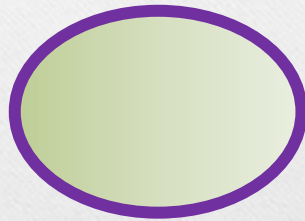


Neutrinos, Cosmological Phase-Transition and the Matter-Antimatter Asymmetry



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Rémi Faure, in collaboration with Stéphane
Lavignac (IPhT, Saclay)

- Why is there no « anti »-version of us?
- Almost only particles are observed at all scales (no anti-particles)
- The CMB gives us the ratio baryons/photons:

$$\eta_B \equiv \frac{n_B - n_{\bar{B}}}{n_\gamma} \simeq \frac{n_B}{n_\gamma} = 6 \cdot 10^{-10} \neq 0$$



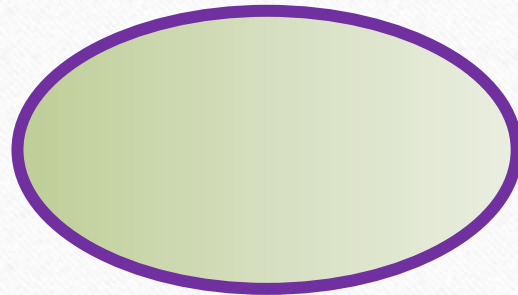
Need for New Physics: **Sterile Neutrinos**



Outline

1. Phase-Transition
2. Propagation with time-dependent masses
3. The role of flavor
4. Lepton asymmetry

1. Phase Transition: going out of equilibrium



Sakharov conditions for generation of matter-antimatter asymmetry (1967):

- **Baryon/Lepton number violation**
- **C and CP violation**
- **Out-of-Equilibrium**



Andrei Sakharov

$$L = L_{SM} + i\bar{N}\gamma^\mu\partial_\mu N - M_N^I \bar{N}_I^c N_I + Y_{I\alpha} N_I \bar{l}_\alpha \tilde{\phi} + h.c.$$

Massive Neutrinos
Majorana mass M

→ Violates **Lepton number**

The Neutrinos interact
with the Standard Model

→ Violates **CP
symmetry**

$$L = L_{SM} + i\bar{N}\gamma^\mu\partial_\mu N - M_N^I \bar{N}_I^c N_I + Y_{I\alpha} N_I \bar{l}_\alpha \tilde{\phi} + h.c.$$

[Khoze, Ro, 2013]
[Fischer, Lindner, van der Woude, 2021]



[Rosauero-Alcaraz (2021)]
[Huang, Xie (2022)]

$$L = L_{SM} + L_S + i\bar{N}\gamma^\mu\partial_\mu N - \lambda_{NS}^I S \bar{N}_I^c N_I + Y_{I\alpha} N_I \bar{l}_\alpha \tilde{\phi} + h.c.$$

$$M_N^I = \lambda_{NS}^I S$$

$$L = L_{SM} + i\bar{N}\gamma^\mu\partial_\mu N - M_N^I \bar{N}_I^c N_I + Y_{I\alpha} N_I \bar{l}_\alpha \tilde{\phi} + h.c.$$

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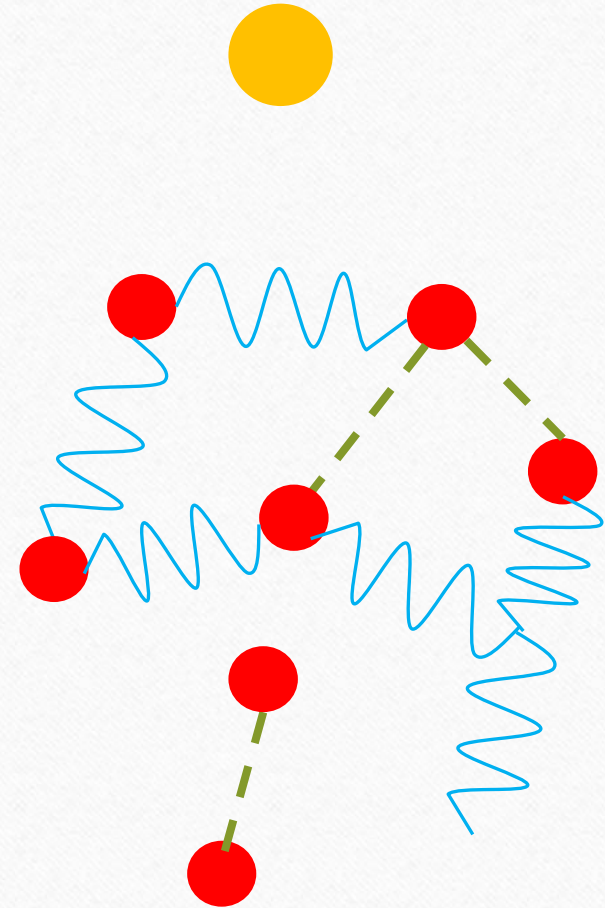
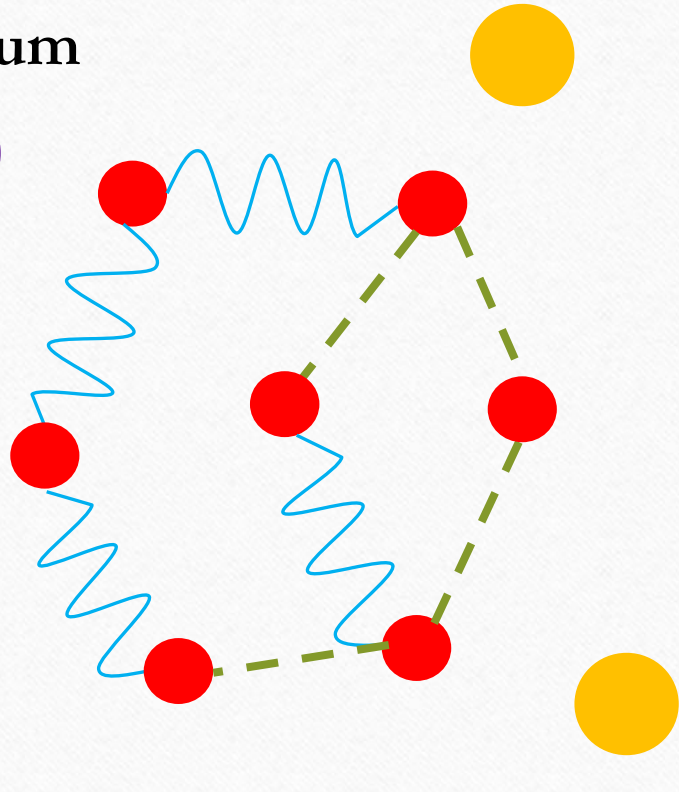
$$M_N^I = \lambda_{NS}^I S$$

