

Winter school analogue gravity/cosmology in Benasque in January 2026!



Experiments on rotating geometries with fluids of light

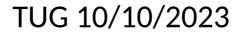
Quantum Optics Group Laboratoire Kastler Brossel, Paris

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In a (quantum) fluid m - massFluid velocity v= $(\hbar/m)\nabla\phi$ Speed of sound c_s $\propto \sqrt{\frac{g\rho_0}{m}}$ q – interaction constant ρ_0 – density Wave eq for collective excitations of (super)fluid $\psi = \psi_0 + \epsilon_1 \psi_1$ $-\partial_t \left(\frac{\rho_0}{c_*^2} (\partial_t \rho_1 + \boldsymbol{v_0} \nabla \rho_1) \right) + \nabla \left(\rho_0 \nabla \rho_1 - \frac{\rho_0 \boldsymbol{v_0}}{c_*^2} \partial_t \rho_1 + \boldsymbol{v_0} \nabla \rho_1 \right) = 0$ Relavistic form of wave eq for collective excitations: $\Delta \rho_1 = \frac{1}{\sqrt{-\eta}} \partial_\mu (\sqrt{-\eta} \eta^{\mu\nu} \partial_\nu \rho_1) = 0$ with $\eta_{\mu\nu} = \begin{pmatrix} -(c_s^2 - \boldsymbol{v_0}^2) & -v_o^x & -v_o^y \\ -v_o^x & 1 & 0 \\ -v_o^y & 0 & 1 \end{pmatrix}$

Motion of collective excitations in inhomogeneous fluid flow ↔ scalar field on curved spacetime

Control parameters: **v**₀, **c**_s

In a (quantum) fluid

Fluid velocity $\mathbf{v} = (\hbar/m) \nabla \phi$

Speed of sound $\mathrm{c}_s \propto \sqrt{rac{g
ho_0}{m}}$

m – mass g – interaction constant ρ_0 – density

Possible geometries with
$$\eta_{\mu\nu} = \begin{pmatrix} -(c_s^2 - v_0^2) & -v_o^x & -v_o^y \\ -v_o^x & 1 & 0 \\ -v_o^y & 0 & 1 \end{pmatrix}$$

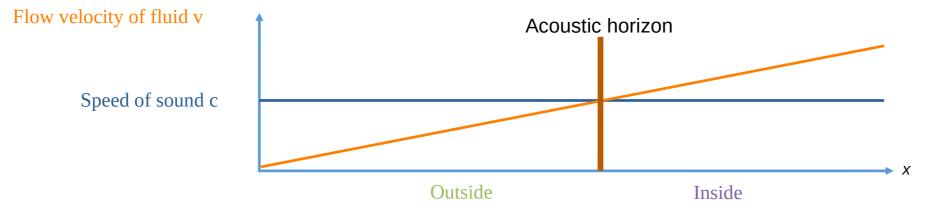
(i) accelerating flow along 1 spatial dimension $\, \scriptscriptstyle \to \,$ static 1D spacetime Horizon where $v_0 = c_s$

(ii) radially accelerating flow in 2 spatial dimensions $\,_{\rightarrow}\,$ static spherically symmetric 2D spacetime Horizon where $\,\,v_r=c_s$

(iii) radially and azimuthally accelerating flow in 2 spatial dimensions \rightarrow static rotating spacetime Horizon where $v_r = c_s$ Ergosurface where $|v_0| = c_s$

Unruh PRL 46 1351 (1981), Visser Class Quant Grav 15 1767 (1998)

Static 1D geometry \leftrightarrow waterfall geometry



Quantised acoustic field:

in:
$$\phi = \int d\omega \left(a_{\omega} f_{\omega} + a_{\omega}^{\dagger} f_{\omega}^{*} \right) \quad a |0\rangle = 0$$
 Express out modes in terms of in modes:
out: $\phi = \int d\omega \left(\bar{a}_{\omega} F_{\omega} + \bar{a}_{\omega}^{\dagger} F_{\omega}^{*} \right) \quad \bar{a} |\bar{0}\rangle = 0$

