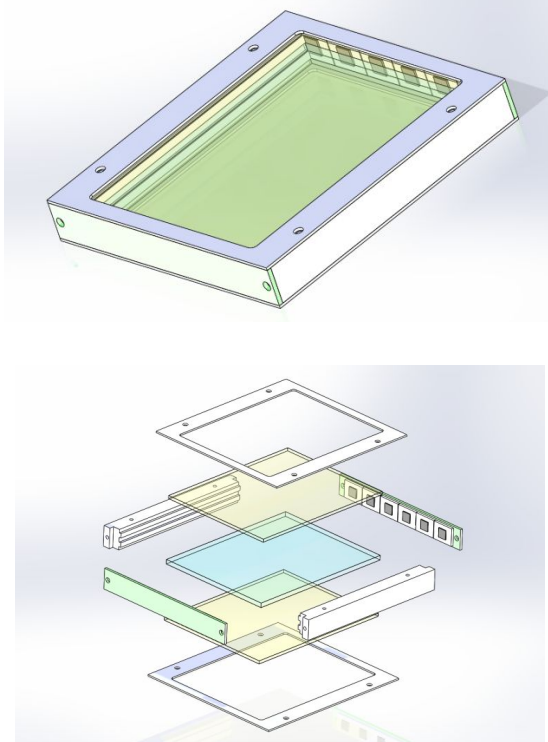


Detection

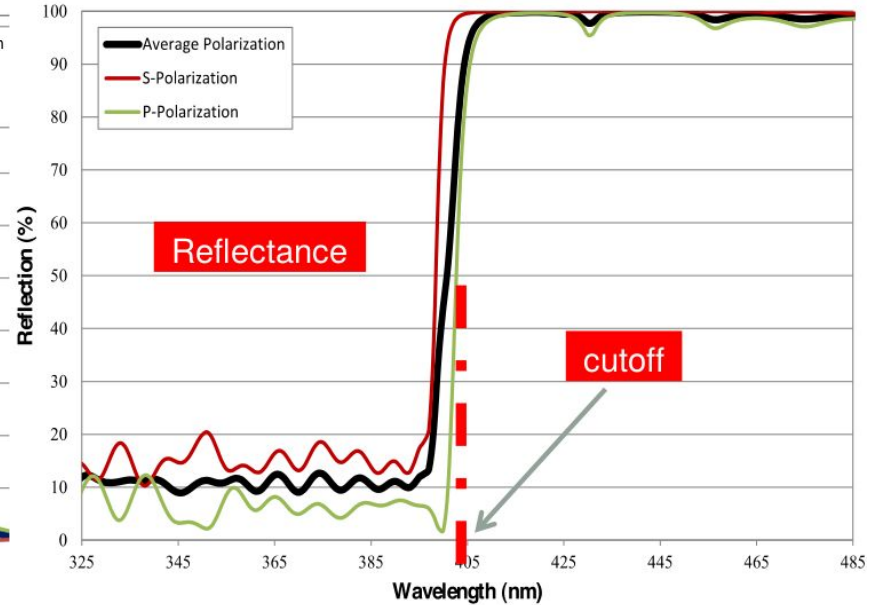
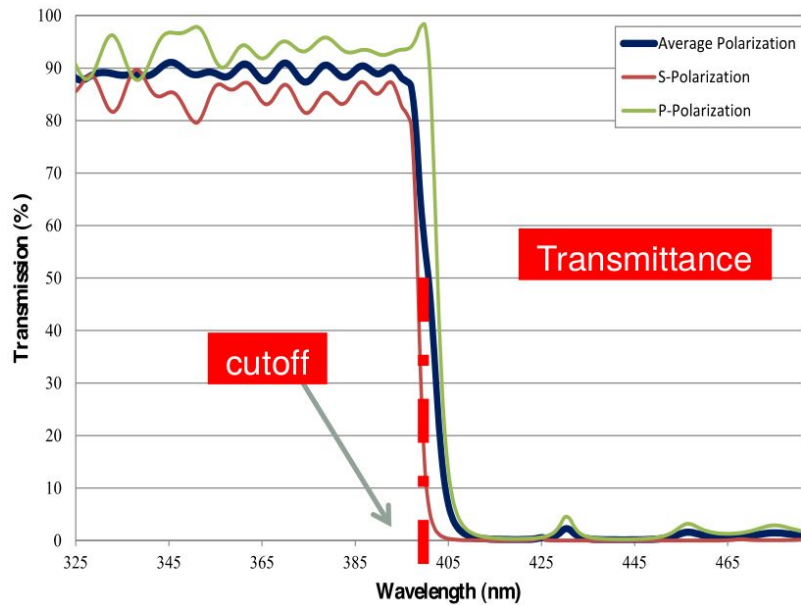
- Photon detection system (PDS) motivation
- X-Arapuca working principle

Photon Detection system - PDS

- Detecting 127 nm light is challenging. Besides, HD and VD requires that the photon detectors must have no more than 2 cm in thickness



X-Arapuca - Working principle

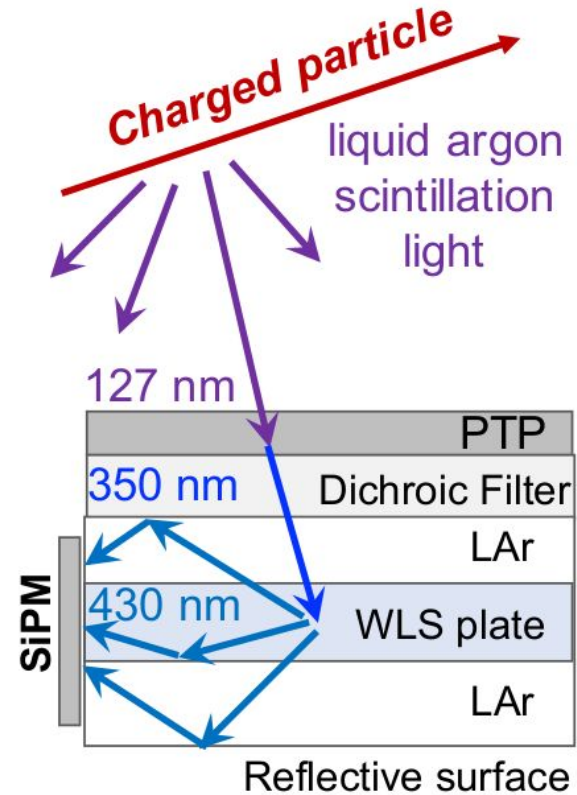
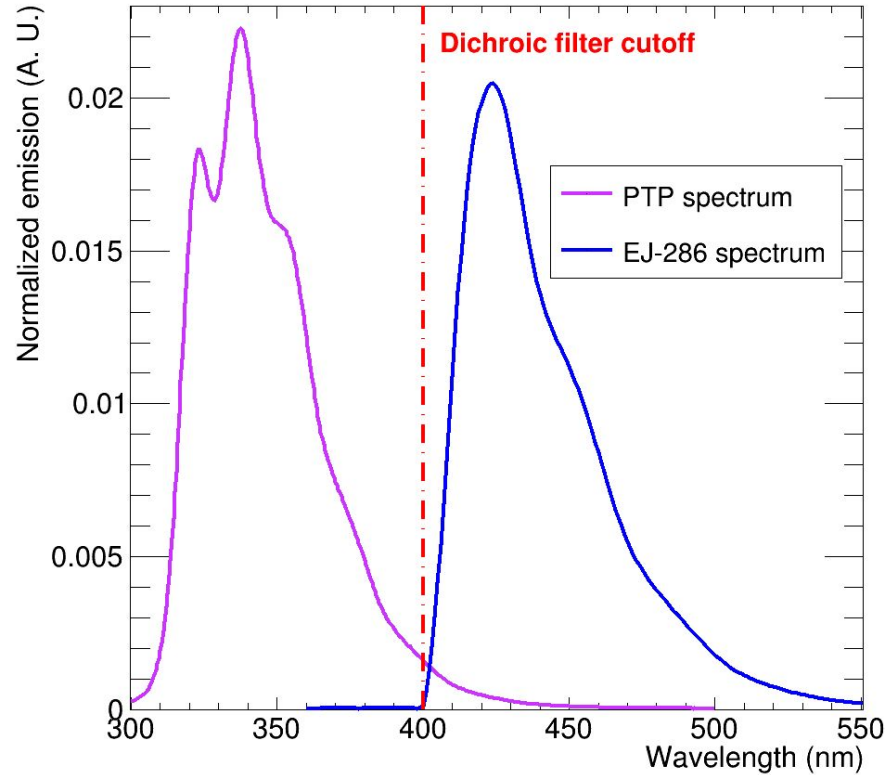


A.A. Machado and E. Segreto 2016 *JINST* 11 C02004

- The device makes use of a dichroic filter in combination with two wavelength shifters (WLS)

X-Arapuca - Working principle

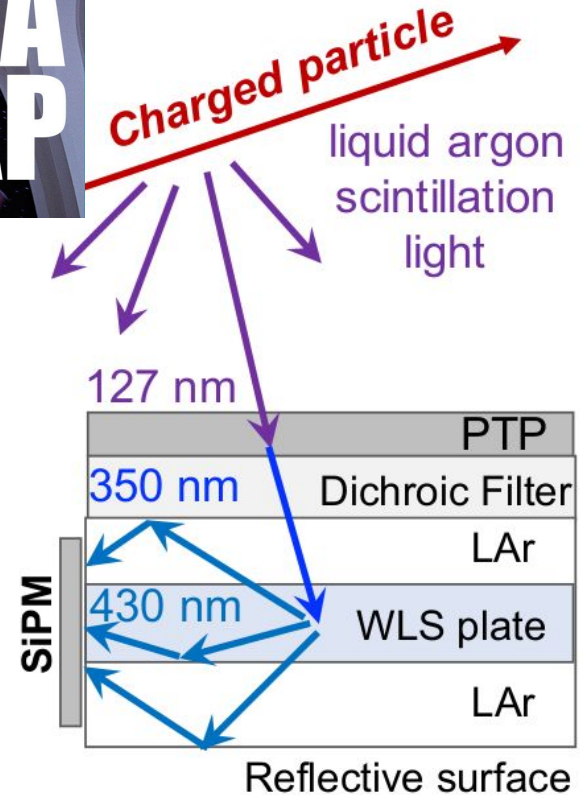
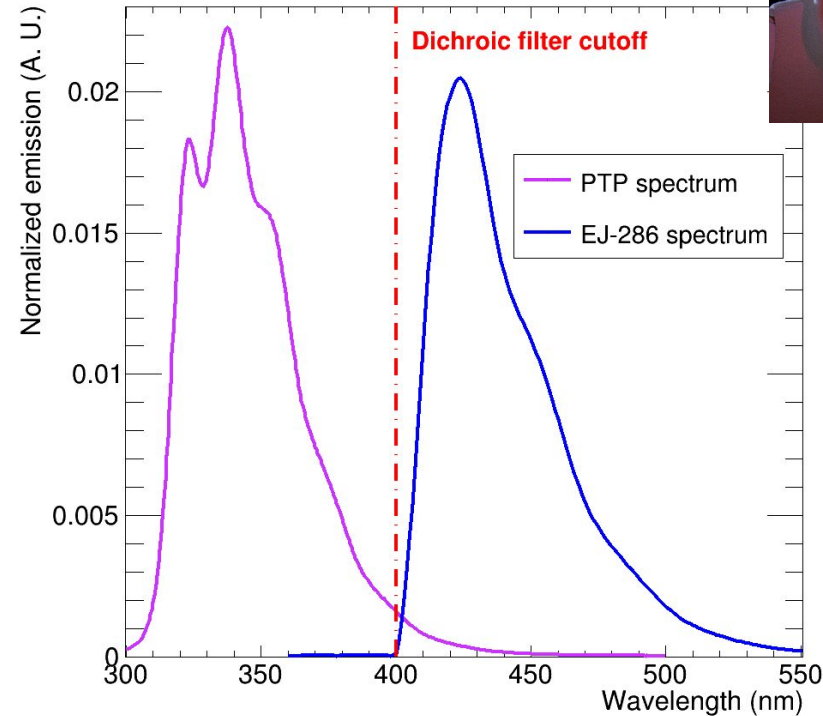
PTP → p-Terphenyl
SiPM → Silicon photomultiplier