New dynamics for the sterile sector: **phase-transition**

False vacuum

$$\langle S \rangle = 0$$

$$M = 0$$



Sterile Neutrino

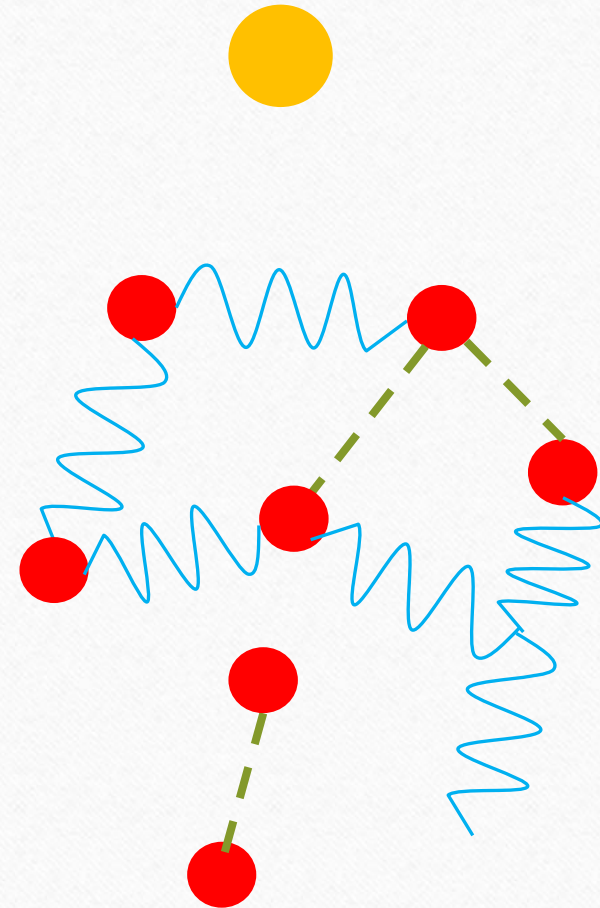
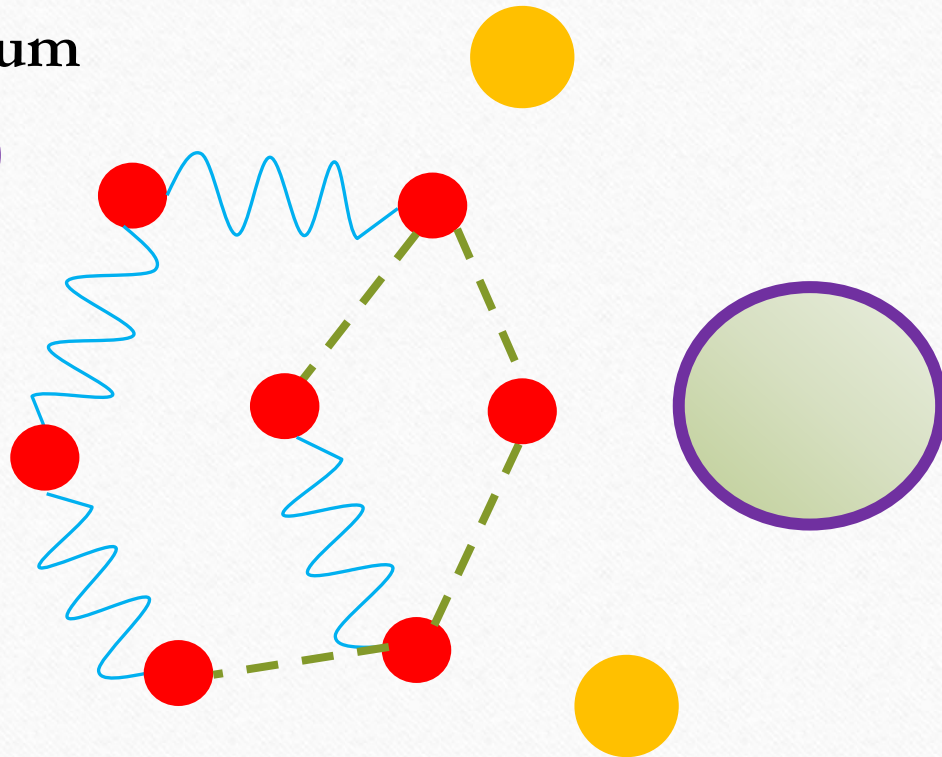


