

Direct search for dark matter axion with MADMAX



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- 1- Scientific context
- 2- MADMAX: principles and prototyping phase
- 3- Technological developments (inc. IN2P3 participation)
- 4- Conclusion

Journées R&T IN2P3, 07 novembre 2023

How to see the axion ?

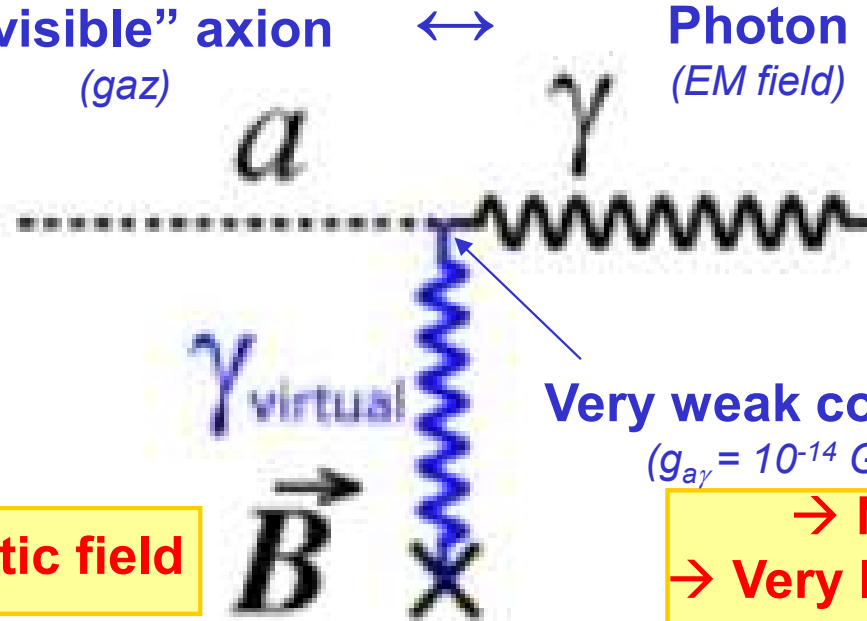
□ Convert it to photon in a magnetic field

Exp. constraints

$\mu\text{eV} \approx 10^{-15} \text{ M}(\text{proton})$
"Invisible" axion
(gaz)

GHz
Photon
(EM field)

$\approx \text{RF, micro-wave}$



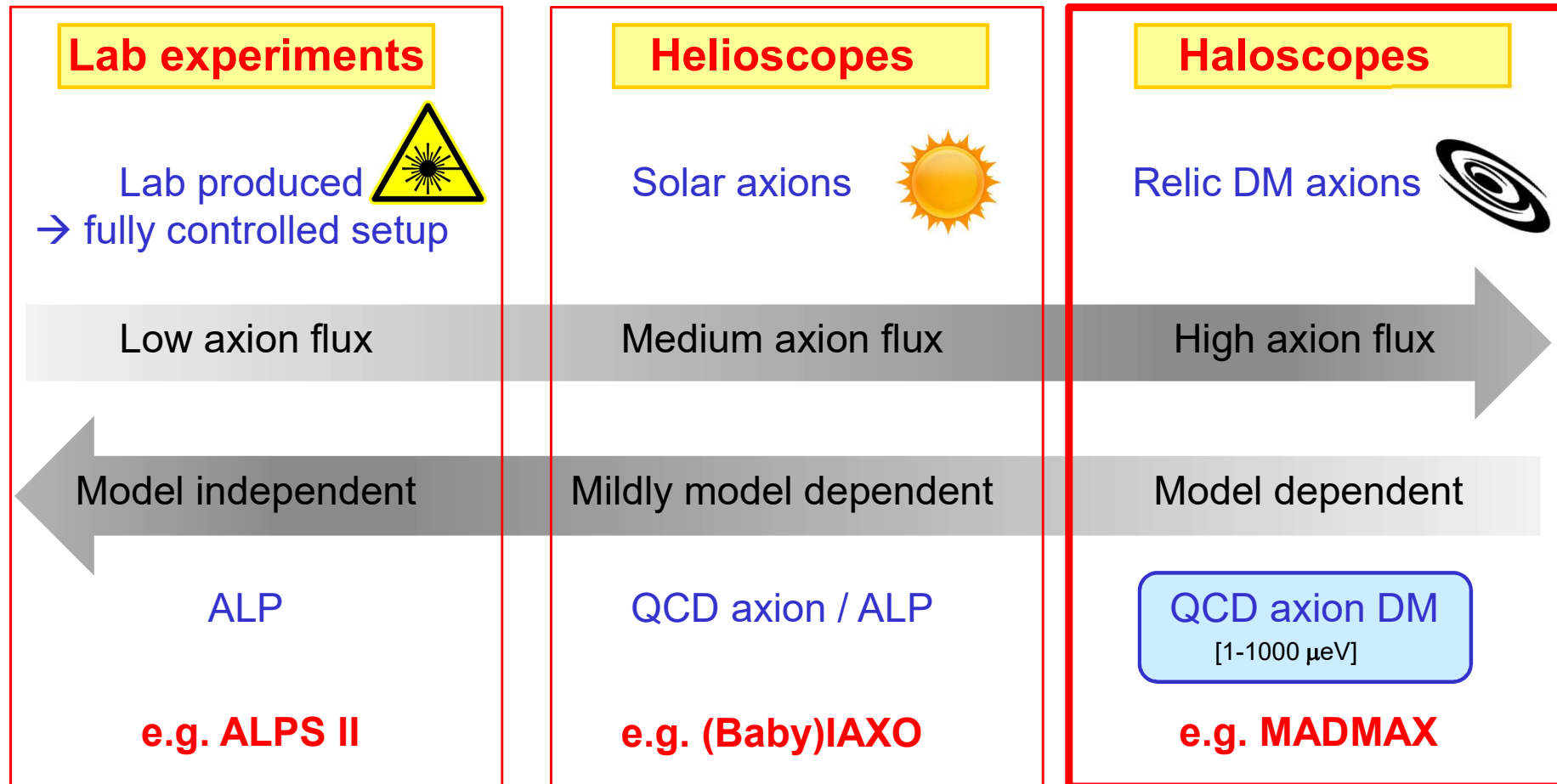
High Magnetic field

Very weak coupling: feeble signal
($g_{a\gamma} = 10^{-14} \text{ GeV}^{-1}$ for $m_a = 100 \mu\text{eV}$)

- Need boost
- Very low temperature
- Low Noise Amplifier

Axion search very rich in experimental challenges

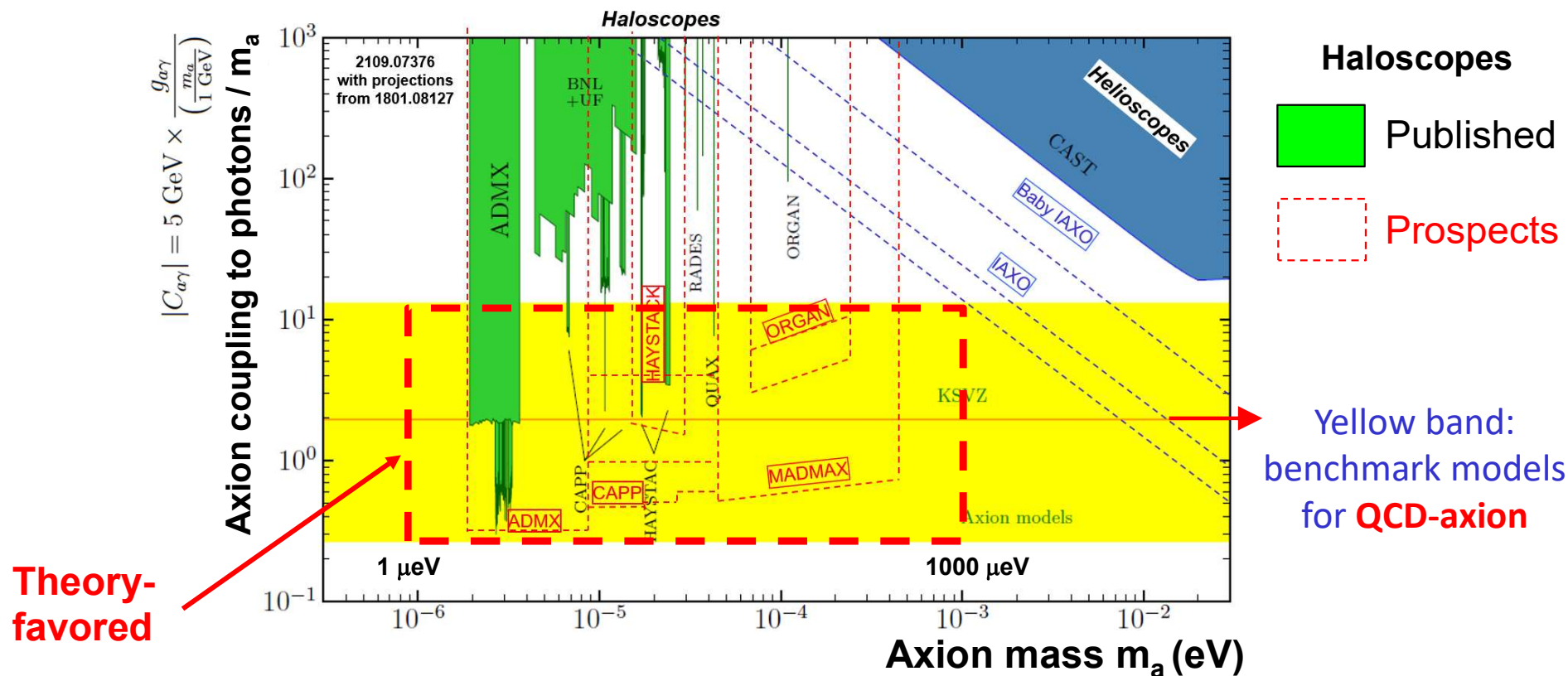
Axion/ALP direct searches



Complementary experimental approaches (all present in DESY axion hub)

DM axion search: status / prospects

□ Haloscopes = main way to search for Dark Matter axion



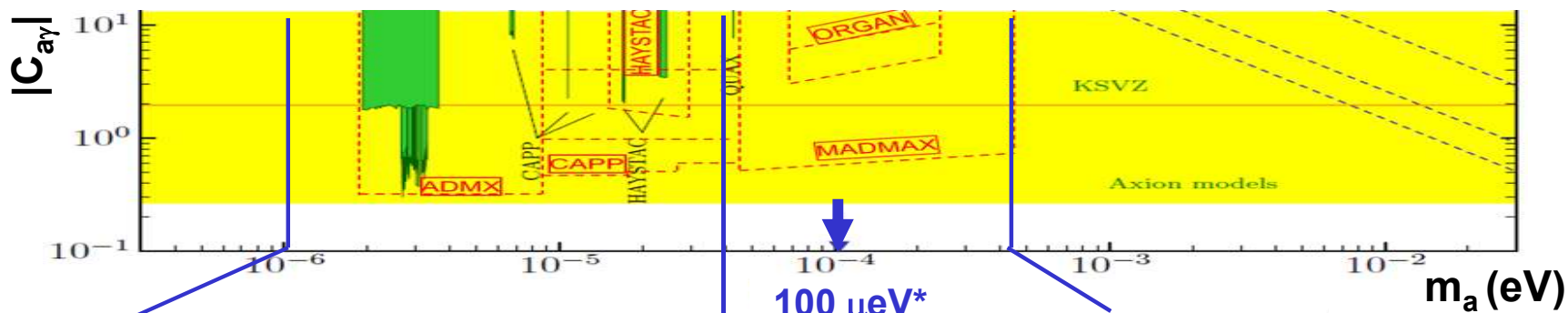
- Only very few experiments **currently** probe a (very small) part of the QCD axion phase space
- Vast **R&D program** to improve signal sensitivity and expand range of axion mass search

Rising interest (techno improvements + LHC/WIMP results) : next decades promising

DM axion search: how?

