

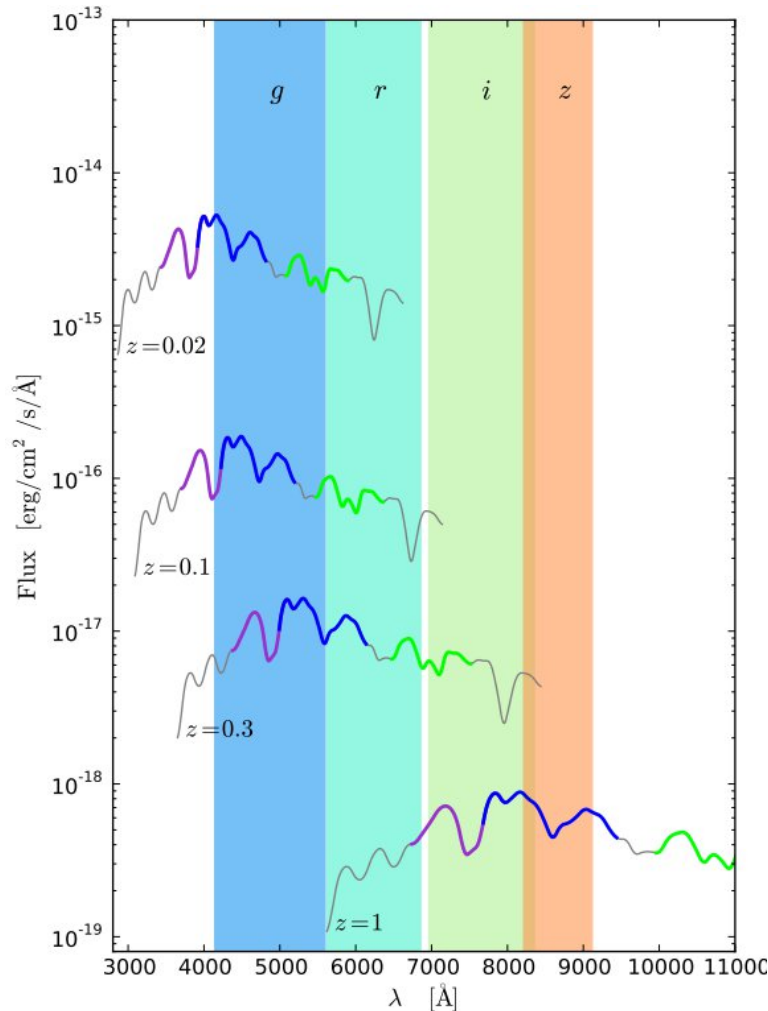


# StarDICE status overview

LSST-France

LPSC June 08th 2023

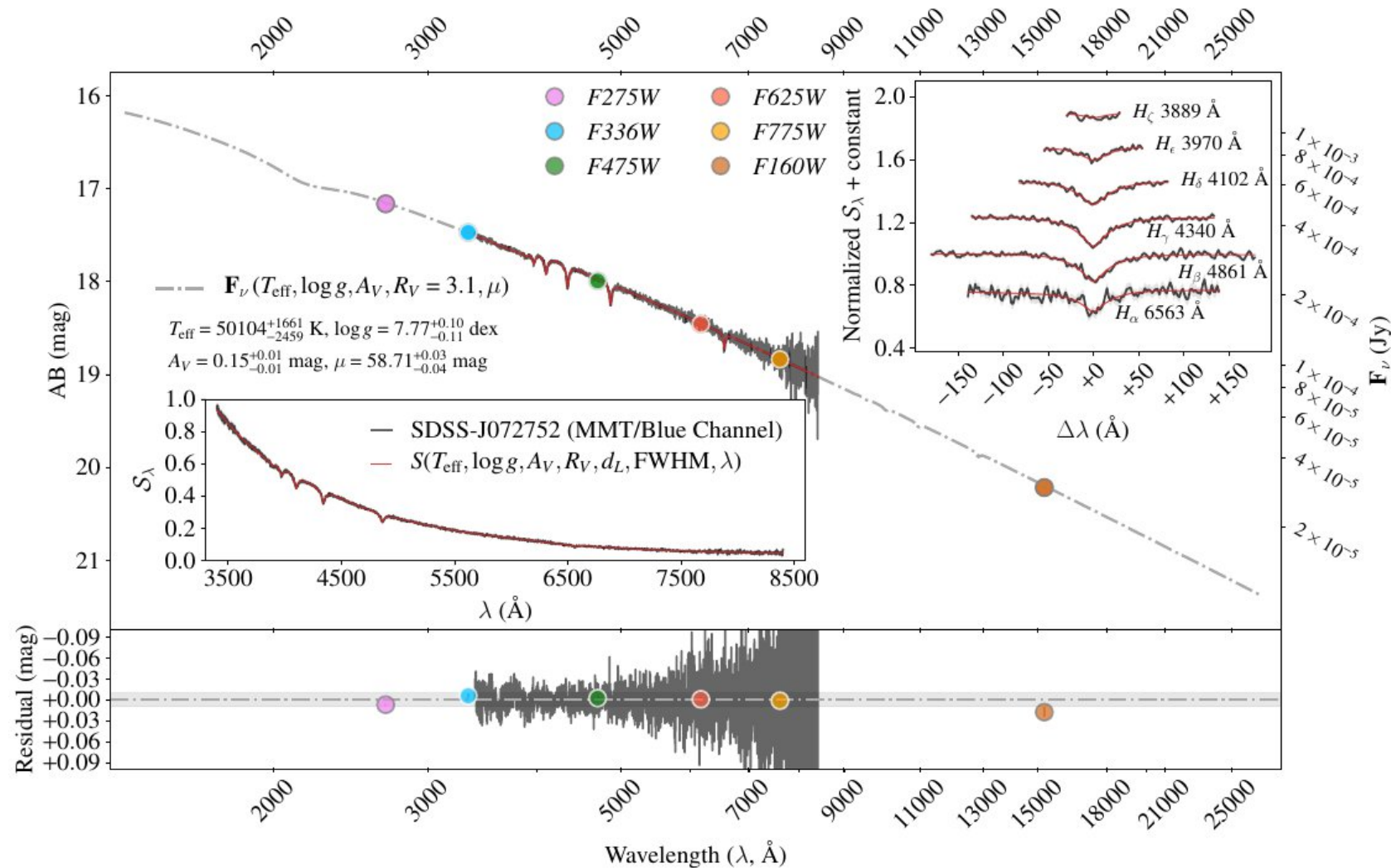
# Providing flux references for Supernovae distances



- Apparent flux of supernovae at different redshifts gives a noisy ( $\sim 7\%$ ) measurement of distance  $\rightarrow$  Hubble diagram
- VRO will provide thousands of SNe to beat this noise
- All SNe at a given redshift share one noise in common : The error on the survey passbands
- Large survey needs exquisite calibration references (standard stars)
- The goal of StarDICE is to provide such standards

