

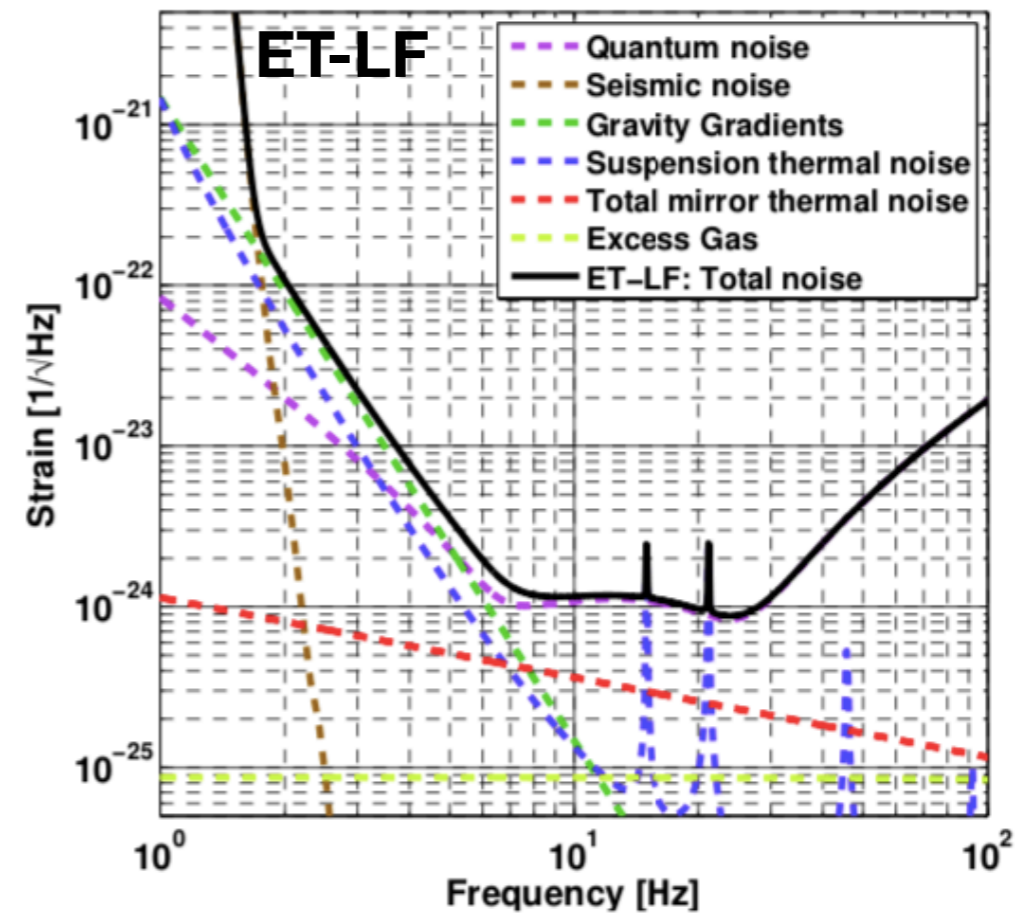
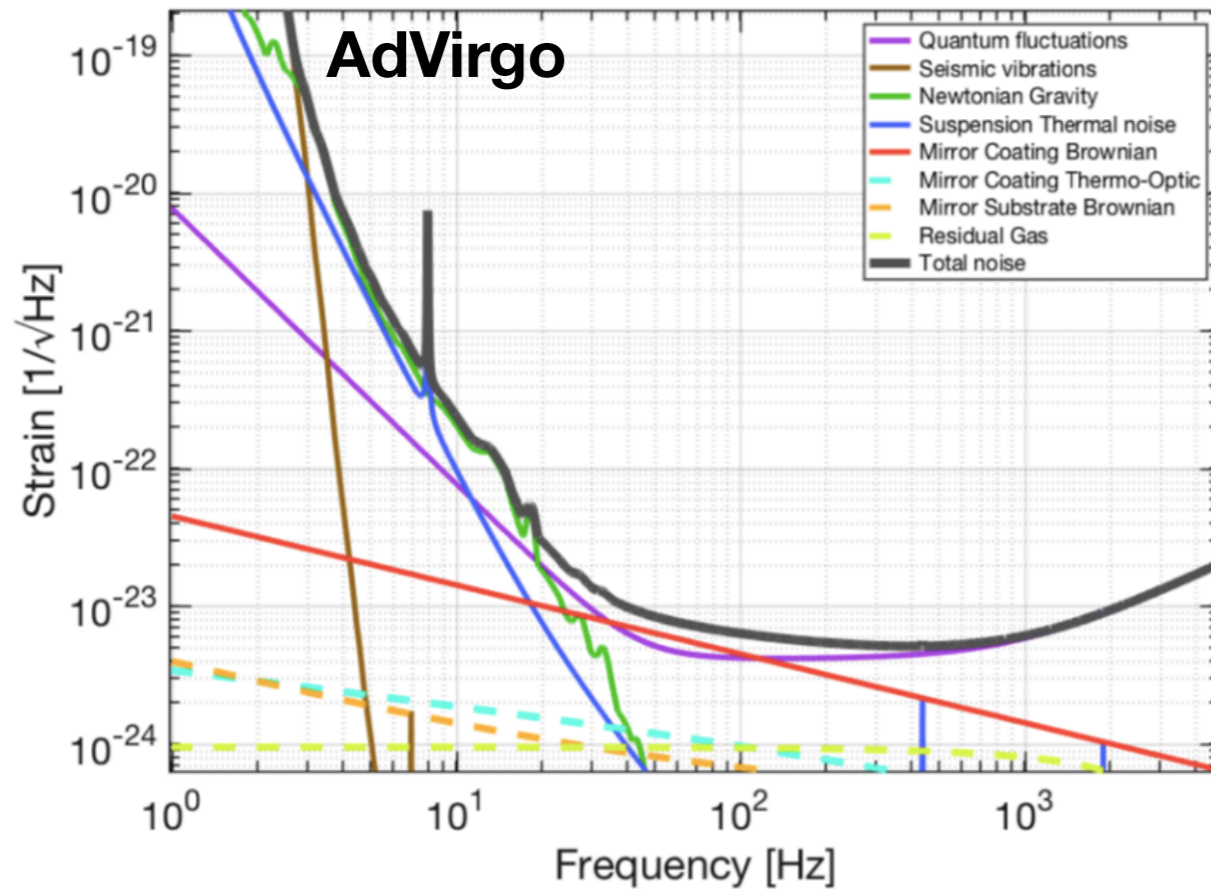


