

空间科学与物理学院 \_\_\_\_\_\_\_\_ School of Space Science and Physics

Earth: photo credit

# **Theoretical studies on the chirality and wobbling in SDU**

### Bin Qi, Shandong University, China

# CWAN'23

International Conference on Chirality and Wobbling in Atomic Nuclei Huizhou (China), July 10 - 14, 2023



空间科学与物理学院 School of Space Science and Physics

Outline





<mark>Precession(进动)</mark> the angle of Precession is changing --Nutation(章动) 空间科学与物理学院 School of Space Science and Physics

## **Even-Even nuclei**



Rotational angular momentum for a triaxial nucleus is not aligned along the axis with the largest moment of inertia, but precesses and wobbles

A. Bohr and B. R. Mottelson, Nuclear Structure Vol. II. (1975)



空间科学与物理学院 School of Space Science and Physics

## **Even-Even nuclei**

#### In a triaxial deformed even-even nucleus



#### Harmonic approximation (HA)

$$\hat{H}_{\rm rot} = \frac{\hat{I}_1^2}{2\mathcal{J}_1} + \frac{\hat{I}_2^2}{2\mathcal{J}_2} + \frac{\hat{I}_3^2}{2\mathcal{J}_3}$$

Considering the **approximation** 

$$[I_{-}, I_{+}] = 2I_3 \approx 2I \quad (I_{\pm} = I_2 \pm iI_1)$$

$$E(I, \mathbf{n}) = \frac{I(I+1)}{2\mathcal{J}_3} + (\mathbf{n} + \frac{1}{2})\hbar\Omega_{\text{wob}}$$
$$\hbar\Omega_{\text{wob}} = 2I\sqrt{\left(\frac{\hbar^2}{2\mathcal{J}_1} - \frac{\hbar^2}{2\mathcal{J}_3}\right)\left(\frac{\hbar^2}{2\mathcal{J}_2} - \frac{\hbar^2}{2\mathcal{J}_3}\right)} \propto I$$

 $I\uparrow, \ \hbar\Omega_{
m wob}\uparrow$ 

A. Bohr and B. R. Mottelson, Nuclear Structure Vol. II. (1975)



空间科学与物理学院\_\_School of Space Science and Physics

## **Even-Even nuclei**

$$\mathcal{J}_k = \mathcal{J}_0 \sin^2(\gamma - \frac{2}{3}\pi k).$$



Hydrodynamical MoI of the three principal axes as functions of triaxial parameter  $\gamma$ . The unit is taken as  $J_0$ .

$$\hbar\omega = I\left[\left(\frac{1}{\mathcal{J}_2} - \frac{1}{\mathcal{J}_1}\right)\left(\frac{1}{\mathcal{J}_3} - \frac{1}{\mathcal{J}_1}\right)\right]^{1/2}$$



calculated by HA equation (Line) and triaxial rotor model (Dot)

$$E_{\text{wob}} = E(n, I) - \frac{1}{2}[E(0, I-1) + E(0, I+1)]$$

n: phonon number , here n=1



空间科学与物理学院

School of Space Science and Physics



wobbling energies, intraband and interband B(E2) values calculated by HA (Line) and rotor model (Dot)



**Even-Even nuclei** 

Energy level scheme calculated by the rotor model for ground band and n = 1, 2 wobbling bands

BQ\*, Zhang, Wang, Q.B.Chen\*, JPG 48 (2021) 055102



空间科学与物理学院 \_School of Space Science and Physics

## **Even-Even nuclei**

θ

φ

m

To further illustrate the angular momentum geometry with different  $\gamma$ , the probability distribution of angular momentum on the  $(\theta, \phi)$  plane, i.e., azimuthal plot, is calculated.

$$\mathcal{P}^{(\nu)}(\theta,\varphi) = \langle I, \theta\varphi \mid II\nu \rangle^2 = \frac{2I+1}{8\pi} \sum_{KK'} D_{KI}^{I*}(\theta,\varphi,0) \rho_{KK'}^{(\nu)} D_{K'I}^{I}(\theta,\varphi,0)$$

S. Frauendorf, Report on "International conference on Chiral bands in nuclei, KTH, Stockholm", 2016.04 F. Q. Chen, Q. B. Chen, et al, PRC 96, 051303(R) (2017). Q. B. Chen and J. Meng, PRC 98, 031303(R) (2018).





J. 4 7. 3 (100 300)

Precession and tunnelingare two aspects of the quantum wobbling motion.