Experimental challenges for haloscopes

- Convert axions into photons [E field of $O(10^{-12} \cdot \frac{B}{10 T})$ V/m] \rightarrow high B_{field} [$B \gg 1 T$]
- Boost E_{field} [up to detectable $P \sim 10^{-22} W$] \rightarrow resonant set-up or large area
- Scan over range of axion mass \rightarrow tunable set-up [precision mechanics]



1 μeV [0.25 GHz]

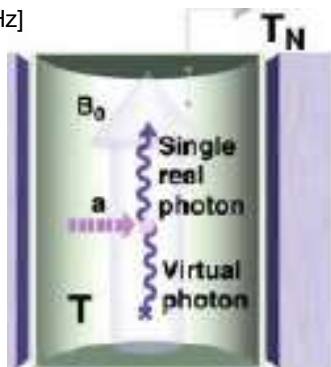
40 μeV [10 GHz]

400 μeV [100 GHz] [$v_a = v_\gamma$]

\rightarrow RF / Microwave regime

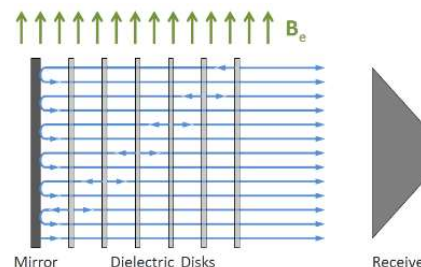
Cavities

[baseline concept 1983]
PRL51 (1983) 1415



- \triangleright Very high B
- \triangleright Quantum noise limit
- \triangleright Higher Q
- \triangleright Multi-cavities

Cavity too small
+ high noise



MADMAX

e.g. Dielectric haloscopes

[novel concept 2013]
PRD88 (2013) 115002

\rightarrow MADMAX can probe the theory-favored range $m_a \sim O(100) \mu\text{eV}^*$

*Nat. Com. 13 (2022) 1, 1049 : $40 < m_a [\mu\text{eV}] < 180$

MADMAX (1/2)

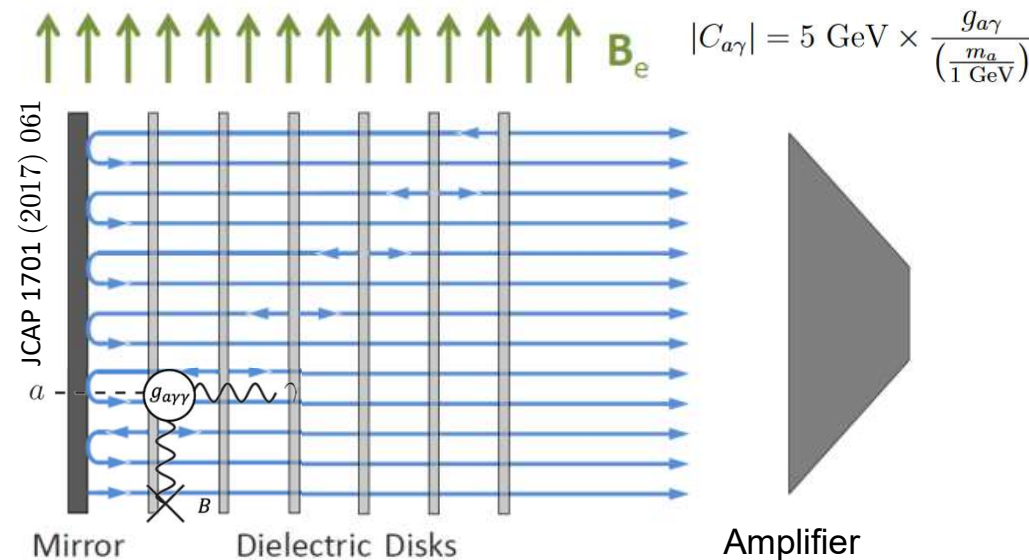
White Paper [EPJC 79 (2019) 186, 1901.07401]

□ A novel experimental concept: dielectric haloscope

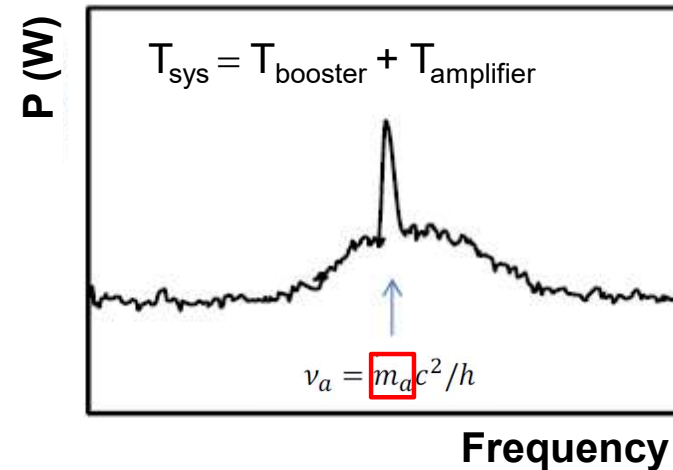
- **Constructive interference** of coherent photon emissions at dielectric layer surfaces
+ **resonant** enhancement (~leaky resonant cavities) : **boost (β^2)** signal wrt mirror only

$$P_{sig} = 10^{-22} \text{ W} \times \left(\frac{\beta^2}{50000}\right) \times \left(\frac{B_e}{10 \text{ T}}\right)^2 \times \left(\frac{A}{1 \text{ m}^2}\right) \times C_{a\gamma}^2$$

$$P_{sig}^{detect.} = 10^{-22} \text{ W} \times \left(\frac{SNR}{5}\right) \times \left(\frac{T_{sys}}{4 \text{ K}}\right) \times \left(\frac{4 \text{ days}}{t}\right)^{1/2}$$

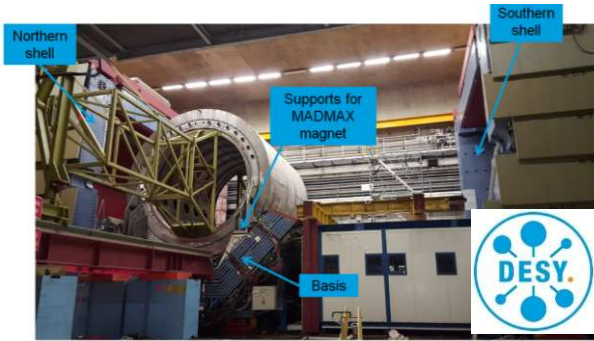


Thermal Noise

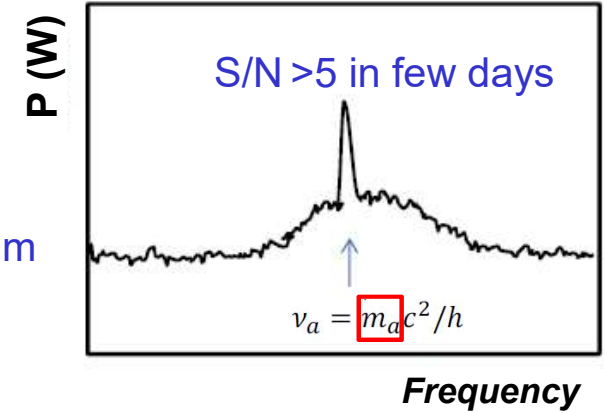
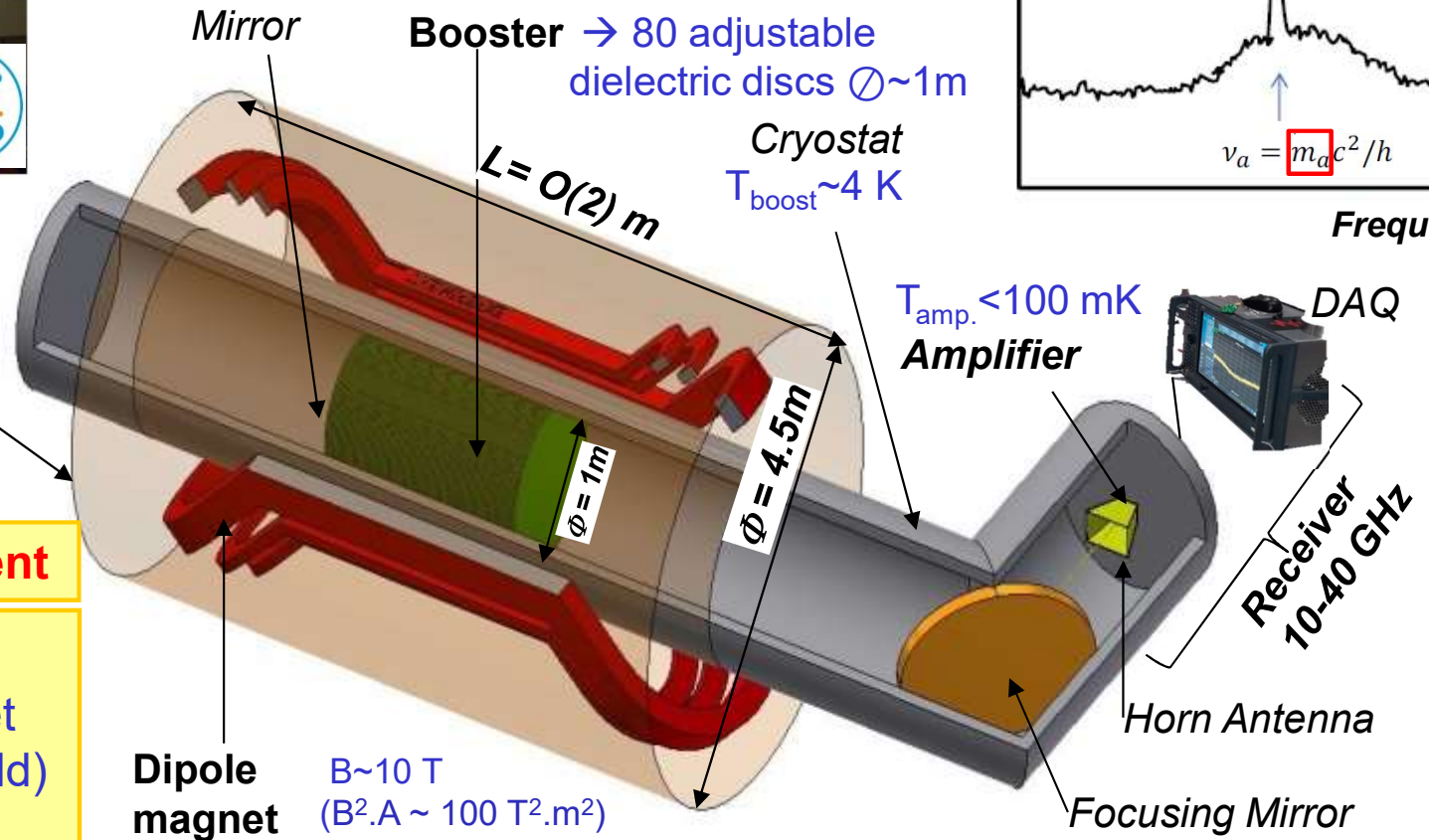


MADMAX (2/2)

Formed in 2017. 10 institutes: French (2), German (6), Spanish (1) and US (1) → ~50 people



Experiment location: HERA
in former H1 iron yoke



1st generation experiment

3 main challenges :

- High field dipole magnet
- Receiver (10's GHz, cold)
- Booster (cold, B field)

→ Start with prototyping phase to validate booster concept: cutting-edge R&D

Prototyping phase strategy

❑ **CERN lends us the world largest warm bore dipole magnet** [Morpurgo]

