Large *N* Partition Functions, Holography, and Black Holes

Nikolay Bobev

Instituut voor Theoretische Fysica, KU Leuven

ENS Summer Institute

July 4 2023

 $2203.14981 + 2210.09318 + 2210.15326 + 2304.01734 +$ to appear

The large *N* team

Junho Hong (Leuven)

Valentin Reys (Saclay)

Supersymmetric localization allows for the exact calculation of physical observables in supersymmetric QFTs.

Apply this tool to SCFTs with holographic duals in string and M-theory.

 $Z_{\text{CFT}}[J] = Z_{\text{string/M}}[\phi]$

Supersymmetric localization allows for the exact calculation of physical observables in supersymmetric QFTs.

Apply this tool to SCFTs with holographic duals in string and M-theory.

$$
Z_{\text{CFT}}[J] = Z_{\text{string/M}}[\phi]
$$

Focus on subleading terms in the large *N* expansion to learn about quantum corrections to the supergravity approximation.

Supersymmetric localization allows for the exact calculation of physical observables in supersymmetric QFTs.

Apply this tool to SCFTs with holographic duals in string and M-theory.

$$
Z_{\text{CFT}}[J] = Z_{\text{string/M}}[\phi]
$$

Focus on subleading terms in the large *N* expansion to learn about quantum corrections to the supergravity approximation.

A new handle on AdS vacua of string and M-theory with non-trivial fluxes.

Learn about quantum corrections to black hole thermodynamics.

Supersymmetric localization allows for the exact calculation of physical observables in supersymmetric QFTs.

Apply this tool to SCFTs with holographic duals in string and M-theory.

$$
Z_{\text{CFT}}[J] = Z_{\text{string/M}}[\phi]
$$

Focus on subleading terms in the large *N* expansion to learn about quantum corrections to the supergravity approximation.

A new handle on AdS vacua of string and M-theory with non-trivial fluxes.

Learn about quantum corrections to black hole thermodynamics.

Goal: Describe recent progress on these topics for 3d SCFTs with $AdS₄$ duals in M-theory.

Plan

- Motivation √
- The ABJM theory on S^3
- **•** The ABJM topologically twisted index
- **•** Holography and black holes
- Other 3d $\mathcal{N}=2$ SCFTs
- The superconformal index
- **•** Thermal observables

The ABJM theory on *S* 3

ABJM and holography

The ABJM theory: $U(N)_k \times U(N)_{-k}$ CS-matter theory with two pairs of $\overline{\text{bi-fundamental chirals} } (A_{1,2}, B_{1,2})$ and superpotential

 $W = \text{Tr}(A_1B_1A_2B_2 - A_1B_2A_2B_1)$.

For $k > 2$ it has $\mathcal{N} = 6$ supersymmetry and $SU(4)_R \times U(1)_b$ global symmetry. Describes N M2-branes on $\mathbb{C}^4/\mathbb{Z}_k$.

ABJM and holography

The ABJM theory: $U(N)_k \times U(N)_{-k}$ CS-matter theory with two pairs of $\overline{\text{bi-fundamental chirals} } (A_{1,2}, B_{1,2})$ and superpotential

$$
W = \text{Tr}(A_1B_1A_2B_2 - A_1B_2A_2B_1).
$$

For $k > 2$ it has $\mathcal{N} = 6$ supersymmetry and $SU(4)_R \times U(1)_b$ global symmetry. Describes N M2-branes on $\mathbb{C}^4/\mathbb{Z}_k$.

 \bullet In the limit of fixed *k* and large *N*, the ABJM theory is dual to the M-theory background $\text{AdS}_4 \times S^7/\mathbb{Z}_k$

$$
(L/\ell_{\rm P})^6 \sim k\,N\,.
$$

ABJM and holography

The ABJM theory: $U(N)_k \times U(N)_{-k}$ CS-matter theory with two pairs of $\overline{\text{bi-fundamental chirals} } (A_{1,2}, B_{1,2})$ and superpotential

$$
W = \text{Tr}(A_1B_1A_2B_2 - A_1B_2A_2B_1).
$$

For $k > 2$ it has $\mathcal{N} = 6$ supersymmetry and $SU(4)_R \times U(1)_b$ global symmetry. Describes N M2-branes on $\mathbb{C}^4/\mathbb{Z}_k$.

 \bullet In the limit of fixed *k* and large *N*, the ABJM theory is dual to the M-theory background $\text{AdS}_4 \times S^7/\mathbb{Z}_k$

$$
(L/\ell_{\rm P})^6 \sim k\,N\,.
$$

At large *k* and fixed 't Hooft coupling $\lambda = N/k$ the theory is dual to type IIA string theory on $AdS_4 \times \mathbb{CP}^3$

$$
k g_{\rm st} = L/\ell_{\rm s} \sim \lambda^{1/4}.
$$

Perturbative type IIA string theory at large *k* and small g_{st} , i.e. fixed λ and large *N*.

$ABJM$ on S^3

The path integral on S^3 can be computed by supersymmetric localization and reduces to a matrix model [Kapustin-Willett-Yaakov]

$$
Z(N,k) = \frac{1}{N!^2} \int \frac{d^N \mu}{(2\pi)^N} \frac{d^N \nu}{(2\pi)^N} \exp\left[\frac{ik}{4\pi} \sum_{i=1}^N (\mu_i^2 - \nu_i^2)\right]
$$

$$
\frac{\Pi_{i
$$

$ABJM$ on S^3

The path integral on S^3 can be computed by supersymmetric localization and reduces to a matrix model [Kapustin-Willett-Yaakov]

$$
Z(N,k) = \frac{1}{N!^2} \int \frac{d^N \mu}{(2\pi)^N} \frac{d^N \nu}{(2\pi)^N} \exp\left[\frac{ik}{4\pi} \sum_{i=1}^N (\mu_i^2 - \nu_i^2)\right]
$$

$$
\frac{\Pi_{i
$$

Three methods have been used to study *Z*(*N, k*) at large *N*

- \blacksquare Map to CS theory on S^3/\mathbb{Z}_2 (or topological strings on $\mathbb{P}^1\times \mathbb{P}^1)$ and solve with large *N* techniques. Applies at large *N*, fixed *N/k*.[Drukker-Mariño-Putrov]
- 2 Study the large N limit at fixed k numerically. [Herzog-Klebanov-Pufu-Tesileanu]
- ³ Map the problem to a free Fermi gas on the real line with non-standard kinetic term. Valid at large *N* and finite *k*.[Mariño-Putrov]

ABJM at large *N* - An Airy tale

At large N and fixed k the S^3 partition function of the ABJM theory is [Mariño-Putrov], [Fuji-Hirano-Moriyama]

$$
Z_{S^3} = e^{\mathcal{A}(k)} C^{-\frac{1}{3}} \text{Ai}[C^{-\frac{1}{3}}(N-B)] + \mathcal{O}(e^{-\sqrt{N}}),
$$

with

$$
B = \frac{k}{24} + \frac{1}{3k}, \qquad C = \frac{2}{\pi^2 k},
$$

and

$$
\mathcal{A}(k) = \frac{2\zeta(3)}{\pi^2 k} \left(1 - \frac{k^3}{16} \right) + \frac{k^2}{\pi^2} \int_0^\infty \frac{x \log\left(1 - e^{-2x} \right)}{e^{kx} - 1} dx
$$

=
$$
- \frac{\zeta(3)}{8\pi^2} k^2 + 2\zeta'(-1) + \frac{1}{6} \log \frac{4\pi}{k} + \sum_{n \ge 2} \left(\frac{2\pi}{k} \right)^{2n-2} \frac{(-4)^{n-1} B_{2n} B_{2n-2}}{n(2n-2)(2n-2)!}.
$$

ABJM at large *N* - An Airy tale

At large N and fixed k the S^3 partition function of the ABJM theory is [Mariño-Putrov], [Fuji-Hirano-Moriyama]

$$
Z_{S^3} = e^{\mathcal{A}(k)} C^{-\frac{1}{3}} \text{Ai}[C^{-\frac{1}{3}}(N-B)] + \mathcal{O}(e^{-\sqrt{N}}),
$$

with

$$
B = \frac{k}{24} + \frac{1}{3k}, \qquad C = \frac{2}{\pi^2 k},
$$

and

$$
\mathcal{A}(k) = \frac{2\zeta(3)}{\pi^2 k} \left(1 - \frac{k^3}{16} \right) + \frac{k^2}{\pi^2} \int_0^\infty \frac{x \log(1 - e^{-2x})}{e^{kx} - 1} dx
$$