# The only practical references to date are pure hydrogen WD

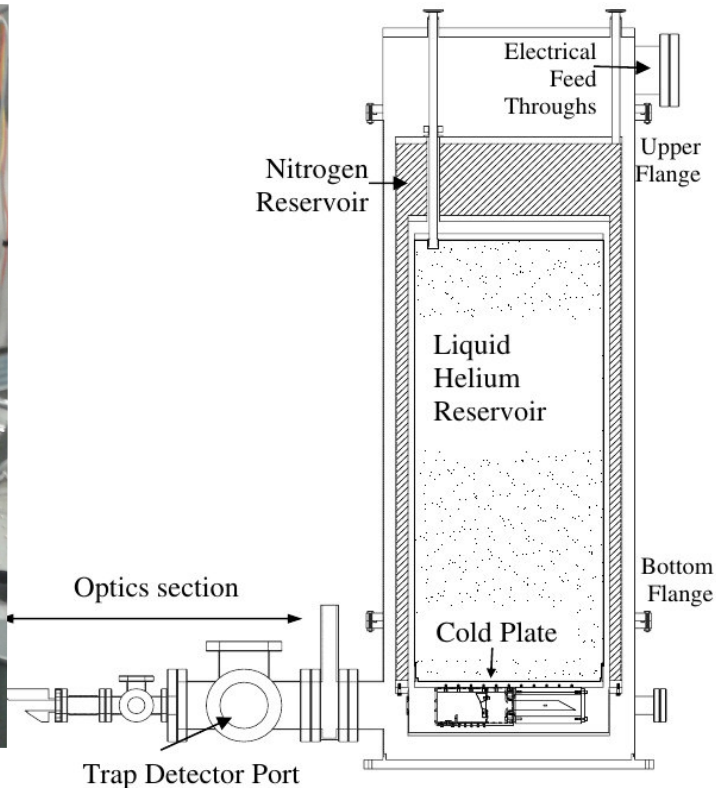
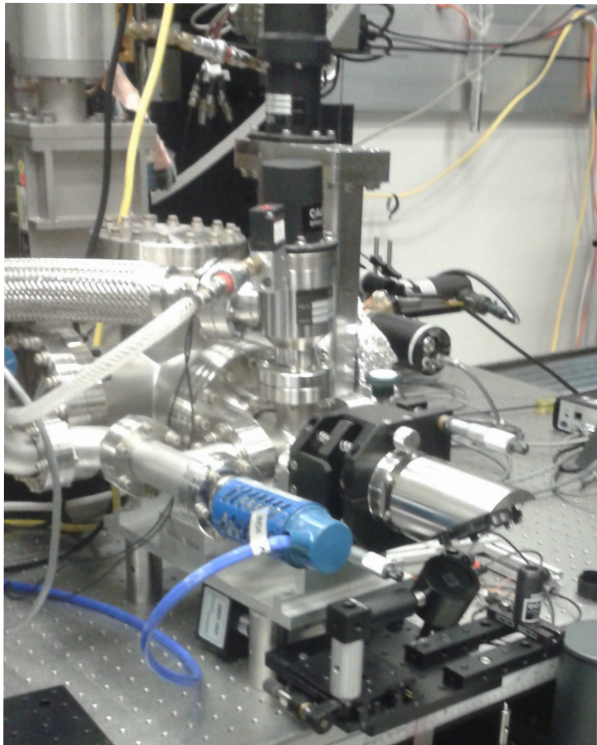
- Numerical model of the radiative transfer in the hydrogen atmosphere
- Model parameters inferred from measurement of H profile in high-resolution spectroscopy
- Largest implementation to date : The cosmic flux standards program (Narayan et al. 2022)



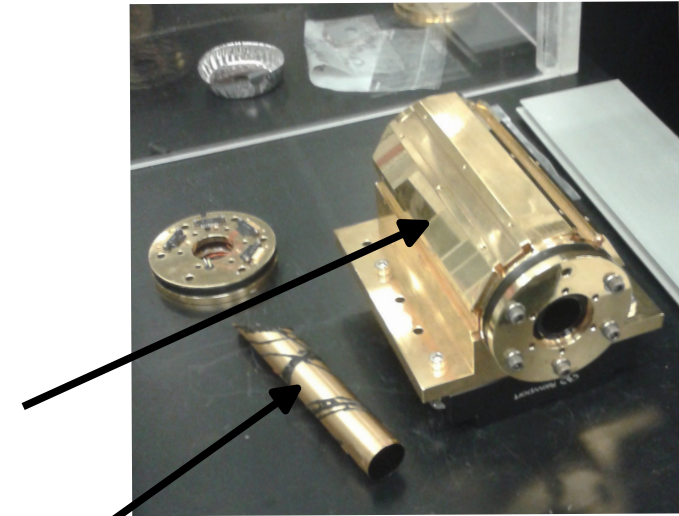
The proposal of the StarDICE experiment is to build an instrumental check of this model

# The alternative standard

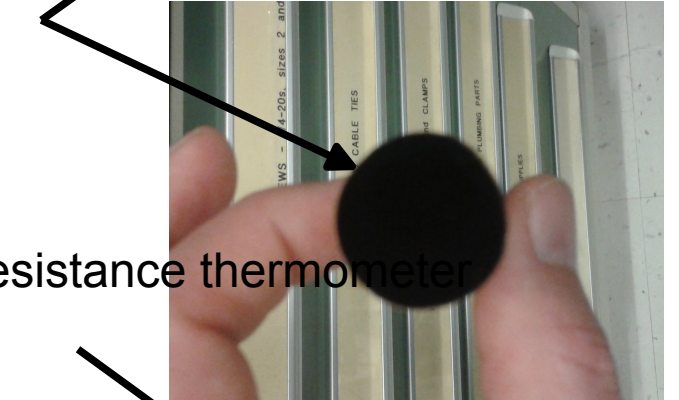
POWR: the Primary Optical Watt Radiometer  
(Brown et al. 2006, Houston et al. 2006)  
high-accuracy electrical substitution cryogenic radiometer