SM fermion

False vacuum

$$\langle S \rangle = 0$$

$$M = 0$$



● Sterile Neutrino ● SM fermion

False vacuum

$$\langle S \rangle = 0$$

$$M = 0$$

True vacuum

$$\langle S \rangle \neq 0$$

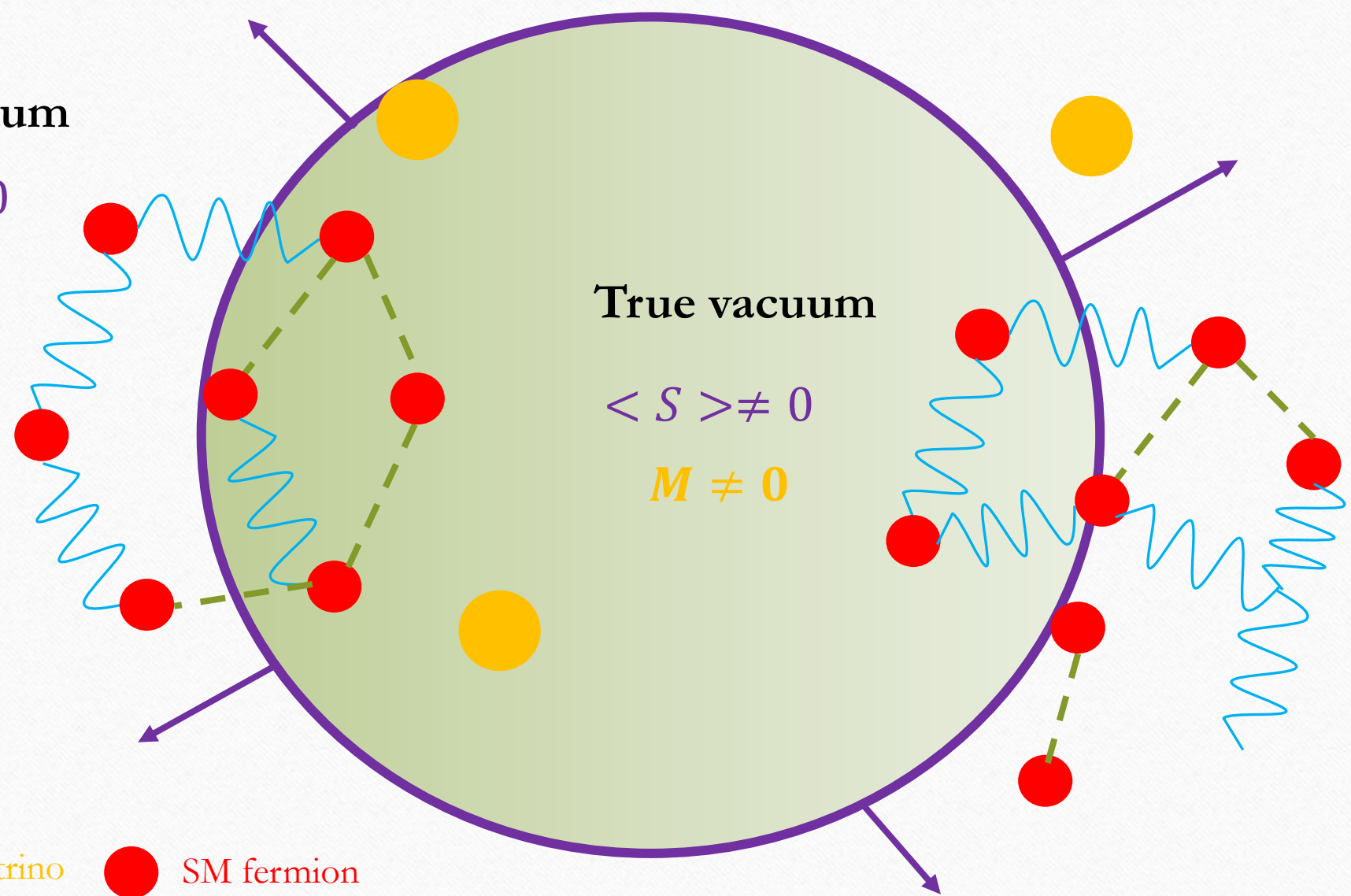
$$M \neq 0$$



Sterile Neutrino



SM fermion



False vacuum

$$\langle S \rangle = 0$$

$$M = 0$$

True vacuum

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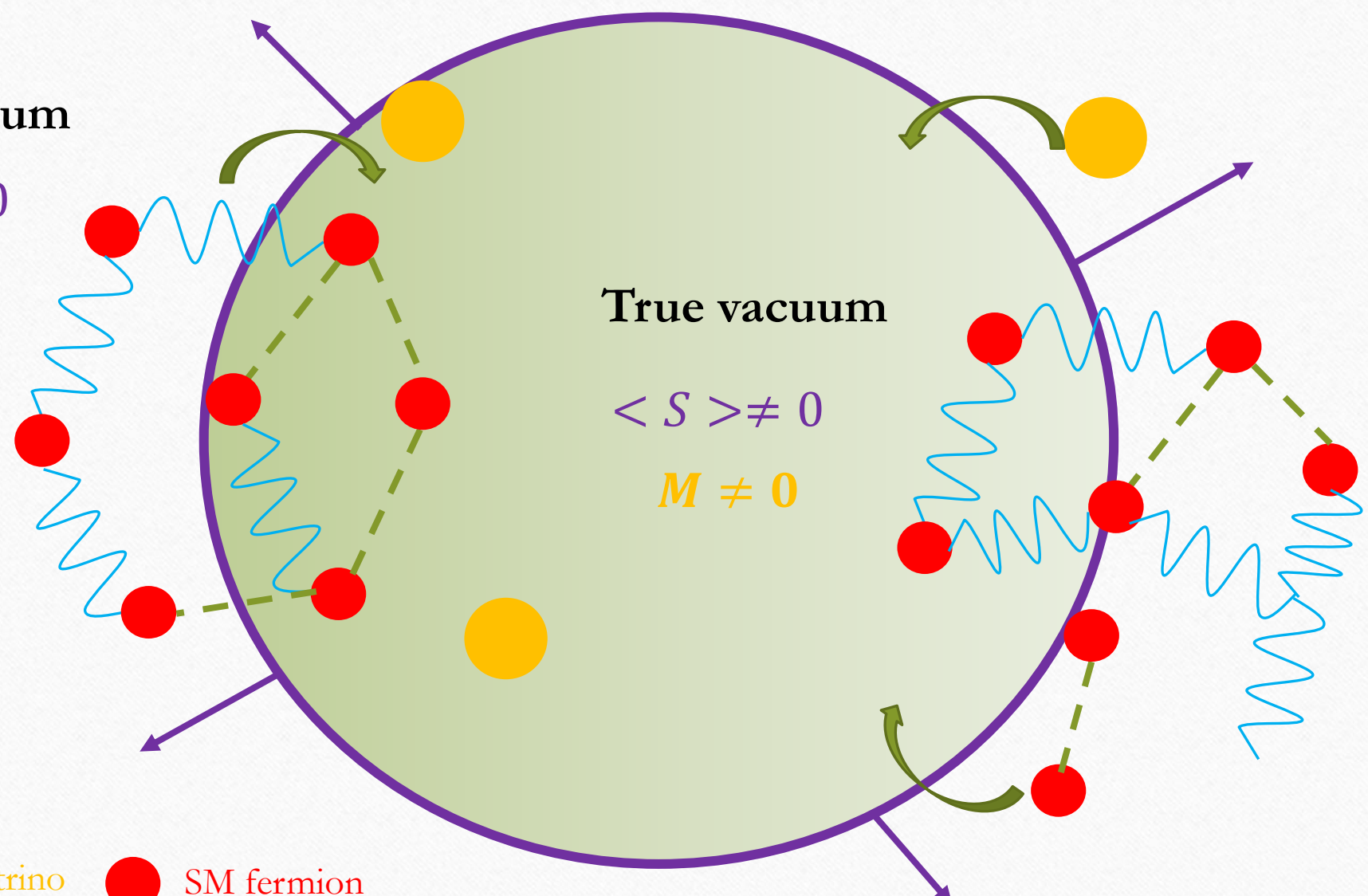
$$M \neq 0$$



