



# Feedback on R2D2 Rare Decay Radial Detector

**Pierre Charpentier** 

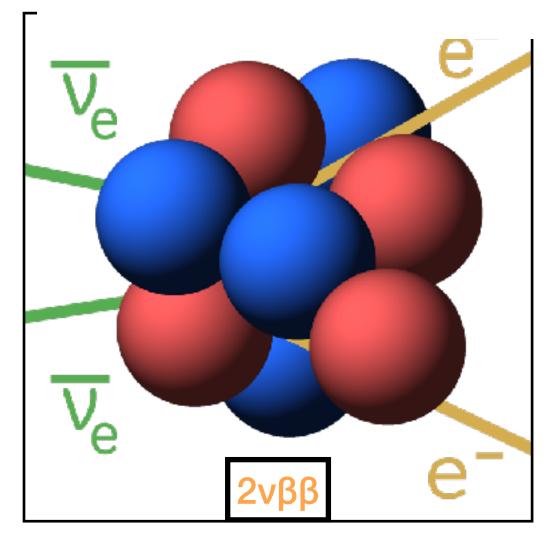
Journée Atelier Détecteur Gazeux

## université Bordeaux



# Neutrino, Majorana and *BB* Decay

- force in the exploration of outside standard model physics.
- observation of the  $0\nu\beta\beta$  decay
  - $2\nu\beta\beta$  decay :  $(A, \Xi)$

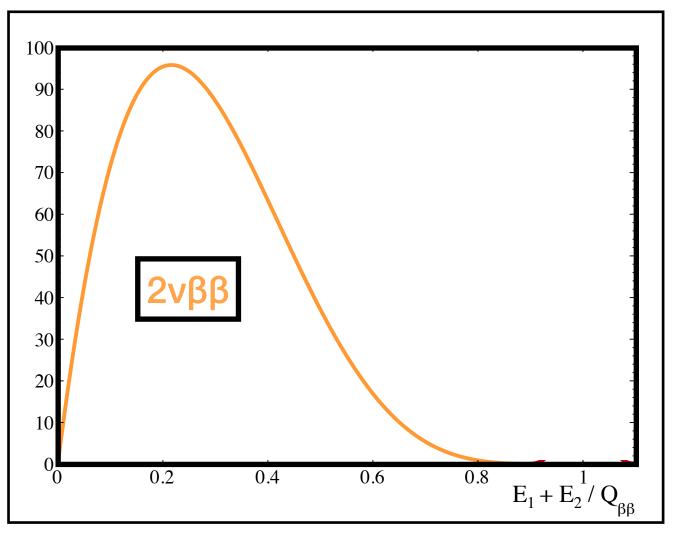


### **Pierre Charpentier**

• Since the massive neutrino postulate, neutrino research have been one of the leading

In the neutrino nature, Dirac or Majorana, mobilises global efforts and resources. The current most sensitive experimental proof of the Majorana nature of neutrino is the





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# Introduction Neutrino, Majorana and ßß Decay

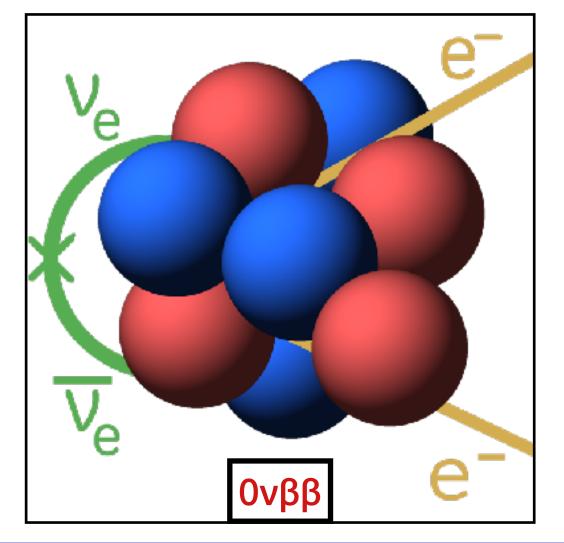
- force in the exploration of outside standard model physics.
- observation of the  $0\nu\beta\beta$  decay.
  - ►  $2\nu\beta\beta$  decay :  $(A,Z) \rightarrow (A,Z+2) + 2e^- + 2\overline{\nu}_{\rho}$  first direct observation in 1987.
  - $0\nu\beta\beta$  decay :  $(A,Z) \rightarrow (A,Z+2) + 2e^{-1}$  postulated in 1939.
    - Only possible if neutrino are Majorana particles
    - Violation of total lepton number
    - Total decay energy,  $Q_{\beta\beta}$ , share by the two electrons.

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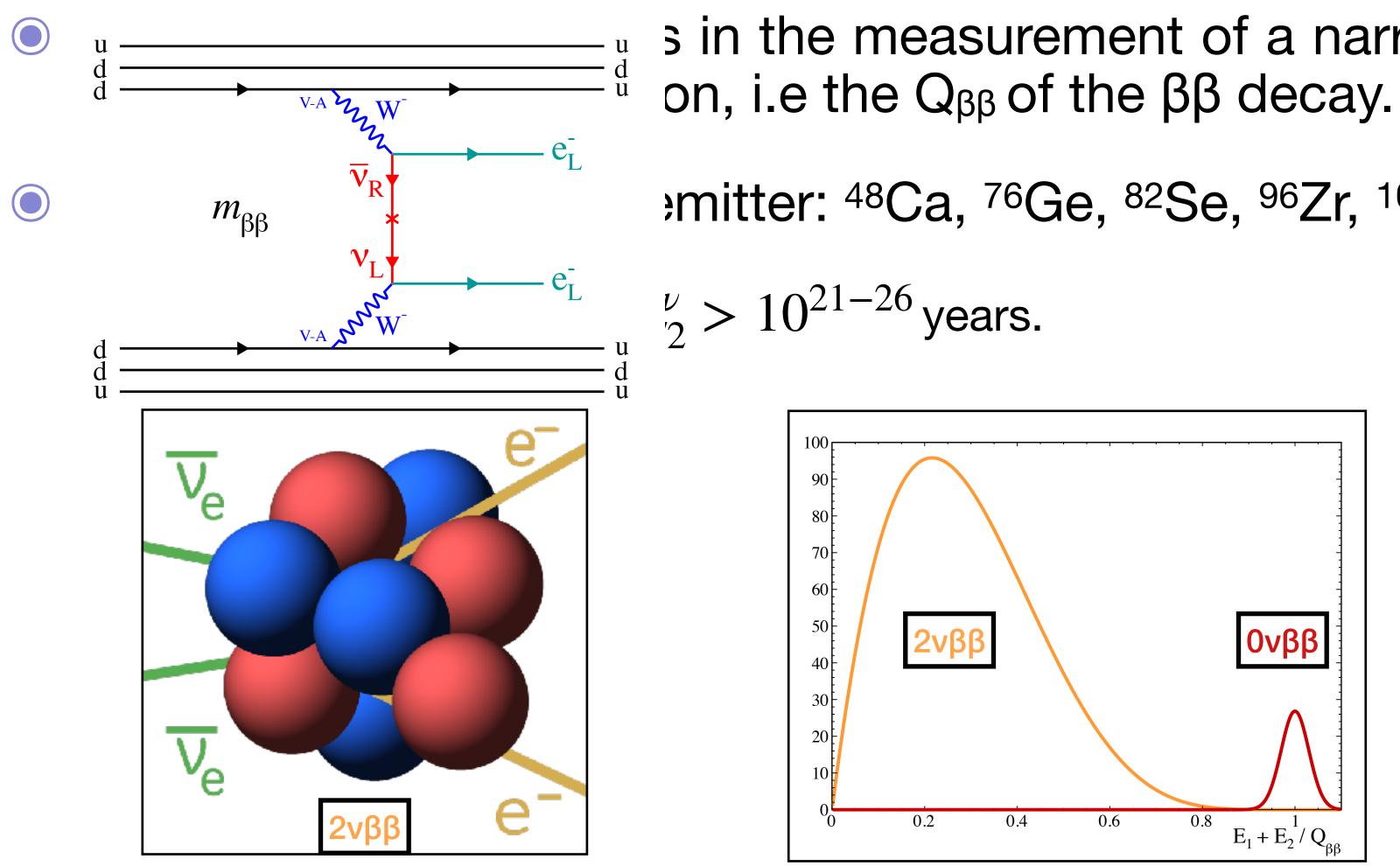








# **Ovßß Observation**

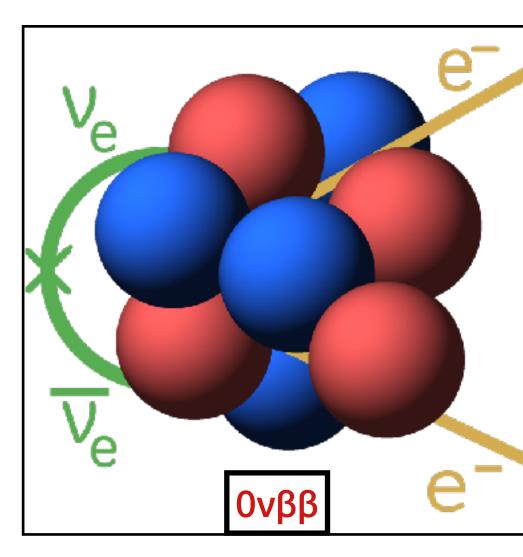


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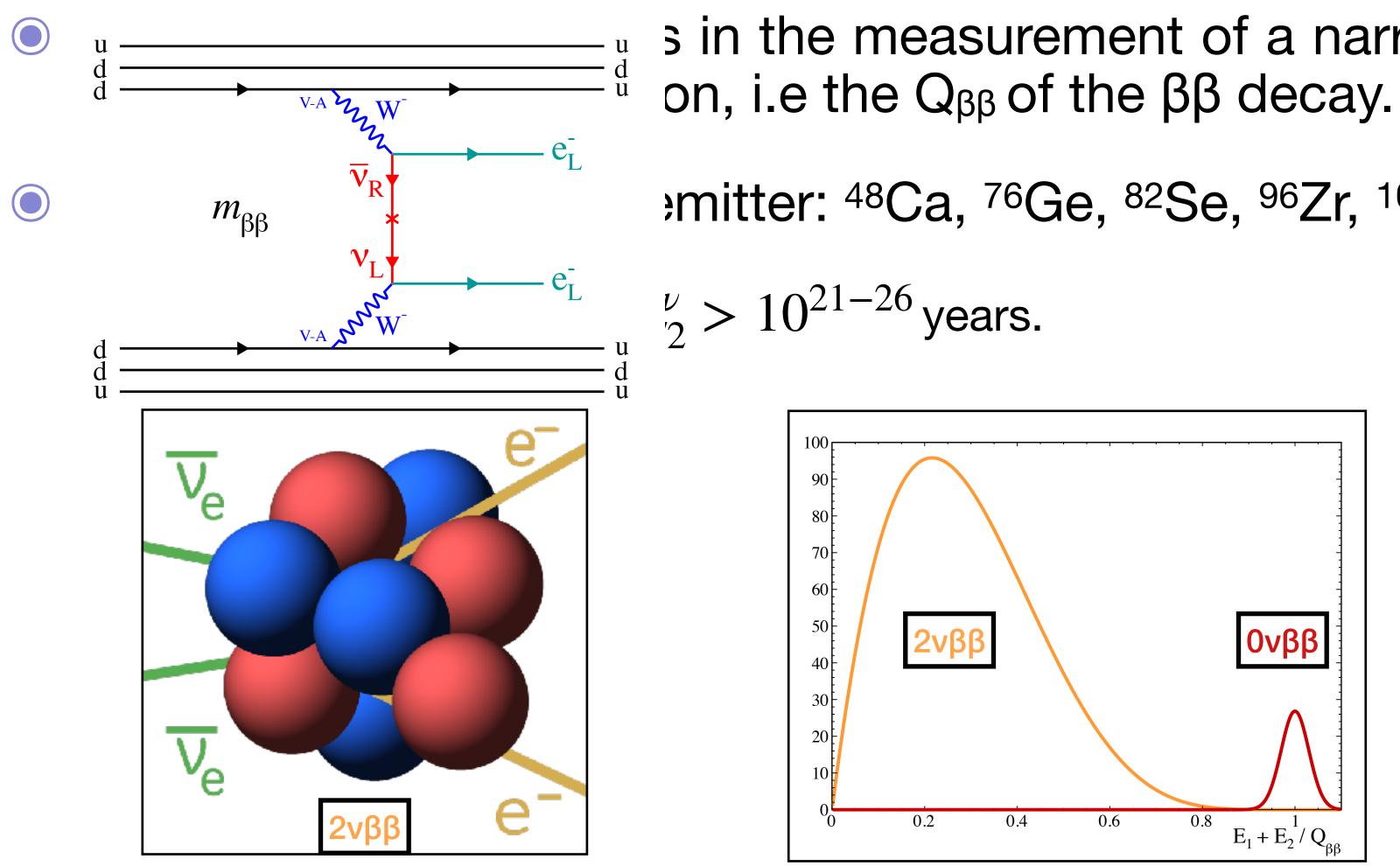
mitter: <sup>48</sup>Ca, <sup>76</sup>Ge, <sup>82</sup>Se, <sup>96</sup>Zr, <sup>100</sup>Mo, <sup>116</sup>Cd, <sup>128</sup>Te, <sup>136</sup>Xe, <sup>150</sup>Nd