Squeezing for future GW detectors

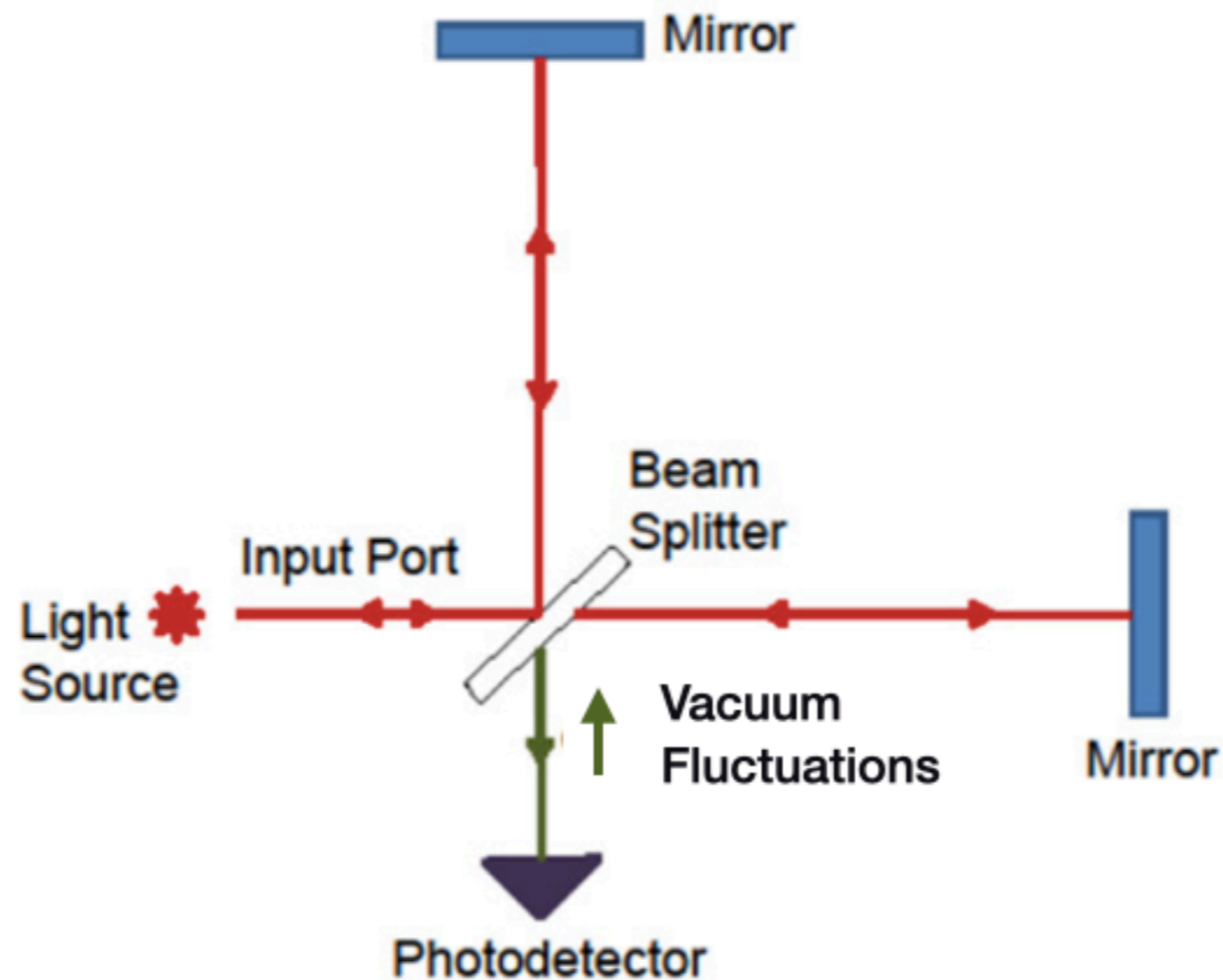
Eleonora Capocasa

Context: Quantum noise in GW detectors

- Main limiting noise of current and future GW detectors

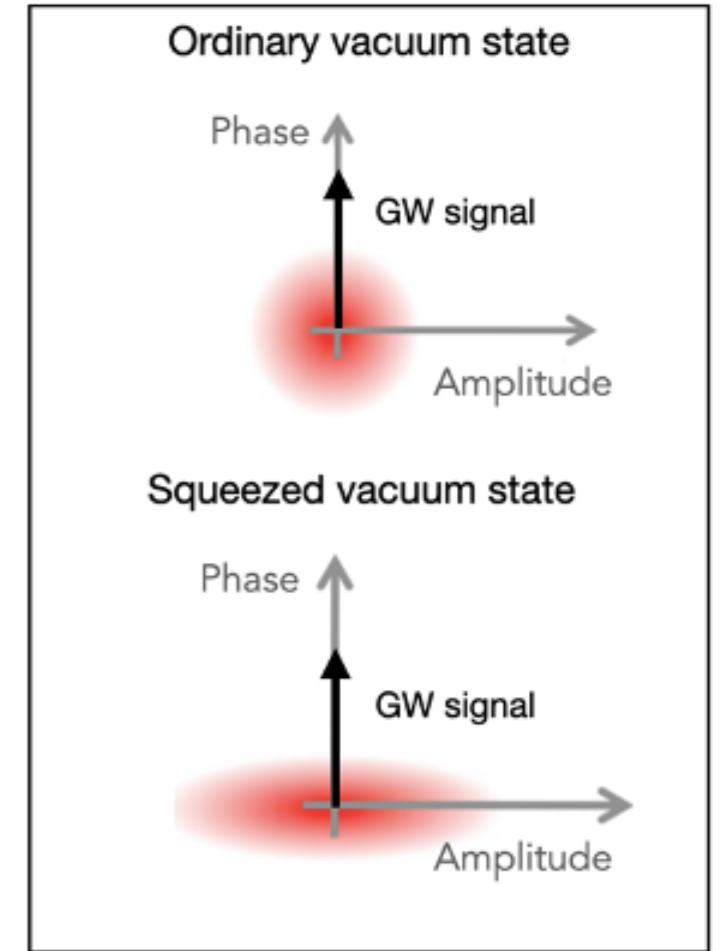