= $-\frac{\zeta(3)}{8\pi^2} k^2 + 2\zeta'(-1) + \frac{1}{6} \log \frac{4\pi}{k} + \sum_{n \ge 2} \left(\frac{2\pi}{k} \right)^{2n-2} \frac{(-4)^{n-1} B_{2n} B_{2n-2}}{n(2n-2)(2n-2)!}.$

The large *N* expansion takes the explicit form

$$
-\log Z_{S^3} = \frac{2}{3\sqrt{C}} N^{\frac{3}{2}} - \frac{B}{\sqrt{C}} N^{\frac{1}{2}} + \frac{1}{4} \log N - \mathcal{A}(k) + \frac{1}{4} \log \frac{32}{k} + \mathcal{O}(N^{-\frac{1}{2}}).
$$

ABJM at large *N* - An Airy tale

This can be reorganized *à la 't Hooft* into a type IIA string theory expansion

$$
F_{S^3} = -\sum_{g\geq 0} (2\pi i \lambda)^{2g-2} F_g(\lambda) N^{2-2g}.
$$

The genus g type IIA free energies can be computed systematically (up to $e^{-\sqrt{\lambda}}$ corrections) and read (agrees with topological string results)

$$
F_0(\lambda) = \frac{4\pi^3\sqrt{2}}{3}\hat{\lambda}^{\frac{3}{2}} + \frac{\zeta(3)}{2},
$$

\n
$$
F_1(\lambda) = \frac{\pi}{3\sqrt{2}}\hat{\lambda}^{\frac{1}{2}} - \frac{1}{4}\log\hat{\lambda} + \frac{1}{6}\log\lambda + \frac{1}{12}\log\frac{\pi^2}{32} + 2\zeta'(-1) - \frac{1}{2}\log 2,
$$

\n
$$
F_2(\lambda) = \frac{5\hat{\lambda}^{-\frac{3}{2}}}{96\pi^3\sqrt{2}} - \frac{\hat{\lambda}^{-1}}{48\pi^2} + \frac{\hat{\lambda}^{-\frac{1}{2}}}{144\pi\sqrt{2}} - \frac{1}{360},
$$

\n
$$
F_3(\lambda) = \frac{5\hat{\lambda}^{-3}}{512\pi^6} - \frac{5\hat{\lambda}^{-\frac{5}{2}}}{768\pi^5\sqrt{2}} + \frac{\hat{\lambda}^{-2}}{1152\pi^4} - \frac{\hat{\lambda}^{-\frac{3}{2}}}{10368\pi^3\sqrt{2}} - \frac{1}{22680},
$$

where

$$
\hat{\lambda} = \lambda - \frac{1}{24} \, .
$$

These results are prime targets for string/M-theory and AdS_4 holography!

$$
-\log Z_{S^3} = \frac{\pi \sqrt{2k}}{3} \, N^{\frac{3}{2}} - \frac{\pi}{3\sqrt{2}} \left(\frac{1}{8} k^{\frac{3}{2}} + k^{-\frac{1}{2}} \right) \, N^{\frac{1}{2}} + \frac{1}{4} \log N + \dots.
$$

These results are prime targets for string/M-theory and AdS₄ holography!

$$
-\log Z_{S^3} = \frac{\pi\sqrt{2k}}{3} N^{\frac{3}{2}} - \frac{\pi}{3\sqrt{2}} \left(\frac{1}{8}k^{\frac{3}{2}} + k^{-\frac{1}{2}}\right) N^{\frac{1}{2}} + \frac{1}{4}\log N + \dots
$$

From the on-shell action of the $AdS_4 \times X_7$ supergravity solution one finds

$$
F_{S^3} = -\log Z_{S^3} = \sqrt{\frac{2\pi^6}{27 \text{vol}(X_7)}} N^{\frac{3}{2}}.
$$

Plug in $\text{vol}(S^7/\mathbb{Z}_k) = \frac{\pi^4}{3k}$ $\frac{\pi^2}{3k}$ to find a match with the localization result.

These results are prime targets for string/M-theory and $AdS₄$ holography!

$$
-\log Z_{S^3} = \frac{\pi\sqrt{2k}}{3} N^{\frac{3}{2}} - \frac{\pi}{3\sqrt{2}} \left(\frac{1}{8}k^{\frac{3}{2}} + k^{-\frac{1}{2}}\right) N^{\frac{1}{2}} + \frac{1}{4}\log N + \dots
$$

From the on-shell action of the $AdS_4 \times X_7$ supergravity solution one finds

$$
F_{S^3} = -\log Z_{S^3} = \sqrt{\frac{2\pi^6}{27 \text{vol}(X_7)}} N^{\frac{3}{2}}.
$$

Plug in $\text{vol}(S^7/\mathbb{Z}_k) = \frac{\pi^4}{3k}$ $\frac{\pi^2}{3k}$ to find a match with the localization result.

The $N^{\frac{1}{2}}$ term can be derived using the on-shell action of four-derivative gauged supergravity.[NPB-Charles-Hristov-Reys]

These results are prime targets for string/M-theory and $AdS₄$ holography!

$$
-\log Z_{S^3} = \frac{\pi\sqrt{2k}}{3} N^{\frac{3}{2}} - \frac{\pi}{3\sqrt{2}} \left(\frac{1}{8}k^{\frac{3}{2}} + k^{-\frac{1}{2}}\right) N^{\frac{1}{2}} + \frac{1}{4}\log N + \dots
$$

From the on-shell action of the $AdS_4 \times X_7$ supergravity solution one finds

$$
F_{S^3} = -\log Z_{S^3} = \sqrt{\frac{2\pi^6}{27 \text{vol}(X_7)}} N^{\frac{3}{2}}.
$$

Plug in $\text{vol}(S^7/\mathbb{Z}_k) = \frac{\pi^4}{3k}$ $\frac{\pi^2}{3k}$ to find a match with the localization result.

The $N^{\frac{1}{2}}$ term can be derived using the on-shell action of four-derivative gauged supergravity.[NPB-Charles-Hristov-Reys]

The log *N* term is obtained by summing over the KK modes around $\mathsf{AdS}_4\times S^7$. [Bhattacharyya-Grassi-Mariño-Sen]

These results are prime targets for string/M-theory and $AdS₄$ holography!

$$
-\log Z_{S^3} = \frac{\pi\sqrt{2k}}{3} N^{\frac{3}{2}} - \frac{\pi}{3\sqrt{2}} \left(\frac{1}{8}k^{\frac{3}{2}} + k^{-\frac{1}{2}}\right) N^{\frac{1}{2}} + \frac{1}{4}\log N + \dots
$$

From the on-shell action of the $AdS_4 \times X_7$ supergravity solution one finds

$$
F_{S^3} = -\log Z_{S^3} = \sqrt{\frac{2\pi^6}{27 \text{vol}(X_7)}} N^{\frac{3}{2}}.
$$

Plug in $\text{vol}(S^7/\mathbb{Z}_k) = \frac{\pi^4}{3k}$ $\frac{\pi^2}{3k}$ to find a match with the localization result.

The $N^{\frac{1}{2}}$ term can be derived using the on-shell action of four-derivative gauged supergravity.[NPB-Charles-Hristov-Reys]

The log *N* term is obtained by summing over the KK modes around $\mathsf{AdS}_4\times S^7$. [Bhattacharyya-Grassi-Mariño-Sen]

Note: Derive the full Airy function using supersymmetric localization in supergravity on AdS₄?[Dabholkar-Drukker-Gomes]

Natural to consider deformations of the theory that preserve supersymmetry and break conformal invariance.

Natural to consider deformations of the theory that preserve supersymmetry and break conformal invariance.