Zhang, BQ\*, Wang, Jia, Wang, PRC 105, 034339 (2022)

空间科学与物理学院



空间科学与物理学院 School of Space Science and Physics

Outline





空间科学与物理学院 School of Space Science and Physics

## **Even-Even nuclei**

 $Y_{20} + Y_{22} + Y_{30}$ 



- quadrupole-octupole deformed nucleus
- ➢ P: space reflection,
- >  $R_1$ : rotation through angle  $\pi$  about 1-axis,
- >  $S_1$ : reflection with respect to the 2-3 plane.

multip	lication	table	of	$C_{2v}$
			-	2 V

	E	$S_1$	$S_2$	$\mathcal{R}_3$
E	E	$\mathcal{S}_1$	$\mathcal{S}_2$	$\mathcal{R}_3$
$S_1$	$S_1$	E	$\mathcal{R}_3$	$S_2$
$S_2$	$S_2$	$\mathcal{R}_3$	E	$S_1$
$\mathcal{R}_3$	$\mathcal{R}_3$	$S_2$	$S_1$	E

**Symmetry of nuclear density distribution:** C<sub>2v</sub> point group



空间科学与物理学院 \_School of Space Science and Physics

## **Even-Even nuclei**

## Experimental data in <sup>148</sup>Ce





空间科学与物理学院 \_ School of Space Science and Physics

## **Even-Even nuclei**

	E	$\mathcal{R}_1$	$\mathcal{R}_2$	$\mathcal{R}_3$	$\mathcal{S}_1$	$\mathcal{S}_2$	$S_3$	$\mathcal{P}$
E	E	$\mathcal{R}_1$	$\mathcal{R}_2$	$\mathcal{R}_3$	$S_1$	$S_2$	$S_3$	$\mathcal{P}$
$\mathcal{R}_1$	$\mathcal{R}_1$	E	$\mathcal{R}_3$	$\mathcal{R}_2$	$\mathcal{P}$	$S_3$	$S_2$	$S_1$
$\mathcal{R}_2$	$\mathcal{R}_2$	$\mathcal{R}_3$	E	$\mathcal{R}_1$	$S_3$	$\mathcal{P}$	$S_1$	$S_2$
$\mathcal{R}_3$	$\mathcal{R}_3$	$\mathcal{R}_2$	$\mathcal{R}_1$	E	$S_2$	$S_1$	$\mathcal{P}$	$S_3$
$\mathcal{S}_1$	$S_1$	$\mathcal{P}$	$S_3$	$\mathcal{S}_2$	E	$\mathcal{R}_3$	$\mathcal{R}_2$	$\mathcal{R}_1$
$S_2$	$S_2$	$S_3$	$\mathcal{P}$	$\mathcal{S}_1$	$\mathcal{R}_3$	E	$\mathcal{R}_1$	$\mathcal{R}_2$
$\mathcal{S}_3$	$S_3$	$S_2$	$S_1$	$\mathcal{P}$	$\mathcal{R}_2$	$\mathcal{R}_1$	E	$\mathcal{R}_3$
$\mathcal{P}$	$\mathcal{P}$	$\mathcal{S}_1$	$S_2$	$\mathcal{S}_3$	$\mathcal{R}_1$	$\mathcal{R}_2$	$\mathcal{R}_3$	E
	E	$\mathcal{R}_1$	$\mathcal{R}_2$	$\mathcal{R}_3$	$S_1$	$S_2$	$S_3$	$\mathcal{P}$
Ag	<u>E</u> 1	$\frac{\mathcal{R}_1}{1}$	$\frac{\mathcal{R}_2}{1}$	$\frac{\mathcal{R}_3}{1}$	$\frac{S_1}{1}$	$\frac{S_2}{1}$	$\frac{S_3}{1}$	$\frac{\mathcal{P}}{1}$
Ag B <sub>1g</sub>	<u>Е</u> 1 1	$\frac{\mathcal{R}_1}{1}$	$\frac{\mathcal{R}_2}{1} \\ -1$	$\frac{\mathcal{R}_3}{1}$	$\frac{S_1}{1}$	$\frac{S_2}{1}$	$\frac{S_3}{1}$	P 1 1
Ag B <sub>1g</sub> B <sub>2g</sub>	<i>E</i> 1 1 1	$\frac{\mathcal{R}_1}{1}\\-1$	$\frac{\mathcal{R}_2}{1} \\ -1 \\ 1$	$\frac{\mathcal{R}_3}{1} \\ -1 \\ -1$	$\frac{S_1}{1}$ -1	$\frac{S_2}{1} - 1 \\ 1$	$\frac{S_3}{1} \\ -1 \\ -1$	<i>P</i> 1 1 1 1
$\begin{array}{c} A_g\\ B_{1g}\\ B_{2g}\\ B_{3g} \end{array}$	E 1 1 1 1 1	$\begin{array}{c} \mathcal{R}_1 \\ 1 \\ -1 \\ -1 \\ -1 \end{array}$		$     \begin{array}{r} \mathcal{R}_3 \\     1 \\     -1 \\     -1 \\     1 \\     1   \end{array} $	$\frac{S_1}{1} \\ -1 \\ -1 \\ -1$			P 1 1 1 1
$\begin{array}{c} A_g\\ B_{1g}\\ B_{2g}\\ B_{3g}\\ A_u \end{array}$	E 1 1 1 1 1	$\begin{array}{c} \mathcal{R}_{1} \\ 1 \\ -1 \\ -1 \\ 1 \\ 1 \end{array}$			$\frac{S_1}{1} \\ -1 \\ -1 \\ -1 \\ -1 \\ -1 \\ -1 \\ -1 \\ $			P 1 1 1 1 1 -1
$\begin{array}{c} A_g\\ B_{1g}\\ B_{2g}\\ B_{3g}\\ A_u\\ B_{1u} \end{array}$	E 1 1 1 1 1 1	$\begin{array}{c} \mathcal{R}_1 \\ 1 \\ -1 \\ -1 \\ 1 \\ 1 \\ 1 \end{array}$		$\begin{array}{c} \mathcal{R}_{3} \\ 1 \\ -1 \\ -1 \\ 1 \\ 1 \\ -1 \end{array}$	$S_1$ 1 -1 -1 -1 -1 -1 -1			P 1 1 1 1 -1 -1
$\begin{array}{c} A_g\\ B_{1g}\\ B_{2g}\\ B_{3g}\\ A_u\\ B_{1u}\\ B_{2u} \end{array}$	E 1 1 1 1 1 1 1 1	$\begin{array}{c} \mathcal{R}_1 \\ 1 \\ -1 \\ -1 \\ 1 \\ 1 \\ -1 \\ 1 \\ -1 \end{array}$	$\begin{array}{c} \mathcal{R}_2 \\ 1 \\ -1 \\ 1 \\ -1 \\ 1 \\ -1 \\ 1 \\ 1 \end{array}$	$\begin{array}{c} \mathcal{R}_{3} \\ 1 \\ -1 \\ -1 \\ 1 \\ 1 \\ -1 \\ -1 \\ -1 \end{array}$	$S_1$ 1 -1 -1 -1 -1 -1 1	$S_2$ 1 -1 1 -1 -1 1 -1 1 -1 -1		P 1 1 1 1 -1 -1 -1 -1