X-Arapuca - Working principle

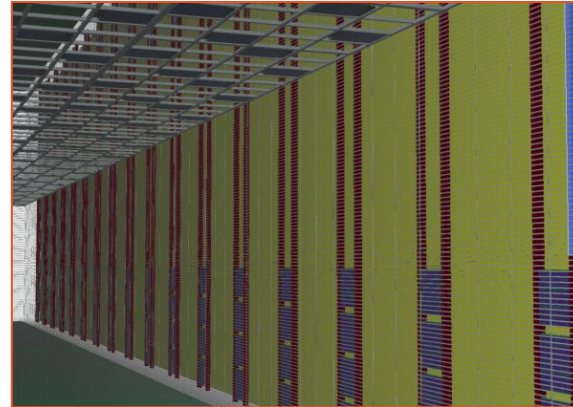
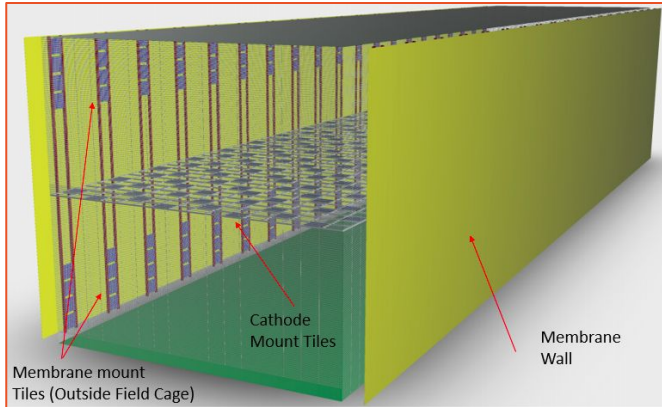
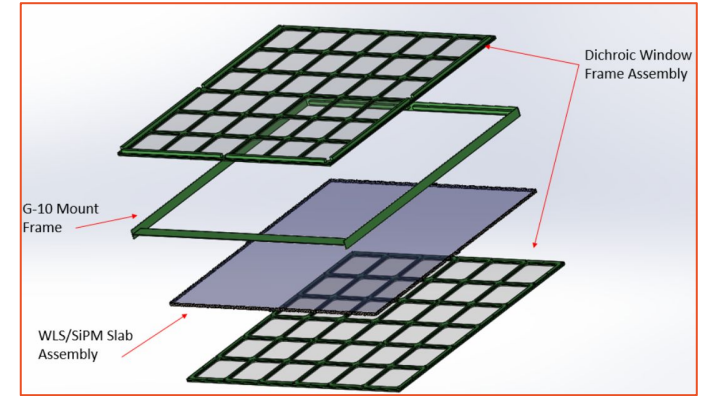


PTP → p-Terphenyl
SiPM → Silicon photomultiplier



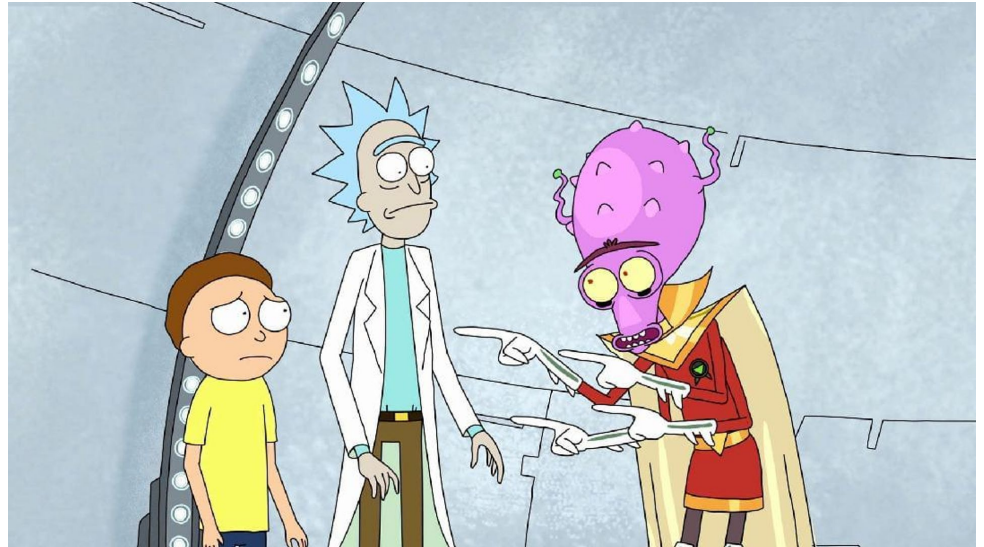
Photon Detection system - PDS

- The PDS is based on the X-Arapuca device
- A total of 2 x 80 Silicon Photomultipliers (SiPMs) per module
- 2x36 dichroic filters coated with pTP
- These devices are installed on the Cathode at -300 kV
 - Power supply and signal must be transmitted over non-conducting materials (not this talk)



Simulation

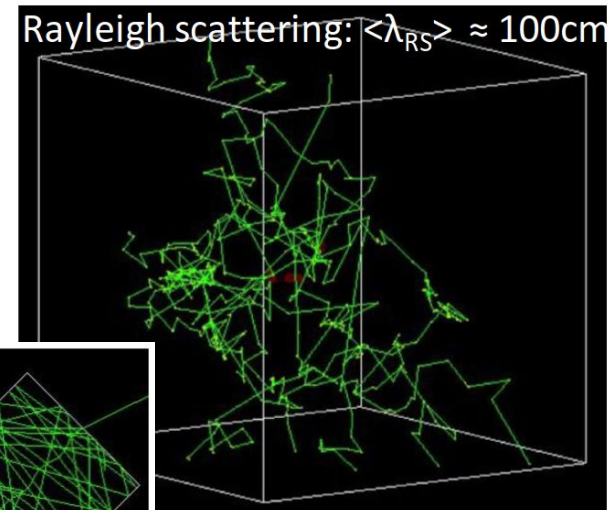
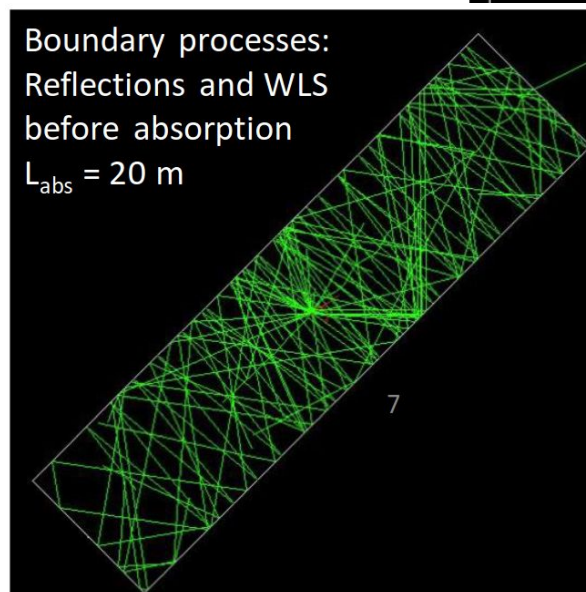
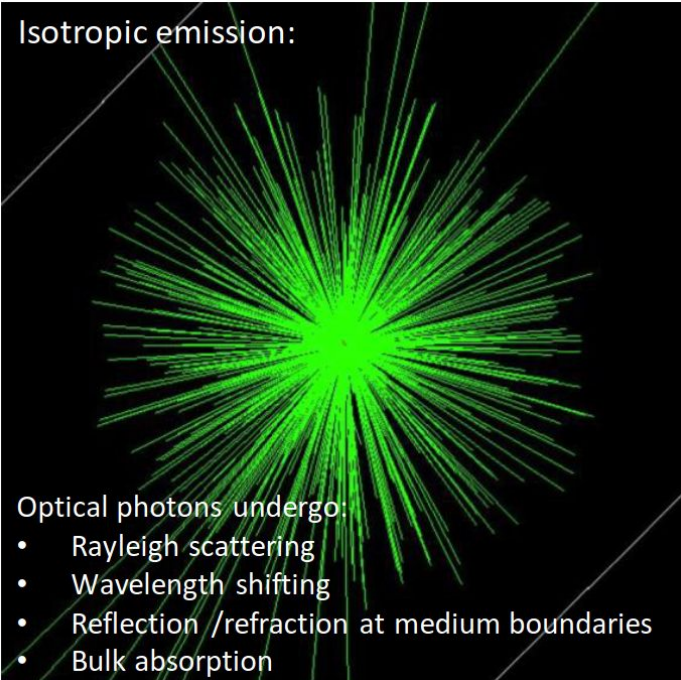
- What about the simulation?
- It takes into account everything said up to here:
 - Emission spectrum
 - Time response
 - Scintillation yield
 - Propagation
 - Detection



Simulation

- Different modes of simulation
 - Full optical simulation (extremely slow)
 - Requires definition of all optical properties
 - Fast optical simulation (faster, but less precise)
 - Still need to run full optical at least once
 - Majority of optical properties "burned in"
 - Semi-analytic and optical library
- Brief description of LArSoft output

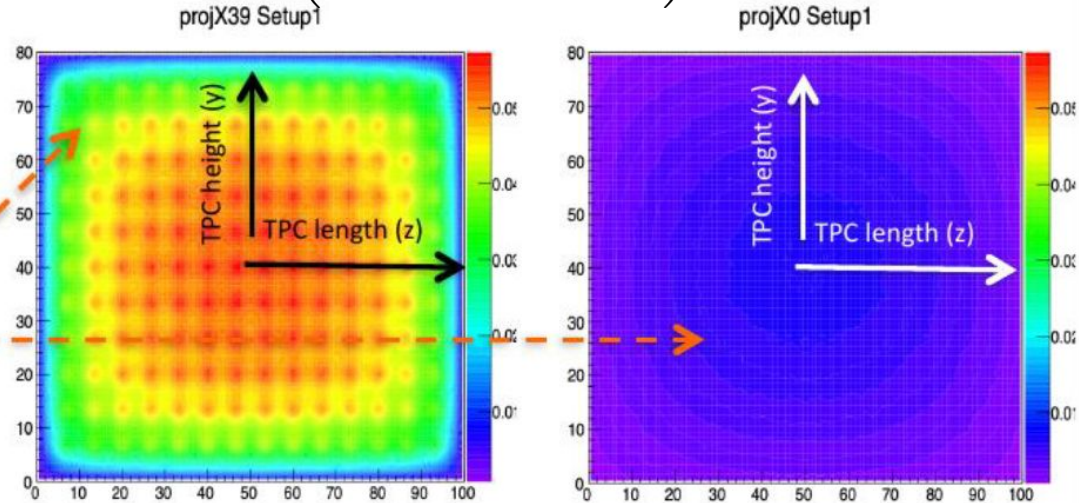
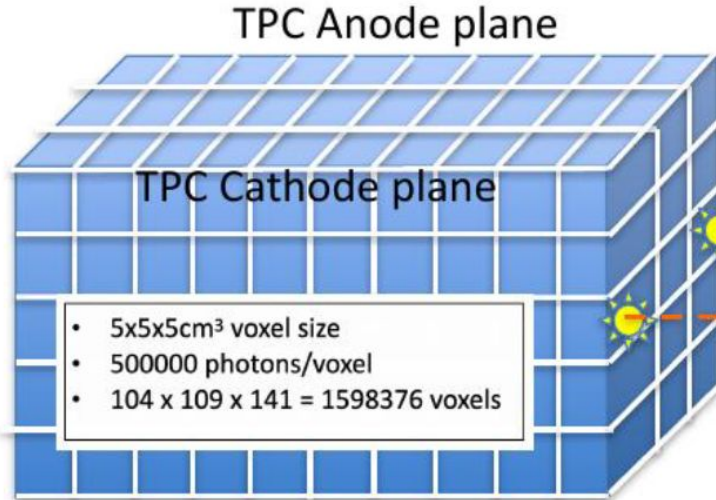
Full optical light simulation



From Andrzej Szalc presentation

Fast optical model: Optical Library

$$\langle N \rangle_{\text{Det-hits}} = \left(\frac{dE}{dx} \times \text{Length}_{\text{step}} \right) \times \text{LY} \times \text{visibily}_{\text{step}}^{\text{Det}}$$

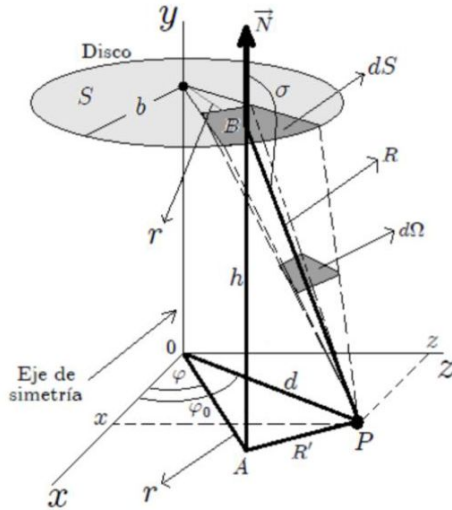


From Andrzej Szalc presentation

- Resolution depends on voxel sizes:
 - granularity effects at short distances
- Optical library size scales with detector size and number of photon detectors
 - Difficult to get working in DUNE

Fast optical model: Semi-Analytic

From Andrzej Szelc presentation



$$\Omega = h \int_0^{2\pi} \int_0^b \frac{r}{[h^2 + r^2 + d^2 - 2rd \cos(\varphi_0 - \varphi)]^{3/2}} dr d\varphi$$

Eur. Phys. J. C 81, 349 (2021)

- Given a dE/dx in a point (x, y, z) we want to predict the number of hits in our optical detector (x_i, y_i, z_i)
- Isotropic scintillation emission makes the problem “almost” geometric

