

Spin dynamics of **triaxial** nuclei with quasiparticle alignments



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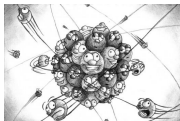
Wobbling - phenomenological interpretation of the excited rotational states [A. Bohr and B. R. Mottelson, *Nuclear Structure* (1975)]

Chirality - handedness of the trihedral corners formed by three spin vectors with respect to the total angular momentum vector.

[S. Frauendorf and J. Meng, *Nucl. Phys. A* **617**, 131 (1997)]

Particle-rotor Hamiltonian

$$H = H_R + H_{sp}$$



$$H_R = \sum_{k=1,2,3} A_k (\hat{I}_k - \hat{j}_k - \hat{j}'_k)^2 \text{ triaxial rotor Hamiltonian}$$

$$\vec{R} = \vec{I} - \vec{j} - \vec{j}' \text{ core angular momentum}$$

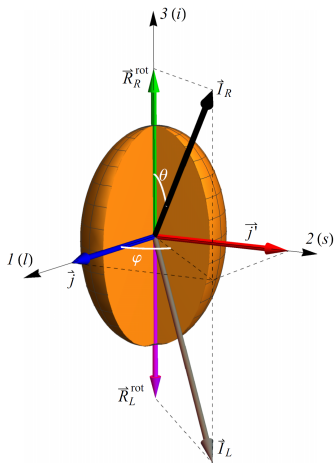
$$j \ \& \ j' \text{ single-particle spins } j(j') = 11/2$$

$$A_k \text{ the inertial parameters related to the MOI by } A_k = \frac{1}{2\mathcal{J}_k}$$

- Hydrodynamic MOI assumption: $\mathcal{J}_k = \frac{4}{3} \mathcal{J}_0 \sin^2 \left(\gamma - \frac{2}{3} k\pi \right)$
- Axes 1,2,3 become long, short, intermediate $\implies \gamma \in (60^\circ, 120^\circ)$

$$R_k = R_0 \left[1 + \sqrt{\frac{5}{4\pi}} \beta \cos \left(\gamma - \frac{2\pi}{3} k \right) \right]$$

- (1) long axis - alignment of the hole angular momentum
- (2) short axis - alignment of the particle angular momentum
- (3) medium axis - core rotation with maximal MOI



Rigidly aligned quasiparticles

$$\hat{j}_1 = j_1 \cos \alpha, \quad \hat{j}'_1 = 0,$$

$$\hat{j}_2 = 0, \quad \hat{j}'_2 = j' \cos \alpha',$$

$$\hat{j}_3 = j \sin \alpha, \quad \hat{j}'_3 = j' \sin \alpha'$$

The relevant Hamiltonian to be treated

$$H_{align} =$$

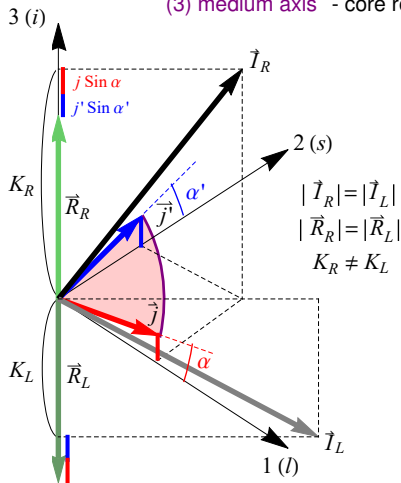
$$A_1 \hat{I}_1^2 + A_2 \hat{I}_2^2 + A_3 \hat{I}_3^2$$

$$-2A_1 j \cos \alpha \hat{I}_1 - 2A_2 j' \cos \alpha' \hat{I}_2$$

$$-2A_3 (j \sin \alpha + j' \sin \alpha') \hat{I}_3$$

Transverse wobbling 🖐 $j = 0$ and small α'

- (1) long axis - alignment of the hole angular momentum
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Rigidly aligned quasiparticles

$$\begin{aligned} \hat{j}_1 &= j_1 \cos \alpha, \quad \hat{j}'_1 = 0, \\ \hat{j}_2 &= 0, \quad \hat{j}'_2 = j' \cos \alpha', \\ \hat{j}_3 &= j \sin \alpha, \quad \hat{j}'_3 = j' \sin \alpha' \end{aligned}$$