- Usage by MADMAX during YETS approved by CERN RRB under CPPM impulse

❑ **Address the two main challenges to develop booster concept**

- Understand Radio Frequency (RF) response in O(10) GHz regime → Calibrate boost factor
- Move the disks at μm level precision at cold and under high B-field

| Name | Goal | Type | Made of | Avail. | Test | Room Temp. Cold (10 K) |
|-------|--------------------------------------|-------------------|--|--------|---------------------|---------------------------|
| P200 | Piezo-motor + mechanics | Open booster | 1 moveable disk $\phi = 200$ mm | 2022 | 2022 | |
| CB100 | RF studies + First physics | Closed booster | 3 fixed disks $\phi = 100$ mm | 2021 | 2022, 23, 24 | |
| CB200 | RF studies + First physics | Closed booster | 3 fixed disks $\phi = 200$ mm | 2023 | | 24 |
| OB300 | Scan ALP around $80 \mu\text{eV}$ | Open booster | 3-20 moveable disks $\phi = 300$ mm | 2024 | | 25, 26? |

➔ **Gradually build the final booster design + do physics**

RF studies + ALP Physics (1/2)

| Name | Goal | Concept | Made of | Avail. | CERN test |
|-------|-------------------------------|----------------|---|--------|------------------------------|
| CB100 | RF studies + First physics | Closed booster | 3 fixed disks $\phi = 100$ mm | 2021 | 2022, 2023 10 hr, 21 days |

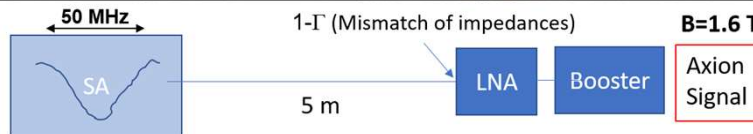
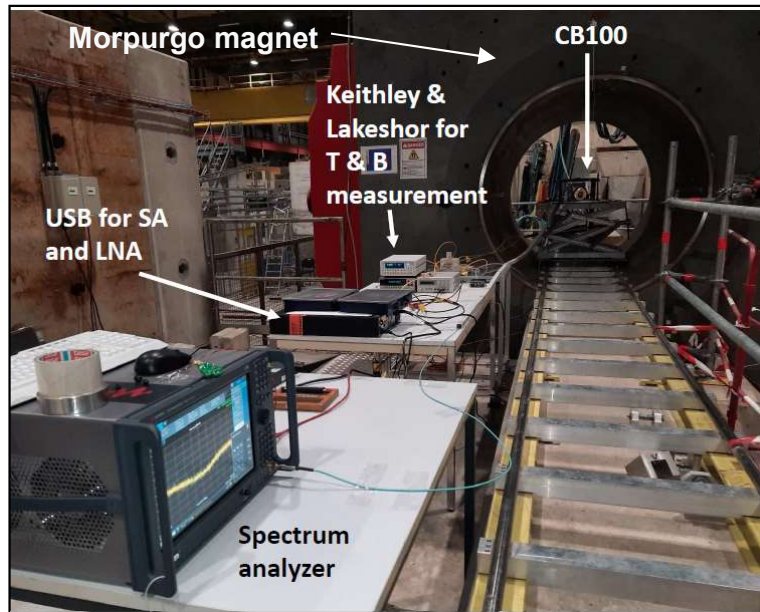
Booster

Mirror
+
3
sapphire
disks

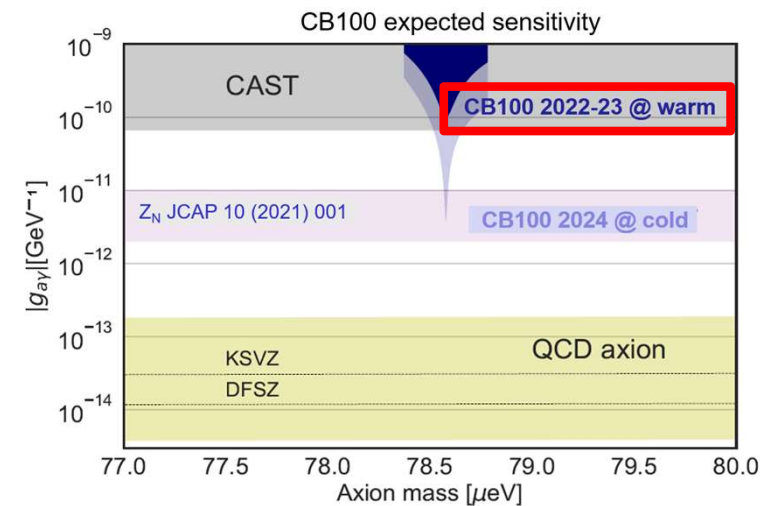
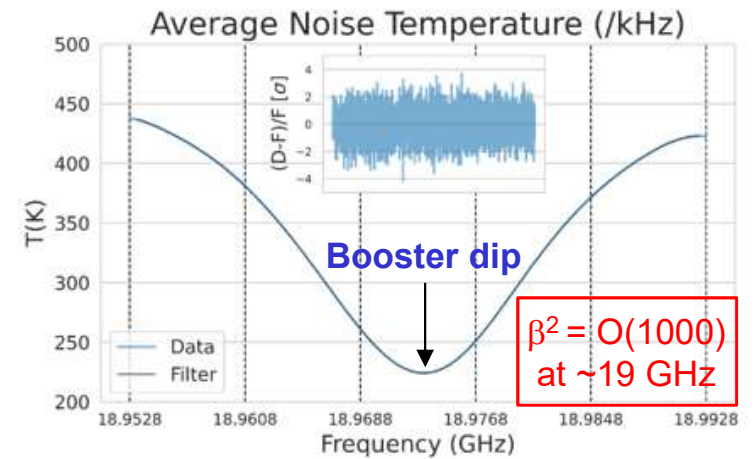
Parabolic
taper

Cylindrical
wave guide

Low Noise
Amplifier
(LNA)



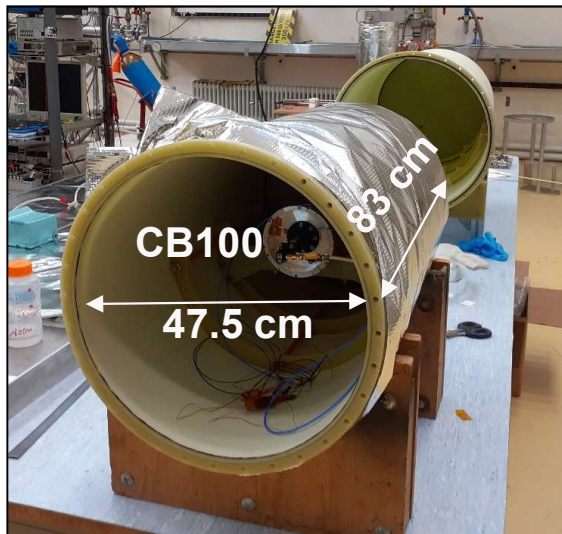
→ First ALP limit [paper in preparation]



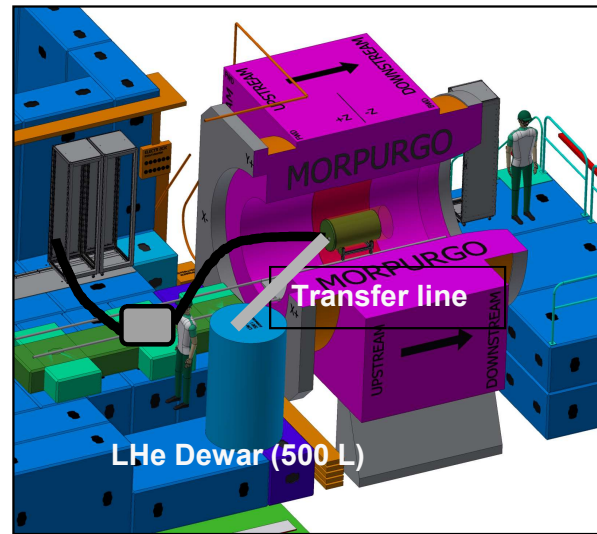
RF studies + ALP Physics (2/2)

| Name | Goal | Concept | Made of | Avail. | CERN test |
|-------|-------------------------------|-------------------|---|--------|-------------|
| CB100 | RF studies + First physics | Closed booster | 3 fixed disks $\phi = 100$ mm | 2021 | 2024 |

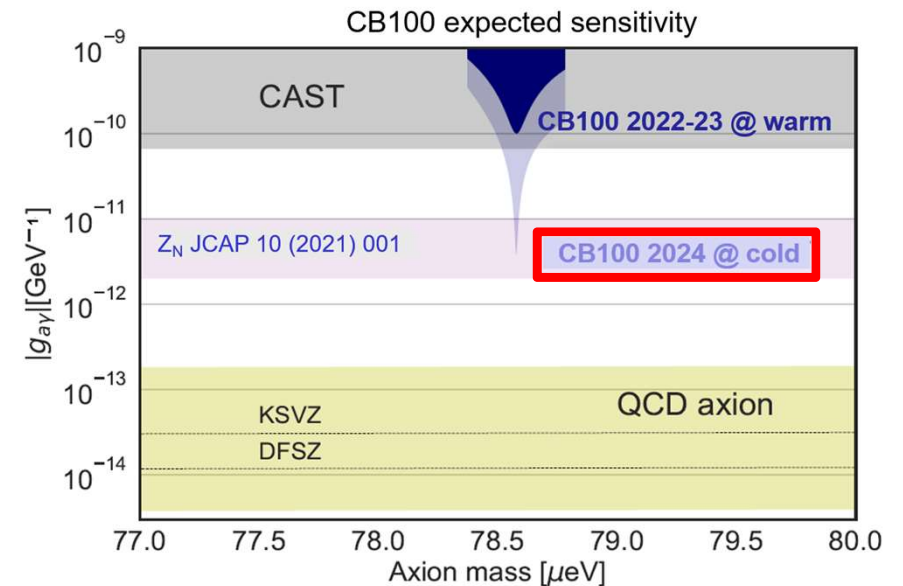
Develop a 'cheap' cryostat with CERN cryolab to cool the booster + LNA \rightarrow Validated the principle in 2023



Vacuum between Inner and Outer vessel



He flow from dewar to cryo
 \rightarrow cooling in $\frac{1}{2}$ day
 \rightarrow Stable <10 K during >10 hours

