

JEMU

A BRIEF INTRODUCTION TO ANOTHER SIMPLE & NON-OFFICIAL **P(K,Z) CLASS EMULATOR** DESIGNED IN THE CONTEXT OF **JAX-COSMO** LIB.

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JAX-COSMO LIBRARY*

Cosmological library...

```
! pip install -q jax_cosmo
import jax_cosmo as jc
```

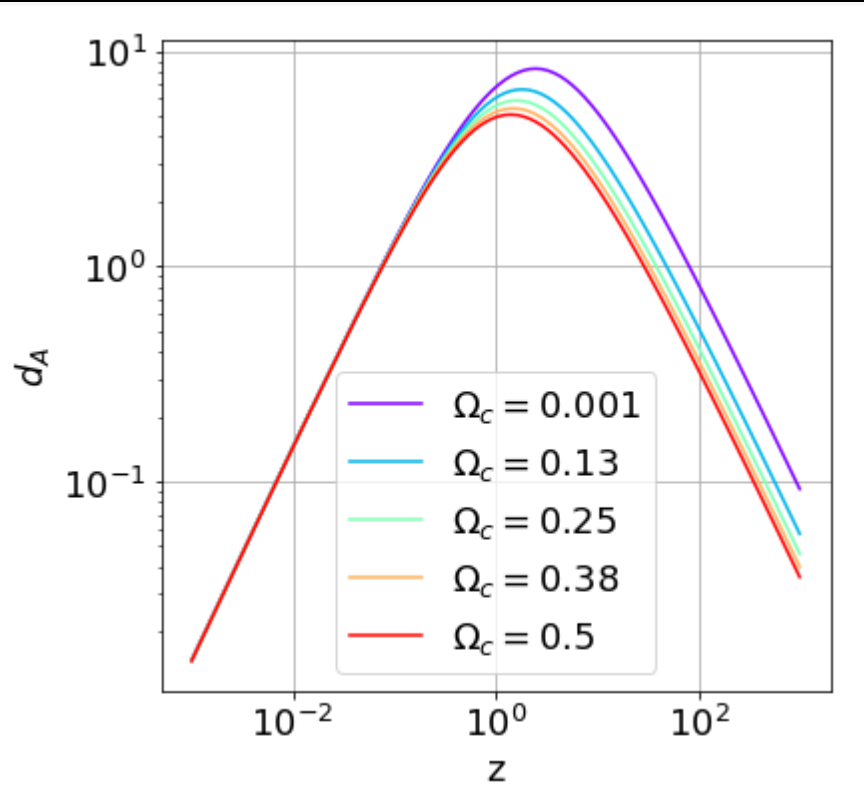
```
# Cosmologie "Standard"
cosmo_jax = jc.Planck15()
```

Cosmology as PyTree structure

Parameters:

- Omega_c, float
Cold dark matter density fraction.
- Omega_b, float
Baryonic matter density fraction.
- h, float
Hubble constant divided by 100 km/s/Mpc; unitless.
- n_s, float
Primordial scalar perturbation spectral index.
- sigma8, float
Variance of matter density perturbations at an 8 Mpc/h scale
- Omega_k, float
Curvature density fraction.
- w0, float
First order term of dark energy equation
- wa, float
Second order term of dark energy equation of state
- gamma: float
Index of the growth rate (optional)

- ✓ automatic differentiation
- ✓ Just in time compilation
- ✓ Vectorization on single & multi devices
- ✓ Same code for CPU & GPU/TPU