Cryogenic shelter

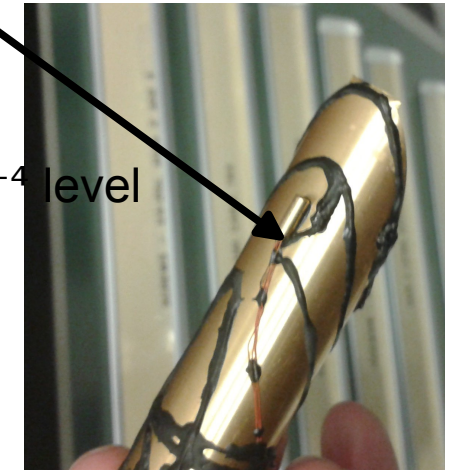


Black absorbing cavity

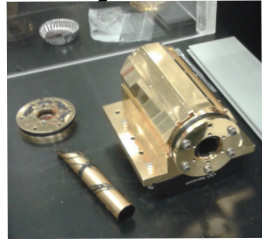


Germanium resistance thermometer

Claimed accuracy at the  $10^{-4}$  level



# The original concept proposal for a path finder (2016-2019)

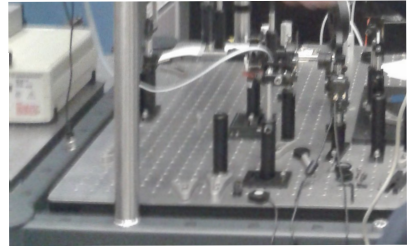


NIST POWR

$10^{-3}W$

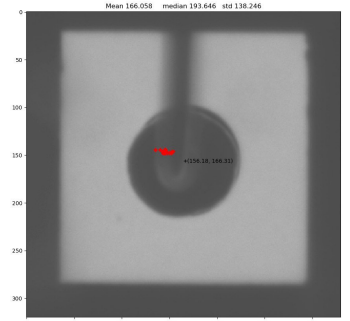
$10^{-6}W$

NIST SCF



NIST Photodiodes

Narrow spectrum LEDs with temperature monitoring

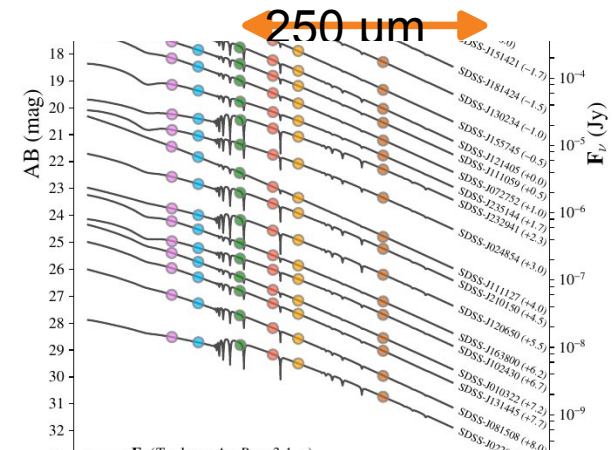


Small aperture telescope

$10^{-17} W/cm^2/nm$

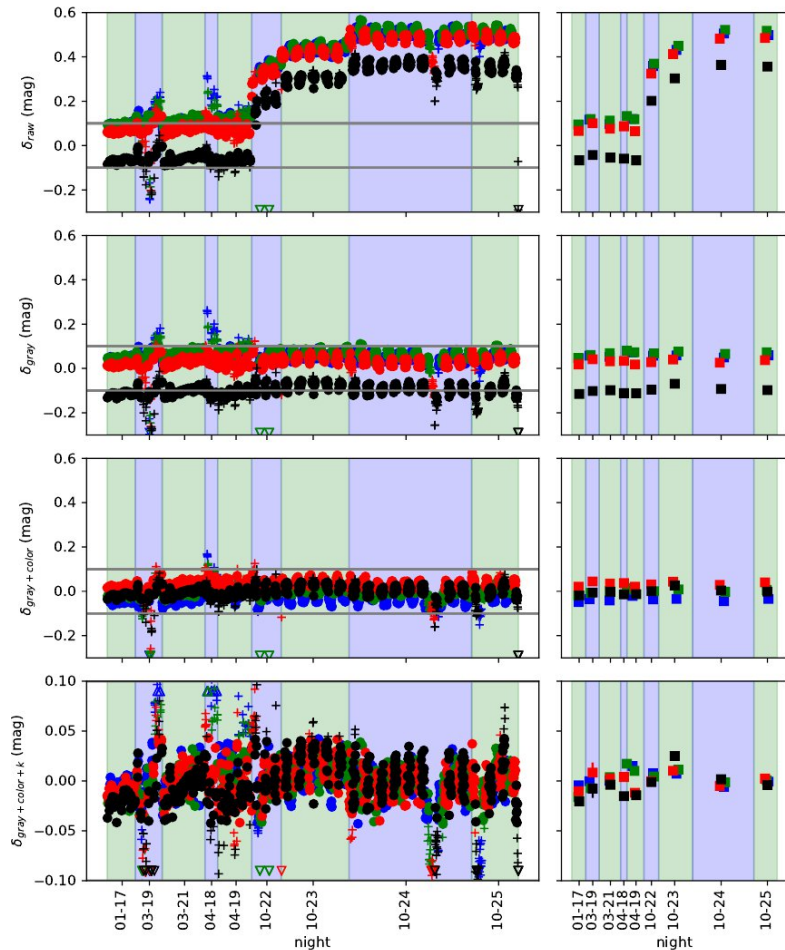
$10^{-19}W/cm^2/nm$

Spectrophotometric standard stars



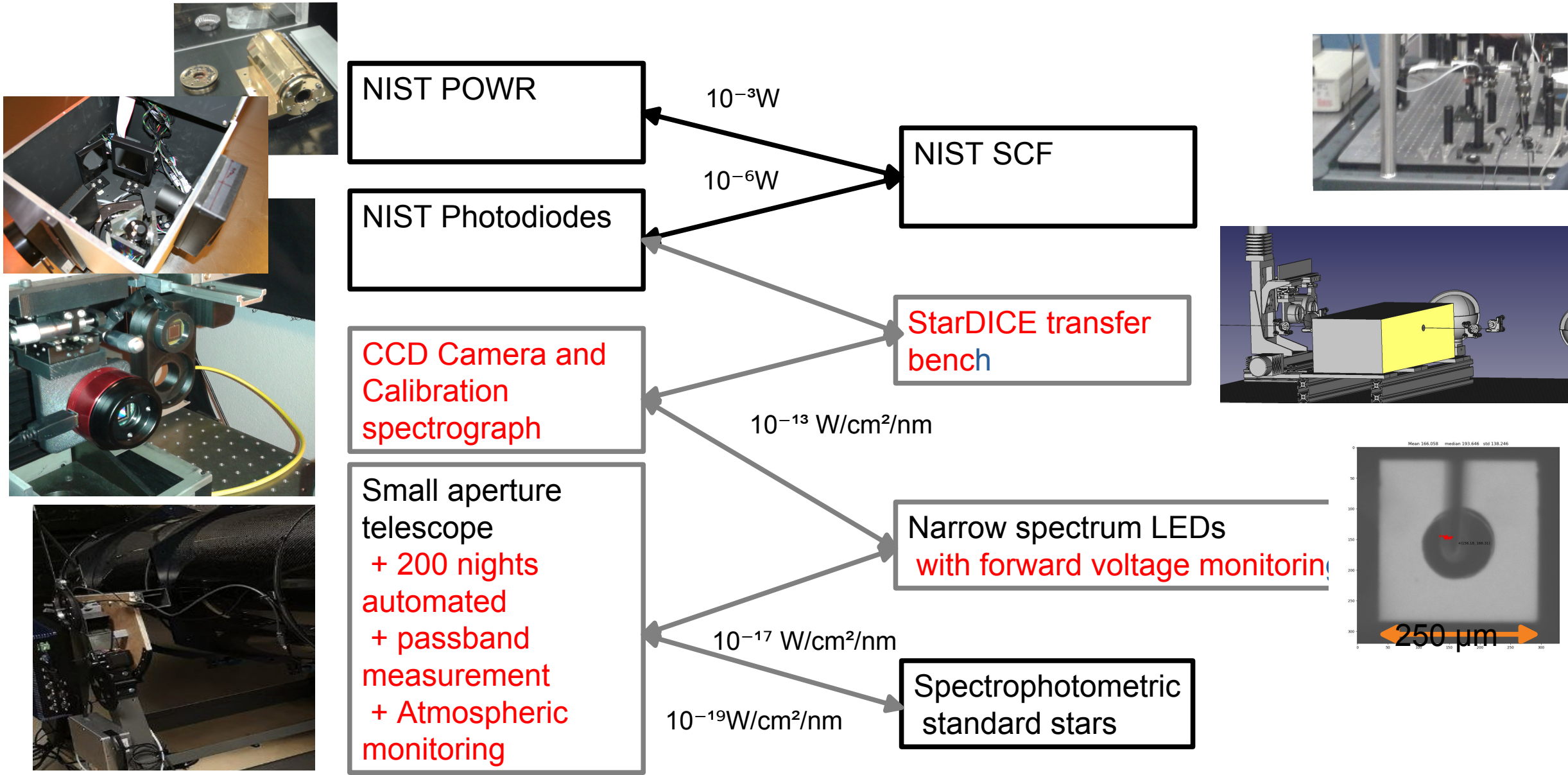
250 um

# Conclusion from F. Hazenberg thesis (2019)



- Optically perfect : Full pupil and point like which makes the result absolutely Independent of any optics model
- Fast : full telescope calibration in 20min
- Stable : consistent results over 6 monthes at the % level
- Lessons learned for the **upgrade**:
- Current T monitoring accurate at .5%: **new T proxy**
- One step missing in the transfer: **Calibration transfer bench**
- Transfer hindered by clouds: **IR instrument**
- Estimated number of nights 200-400: **Robotic Telescope**
- Accurate determination of mirror reflectivity: **Collaboration with CBP**
- LEDs spectra at nominal (extremely low) flux not acquired: **Spectrograph**