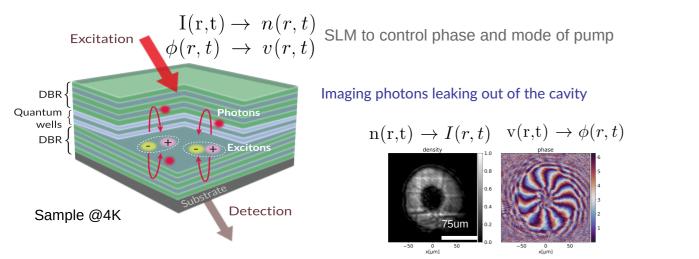
$$F_{\omega} = \int d\omega' \left(\alpha_{\omega\omega'} f_{\omega'} + \beta_{\omega\omega'} f_{\omega'}^{*} \right) \quad \bar{a} |\bar{0}\rangle = 0$$

Different speeds on either side of the horizon $\Rightarrow |\bar{0}\rangle \neq |0\rangle \Rightarrow \beta_{\omega\omega'} \neq 0$

Mixing of positive and negative frequency waves \Rightarrow mixing of creation and annihilation operators

a $|\bar{0}\rangle = \sum_{\omega'} \beta_{\omega\omega'} |\bar{1}\rangle > 0$



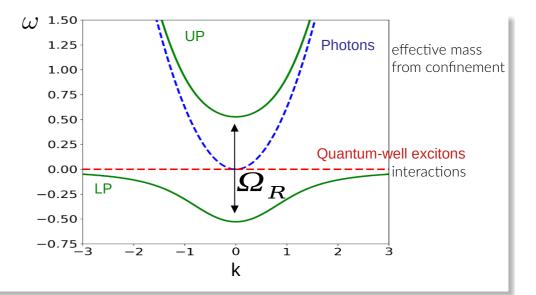


Polaritons= photons dressed with material excitations that live in the cavity plane

Dynamics in the cavity plane described by Gross-Pitaevskii (Nonlinear Schrödinger) equation:

$$\mathrm{i}\hbar\frac{\partial\psi}{\partial t} = \left(-\frac{\hbar^2\nabla^2}{2m_{LP}^*} + gn\right)\psi - \frac{i\hbar\gamma}{2}\psi + P(r,t)$$

Driven-dissipative dynamics \rightarrow Out-of-equilibrium system



q polariton-polariton interaction constant



Collective excitations of polariton fluid

GPE:
$$i\hbar \frac{\partial \psi}{\partial t} = \left(-\frac{\hbar^2 \nabla^2}{2m_{LP}^*} + gn\right)\psi - \frac{i\hbar\gamma}{2}\psi + P(r,t)$$

Bogoliubov theory:

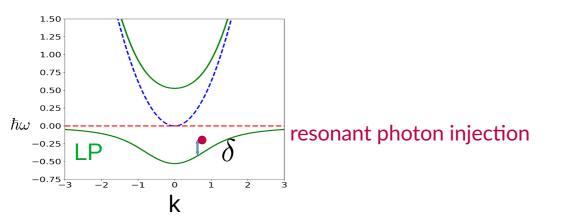
1. Linearise GPE around steady-state solution $\psi(r,t) = \psi_0(r,t) + \delta \psi(r,t)$

2. Equation of motion of weak perturbations

3. Eigenvalues of $L_{Bog} ==$ dispersion relation

$$\hbar\omega(k) = \pm \sqrt{\left(\frac{\hbar^2 k^2}{2m} - \delta + 2gn\right)^2 - (gn)^2} - i\gamma$$

 $i\hbar\frac{\partial}{\partial t} \begin{pmatrix} \delta\psi(r,t) \\ \delta\psi^*(r,t) \end{pmatrix} = L_{\text{Bog}} \begin{pmatrix} \delta\psi(r,t) \\ \delta\psi^*(r,t) \end{pmatrix}$