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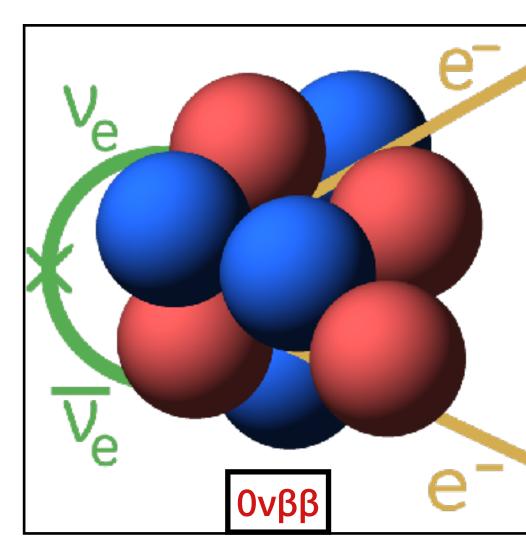


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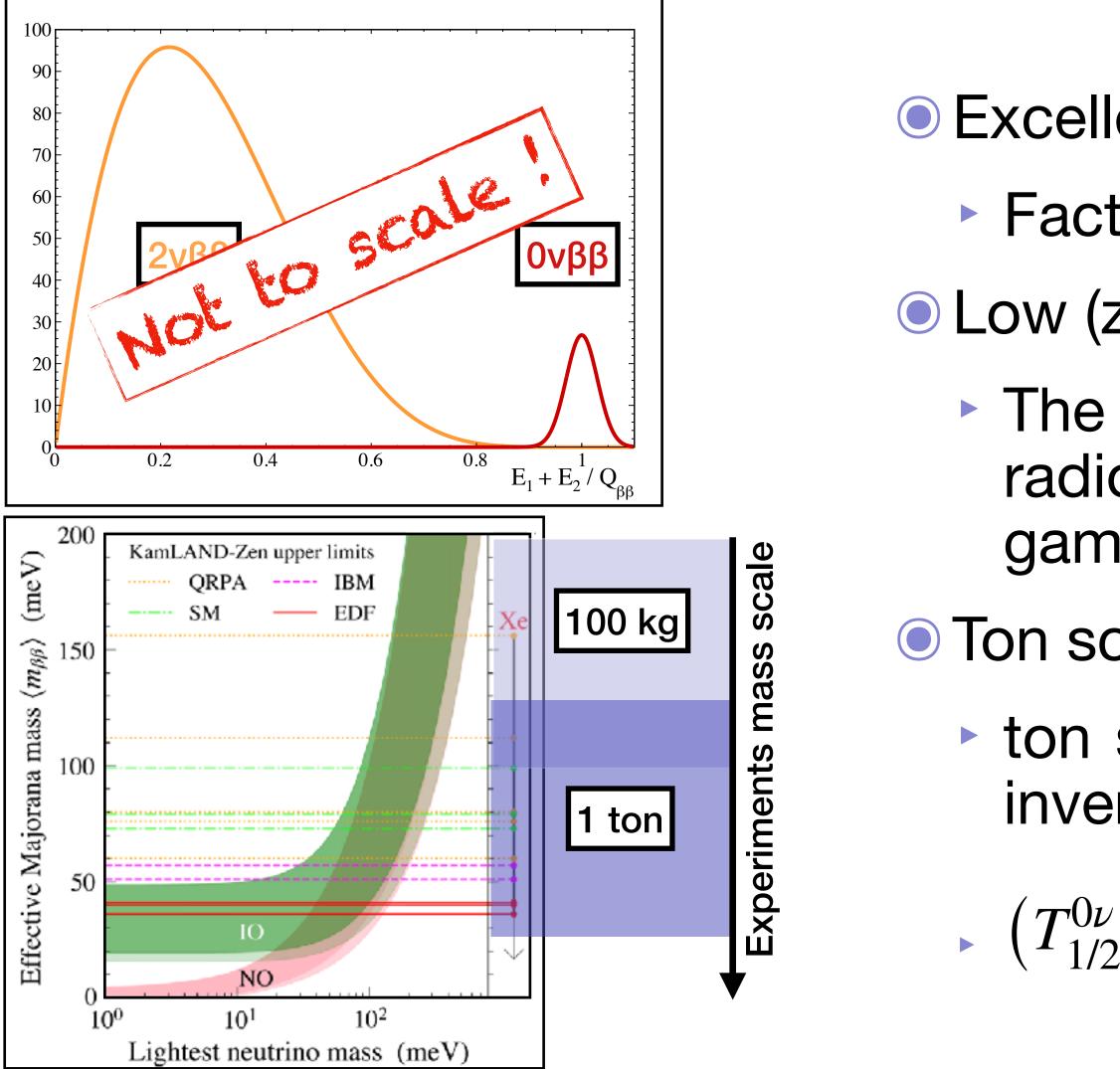
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# Requirements



### **Pierre Charpentier**

Excellent energy resolution

- Factor >10<sup>5</sup> years between  $2\nu\beta\beta$  and  $0\nu\beta\beta$  half-life.
- Low (zero) background.
  - The region of interest is surrounded by natural radioactivity. Ex: <sup>136</sup>Xe Q<sub>BB</sub> is 2.458 MeV and <sup>208</sup>TI gamma is at 2.615 MeV
- Ton scale experiment.
  - ton scale experiment is required to fully cover the inverse mass hierarchy region.

$$\binom{2}{2}^{-1} \propto \left| m_{\beta\beta} \right|^2$$







# **Curent Status**

	Energy resolution	Low background	Large isotope	
Solid state detectors			Large number of electronics c Difficult scalability t	
Liquid Xenon	Order of 4% at Q	Far from zero	Ton scale easily	
experiments	value	background		
Gaseous Xenon	Order of 1% at Q value	Far from zero	Complex detector I	
experiments		background	scale	

R2D2 Goals : 

### <sup>1</sup>HP-TPC: High Pressure Time Projection Chamber

### **Pierre Charpentier**

of crystals/ channels to large masses
y achievable
• Feasible at ton e?

e masses

• 
$$T_{1/2}^{2\nu} = 2.16 \pm .02 \cdot 10^{21}$$
 years.

• 
$$T_{1/2}^{0\nu} > 2.3 \pm .02 \cdot 10^{26}$$
 years.

- One of the most abundant, 8.86 %, and easiest to enrich.
- R2D2 is an R&D that explores a single anode HP-TPC<sup>1</sup> solution for  $0\nu\beta\beta$  search.
  - Meet all those requirements.
    - While also being able to perform tracking.

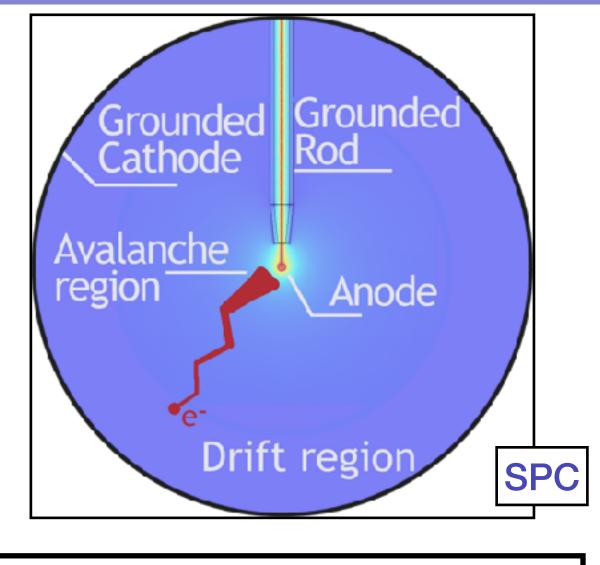








## Detectors



anode.

Design based on dark matter experiment



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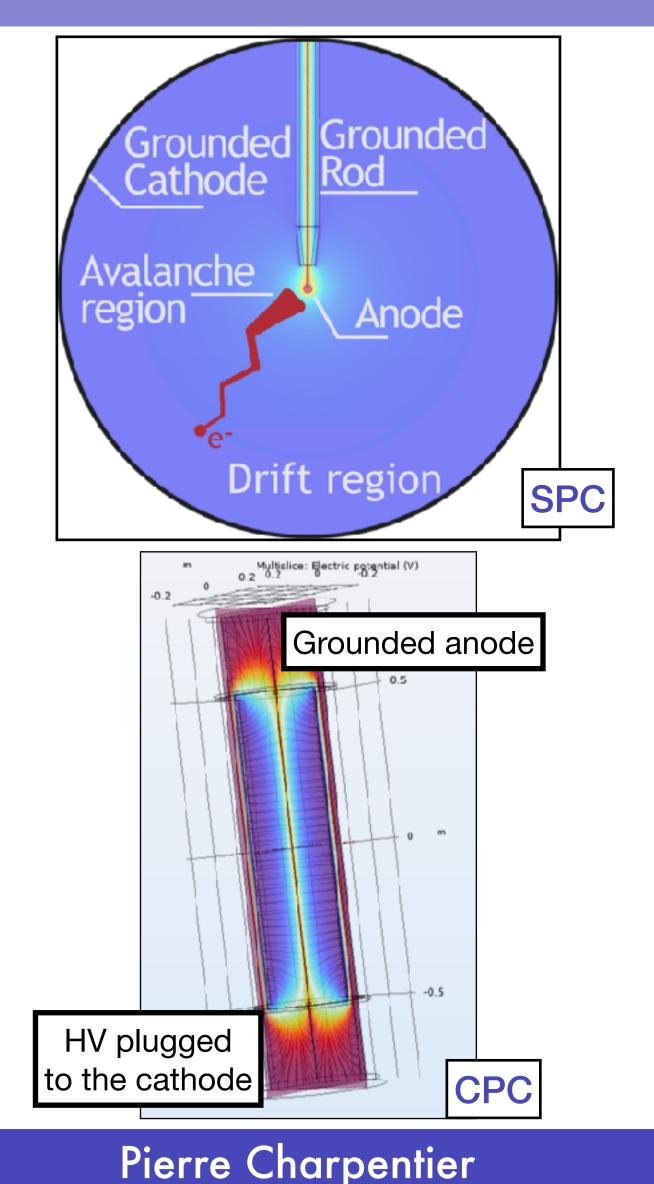
- The R&D is currently considering two detectors options:
  - Spherical Proportional Counter (SPC):
    - Spherical grounded cathode with a central
    - Both signal and tension goes through the anode.







## Detectors



- Spherical Proportional Counter (SPC):
  - Spherical grounded cathode with a central anode.
  - Both signal and tension goes through the anode.
- Optimized Proportional Counter (CPC):
  - Grounded tungsten wire as the central anode of a cylindrical copper cathode.
  - Signal is red through the anode and tension with negative polarisation is applied on the cathode.

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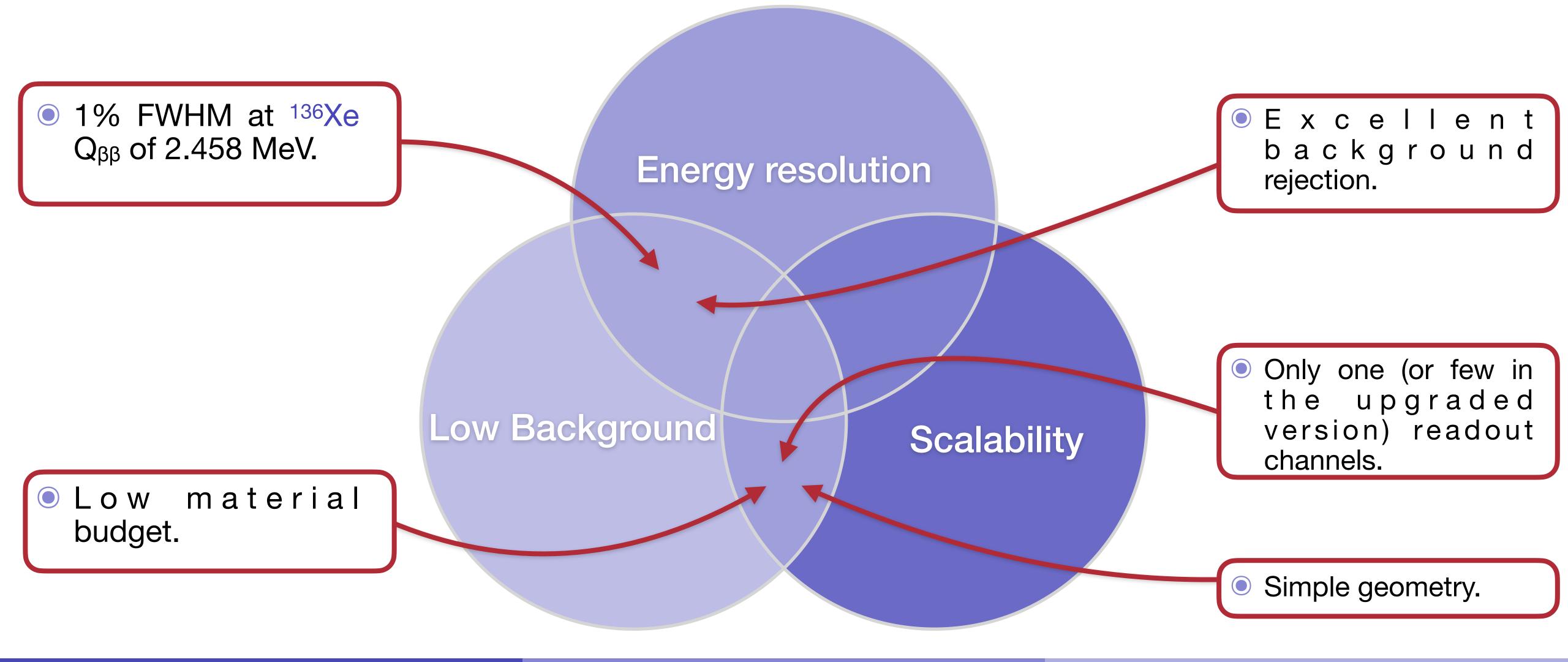
The R&D is currently considering two detectors options:





## Features

**R2D2** 



### **Pierre Charpentier**

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## **R2D2**

# **Current R&D Phases**

<sup>210</sup>Po as  $\alpha$  source.