The relevant Hamiltonian to be treated

$$\begin{aligned} H_{align} = & \\ & A_1 \hat{I}_1^2 + A_2 \hat{I}_2^2 + A_3 \hat{I}_3^2 \\ & - 2A_1 j \cos \alpha \hat{I}_1 - 2A_2 j' \cos \alpha' \hat{I}_2 \\ & - 2A_3 (j \sin \alpha + j' \sin \alpha') \hat{I}_3 \end{aligned}$$

Transverse wobbling 🖐️ $j = 0$ and small α'

- A description by means of only few classical variables associated to some particular dynamics of the quantum system - **is desired**

↓ Solution 🖱️ **The semiclassical approach**

[RB, Phys. Rev. C **97**, 024302 (2018); **98**, 014303 (2018), RB, Phys. Lett. B **797**, 134853 (2019); **817**, 136308 (2021)]

- Relies on a time-dependent variational principle applied to a variational state which is constructed according to the problem.

$$\delta \int_0^t \langle \psi | H_{align} - \frac{\partial}{\partial t'} | \psi \rangle dt' = 0$$

- The variational principle provides the time-dependence of some restricted set of variables which parametrize the variational state:

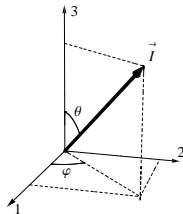
$$\begin{aligned} |\psi_{IM}(\theta, \varphi)\rangle &= \sum_{K=-I}^I \frac{1}{2^I} \sqrt{\frac{(2I)!}{(I-K)!(I+K)!}} (1 + \cos \theta)^{\frac{I-K}{2}} (1 - \cos \theta)^{\frac{I+K}{2}} e^{i\varphi(I+K)} |IMK\rangle \\ &= \frac{1}{\left(1 + \tan^2 \frac{\theta}{2}\right)^I} e^{\tan \frac{\theta}{2} e^{i\varphi} \hat{I}_-} |IMI\rangle \end{aligned}$$

Stereographic representation:

$$0 \leq \theta < \pi, \quad 0 \leq \varphi < 2\pi$$

Projection variable $x = I \cos \theta$, $-I < x \leq I$

$\equiv K$ projection on axis 3



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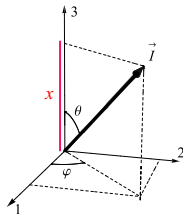
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≡ K projection on axis 3



- Solving the equations provided by the variational principle is equivalent to solving the eigenvalue equation associated to the quantum Hamiltonian H_{align} .

The full structure of the classical system is reproduced if the variables are canonical:

$$\frac{\partial \mathcal{H}}{\partial x} = \dot{\varphi}, \quad \frac{\partial \mathcal{H}}{\partial \varphi} = -\dot{x} \quad \text{or} \quad \{x, \mathcal{H}\} = \dot{x}, \quad \{\varphi, \mathcal{H}\} = \dot{\varphi}, \quad \{f, g\} = \frac{\partial f}{\partial \varphi} \frac{\partial g}{\partial x} - \frac{\partial f}{\partial x} \frac{\partial g}{\partial \varphi}$$

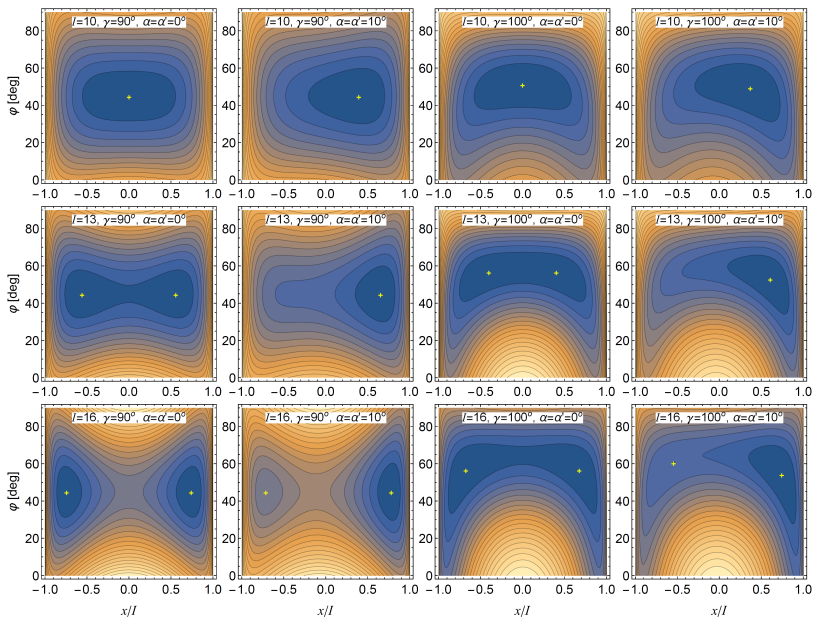
$$\{\varphi, x\} = 1 \quad \begin{array}{l} \varphi \text{ the generalized coordinate} \\ x \text{ the generalized momentum} \end{array} \quad \text{Poisson bracket}$$