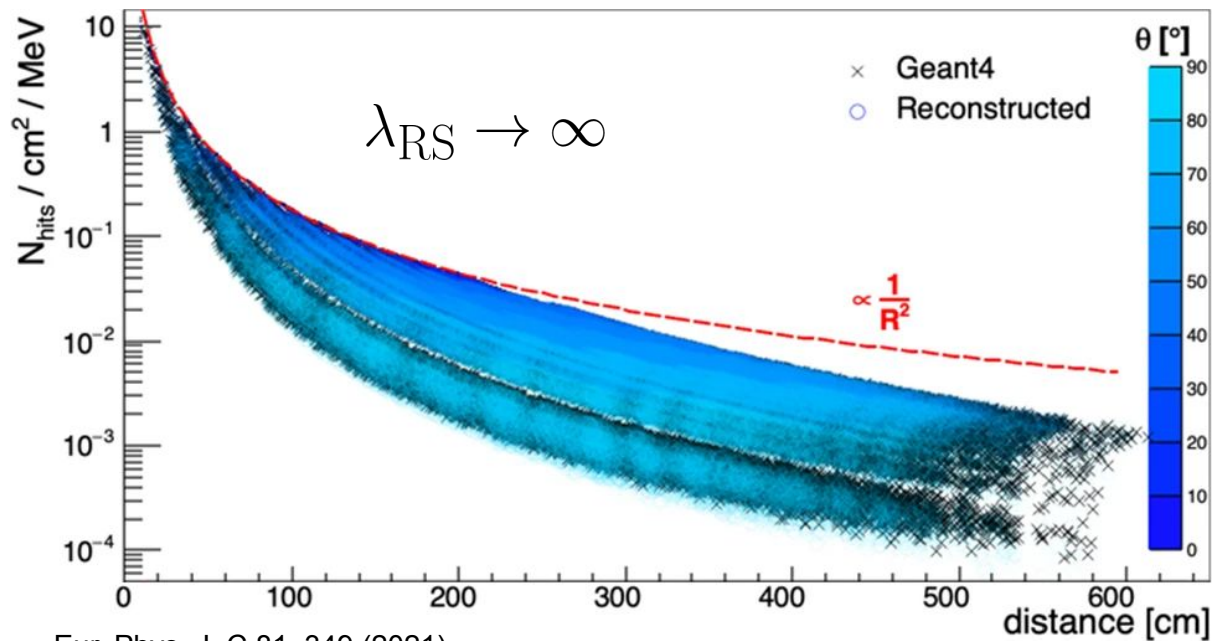
$$N_{\Omega} = e^{-\frac{d}{\lambda_{\text{abs}}}} \times \Delta E \times S_{\gamma}(\mathcal{E}) \times \frac{\Omega}{4\pi}$$

λ_{abs} = LAr absorption

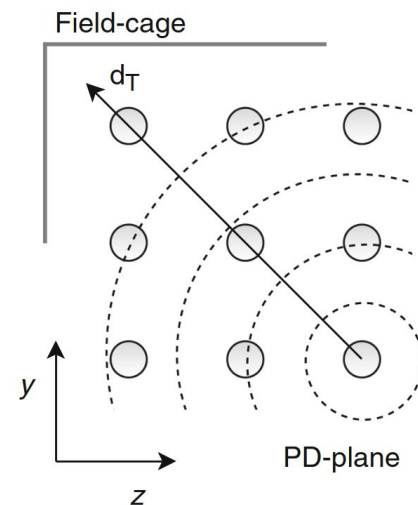
$S_{\gamma}(\mathcal{E})$ = Scintillation Yield as function of electric field

ΔE = Energy deposited

Fast optical model: Semi-Analytic



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In the DUNE-like geometry, energy depositions are spaced in approximately 25 cm steps in the drift direction and 100 cm steps in d_T

λ_{abs} = LAr absorption

$$N_{\Omega} = e^{-\frac{d}{\lambda_{abs}}} \times \Delta E \times S_{\gamma}(\mathcal{E}) \times \frac{\Omega}{4\pi}$$

$S_{\gamma}(\mathcal{E})$ = Scintillation Yield as function of electric field

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Fast optical model: Semi-Analytic

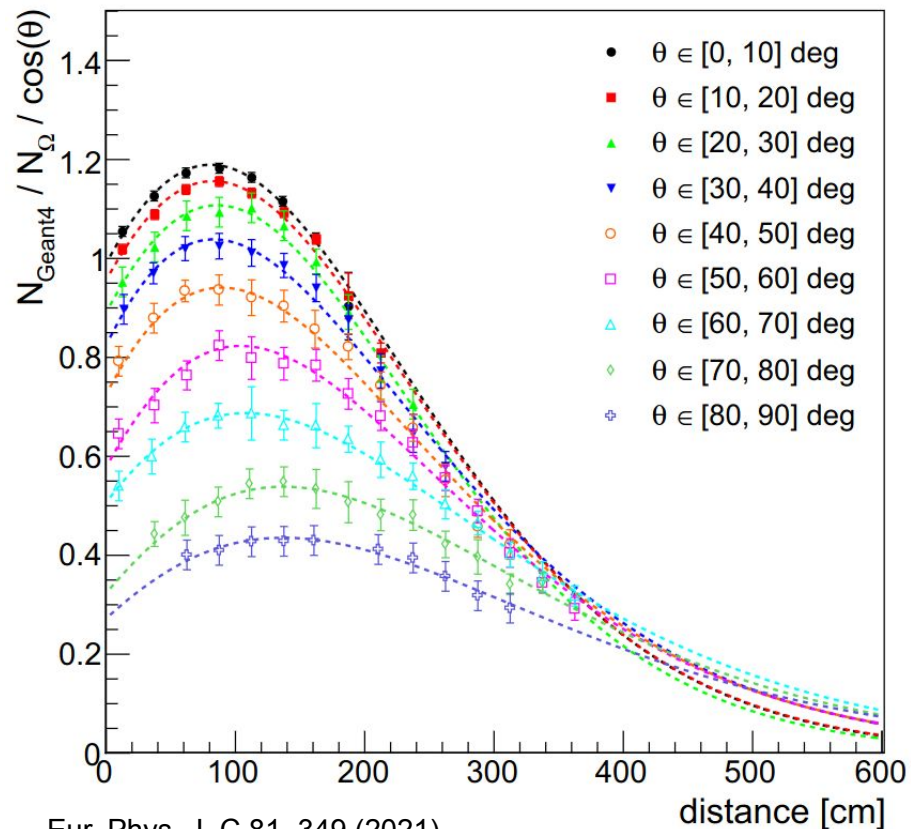
- Implementation of Rayleigh scattering
 - Correction using Geant4 simulation
- Correction for detector size and geometry (not included here)

Gaisser–Hillas (GH) functions:

$$GH(d) = N_{\max} \left(\frac{d - d_0}{d_{\max} - d_0} \right)^{\frac{d_{\max} - d_0}{\Lambda}} e^{-\frac{d_{\max} - d}{\Lambda}}$$

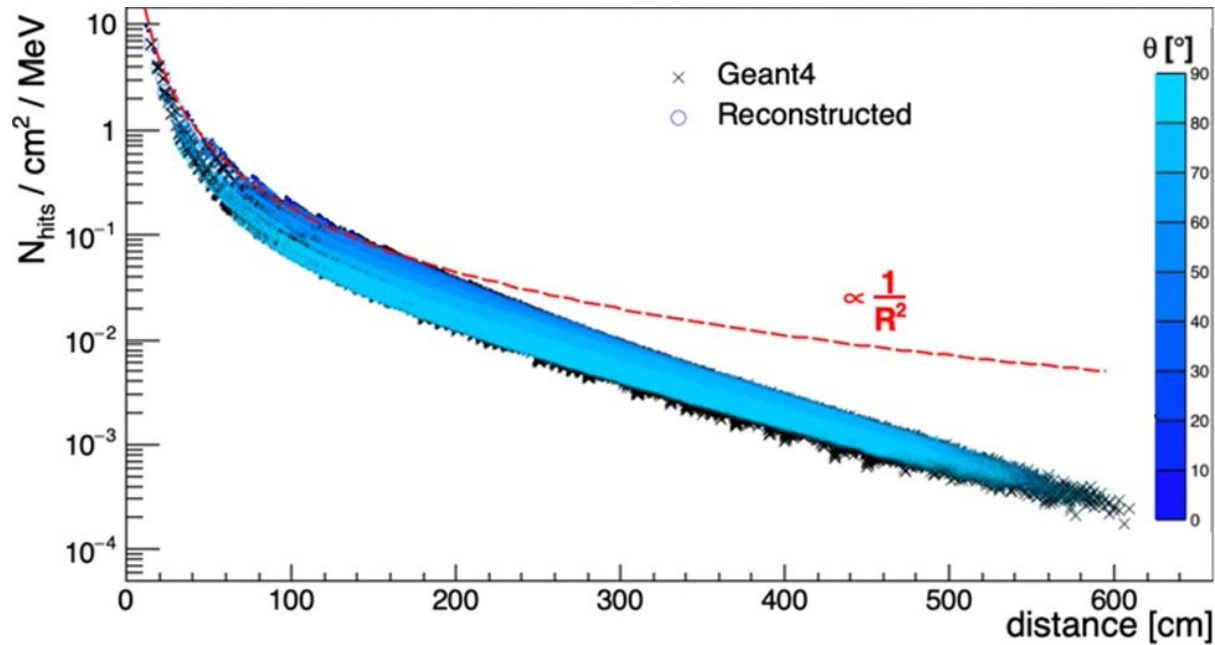
“where N_{\max} is the maximum of the function located at a distance d_{\max} , and d_0 and Λ are parameters describing the width of the distribution”

$$N_{\gamma} = N_{\Omega} \times GH'(d, \theta, d_T) / \cos(\theta)$$



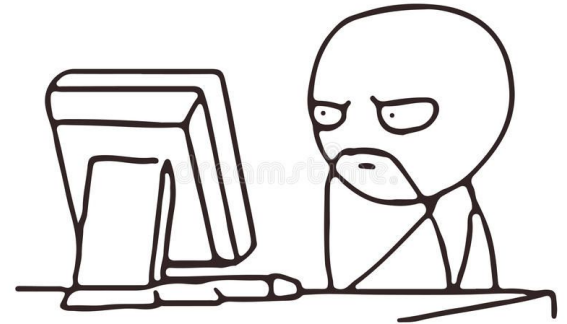
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Fast optical model: Semi-Analytic



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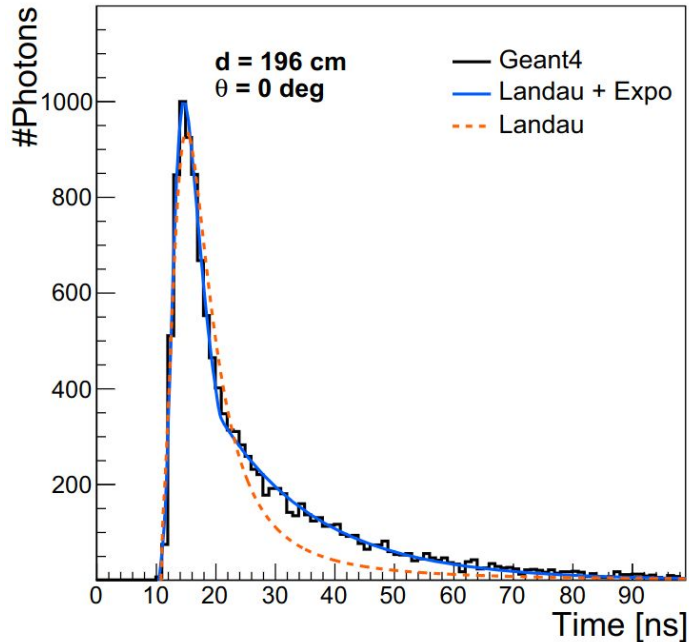
One thing is still missing



$$N_\gamma = N_\Omega \times GH'(d, \theta, d_T) / \cos(\theta)$$

Time structure of detected signals

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Example of the distribution of direct photons arrival times due to only transport effects

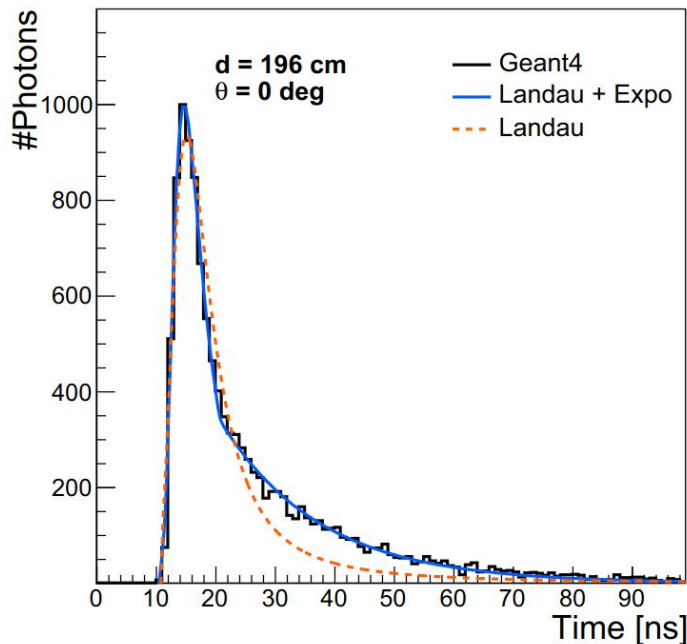
- Empirically described by a Landau and exponential for all emission points

$$t_t(x) = \underbrace{N_1 \frac{1}{\xi} \frac{1}{2\pi i} \int_{c-i\infty}^{c+i\infty} e^{\lambda s + s \log s} ds}_{\text{Landau}} + \underbrace{N_2 e^{\kappa x}}_{\text{Exponential}}$$

“where $\lambda = x - \mu/\xi$, with μ and ξ commonly referred as the landau most probable value and width parameters respectively, κ is the slope of the exponential and N_1 and N_2 are normalisation constants.”