Wang, BQ\*, Liu, Aman, Zhang, PRC 106, 064325 (2022)

Symmetry group of Hamitonian / wave function:  $\{E,R_1,R_2,R_3,S_1,S_2,S_3,P\}$ 

 $\succ$  D<sub>2</sub>×{E,P}

rotational states with one parity for the quadrupoledeformed nucleus are extended to the states with two parities for the octupole deformed nucleus



空间科学与物理学院 \_\_\_\_\_\_school of Space Science and Physics

## **Even-Even nuclei**

$$|I, K, \mathbb{S}\rangle = \frac{1}{\sqrt{1+\delta_{K0}}} \frac{1}{\sqrt{2}} (|IMK\rangle + |IM - K\rangle),$$
$$|I, K, \mathbb{A}\rangle = \frac{1}{\sqrt{2}} (|IMK\rangle - |IM - K\rangle),$$

$(r_1, r_2, r_3)$	D <sub>2</sub> group REP	$ I, K, \mathbb{S}\rangle$		$ I, K, \mathbb{A}\rangle$	
8		Ι	K	Ι	K
(+1, +1, +1)	A	Even	Even	Odd	Even
(+1, -1, -1)	B <sub>1</sub>	Even	Odd	Odd	Odd
(-1, +1, -1)	$B_2$	Odd	Odd	Even	Odd
(-1, -1, +1)	$B_3$	Odd	Even	Even	Even