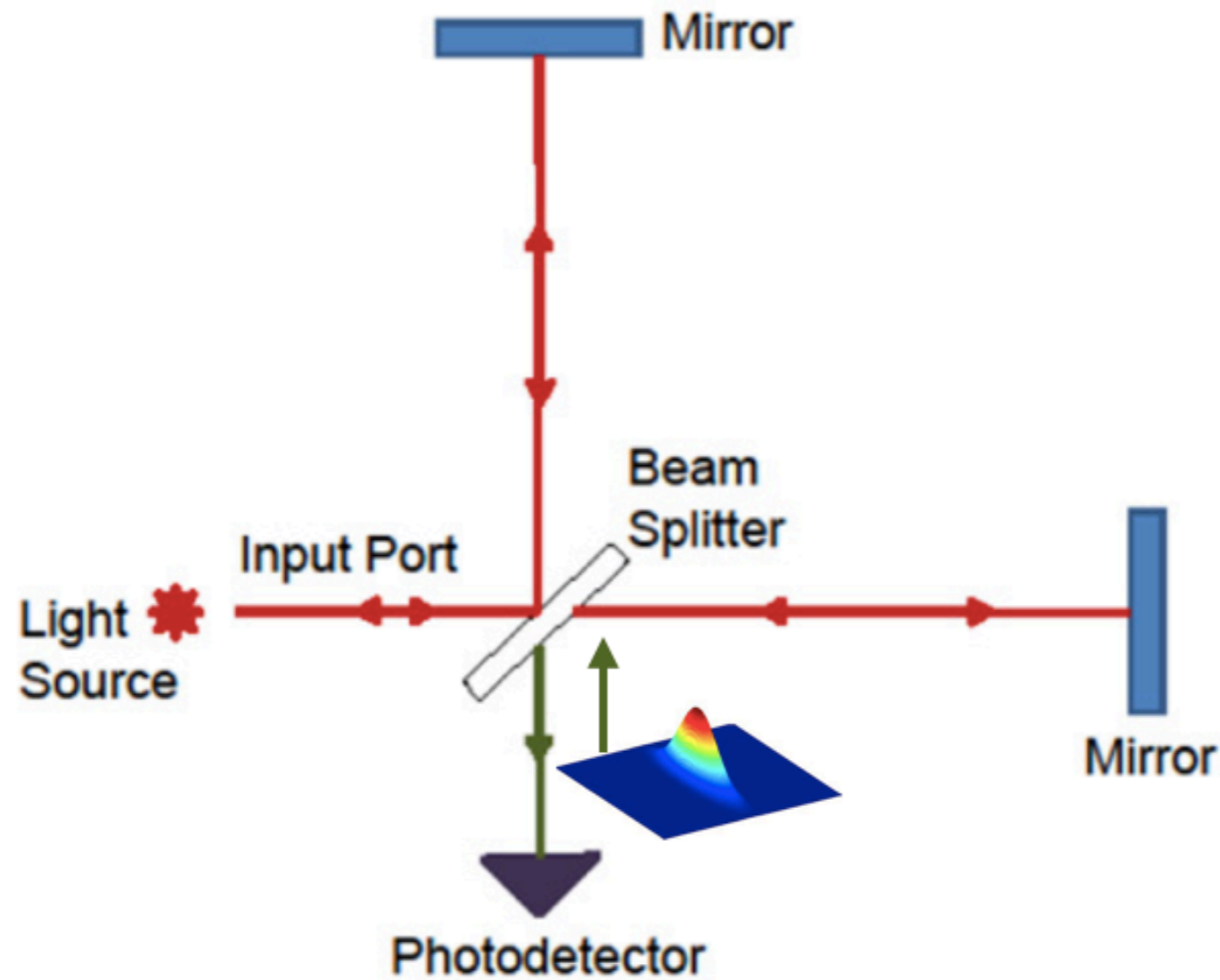


Quantum noise in GW interferometers



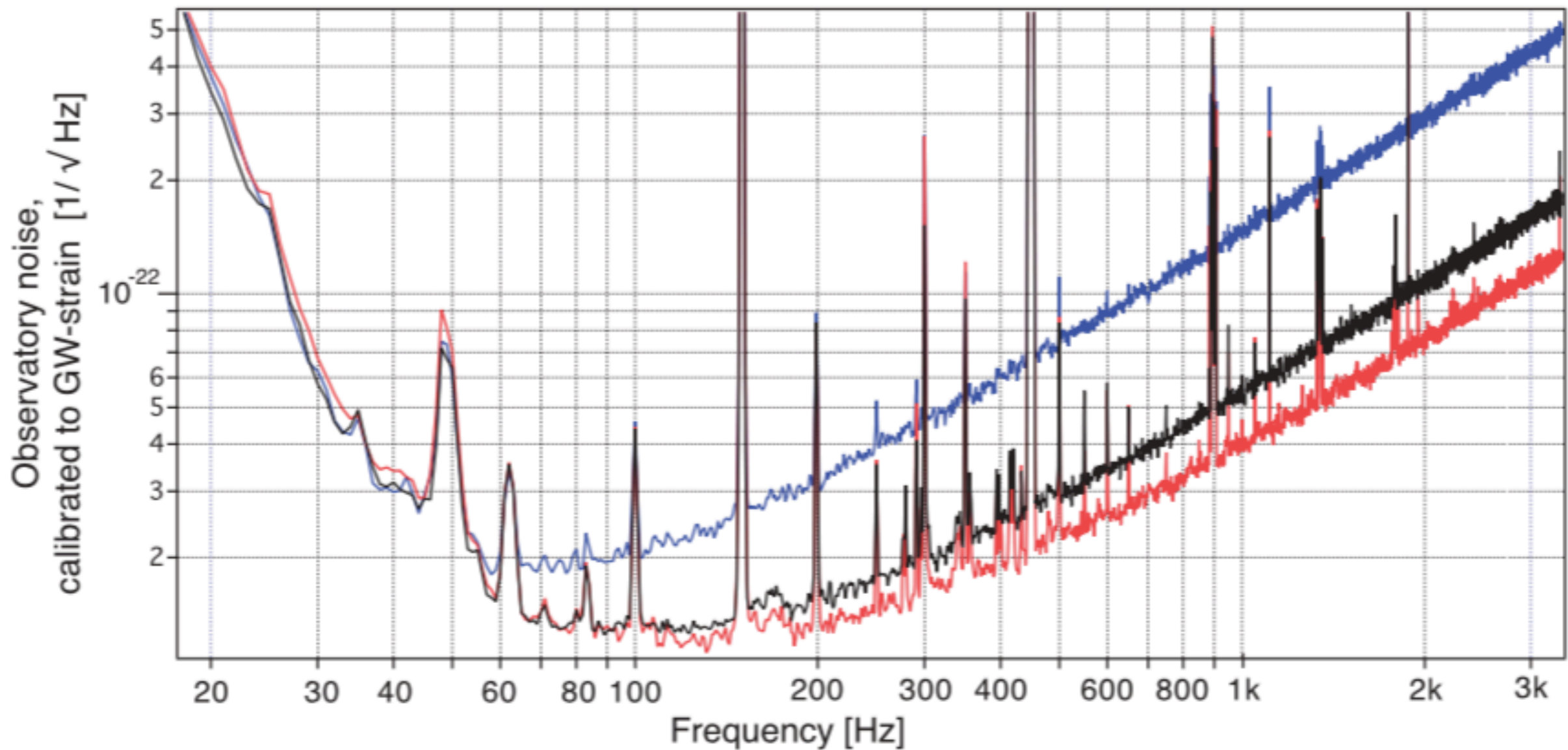
¹C.Caves "Quantum-mechanical noise in an interferometer"
Phys. Rev. D 23 (1981)