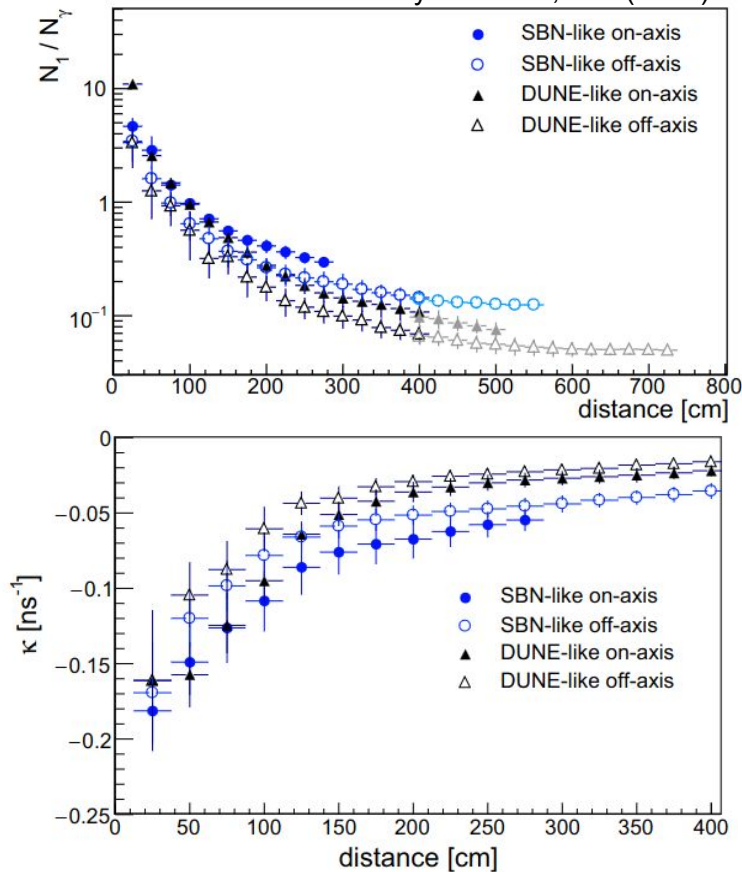
Time structure of detected signals

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- [0, 45] on-axis
- [45, 90] off-axis
- Lighter grey and blue points are landau only approach

Eur. Phys. J. C 81, 349 (2021)

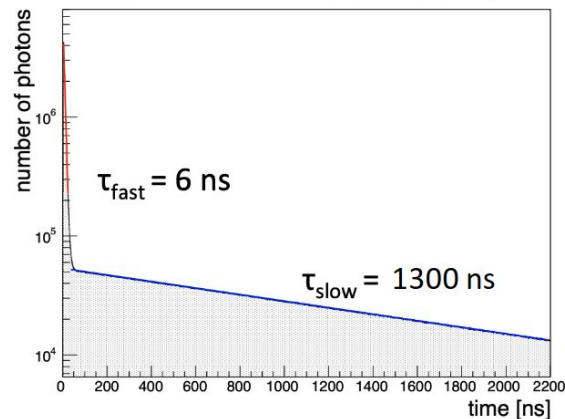


$$t_t(x) = \underbrace{N_1 \frac{1}{\xi} \frac{1}{2\pi i} \int_{c-i\infty}^{c+i\infty} e^{\lambda s + s \log s} ds}_{\text{Landau}} + \underbrace{N_2 e^{\kappa x}}_{\text{Exponential}}$$

Time structure of detected signals

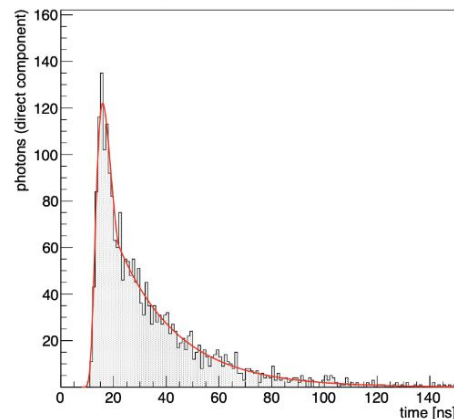
Scintillation (emission):

$$0.3 \times \tau_{\text{fast}} (6 \text{ ns}) + 0.7 \times \tau_{\text{slow}} (1300 \text{ ns})$$

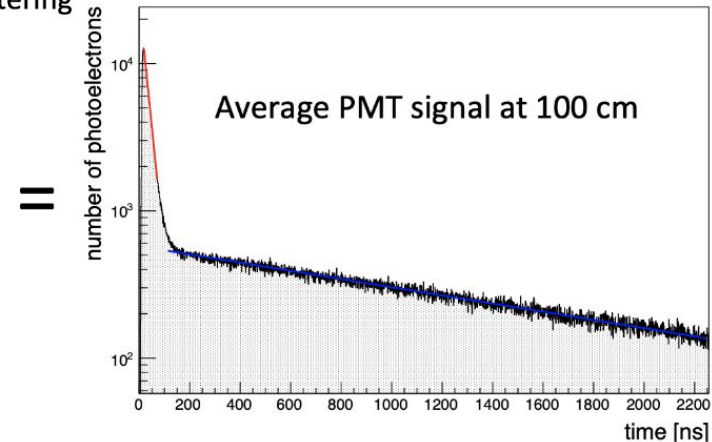


Propagation:

Direct transportation + Rayleigh Scattering



From Andrzej Szecel presentation



- The final time response of the detector will take into account:

- Emission time
- Propagation time

- Wavelength shifter delay
- Detector time

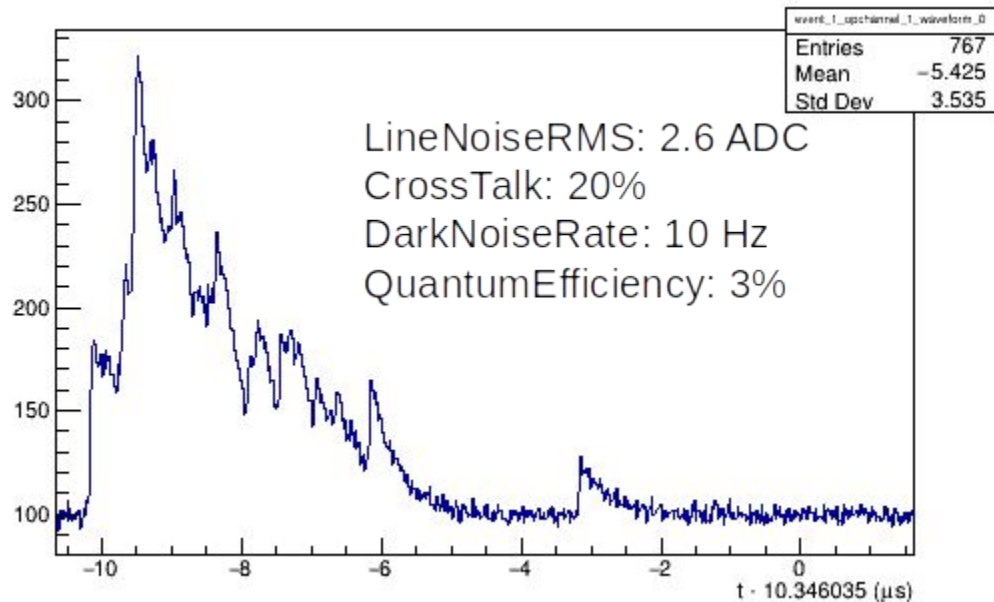
Time structure of detected signals

Time histogram of photons at each PD

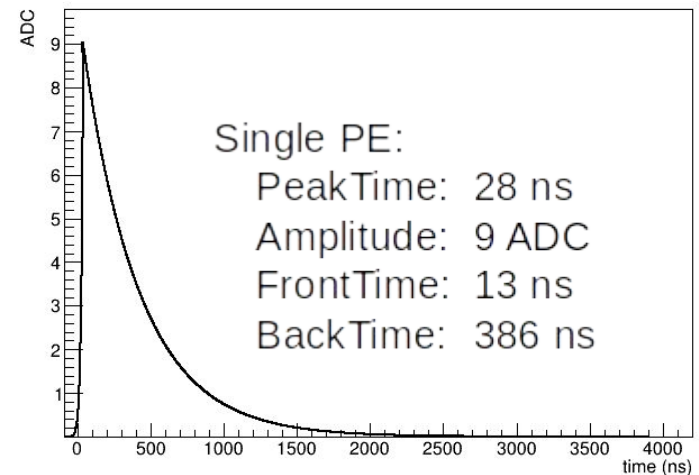


Detector
simulation

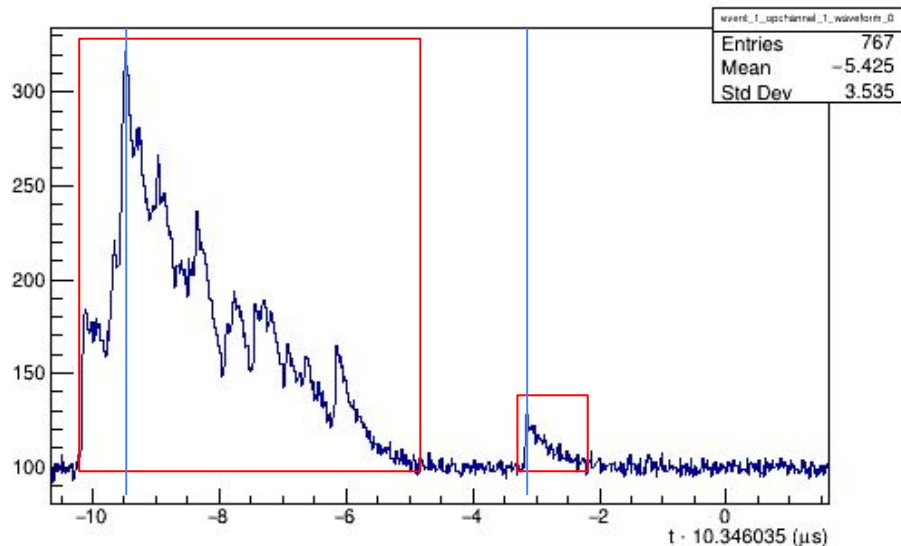
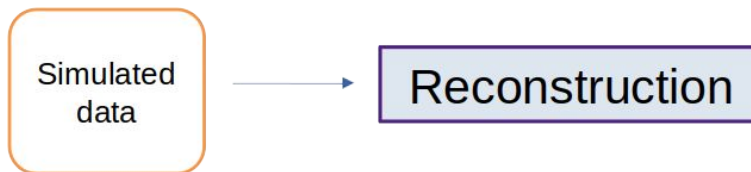
Simulated
data



Courtesy of Laura Paulucci



Reconstruction of events



Courtesy of Laura Paulucci

Reconstruction

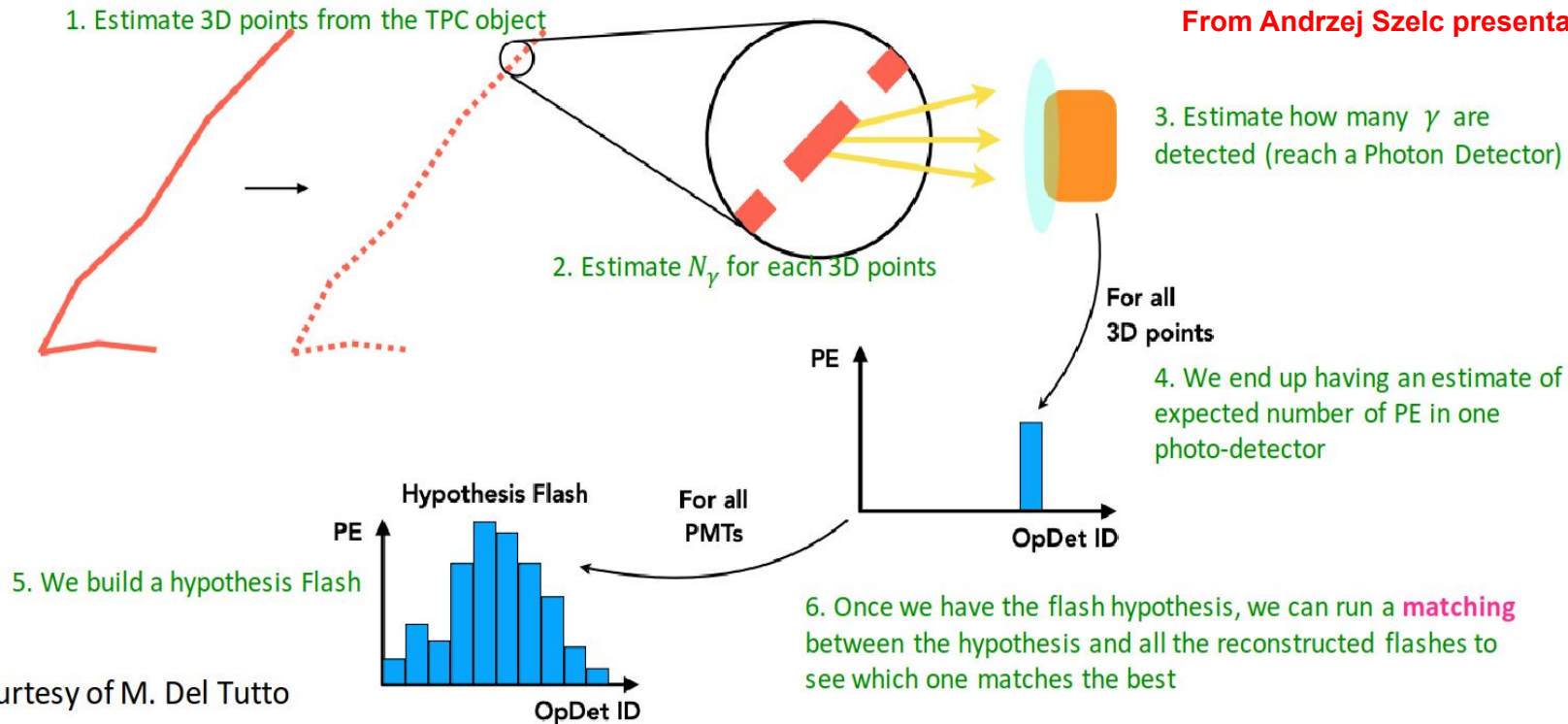
Hit finding:

searches for peaks on individual waveforms channel-by-channel, identifying the time and the total amount of PEs