The ABJM S^3 partition function with a $\mathrm{U}(1)\times\mathrm{U}(1)$ invariant squashing and real mass deformation takes the form [NPB-Hong-Reys], [Nosaka], [Hatsuda], [Hristov],

[Chester-Kalloor-Sharon], [Minahan-Naseer-Thull]

$$
Z_{S^3}(N, k, \Delta, b) = e^{\mathcal{A}(k, \Delta, b)} C_k^{-\frac{1}{3}} \text{Ai}[C_k^{-\frac{1}{3}}(N - B_k)] + \mathcal{O}(e^{-\sqrt{N}})
$$

with

$$
C_k = \frac{2}{\pi^2 k} \frac{(b+b^{-1})^{-4}}{\prod_{a=1}^4 \Delta_a}, \quad B_k = \frac{k}{24} - \frac{1}{12k} \sum_{a=1}^4 \frac{1}{\Delta_a} + \frac{1 - \frac{1}{4} \sum_a \Delta_a^2}{3k(b+b^{-1})^2 \prod_{a=1}^4 \Delta_a},
$$

and

$$
\Delta_1 = \frac{1}{2} - i \frac{m_1 + m_2 + m_3}{b + b^{-1}}, \ \Delta_2 = \frac{1}{2} - i \frac{m_1 - m_2 - m_3}{b + b^{-1}},
$$

$$
\Delta_3 = \frac{1}{2} + i \frac{m_1 + m_2 - m_3}{b + b^{-1}}, \ \Delta_4 = \frac{1}{2} + i \frac{m_1 - m_2 + m_3}{b + b^{-1}},
$$

such that $\sum_a \Delta_a = 2$.

This result encodes integrated correlation functions of the ABJM theory on $\mathbb{R}^3.$

Expand at large *N* and use holography to constrain/compute the higher-derivative corrections to type II string theory and M-theory. [Chester-Pufu-Yin], [Binder-Chester-Pufu], ...

This result encodes integrated correlation functions of the ABJM theory on $\mathbb{R}^3.$

Expand at large *N* and use holography to constrain/compute the higher-derivative corrections to type II string theory and M-theory. [Chester-Pufu-Yin], [Binder-Chester-Pufu], ...

Example:

Squashed S^3 partition function

$$
F_{S_b^3} = \frac{\pi\sqrt{2k}}{12} \left[(b+b^{-1}) \left[N^{\frac{3}{2}} + \left(\frac{1}{k} - \frac{k}{16} \right) N^{\frac{1}{2}} \right] - \frac{6}{k} N^{\frac{1}{2}} \right] + \frac{1}{4} \log N + \mathcal{O}(N^0).
$$

This captures integrated correlators of *Tµν* for the ABJM theory in flat space!

$$
C_T = \frac{32}{\pi^2} \left(\frac{\partial^2 F_{S_b^3}}{\partial b^2} \right)_{b=1} = \frac{64\sqrt{2k}}{3\pi} N^{\frac{3}{2}} + \frac{4(16 - k^2)\sqrt{2}}{3\pi\sqrt{k}} N^{\frac{1}{2}} + \mathcal{O}(N^0) \,.
$$

where

$$
\langle T_{\mu\nu} T_{\rho\sigma} \rangle = \frac{C_T}{(48\pi)^2} \left(P_{\mu\rho} P_{\nu\sigma} + P_{\nu\rho} P_{\mu\sigma} - P_{\mu\nu} P_{\rho\sigma} \right) \frac{1}{\vec{x}^2} \,, \quad P_{\mu\nu} \equiv \delta_{\mu\nu} \partial^2 - \partial_\mu \partial_\nu \,.
$$

What about the ABJM partition function on other 3-manifolds?

The ABJM topologically twisted index

TTI

The topologically twisted index (TTI) is a partition function of a 3d $\mathcal{N}=2$ SCFT on $S^1 \times \Sigma_\mathfrak{g}$. Supersymmetry is preserved by Witten's topological twist on $\Sigma_{\mathfrak{g}}$. The 3d QFT is not topological. Using supersymmetric localization the path integral can be reduced to a matrix integral.[Benini-Zaffaroni], [Closset-Kim]

TTI

The topologically twisted index (TTI) is a partition function of a 3d $\mathcal{N}=2$ SCFT on $S^1 \times \Sigma_\mathfrak{g}$. Supersymmetry is preserved by Witten's topological twist on $\Sigma_{\mathfrak{g}}$. The 3d QFT is not topological. Using supersymmetric localization the path integral can be reduced to a matrix integral.[Benini-Zaffaroni], [Closset-Kim]

For the ABJM theory the result is

$$
Z_{S^1 \times \Sigma_{\mathfrak{g}}}(N, k, \Delta, \mathfrak{n}) = \frac{1}{(N!)^2} \sum_{\mathfrak{m}, \tilde{\mathfrak{m}} \in \mathbb{Z}_N} \oint_C \prod_{i=1}^N \frac{dx_i}{2\pi i x_i} \prod_{j=1}^N \frac{d\tilde{x}_j}{2\pi i \tilde{x}_j} \prod_{i=1}^N x_i^{k \mathfrak{m}_i} \prod_{j=1}^N \tilde{x}_j^{-k \tilde{\mathfrak{m}}_j}
$$

$$
\times (\det \mathbb{B}(N, k, x, \tilde{x}, \Delta))^{\mathfrak{g}} \times \prod_{i \neq j}^N \left(1 - \frac{x_i}{x_j}\right)^{1 - \mathfrak{g}} \left(1 - \frac{\tilde{x}_i}{\tilde{x}_j}\right)^{1 - \mathfrak{g}}
$$

$$
\times \prod_{i,j=1}^N \prod_{a=1,2} \left(\frac{\sqrt{\frac{x_i}{\tilde{x}_j} y_a}}{1 - \frac{x_i}{\tilde{x}_j} y_a}\right)^{\mathfrak{m}_i - \tilde{\mathfrak{m}}_j + 1 - \mathfrak{g} - \mathfrak{n}_a} \prod_{a=3,4} \left(\frac{\sqrt{\frac{\tilde{x}_j}{\tilde{x}_i} y_a}}{1 - \frac{\tilde{x}_i}{\tilde{x}_j} y_a}\right)^{-\mathfrak{m}_i + \tilde{\mathfrak{m}}_j + 1 - \mathfrak{g} - \mathfrak{n}_a}
$$

.

Here $y_a = e^{i\pi \Delta_a}$ and supersymmetry imposes

$$
\sum_{a=1}^4 \Delta_a = 2\,,\qquad \sum_{a=1}^4 \mathfrak{n}_a = 2(1-\mathfrak{g})\,.
$$

TTI

Using (subtle) contour integration the TTI can be rewritten as

$$
Z = \prod_{a=1}^4 y_a^{-\tfrac{N^2}{2}\, \mathfrak{n}_a} \sum_{\{x_i,\tilde{x}_j\}} \left[\frac{1}{\det \mathbb{B}}\, \frac{\prod_{i=1}^N x_i^N \tilde{x}_i^N \prod_{i \neq j}^N (1-\frac{x_i}{x_j})(1-\frac{\tilde{x}_i}{\tilde{x}_j})}{\prod_{i,j=1}^N \prod_{a=1}^2 (\tilde{x}_j-x_i y_a)^{1-\tfrac{\mathfrak{n}_a}{1-\mathfrak{g}}} \prod_{a=3}^4 (x_i-\tilde{x}_j y_a)^{1-\tfrac{\mathfrak{n}_a}{1-\mathfrak{g}}} } \right]^{1-\mathfrak{g}}
$$

Where x_i and \tilde{x}_i are solutions to the following "Bethe Ansatz Equations"

$$
\begin{aligned} e^{\mathrm{i} B_i} &\equiv x_i^k \prod_{j=1}^N \frac{(1-y_3 \frac{\tilde{x}_j}{x_i})(1-y_4 \frac{\tilde{x}_j}{x_i})}{(1-y_1^{-1} \frac{\tilde{x}_j}{x_i})(1-y_2^{-1} \frac{\tilde{x}_j}{x_i})} = 1\,,\\ e^{\mathrm{i} \tilde{B}_j} &\equiv \tilde{x}_j^k \prod_{i=1}^N \frac{(1-y_3 \frac{\tilde{x}_j}{x_i})(1-y_4 \frac{\tilde{x}_j}{x_i})}{(1-y_1^{-1} \frac{\tilde{x}_j}{x_i})(1-y_2^{-1} \frac{\tilde{x}_j}{x_i})} = 1\,, \end{aligned}
$$

and the Jacobian matrix $\mathbb B$ is given by

$$
\mathbb{B} = \frac{\partial (e^{iB_1}, \cdots, e^{iB_N}, e^{i\tilde{B}_1}, \cdots, e^{i\tilde{B}_N})}{\partial (\log x_1, \cdots, \log x_N, \log \tilde{x}_1, \cdots, \log \tilde{x}_N)}.
$$

TTI at large N

The BAE solution in the large *N* limit takes the form [Benini-Hristov-Zaffaroni]

$$
\log x_i = N^{\frac{1}{2}} t_i - \mathrm{i} v_i \,, \quad \log \tilde{x}_j = N^{\frac{1}{2}} t_j - \mathrm{i} \tilde{v}_j \,.
$$

TTI at large N

The BAE solution in the large *N* limit takes the form [Benini-Hristov-Zaffaroni]

$$
\log x_i = N^{\frac{1}{2}} t_i - \mathrm{i} v_i \,, \quad \log \tilde{x}_j = N^{\frac{1}{2}} t_j - \mathrm{i} \tilde{v}_j \,.
$$

Our approach: Use this solution as a starting point to numerically solve the BAE and calculate the index. The numerical results are very precise and led us to an analytic form for the TTI valid to all orders in 1*/N*!