Case	Wave function	K	Even I		Odd I	
			p = +1	p = -1	p = +1	p = -1
Ι	$ I, K, \mathbb{S}\rangle$	Odd	$B_{1g}$	B <sub>1u</sub>	B <sub>2g</sub>	B <sub>2u</sub>
II	$ I, K, \mathbb{A}\rangle$	Odd	B <sub>2g</sub>	B <sub>2u</sub>	B <sub>1g</sub>	B <sub>1u</sub>
III	$ I, K, \mathbb{A}\rangle$	Even	B <sub>3g</sub>	B <sub>3u</sub>	Ag	A <sub>u</sub>
IV	$ I, K, \mathbb{S}\rangle$	Even	Ag	A <sub>u</sub>	B <sub>3g</sub>	B <sub>3u</sub>



s in the figure means  $s_{1.}$ For two group representations  $A_g$  and  $B_{3u}$ ,  $s_1=s_2$  The reasonable wave functions for excited band are selected by agreeing with the experimental transition properties.



空间科学与物理学院 School of Space Science and Physics

## **Even-Even nuclei**

- The reasonable wave functions for excited band are selected by agreeing with the experimental transition properties.
- Parity + sequence in both ground and excited band: Ag,  $(r_1,s_1,p)=(+1,+1,+1)$ Parity - sequence in both ground and excited band:  $B_{3u}$ ,  $(r_1,s_1,p)=(-1,+1,-1)$
- The selected wave functions are also consistent with the conclusion obtained from the perspective of symmetry restoration

In such an assignment, the states occur with two values  $\pm 1$  of the quantum number  $r_1$ ,  $r_2$ ,  $s_3$ , and p, which corresponds to the violation of  $R_1$ ,  $R_2$ ,  $S_3$ , and P in the intrinsic frame.

While the states occur with single value of  $r_3$ ,  $s_1$ , and  $s_2$  in the laboratory frame which corresponds to  $R_3$ ,  $S_1$ , and  $S_2$  invariance in the intrinsic frame.



空间科学与物理学院 \_school of Space Science and Physics

## **Even-Even nuclei**





The excited band originates from the wobbling excitation of ground band.

 $\frac{\text{Wobbling}}{\text{in } Y_{20} + Y_{22} + Y_{30} \text{ nucleus}}$ 

Wang, BQ\*, Liu, Aman, Zhang, PRC 106, 064325 (2022)



空间科学与物理学院 \_ School of Space Science and Physics

Outline





空间科学与物理学院 \_\_School of Space Science and Physics

## odd-A nuclei





Ødegård *et al*. PRL 86, 5866 (2001) Jensen *et al*. PRL 89.142503 (2002)

8



空间科学与物理学院 \_\_\_\_\_\_school of Space Science and Physics

## odd-A nuclei

Nuclei	Z	Ν	configur ation	β	γ	mass rigen	Reference
<sup>105</sup> Pd	46	59	$vh_{11/2}$	0.27	25	100	J. Timár, et al., PRL. 122, 062501 (2019).
<sup>127</sup> Xe	54	73	vh <sub>11/2</sub>				S. Chakraborty, et al. PLB 811, 135854 (2020).
<sup>130</sup> Ba	<mark>56</mark>	<mark>74</mark>	$\pi(h_{11/2})^2$	<mark>0.24</mark>	<mark>21.5</mark>		Q. B. Chen, et al. PRC 100, 061301 (2019).
<sup>133</sup> Ba	56	77	$vh_{11/2}$			130	D. K. Rojeeta, et al. PLB 823, 136756 (2021).
<sup>133</sup> La	57	76	$\pi h_{11/2}$	0.17	26		S. Biswas, et al., EPJA 55: 159(2019).
<sup>135</sup> Pr	59	76	$\pi h_{11/2}$	0.17	26		J. T. Matta, et al., PRL 114, 082501 (2015).
<sup>161</sup> Lu	71	90	$\pi i_{13/2}$	0.42	20		P. Bringel, et al., EPJA 24, 167 (2005).
<sup>163</sup> Lu	71	92	$\pi i_{13/2}$	0.42	20		S. W. Ødegård, et al., PRL. 86, 5866 (2001).
<sup>165</sup> Lu	71	94	$\pi i_{13/2}$	0.42	20	160	G. Schönwaßer, et al., PLB 552, 9 (2003).
<sup>167</sup> Lu	71	96	$\pi i_{13/2}$	0.43	19		H. Amro, et al., PLB553, 197 (2003).
<sup>167</sup> Ta	73	94	$\pi i_{13/2}$	0.41	20		D. J. Hartley, et al., PRC 80, 041304(R) (2009).
192 •	70	104	$\pi i_{13/2}$	0.29	21.4		
<sup>105</sup> Au	/9	104	$\pi h_{9/2}$	0.3	20	190	S. Nandi, <i>et al.</i> , PKL 125, 132501 (2020).
<sup>187</sup> Au	79	108	$\pi h_{9/2}$	0.23	23		N. Sensharma, et al., PRL 124, 052501 (2020).