- Exploring ionisation and proportional mode.
- Electronics and data acquisition.
- Sensor characterisation and improvement.
- Light readout<sup>1</sup>.
- Gas purity development.
- Gas recirculation and recovery.

<sup>1</sup> Nucl.Instrum.Meth.A 1028 (2022) 166382 [arXiv:2201.12621]

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### Current prototype goal: Achieve 1% FWHM energy resolution at 2.458 MeV, $^{136}Xe Q_{\beta\beta}$ .







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2 First with Argon then with Xenon

9/06/2023





# **Experimental setup overview Prototype Setup Evolution At LP2I Bordeaux**



### **SPC-1** (2018) 40 cm Ø Up to 1 bar<sup>1</sup>

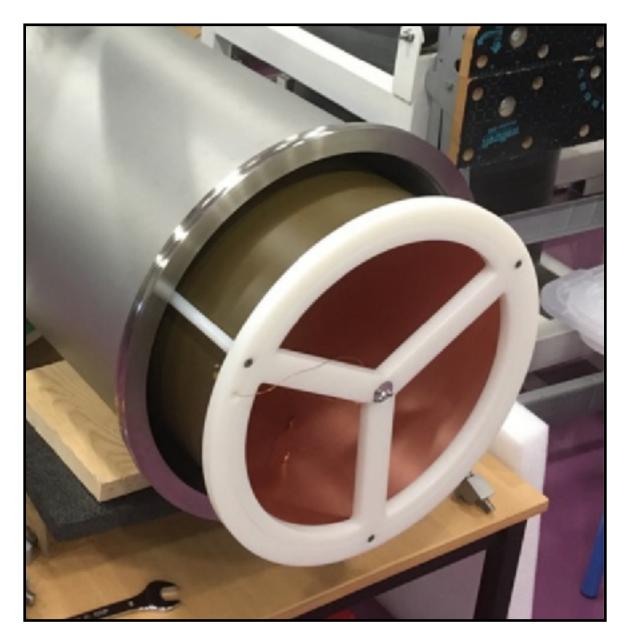


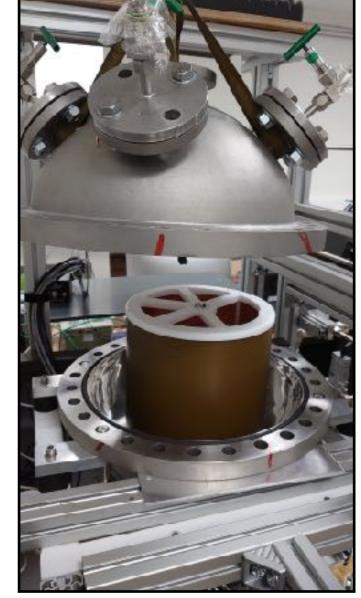
**SPC-2** (2021) 40 cm Ø Up to 40 bar<sup>2</sup>

<sup>1</sup> No Pressure certification <sup>2</sup> Pressure certified

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### **CPC-1** (2022) 1m x 37 cm Ø Up to 1 bar<sup>1</sup>

**CPC-2** (2023) 27cm x 20 cm Ø Up to 40 bar<sup>2</sup>

CPC Made at

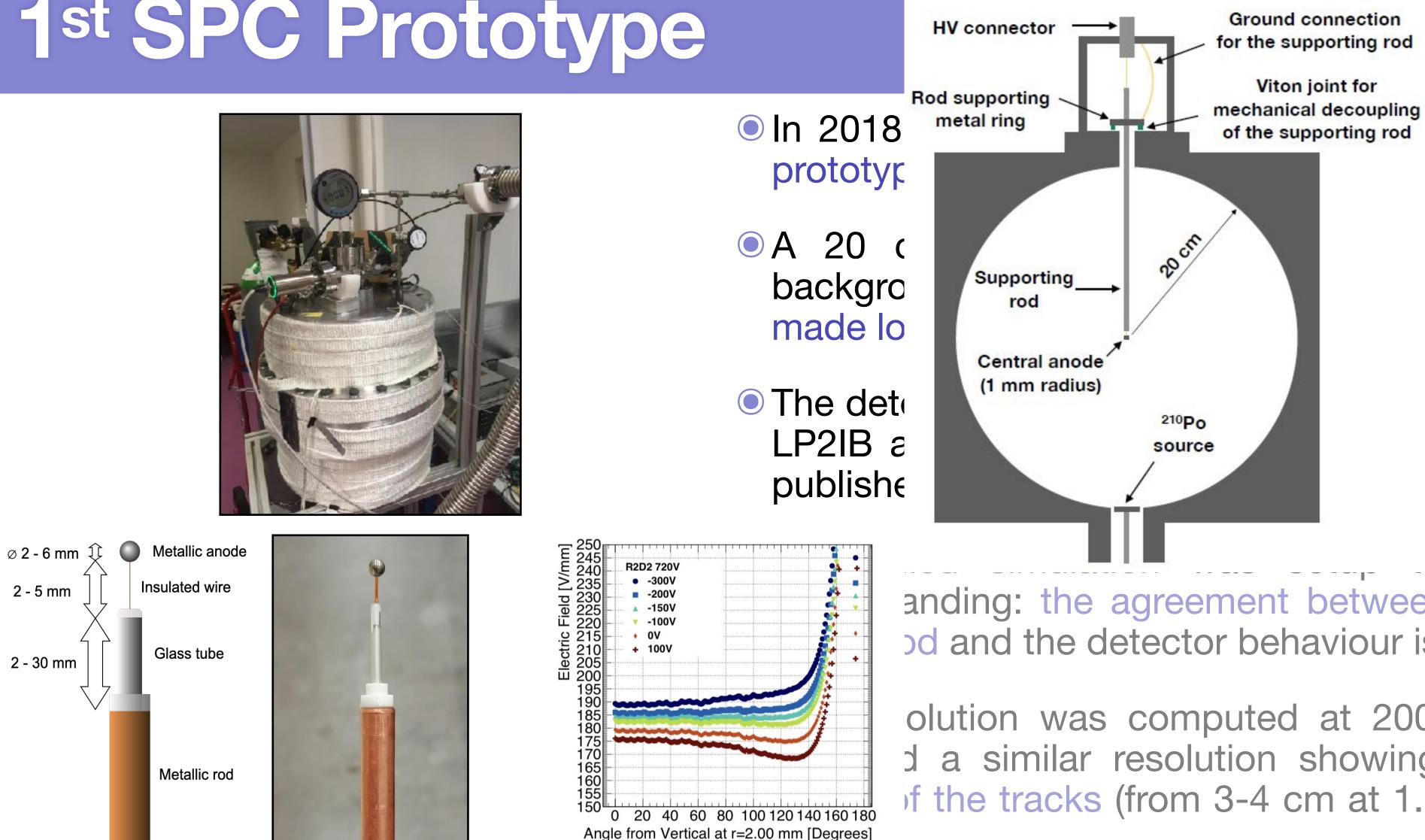


9/06/2023



 $\mathbf{O}$ 

## **Experimental setup overview** 1st SPC Prototype



### <sup>1</sup> JINST 16 (2021) 03, P03012 [arXiv:2007.02570] **Pierre Charpentier**

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# **Experimental setup overview** 1st SPC Prototype



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- mbar).

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### **Pierre Charpentier**

● In 2018 the R2D2 was funded as R&D by the IN2P3: 1<sup>st</sup> SCP prototype was built.

● A 20 cm radius sphere made of Aluminium (i.e. no low background but much cheaper) was built at LP2IB and a custom made low noise electronics (OWEN project) was developed.

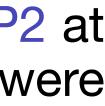
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A detailed simulation was setup to confirm our detector understanding: the agreement between data and simulation is very good and the detector behaviour is well understood.

The resolution was computed at 200 mbar and 1.1 bar: we obtained a similar resolution showing no impact due to the length of the tracks (from 3-4 cm at 1.1 bar to 15-20 cm at 200

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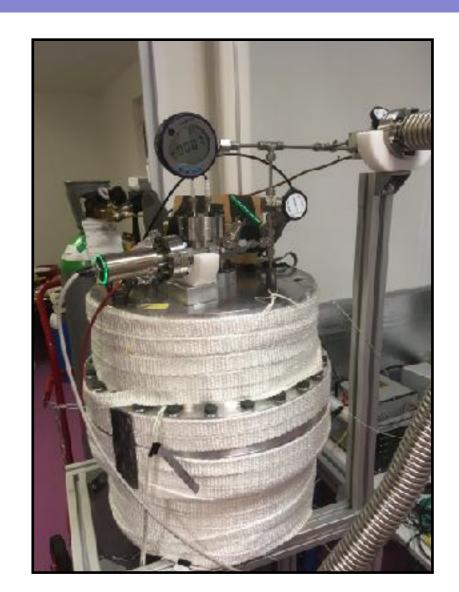




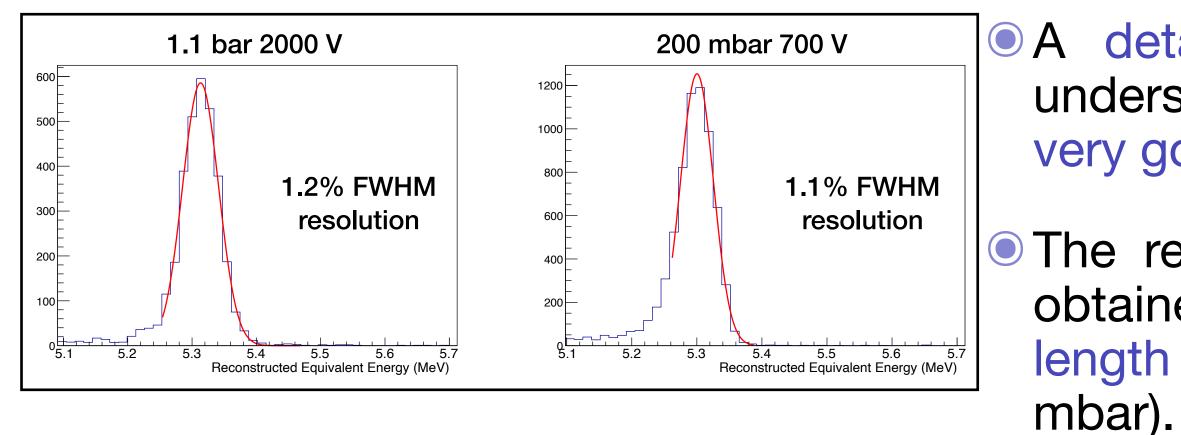




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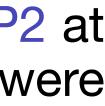
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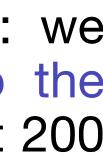
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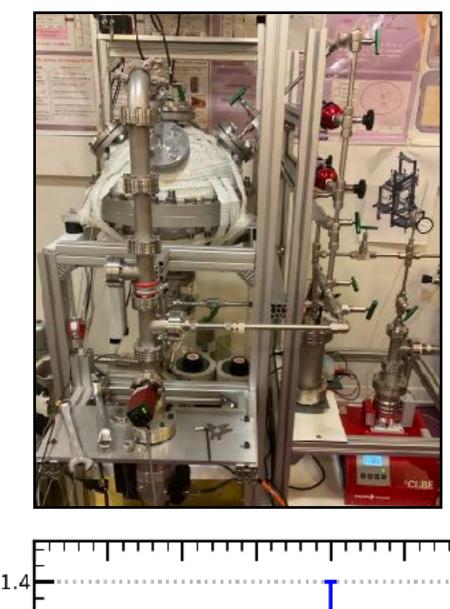


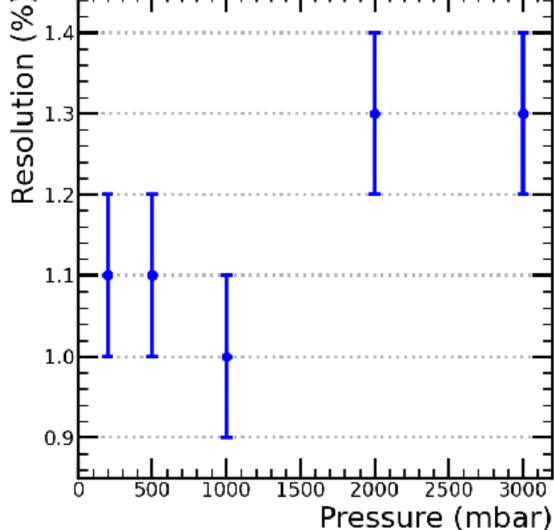






# **Experimental setup overview** 2<sup>nd</sup> SPC Prototype





**Pierre Charpentier** 

company.

In the meantime the xenon recirculation and recuperation system was finalized and commissioned.

In 2022 the detector was operated with ArP2 with pressures up to 3 bar. A set of measurement was carried out with a resolution below 1.4% up to 3 bar.

### In 2021 the second SPC prototype, certified to be operated up to 40 bar, was built by RAVANAT









## **Experimental setup overview**

# 1st CPC Prototype

Since May 2022 a new prototype is under study. A CPC exploiting the existing electronic chain, pumping and gas management system. First validation in ArP2 showed a resolution comparable with SPC (1.2% at 1 bar) at lower voltage.

- nox Tube: 1m50 x 40cm Ø.
- Copper cathode: 1m x 35 cm Ø.
- Tungsten anode: 50  $\mu$ m Ø.
- <sup>210</sup>Po source.