TTI at large N

The BAE solution in the large *N* limit takes the form [Benini-Hristov-Zaffaroni]

$$
\log x_i = N^{\frac{1}{2}} t_i - \mathrm{i} v_i \,, \quad \log \tilde{x}_j = N^{\frac{1}{2}} t_j - \mathrm{i} \tilde{v}_j \,.
$$

Our approach: Use this solution as a starting point to numerically solve the BAE and calculate the index. The numerical results are very precise and led us to an analytic form for the TTI valid to all orders in 1*/N*!

To write the result compactly define

$$
\hat{N}_{\Delta} \equiv N - \frac{k}{24} + \frac{1}{12k} \sum_{a=1}^{4} \frac{1}{\Delta_a},
$$

in terms of which $F_{S^1\times \Sigma_\mathfrak{g}}=-\log Z_{S^1\times \Sigma_\mathfrak{g}}$, takes the simple form:

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} = \frac{\pi \sqrt{2k \Delta_1 \Delta_2 \Delta_3 \Delta_4}}{3} \sum_{a=1}^4 \frac{\mathfrak{n}_a}{\Delta_a} \left(\hat{N}_{\Delta}^{\frac{3}{2}} - \frac{\mathfrak{c}_a}{k} \hat{N}_{\Delta}^{\frac{1}{2}} \right) + \frac{1-\mathfrak{g}}{2} \log \hat{N}_{\Delta} - \hat{f}_0(k, \Delta, \mathfrak{n}),
$$

where c*^a* are given by

$$
\mathfrak{c}_a = \frac{\prod_{b \neq a} (\Delta_a + \Delta_b)}{8 \Delta_1 \Delta_2 \Delta_3 \Delta_4} \sum_{b \neq a} \Delta_b.
$$

The universal index is defined by setting $\Delta_a = \frac{1}{2}$ and $\mathfrak{n}_a = \frac{1-\mathfrak{g}}{2}$. We then define $\hat{N}=N-\frac{k}{24}+\frac{2}{3k}$ to obtain

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1-\mathfrak{g}} = \frac{\pi \sqrt{2k}}{3} \left(\hat{N}^{\frac{3}{2}} - \frac{3}{k} \hat{N}^{\frac{1}{2}} \right) + \frac{1}{2} \log \hat{N} - \hat{f}_0(k) \,.
$$

The universal index is defined by setting $\Delta_a = \frac{1}{2}$ and $\mathfrak{n}_a = \frac{1-\mathfrak{g}}{2}$. We then define $\hat{N}=N-\frac{k}{24}+\frac{2}{3k}$ to obtain

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1-\mathfrak{g}} = \frac{\pi \sqrt{2k}}{3} \left(\hat{N}^{\frac{3}{2}} - \frac{3}{k} \hat{N}^{\frac{1}{2}} \right) + \frac{1}{2} \log \hat{N} - \hat{f}_0(k) \,.
$$

No closed form expression for $\hat{f}_0(k)$ but at large *k* we find

$$
\hat{f}_0(k) = -\frac{3\zeta(3)}{8\pi^2}k^2 + \frac{7}{6}\log k + \mathfrak{f}_0 + \sum_{n=1}^5 \left(\frac{2\pi}{k}\right)^{2n} \frac{\mathfrak{f}_{2n}}{3^{n+2}} + \mathcal{O}(k^{-12}),
$$

with
$$
\{f_{2n}\} = \left\{-\frac{6}{5}, \frac{19}{70}, -\frac{41}{175}, \frac{279}{700}, -\frac{964636}{875875}\right\}
$$
 and $f_0 = -2.096848299$.

The universal index is defined by setting $\Delta_a = \frac{1}{2}$ and $\mathfrak{n}_a = \frac{1-\mathfrak{g}}{2}$. We then define $\hat{N}=N-\frac{k}{24}+\frac{2}{3k}$ to obtain

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1 - \mathfrak{g}} = \frac{\pi \sqrt{2k}}{3} \left(\hat{N}^{\frac{3}{2}} - \frac{3}{k} \hat{N}^{\frac{1}{2}} \right) + \frac{1}{2} \log \hat{N} - \hat{f}_0(k) .
$$

No closed form expression for $\hat{f}_0(k)$ but at large k we find

$$
\hat{f}_0(k) = -\frac{3\zeta(3)}{8\pi^2}k^2 + \frac{7}{6}\log k + \mathfrak{f}_0 + \sum_{n=1}^5 \left(\frac{2\pi}{k}\right)^{2n} \frac{\mathfrak{f}_{2n}}{3^{n+2}} + \mathcal{O}(k^{-12}),
$$

with $\{f_{2n}\}=\left\{-\frac{6}{5},\frac{19}{70},-\frac{41}{175},\frac{279}{700},-\frac{964636}{875875}\right\}$ and $\mathfrak{f}_0=-2.096848299$. For low values of k , $\hat f_0(k)$ can be determined numerically with very good precision:

$$
\hat{f}_0(1) = -3.045951311
$$
, $\hat{f}_0(2) = -1.786597534$,
\n $\hat{f}_0(3) = -1.386373044$, $\hat{f}_0(4) = -1.306589553$.

The universal index is defined by setting $\Delta_a = \frac{1}{2}$ and $\mathfrak{n}_a = \frac{1-\mathfrak{g}}{2}$. We then define $\hat{N}=N-\frac{k}{24}+\frac{2}{3k}$ to obtain

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1 - \mathfrak{g}} = \frac{\pi \sqrt{2k}}{3} \left(\hat{N}^{\frac{3}{2}} - \frac{3}{k} \hat{N}^{\frac{1}{2}} \right) + \frac{1}{2} \log \hat{N} - \hat{f}_0(k) .
$$

No closed form expression for $\hat{f}_0(k)$ but at large k we find

$$
\hat{f}_0(k) = -\frac{3\zeta(3)}{8\pi^2}k^2 + \frac{7}{6}\log k + \mathfrak{f}_0 + \sum_{n=1}^5 \left(\frac{2\pi}{k}\right)^{2n} \frac{\mathfrak{f}_{2n}}{3^{n+2}} + \mathcal{O}(k^{-12}),
$$

with $\{f_{2n}\}=\left\{-\frac{6}{5},\frac{19}{70},-\frac{41}{175},\frac{279}{700},-\frac{964636}{875875}\right\}$ and $\mathfrak{f}_0=-2.096848299$. For low values of k , $\hat f_0(k)$ can be determined numerically with very good precision:

$$
\hat{f}_0(1) = -3.045951311
$$
, $\hat{f}_0(2) = -1.786597534$,
\n $\hat{f}_0(3) = -1.386373044$, $\hat{f}_0(4) = -1.306589553$.

We have checked these results with ${\rm ext}$ extensive numerical calculations to great accuracy. They are exact up to $e^{-\sqrt{N}}$ corrections.

This result can be reorganized as a type IIA string theory expansion

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} = -\sum_{g \ge 0} (2\pi i \lambda)^{2g-2} F_g(\lambda) N^{2-2g}.
$$

For low genera we find

$$
\frac{F_0(\lambda)}{1 - \mathfrak{g}} = \frac{4\pi^3 \sqrt{2}}{3} \hat{\lambda}^{\frac{3}{2}} + \frac{3\zeta(3)}{2},
$$

\n
$$
\frac{F_1(\lambda)}{1 - \mathfrak{g}} = \frac{2\pi \sqrt{2}}{3} \hat{\lambda}^{\frac{1}{2}} - \frac{1}{2} \log \hat{\lambda} - \frac{2}{3} \log \lambda + \mathfrak{h},
$$

\n
$$
\frac{F_2(\lambda)}{1 - \mathfrak{g}} = \frac{\hat{\lambda}^{-1}}{12\pi^2} - \frac{5\hat{\lambda}^{-\frac{1}{2}}}{36\sqrt{2}\pi} + \frac{2}{45},
$$

\n
$$
\frac{F_3(\lambda)}{1 - \mathfrak{g}} = \frac{\hat{\lambda}^{-2}}{144\pi^4} - \frac{\hat{\lambda}^{-\frac{3}{2}}}{162\sqrt{2}\pi^3} + \frac{19}{5670}.
$$

where $\hat{\lambda} = \lambda - \frac{1}{24}$.