# The revised StarDICE metrology chain



# StarDICE Milestones

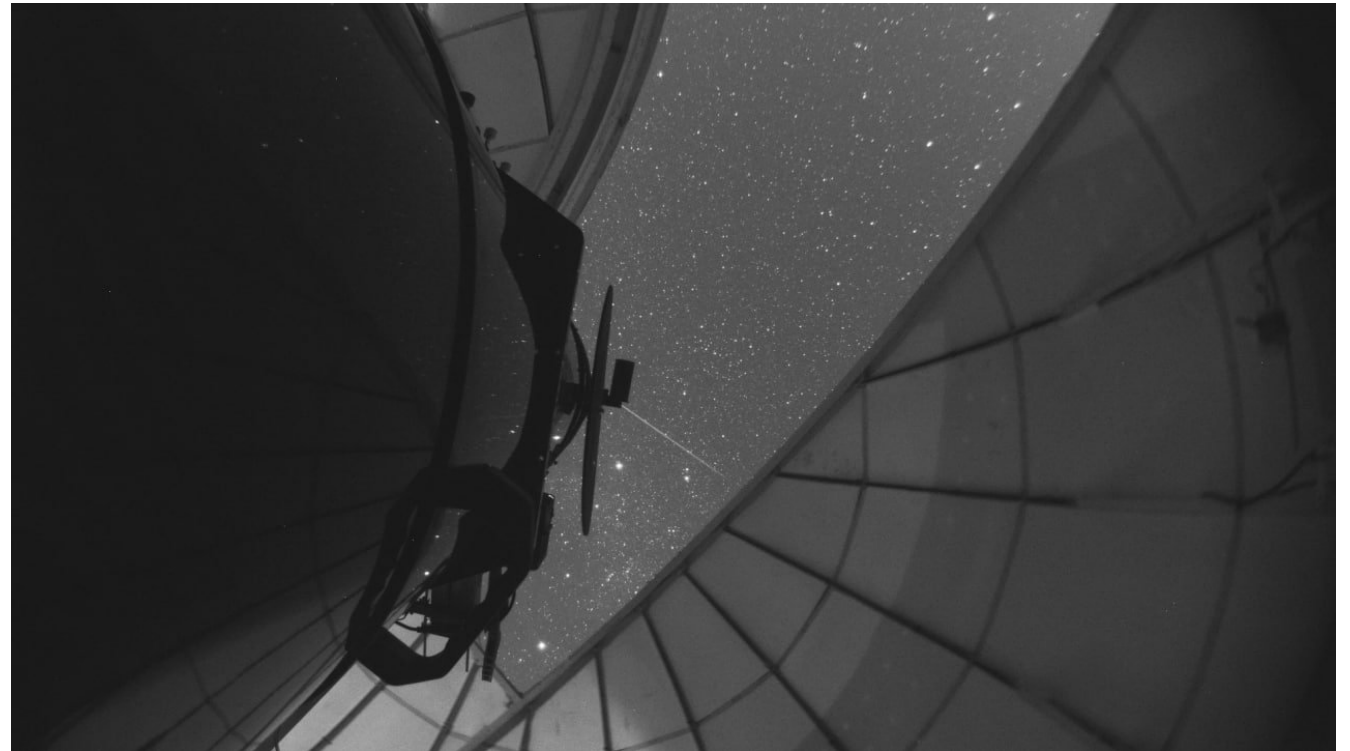
- Sept. 2019: F. Hazenberg PhD thesis. The complete experiment is designed and requires ~200 - 400 nights
- 2020 Sept. 25 : MOU with OHP to host the experiment in one of its jumelé coupola, parallel work starts on the building, mount, optical instrument, calibration bench and artificial star.
- 2022 Mar. : CBP measurement of the optical instrument complete
- 2022 Nov. 11 : Acceptance of the StarDICE bench demonstration paper
- 2023 Mar. : Robotic telescope complete with optical and IR instrument
- 2023 Mai 28 : OHP review grants green light for remote operations



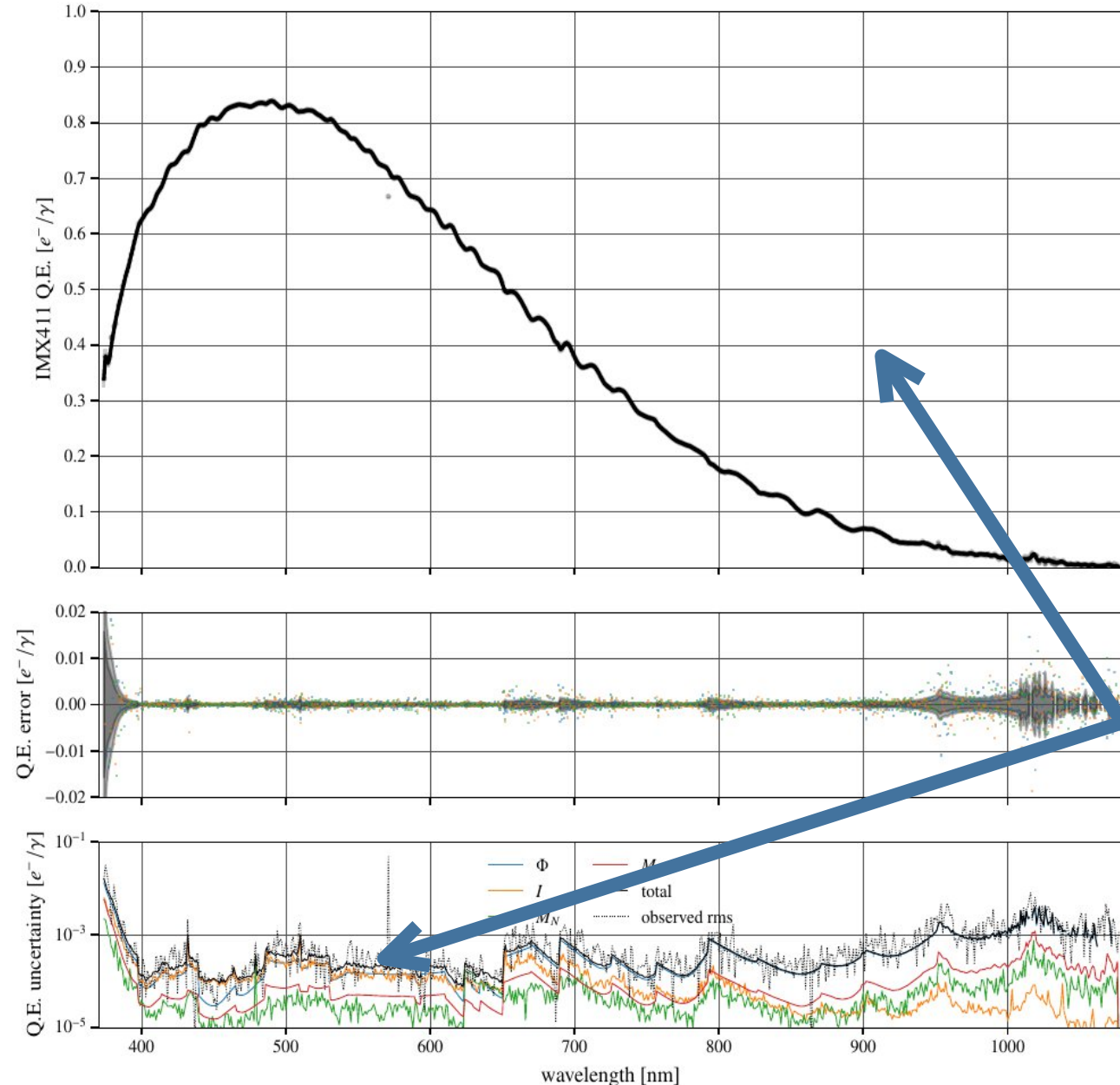


# From then to now

- CPPM** : O. Angelini, S. Beurthey, S. Deguero, F. Feinstein
- LPNHE** : P. Antilogus, Ph. Bailly, E. Barrelet, M. Betoule, S. Bongard, J. Coridian, M. Dellhot, P. Ghislain, A. Guyonnet, F. Hazenberg, C. Juramy, H. Lebbolo, L. Le Guillou, E. Pierre, N. Regnault, Ph. Repain, M. Roynel, K. Schahmaneche, E. Sepulveda, T. Souverin
- LUPM** : J. Cohen-Tanugi, Eric Nuss, B. Plez, Kélian Summer
- IJCLab** : J. Neveu, S. Dagoret-Campagne, M. Moniez
- OHP** : Pierre-Eric Blanc, Auguste Le Van Suu, Jean-Paul Payan, Jean Claude Brunel, François Dolon, Marc Ferrari