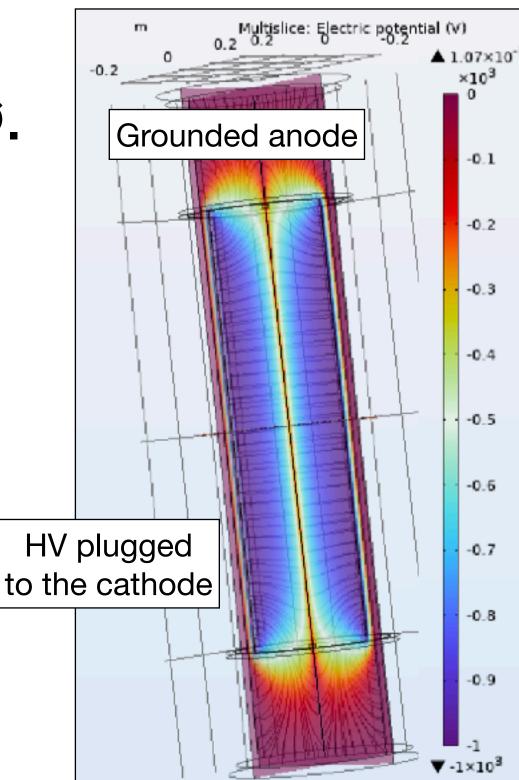
### **Electric Field:**

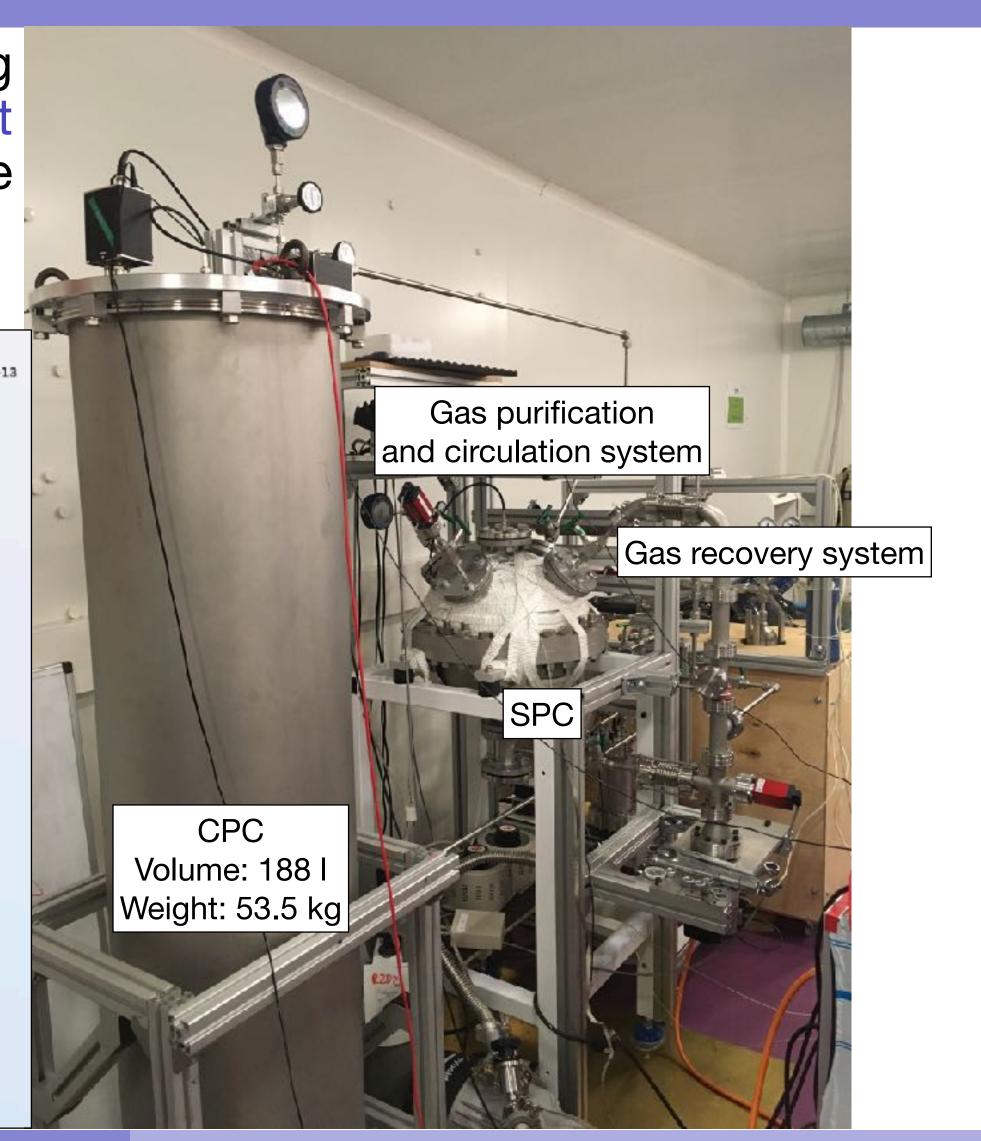
• SPC: 
$$\propto \frac{1}{r^2}$$

• CPC:  $\propto \frac{1}{r}$  (far from the edges)







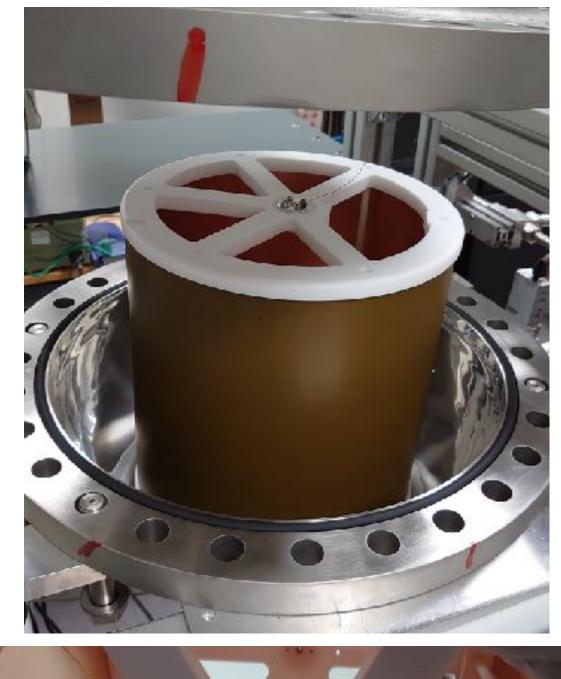


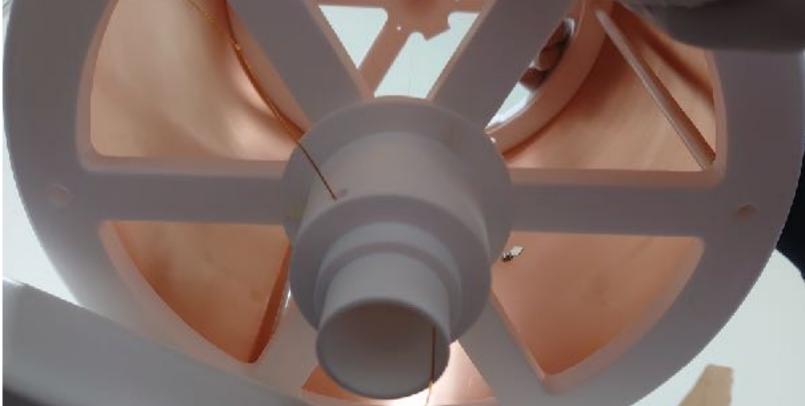
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9/06/2023

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# **Experimental setup overview** 2nd CPC Prototype





### **Pierre Charpentier**

high pressure.

bars as well.

- Early 2023 the second prototype for CPC was conceived and built at Substech.
- The CPC is designed to be operated inside the sphere (SPC 2nd prototype) in order to test the detector at
- First test in ArP2 showed a good behavior of the CPC up to 15 bars. Tests in xenon were carried out up to 3
- The limiting factor is still the gas purity. The hot getter was received in May 2023 and tests are in progress.



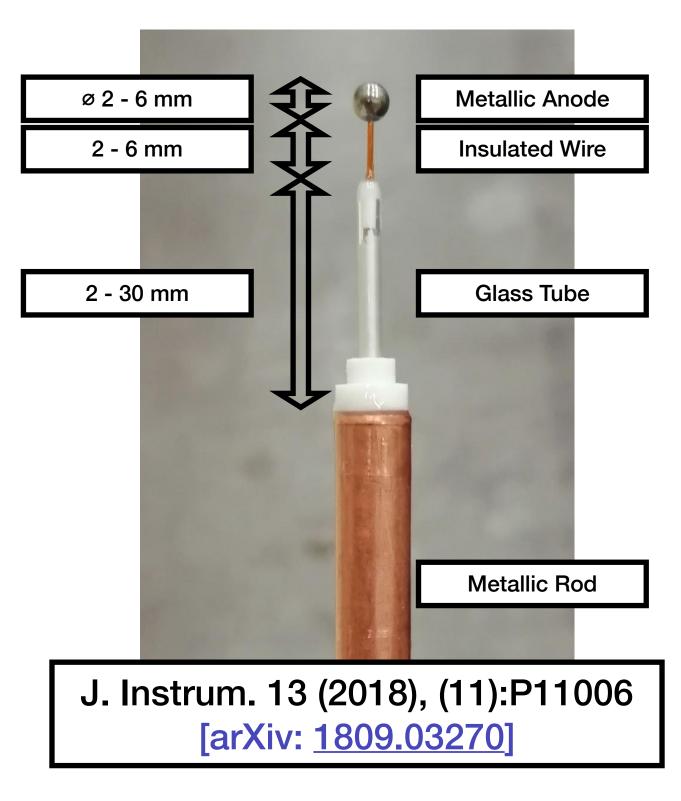






## **Technical development**

# SPC Sensor





### **Pierre Charpentier**



Multi channel sensor (Achinos) was also tested but channel equalisation is currently an issue for resolution.

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(1 mm radius)

210Po source

The sensor is the key point of the SPC detector.

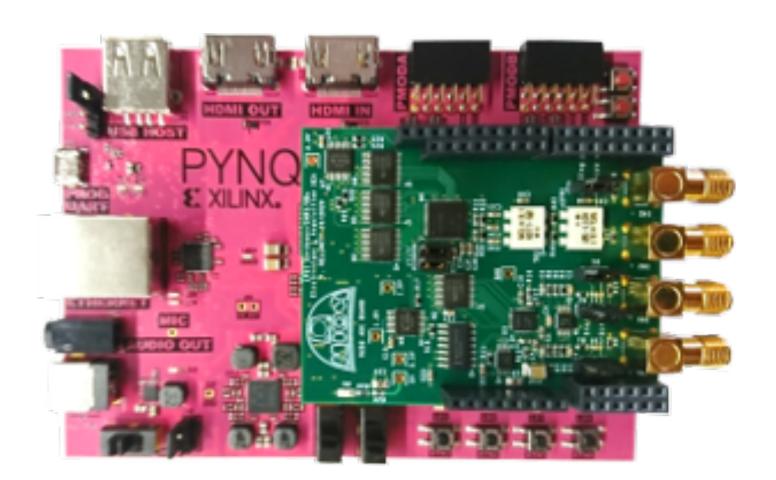
ith the ongoing R&D we learned a lot from the <sup>2</sup>detector functioning and we tested different option podifying the distance between the anode and the

he anode soldering to the wire is still a critical point since any imperfection results into a field distortion. soldering without of the anode and compromise



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# **Technical development Electronics and DAQ**





To achieve the energy resolution requirement, a dedicated low noise electronics chain is essential.

- today in NEWS-G.
- processing.

### **Pierre Charpentier**

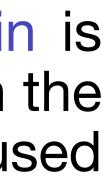
A low noise preamplifier was developed and a DAQ chain is under development (foreseen for end-2023) while we use in the meantime the CALI card developed for EDELWEISS and used

This work is carried out in the framework of the OWEN project (IdEX Emergence Université Bordeaux) which includes a dedicated development of onboard technology for a fast data

A specific work related to Artificial Intelligence is also ongoing in synergy with the THINK project of IN2P3 both for the final onboard technology and for the offline waveform processing in order to analyse signal and possibly reconstruct two-electrons tracks signature in the signal.