This result can be reorganized as a type IIA string theory expansion

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} = -\sum_{\mathsf{g} \geq 0} (2\pi i \lambda)^{2\mathsf{g}-2} F_{\mathsf{g}}(\lambda) N^{2-2\mathsf{g}}.
$$

For low genera we find

$$
\frac{F_0(\lambda)}{1-\mathfrak{g}} = \frac{4\pi^3\sqrt{2}}{3}\hat{\lambda}^{\frac{3}{2}} + \frac{3\zeta(3)}{2},
$$

\n
$$
\frac{F_1(\lambda)}{1-\mathfrak{g}} = \frac{2\pi\sqrt{2}}{3}\hat{\lambda}^{\frac{1}{2}} - \frac{1}{2}\log\hat{\lambda} - \frac{2}{3}\log\lambda + \mathfrak{f}_0,
$$

\n
$$
\frac{F_2(\lambda)}{1-\mathfrak{g}} = \frac{\hat{\lambda}^{-1}}{12\pi^2} - \frac{5\hat{\lambda}^{-\frac{1}{2}}}{36\sqrt{2}\pi} + \frac{2}{45},
$$

\n
$$
\frac{F_3(\lambda)}{1-\mathfrak{g}} = \frac{\hat{\lambda}^{-2}}{144\pi^4} - \frac{\hat{\lambda}^{-\frac{3}{2}}}{162\sqrt{2}\pi^3} + \frac{19}{5670}.
$$

where $\hat{\lambda} = \lambda - \frac{1}{24}$.

How can we derive this from type IIA string theory?

Holography and black holes

The following is a supersymmetric Euclidean solution of 4d $\mathcal{N}=2$ gauged supergravity [Romans], [Benetti Genolini-Ipiña-Sparks], [NPB-Charles-Min], ...

$$
ds_4^2 = U(r)d\tau^2 + \frac{dr^2}{U(r)} + r^2 ds_{\Sigma_{\mathfrak{g}}}^2, \qquad F = \frac{Q}{r^2}d\tau \wedge dr - \frac{\kappa}{g}\text{vol}(\Sigma_{\mathfrak{g}}),
$$

$$
U(r) = \left(\sqrt{2}gr + \frac{\kappa}{2\sqrt{2}gr}\right)^2 - \frac{Q^2}{8r^2}, \qquad \kappa = 1, 0, -1.
$$

The following is a supersymmetric Euclidean solution of 4d $\mathcal{N}=2$ gauged supergravity [Romans], [Benetti Genolini-Ipiña-Sparks], [NPB-Charles-Min], ...

$$
ds_4^2 = U(r)d\tau^2 + \frac{dr^2}{U(r)} + r^2 ds_{\Sigma_{\mathfrak{g}}}^2, \qquad F = \frac{Q}{r^2}d\tau \wedge dr - \frac{\kappa}{g}\text{vol}(\Sigma_{\mathfrak{g}}),
$$

$$
U(r) = \left(\sqrt{2}gr + \frac{\kappa}{2\sqrt{2}gr}\right)^2 - \frac{Q^2}{8r^2}, \qquad \kappa = 1, 0, -1.
$$

The solution is smooth for $q|Q| > \kappa$. The periodicity of τ is

$$
\beta_{\tau} = \frac{\pi\sqrt{-\kappa + g|Q|}}{g^2|Q|}
$$

.

The following is a supersymmetric Euclidean solution of 4d $\mathcal{N}=2$ gauged supergravity [Romans], [Benetti Genolini-Ipiña-Sparks], [NPB-Charles-Min], ...

$$
ds_4^2 = U(r)d\tau^2 + \frac{dr^2}{U(r)} + r^2 ds_{\Sigma_{\mathfrak{g}}}^2, \qquad F = \frac{Q}{r^2}d\tau \wedge dr - \frac{\kappa}{g}\text{vol}(\Sigma_{\mathfrak{g}}),
$$

$$
U(r) = \left(\sqrt{2}gr + \frac{\kappa}{2\sqrt{2}gr}\right)^2 - \frac{Q^2}{8r^2}, \qquad \kappa = 1, 0, -1.
$$

The solution is smooth for $g|Q| > \kappa$. The periodicity of τ is

$$
\beta_{\tau} = \frac{\pi\sqrt{-\kappa + g|Q|}}{g^2|Q|}
$$

.

The smooth Lorentzian black hole is obtained by taking $Q \rightarrow 0$ and exists only for $\kappa = -1$, i.e. $g > 1$. The regularized on-shell action for any β_{τ} is

$$
I=-\frac{\pi}{4g^2G_N}(\mathfrak{g}-1)\,.
$$

For the black hole solution this leads to $S_{\text{BH}} = -I$. [Gibbons-Hawking],...

The following is a supersymmetric Euclidean solution of 4d $\mathcal{N}=2$ gauged supergravity [Romans], [Benetti Genolini-Ipiña-Sparks], [NPB-Charles-Min], ...

$$
ds_4^2 = U(r)d\tau^2 + \frac{dr^2}{U(r)} + r^2 ds_{\Sigma_\mathfrak{g}}^2, \qquad F = \frac{Q}{r^2}d\tau \wedge dr - \frac{\kappa}{g}\text{vol}(\Sigma_\mathfrak{g}),
$$

$$
U(r) = \left(\sqrt{2}gr + \frac{\kappa}{2\sqrt{2}gr}\right)^2 - \frac{Q^2}{8r^2}, \qquad \kappa = 1, 0, -1.
$$

The solution is smooth for $g|Q| > \kappa$. The periodicity of τ is

$$
\beta_{\tau} = \frac{\pi\sqrt{-\kappa + g|Q|}}{g^2|Q|}
$$

.

The smooth Lorentzian black hole is obtained by taking $Q \rightarrow 0$ and exists only for $\kappa = -1$, i.e. $\mathfrak{g} > 1$. The regularized on-shell action for any β_{τ} is

$$
I=-\frac{\pi}{4g^2G_N}(\mathfrak{g}-1)\,.
$$

For the black hole solution this leads to $S_{\text{BH}} = -I$. [Gibbons-Hawking],...

This is the extremal magnetic Reissner-Nordström black hole in $AdS₄!$

An 11d uplift of this solution is given by

$$
ds_{11}^2 = \frac{1}{4}ds_4^2 + ds_{\mathbb{CP}^3}^2 + \left(d\psi + \sigma_{\mathbb{CP}^3} + \frac{1}{4}A\right)^2,
$$

\n
$$
G_4 = \frac{3}{8}\text{vol}_4 - \frac{1}{4} \star_4 F \wedge J_{\mathbb{CP}^3}, \quad \text{with} \quad d\sigma_{\mathbb{CP}^3} = 2J_{\mathbb{CP}^3}.
$$

Our QFT result for the TTI amounts to a prediction for the path integral of M-theory on this background to all orders in the 1*/N* expansion.

An 11d uplift of this solution is given by

$$
ds_{11}^2 = \frac{1}{4}ds_4^2 + ds_{\mathbb{CP}^3}^2 + \left(d\psi + \sigma_{\mathbb{CP}^3} + \frac{1}{4}A\right)^2,
$$

\n
$$
G_4 = \frac{3}{8}\text{vol}_4 - \frac{1}{4} \star_4 F \wedge J_{\mathbb{CP}^3}, \quad \text{with} \quad d\sigma_{\mathbb{CP}^3} = 2J_{\mathbb{CP}^3}.
$$

Our QFT result for the TTI amounts to a prediction for the path integral of M-theory on this background to all orders in the 1*/N* expansion.