Quantum noise reduction using squeezed light



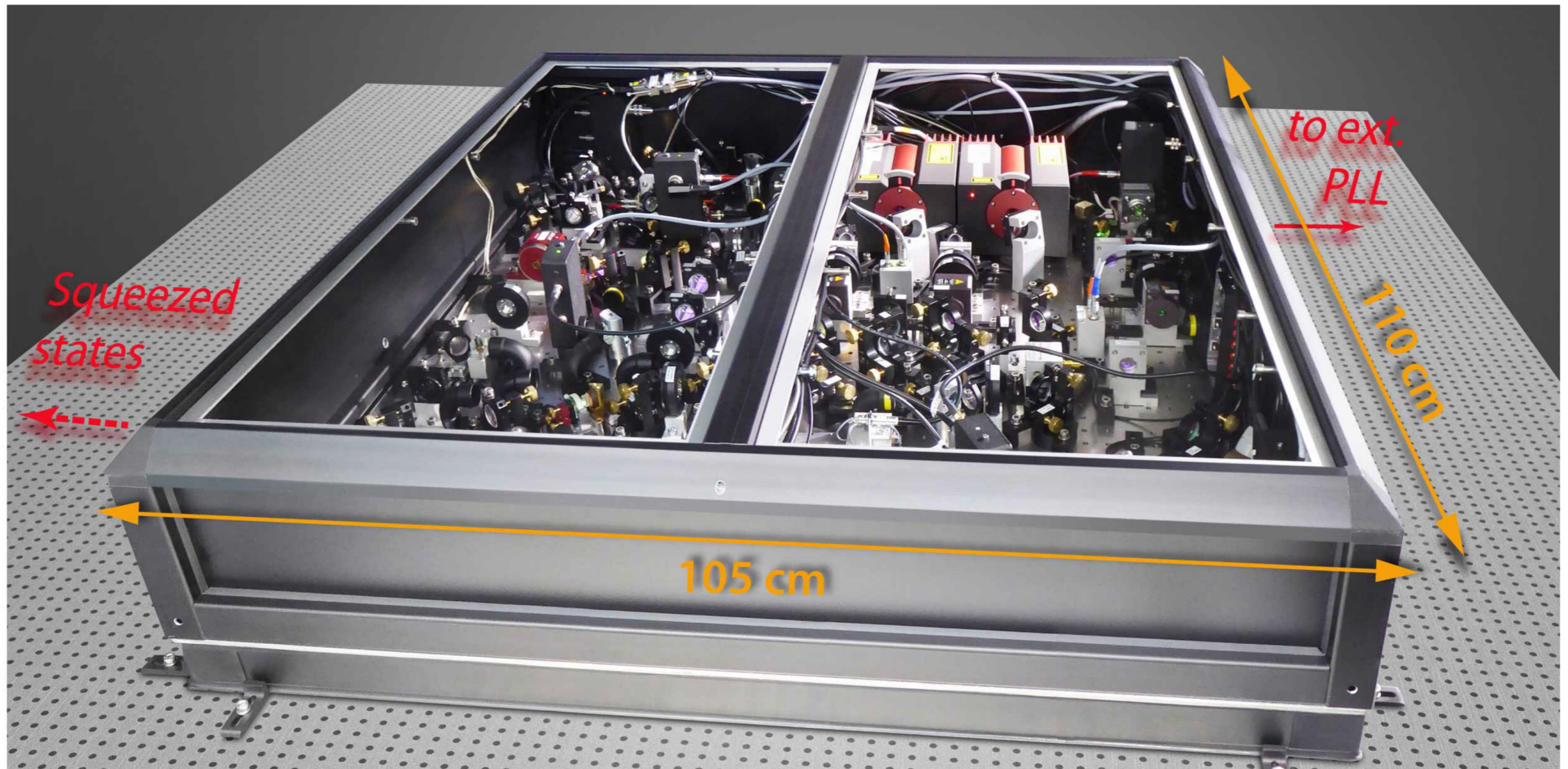
Squeezed source integrated in GW detectors