空间科学与物理学院 School of Space Science and Physics

## odd-A nuclei



Fingerprint of wobbling band:

- > Sequences of  $\Delta I=2$  rotational bands
- Exhibit similar moments of inertia, spin alignments and in-band B(E2) values for ground (n=0) and excited (n=1,2) band
- ➤ the interband  $\Delta I = 1, n \rightarrow n 1$  transitions are dominated by the E2 component
- the wobbling energy decreases with spin I, contrary to the behavior expected for even-even nuclei

J. Timár, et al., Phys. Rev. Lett. 122, 062501 (2019)



空间科学与物理学院 \_\_School of Space Science and Physics

## odd-A nuclei



#### Transverse wobbler (TW)



				1/2
to j	$\begin{bmatrix} 1 & J \\ J \end{bmatrix}$	1)][1 .	$J_{J_3}$	1)][
$n_{\rm Mob} = \overline{T_2}$	$1 + \frac{1}{i} (\frac{1}{7_1})$	- 1)  1 + -	$\overline{i}$ $\overline{7_0}$	1) [
03	- J .01		J 102	·-)

1/2
211

	the orientation of $j$ (// $I$ )	wobbling energy	quasiparticle orbital of j shell
LW	aligned parallel to the m-axis	increases with spin	middle (m-axis)
TW	aligned perpendicular to the m-axis	decreases with spin	bottom (s-axis) top (l-axis)

1

Frauendorf & Dönau, PRC 89, 014322 (2014)



空间科学与物理学院\_\_school of Space Science and Physics

## odd-A nuclei

Stability of the wobbling motion in an odd-mass nucleus and the analysis of <sup>135</sup>Pr Tanabe K and Sugawara-Tanabe K 2017 Phys. Rev. C 95 064315

Comment on "Stability of the wobbling motion in an odd-mass nucleus and the analysis of <sup>135</sup>Pr" Frauendorf S 2018 Phys. Rev. C 97 069801

Reply to "Comment on 'Stability of the wobbling motion in an odd-mass nucleus and the analysis of <sup>135</sup>Pr'" Tanabe K 2018 Phys. Rev. C 97 069802

#### Tilted precession and wobbling in triaxial nuclei,

Lawrie, Shirinda and Petrache 2020 Phys. Rev. C 101 034306

Eur. Phys. J. A (2022) 58:75 https://doi.org/10.1140/epja/s10050-022-00727-5

Regular Article - Theoretical Physics

THE EUROPEAN PHYSICAL JOURNAL A Check for updates

#### Study of wobbling modes by means of spin coherent state maps

#### Q. B. Chen<sup>1,2</sup>, S. Frauendorf<sup>3,a</sup>

<sup>1</sup> Department of Physics, East China Normal University, Shanghai 200241, China

<sup>2</sup> Physik-Department, Technische Universität München, 85747 Garching, Germany

<sup>3</sup> Physics Department, University of Notre Dame, Notre Dame, IN 46556, USA



空间科学与物理学院 \_\_\_\_\_\_school of Space Science and Physics

Employing different parameter sets of moment of inertia

## odd-A nuclei

(MOI), several calculated results for <sup>105</sup>Pd could be in good

agreement with the experimental data

configuration	$v(1h_{11/2})^1$
deformation	β~0.27; γ~25°
a footor	$g_{R}=0.43$
g-factor	$g_n = -0.21$
quadrupole moments	Q=3.0eb

$$\mathcal{J}_k = a_k \sqrt{1 + bI(I+1)}$$

J. Timár, et al., Phys. Rev. Lett. 122, 062501 (2019)

参数组	$a_m$	$a_s$	$a_l$	b	$\mathcal{J}_m:\mathcal{J}_s:\mathcal{J}_l$
(A)	6.0	5.4	1.8	0.016	1:0.9:0.3
(B)	6.0	4.2	1.2	0.023	1:0.7:0.2
(C)	6.0	3.0	1.0	0.026	1:0.5:0.17
(D)	12.0	3.6	1.0	0.008	1:0.3:0.08