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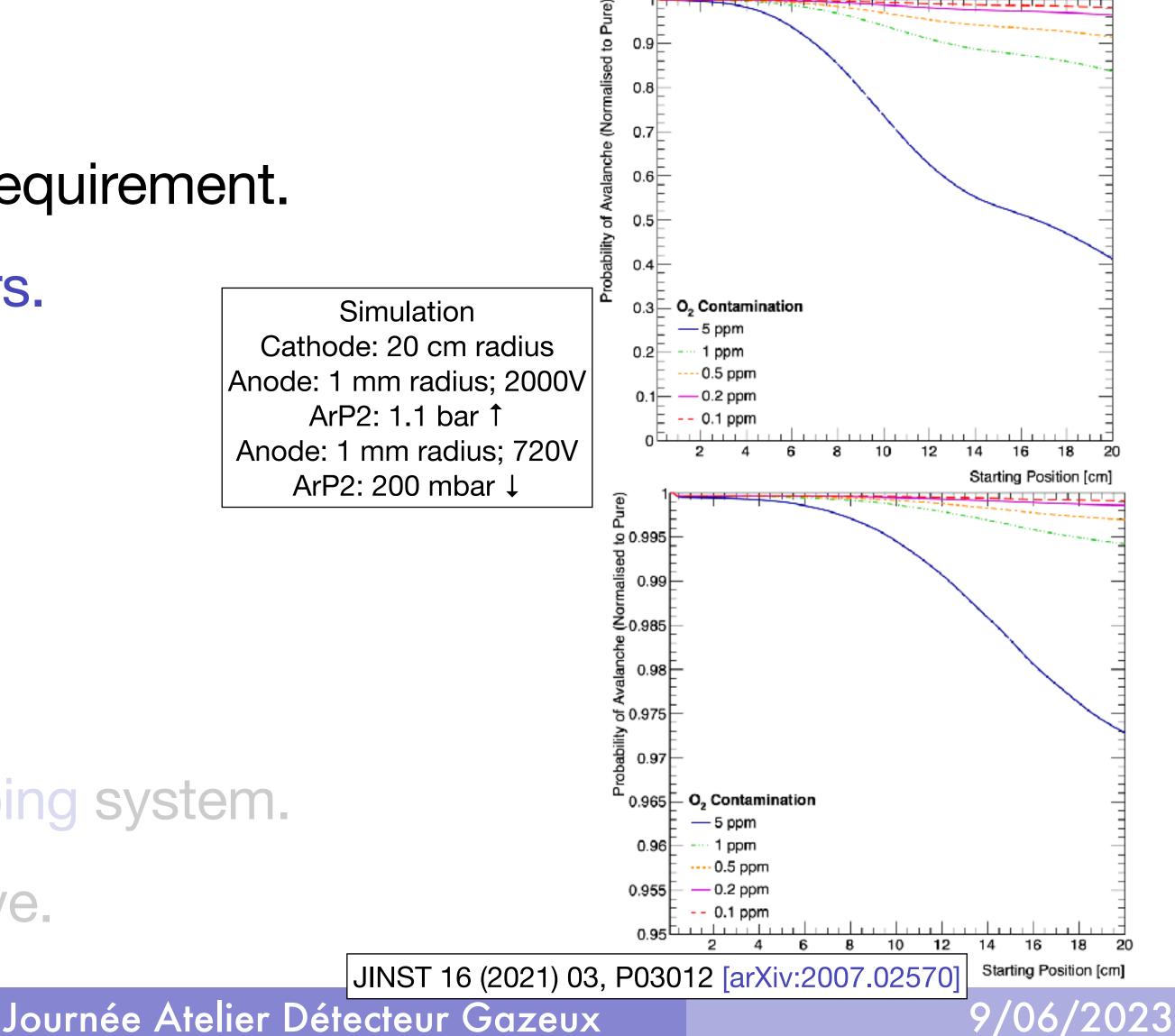




## Ourification:

- High purity is a strong requirement.
- Circulation inside getters.
- Recirculation:
  - Recirculation system.
  - Controlled flow.
- Recovery: First design by
  - Creation of a cryopumping system.
  - Pressure controlled valve.

### **Pierre Charpentier**



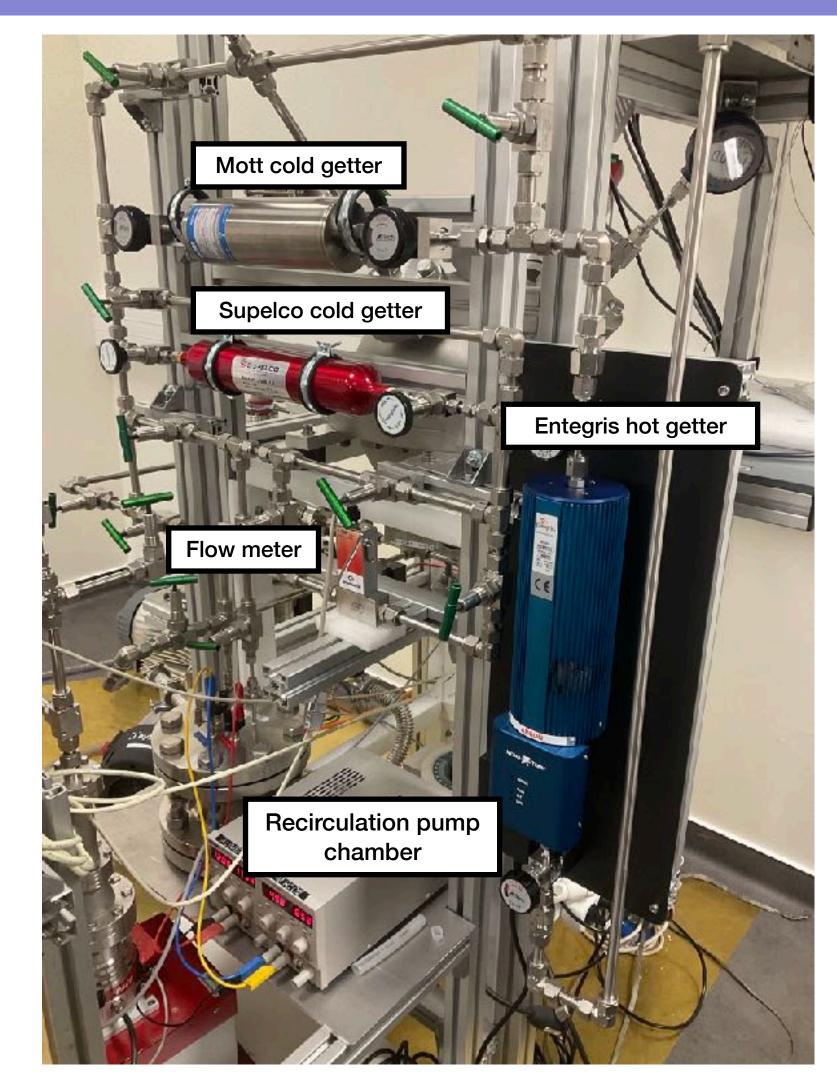


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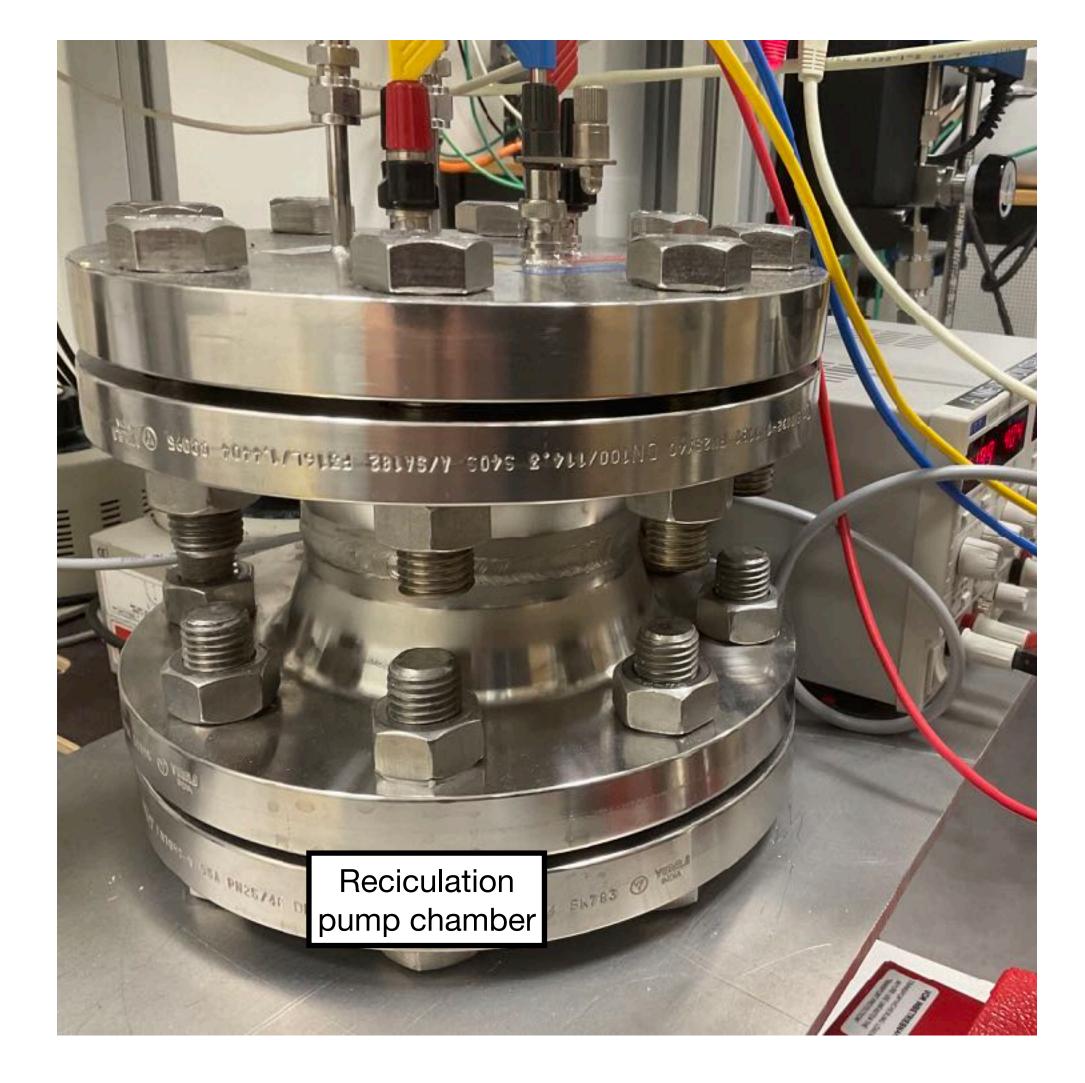


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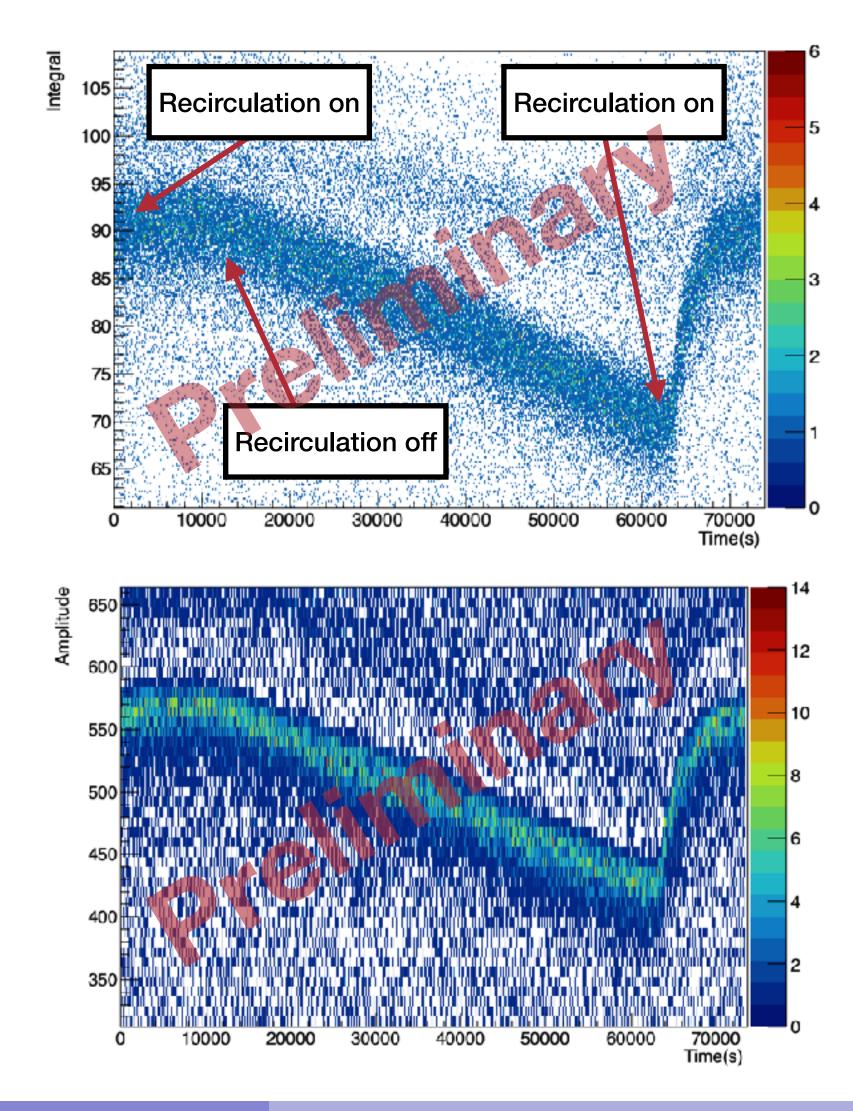