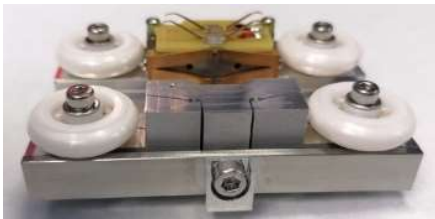


First short cold run in 2024

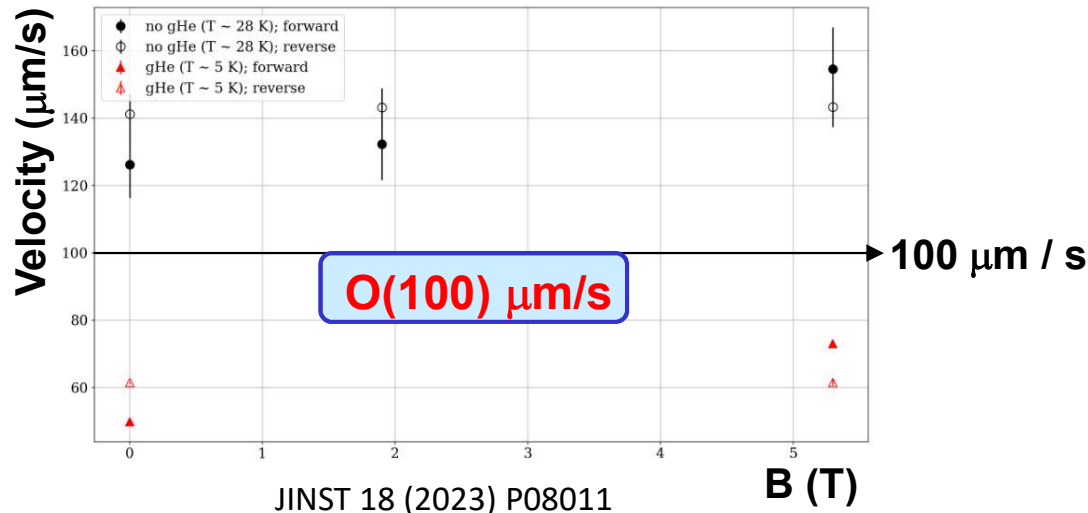
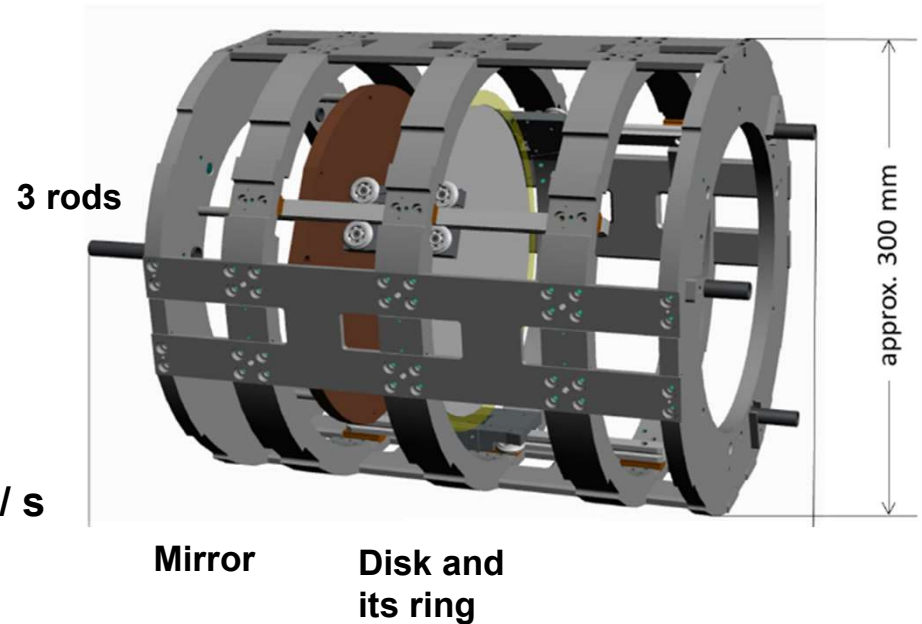
Set-up with moveable disk (1/2)

| Name | Goal | Concept | Made of | Avail. | DESY magnet test |
|------|-------------------------|--------------|---|--------|------------------|
| P200 | Piezo-motor + mechanics | Open Booster | 1 moveable disk $\phi = 200$ mm | 2021 | 2022 |

Successful test of JPE piezo motor at 5 K and 5.3 T (*ALP magnet in DESY*)



Build full mechanical structure of Open Booster and insert 1 mirror + 1 disk (3 piezo motors)

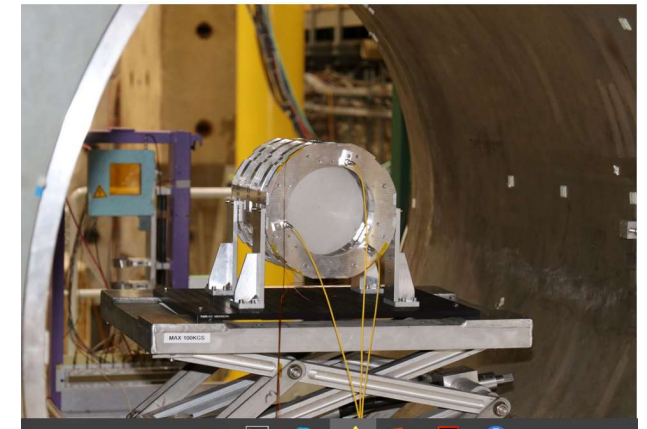
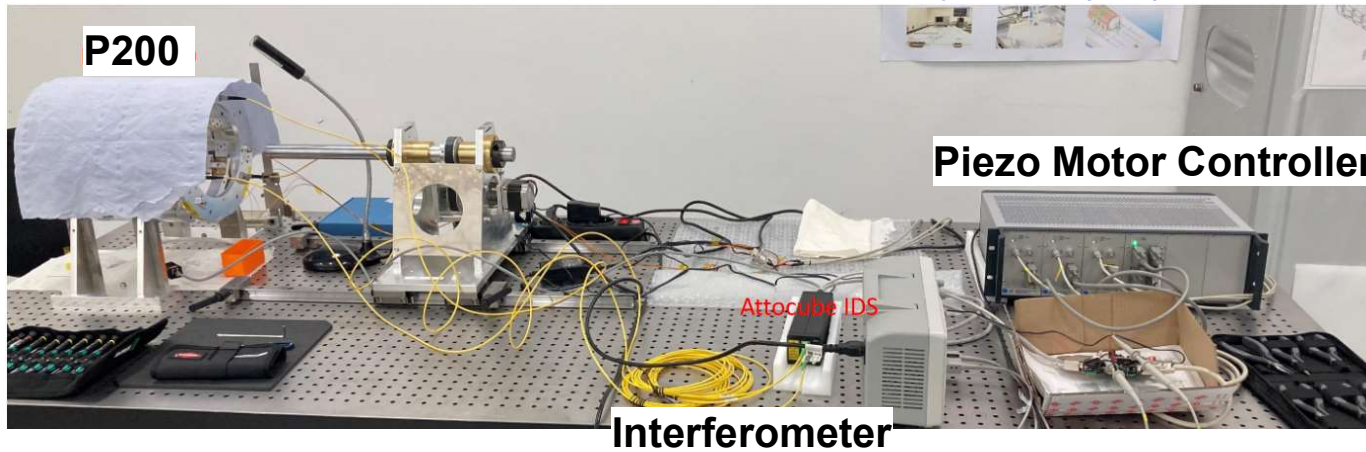


Set-up with moveable disk (2/2)

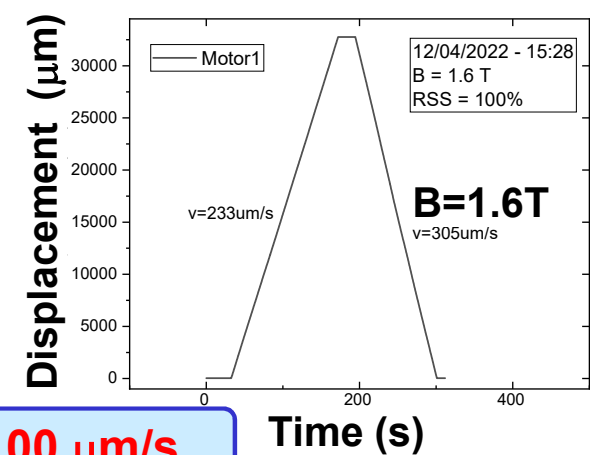
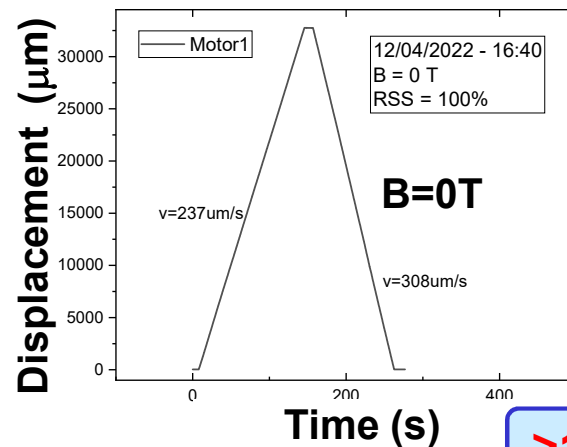
| Name | Goal | Concept | Made of | Avail. | Morpurgo test |
|------|-------------------------|--------------|---|--------|---------------|
| P200 | Piezo-motor + mechanics | Open Booster | 1 moveable disk $\phi = 200$ mm | 2021 | 2022 |

Tested in the lab, in a CERN cryostat (4K) ...

... and under 1.6 T



Validate piezo motors + booster mechanics

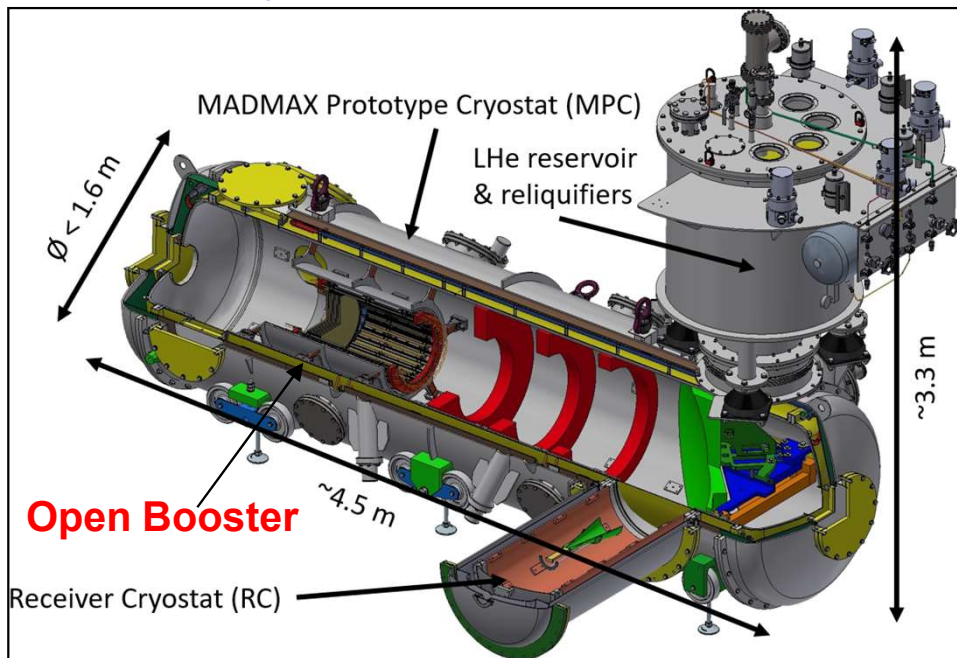


>100 µm/s

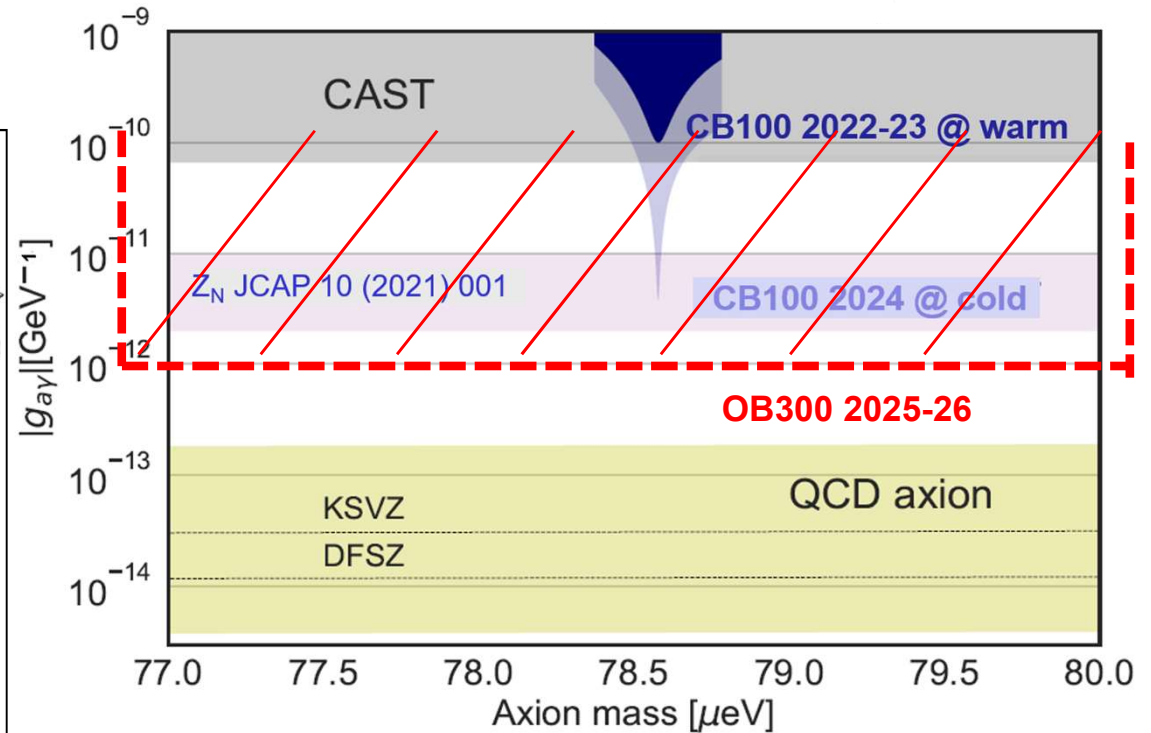
Final prototype + ALP physics

| Name | Goal | Concept | Made of | Avail. | CERN test |
|-------|-----------------------------------|--------------|---|--------|------------------|
| OB300 | Scan ALP around $80 \mu\text{eV}$ | Open booster | 3-20 moveable disks $\phi = 300 \text{ mm}$ | 2024 | 2025, 26? |