Collective excitations of polariton fluid

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 $\delta = gn$

$$\omega(k) = \pm \sqrt{\frac{\hbar k^2}{2m} \left(\frac{\hbar k^2}{2m} + 2gn\right)}$$

At low *k*, dispersion is linear \rightarrow excitations are phononic with "speed of sound" $c_s = \sqrt{\hbar g n}/m$

Experimental scheme LKB $\nabla \phi_{SLM} = \frac{C}{r} \boldsymbol{u}_{\boldsymbol{\theta}} - \frac{D}{r} \boldsymbol{u}_{\boldsymbol{r}}$ $\phi_{SLM} = C\boldsymbol{\theta} - Dln(r)$ SLM Ti:Sa CL CL PBS slit HWP Reference beam sample in cryostat 120µm camera Real space C=6 NPBS

NPBS

Momentum space

C=12

D=0

camera

spectrometer

6

- 5

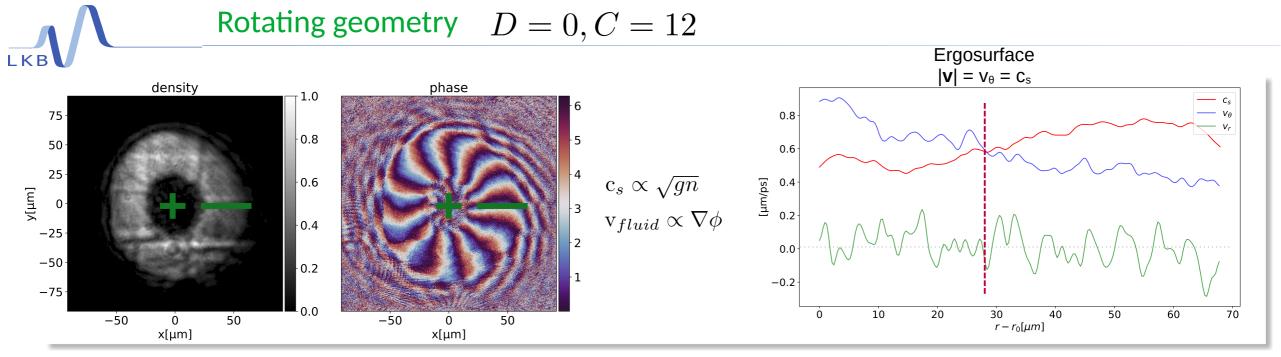
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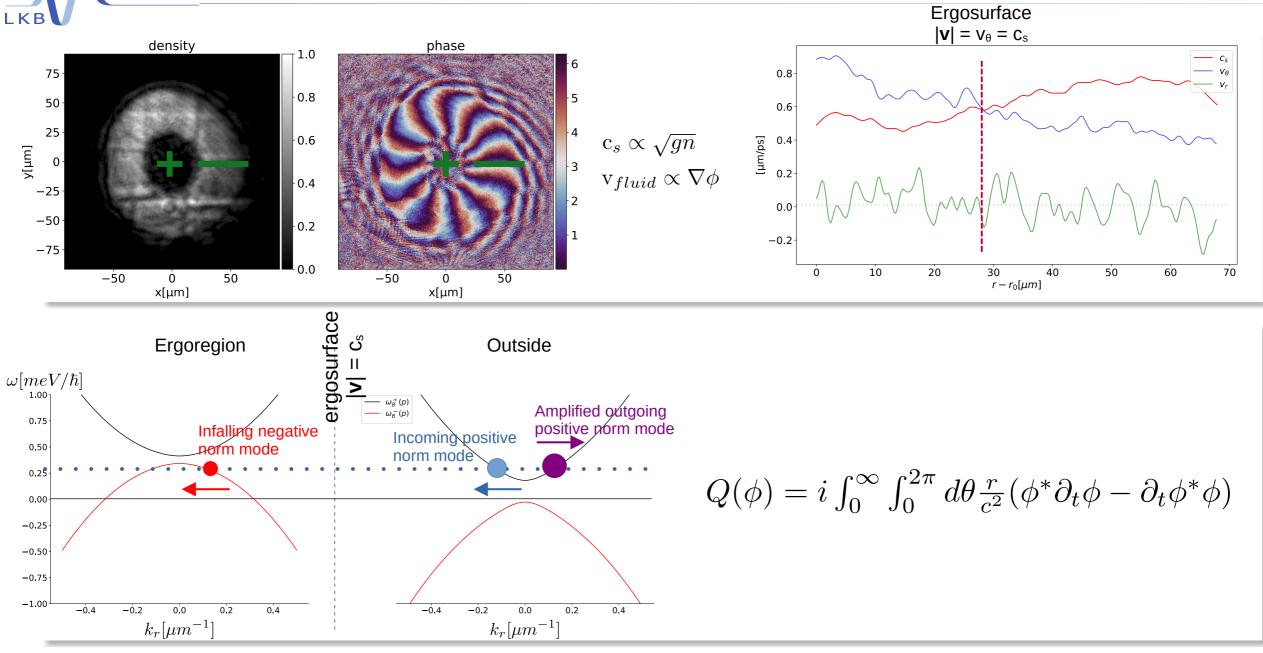
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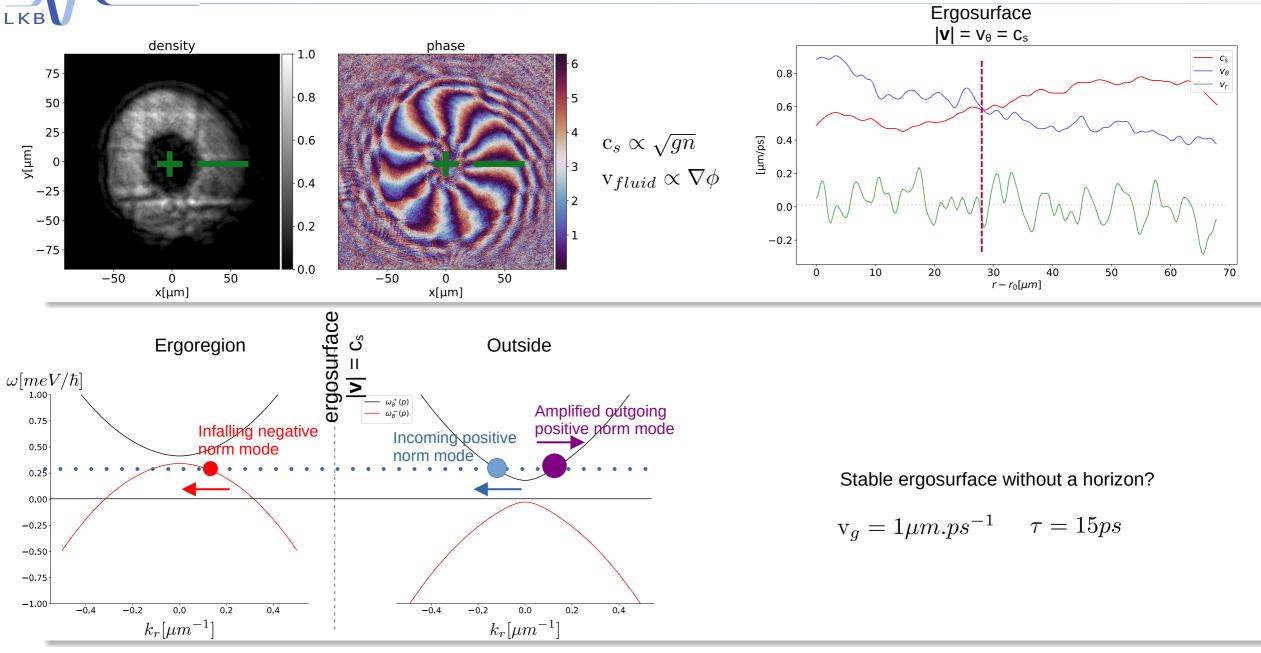
D=9