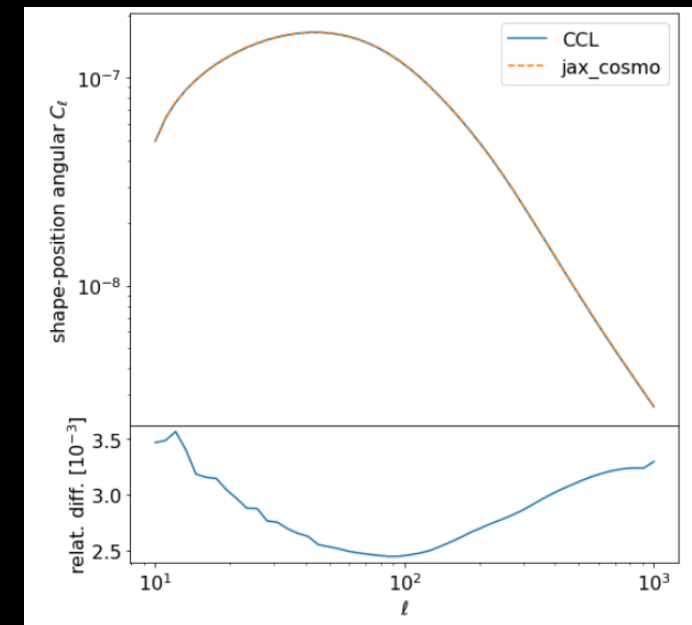
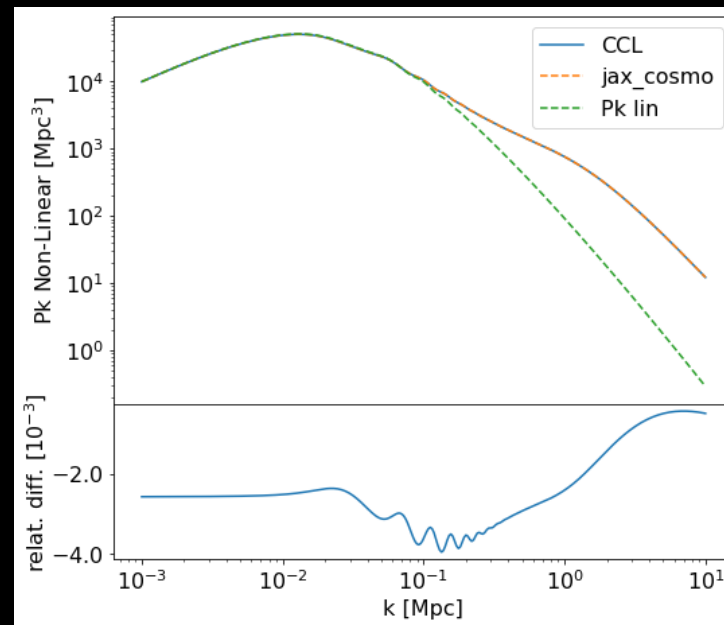


*) JEC, F. Lanusse, J. Zuntz et al [arxiv:2302.05163](https://arxiv.org/abs/2302.05163) (2023), OJAp, 6, 15

JAX-COSMO LIBRARY

Currently: light implementation

- Follow **CCL**'s tracer mechanism
- Unit tested against CCL
- **P(k)** (Eisenstein-Hu, Halofit),
- **Cls** (Lensing including NLA, multiplicative bias, number counts)
- Limber approximation



Idea: change the present $P(k)$ by an emulator

$$\tilde{D}(k, z) := \frac{P_{lin}(k, z)}{P_{lin}(k, 0)}$$

$$\tilde{Q}(k, z) := \frac{P_{nl}(k, z)}{P_{nl}(k, 0)}$$

Different from original schema

$$y_{ij}^X$$

$$X = \{P_{lin}(k, z = 0), P_{nl}(k, z = 0), \tilde{D}(k, z), \tilde{Q}(k, z)\}$$

RBF

Matern12

(k_i, z_j) log-lin grid of 120x20 pts
One GP with polynomial mean per (i,j)