Take set (A) of MOI results as an example:



azimuthal plot see Refs. Chen, Chen, Luo, Meng and Zhang PRC 96 051303(2017) Chen and Meng, PRC 98, 031303(2018)

#### School of Space Science and Physics Odd-A nuclei

#### close to the rigid-body model

空间科学与物理学院

参数组	$a_m$	$a_s$	$a_l$	b	$\mathcal{J}_m:\mathcal{J}_s:\mathcal{J}_l$
(A)	6.0	5.4	1.8	0.016	1:0.9:0.3

The probability distribution of angular momentum on the  $(\theta, \phi)$  plane, i.e., azimuthal plot, is shown and corresponding schematic diagram is provided.





空间科学与物理学院 \_\_school of Space Science and Physics

## odd-A nuclei

参数组	$a_m$	$a_s$	$a_l$	b	$\mathcal{J}_m:\mathcal{J}_s:\mathcal{J}_l$	
(A)	6.0	5.4	1.8	0.016	1:0.9:0.3	close to the rigid-body model
(B)	6.0	4.2	1.2	0.023	1:0.7:0.2	
(C)	6.0	3.0	1.0	0.026	1:0.5:0.17	
(D)	12.0	3.6	1.0	0.008	1:0.3:0.08	close to the hydrodynamical model

corresponding angular momentum geometry

show distinct modes of rotational excitation



With the increasing of the ratio between the MOI at the *m* and *s* axis, namely Jm/Js, the rotational modes gradually changes from Mode I to Mode II and then to Mode III.



- distinct modes of rotational excitation are shown when Employing different parameter sets of moment of inertia (MOI),
- TW mode is sensitive to the ratio Jm/Js
- ideal precession does not appear, the tunneling between two orientations of angular momentum may be preferable

Zhang, *BQ*\*, Wang, Liu, Wang, PRC105, 034339 (2022)



空间科学与物理学院 \_ School of Space Science and Physics

Outline







空间科学与物理学院

#### chiral partner

yrast band

#### wobbling band

First Observation for Chiral-Wobbler in Nuclei, R J Guo, et al. under Review



间科学与物理学院 chool of Space Science and Physics

## odd-odd nuclei

Parameters of particle rotor model :

 $\pi$  g\_{9/2} (particle-like) v g\_{9/2} (midle subshell) pairing gap 1.40 MeV.

Deformation parameters  $(\beta, \gamma) = (0.45, 27.5^{\circ})$  from RMF

moments of inertia ~hydrodynamical

The calculated results reproduced the corresponding experimental data well



空间科学与物理学院 School of Space Science and Physics

Outline



Thank you for your attention!



# Appendix







空间科学与物理学院 \_School of Space Science and Physics

#### 理论框架

多粒子转子模型哈密顿量:

分:  

$$H = \hat{H}_{coll} + \hat{H}_{intr}$$

$$H_{coll} = \sum_{1}^{3} \frac{\hat{R}_{i}^{2}}{2\mathcal{J}_{i}} = \sum_{1}^{3} \frac{\hat{I}_{i}^{2} - \hat{j}_{i}^{2}}{2\mathcal{J}_{i}}$$

内禀部分:

集体部

$$\hat{H}_{intr} = \sum_{\nu} \varepsilon_{p,\nu} a_{p,\nu}^+ a_{p,\nu} + \sum_{\nu'} \varepsilon_{n,\nu'} a_{n,\nu'}^+ a_{n,\nu'}$$



价核子哈密顿量用单-j哈密顿量给出:  $h_{sp} = \pm \frac{1}{2} C \{ \cos \gamma (j_3^2 - \frac{j(j+1)}{3}) + \frac{\sin \gamma}{2\sqrt{3}} (j_+^2 + j_-^2) \}$ 对于z个质子和n个中子的体系,内禀波函数  $|\varphi\rangle = (\prod_{i=1}^{z_1} a_{p,\nu_i}^{\dagger}) (\prod_{i=1}^{z_2} a_{p,\bar{\mu}_i}^{\dagger}) (\prod_{i=1}^{z_1} a_{p,\nu_i'}^{\dagger}) (\prod_{i=1}^{z_2} a_{p,\bar{\mu}_i'}^{\dagger}) |0\rangle$ 