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} = (1 - \mathfrak{g}) \frac{\pi \sqrt{2k}}{3} \left(N^{\frac{3}{2}} - \frac{32 + k^2}{16k} N^{\frac{1}{2}} \right) + \frac{(1 - \mathfrak{g})}{2} \log N + \dots
$$

The $N^{\frac{3}{2}}$ term comes from the 2-der on-shell action.[Benini-Hristov-Zaffaroni], [Azzurli-NPB-Crichigno-Min-Zaffaroni]

An 11d uplift of this solution is given by

$$
ds_{11}^2 = \frac{1}{4}ds_4^2 + ds_{\mathbb{CP}^3}^2 + \left(d\psi + \sigma_{\mathbb{CP}^3} + \frac{1}{4}A\right)^2,
$$

\n
$$
G_4 = \frac{3}{8}\text{vol}_4 - \frac{1}{4} \star_4 F \wedge J_{\mathbb{CP}^3}, \quad \text{with} \quad d\sigma_{\mathbb{CP}^3} = 2J_{\mathbb{CP}^3}.
$$

Our QFT result for the TTI amounts to a prediction for the path integral of M-theory on this background to all orders in the 1*/N* expansion.

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} = (1 - \mathfrak{g}) \frac{\pi \sqrt{2k}}{3} \left(N^{\frac{3}{2}} - \frac{32 + k^2}{16k} N^{\frac{1}{2}} \right) + \frac{(1 - \mathfrak{g})}{2} \log N + \dots
$$

The $N^{\frac{3}{2}}$ term comes from the 2-der on-shell action.[Benini-Hristov-Zaffaroni], [Azzurli-NPB-Crichigno-Min-Zaffaroni]

The $N^{\frac{1}{2}}$ term comes from the 4-der supergravity on-shell action. [NPB-Charles-Hristov-Reys]

An 11d uplift of this solution is given by

$$
ds_{11}^2 = \frac{1}{4}ds_4^2 + ds_{\mathbb{CP}^3}^2 + \left(d\psi + \sigma_{\mathbb{CP}^3} + \frac{1}{4}A\right)^2,
$$

\n
$$
G_4 = \frac{3}{8}\text{vol}_4 - \frac{1}{4} \star_4 F \wedge J_{\mathbb{CP}^3}, \quad \text{with} \quad d\sigma_{\mathbb{CP}^3} = 2J_{\mathbb{CP}^3}.
$$

Our QFT result for the TTI amounts to a prediction for the path integral of M-theory on this background to all orders in the 1*/N* expansion.

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} = (1 - \mathfrak{g}) \frac{\pi \sqrt{2k}}{3} \left(N^{\frac{3}{2}} - \frac{32 + k^2}{16k} N^{\frac{1}{2}} \right) + \frac{(1 - \mathfrak{g})}{2} \log N + \dots
$$

The $N^{\frac{3}{2}}$ term comes from the 2-der on-shell action.[Benini-Hristov-Zaffaroni], [Azzurli-NPB-Crichigno-Min-Zaffaroni]

The $N^{\frac{1}{2}}$ term comes from the 4-der supergravity on-shell action. [NPB-Charles-Hristov-Reys]

The log *N* term comes from 1-loop contributions of the 11d KK modes. [Liu-Pando Zayas-Rathee-Zhao]

An 11d uplift of this solution is given by

$$
ds_{11}^2 = \frac{1}{4}ds_4^2 + ds_{\mathbb{CP}^3}^2 + \left(d\psi + \sigma_{\mathbb{CP}^3} + \frac{1}{4}A\right)^2,
$$

\n
$$
G_4 = \frac{3}{8}\text{vol}_4 - \frac{1}{4} \star_4 F \wedge J_{\mathbb{CP}^3}, \quad \text{with} \quad d\sigma_{\mathbb{CP}^3} = 2J_{\mathbb{CP}^3}.
$$

Our QFT result for the TTI amounts to a prediction for the path integral of M-theory on this background to all orders in the 1*/N* expansion.

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} = (1 - \mathfrak{g}) \frac{\pi \sqrt{2k}}{3} \left(N^{\frac{3}{2}} - \frac{32 + k^2}{16k} N^{\frac{1}{2}} \right) + \frac{(1 - \mathfrak{g})}{2} \log N + \dots
$$

The $N^{\frac{3}{2}}$ term comes from the 2-der on-shell action.[Benini-Hristov-Zaffaroni], [Azzurli-NPB-Crichigno-Min-Zaffaroni]

The $N^{\frac{1}{2}}$ term comes from the 4-der supergravity on-shell action. [NPB-Charles-Hristov-Reys]

The log *N* term comes from 1-loop contributions of the 11d KK modes. [Liu-Pando Zayas-Rathee-Zhao]

An all-order prediction for the entropy of this black hole?!?

There are more general supersymmetric Euclidean "black saddle" solutions in the STU model of 4d $\mathcal{N}=2$ supergravity (gravity $+3$ vector multiplets)

$$
ds_4^2 = e^{2f_1(r)}d\tau^2 + e^{2f_2(r)}dr^2 + e^{2f_3(r)}ds_{\Sigma_{\mathfrak{g}}}^2, \qquad A^I = v^I(r)d\tau + p^I\omega_{\Sigma_{\mathfrak{g}}} z_\alpha(r), \quad \tilde{z}_\alpha(r), \qquad \alpha = 1, 2, 3 \quad \text{and} \quad I = 0, 1, 2, 3.
$$

There are more general supersymmetric Euclidean "black saddle" solutions in the STU model of 4d $\mathcal{N} = 2$ supergravity (gravity $+3$ vector multiplets)

$$
ds_4^2 = e^{2f_1(r)}d\tau^2 + e^{2f_2(r)}dr^2 + e^{2f_3(r)}ds_{\Sigma_{\mathfrak{g}}}^2, \qquad A^I = v^I(r)d\tau + p^I\omega_{\Sigma_{\mathfrak{g}}}^2
$$

 $z_\alpha(r), \quad \tilde{z}_\alpha(r), \qquad \alpha = 1, 2, 3 \quad \text{and} \quad I = 0, 1, 2, 3.$

The on-shell action of these solutions agrees with the $N^{\frac{3}{2}}$ term in the TTI for general fugacities Δ_a and magnetic fluxes \mathfrak{n}_a . [NPB-Charles-Min]

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} \approx \frac{\pi \sqrt{2k \Delta_1 \Delta_2 \Delta_3 \Delta_4}}{3} \sum_{a=1}^4 \frac{\mathfrak{n}_a}{\Delta_a} N^{\frac{3}{2}} \,.
$$

There are more general supersymmetric Euclidean "black saddle" solutions in the STU model of 4d $\mathcal{N} = 2$ supergravity (gravity $+3$ vector multiplets)

$$
ds_4^2 = e^{2f_1(r)}d\tau^2 + e^{2f_2(r)}dr^2 + e^{2f_3(r)}ds_{\Sigma_{\mathfrak{g}}}^2, \qquad A^I = v^I(r)d\tau + p^I\omega_{\Sigma_{\mathfrak{g}}}^2
$$

 $z_\alpha(r), \quad \tilde{z}_\alpha(r), \qquad \alpha = 1, 2, 3 \quad \text{and} \quad I = 0, 1, 2, 3.$

The on-shell action of these solutions agrees with the $N^{\frac{3}{2}}$ term in the TTI for general fugacities Δ_a and magnetic fluxes \mathfrak{n}_a . [NPB-Charles-Min]

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} \approx \frac{\pi \sqrt{2k \Delta_1 \Delta_2 \Delta_3 \Delta_4}}{3} \sum_{a=1}^4 \frac{\mathfrak{n}_a}{\Delta_a} N^{\frac{3}{2}}.
$$

The exact TTI is a prediction for the string/M-theory path integral on these backgrounds!

There are more general supersymmetric Euclidean "black saddle" solutions in the STU model of 4d $\mathcal{N} = 2$ supergravity (gravity $+3$ vector multiplets)

$$
ds_4^2 = e^{2f_1(r)}d\tau^2 + e^{2f_2(r)}dr^2 + e^{2f_3(r)}ds_{\Sigma_{\mathfrak{g}}}^2, \qquad A^I = v^I(r)d\tau + p^I\omega_{\Sigma_{\mathfrak{g}}}^2
$$

 $z_\alpha(r), \quad \tilde{z}_\alpha(r), \qquad \alpha = 1, 2, 3 \quad \text{and} \quad I = 0, 1, 2, 3.$

The on-shell action of these solutions agrees with the $N^{\frac{3}{2}}$ term in the TTI for general fugacities Δ_a and magnetic fluxes \mathfrak{n}_a . [NPB-Charles-Min]

$$
F_{S^1 \times \Sigma_{\mathfrak{g}}} \approx \frac{\pi \sqrt{2k \Delta_1 \Delta_2 \Delta_3 \Delta_4}}{3} \sum_{a=1}^4 \frac{\mathfrak{n}_a}{\Delta_a} N^{\frac{3}{2}}.
$$

The exact TTI is a prediction for the string/M-theory path integral on these backgrounds!