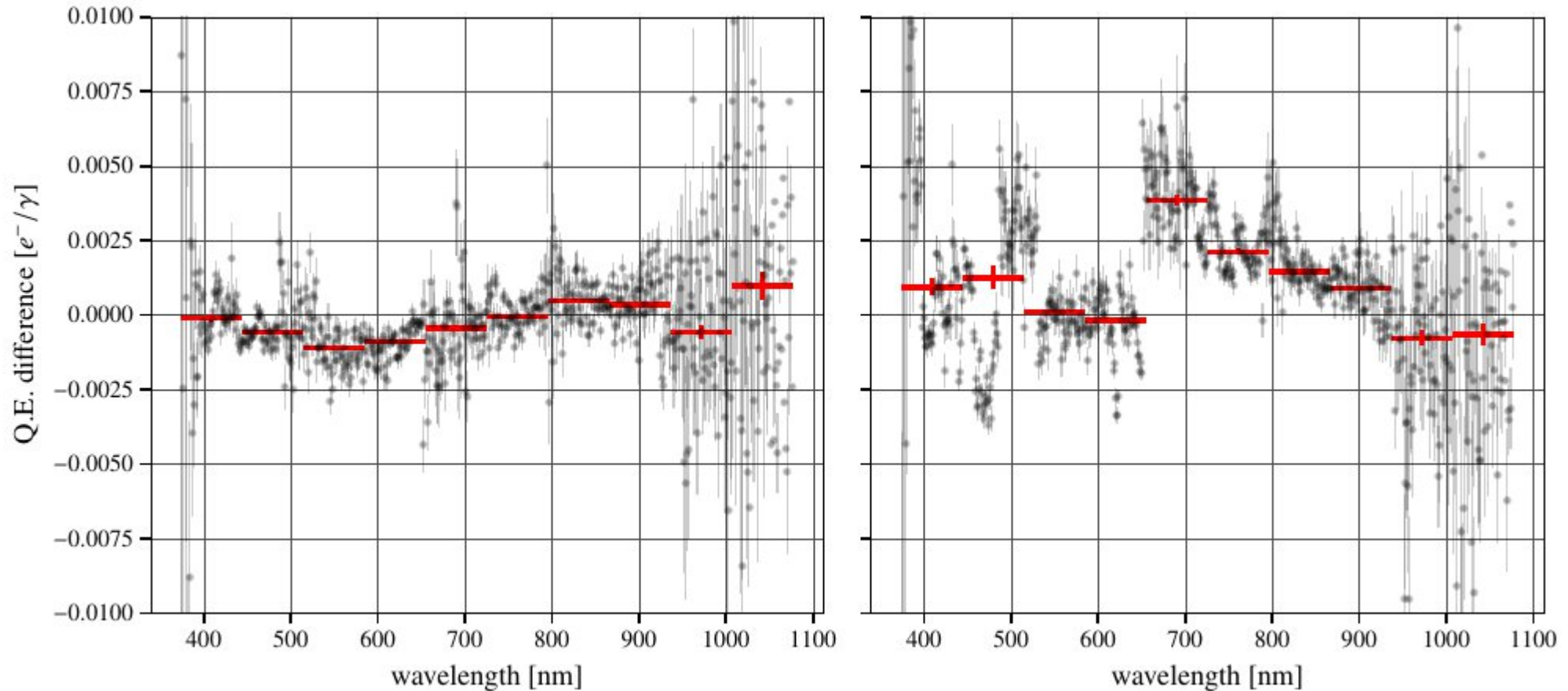


# Calibration bench



- The game is to calibrate a CCD sensor relative to the NIST Photodiodes at  $10^{-3}$  over the range 400-1000 nm
- Already validated performance on v2 are sufficient for the validation survey
- A&A 2022\_44973
- Work on v3 has started
  - Optimization of photodiode reading
  - Optimization of photocurrent extraction
  - Rework of the mechanical components
  - New CCD camera