Open Booster inserted in a Stainless Steel cryostat (to be delivered in Mar 2024)

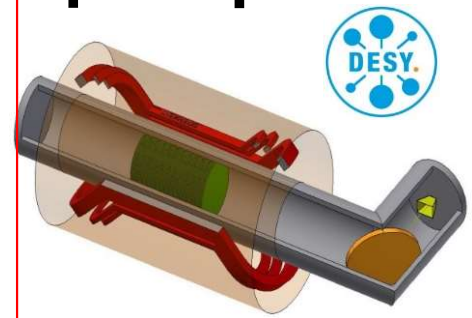
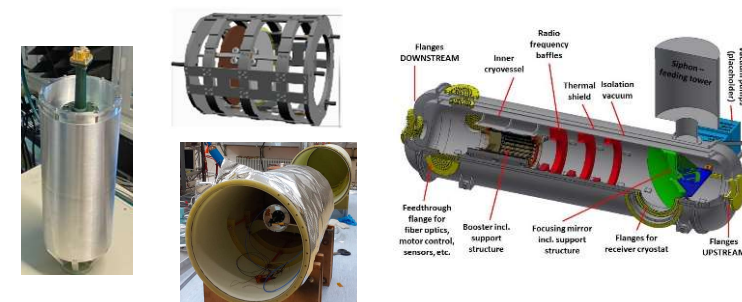
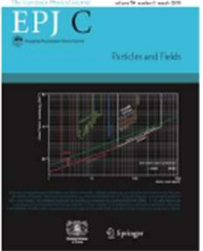
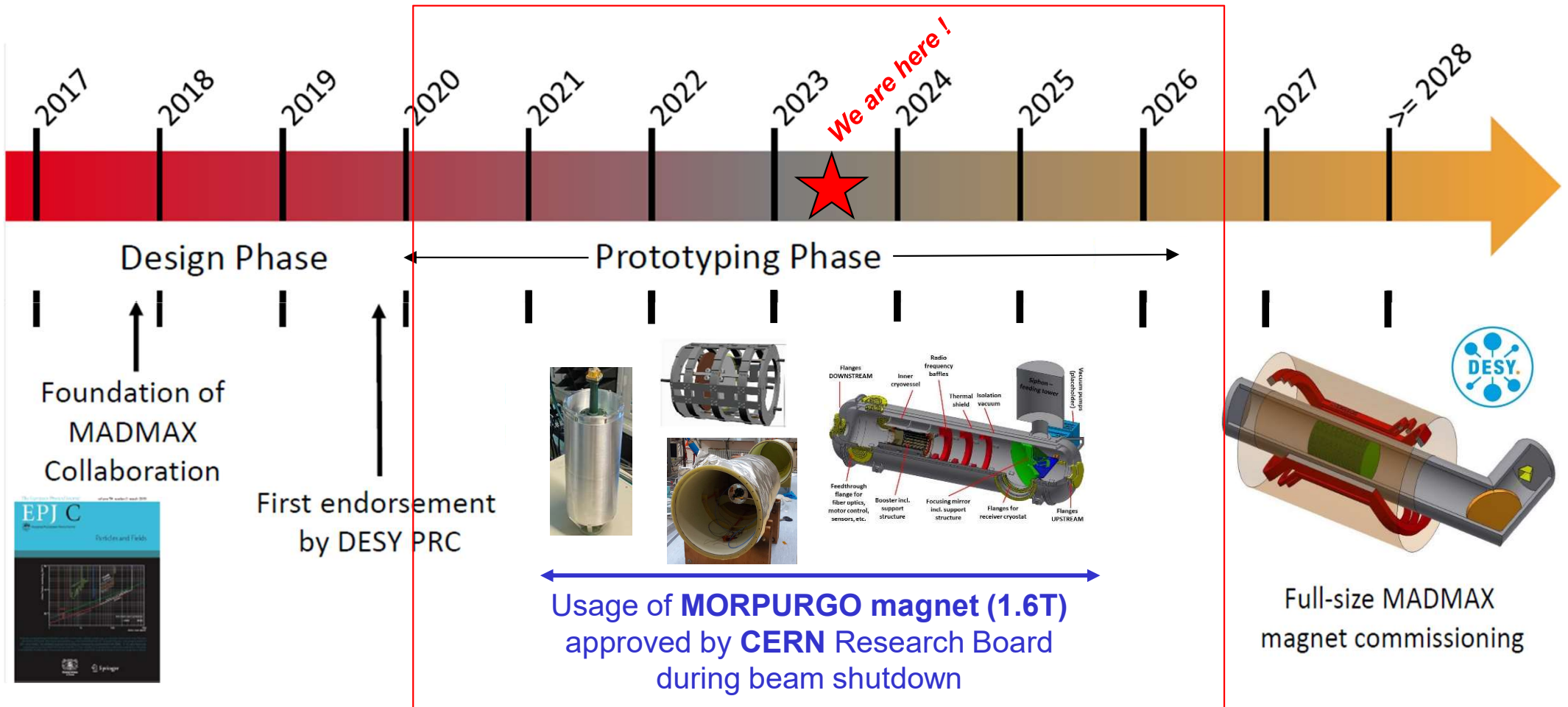


Open Booster built with P200 and CB100 experience



Long cold run + mass scan in 2025 (26?)

MADMAX timescale



$$g_{\alpha\gamma} \propto \left(\frac{1000-10000}{50000 \beta^2}\right)^{1/2} \times \left(\frac{10-220}{4 \text{ K}} T_{sys}\right)^{1/2} \times \left(\frac{1.6}{B_e}\right) \times \left(\frac{0.01-0.1}{A}\right)^{1/2} \times \left(\frac{4 \text{ days}}{t}\right)^{1/4} \times \left(\frac{\text{SNR}}{5}\right)^{1/2}$$



MADMAX & IN2P3

□ Pionnering experimental work at IN2P3 on DM axion search

- **CPPM** joined MADMAX in 2020 (*2 physicists, 1 PhD, 4 IT* → 4 FTE). **IJCLab**: will join in 2024
- **Master Project** at IN2P3 since 2023
- Presentation at IN2P3 Conseil Scientifique on Dark Matter (23-Oct 2023)
- Remark: **CNRS IRL “DMLab” (with Helmholtz centers)** → MADMAX is a central project

Responsibilities in all executive bodies of the MADMAX collaboration



- Member collaboration board
- Member executive board

+ IN2P3 master project coordinator

+ CPPM scientific coordinator

- Member physics board
- CERN tests coordinator
- Member executive board

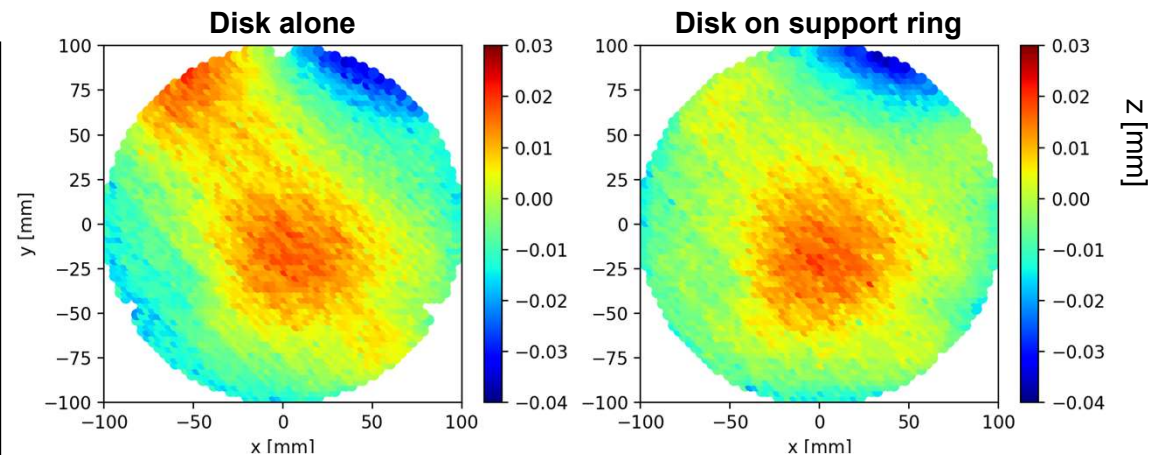
- **Technical coordinator** (2020, 2023-)
- Member executive board

+ CPPM technical coordinator

MADMAX & CPPM (1/2)

□ Precision mechanics for the prototype boosters

- Precision 3D measurements $O(\mu\text{m})$ for geometry control of the disks
 - ✓ CPPM expertise/infrastructure for precision measurements (e.g. ATLAS pixels)
- Conception/fabrication of disk support rings
 - ✓ Interfaces between disks, piezo motors and interferometer system
 - ✓ Cutting edge and challenging R&D → Optimisation of fabrication process to obtain best planarity ($<10\mu\text{m}$)

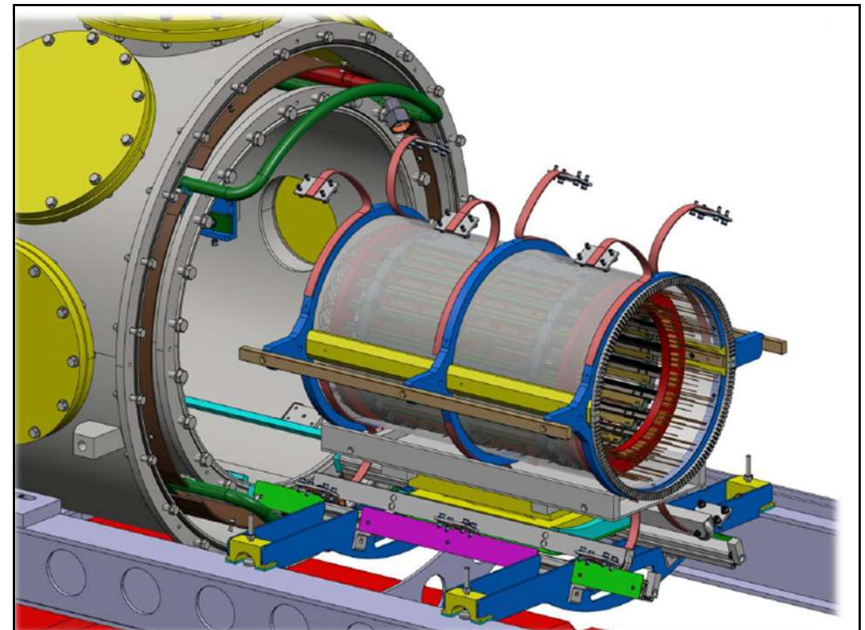


<10 μm RMS disk planarity alone/on the ring

MADMAX & CPPM (2/2)