Some of these Euclidean solutions admit Lorentzian interpretation as supersymmetric black holes with an $AdS_2 \times \Sigma_{\mathfrak{q}}$ near horizon limit.[Gauntlett-Kim-Pakis-Waldram], [Cacciatori-Klemm], [NPB-Min-Pilch], [NPB-Charles-Min]

The entropy of these black holes is computed by the TTI after a Legendre transform and I -extremization. [Benini-Hristov-Zaffaroni], [NPB-Min-Pilch], ...

Other 3d $\mathcal{N}=2$ SCFTs

3d $\mathcal{N} = 4$ SYM

There is another simple 3d $\mathcal{N} = 4$ holographic SCFT we can study with these tools - 3d $\mathcal{N} = 4$ SYM with $U(N)$ gauge theory with 1 adjoint and N_f fundamental hypermultiplets and no CS term.

3d $\mathcal{N} = 4$ SYM

There is another simple 3d $\mathcal{N} = 4$ holographic SCFT we can study with these tools - 3d $\mathcal{N} = 4$ SYM with $U(N)$ gauge theory with 1 adjoint and N_f fundamental hypermultiplets and no CS term.

The bulk dual is an $\mathsf{AdS}_4\times S^7/\mathbb{Z}_{N_f}$ solution of 11 d supergravity. The orbifold acts on \mathbb{C}^4 as

$$
(z_1, z_2, z_3, z_4) \rightarrow (z_1, z_2, e^{2\pi i/N_f} z_3, e^{2\pi i/N_f} z_4).
$$

3d $\mathcal{N} = 4$ SYM

There is another simple 3d $\mathcal{N}=4$ holographic SCFT we can study with these tools - 3d $\mathcal{N} = 4$ SYM with $U(N)$ gauge theory with 1 adjoint and N_f fundamental hypermultiplets and no CS term.

The bulk dual is an $\mathsf{AdS}_4\times S^7/\mathbb{Z}_{N_f}$ solution of 11 d supergravity. The orbifold acts on \mathbb{C}^4 as

$$
(z_1, z_2, z_3, z_4) \rightarrow (z_1, z_2, e^{2\pi i/N_f} z_3, e^{2\pi i/N_f} z_4).
$$

We can apply the same numerical method to compute the TTI to find

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1-\mathfrak{g}} = \frac{\pi \sqrt{2}}{3} N_f^{\frac{1}{2}} \left[\hat{N}^{\frac{3}{2}} - \left(\frac{N_f}{2} + \frac{5}{2N_f} \right) \hat{N}^{\frac{1}{2}} \right] + \frac{1}{2} \log \hat{N} - \hat{f}_0(N_f),
$$

where

$$
\hat{N} = N + \frac{7N_f}{24} + \frac{1}{3N_f} \, .
$$

There is a corresponding BPS black hole in M-theory for which this index computes the entropy.

Other examples

Similar results can be obtained for other 3d holographic SCFTs (no known Airy function on S^3 for most of these!)

• 3d
$$
\mathcal{N} = 2
$$
 mABJM theory. We have $\hat{N} = N + \frac{19}{24}$

$$
\frac{F_{S^1\times \Sigma_{\mathfrak{g}}}}{1-\mathfrak{g}} = \frac{4\pi\sqrt{2}}{9\sqrt{3}}\Big[\hat{N}^{\frac{3}{2}}-\frac{9}{2}\hat{N}^{\frac{1}{2}}\Big]+ \frac{1}{2}\log \hat{N}-\hat{f}_0\,,
$$

3d $\mathcal{N}=2$ $V^{5,2}$ theory. We have $\hat{N}=N+\frac{k}{6}+\frac{1}{4k}$ and

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1-\mathfrak{g}} = \frac{16\pi\sqrt{k}}{27} \left[\hat{N}^{\frac{3}{2}} - \left(\frac{9k}{16} + \frac{27}{16k} \right) \hat{N}^{\frac{1}{2}} \right] + \frac{1}{2} \log \hat{N} - \hat{f}_0(k) ,
$$

3d $\mathcal{N}=2$ Q^{111} theory. We have $\hat{N}=N+\frac{k}{6}$ and

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1 - \mathfrak{g}} = \frac{4\pi\sqrt{k}}{3\sqrt{3}} \left[\hat{N}^{\frac{3}{2}} - \left(\frac{k}{4} + \frac{3}{4k} \right) \hat{N}^{\frac{1}{2}} \right] + \frac{1}{2} \log \hat{N} - \hat{f}_0(k) ,
$$

3d $\mathcal{N}=3$ N^{010} theory. We have $\hat{N}=N+\frac{k}{12}+\frac{1}{3k}$ and

$$
\frac{F_{S^1 \times \Sigma_{\mathfrak{g}}}}{1 - \mathfrak{g}} = \frac{4\pi\sqrt{k}}{3\sqrt{3}} \left[\hat{N}^{\frac{3}{2}} - \left(\frac{k}{4} + \frac{5}{4k} \right) \hat{N}^{\frac{1}{2}} \right] + \frac{1}{2} \log \hat{N} - \hat{f}_0(k) ,
$$

In each of these cases there is a BPS black hole solution for which the index accounts for the entropy.

The superconformal index [with Sunjin Choi]

The superconformal index (SCI), or $S^1 \times_w S^2$ partition function, counts certain BPS operators in 3d $\mathcal{N} = 2$ SCFTs.

The superconformal index [with Sunjin Choi]

The superconformal index (SCI), or $S^1 \times_w S^2$ partition function, counts certain BPS operators in 3d $\mathcal{N}=2$ SCFTs.

It is useful to consider the Cardy-like limit *ω* → 0. The SCI can then be analyzed with similar tools as the TTI.

The superconformal index $\frac{1}{2}$ [with Sunjin Choi]

The superconformal index (SCI), or $S^1 \times_w S^2$ partition function, counts certain BPS operators in 3d $\mathcal{N}=2$ SCFTs.

It is useful to consider the Cardy-like limit *ω* → 0. The SCI can then be analyzed with similar tools as the TTI.

For the ABJM theory in the M-theory limit we find the following ω^{-1} and ω^0 results

$$
\log \mathcal{I}_{\text{ABJM}}(N, k, \omega) = -\frac{\pi \sqrt{2k}}{3} \left[\left(\frac{1}{2\omega} + 1 \right) \left(N - \frac{k}{24} + \frac{2}{3k} \right)^{\frac{3}{2}} - \frac{3}{k} \left(N - \frac{k}{24} + \frac{2}{3k} \right)^{\frac{1}{2}} \right] - \frac{2}{\omega} \hat{g}_0(k) - \frac{1}{2} \log \left(N - \frac{k}{24} + \frac{2}{3k} \right) + \hat{f}_0(k) + \mathcal{O}(e^{-\sqrt{N}}) + \mathcal{O}(\omega).
$$

The superconformal index $\frac{1}{2}$ [with Sunjin Choi]

The superconformal index (SCI), or $S^1 \times_w S^2$ partition function, counts certain BPS operators in 3d $\mathcal{N}=2$ SCFTs.

It is useful to consider the Cardy-like limit *ω* → 0. The SCI can then be analyzed with similar tools as the TTI.

For the ABJM theory in the M-theory limit we find the following ω^{-1} and ω^0 results

$$
\log \mathcal{I}_{\text{ABJM}}(N, k, \omega) = -\frac{\pi \sqrt{2k}}{3} \left[\left(\frac{1}{2\omega} + 1 \right) \left(N - \frac{k}{24} + \frac{2}{3k} \right)^{\frac{3}{2}} - \frac{3}{k} \left(N - \frac{k}{24} + \frac{2}{3k} \right)^{\frac{1}{2}} \right] - \frac{2}{\omega} \hat{g}_0(k) - \frac{1}{2} \log \left(N - \frac{k}{24} + \frac{2}{3k} \right) + \hat{f}_0(k) + \mathcal{O}(e^{-\sqrt{N}}) + \mathcal{O}(\omega).
$$

Captures the entropy of the dual supersymmetric $AdS₄$ Kerr-Newman black hole.