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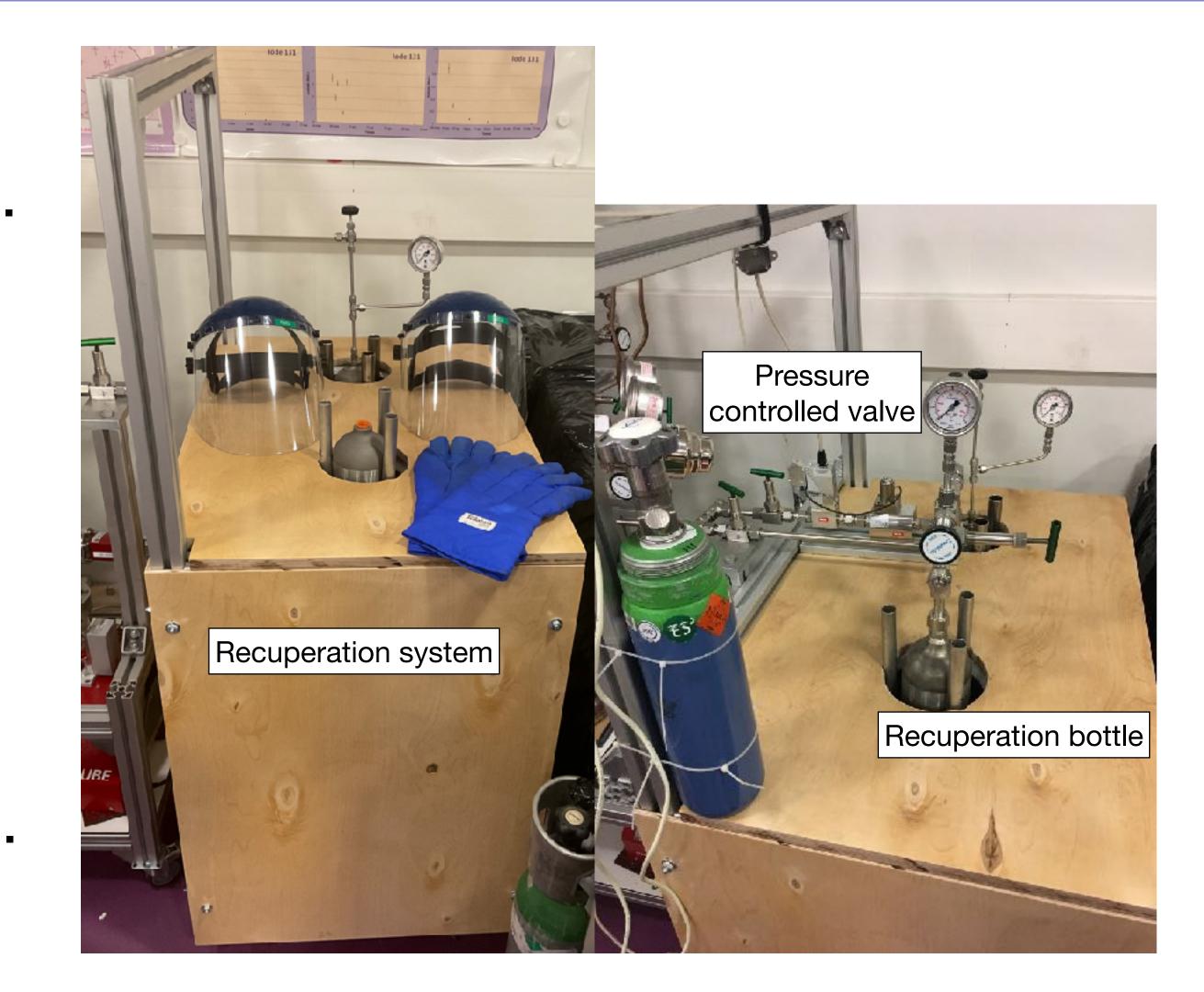
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### **Pierre Charpentier**



### Journée Atelier Détecteur Gazeux



## **Technical development**

## Results

Gas	Detector	Anode (radius)	Pressure (mbar)	HV (V)	Noise (ADU)	Gain	Resolution (%)
	SPC	1 mm	200	800	4.3	45	1.1
<b>ArP2</b>			500	1300	4.1	34	1.1
			1000	1900	4.2	30	0.9
			2000	2700	4.5	10	1.3
			3000	3900	4.8	10	1.3
		3 mm	1000	700	4.4	1	8.2
	CPC	10 µm	1000	200	3.6	1	4.9
			1000	900	3.9	9	1.2
Xe		3 mm	250	1300	4.5	1.3	3.8
	SPC		900	1300	4.4	1	7.2
	CPC	10 µm	500	900	3.8	20	1.8
			1000	1200	3.9	14	1.8

### **Pierre Charpentier**

SPC in proportional mode is limited at HV since noise increases with HV. In ionization mode the limit is given by the electronic noise.

Noise in the CPC is smaller (independent on the HV) and resolution in ionization mode is better.

 OPC in proportional mode
 requires smaller HV and resolution is good enough. At high pressure the limit is given by the gas purity.











- The R2D2, R&D proud of its 6 years of experience, has achieved a specific knowledge of the detector.
- Those efforts yield to a validation of the detector up to 15 bar in ArP2 and up to 3 bar in xenon.
- The current gas purity limits are expected to be improved with new equipment.
- CPC is easier to operate and achieved better results. Therefore, it will be the design baseline for future steps.
- A dedicated low noise electronics has to be developed in particular to work in ionisation mode and to read the sensor at both ends to reconstruct longitudinal position.









### **Pierre Charpentier**

### Journée Atelier Détecteur Gazeux



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# Backups

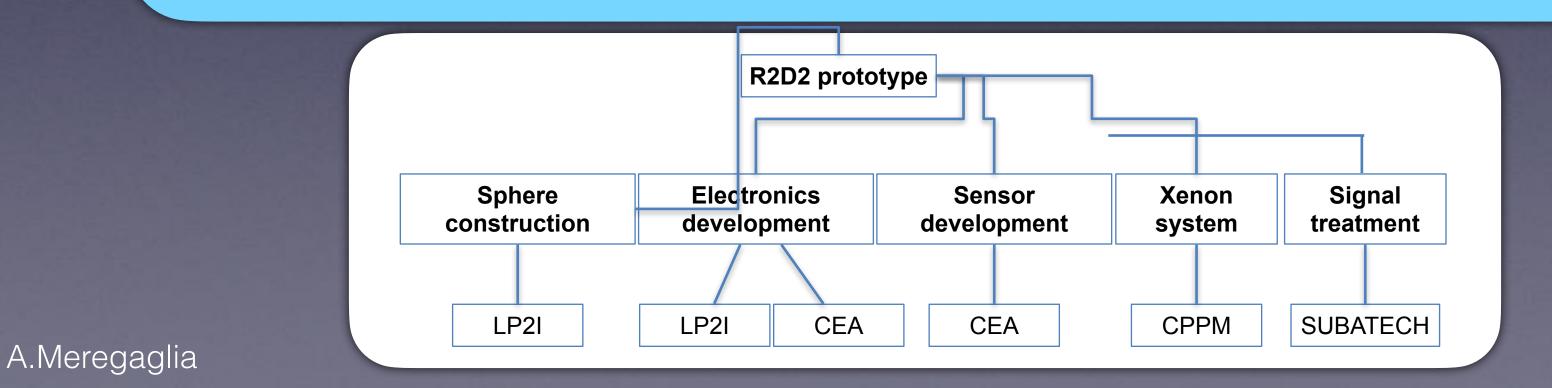


# R2D2 collaboration

- 2023).
- $\bullet$ energy resolution which is the major showstopper.

**R.** Bouet,<sup>*a*</sup> J. Busto,<sup>*b*</sup> V. Cecchini,<sup>*a*</sup> C. Cerna,<sup>*a*</sup> P. Charpentier,<sup>*a*</sup> A. Dastgheibi-Fard,<sup>*c*</sup> F. Druillole,<sup>*a*</sup> C. Jollet,<sup>*a*,1</sup> P. Hellmuth,<sup>*a*</sup> I. Katsioulas,<sup>*d*</sup> P. Knights,<sup>*d*,*e*</sup> I. Giomataris,<sup>*e*</sup> M. Gros,<sup>e</sup> P. Lautridou,<sup>f</sup> A. Meregaglia,<sup>a</sup> X. F. Navick,<sup>e</sup> T. Neep,<sup>d</sup> K. Nikolopoulos,<sup>d</sup> F. Perrot,<sup>a</sup> F. Piquemal,<sup>a</sup> M. Roche,<sup>a</sup> B. Thomas,<sup>a</sup> R. Ward<sup>d</sup>

<sup>a</sup>LP2I Bordeaux, Université de Bordeaux, CNRS/IN2P3, F-33175 Gradignan, France <sup>b</sup>CPPM, Université d'Aix-Marseille, CNRS/IN2P3, F-13288 Marseille, France <sup>c</sup>LSM, CNRS/IN2P3, Université Grenoble-Alpes, Modane, France <sup>d</sup>School of Physics and Astronomy, University of Birmingham, B15 2TT, United Kingdom <sup>e</sup>IRFU, CEA, Université Paris-Saclay, F-91191 Gif-sur-Yvette, France <sup>f</sup>SUBATECH, IMT-Atlantique, Université de Nantes, CNRS-IN2P3, France



A proto-collaboration has been formed (Czech colleagues from Prague joined R2D2 in April

R2D2 is today approved as IN2P3 R&D to assess in particular the possibility to reach the desired

# Sensitivity studies

• A **full Monte Carlo simulation** was developed to assess our capability to reject background and to evaluate the possible sensitivity on the searched signal.

• We considered a geometry including active and passive veto and a small mass of 50 kg of xenon corresponding to the foreseen prototype.

### Xenon active volume

Mass of 50 kg

Radius of 37 cm

Pressure of 40 bar

This choice, based on the results of a pressure and radius scan, is driven by the need of containing at least 80% of the  $\beta\beta0\nu$  electrons.

### Liquid scintillator volume

Thickness of 1.5 m Assumed to be LAB

The thickness is chosen in order to have a background rate below 0.1 events per year from the <sup>208</sup>TI contamination of the liquid scintillator vessel.

### A.Meregaglia

37 cm radius inner volume of Xe gas

0.5 cm thick Cu structure

1.5 m thick liquid scintillator

2 cm thick Cu structure

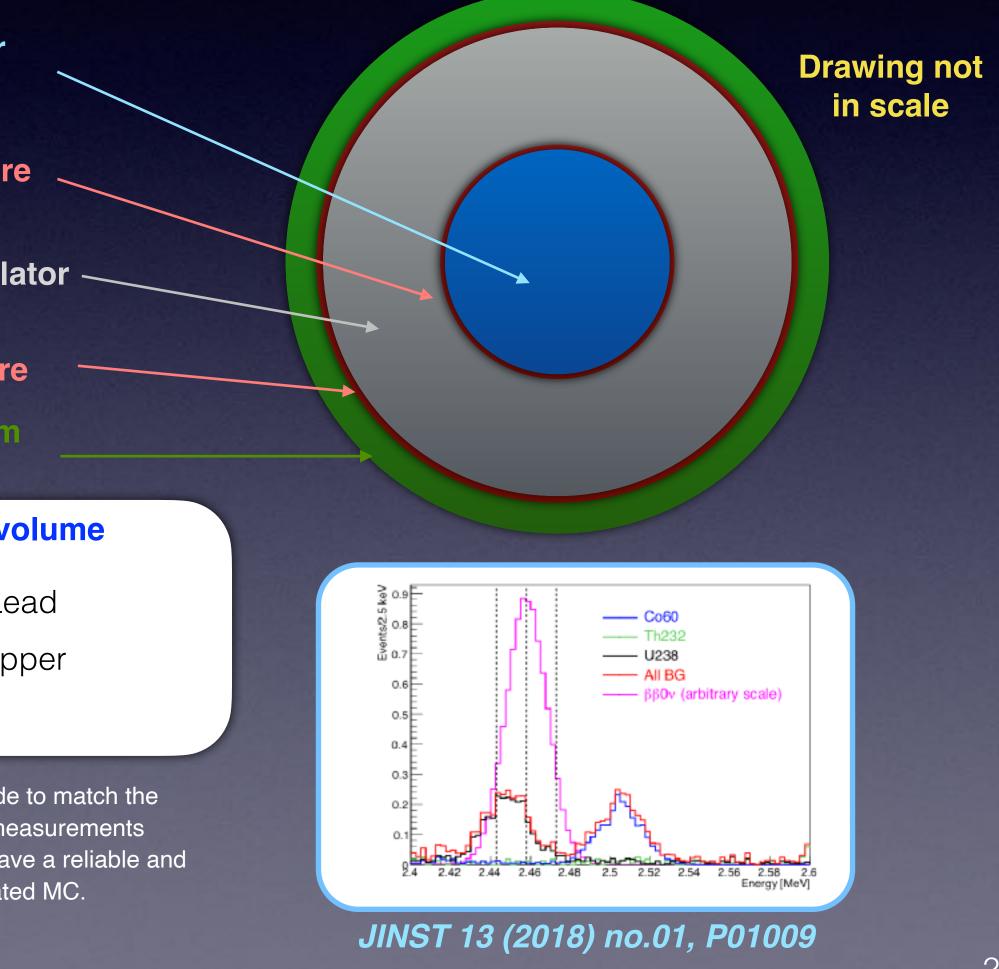
20 cm thick Pb + 5 cm thick Cu shielding

### **Shielding volume**

20 cm Lead

5 cm Copper

The choice was made to match the shielding used in measurements performed at LSM to have a reliable and less complicated MC.



# Released results

### PAPERS

- "Performance of a spherical high pressure gas TPC for neutrino magnetic moment measurement" R. Bouet et al. JINST 18 (2023) 03, P03031 [arXiv:2201.12621]
- "Simultaneous scintillation light and charge readout of a pure argon filled Spherical Proportional Counter" R. Bouet et al. Nucl.Instrum.Meth.A 1028 🛶 (2022) 166382 "[arXiv:2201.12621]
- "R2D2 spherical TPC: first energy resolution results" R. Bouet et al. JINST 16 (2021) 03, P03012 "[arXiv:2007.02570]
- "Study of a spherical Xenon gas TPC for neutrinoless double beta detection" A.Meregaglia et al. JINST 13 (2018) no.01, P01009 "[arXiv:1710.04536] 🛶

### We have presented the obtained results at various international conferences. $\bullet$

### TALKS

- Journé Matière Sombre France 2017 Paris 2017: "The R2D2 project" A.Meregaglia.
- Double Beta France workshop Paris 2018: "Status of the R2D2 project" A.Meregaglia.
- GET workshop Bordeaux 2018: "The R2D2 project" A.Meregaglia.
- 9th Symposium on Large TPCs for low-energy rare event detection Paris 2018: "A new neutrinoless double beta decay experiment: R2D2" A.Meregaglia
- Low Radioactivity Techniques Canfranc 2019: "A new neutrinoless double beta decay experiment: R2D2" A.Meregaglia.
- TAUP 2019 Toyama 2019: "A new neutrinoless double beta decay experiment: R2D2" C.Jol et (Talk given by G.Gerbier).
- ICHEP2020 Prague 2020: "First results of the R2D2 project" A.Meregaglia
- XIX International workshop on Neutrino Telescopes Venice 2021: "Latest results of the R2D2 project" T.Neep.
- TIPP2021 Virtual 2021: "Latest results of the R2D2 project" A.Meregaglia
- TAUP2021 Virtual 2021: "Status of the R2D2 project A future Ovββ experiment" I.Katsioulas.
- 10th LTPC symposium Paris 2021: "R2D2: An R&D program for the research of 2b0n decay with a SPC" PJ autridou.
- XeSAT2022 Coimbra 2022: "R2D2: a xenon TPC for neutrinoless double beta decay search" A.Meregag ia

### POSTERS

- Neutrino2020 Chicago 2020: "R2D2: a spherical high pressure TPC for the neutrinoless double beta decay search" V.Cecchini
- Neutrino2022 Seoul Virtual 2022: "R2D2: a xenon TPC for the neutrinoless double beta decay search" RCharpentier.