Rotating geometry D = 0, C = 12

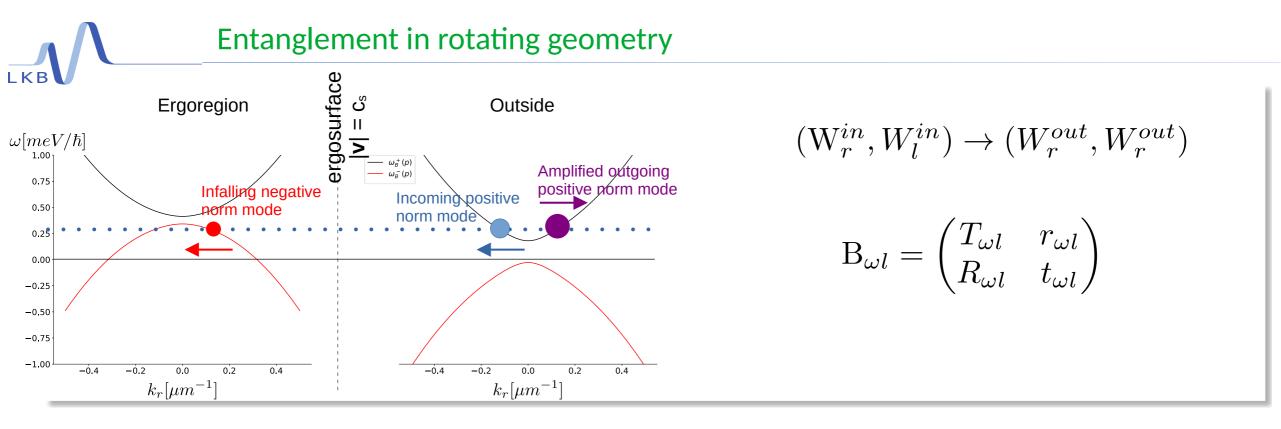


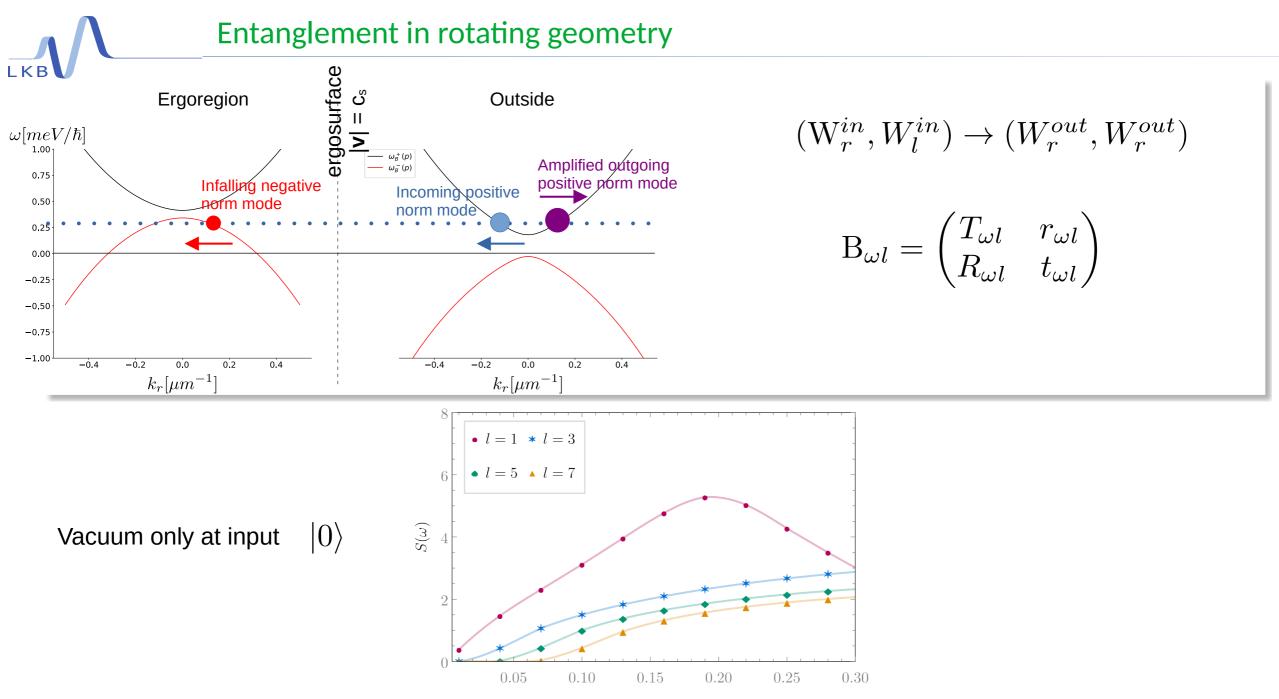