Sterile Neutrino

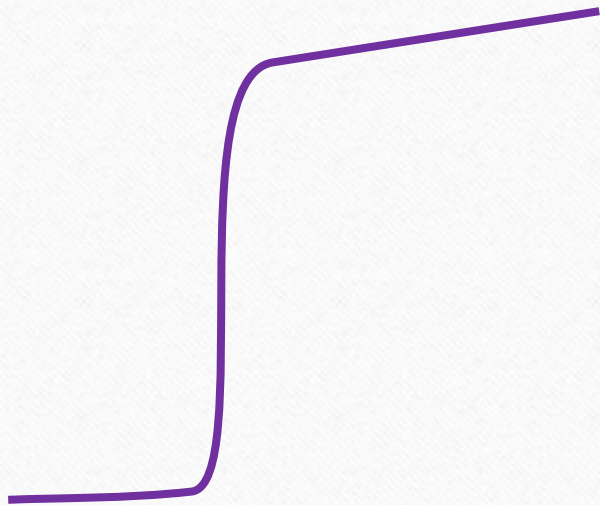


SM fermion



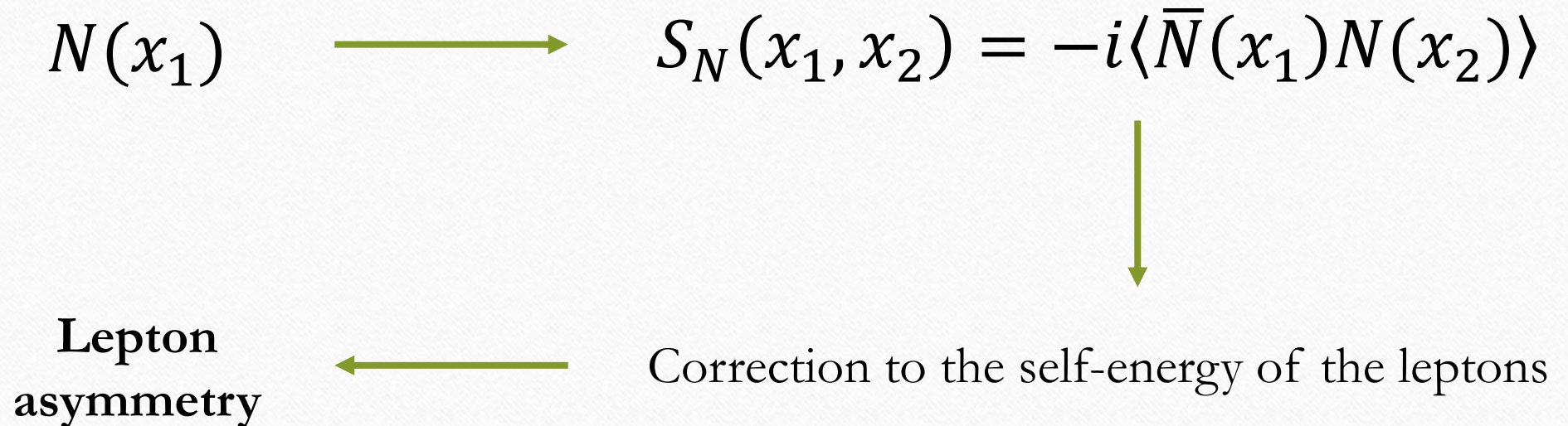
During the **phase-transition**, the masses of the Neutrinos are **time-dependent**. All quantities will have an explicit time-dependence along the **wall**.

$\langle S \rangle (t)$

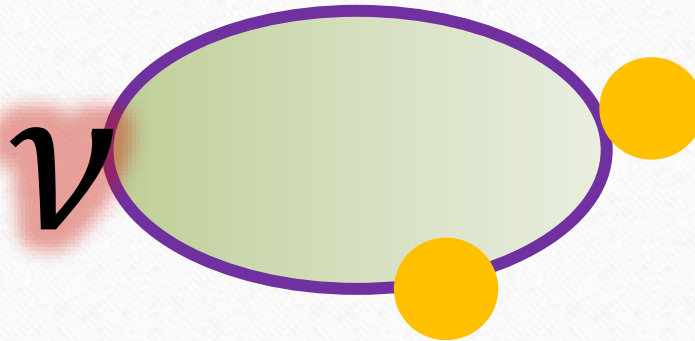


$$M_N^I = \lambda_{NS}^I \langle S \rangle (t) = M_N^I(t)$$

A direct brut-force QFT calculation turns out to be complicated. Our strategy will rather be to consider the **Dirac equation** and solve it directly.



2. Propagation with time-dependent masses



Consider the left- and right-handed parts of the Majorana Neutrino (one flavor for simplicity):

$\sigma^0 = Id, \sigma^i = \text{Pauli matrices}$

$$N = \begin{pmatrix} N_L^c = i \sigma^2 N_R \\ N_R \end{pmatrix}$$

$$i \sigma^\mu \partial_\mu N_R - M_N(t) N_L^c = 0$$

$$i \bar{\sigma}^\mu \partial_\mu N_L^c - M_N(t) N_R = 0$$

We decompose the Majorana field in (spatial) momentum modes:

$$N(\mathbf{x}, t) = \sum_{h=\pm} \int d^3\mathbf{k} e^{i\mathbf{k}\cdot\mathbf{x}} \left(\begin{pmatrix} L_h \\ R_h \end{pmatrix} \otimes \xi_{\mathbf{k},h} a_{\mathbf{k},h} + \begin{pmatrix} h R_h^* \\ -h L_h^* \end{pmatrix} \otimes \xi_{-\mathbf{k},-h} a_{-\mathbf{k},h}^\dagger \right)$$

$\xi_{\mathbf{k},h}$ = helicity eigenvectors

Free case: $L_h, R_h \propto e^{-i\omega t}$, $L_h^*, R_h^* \propto e^{+i\omega t}$
(valid at $t = -\infty$)

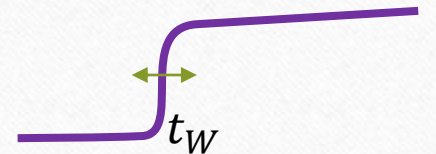
$a_{\mathbf{k},h}, a_{\mathbf{k},h}^\dagger$ = annihilation
and creation operators
defined at $t = -\infty$

$$i \partial_t L_h + h k L_h - M_N(t) R_h = 0$$

$$i \partial_t R_h - h k R_h - M_N(t) L_h = 0$$

For a time-dependence of the mass $M_N(t) = M_N^0(1 + \tanh(t/t_W))/2$, one can get analytical results.