□ Coordination of prototype tests at CERN in Morpurgo magnet

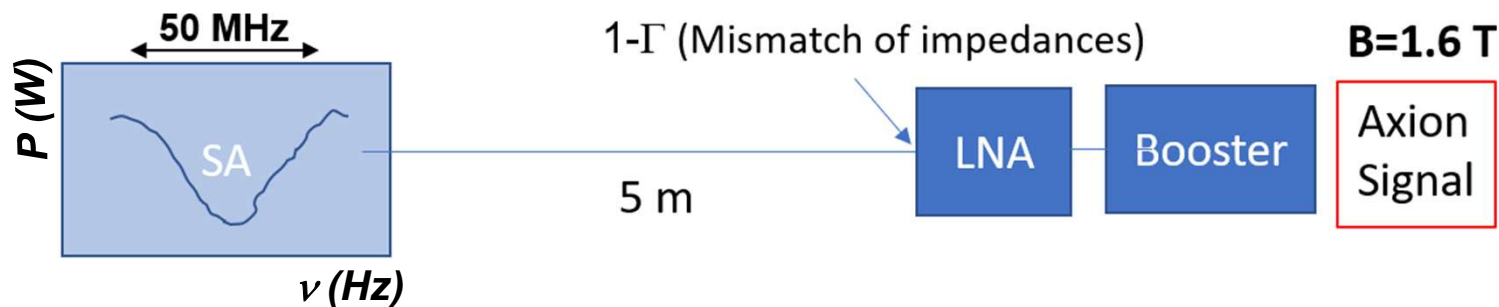
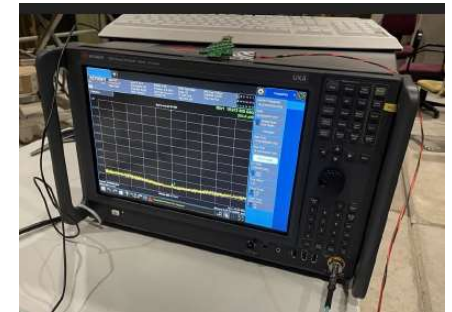
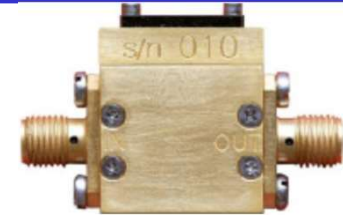
- Impulsion for **magnet choice**, approved by CERN RRB in 2020 for 2021-25 YETS (~1 month/yr)
- Conception, fabrication and installation of **mechanical infrastructures** around the magnet (*Rails for electric racks, supports for prototypes, rails for big test cryostat, ...*)
- Design and construction of **mechanical structure** to align OB300 booster in cryostat and of **integration tools** (at DESY and CERN)
- Simulation and data analysis



Receiver (1/2)

□ Composed of

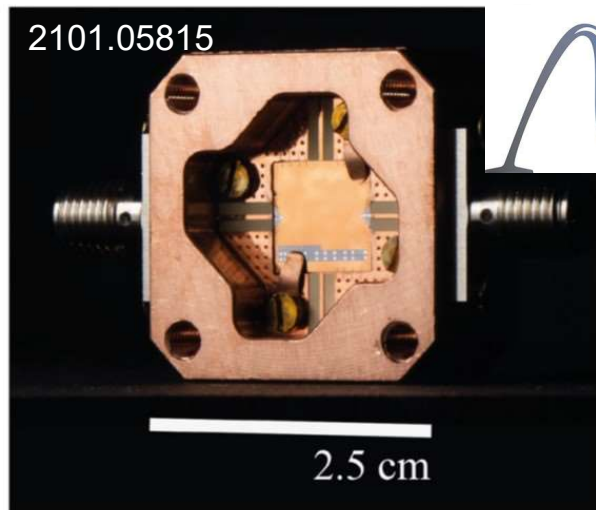
- **Low Noise Amplifier (LNA) ...**
 - ✓ “Classic” HEMT, $G=33$ dB, 4 K added noise
- **... connected to custom-made receiver**
 - ✓ Three mixing stages to down sample from 20 GHz to 50 MHz (*heterodyne mixing*)
 - ✓ Fast Fourier Transform in 4 samplers \rightarrow 1% dead time
 - ✓ Tested at CERN in 2022 but difficult to move + some saturation & time instability
- **... connected to commercial spectrum analyzer (SA)**
 - ✓ Tested at CERN in 2023 : stable, no saturation but higher dead time*
 - ✓ Just bought a new SA with data streaming (~ 0 % dead time)



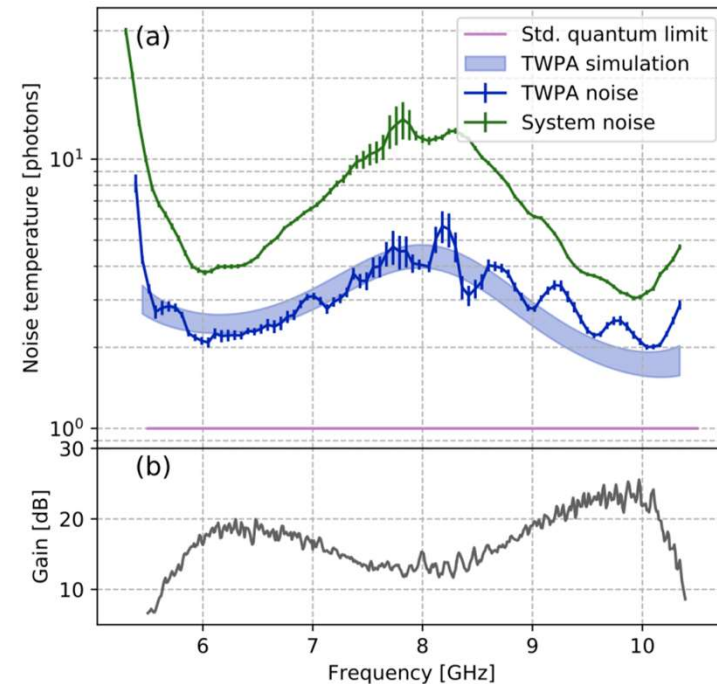
Receiver (2/2)

Progresses on Low Noise amplifier

- Josephson Junction being developed to further minimize noise (*quantum limit*)



TWPA prototype with $G > 20$ dB and 1K added noise at 10 GHz



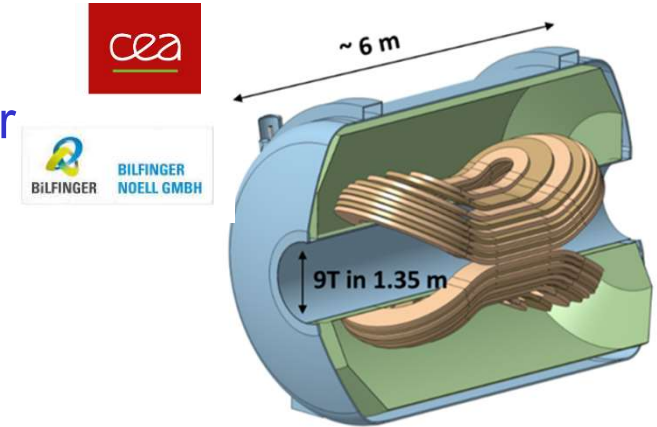
[Reversed Kerr TWPA arXiv:2101.05815]

- Next:** > 40 GHz techno. to be developed

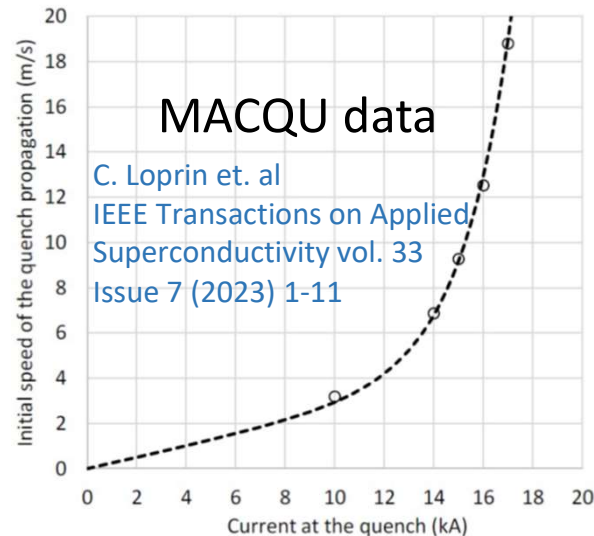
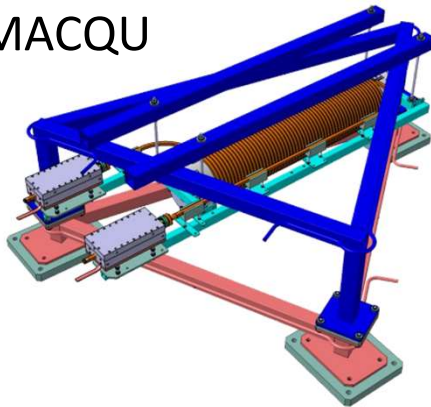
Magnet

Progresses on final magnet

- Design completed: 2x9 skateboard coils with novel copper CICC conductor
[NbTi with Cu jacket @ 1.8K]
- Recently demonstrated that coils will be safe in terms of quench protection (**MA**dmax **C**oil for **Q**uench **U**nderstanding)



MACQU



- Next** : Design, manufacture and test a small coil (6T)

Conclusions

- ❑ **Axion = DM candidate motivated by particle physics since 40 years**
- ❑ **DM axion direct search: rising interest, next decades promising**
 - Resonant cavity sensitivity starts to scratch the QCD axion phase space ($\sim 1 \mu\text{eV}$)
 - Will be extended to most of the interesting mass range (1-1000 μeV) with novel experiments
- ❑ **MADMAX = novel exp. approach to cover theory-favored phase space**
 - Needs for precise (μm) instrumentation in extreme conditions (high B, 4 K, 10's GHz)
 - Prototyping phase at CERN 2021-2026 to validate concept \rightarrow ALP competitive searches
 - CPPM in MadMax since 2020 \rightarrow construction, simulation, test and data analysis of protos

Technological developments on mechanics (precision, piezo motors), magnet, low noise amplifier, cooling

BACKUP

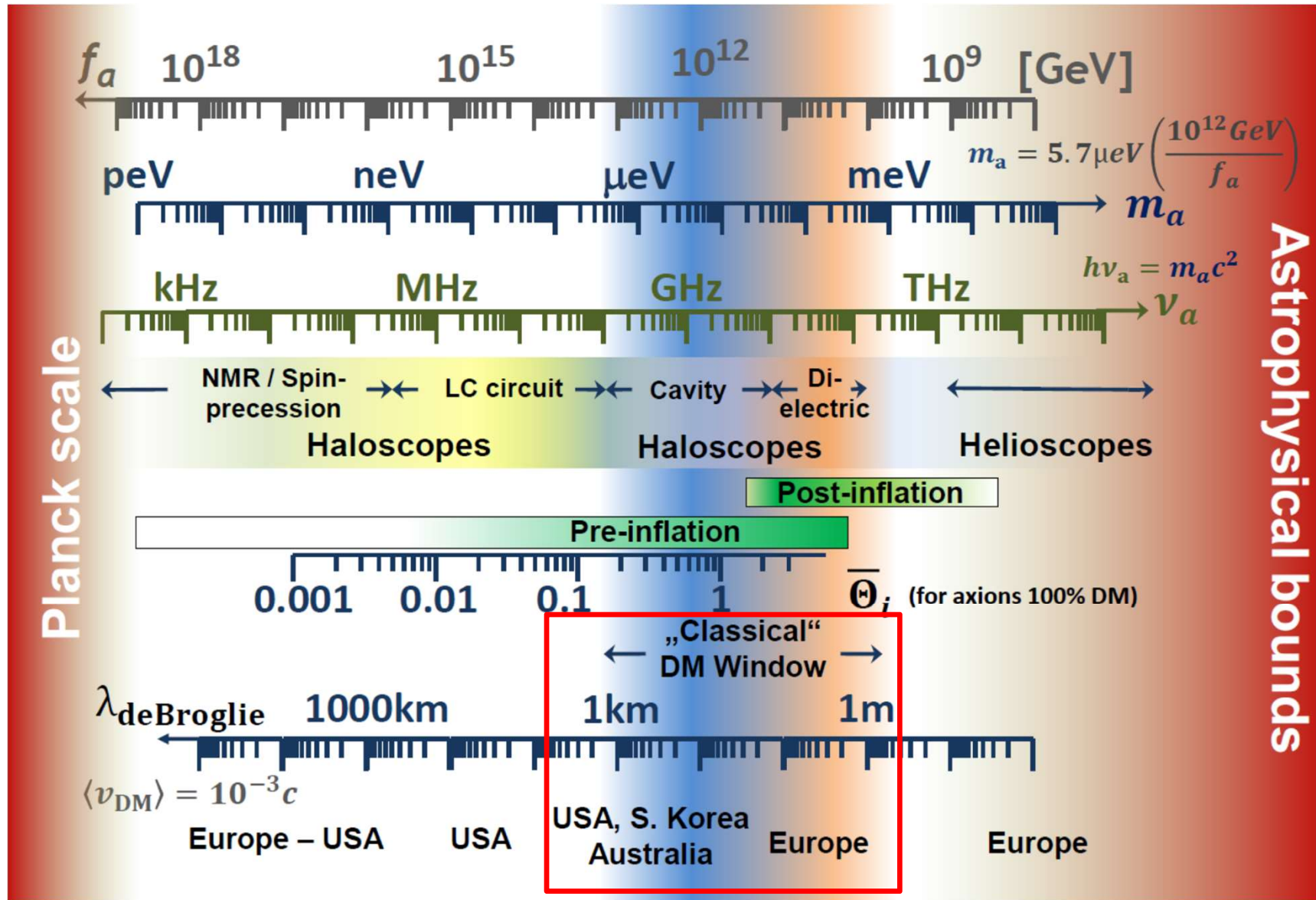
Collab Week at CPPM (Apr 2023)