θ : Cosmo param. 10^3 Training on a Latin Hypercube

Model hyper-param

$$y(\theta) = \Phi(\theta)\beta + \mathcal{GP}(0, K(a, \{\ell\}))$$

$$\alpha_{ij}^X = (\beta, a, \{\ell\})_{ij}^X$$

$$\Phi(\theta) = \{1, \theta, \theta^2\}$$

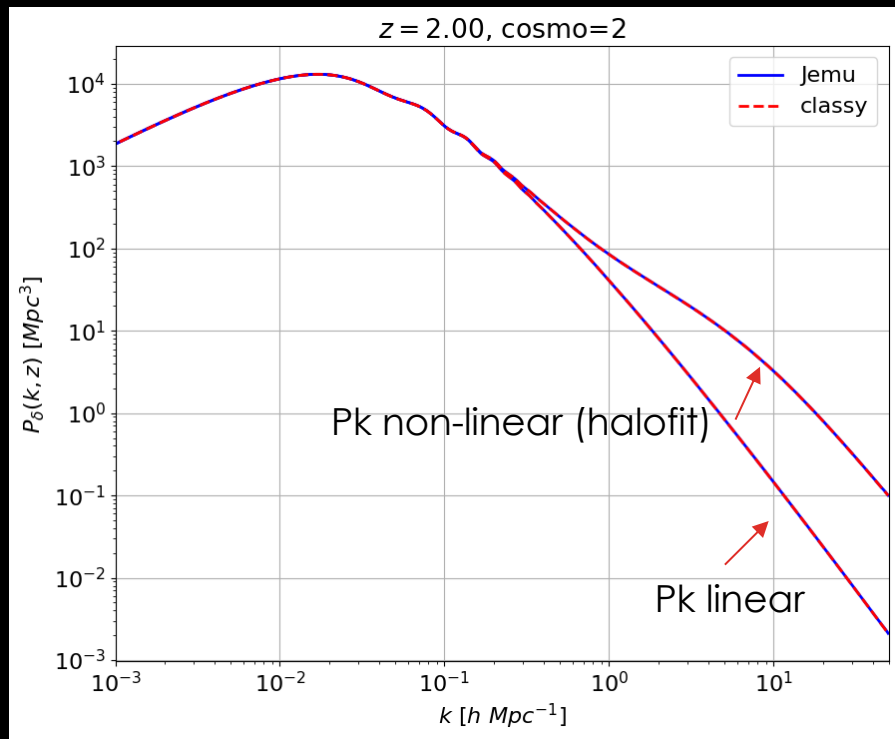
Max. of marginal likelihood with β gaussian prior (Scipy/L-BFGS-B)