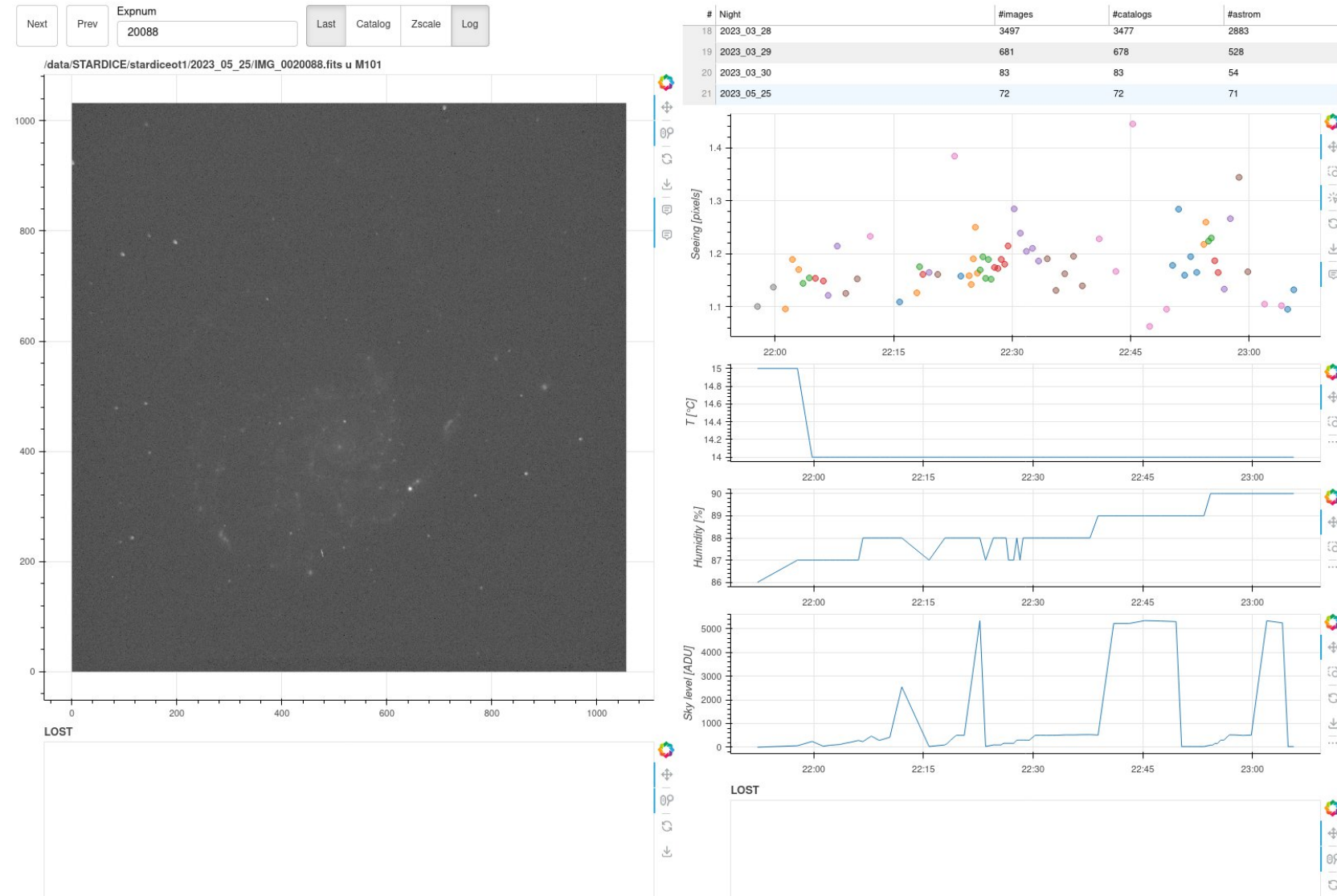
# Calibration bench systematic



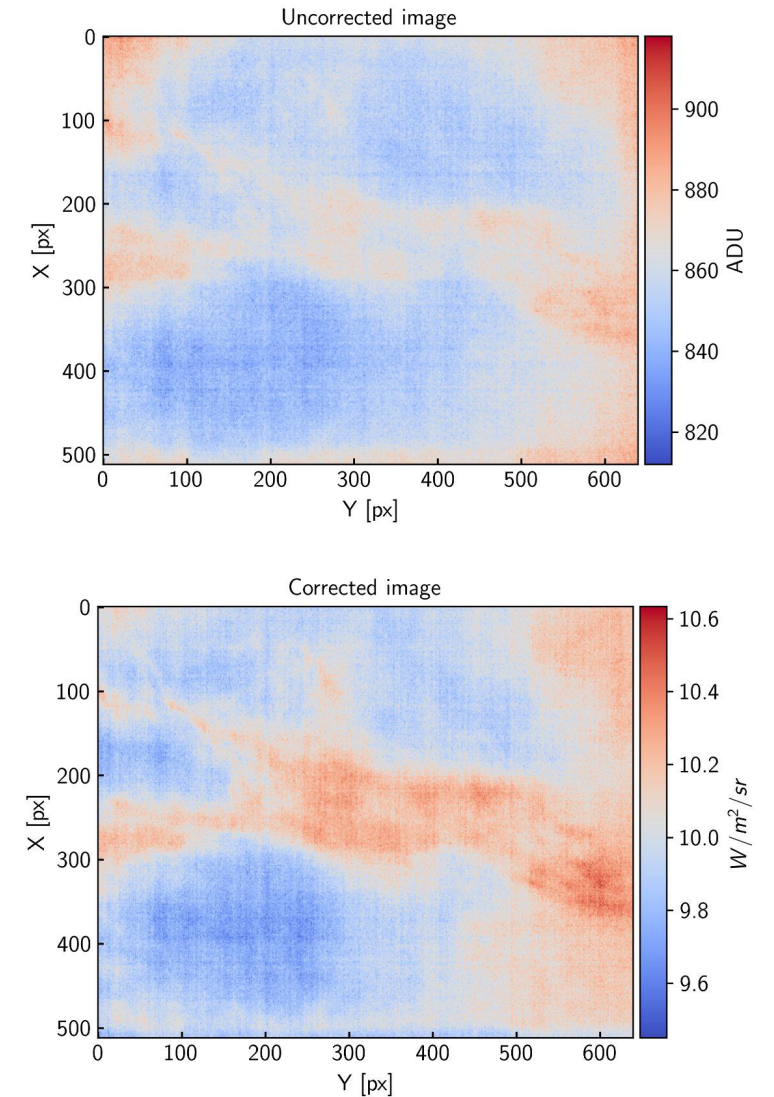
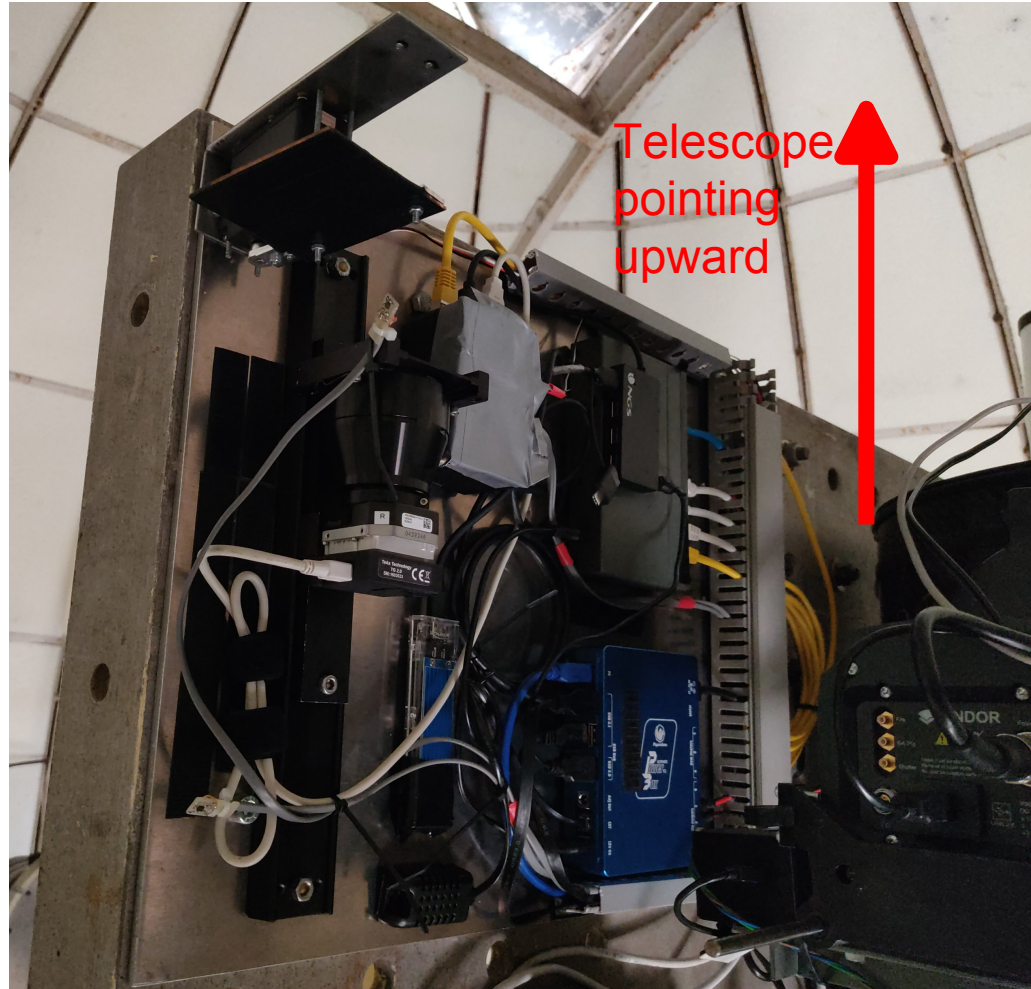
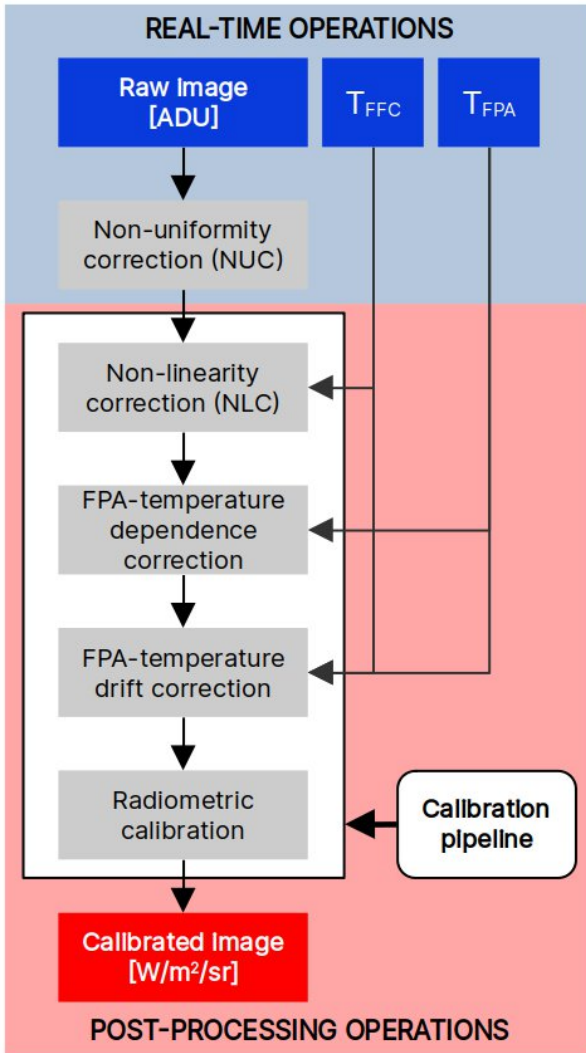
- Main identified systematics is the linearity
- But at a non-critical level for the survey validation
- Still a bit more work to have a fully calibrated artificial star

# Robotic instrument is ready

- Screenshot from the remote control
- 6 good nights accumulated
- Start a 20 night validation survey as soon as weather (and network) cooperate.
- See Jeremy's talk for details on the transmission measurement of the optical instrument
- And Thierry's talk for a first evaluation of the performances of the optical instrument