Classical energy function in terms of the canonical variables:

$$\begin{aligned} \mathcal{H}(x, \varphi) &= \frac{I}{2}(A_1 + A_2) + A_3 I^2 \\ &+ \frac{(2I - 1)(I^2 - x^2)}{2I} (A_1 \cos^2 \varphi + A_2 \sin^2 \varphi - A_3) \\ &- 2\sqrt{I^2 - x^2} (A_1 j \cos \alpha \cos \varphi + A_2 j' \cos \alpha' \sin \varphi) \\ &- 2A_3 x (j \sin \alpha + j' \sin \alpha'). \end{aligned}$$

- The classical orbits are closed curves in the phase space of the canonical coordinates which are concentrically positioned around the stationary points of the constant energy surface.

Classical energy function (chiral configuration $j = j' = 11/2$)



- The classical trajectory of the angular momentum vector \vec{I} is a curve in the space of its classical projections

$$\mathcal{I}_1 = \langle \hat{I}_1 \rangle = \sqrt{I^2 - x^2} \cos \varphi,$$

$$\mathcal{I}_2 = \langle \hat{I}_2 \rangle = \sqrt{I^2 - x^2} \sin \varphi,$$

$$\mathcal{I}_3 = \langle \hat{I}_3 \rangle = x.$$

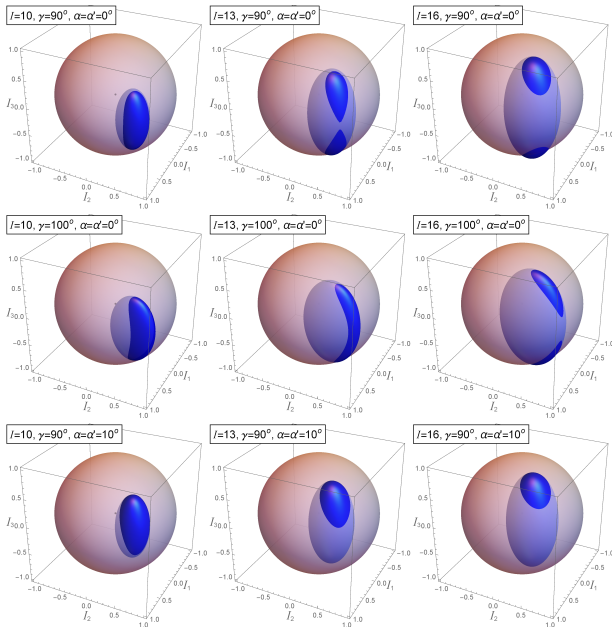
$$\sum_k \langle \hat{I}_k^2 \rangle = I(I + 1)$$

- It is determined by the intersection of the constant energy surfaces provided by the constants of motions:

$$\begin{aligned} \text{Shifted ellipsoid } \mathcal{H} = & A_1 \mathcal{I}_1^2 + A_2 \mathcal{I}_2^2 + A_3 \mathcal{I}_3^2 \\ & - 2A_1 \mathcal{I}_1 j \cos \alpha - 2A_2 \mathcal{I}_2 j' \cos \alpha' \\ & - 2A_3 \mathcal{I}_3 (j \sin \alpha + j' \sin \alpha'), \end{aligned}$$

$$\text{Sphere } I^2 = \mathcal{I}_1^2 + \mathcal{I}_2^2 + \mathcal{I}_3^2.$$

Classical trajectories (chiral configuration $j = j' = 11/2$)



Stationary points for chiral configuration (no tilting)

- The stationary points where $\dot{\varphi} = \dot{x} = 0$ and are stable against fluctuations are those which minimize the classical energy.