Axion scales

APPEC Committee Report

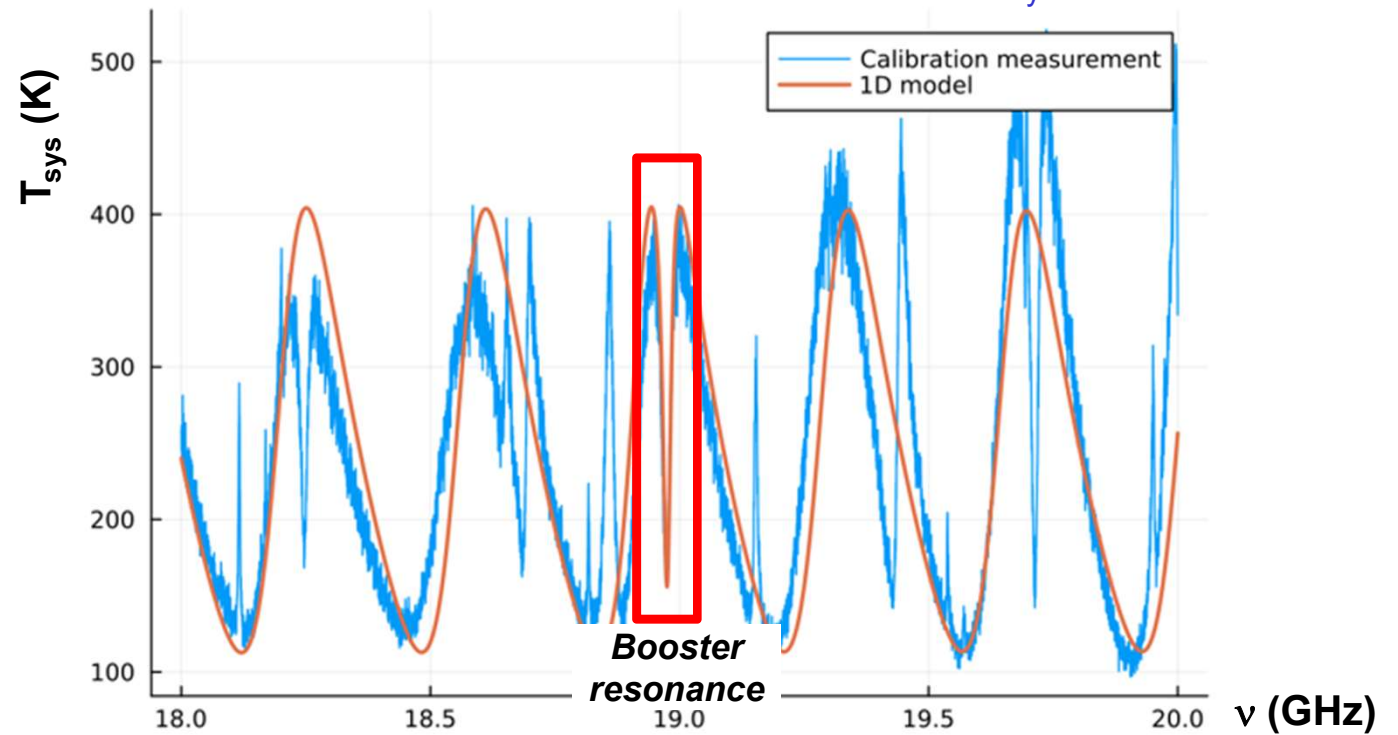
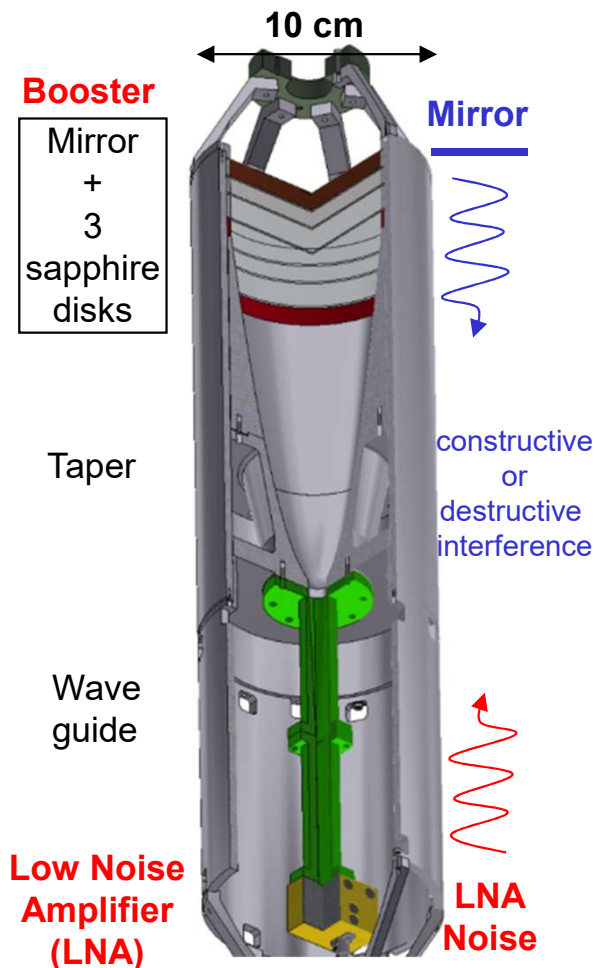
Rept. Prog. Phys., 85(5):056201, 2022, 2104.07634



RF (1/3)

| Name | Goal | Concept | Made of | Avail. |
|-------|------------|----------------|---|--------|
| CB100 | RF studies | Closed booster | 3 fixed disks $\phi = 100\text{mm}$ | 2021 |

Simulate **LNA** (ADS) and **Booster+taper** system (COMSOL)
 → compare with measured system temperature (T_{sys}) in 18-20 GHz



Model the RF behavior (esp. at the booster resonance)

RF (2/3)

| Name | Goal | Concept | Made of | Avail. | Morpurgo test |
|-------|------------|----------------|---|--------|---------------|
| CB100 | RF studies | Closed booster | 3 fixed disks $\phi = 100\text{mm}$ | 2021 | 2022 |

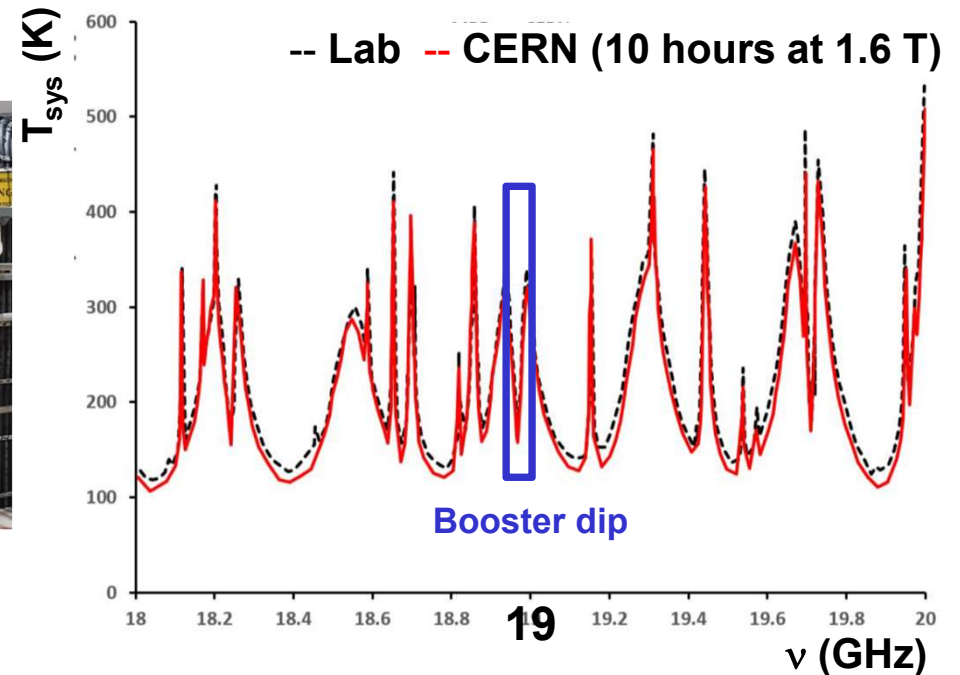
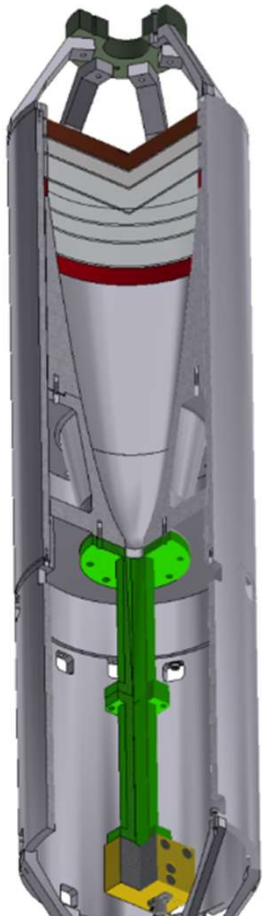
Booster