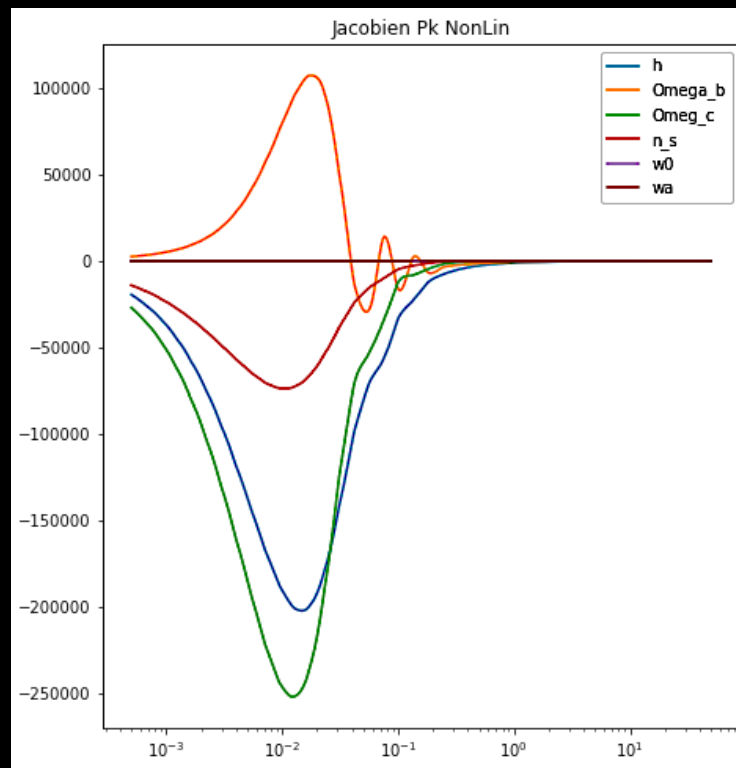
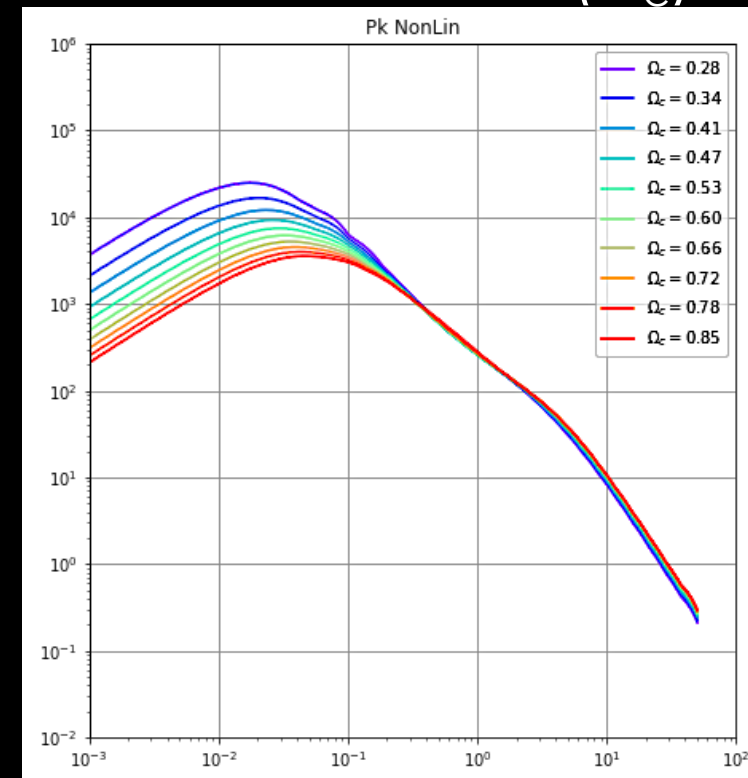
$$y_{ij}^X(\theta_*; \hat{\alpha}_{ij}^X)$$