Bohr, Mottelson, Nuclear Structure, Vol. 2 (1975); QI, PLB (2009)



空间科学与物理学院 \_\_\_\_\_\_school of Space Science and Physics

附录

#### 理论框架

体系波函数: 
$$|IM\rangle = \sum_{K\varphi} c_{K\varphi} |IMK\varphi\rangle$$
  
 $|IMK\varphi\rangle = \frac{1}{\sqrt{2(1 + \delta_{K0}\delta_{\varphi,\bar{\varphi}})}} (|IMK\rangle|\varphi\rangle + (-1)^{I-K}|IM-K\rangle|\bar{\varphi}\rangle)$ 

约化电磁跃迁几率:

$$B(\sigma\lambda, I' \to I) = \frac{1}{2I+1} \sum_{\mu M} \left| \left\langle IM \left| \hat{M}(\sigma\lambda, \mu) \right| I'M' \right\rangle \right|^2$$
$$\hat{M}(M1, \mu) = \sqrt{\frac{3}{4\pi}} \frac{e}{2Mc} \left[ (g_p - g_R) \hat{j}_{p\mu} + (g_n - g_R) \hat{j}_{n\mu} \right]$$
$$\hat{M}(E2, \mu) = \sqrt{\frac{5}{16\pi}} \hat{Q}_{2\mu}$$

B. Qi et al., Phys. Lett. B 675, 175 (2009).

A. D. Ayangeakaa et al., Phys. Rev. Lett. 110, 172504 (2013).I. Kuti et al., Phys. Rev. Lett. 113, 03 (2014)



空间科学与物理学院 \_\_\_\_\_school of Space Science and Physics

形状相变+摇摆



Critical point symmetry for odd-odd nuclei and collective multiple chiral doublet bands Y. Zhang, BQ, and S.Q. Zhang SCIENCE CHINA, 64 ,122011 (2021)



空间科学与物理学院 \_School of Space Science and Physics

## 偶偶核的摇摆

## 公式证明:

Introduce the parameter  $A_k$ :  $\mathcal{J}_k = \frac{\hbar^2}{2A_k}$ 

$$\begin{split} H &= A_1 I_1^2 + A_2 I_2^2 + A_3 I_3^2 = A_3 I^2 + H' \\ H' &= \frac{1}{2} (A_2 + A_1 - 2A_3) (I_2^2 + I_1^2) + \frac{1}{2} (A_2 - A_1) (I_2^2 - I_1^2) = \frac{1}{2} \alpha \frac{I_2^2 + I_1^2}{I} + \frac{1}{2} \beta \frac{I_2^2 - I_1^2}{I} \\ \alpha &= (A_2 + A_1 - 2A_3) I, \quad \beta = (A_2 - A_1) I \end{split}$$

Considering the **approximation** 

$$[I_{-}, I_{+}] = 2I_3 \approx 2I \quad (I_{\pm} = I_2 \pm iI_1)$$

We introduce the operator

$$c^{\dagger} = \frac{1}{\sqrt{2I}}I_+, c = \frac{1}{\sqrt{2I}}I_-$$

Then  $H' = \frac{1}{2}\alpha(c^{\dagger}c + cc^{\dagger}) + \frac{1}{2}\beta(c^{\dagger}c^{\dagger} + cc)$ 



空间科学与物理学院 \_School of Space Science and Physics



## 公式证明: Introduce $c^{\dagger} = x\hat{c}^{\dagger} + y\hat{c}, \quad \hat{c}^{\dagger} = xc^{\dagger} - yc$ $x, y = \left[\frac{1}{2}\left(\frac{\alpha}{(\alpha^2 - \beta^2)}\right)^{1/2} \pm 1\right)\right]^{1/2}, \ x^2 - y^2 = 1$ $H' = \frac{1}{2}\alpha(c^{\dagger}c + cc^{\dagger}) + \frac{1}{2}\beta(c^{\dagger}c^{\dagger} + cc)$ $=\sqrt{\alpha^2 - \beta^2} [\hat{c}^{\dagger}\hat{c} + \frac{1}{2}] = \hbar\omega(\hat{n} + \frac{1}{2})$ $\hbar\omega = \sqrt{\alpha^2 - \beta^2} = \sqrt{(A_2 + A_1 - 2A_3)^2 I^2 - (A_2 - A_1)^2 I^2}$ $= 2I[(A_2 - A_3)(A_1 - A_3)]^{1/2}$ $= 2I \left[ \left( \frac{\hbar^2}{2 \tau_2} - \frac{\hbar^2}{2 \tau_2} \right) \left( \frac{\hbar^2}{2 \tau_1} - \frac{\hbar^2}{2 \tau_2} \right) \right]^{1/2}$