Reconstruction of events

From Andrzej Szecel presentation



Courtesy of M. Del Tutto

Analysis

- Where you can find some actual data
- The structure of this data
- Some of the main analysis performed up to now

Some actual data

- Since Dec. 2021 we have been collecting data with the coldbox.
 - Unfortunately, one need to understand the setup by asking / tracing back old slides as the configuration changed quite often there
 - Besides, there was no simulation implemented for the coldbox and, at this point, Module-0 will soon enough collect data to be analyzed.
 - However if anyone interest, please let me know :D
 - Nevertheless, the data of past coldbox runs can be found in Ixplus:
`/eos/experiment/neutplatform/protodune/experiments/ColdBoxVD`
 - I will quickly show how the data was collected and the main analysis up to now (if there is time hehe)
- Future data of Module-0
 - Where? Don't know
 - Format? Probably binary in similar way of what I am going to show now

Some actual data

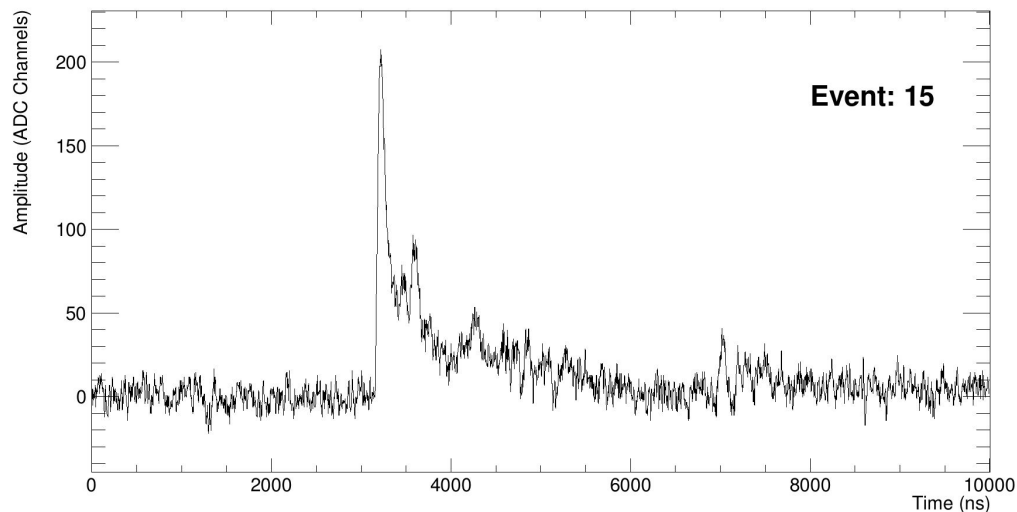
Data acquisition done with CAEN Digizer DT5730SB (2 Vpp 14 bits 250* MS/s):

- Data stored at Ixplus:

`/eos/experiment/neutplatform/protodune/experiments/ColdBoxVD/`

- To access data, please, register in the np-comp e-group:

<https://e-groups.cern.ch/e-groups/EgroupsSearchForm.do>



Data format

```
> 20220615_LED_calibration_cathode_off
```

```
> run11_v4_275nm_20ns_4V7
```

```
> run12_v5_275nm_20ns_5V5
```

Run folder with brief description
(README file available)

```
0_wave3_v4_275nm_20ns_4V7.dat
```

```
0_wave4_v4_275nm_20ns_4V7.dat
```

```
1_wave3_v4_275nm_20ns_4V7.dat
```

```
1_wave4_v4_275nm_20ns_4V7.dat
```

Data format

```
> 20220615_LED_calibration_cathode_off
```

```
> run11_v4_275nm_20ns_4V7
```

```
> run12_v5_275nm_20ns_5V5
```

```
0_wave3_v4_275nm_20ns_4V7.dat
```

```
0_wave4_v4_275nm_20ns_4V7.dat
```

```
1_wave3_v4_275nm_20ns_4V7.dat
```

```
1_wave4_v4_275nm_20ns_4V7.dat
```

Run number, it refers to the order in which data were taken

Some block of information relevant for that specific run.

A README file helps to understand this block of information

Data format

```
> 20220615_LED_calibration_cathode_off
```

```
> run11_v4_275nm_20ns_4V7
```

```
> run12_v5_275nm_20ns_5V5
```

X_waveY

X are subruns with 10k events each
Y is the channel number

```
0_wave3_v4_275nm_20ns_4V7.dat
```

```
0_wave4_v4_275nm_20ns_4V7.dat
```

```
1_wave3_v4_275nm_20ns_4V7.dat
```

```
1_wave4_v4_275nm_20ns_4V7.dat
```

Data format

Example:

wave0: xArapuca A4ch2 (light blue fiber)

wave1: xArapuca A1ch1 (white fiber)

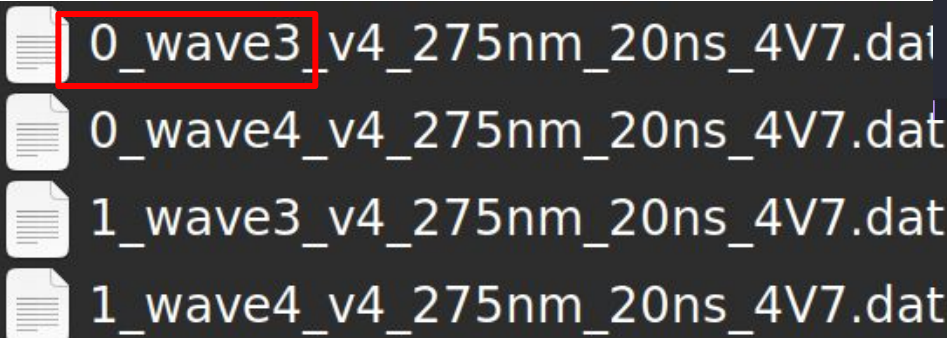
wave2: miniArapuca A4ch1 (green fiber)

wave3: miniArapuca A1ch2 (blue fiber)

X_waveY

X are subruns with 10k events each

Y is the channel number



0_wave3_v4_275nm_20ns_4V7.dat
0_wave4_v4_275nm_20ns_4V7.dat
1_wave3_v4_275nm_20ns_4V7.dat
1_wave4_v4_275nm_20ns_4V7.dat

```
February2023/README.org - Doom Emacs
February2023/README.org x  March2023run/README.org x +
14 :PROPERTIES:...
11
10 #+title: Data taking log
9
8 ● General description
7
6 Digitizer: CAEN DT5730SB \\
5 Sample rate: 250 MSamples/s (4 ns step) (unless specified) \\
4 Samples: 5000 (20 us total) unless specified \\
3 Digitizer: 2 Vpp 14 Bits (0.122 mV/ADC) \\
2 DC offset: 10% \\
1 Pretrigger: 50% or 30% \\
15 Acquisition done with Wavedump. \\
1 Each binary file has 10,000 events (unless specified). The data are
save with 6 headers and nsamples of data. 4 bytes per header and 2
bytes per sample. Please, refer to the wavedump manual. \\
2
3
4 | CAEN ADC | Device |
5 | Ch0 | miniArapuca 37V Argon2x2 |
6 | Ch1 | miniArapuca 37V Argon4 |
7 | Ch2 | DCem1.0VD |
8 | Ch3 | xArapuca v4 Ch2 |
9 | Ch4 | xArapuca v4 Ch1 |
10 | Ch5 | xArapuca v5 ch2 |
11 | Ch6 | xArapuca v5 ch1 |
12 | Ch7 | DCem1.2VD |
13
14
15 NOTE: Channels label were swaped in the Koheron for v4 and v5. The
values labeled here are correct
2.8k February2023/README.org 15:33 Top Org (+1)
```

Data is saved as binary (faster and lighter).

1 waveform consist of 6 headers and **n** samples

The HEADER is so composed (for all digitizer families except the 742 one):

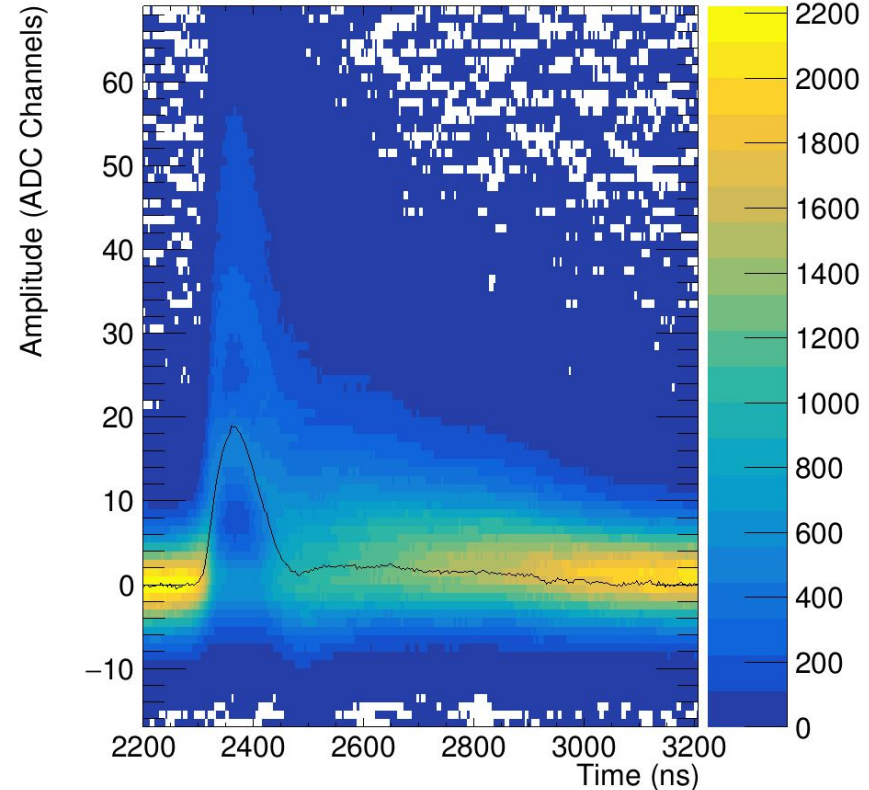
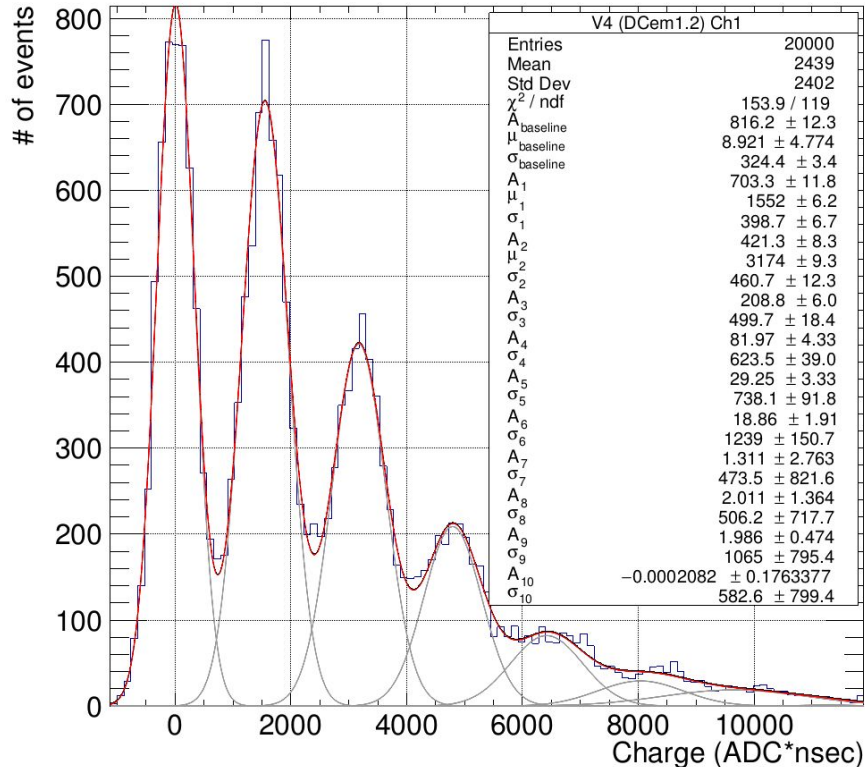
4 bytes	<header0> Event Size (i.e. header + samples)	This number is given in bytes. So if we have 5000 samples, this number will be:
4 bytes	<header1> Board ID	Headers = 24 bytes
4 bytes	<header2> Pattern (meaningful only for VME boards)	+ samples = 5000*2
4 bytes	<header3> Channel	Total = 10024 bytes
4 bytes	<header4> Event Counter	
4 bytes	<header5> Trigger Time Tag	

N * 2 bytes <N samples>

Let me know if you need an example code to read the data with Root.