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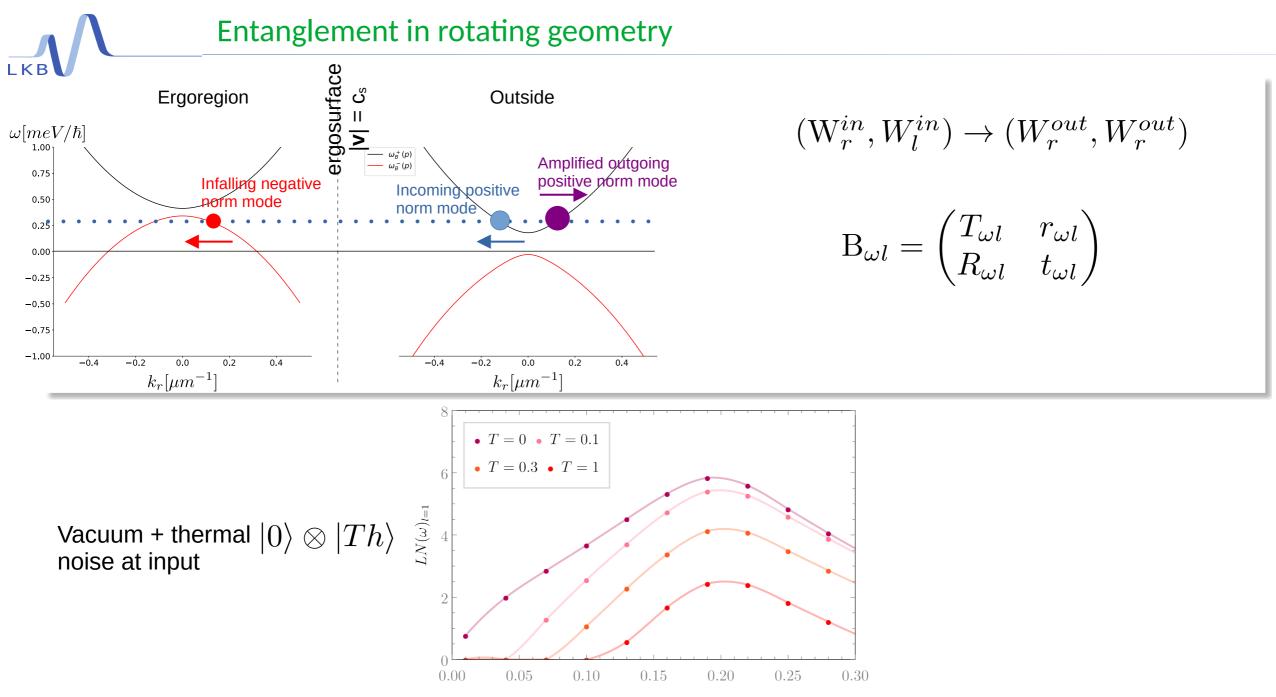