空间科学与物理学院 \_\_\_\_\_school of Space Science and Physics

三轴自由度的影响



For TRM,  $K_m = I-n$  for  $\gamma = 30^\circ$  exactly



- 1. Collective Hamiltonian for wobbling modes Q. B. Chen, S. Q. Zhang, P. W. Zhao, and J. Meng, Phys. Rev. C (2014) 90, 044306.
- 2. Collective Hamiltonian and its applications for chiral and wobbling modes, Q. B. Chen, Acta Phys. Pol. B (2015) 8, 545.
- 3. Wobbling geometry in a simple triaxial rotor, W. X. Shi and Q. B. Chen, Chinese Physics C (2015) 39, 054105.
- 4. Wobbling motion in <sup>135</sup>Pr within a collective Hamiltonian, Q. B. Chen, S. Q. Zhang, and J. Meng, Phys. Rev. C (2016) 94, 054308.
- 5. *Two-dimensional collective Hamiltonian for chiral and wobbling modes*, Q. B. Chen, S. Q. Zhang, P. W. Zhao, R. V. Jolos, and J. Meng Phys. Rev. C (2016) 94, 044301.
- 6. Collective Hamiltonian for chiral and wobbling modes: form one- to two-dimensional, Q.B. Chen, Acta Phys. Pol. B (2017) 10, 27.
- 7. Behavior of the collective rotor in wobbling motion, E. Streck, Q. B. Chen, N. Kaiser, et al., Phys. Rev. C (2018) 98, 044314.
- 8. *Two-dimensional collective Hamiltonian for chiral and wobbling modes. II. Electromagnetic transitions*, X. H. Wu, Q. B. Chen, P. W. Zhao, S. Q. Zhang, and J. Meng, Phys. Rev. C (2018) 98, 064302.
- 9. Experimental Evidence for Transverse Wobbling in <sup>105</sup>Pd J. Timár, Q. B. Chen, et al., Phys. Rev. L (2019) 122, 062501.
- 10. Transverse wobbling in an even-even nucleus, Q. B. Chen, S. Frauendorf, and C. M. Petrache, Phys. Rev. C (2019) 100, 061301(R).
- 11. First Observation of Multiple Transverse Wobbling Bands of Different Kinds in <sup>183</sup>Au, Nandi, Mukherjee, Q. B. Chen, et al., Phys. Rev. L (2020) 125, 132501.
- 12. g-factor and static quadrupole moment for the wobbling mode in <sup>133</sup>La, Q.B. Chen, et al., Phys. Lett. B (2020) 807, 135596.
- 13. Two quasiparticle wobbling in the even-even nucleus <sup>130</sup>Ba Y.K.Wang, F.Q.Chen, P.W.Zhao, Phys. Lett. B802(2020)135246.
- Microscopic investigation on the existence of transverse wobbling under the effect of rotational alignment: The <sup>136</sup>Nd case F. Q. Chen and C. M. Petrache, Phys. Rev. C (2021) 103, 064319.
- 15. Study of wobbling modes by means of spin coherent state maps, Q. B. Chen, S. Frauendorf, Eur. Phys. J. A (2022) 58, 75.
- 16. Dynamics of rotation in chiral nuclei, Z. X. Ren, P. W. Zhao, and J. Meng, Phys. Rev. C (2022) 105, L011301.

••••••



空间科学与物理学院\_\_School of Space Science and Physics

建议的摇摆带

#### 将31+和22+态建议为候选摇摆带的带首





空间科学与物理学院 \_school of Space Science and Physics



BQ\*, Zhang, Wang, Q.B.Chen\*, JPG 48 (2021) 055102





空间科学与物理学院 \_\_\_\_\_\_school of Space Science and Physics



41



\*





间科学与物理学院

hool of Space Science and Physics





空间科学与物理学院 \_\_School of Space Science and Physics

手征摇摆共存





空间科学与物理学院 \_\_\_\_\_\_school of Space Science and Physics

## odd-odd nuclei

手征摇摆共存



带3&4	相比	带1&2
有摇摆	激发	的运动

手征和摇摆如何共存 还需讨论和澄清!

Jia, Wang\*, BQ, Liu, Zhu, PLB 833, 137303 (2022)