- Operating in both LIGO and Virgo since the beginning of O3

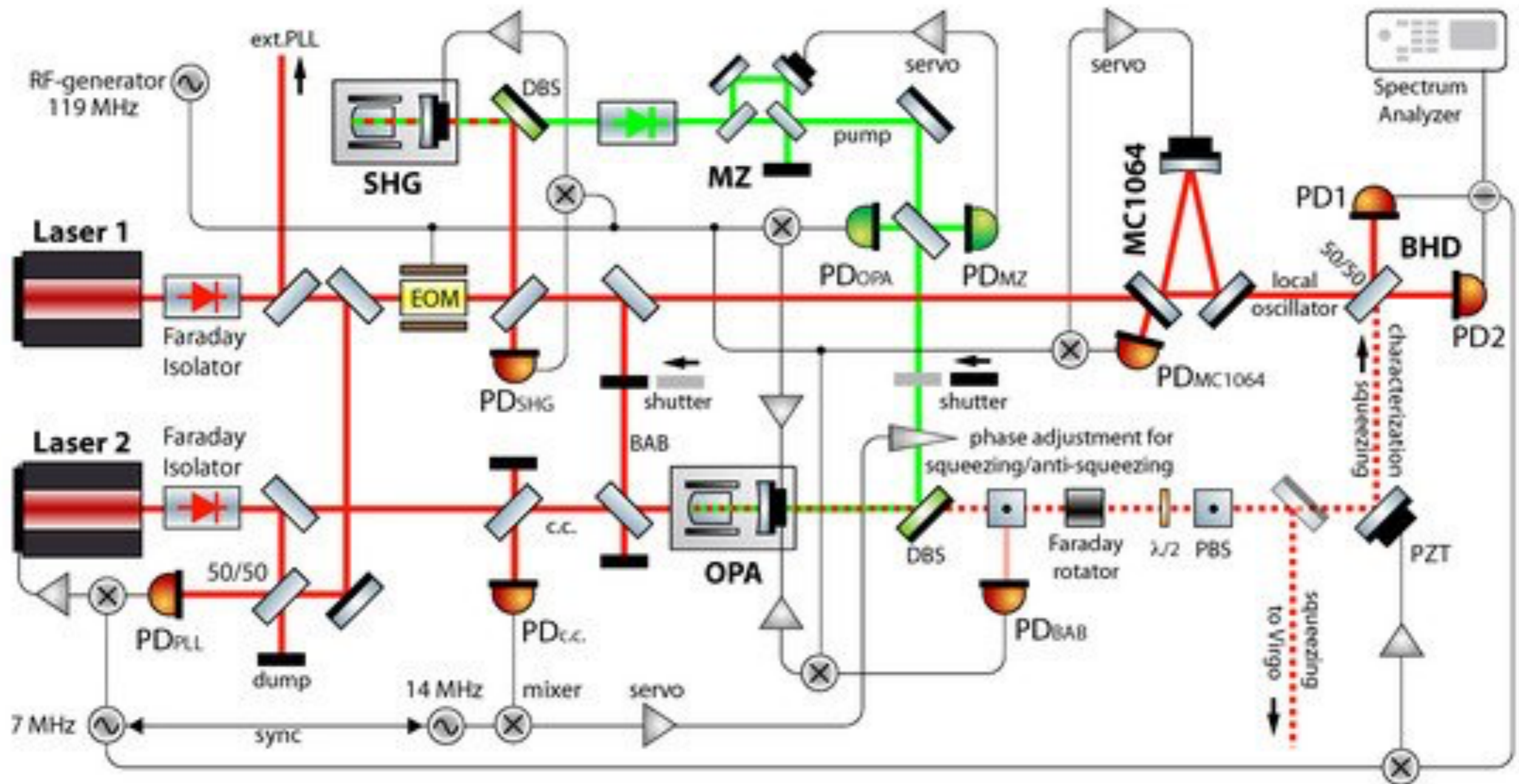


- About 3 dB of squeezing measured
- Between 25-40% of losses measured

Vacuum squeezed source

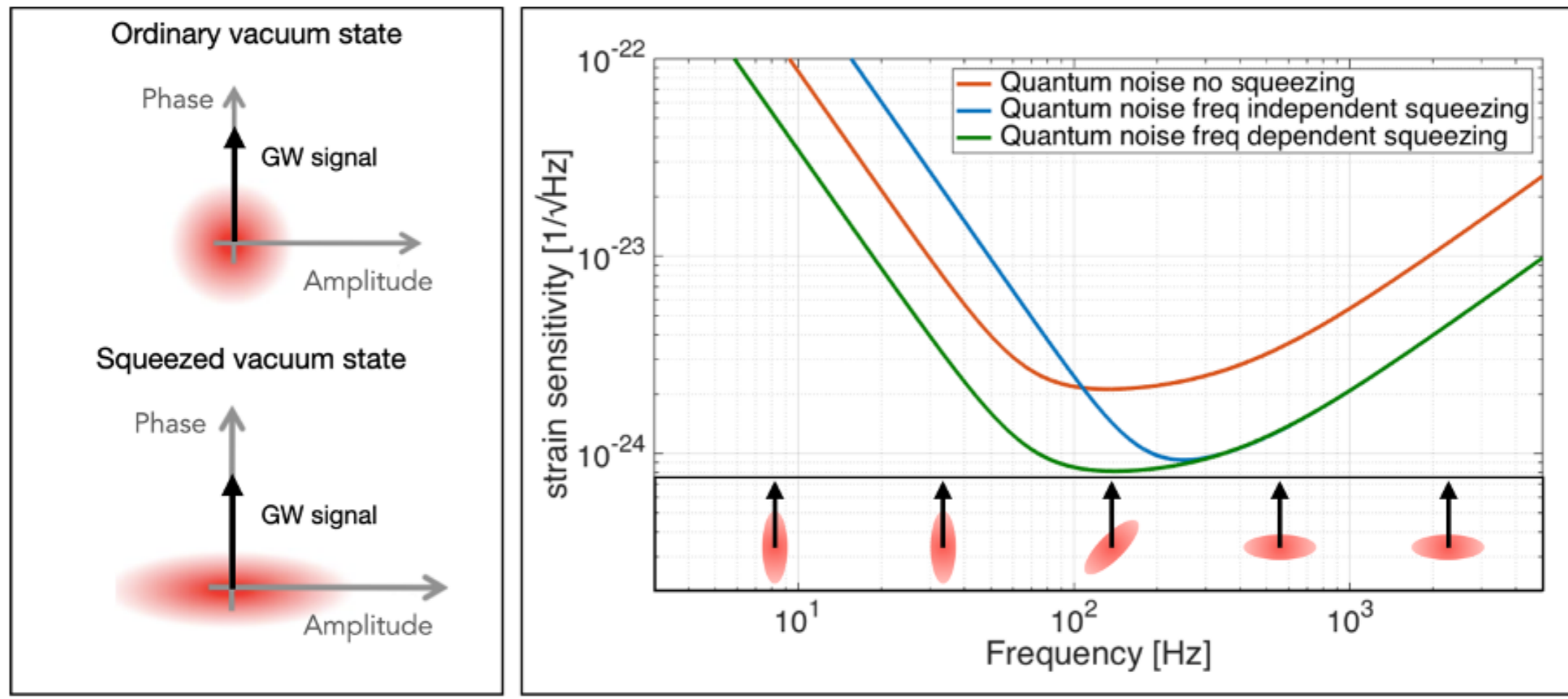