[Prokopec, Schmidt, Weenink, 2013]



$$Z \equiv \frac{1 + \tanh(t/t_W)}{2}$$

$$\gamma \equiv 1/t_W \quad \frac{d}{dt} = \frac{dZ}{dt} \frac{d}{dZ} = \frac{1}{2 t_W} (1 - \tanh(t/t_W))^2 \frac{d}{dZ} = 2 \gamma Z (1 - Z) \frac{d}{dZ}$$

$$\begin{aligned}
 i \partial_t L_h + h k L_h - M_N(t) R_h &= 0 \\
 i \partial_t R_h - h k R_h - M_N(t) L_h &= 0
 \end{aligned}
 \longrightarrow
 \left[\partial_t^2 + 2ihk \frac{\partial_t M_N}{M_N} \partial_t + (k^2 + M_N(t)^2) \right] L_h = 0$$

$$L_h \equiv Z^\alpha (1 - Z)^\beta \chi_h(Z)$$



$$Z(1 - Z)\chi_h'' + (c_h - (a_h + b_h + 1)Z)\chi_h' - a_h b_h \chi_h = 0$$

$$\begin{aligned} i \partial_t L_h + h k L_h - M_N(t) R_h &= 0 \\ i \partial_t R_h - h k R_h - M_N(t) L_h &= 0 \end{aligned} \quad \longrightarrow \quad \left[\partial_t^2 + 2ihk \frac{\partial_t M_N}{M_N} \partial_t + (k^2 + M_N(t)^2) \right] L_h = 0$$

$$L_h \equiv Z^\alpha (1 - Z)^\beta \chi_h(Z)$$



$$Z(1 - Z)\chi_h'' + (c_h - (a_h + b_h + 1)Z)\chi_h' - a_h b_h \chi_h = 0$$

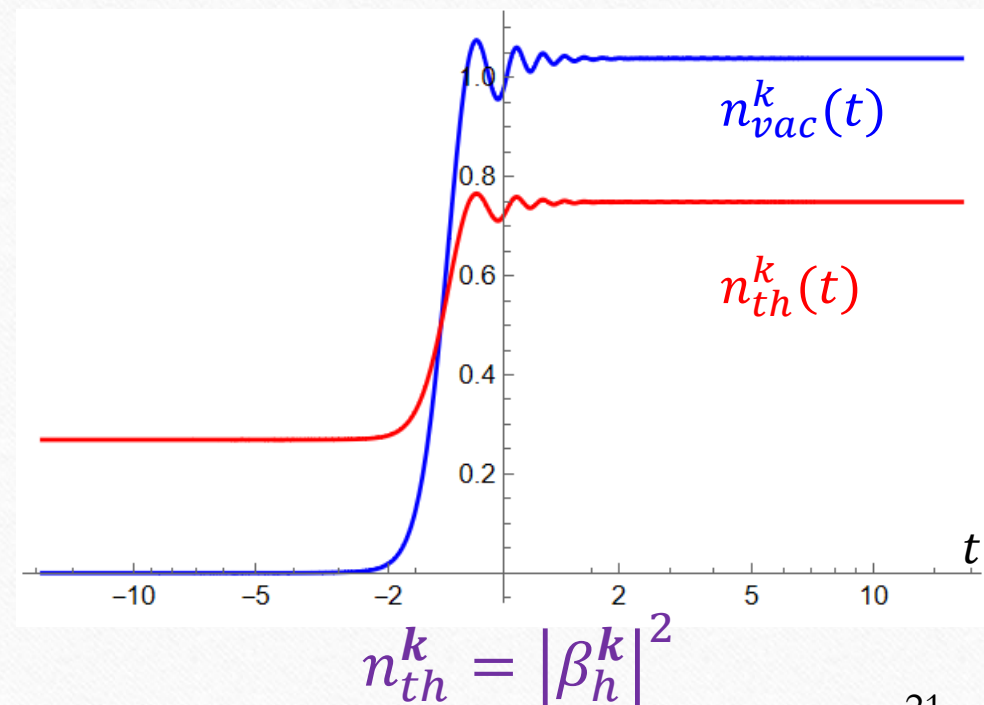
$$\dots \quad L_h \sim \lambda_h {}_2F_1(a_h, b_h, c_h, Z) + \mu_h {}_2F_1(a'_h, b'_h, c'_h, Z)$$

(Gaussian hypergeometric function, with a_h, b_h, c_h functions of k and M_N)

$$N_R(\mathbf{x}, t) = \sum_{h=\pm} \int d^3\mathbf{k} e^{i\mathbf{k}\cdot\mathbf{x}} (\mathbf{L}_h \otimes \xi_{\mathbf{k},h} a_{\mathbf{k},h} + h \mathbf{R}_h^* \otimes \xi_{-\mathbf{k},-h} a_{-\mathbf{k},h}^\dagger)$$

$$L_h \sim {}_2F_1(\dots) \sim_{t \rightarrow +\infty} \alpha_h^{\mathbf{k}} e^{-i\omega_+ t} + \beta_h^{\mathbf{k}} e^{+i\omega_+ t} \neq \text{Free case}$$

The modes that correspond to positive frequencies at initial times ($-\infty$) end up being a **linear combination** of positive and negative frequencies in the far future ($+\infty$). This corresponds to making a **Bogolyubov transformation** of the annihilation and creation operators.



3. The role of flavor

If we add interactions to the story,

$$i \sigma^\mu \partial_\mu N_{RI} - M_N^I(t) N_{LI}^c = Y_{I\alpha} \bar{l}_\alpha \tilde{\phi}$$

$$i \bar{\sigma}^\mu \partial_\mu N_{LI}^c - M_N^I(t) N_{RI} = Y_{I\alpha}^* \bar{l}_\alpha^c \phi$$

...(thermal average)

$$\left[\partial_t^2 + 2ihk \frac{\partial_t M_N^I}{M_N^I} \partial_t + (k^2 + M_N^I(t)^2) \right] L_h^{IJ} + M_{th,IK}^2 L_h^{KJ} = 0$$

$$M_{th,IK}^2 = (Y^* Y^T)_{IK} T^2 / 12$$

Order-by-order: $M_{th,IK}^2$ is small

$$L_h^{IJ} = L_h^{IJ(0)} + L_h^{IJ(1)} + \dots$$

$$\left[\partial_t^2 + 2ihk \frac{\partial_t M_N^I}{M_N^I} \partial_t + (k^2 + M_N^I(t)^2) \right] L_h^{IJ(0)} = 0$$