### A.Meregaglia

We have published 4 papers and recently submitted a new one on the first results on xenon.



Sensitivity to neutrino magnetic moment **Observation of scintillation light First results in ArP2 First sensitivity studies** 

## Technical developments **High Voltage**

- High voltages at the level of more than 20 kV might be needed when working at 40 bars (a possible back up option is to work in ionisation mode).
- At the moment we need however a feed through with several features:
  - Good for vacuum and high pressure
  - Good up to 10 kV (possibly more in the future)
  - Good for temperatures up to 100 degrees for detector heating.

### Low noise

- We tested several commercial options but each feed through has to be welded by a certified company and the behavior in terms of noise is not guarantee to be the same after and before.
- Discussion ongoing with AXON company (already collaborating) in JUNO) and prototype expected in 2023.

A.Meregaglia



Leakage current



Not shielded (noise)



Not certified for HP

Note: This is an issue only for SPC

## OWEN **Optimal Waveform recognition Electronic Node**

### Hardware developments:

- Very low noise front end
- Optimized waveform digitization with High resolution (18 bits)
- Embedded processor in integrated shape @ 1Gb/s

### **On-Line Embedded Artificial Intelligence:** $\bullet$

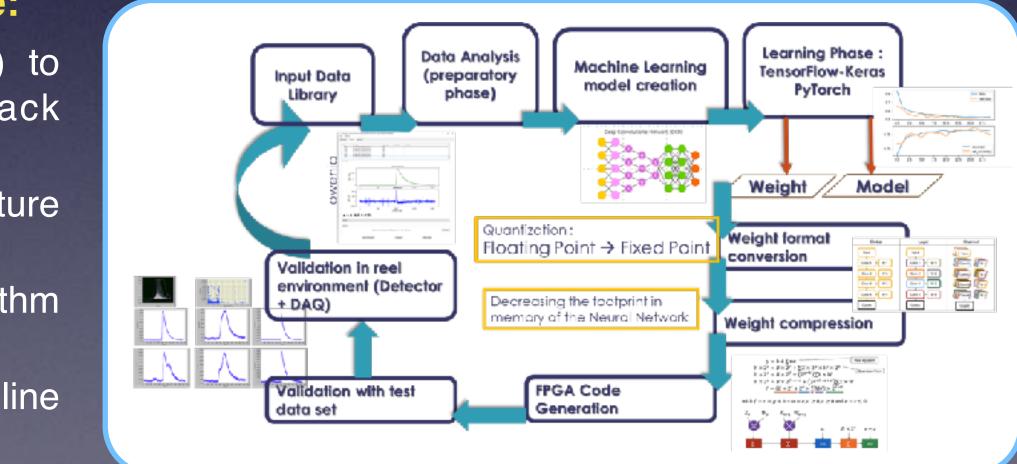
- Offline classification waveform (classic AI) to possibly reconstruct two-electrons track signature
- Research of a good neural network architecture to fulfill R2D2 needs
- Research of a process to integrate AI algorithm in embedded system
- Digital signal processing to tag events online (with embedded AI)

### Full system expected to be ready for experiments in 2023

### A.Meregaglia







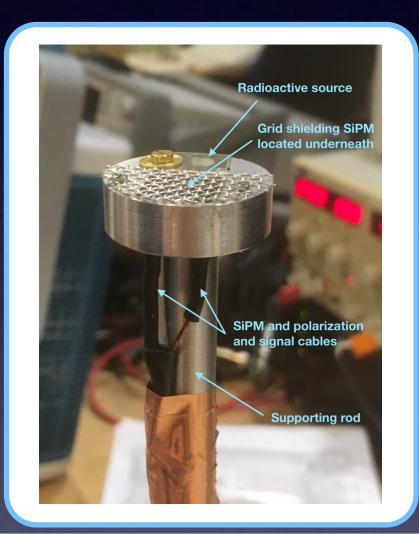
## Technical developments Light readout

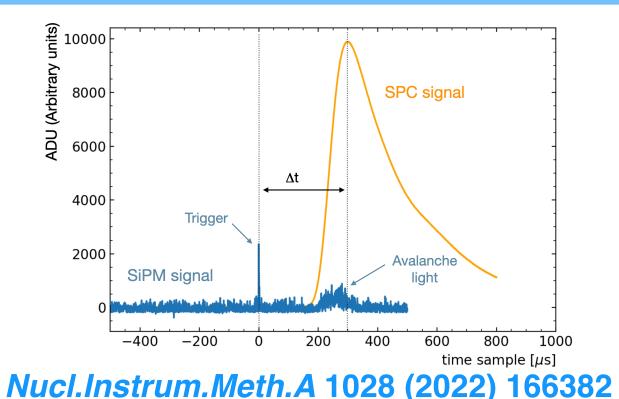
- $\bullet$ the radial position of the deposited energy with a precision of the order of few cm.
- (electron/gamma rejection).
- We run the detector in pure argon to observe the  $\bullet$ scintillation light and use it as trigger for the first time in a SPC detector.
- We used a 6x6 mm<sup>2</sup> SiPM from Hamamatsu with a 15% QE at 128 nm.
- We observed two signals on the SiPM: a trigger given by the scintillation light (S1) and a second signal on time with the SPC signal due to the light emitted in the avalanche (S2).
- The time between the S1 and S2 gives the electrons drift time and can be used to validate the Garfield++ simulation. An excellent agreement is found for alphas emitted at about 19 cm from the anode as expected.

### A.Meregaglia

So far spherical TPC (NEWS-G or SEDINE detectors) used the waveform rise time to reconstruct

Having an event trigger ( $T_0$ ) and knowing the drift velocity a sub-cm precision can be reached which is important for any fiducialization of the volume or to identify multiple energy depositions

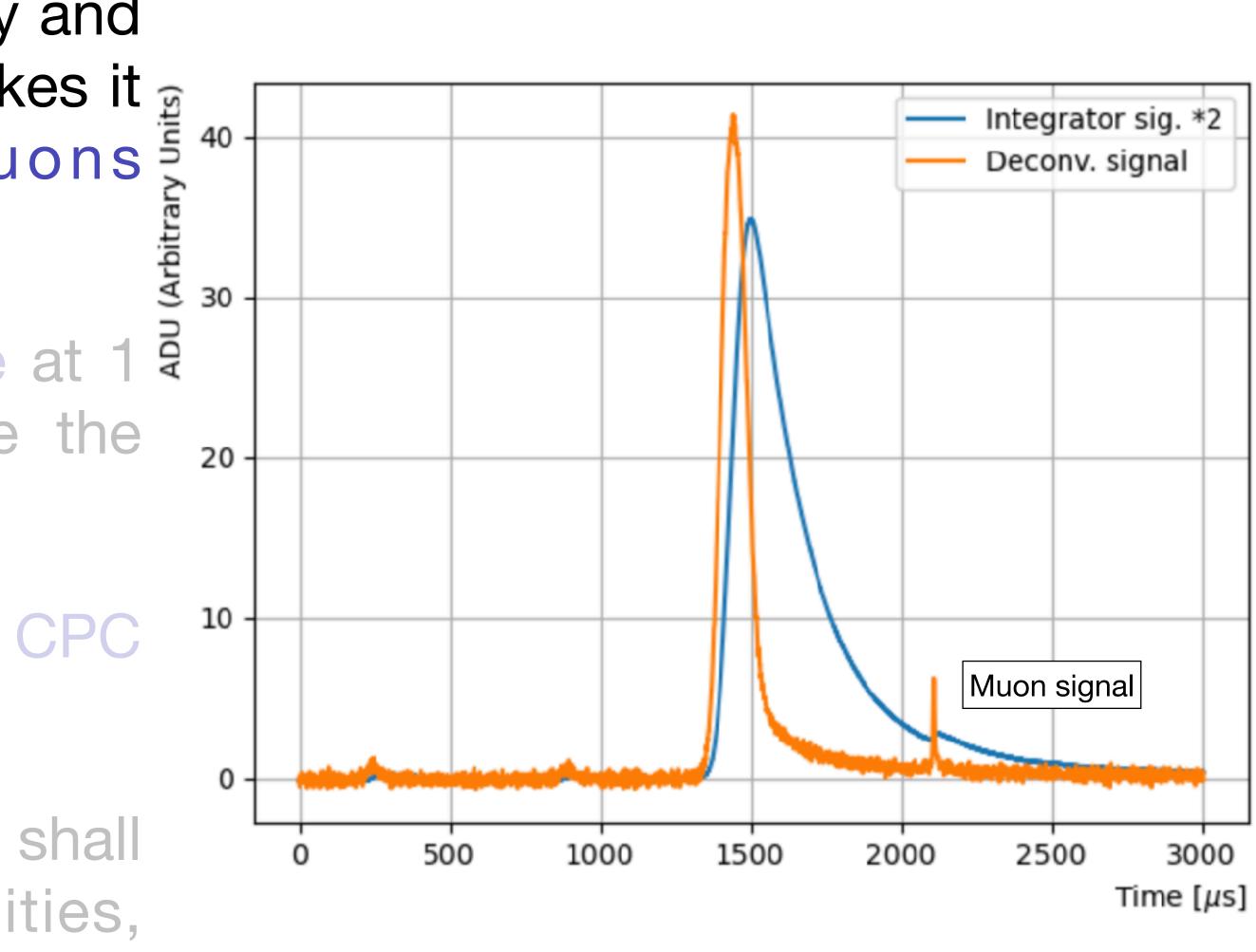




# **First Xenon Results CPC Xenon: Cosmic Background**

- Output Content SPC, the geometry and orientation of our CPC prototype makes it more sensible to cosmic muons background.
- The energy deposit of a muon in Xe at 1 \u00e4 bar is significantly enough degrade the energy resolution of the  $\alpha$  particles.
- This explain the right hand tail of the CPC reconstructed integral distribution.
- Nevertheless the final experiment shall take place in underground facilities, avoiding such inconveniences.

Pierre Charpentier



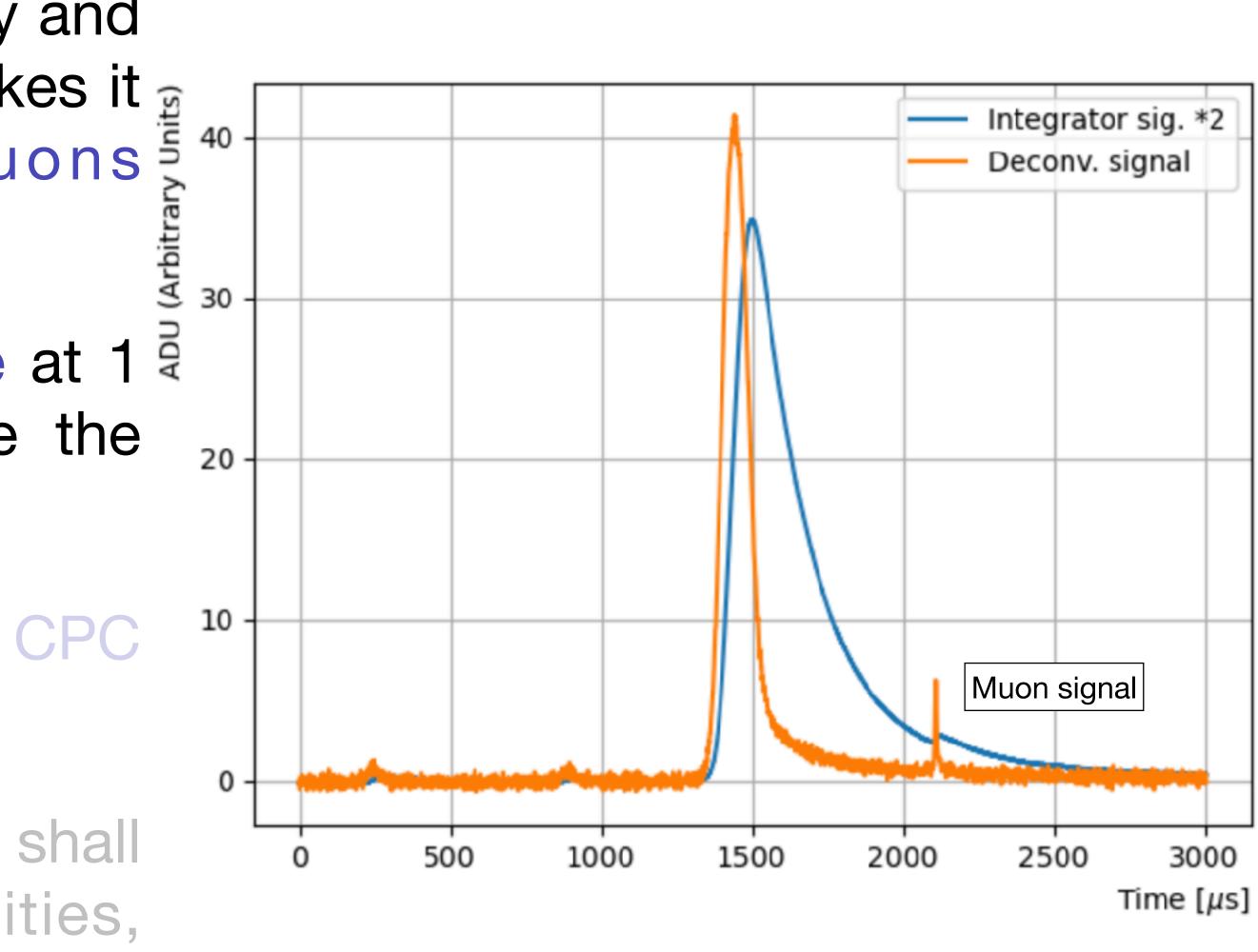
**IRN Neutrino** 

17/11/2022



# **First Xenon Results CPC Xenon: Cosmic Background**

- Output Content SPC, the geometry and orientation of our CPC prototype makes it  $\widehat{\mathbf{z}}$ more sensible to cosmic muons background.
- The energy deposit of a muon in Xe at  $1 \overline{4}$ bar is significantly enough degrade the energy resolution of the  $\alpha$  particles.
- This explain the right hand tail of the CPC reconstructed integral distribution.
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### **IRN Neutrino**