Vacuum squeezed source



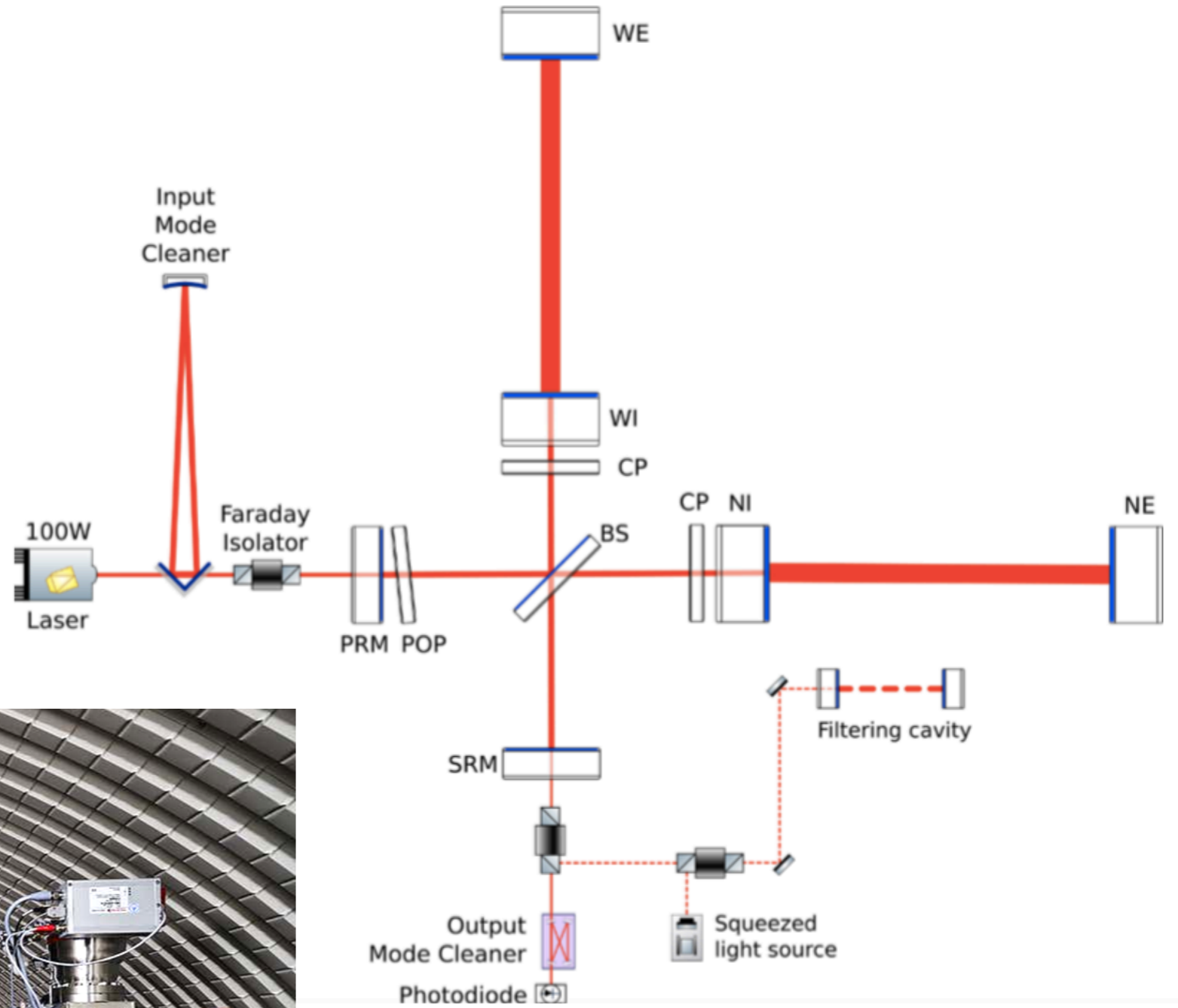
Broadband quantum noise reduction

- Squeezing ellipse undergoes a rotation inside the interferometer
- Squeezing angle should change with the frequency for optimal noise reduction