Diagonal
in flavor

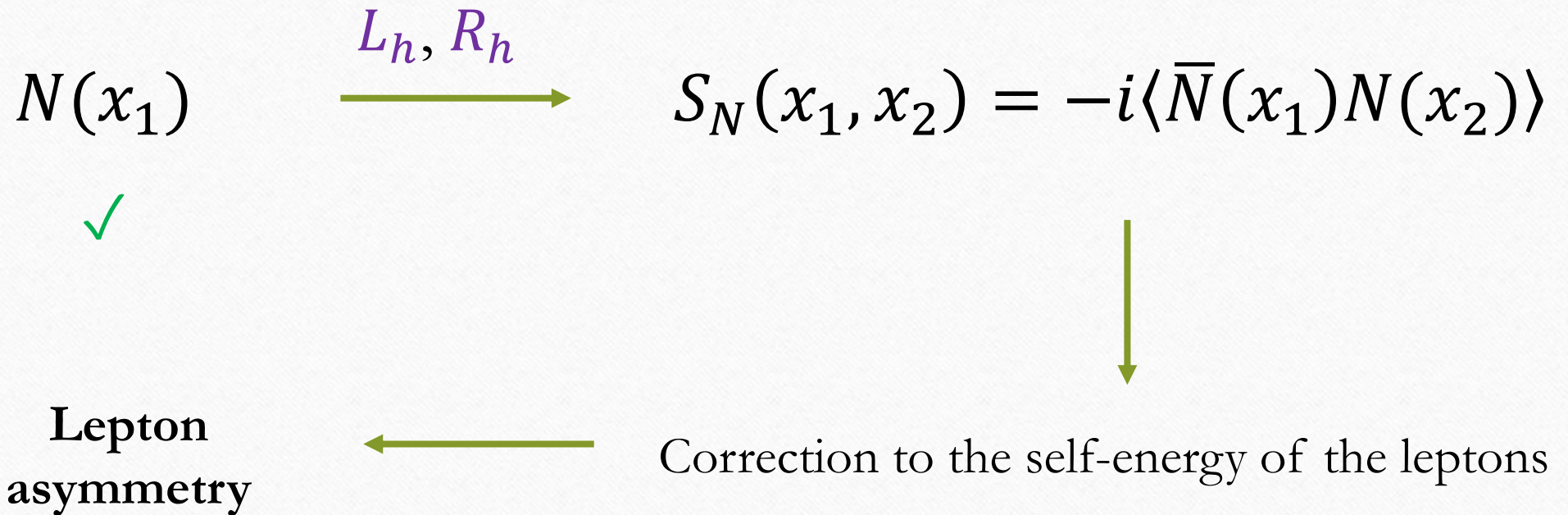
$$\left[\partial_t^2 + 2ihk \frac{\partial_t M_N^I}{M_N^I} \partial_t + (k^2 + M_N^I(t)^2) \right] L_h^{IJ(1)} = -M_{th,IK}^2 L_h^{KJ(0)}$$

Same homogeneous part \Rightarrow hypergeometric functions

4. Lepton asymmetry

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The propagator can equivalently be written with the modes or with **phase-space distributions**:

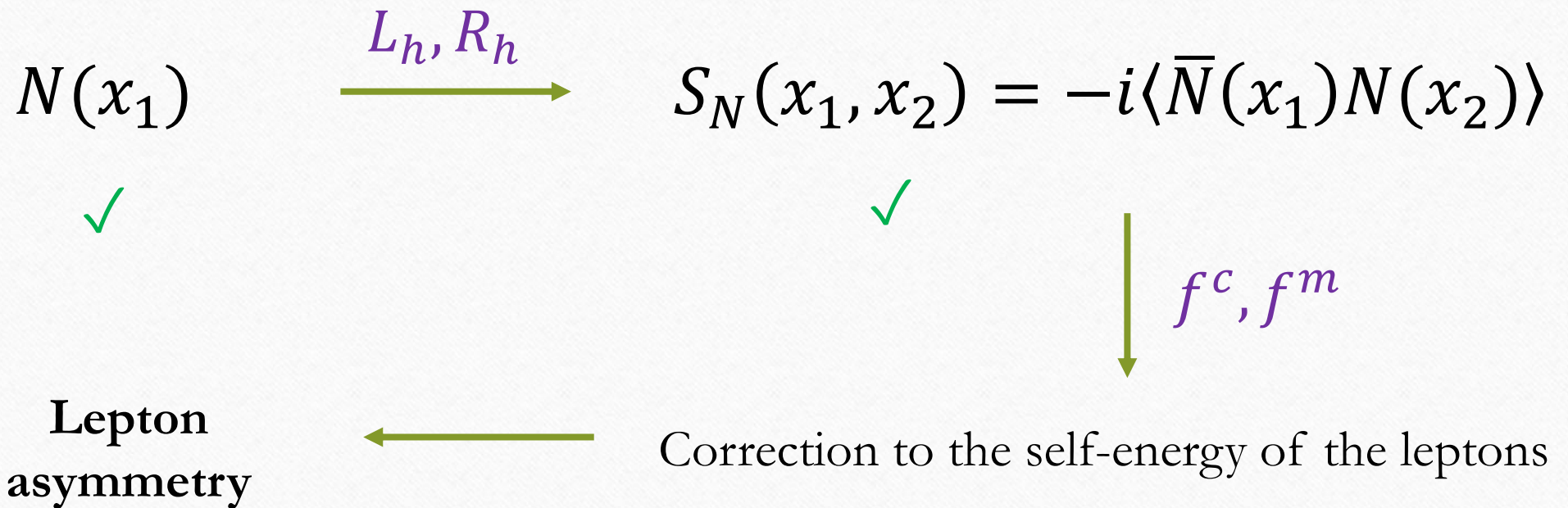
$$\begin{aligned}
 S_N &= \mathbb{1} - \sum_h \begin{pmatrix} L_h L_h^\dagger & L_h R_h^\dagger \\ R_h L_h^\dagger & R_h R_h^\dagger \end{pmatrix} \otimes P_{\mathbf{k},h} \\
 &= \sum_{h,s} \underbrace{P_{\mathbf{k},h} P_I^s \gamma^0 P_J^s}_{\mathcal{P}_{\mathbf{k},h}^{m,s}} f_{IJ}^{m,s} + \underbrace{P_{\mathbf{k},h} P_I^s \gamma^0 P_J^{-s}}_{\mathcal{P}_{\mathbf{k},h}^{c,s}} f_{IJ}^{c,s} \\
 &= \left[\sum_{h,s} \mathcal{P}_{\mathbf{k},h}^{m,s} f_h^{m,s} + \mathcal{P}_{\mathbf{k},h}^{c,s} f_h^{c,s} \right]_{IJ}
 \end{aligned}$$