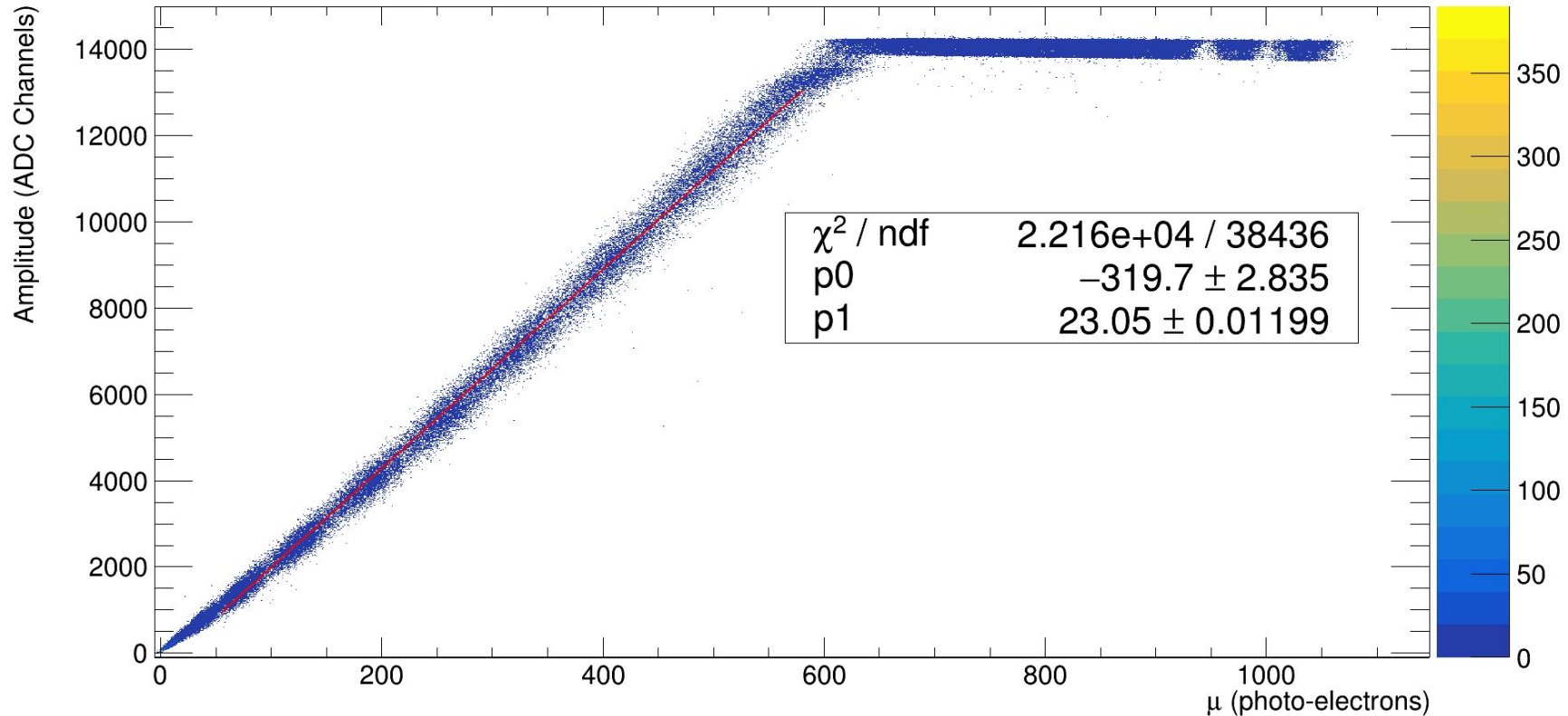
Main analysis up to now

Single photo-electron (SPE): uses low intensity LED flashed of light to detect one or more photons. Which results in what we call `SPE spectrum`



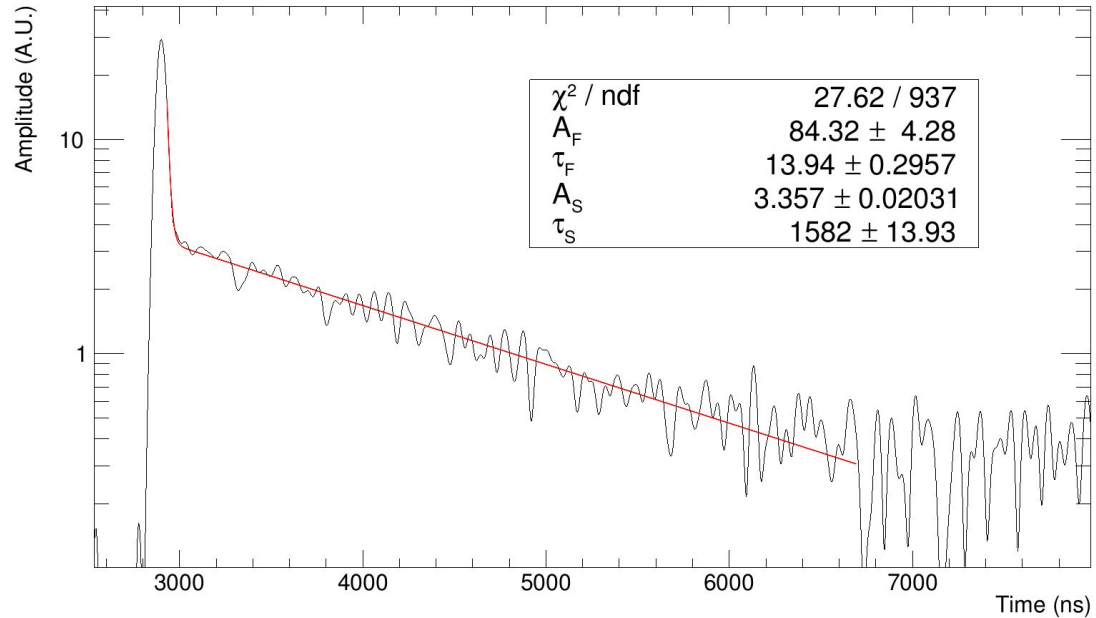
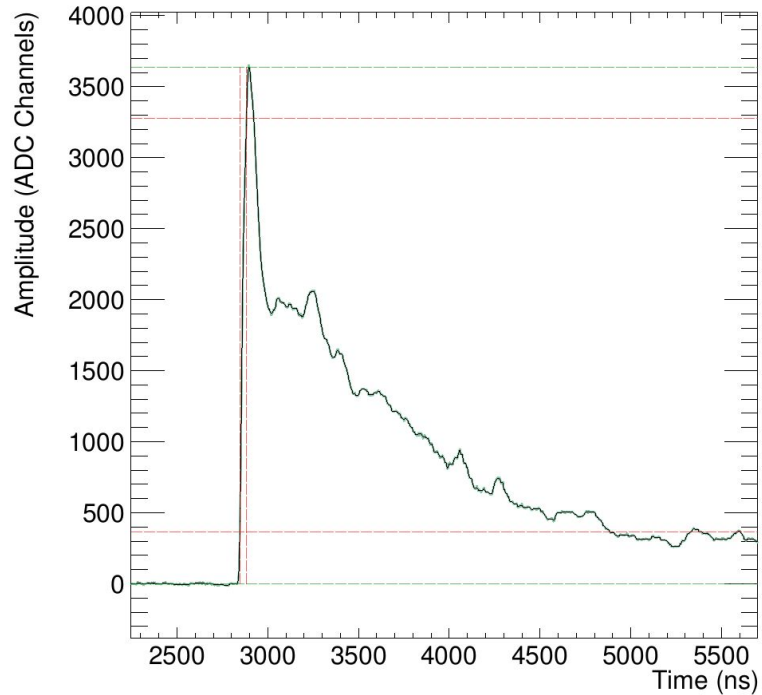
Main analysis up to now

Linearity and dynamic range: uses LED to check linear behaviour of detector/electronics over the entire dynamic range of the device.



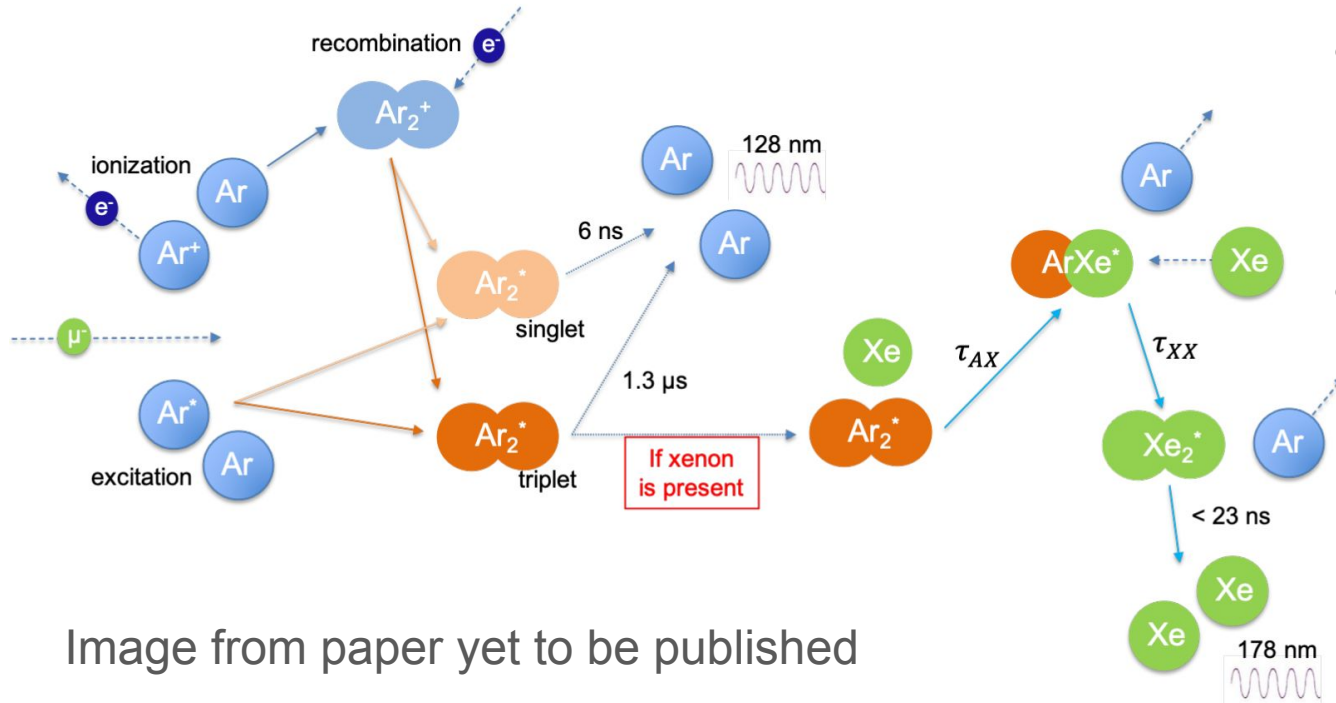
Main analysis up to now

Overall pulse shape: undershoot, overshoot, rise and fall time characterization of signals with LED and Cosmic (self-trigger data) data



If there is still time...