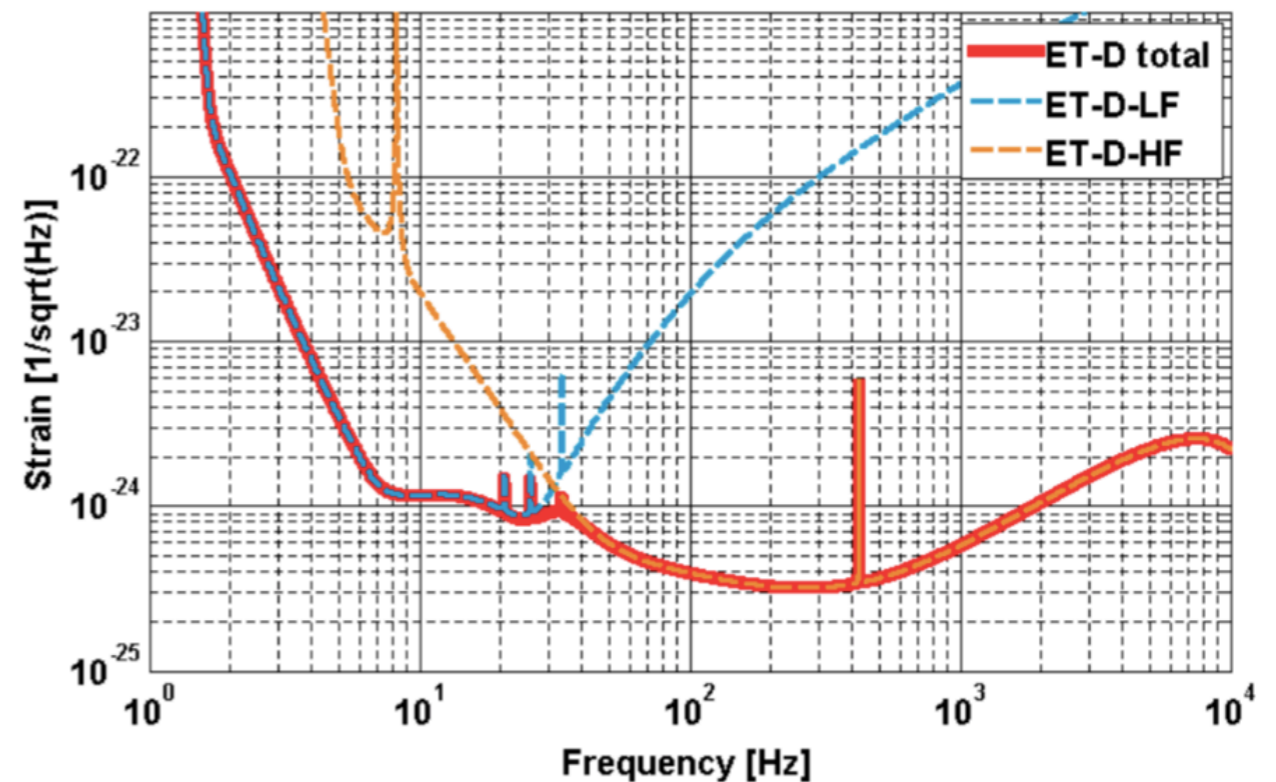
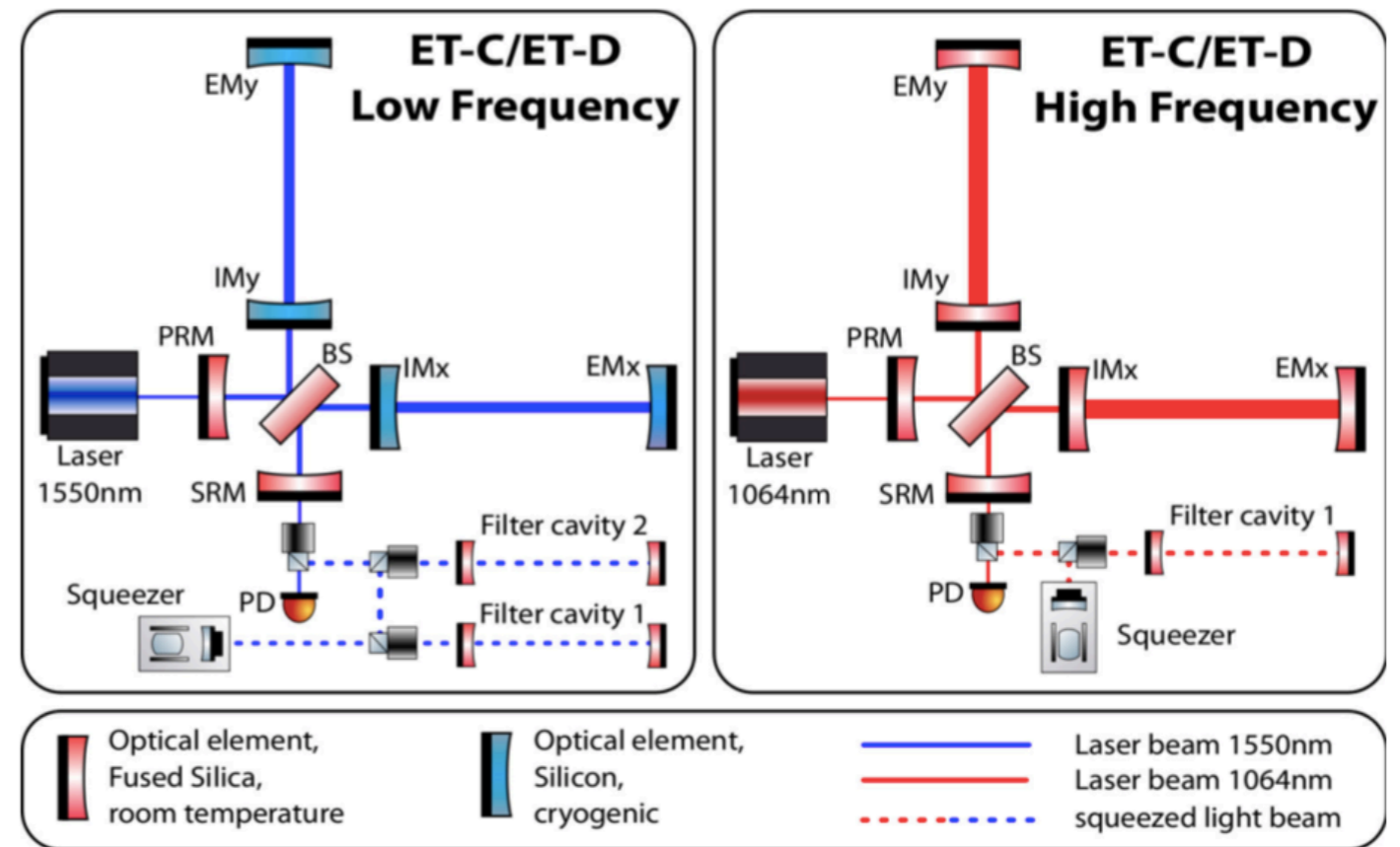
Filter cavity implementation for O4