Entanglement in rotating geometry LKB ergosurface ပိ Ergoregion Outside || |> $\omega [meV/\hbar]$ 1.00 $\omega_B^{+}(p)$ Amplified outgoing $\omega_B^{-}(p)$ 0.75 positive porm mode Infalling negative Incoming positive norm mode 0.50 norm mode 0.25 0.00 -0.25 -0.50-0.75 -1.000.0 0.2 -0.4 -<u>0</u>.2 0.0 0.2 0.4 -0.4 -0.2 0.4 $k_r[\mu m^{-1}]$ $k_r[\mu m^{-1}]$

$$\begin{split} & [(\partial_t - \vec{v}\vec{\nabla})^2 - c^2\Delta]\Phi(t,r,\theta) = 0 \\ & \Phi_{\omega,l}(t,r,\theta) = e^{-i\omega t}e^{il\theta}\phi_{\omega l}(r) \\ & (\frac{1}{r}\frac{d}{dr}(r\frac{d}{dr}) + V_{\omega l}(r))\phi_{\omega l}(r) = 0 \\ & (\mathbf{W}_r^{in}, W_l^{in}) \to (W_r^{out}, W_r^{out}) \end{split}$$

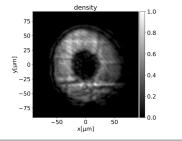






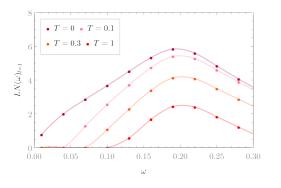


Where do we go from here?



Experiments with polaritons

- Measure Hawking radiation and rotational superradiance independently from one another
- Measure interplay between the two effects → modification of correlations?



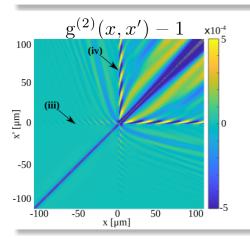
All optical experiments

- Measure phase and density \rightarrow access full field statistics and dynamics
- High resolution spectroscopy in 1 and 2D with and without rotation
- Homodyne detection to enhance signal strength and measure quantum correlations
- Enhance strength of emission and degree of entanglement by probing with squeezed state

I Agullo *et al* PRL **128** 091301 2022

F Claude et al PRL **129** 103601 2022.

PRB 107 174507 2023

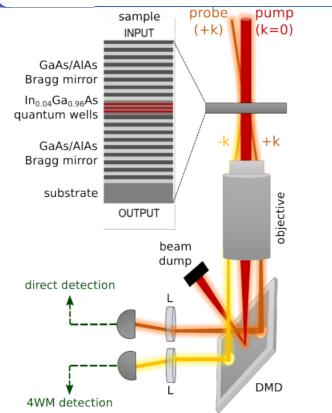


Numerical simulations

- New effect of quantum fields predicted: vacuum excitation of quasi-normal mode of acoustic field
- Good experimental configuration to observe strong correlations