The superconformal index $\frac{1}{2}$ [with Sunjin Choi]

The superconformal index (SCI), or $S^1 \times_w S^2$ partition function, counts certain BPS operators in 3d $\mathcal{N}=2$ SCFTs.

It is useful to consider the Cardy-like limit *ω* → 0. The SCI can then be analyzed with similar tools as the TTI.

For the ABJM theory in the M-theory limit we find the following ω^{-1} and ω^0 results

$$
\log \mathcal{I}_{\text{ABJM}}(N, k, \omega) = -\frac{\pi \sqrt{2k}}{3} \left[\left(\frac{1}{2\omega} + 1 \right) \left(N - \frac{k}{24} + \frac{2}{3k} \right)^{\frac{3}{2}} - \frac{3}{k} \left(N - \frac{k}{24} + \frac{2}{3k} \right)^{\frac{1}{2}} \right] - \frac{2}{\omega} \hat{g}_0(k) - \frac{1}{2} \log \left(N - \frac{k}{24} + \frac{2}{3k} \right) + \hat{f}_0(k) + \mathcal{O}(e^{-\sqrt{N}}) + \mathcal{O}(\omega).
$$

Captures the entropy of the dual supersymmetric $AdS₄$ Kerr-Newman black hole.

Similar results for other 3d $\mathcal{N}=2$ holographic SCFTs.

 $\mathsf{Consider} \ \mathsf{a} \ \mathsf{3d} \ \mathsf{CFT} \ \mathsf{on} \ S^1_{\beta} \times \mathbb{R}^2 \ \mathsf{where} \ \beta = 1/T. \ \mathsf{The} \ \mathsf{1pt} \ \mathsf{function} \ \mathsf{of} \ \mathsf{the}$ energy-momentum tensor and the thermal free energy are

$$
\langle \mathcal{T}^{00} \rangle = \frac{2}{3} \frac{b_{\mathcal{T}}}{\beta^3} \,, \qquad F_{S^1_\beta \times \mathbb{R}^2} = \frac{f_{\mathcal{T}}}{\beta^3} \,, \qquad 3f_{\mathcal{T}} = b_{\mathcal{T}} \,.
$$

 $\mathsf{Consider} \ \mathsf{a} \ \mathsf{3d} \ \mathsf{CFT} \ \mathsf{on} \ S^1_{\beta} \times \mathbb{R}^2 \ \mathsf{where} \ \beta = 1/T. \ \mathsf{The} \ \mathsf{1pt} \ \mathsf{function} \ \mathsf{of} \ \mathsf{the}$ energy-momentum tensor and the thermal free energy are

$$
\langle \mathcal{T}^{00} \rangle = \frac{2}{3} \frac{b_{\mathcal{T}}}{\beta^3} \,, \qquad F_{S^1_\beta \times \mathbb{R}^2} = \frac{f_{\mathcal{T}}}{\beta^3} \,, \qquad 3 f_{\mathcal{T}} = b_{\mathcal{T}} \,.
$$

Q: How do we compute these thermal observables in holographic CFTs?

A: Employ the "unreasonable effectiveness of higher-derivative supergravity in holography"! [NPB-Charles-Hristov-Reys]

 $\mathsf{Consider} \ \mathsf{a} \ \mathsf{3d} \ \mathsf{CFT} \ \mathsf{on} \ S^1_{\beta} \times \mathbb{R}^2 \ \mathsf{where} \ \beta = 1/T. \ \mathsf{The} \ \mathsf{1pt} \ \mathsf{function} \ \mathsf{of} \ \mathsf{the}$ energy-momentum tensor and the thermal free energy are

$$
\langle \mathcal{T}^{00} \rangle = \frac{2}{3} \frac{b \tau}{\beta^3} \,, \qquad F_{S^1_\beta \times \mathbb{R}^2} = \frac{f \tau}{\beta^3} \,, \qquad 3 f \tau = b \tau \,.
$$

Q: How do we compute these thermal observables in holographic CFTs?

A: Employ the "unreasonable effectiveness of higher-derivative supergravity in holography"! [NPB-Charles-Hristov-Reys]

Basic idea: Use the susy localization results above to fix the four-derivative gauged supergravity action. Then use this action in conjunction with holography to compute thermal observables.

 $\mathsf{Consider} \ \mathsf{a} \ \mathsf{3d} \ \mathsf{CFT} \ \mathsf{on} \ S^1_{\beta} \times \mathbb{R}^2 \ \mathsf{where} \ \beta = 1/T. \ \mathsf{The} \ \mathsf{1pt} \ \mathsf{function} \ \mathsf{of} \ \mathsf{the}$ energy-momentum tensor and the thermal free energy are

$$
\langle \mathcal{T}^{00} \rangle = \frac{2}{3} \frac{b_{\mathcal{T}}}{\beta^3} \,, \qquad F_{S^1_\beta \times \mathbb{R}^2} = \frac{f_{\mathcal{T}}}{\beta^3} \,, \qquad 3 f_{\mathcal{T}} = b_{\mathcal{T}} \,.
$$

Q: How do we compute these thermal observables in holographic CFTs?

A: Employ the "unreasonable effectiveness of higher-derivative supergravity in holography"! [NPB-Charles-Hristov-Reys]

Basic idea: Use the susy localization results above to fix the four-derivative gauged supergravity action. Then use this action in conjunction with holography to compute thermal observables.

For the ABJM theory we find

$$
b_{\mathcal{T}} = -\frac{8\pi^2\sqrt{2k}}{27} N^{3/2} + \frac{\pi^2(k^2 - 16)}{27\sqrt{2k}} N^{1/2} + \dots
$$

Somewhat surprisingly we find that to this order at large N $b_{\mathcal{T}} = -\frac{\pi^3}{72} C_T!$

 $\mathsf{Consider} \ \mathsf{a} \ \mathsf{3d} \ \mathsf{CFT} \ \mathsf{on} \ S^1_{\beta} \times \mathbb{R}^2 \ \mathsf{where} \ \beta = 1/T. \ \mathsf{The} \ \mathsf{1pt} \ \mathsf{function} \ \mathsf{of} \ \mathsf{the}$ energy-momentum tensor and the thermal free energy are

$$
\langle \mathcal{T}^{00} \rangle = \frac{2}{3} \frac{b \tau}{\beta^3} \,, \qquad F_{S^1_\beta \times \mathbb{R}^2} = \frac{f \tau}{\beta^3} \,, \qquad 3 f \tau = b \tau \,.
$$

Q: How do we compute these thermal observables in holographic CFTs?

A: Employ the "unreasonable effectiveness of higher-derivative supergravity in holography"! [NPB-Charles-Hristov-Reys]

Basic idea: Use the susy localization results above to fix the four-derivative gauged supergravity action. Then use this action in conjunction with holography to compute thermal observables.

For the ABJM theory we find

$$
b_{\mathcal{T}} = -\frac{8\pi^2\sqrt{2k}}{27} N^{3/2} + \frac{\pi^2(k^2 - 16)}{27\sqrt{2k}} N^{1/2} + \dots
$$

Somewhat surprisingly we find that to this order at large N $b_{\mathcal{T}} = -\frac{\pi^3}{72} C_T!$ Similar results for other 3d $\mathcal{N}=2$ holographic SCFTs.

Summary

- \bullet Presented exact results for the large N limit of the partition function of the ABJM theory on S^3 , $S^1 \times \Sigma_{\mathfrak{g}},$ and $S^1 \times_\omega S^2.$
- **•** Discussed how some of these results can be reproduced by string/M-theory via AdS/CFT.
- All order microscopic prediction for the entropy of the BPS AdS_4 Reissner-Nordström and Kerr-Newman black holes.
- Generalization of these results to some other 3d $\mathcal{N}=2$, $\mathcal{N}=3$, and $\mathcal{N} = 4$ holographic SCFTs.
- Applications of these results to the calculation of thermal observables.

Outlook

- Extend to other 3d $\mathcal{N} = 2$ holographic SCFTs. [in progress]
- Analytic derivation of our results.
- **•** Partition functions on other compact 3-manifolds.
- **O** Understand the shift in N from M-theory. [Bergman-Hirano], [in progress]
- **Implications for the higher-derivative corrections to 4d and 11d** supergravity?
- **•** Supersymmetric localization in 4d supergravity?
- **•** Derivation from (or lessons for) type IIA string theory and M-theory?
- OSV-type conjecture for AdS black holes?