- Length: ~300 m



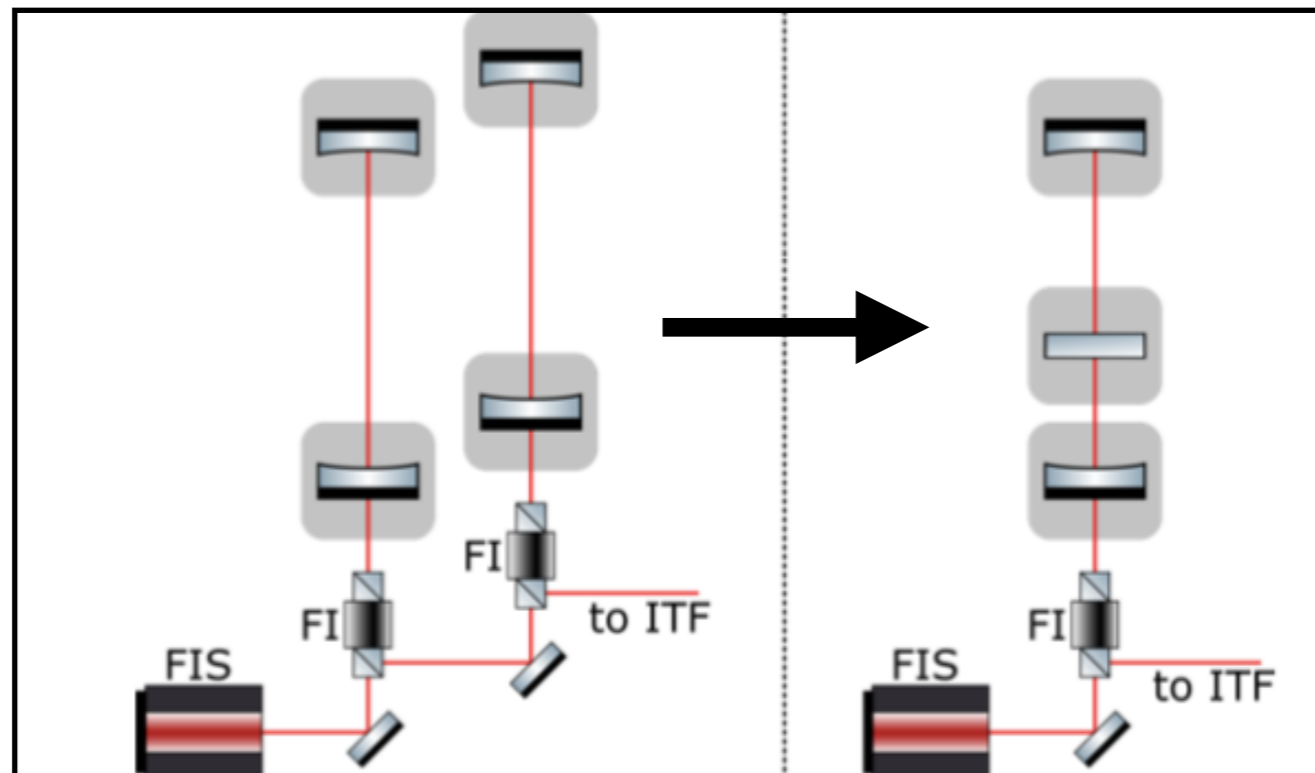
Squeezing for Einstein Telescope

- **Goal:** 10 dB of broadband quantum noise reduction
- 2 filter cavity fo ET-LF



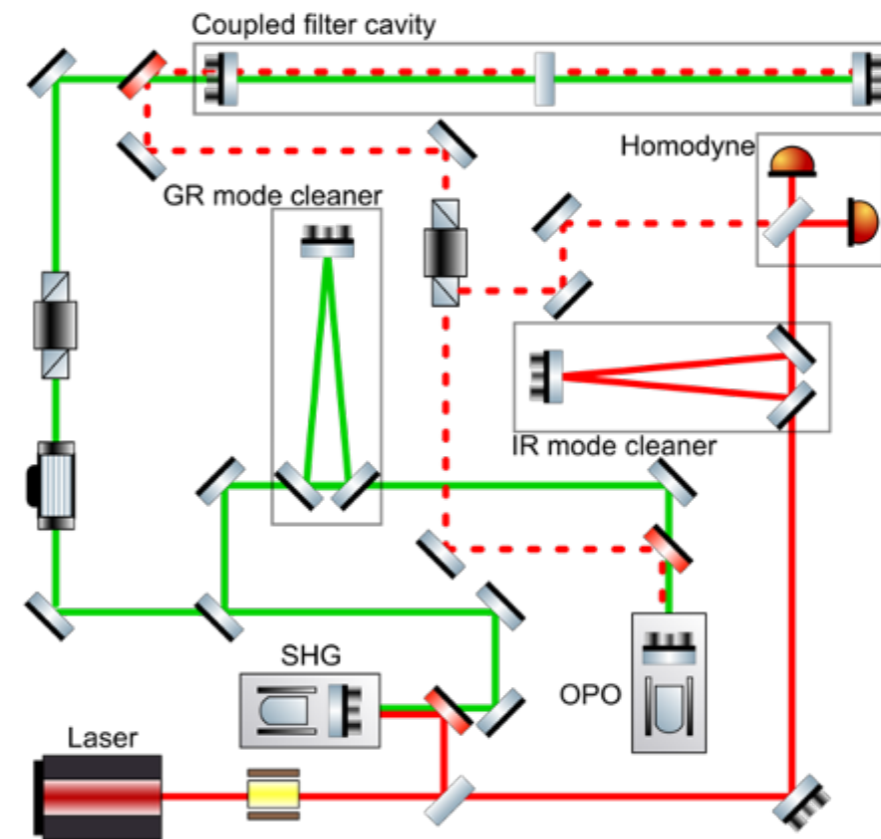
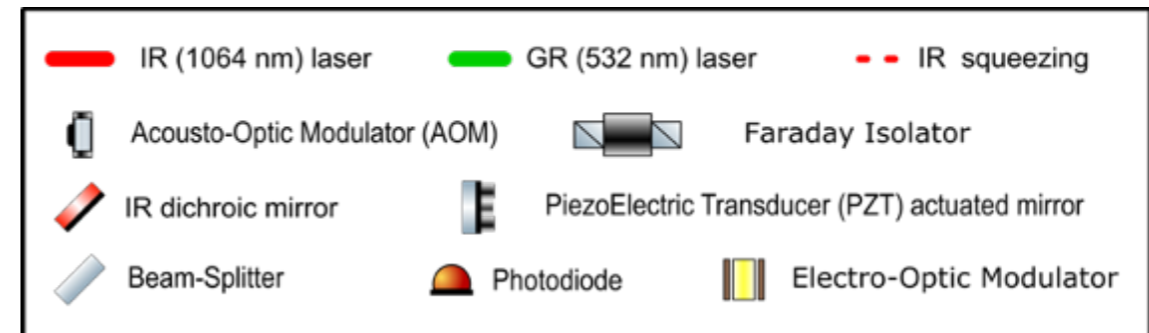
Project: frequency dependent squeezing for ET

- Design of optimal filter cavities
- Possible replacement of two filter cavity with a double filter cavity -> simulation and table top experiment



APC table top experiment milestones

- Experiment design
- Squeezing source realisation
- Double cavity control and characterisation
- Demonstration of FDS with non-trivial rotation



Team

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