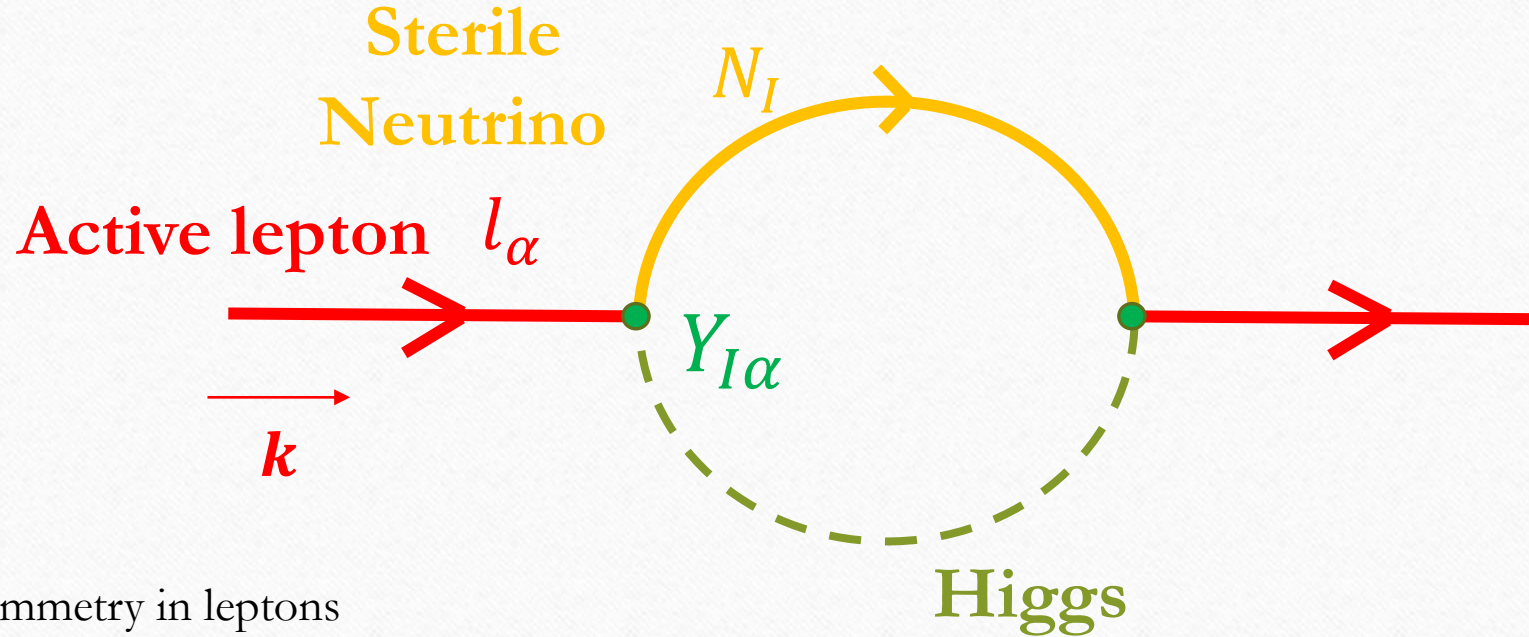
By projecting on the correct subspaces, we have linear relations between the modes L_h/R_h and the distribution functions f^c/f^m ,

$$f_{k,h}^{m,s} = \text{Tr}(\mathcal{P}_{k,h}^{m,s} S_N) \quad f_{k,h}^{c,s} = \text{Tr}(\mathcal{P}_{k,h}^{c,s} S_N)$$

$f_{k,h}^{m,s}$ is called the **mass-shell** distribution function \sim particle-particle transitions $\sim e^{i(\omega_I - \omega_J)t}$ **Slow mode**

$f_{k,h}^{c,s}$ is called the **coherence-shell** distribution function \sim particle-antiparticle transitions $\sim e^{i(\omega_I + \omega_J)t}$ **Fast mode**





$\Delta_L^k =$ asymmetry in leptons

$$\partial_t \Delta_L^k \approx \left[\text{Im}(YY^\dagger)_{IJ} \text{Im}(f_{k,+}^m + f_{k,-}^m)_{JI} + \text{Re}(YY^\dagger)_{IJ} \text{Re}(f_{k,+}^m - f_{k,-}^m)_{JI} \right] \times \hat{\Sigma}_{\mathcal{A}}(\mathbf{k})$$

Phase-space distribution of Neutrinos

[Jukkala, Kainulainen, Rahkila, 2021]

[Drewes, Garbrecht, 2012]

$\hat{\Sigma}_{\mathcal{A}}(\mathbf{k}) =$ self-energy of the Neutrino

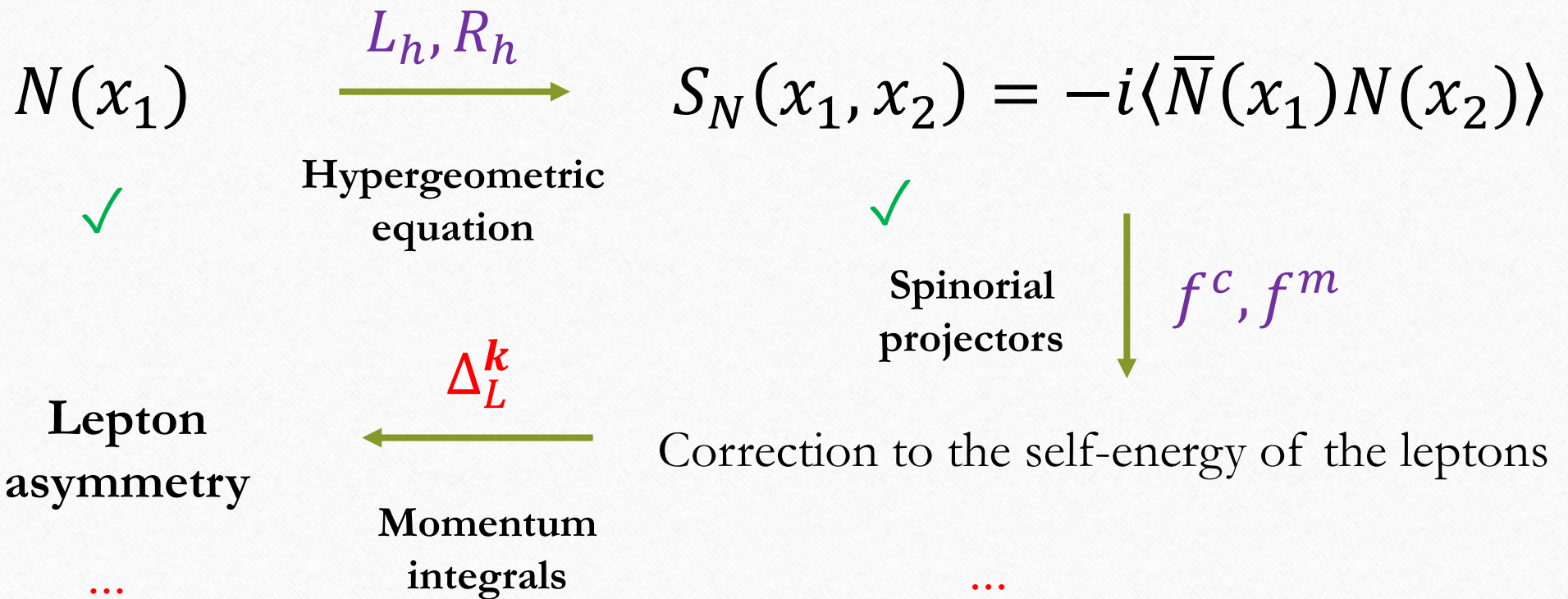
$$\partial_t \Delta_L^k \approx \left[\text{Im}(YY^\dagger)_{IJ} \text{Im}(f_{k,+}^m + f_{k,-}^m)_{JI} + \text{Re}(YY^\dagger)_{IJ} \text{Re}(f_{k,+}^m - f_{k,-}^m)_{JI} \right] \times \hat{\Sigma}_{\mathcal{A}}(\mathbf{k})$$