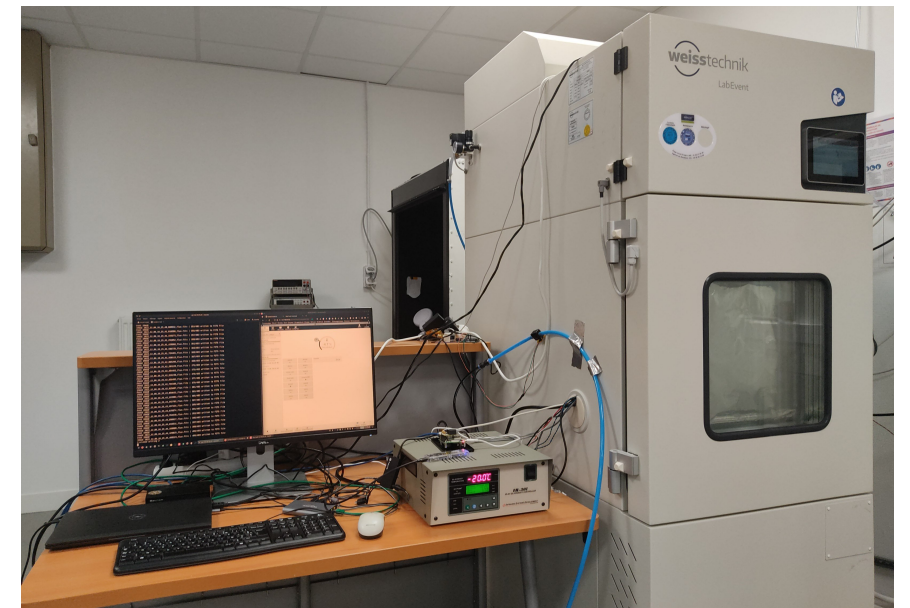
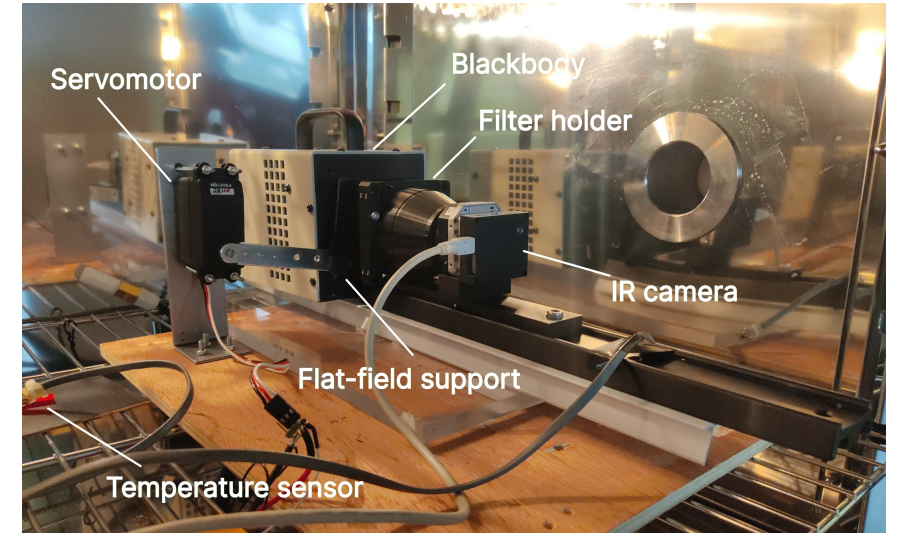
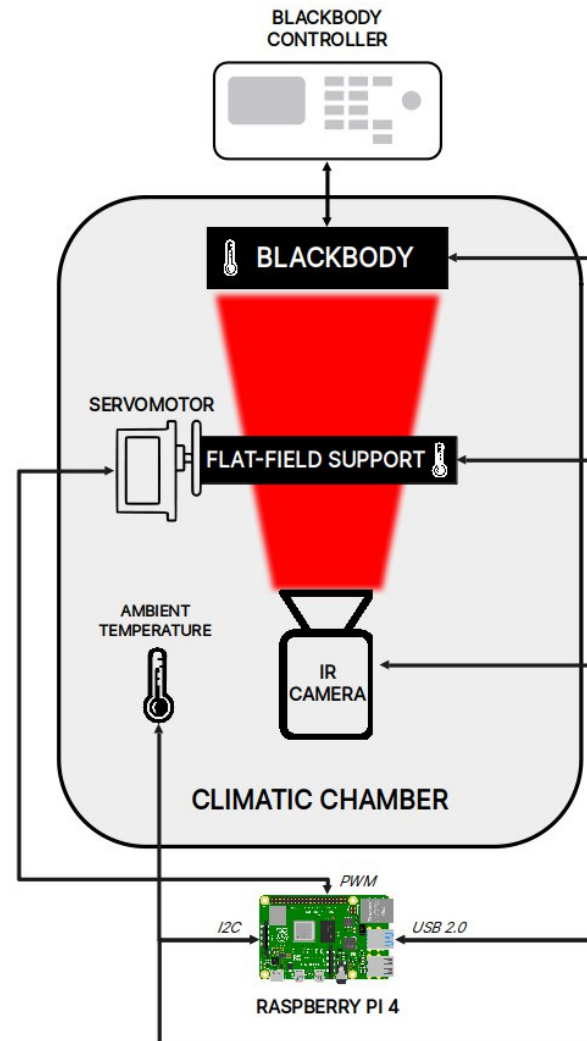


# Installation of the IR instrument at OHP



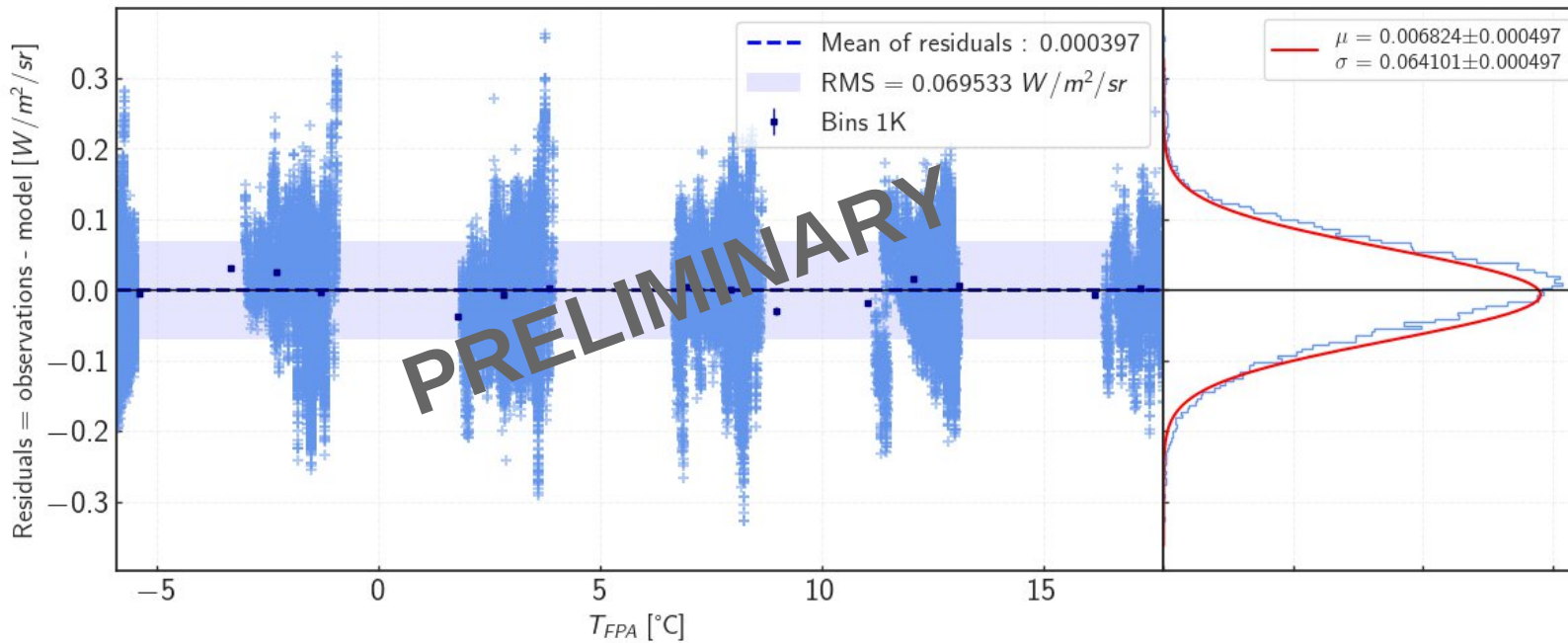
# Calibration of the bolometer array (1/2)

- **December 2022** : calibration data acquisition at IJCLab in ATLAS climatic chamber with S. Dagoret-Campagne and M. Moniez
- **Goal** : imaging a low-temperature/radiance blackbody at different camera operating temperatures
- → correct for various camera defects (non-linearity response, sensor temperature drift...)