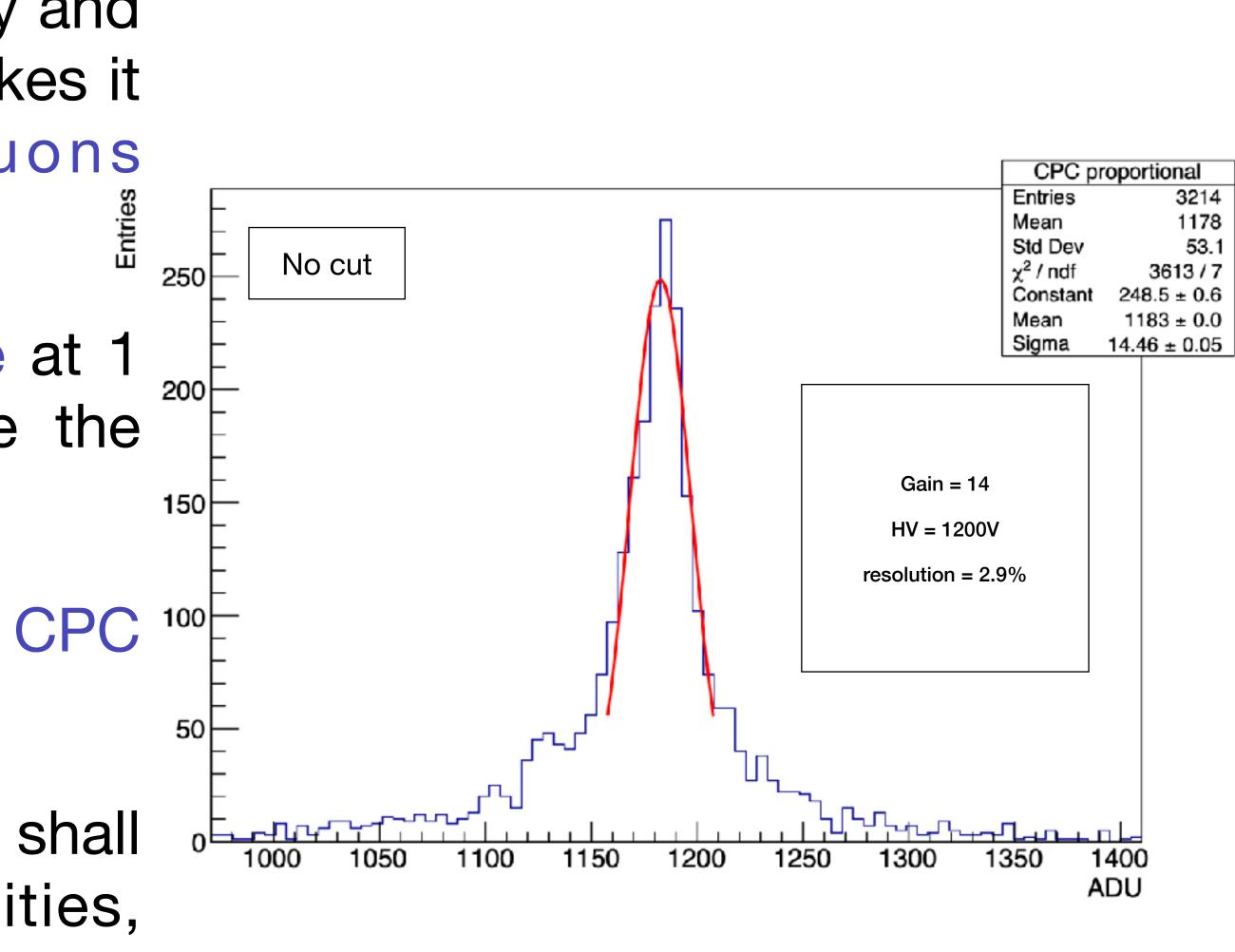
17/11/2022



# **First Xenon Results CPC Xenon: Cosmic Background**

- Output Content SPC, the geometry and orientation of our CPC prototype makes it more sensible to cosmic muons background.
- The energy deposit of a muon in Xe at 1 bar is significantly enough degrade the energy resolution of the  $\alpha$  particles.
- This explain the right hand tail of the CPC <sup>100</sup> reconstructed integral distribution.
- Nevertheless the final experiment shall of take place in underground facilities, avoiding such inconveniences.

**Pierre Charpentier** 



**IRN Neutrino** 

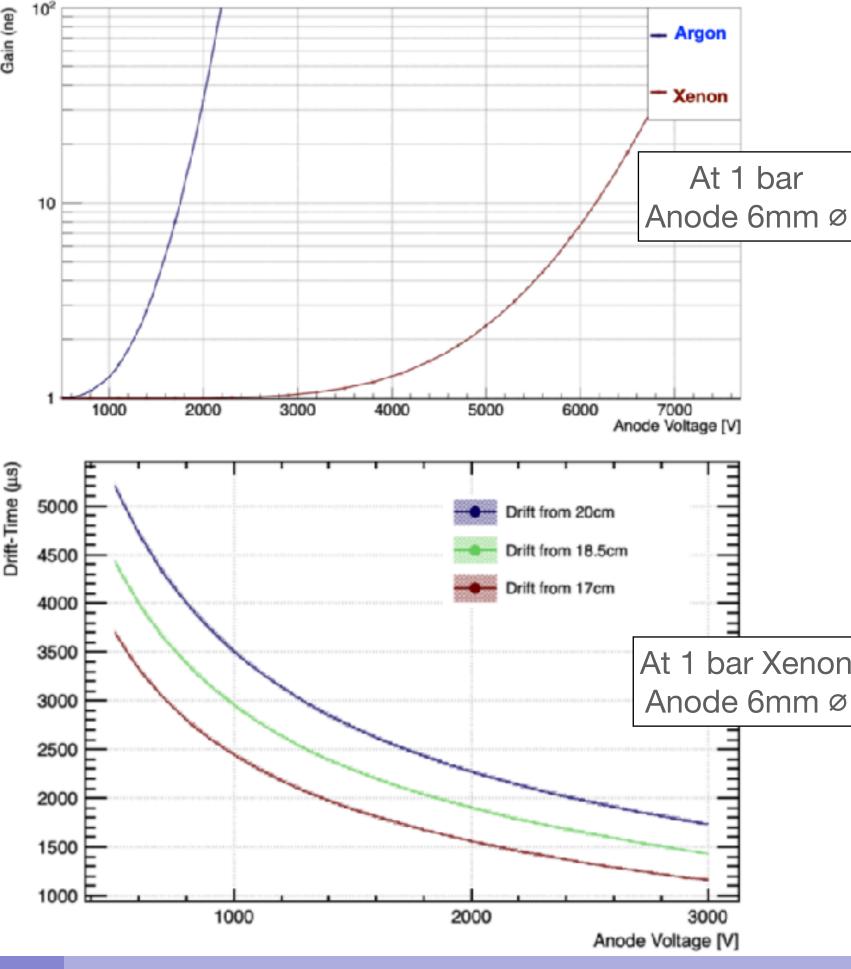
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# **First Xenon Results SPC Xenon: Main Difficulties**

Switching from Ar to Xe implied a lot of challenges to overcome. Aside from the previously discussed technical consideration: Argon

- Xe electrons drift time is one order of magnitude larger than Ar.
- Electronegative impurities become more critical. Purity is paramount.
- A stronger electric field is needed across the whole medium.
  - Higher  $HV \rightarrow$  higher noise.
  - Larger anode  $\rightarrow$  lonisation mode only.



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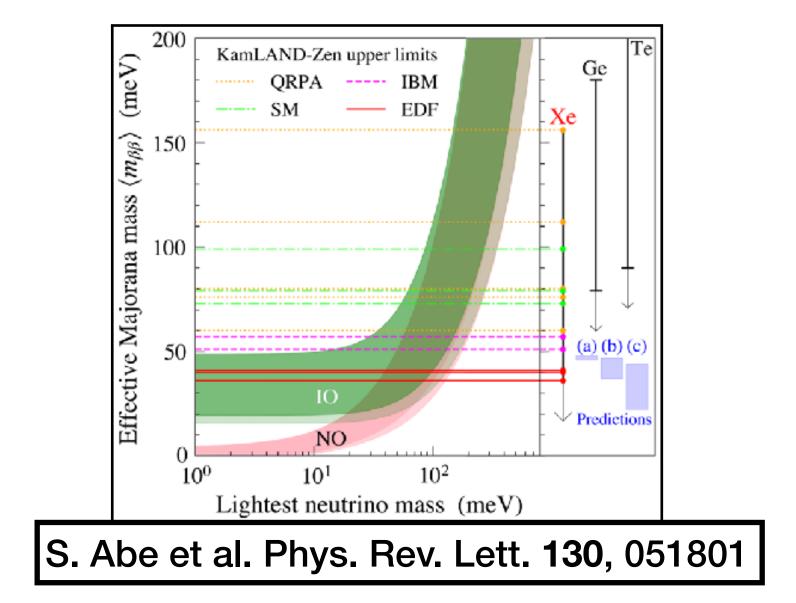
**IRN Neutrino** 





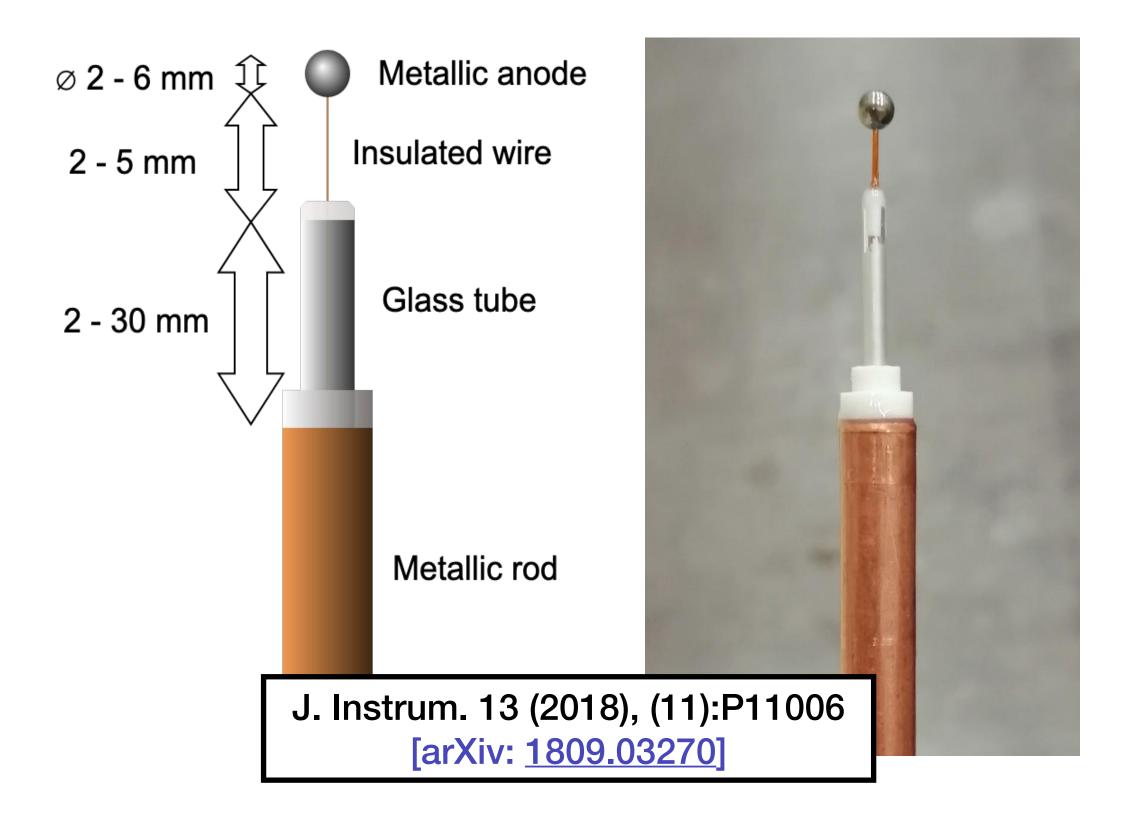






### **Pierre Charpentier**





Journée Atelier Détecteur Gazeux