The phase-space distributions were solved to first-order in YY^\dagger and lead to many terms. As an example,

$$\begin{aligned} & \left[\text{Im}(YY^\dagger)_{IJ} \text{Im}(f_{k,+}^m + f_{k,-}^m)_{JI} + \text{Re}(YY^\dagger)_{IJ} \text{Re}(f_{k,+}^m - f_{k,-}^m)_{JI} \right] \\ & \ni \frac{M_I^2 - M_J^2}{8 k \langle M \rangle_{IJ}} \text{Im} \left[(YY^\dagger)_{IJ}^2 \right] \left[\text{Im} \left(L_+^{(0)} \chi_-^* - L_-^{(0)} \chi_+^* \right)_{JI} - \text{Re} \left(L_+^{(0)} \chi_-^* + L_-^{(0)} \chi_+^* \right)_{JI} \right] \end{aligned}$$

Lepton number violation + CP-violation + Out-of-equilibrium modes = Lepton asymmetry source

Summary and conclusion



Further prospects

- Numerical implementation
- Dependence on the parameters of the Phase Transition (thickness of the wall)
- Washout of the asymmetries before electroweak PT

True vacuum

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Thank you for your attention!



Homogeneous and isotropic Universe

$$ds^2 = dt^2 - a(t)^2 d\vec{x}^2$$

Scale factor = « Size » of the Universe

Its evolution is driven by the **matter content**

$$f_h(\vec{p}, \vec{x}, t) = f_h(\vec{p}, t) \quad n_h = \int d^3\vec{p} f_h(\vec{p}, t) = n_h(t) = n^{eq} + \delta n_h$$

The constants λ_h and μ_h are determined from initial conditions + normalization:

$$N(\mathbf{x}, t) = \sum_{h=\pm} \int d^3\mathbf{k} e^{i\mathbf{k}\cdot\mathbf{x}} \left(\begin{pmatrix} L_h \\ R_h \end{pmatrix} \otimes \xi_{\mathbf{k},h} a_{\mathbf{k},h} + \begin{pmatrix} h R_h^* \\ -h L_h^* \end{pmatrix} \otimes \xi_{\mathbf{k},-h} a_{-\mathbf{k},h}^\dagger \right)$$

$\xi_{\mathbf{k},h}$ = helicity eigenvectors

$$\underline{t = -\infty} \quad L_h, R_h \propto e^{-i\omega_- t}, \quad L_h^*, R_h^* \propto e^{+i\omega_- t}$$

$a_{\mathbf{k},h}, a_{\mathbf{k},h}^\dagger$ = annihilation
and creation operators
defined at $t = -\infty$

Normalization: $|L_h|^2 + |R_h|^2 = 1$

Multiflavor field decomposition:

$$N_I(\mathbf{x}, t) = \sum_{h=\pm} \int d^3\mathbf{k} e^{i\mathbf{k}\cdot\mathbf{x}} \left(\begin{pmatrix} L_h^{IJ} \\ R_h^{IJ} \end{pmatrix} \otimes \xi_{\mathbf{k},h} a_{\mathbf{k},hJ} + h.c. \right)$$

$\xi_{\mathbf{k},h}$ = helicity eigenvectors

$t = -\infty$

$$L_h^{IJ}, R_h^{IJ} \propto e^{-i\omega_- t}$$

$a_{\mathbf{k},h}, a_{\mathbf{k},h}^\dagger$ = annihilation
and creation operators
defined at $t = -\infty$

Normalization:

$$L_h L_h^\dagger + R_h R_h^\dagger = 1$$

$$L_h^{IJ(1)} \equiv -M_{th,IJ}^2 Z^\alpha (1-Z)^{\beta_I} \chi_h^{IJ}(Z)$$

$$L_h^{JJ(0)} \equiv Z^\alpha (1-Z)^{\beta_J} \chi_h^{J(0)}(Z)$$

$$\begin{aligned} Z(1-Z)\chi_h^{IJ'''} + (c_I - (a_I + b_I + 1)Z)\chi_h^{IJ'} - a_I b_I \chi_h^{IJ} \\ = (1-Z)^{\beta_J - \beta_I} \chi_h^{J(0)}(Z) \end{aligned}$$

The general solution is the sum of an **homogeneous** and a **particular** solutions. The particular solution can be found from the source using the **Wronskian**.

$$\chi_h^{IJ} = \chi_p^{IJ} + \chi_{hom}^I$$

Leptogenesis via neutrino oscillations (ARS)

[See for example
Drewes, Garbrecht,
Gueter, Klaric (2010)]

$$T_{osc} \equiv \left(M_{Pl} \Delta M_{2,1}^2 \right)^{1/3}$$

$M_{Pl} \simeq 7 \times 10^{17} \text{ GeV}$
is the Planck Mass

$$T_{osc} = \left(7 \times 10^{17} \times (110^2 - 100^2) \right)^{1/3} \text{ GeV} \simeq 10^7 \text{ GeV} > T_{PT}$$

$$M_1 = 100 \text{ GeV} \quad M_2 = 110 \text{ GeV} \quad T_{PT} = 10^6 \text{ GeV}$$