Mirror
+
3
sapphire
disks

Taper

Wave
guide

LNA

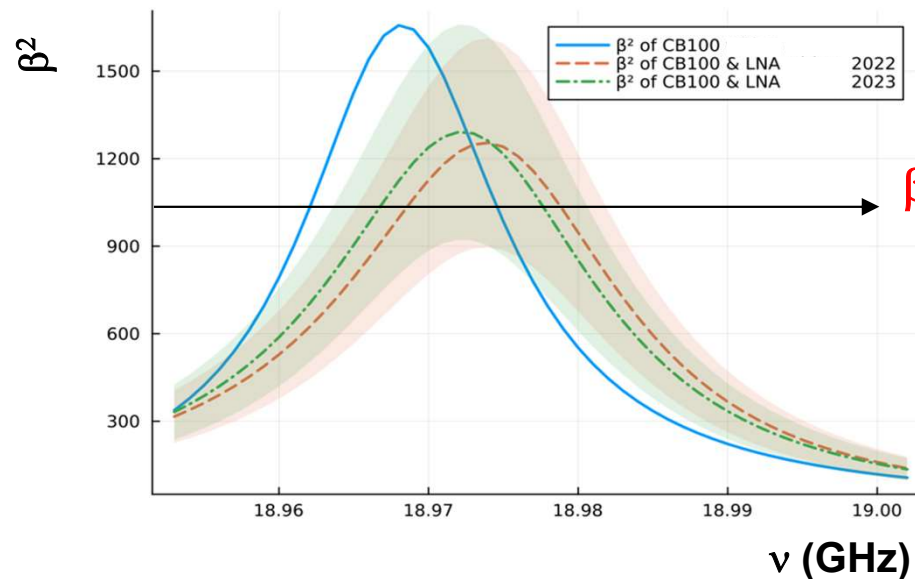
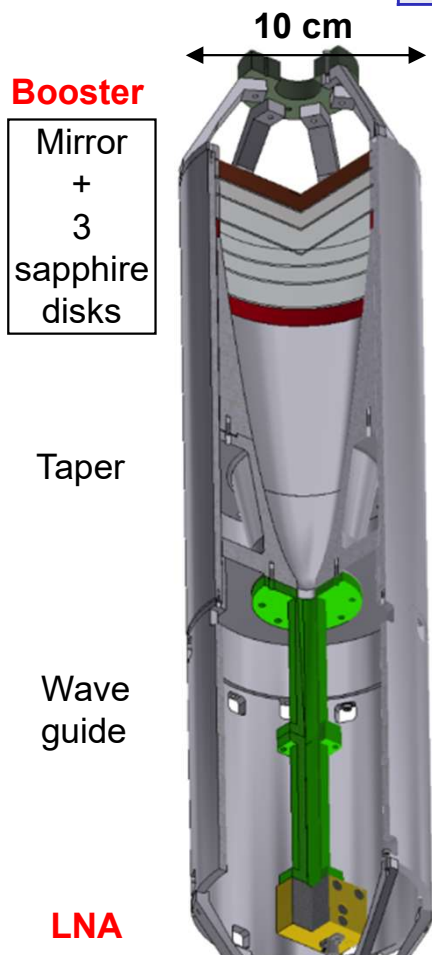


No impact from the CERN environment

RF (3/3)

| Name | Goal | Concept | Made of | Avail. |
|-------|------------|----------------|---|--------|
| CB100 | RF studies | Closed booster | 3 fixed disks $\phi = 100\text{mm}$ | 2021 |

Simulate **LNA** (ADS) and **Booster+taper** system (COMSOL)
 → Extract the boost factor shape

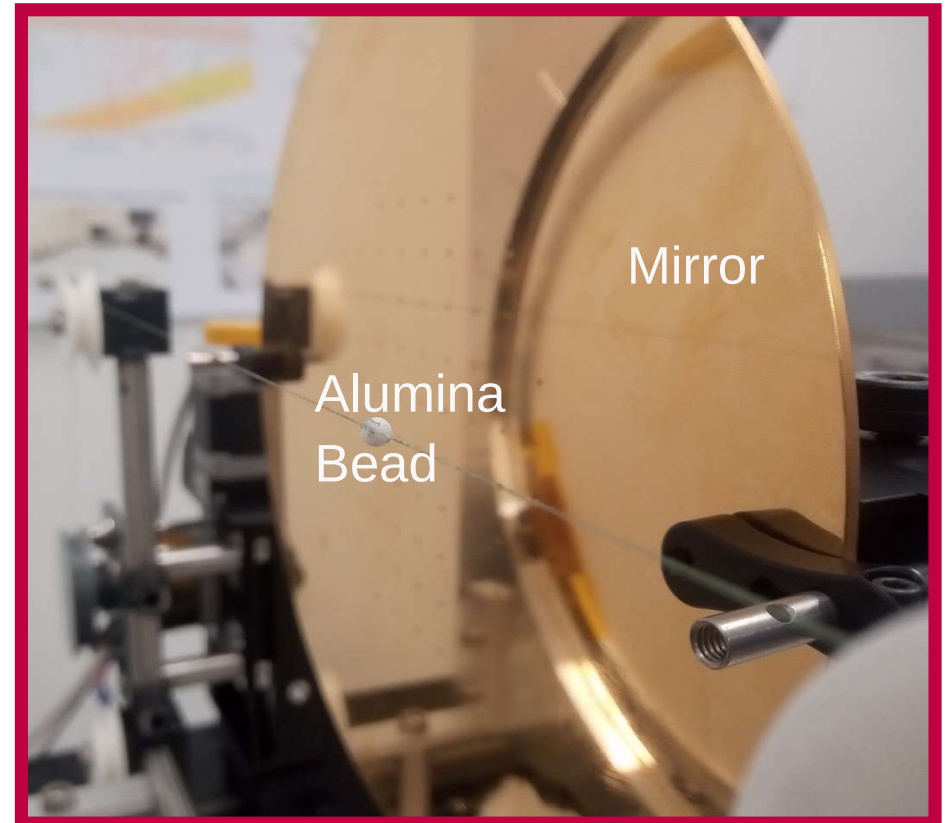
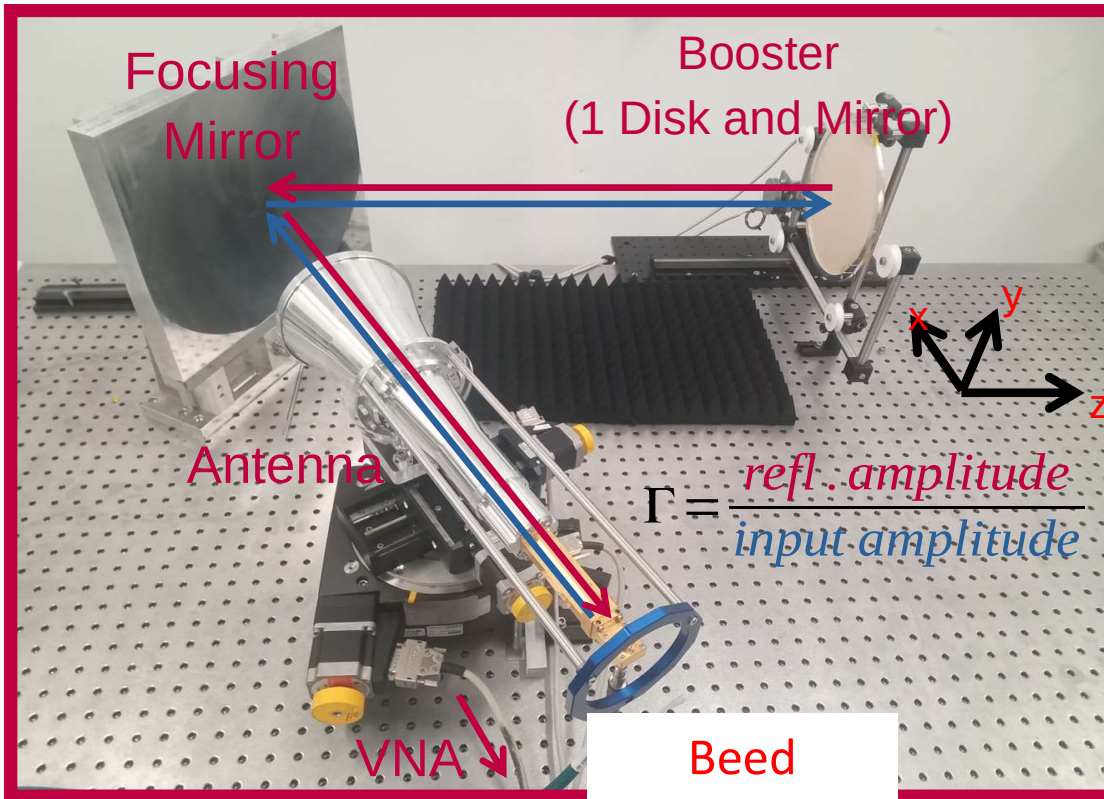


Model the boost factor shape with systematic uncertainties

[paper in preparation]

OB calibration (1/2)

Boost factor determined using Bead Pull Method (non-resonant perturbation theory)
 + Lorentz reciprocity theorem J. Egge, JCAP 04 (2023) 064



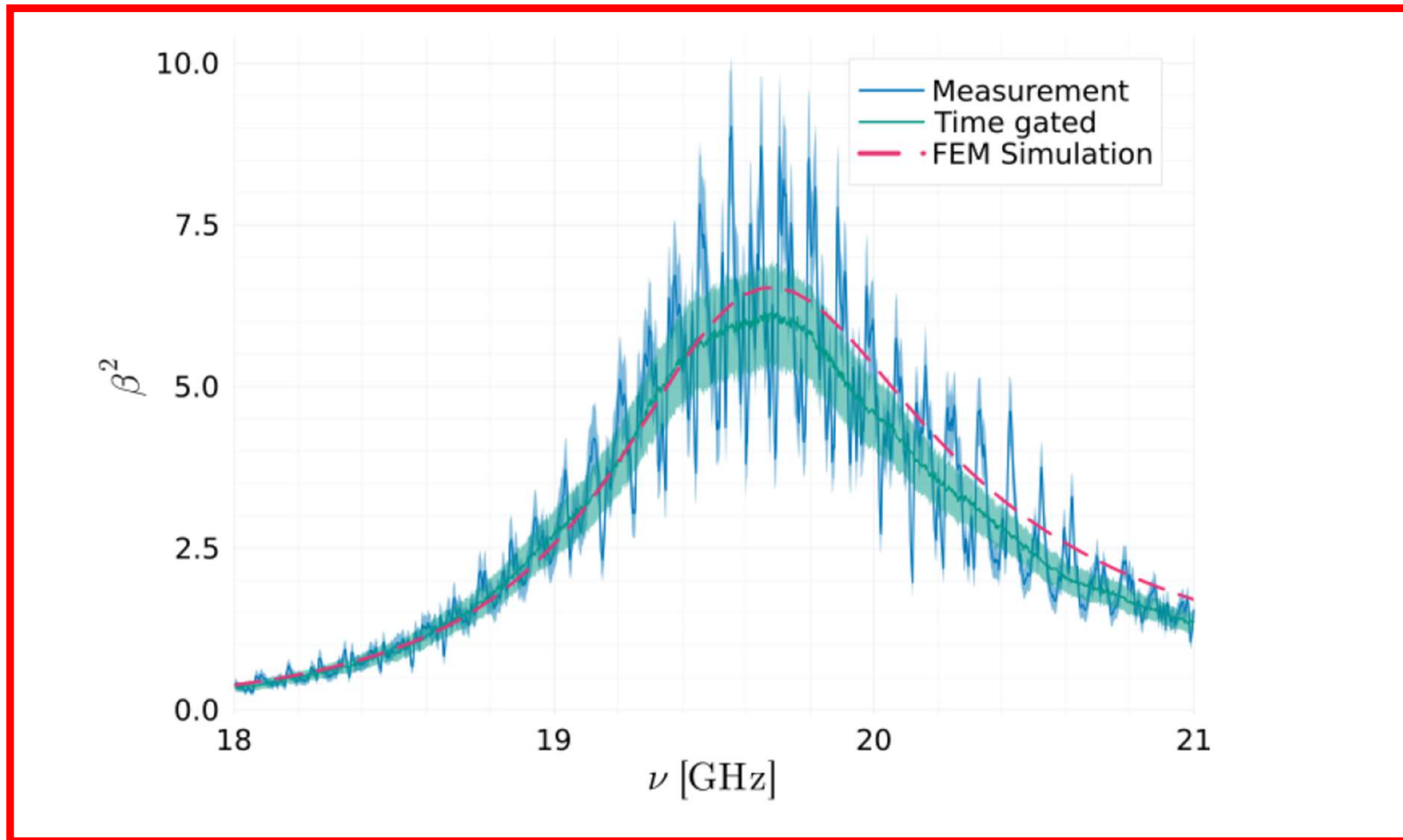
Change in reflection coefficient

$$\Delta\Gamma = \frac{\alpha_e \omega}{4P_{\text{in}}} E_R^2 \rightarrow \text{E field}$$

$$P_{\text{sig}} = \frac{g_{a\gamma}^2}{16P_{\text{in}}} \left| \int_{V_a} dV \mathbf{E}_R \cdot \dot{\mathbf{a}} \mathbf{B}_e \right|^2 \rightarrow \beta^2 = \frac{P_{\text{sig}}}{P_0}$$

OB calibration (2/2)

Test with a single disk + mirror (low boost factor)



Measure boost factor (+ systematics)

[paper in preparation]