- Interesting past analysis with ProtoDUNE-SP



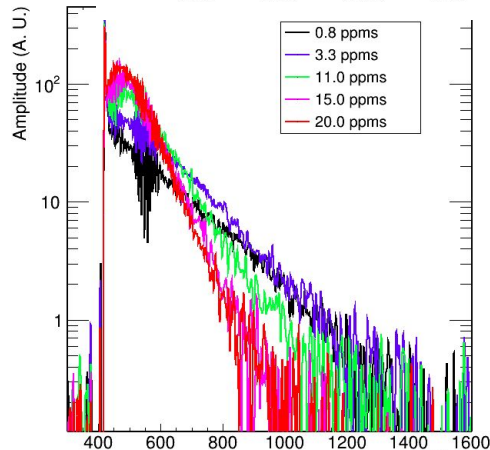
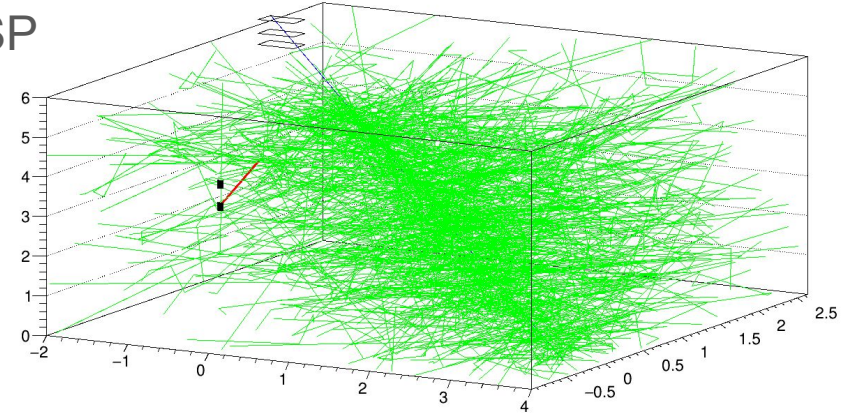
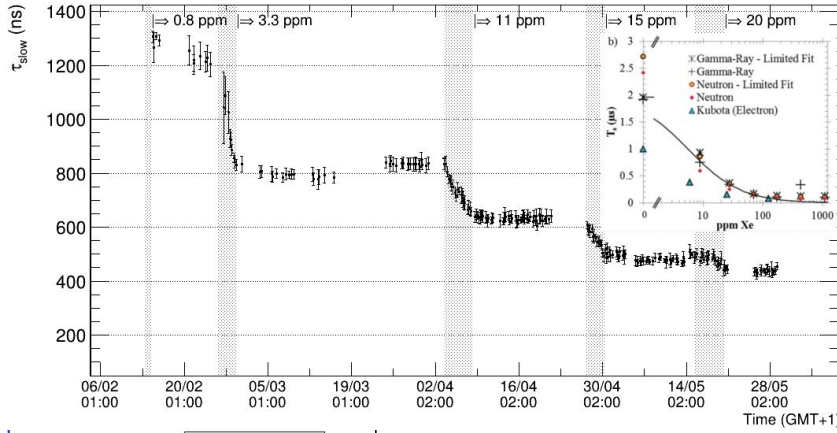
- Xenon doping

- Recover light that would be lost to nitrogen contamination
- Increase the wavelength of the photons:
 - Easier to detect
 - Higher Rayleigh scattering
- Possibly increase PID capability

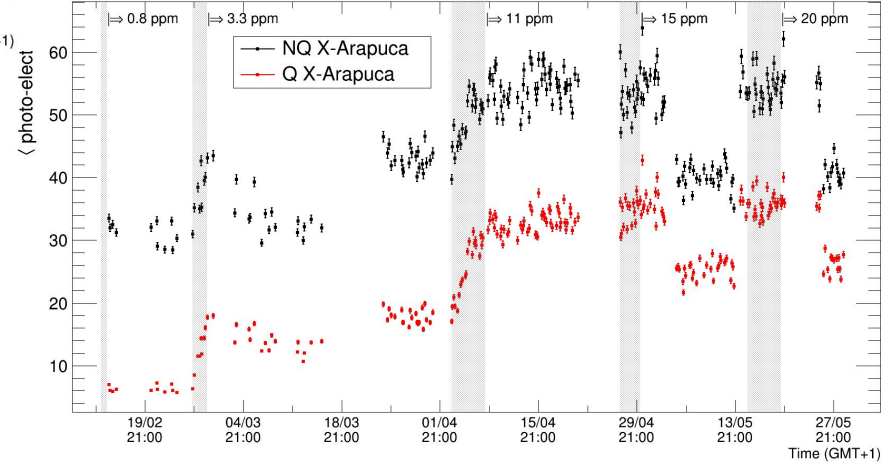
Image from paper yet to be published

If there is still time...

- Interesting past analysis with ProtoDUNE-SP

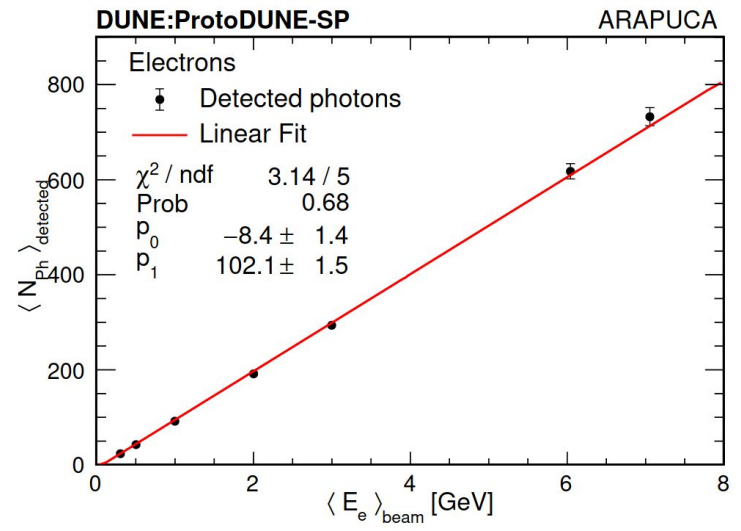
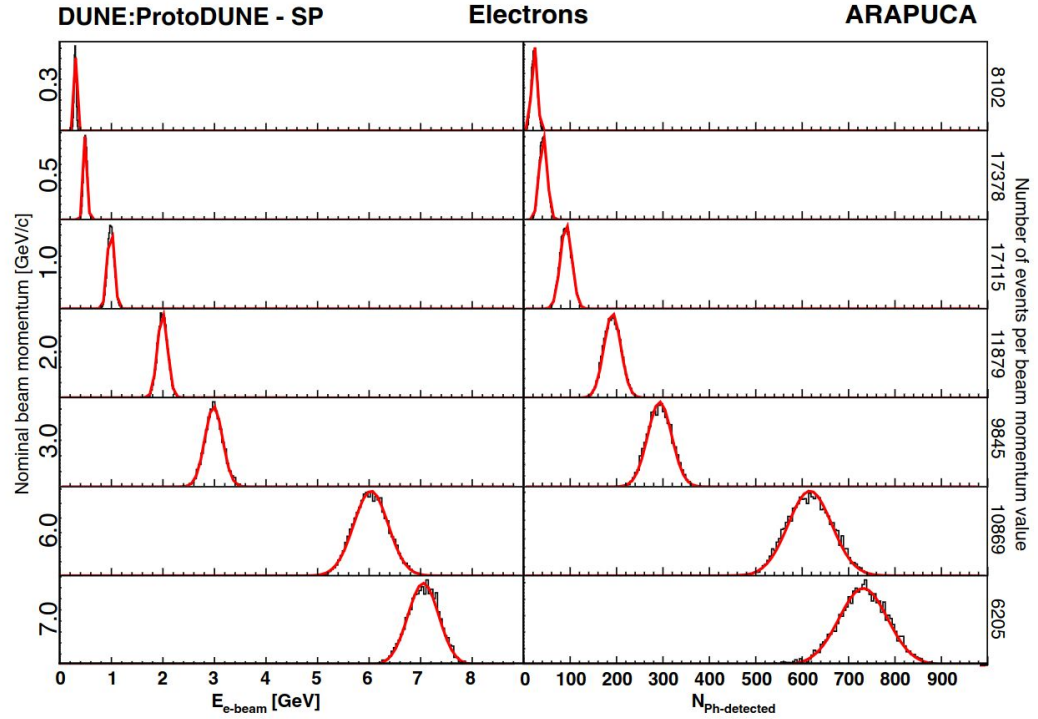


- Xenon doping



If there is still time...

- Interesting past analysis with ProtoDUNE-SP



Thanks :D

- Estimating number of photons:

Number of ionized atoms is proportional to the energy deposited (E_0) by the particle divided by the average energy expected per ion pair ($W_l = 23.6 \pm 0.3$ eV):

$$N_i = E_0 / W_l$$

Assuming that all ionized and excited molecules will produce photons, we have:

$$N_{ph} = N_i + N_{ex} = N_i \cdot (1 + N_{ex}/N_i) = E_0 / W_l \cdot (1 + N_{ex}/N_i)$$

And so:

$$N_{ph} = \frac{E_0}{W_{ph}^{\min}} \quad \text{with} \quad W_{ph}^{\min} = \frac{W_l}{1 + N_{ex}/N_i} = 19.5 \pm 1.0 \text{ eV}$$

So the maximum number of photons produced by MeV is simple $1\text{MeV} / 19.5 \sim 50 \times 10^3$ photons/MeV

If you consider 0.8 factor for muons and 0.6 factor for Electric field, wth have the usual 24×10^3 photons/MeV

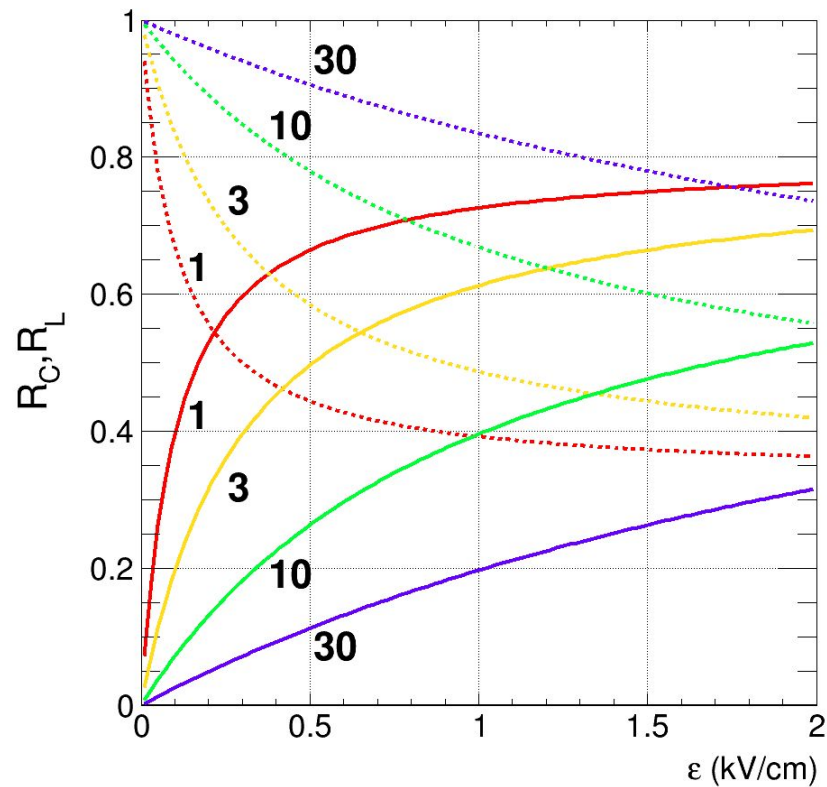
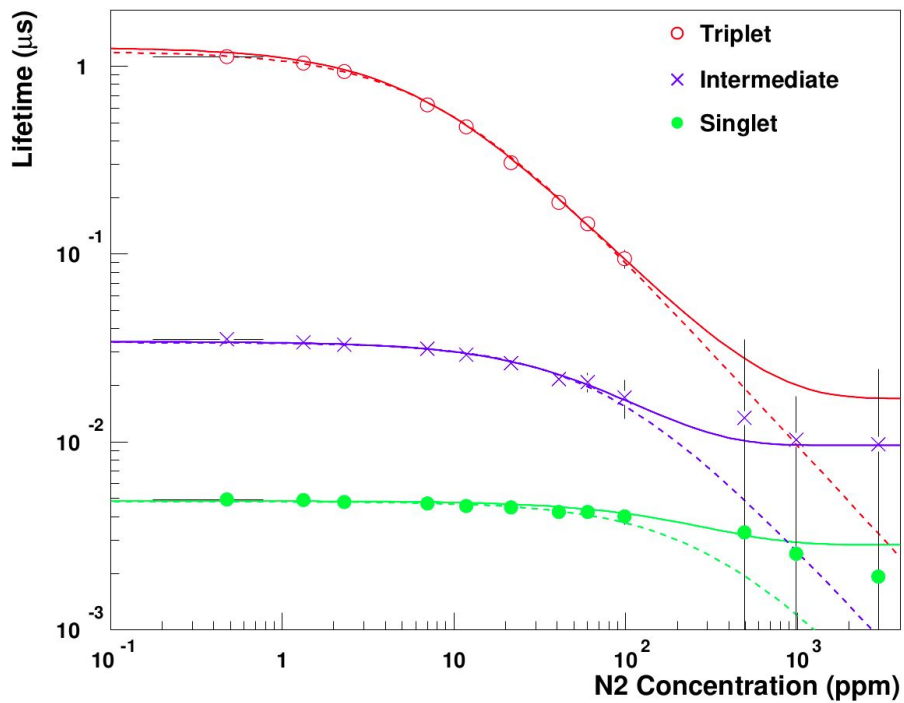
Backup

- Noble gas: electropositive and dielectric (low electron absorbance and high voltage allowed)
- High density
- High radiation length (allows good discrimination between electrons and photons and make it easier to retrieve neutrino vertex)
- Abundant in nature

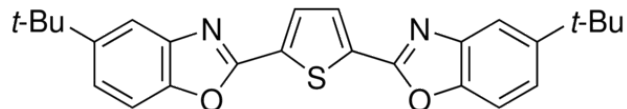
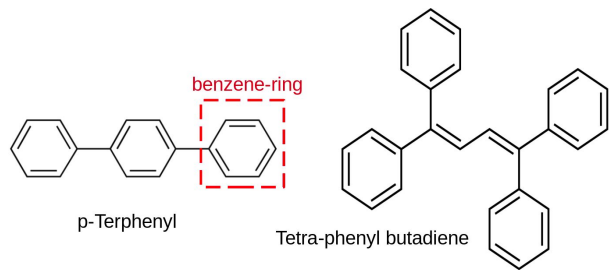
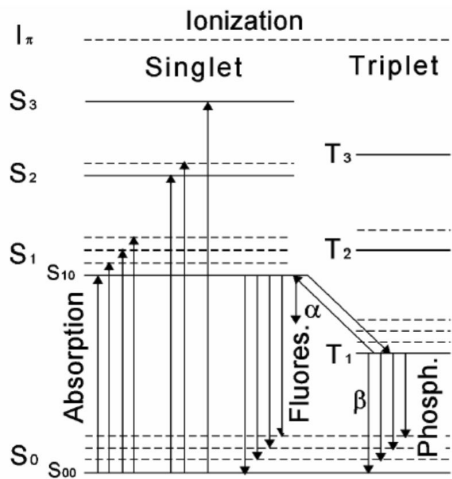
<https://arxiv.org/abs/2112.02967>