Jacquet et al. PRL 130 2023, EPJD 76 2022

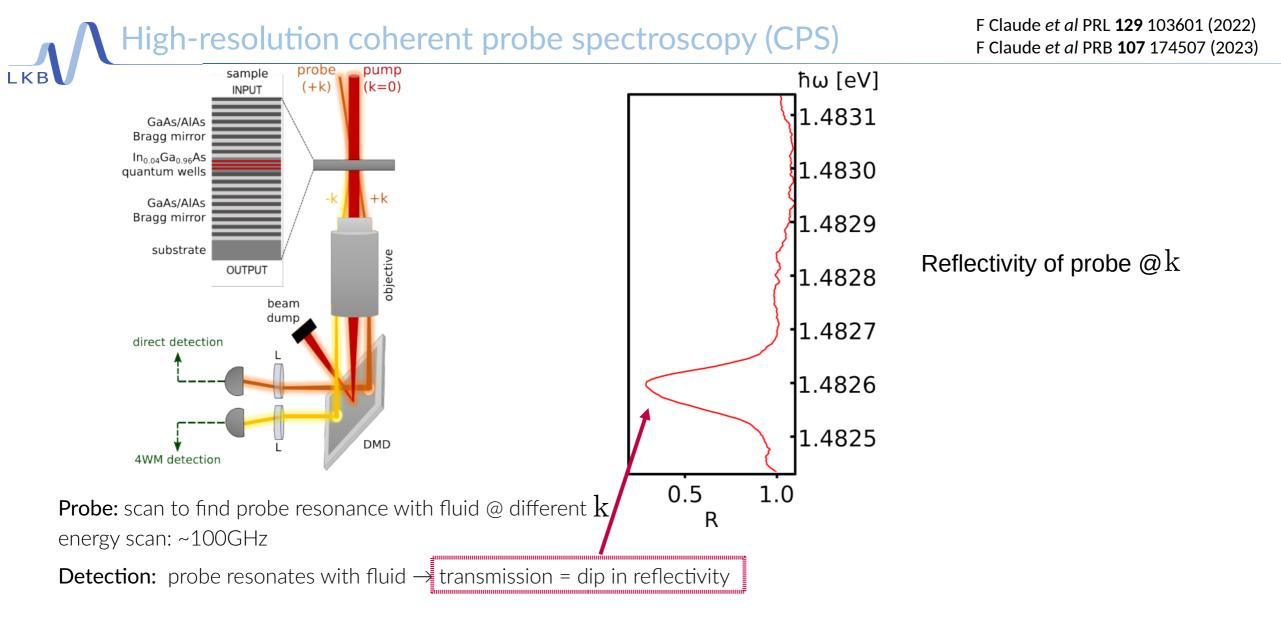
A High-resolution coherent probe spectroscopy (CPS)



Probe: scan to find probe resonance with fluid @ different k energy scan: ~100GHz

Detection: probe resonates with fluid \rightarrow transmission = dip in reflectivity

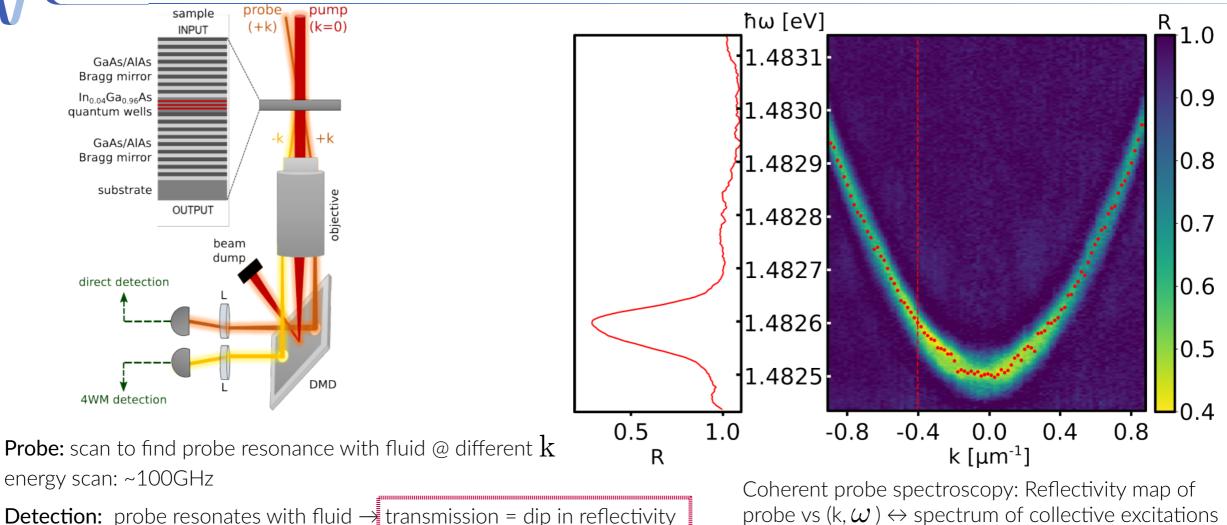
 ω resolution fixed by the probe laser linewidth (<250kHz) k resolution fixed by the k-space filtering (0.02um^-1)



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F Claude *et al* PRL **129** 103601 (2022) F Claude *et al* PRB **107** 174507 (2023)



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