Planar

$$x = 0, \quad \frac{(2I - 1)}{2} (A_2 - A_1) \cos \varphi_p \sin \varphi_p = A_2 j' \cos \varphi_p - A_1 j \sin \varphi_p$$

$$\mathcal{I}_1 = I \cos \varphi_p, \quad \mathcal{I}_2 = I \sin \varphi_p, \quad \mathcal{I}_3 = 0.$$

Aplanar

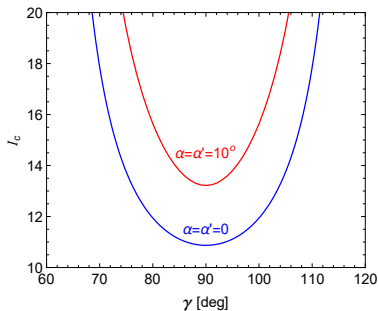
$$x_L^R = \pm I \cos \theta_a^I, \quad \sin \varphi_a = \frac{A_2 j' (A_1 - A_3)}{\sqrt{A_1^2 j^2 (A_2 - A_3)^2 + A_2^2 j'^2 (A_1 - A_3)^2}}$$

$$\mathcal{I}_1 = I \sin \theta_a^I \cos \varphi_a = \frac{2IA_1 j}{(2I - 1)(A_1 - A_3)}, \quad \mathcal{I}_2 = I \sin \theta_a^I \sin \varphi_a = \frac{2IA_2 j'}{(2I - 1)(A_2 - A_3)}$$

$$\mathcal{I}_3 = \pm I \cos \theta_a^I, \quad \sin \theta_a^I = \frac{2\sqrt{A_1^2 j^2 (A_2 - A_3)^2 + A_2^2 j'^2 (A_1 - A_3)^2}}{(2I - 1)(A_1 - A_3)(A_2 - A_3)}$$

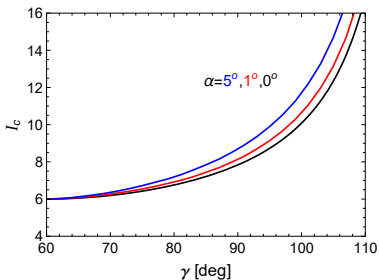
Chiral configuration

$$j = j' = 11/2$$

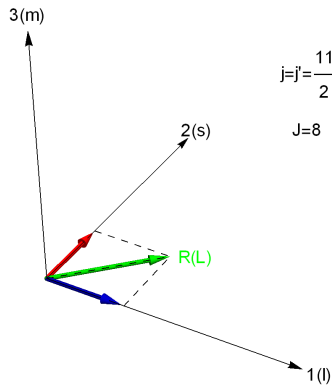


Transverse alignment (wobbling)

$j' = 11/2$, α tilting from the s axis

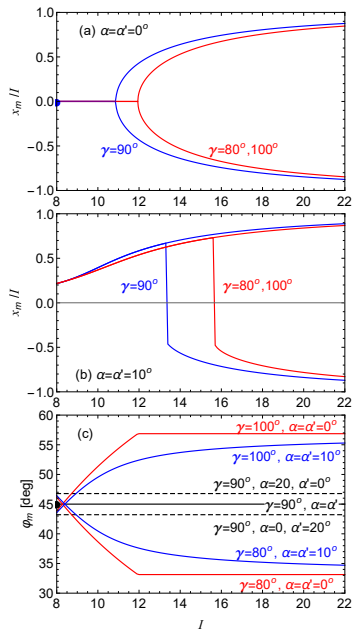


Stationary points

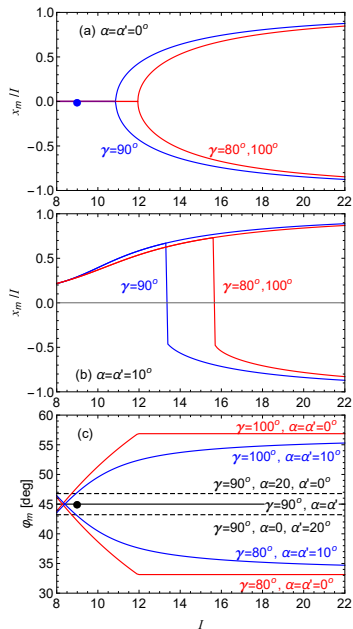
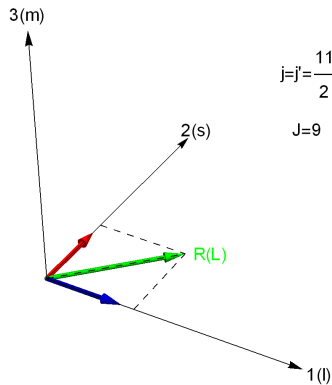