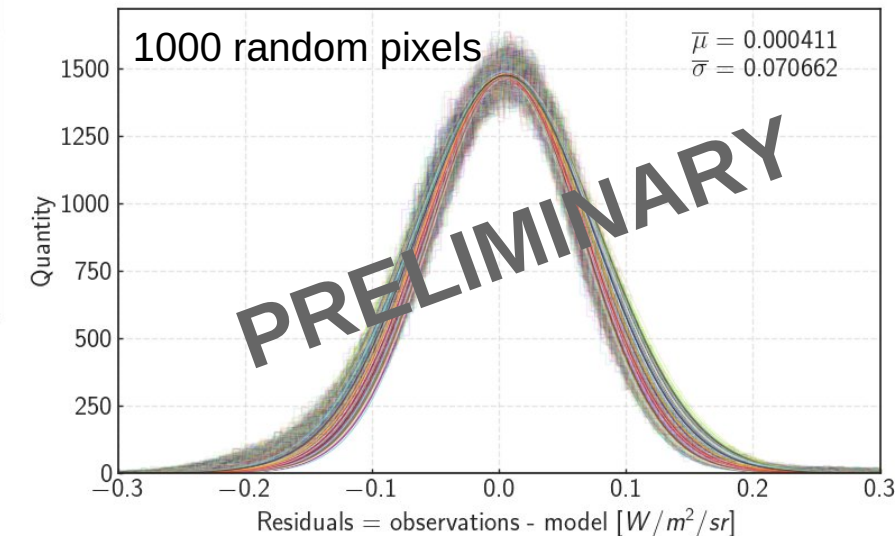
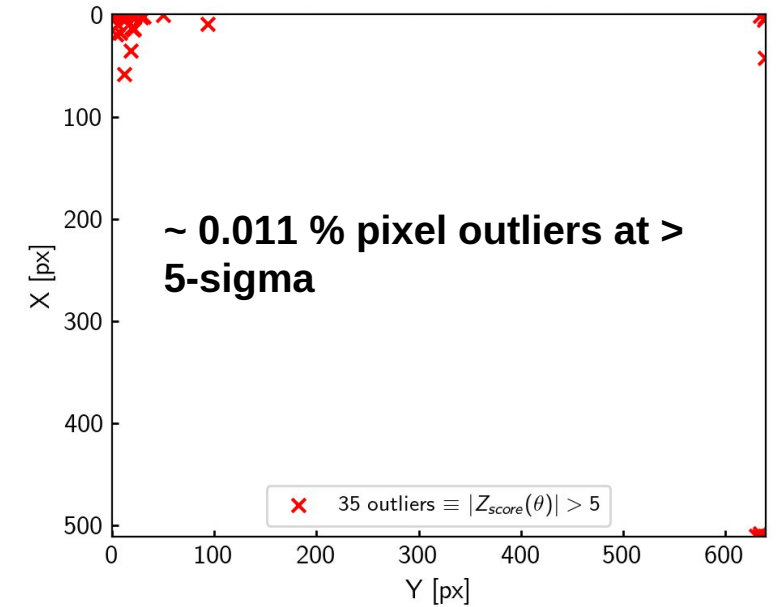


# Calibration of the bolometer array (2/2)

- Preliminary average uncertainty per pixel :
- **+/- 0.07 (stat.)  $W/m^2/sr$**
- **+/- 0.04 (syst. of calibration experiment)  $W/m^2/sr$**

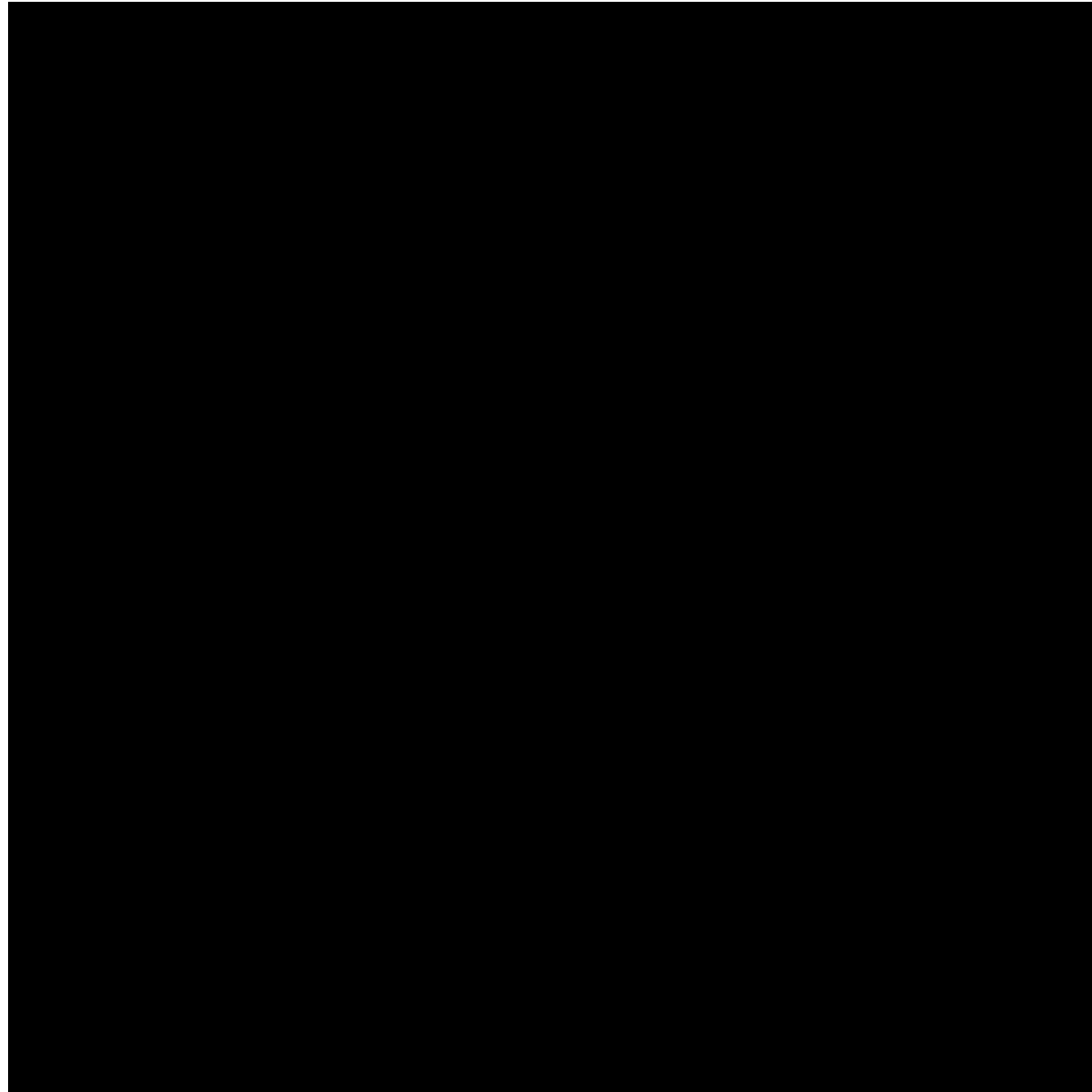


Residuals for the central pixel of the focal plane array between the corrected images and the scene radiance model. Temperature gaps are due to environmental conditions in the climatic chamber. [K. Sommer et al., in prep]





First light at OHP



# Conclusion

- StarDICE robotic installation is up and running
- Up for a 20-night validation survey allowing to build the data analysis chain for the spectrophotometric and IR instruments
- Optimizing the observation strategy to learn the best way to beat the atmospheric noise
- On the hardware side, the focus is switching toward the finalisation of the calibrated artificial star