	Water	He	Ne	Ar	Kr	Xe
Boiling point [K] @ 1 atm	373	4.2	27.1	87.3	120	165
Density [g/cm ³]	1	0.125	1.2	1.4	2.4	3.0
Radiation length [cm]	36.1	755.2	24	14	4.9	2.8
Scintillation [γ /keV]	-	19	30	40	25	42
Scintillation λ [nm]	-	80	78	128	150	175
dE/dx [MeV/cm]	1.9	0.24	1.4	2.1	3.0	3.8
Abundance (Earth atm) [ppm]	25×10 ³	5.2	18.2	9300	1.1	0.09
Electron mobility [cm ² /V·s]	-	< 0.3	< 0.01	~500	~1800	~2200

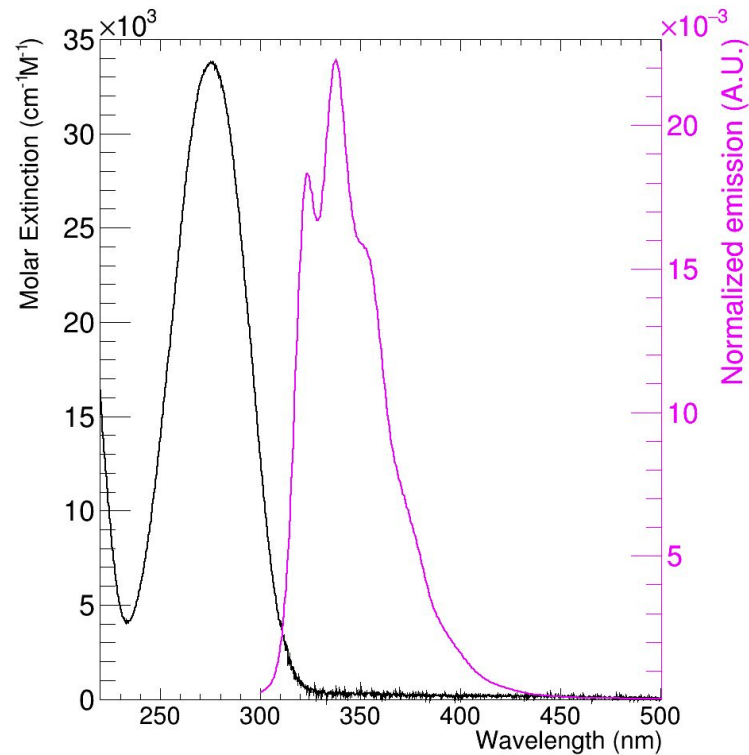
Mean energy loss (mip)	$\langle dE_{\text{mip}}/dx \rangle = 1.519 \text{ MeV}/(\text{g}/\text{cm}^2)$ ^[13]
Average energy for pair production (e^- , Ar^+)	$W_l = 23.6 \pm 0.3$ ^[45, 46]
Excited to ionized atoms ratio	$N_{\text{ex}}/N_i = 0.21$ ^[45, 48, 49]
γ emission spectrum	$\langle \lambda_{\text{scint}} \rangle = 127 \text{ nm}; \sigma_{\text{scint}} \approx 3 \text{ nm}$ ^[57]
Decay time constants	$\tau_S \sim 6 \text{ ns}; \tau_T \sim 1600 \text{ ns}$ ^[37, 57]
Relative intensity	$A_S/A_T = 0.3$ for electrons and muons $= 1.3$ for alpha particles $= 3.0$ for neutrons ^[37, 59, 60]
Average energy for γ production	$W_{ph}^{\text{min}} = 19.5 \pm 1.0 \text{ eV}$ ^[45, 48, 49]
Light Yield [$\epsilon = 0 \text{ V/cm}$] (ideal)	$Y_{ph}^{\text{ideal}} = 5.1 \times 10^4 \text{ } \gamma/\text{MeV}$
[$\epsilon = 0 \text{ V/cm}$] (mip)	$Y_{ph}^{\text{mip}} = 4.1 \times 10^4 \text{ } \gamma/\text{MeV}$
[$\epsilon = 500 \text{ V/cm}$] (mip)	$Y_{ph}^{\text{mip}} = 2.4 \times 10^4 \text{ } \gamma/\text{MeV}$ ^[5, 48]
Rayleigh scattering length ($\lambda_{\text{scint}} = 127 \text{ nm}$)	$99.1 \pm 2.3 \text{ cm}$ ^[64]
Absorption length (for N_2 concentration $< 5 \text{ ppm}$)	$L_A > 20 \text{ m}$ ^[65]
Refractive index	$n_{\text{LAr}} = 1.38$ ^[62]



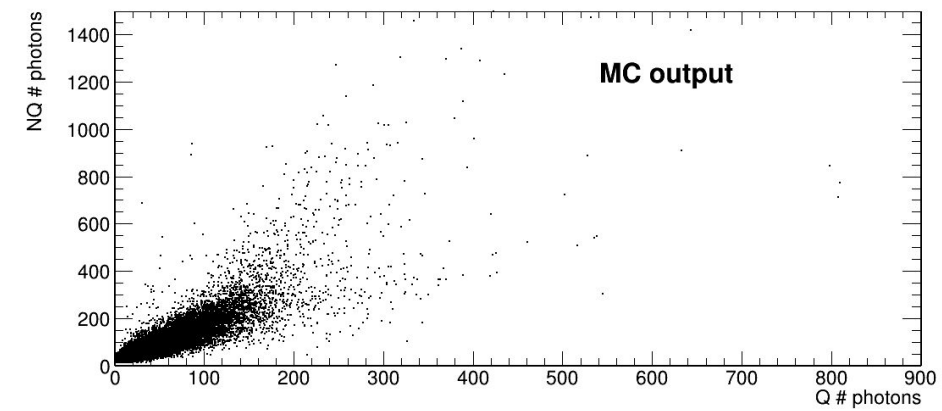
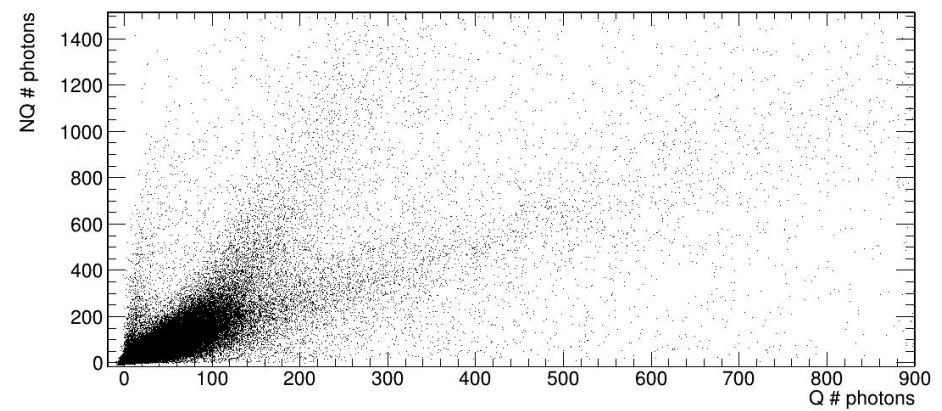
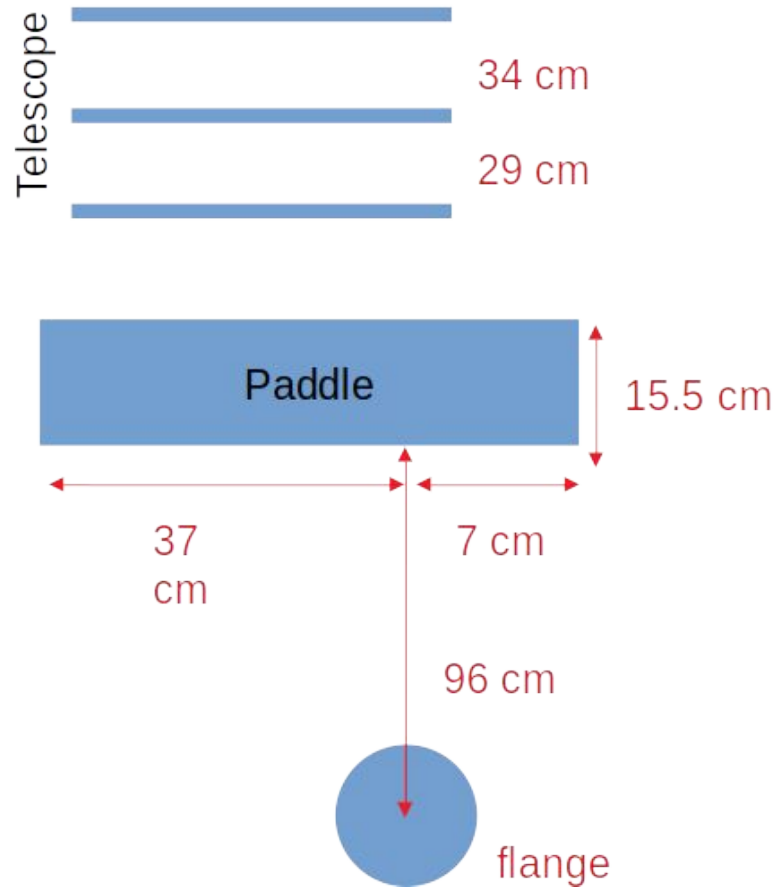
Backup

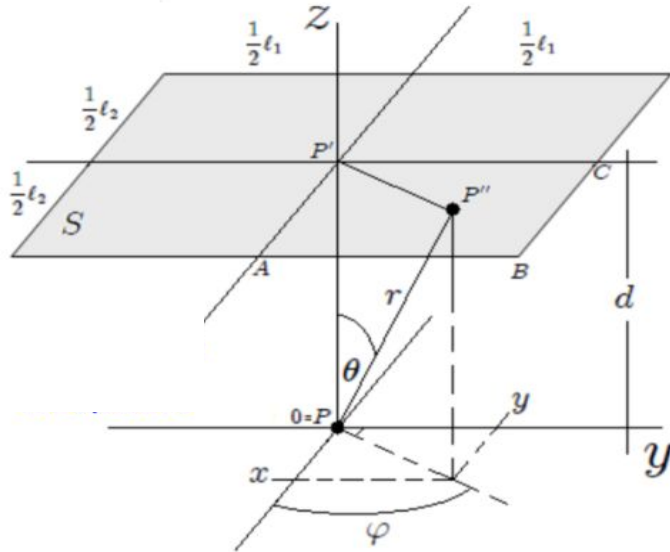


5-Bis(5-tert-butyl-benzoxazol-2-yl)thiophene (BBT).



Backup





Courtesy of Laura Paulucci

```
prod_muminus_0.1-5.0GeV_isotropic_dune10ktvd_1x8x14_gen_g4_detsim_reco.root
├── RootFileDB;1
├── Metadata;1
├── FileIndex;1
├── Parentage;1
├── EventHistory;1
└── Events;1
    ├── EventAuxiliary
    ├── art::TriggerResults_TriggerResults_Reco.
    ├── simb::MCTruths_generator_SinglesGen.
    ├── recob::OpHits_ophit10ppm_Reco.
    ├── sim::SimChannels_tpcrawdecoder_simpleSC_detsim.
    ├── sim::OpDetBacktrackerRecords_PDFastSimArExternal_G4.
    ├── sim::SimEnergyDeposits_IonAndScintExternal_G4.
    ├── sim::OpDetBacktrackerRecords_PDFastSimXeExternal_G4.
    ├── raw::RawDigits_tpcrawdecoder_daq_detsim.
    ├── sim::SimChannels_elecDrift_G4.
    ├── art::RNGsnapshots_ms_detsim.
    ├── recob::Wires_wclsdatanfsp_wiener_Reco.
    ├── sim::SimPhotonsLites_PDFastSimAr_G4.
    ├── art::RNGsnapshots_ms_G4.
    ├── raw::OpDetWaveforms_opdigi10ppm_detsim.
    ├── art::RNGsnapshots_ms_SinglesGen.
    ├── recob::OpFlashes_opflash10ppm_Reco.
    ├── sim::SimEnergyDeposits_IonAndScint_G4.
    ├── recob::OpFlashrecob::OpHitvoidart::Assns_opflash10ppm_Reco.
    ├── art::TriggerResults_TriggerResults_G4.
    ├── sim::OpDetBacktrackerRecords_PDFastSimAr_G4.
    ├── sim::SimEnergyDeposits_largeant_LARg4DetectorServicevolExternalActive_G4.
    ├── sim::OpDetDivRecs_sipmAr10ppm_detsim.
    ├── simb::MCParticles_largeant_G4.
    ├── sim::SimEnergyDeposits_largeant_LARg4DetectorServicevolITPCActive_G4.
    ├── sim::OpDetBacktrackerRecords_PDFastSimXe_G4.
    ├── sim::OpDetDivRecs_sipmXe10ppm_detsim.
    ├── sim::SimPhotonsLites_PDFastSimXeExternal_G4.
    ├── recob::Hits_gaushit_Reco.
    ├── sim::SimPhotonsLites_PDFastSimArExternal_G4.
    ├── sim::SimPhotonsLites_PDFastSimXe_G4.
```

Backup