$$j=j'=\frac{11}{2}$$

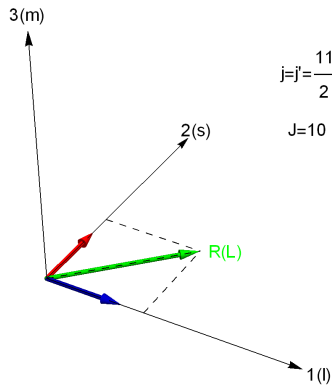
$$J=8$$



Stationary points

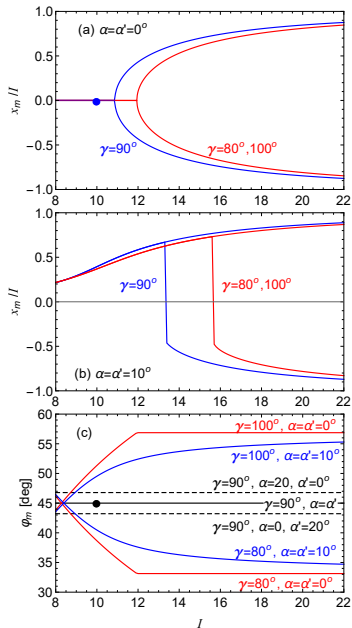


Stationary points

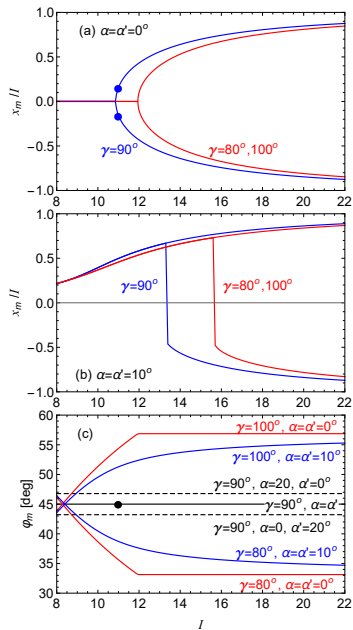
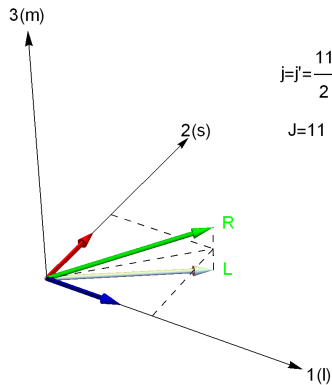


$$j=j'=\frac{11}{2}$$

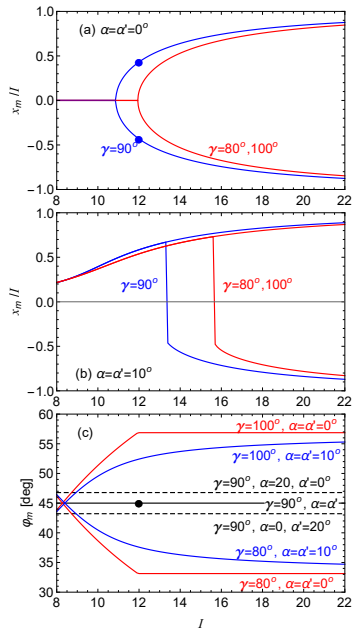
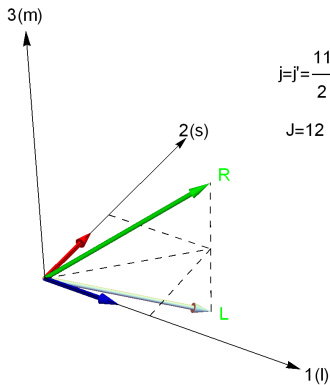
$$J=10$$



