

Charting the landscape of non-supersymmetric strings

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Deconstructing the String Landscape

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Plan

- Introduction
 - Equations of motion without susy
 - Non-susy tachyon-free string theories in 10D

- Flux compactifications
 - Sum rules
 - Freund-Rubin vacua

- A first-order formalism

Introduction

Context: find geometric string vacua without spacetime susy that are (in some appropriate sense) under control.

However, ~~SUSY~~ :



Limited tools.



Limited control (quantum corrections).



Generic instability.

How do we **chart the non-susy landscape?**

Equations of motion without susy

Double expansion in g_s and α'

$$S \sim \sum_{n,m=0}^{\infty} \int e^{(n-2)\phi} (\alpha')^{m-4} \mathcal{O}_{2+2m} \cdot$$

Susy:

- ✓ Protection of terms in the action \rightarrow 2-derivative sugra.
- ✓ First-order equations.
- ✓ Use of spinors: access to more structure.
- ✓ Dynamical obstruction to decays.

What we expect from string theory is a 2-derivative action

$$S \sim \int (e^{-2\phi} + c_R)R + (e^{-2\phi} + c_\phi)4(\partial\phi)^2 - (e^{-2\phi} + c_H)\frac{1}{2}\frac{H^2}{3!} - \Lambda + \dots$$

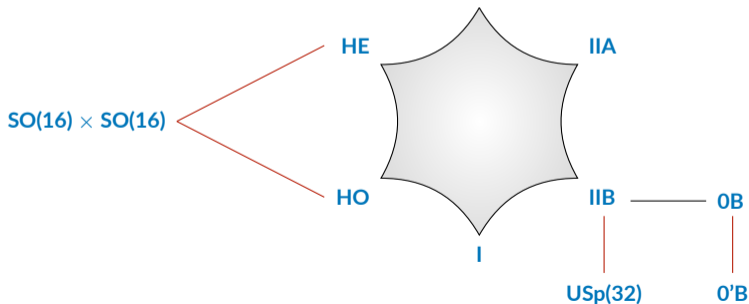
Some comments:

- ▣ Fine as a perturbative expansion, except the new term Λ .
- ▣ Is there a low-energy principle explaining these types of corrections?

Can we compute Λ from string theory? \rightarrow non-susy strings.

Non-susy tachyon-free string theories in 10D

- ① Heterotic: $SO(16) \times SO(16)$ [Alvarez-Gaume, Ginsparg, Moore, Vafa 1986; Dixon, Harvey 1986].
- ② Type IIB with $O9^+$ and 32 $\overline{D9}$: $USp(32)$ [Sugimoto 1999].
- ③ Orientifold of bosonic $O8$: O^+8 [Sagnotti 1995].



Λ = “tadpole” scalar potential

$$\delta S = - \int T e^{\gamma\phi} .$$

- ⇒ Residual NS-NS tension, from sources or vacuum energy.
- ⇒ From worldsheet: **IR divergences** → background shift.
[Fischler, Susskind 1986; Callan, Lovelace, Nappi, Yost 1986-7-8].

Heterotic $SO(16) \times SO(16)$: \sim vacuum energy, positive (more fermions).

Idea: Use δS in flux compactifications.

Flux compactifications

$$S \sim \int e^{-2\phi} [R + 4(\partial\phi)^2] - \sum_k \frac{1}{2} e^{\beta_k \phi} \frac{F_k^2}{k!} - T e^{\gamma\phi}.$$

- Orientifolds: $\gamma = -1$

- $k = (1, 3, 5)$ with $\beta_k = 0$,
- $k = 2$ gauge, with $\beta_k = -1$.

- Heterotic: $\gamma = 0$

- $k = 3$, with $\beta_k = -2$,
- $k = 2$ gauge, with $\beta_k = -2$.

⇒ Not quite the right spectra to treat $T e^{\gamma\phi}$ as a small perturbation.
(However, see [Angelantonj, Armoni 1999] and [Baykara, Robbins, Sethi 2022; Fraiman, Graña, Parra De Freitas, Sethi 2023])

Sum rules

What should we expect?

- ▣➡ Generalization of sum rules in [Maldacena, Nunez 2000]:
no dS or Minkowski [Basile, Lanza 2020].
- ▣➡ No-go can be avoided:
 - Orientifolds, negative tension.
 - Allowing **boundaries**.

Boundaries are particularly puzzling in these models [Dudas, Mourad 2000].

Freund-Rubin vacua

Idea: Balance the tadpole with fluxes.

Focus on heterotic $SO(16) \times SO(16)$. NS-NS F_3 :

- $AdS_7 \times S^3$ with magnetic flux [Mourad, Sagnotti 2016].
 - Perturbatively unstable. We can replace S^3 with M_3 Einstein, perturbative stability with $\mathbb{R}P^3$ [Basile, Mourad, Sagnotti 2018].
 - Non-perturbatively unstable [Basile 2019-20].
- $AdS_4 \times X_3 \times Y_3$ with two magnetic fluxes [Basile 2020; wip].

U(1) flux in the gauge sector:

- $\text{AdS}_8 \times S^2$ with magnetic flux [SR 2022].
 - Stable under perturbations of metric, dilaton, and U(1).
 - Indications of instability in the mixed abelian-nonabelian gauge sector from flat-space limit [Chang, Weiss 1979; Sikivie 1979].
 - Reminiscent of magnetic monopoles.

Take-home message: It is possible to balance the tadpole with fluxes and curvatures. Simple solutions are usually **unstable**.

- ➡ General issue of stability without susy [Ooguri, Vafa 2016].
- ➡ But we are certainly missing tools to address the problem.

A first-order formalism

Possible tools: **fake supergravity**, following [Freedman, Nunez, Schnabl, Skenderis 2003]: define operators D_M and \mathcal{O} such that

$$\left. \begin{array}{l} D_M \varepsilon = 0 \\ \mathcal{O} \varepsilon = 0 \end{array} \right\} + \text{ Bianchi id.} \Rightarrow \text{ EoMs of non-susy strings.}$$

▣ It is a solution-generating technique, in principle there is no physics.

For gravity and the dilaton it is possible [SR 2023]. Simplest possibility:

$$\begin{array}{l} D_M \varepsilon = (\nabla_M + \mathcal{W}(\phi) \Gamma_M) \varepsilon , \\ \mathcal{O} \varepsilon = (d\phi + g(\phi)) \varepsilon . \end{array}$$

Significant progress: **fluxes**. However,

$$D_M \varepsilon = (D_M^{\text{susy}} + \mathcal{W}(\phi) \Gamma_M) \varepsilon$$
$$\mathcal{O} \varepsilon = (\mathcal{O}^{\text{susy}} + g(\phi)) \varepsilon$$

+ Bianchi do **not** imply EoMs.

▣► Change spinor ansatz? Loop corrections?

Subtle points:

▣► Loop and tree-level effects would contribute in the same way.

▣► Non-uniqueness of fake susy equations.

Summary and conclusions

- ⇒ Non-susy strings: can compute Λ .
- ⇒ Use Λ as ingredient in **flux compactifications**.
 - No drastic changes.
 - Balance tadpole with fluxes: Freund-Rubin.
- ⇒ However, we are missing **techniques** to engineer vacua. Fake susy?

The landscape of non-susy strings remains largely uncharted!