Interpolation from the ij-grid

Autodiff & vectorization Interface with jax-cosmo Cosmology



Jacobian

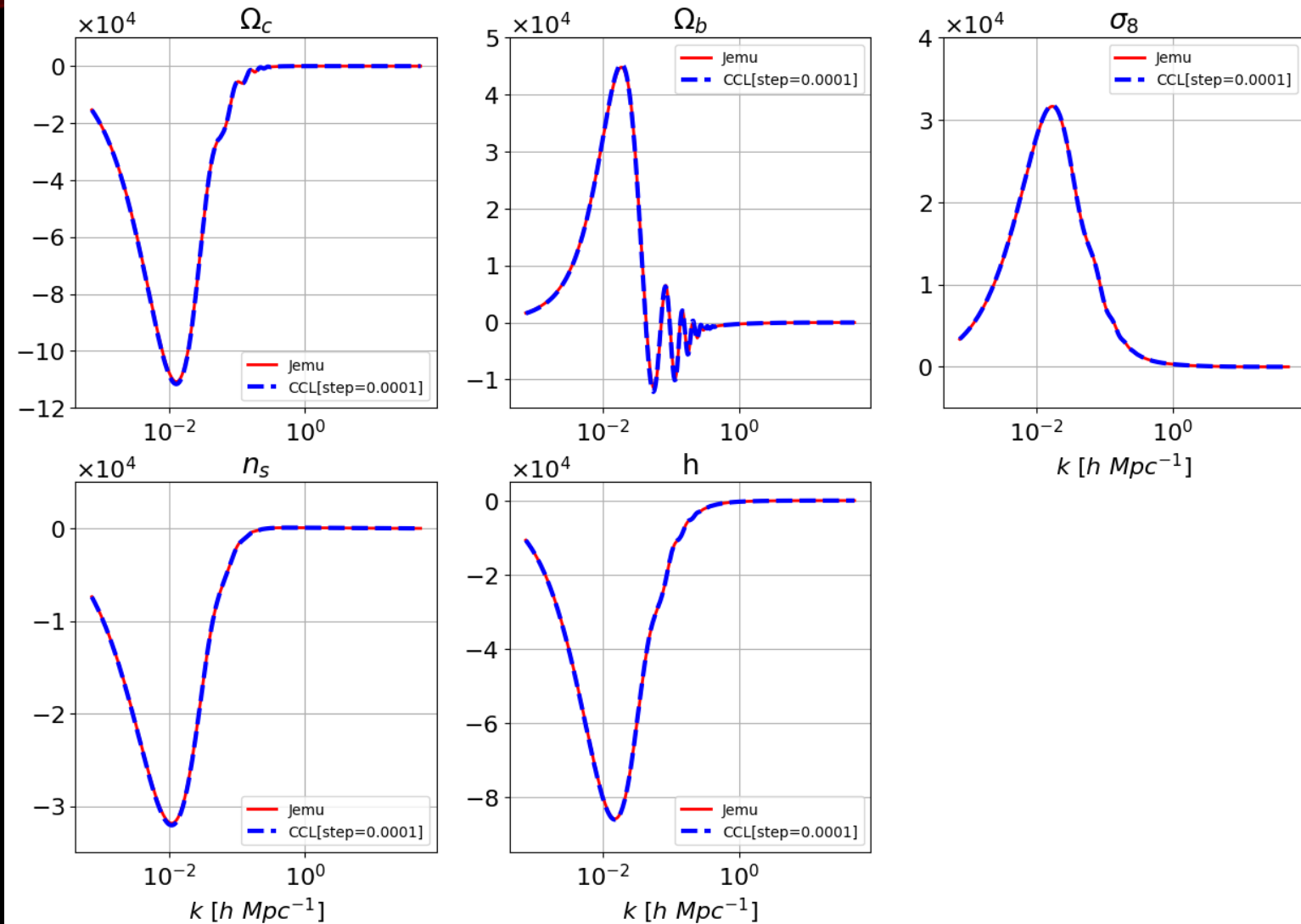
Vectorization (Ω_c)

@ $O(1\%)$ level the differences are not visible

CCL vs Jax-Cosmo with Jemu

$h = 0.7$
 $\Omega_c = 0.4$
 $\Omega_b = 0.045$
 $\sigma_8 = 0.75$
 $n_s = 0.95$

Jacobian of Pk NonLin (z=2.00, cosmo=2)



SUMMARY

✓ Cosmo Parameters

- $\Omega_c h^2$: [0.06, 0.40]
- $\Omega_b h^2$: [0.019, 0.026]
- σ_8 : [0.25, 3.25]
- n_s : [0.7, 1.3]
- h : [0.64, 0.82]

Here the settings used for test but it is user-case dependant
3 neutrino, $M_{tot}=0.06\text{eV}$, $T_{\text{ncdm}}: 0.71611$, $N_{\text{ur}}: 0.00641$

✓ redshift in $[0, \lesssim 3]$, k in $[5 \cdot 10^{-4}, \lesssim 20]$ h/Mpc

✓ $|\Delta P_k/P_k| \approx 1\%$ (nb. More the above ranges are squeezed more the accuracy gets better)

✓ Coded in JAX with interface with `jax-cosmo`* (see [CosmoPower-JAX](#)** for similar aspects and use)

See the project for a JAX standardisation of CLASS emulator interface

[Zelos @ Differentiable Universe Initiative](#)

*: [Jemu Github](#)

** : D. Piras & A. Spurio Mancini [arXiv:2305.06347](#)