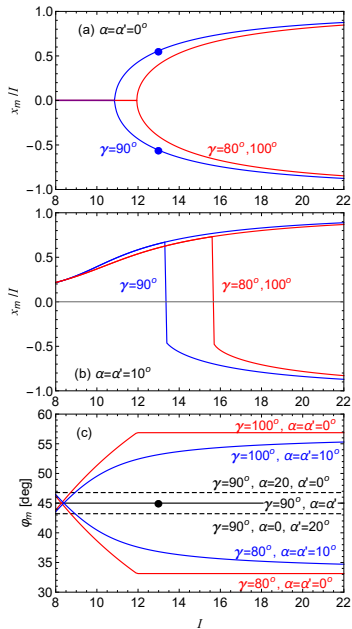
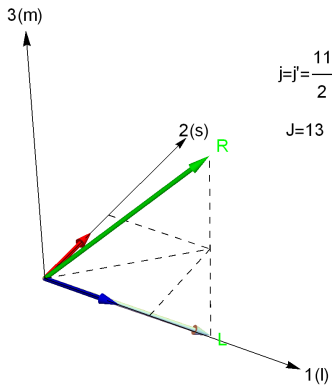
Stationary points



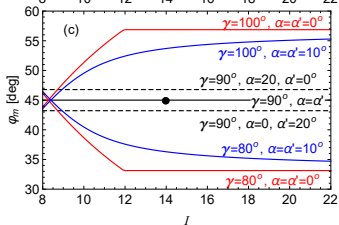
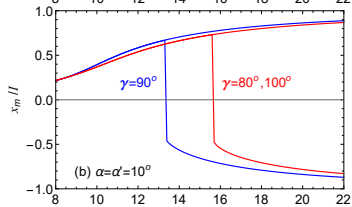
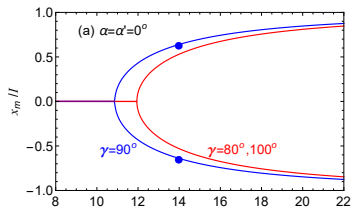
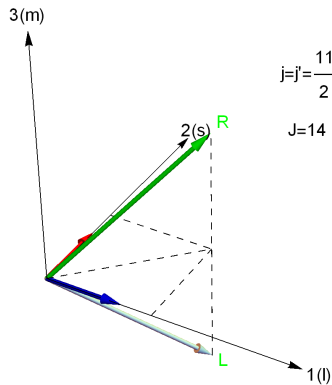
Stationary points



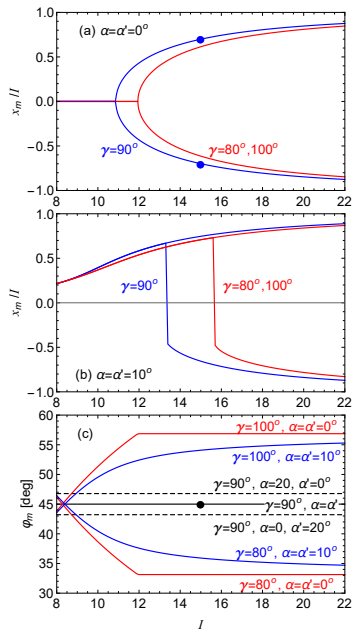
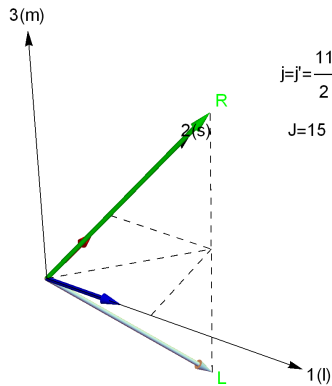
Stationary points



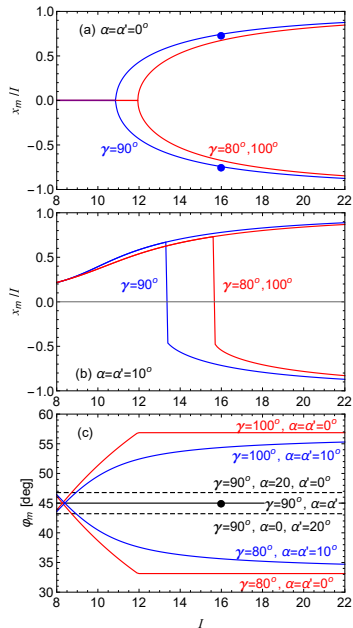
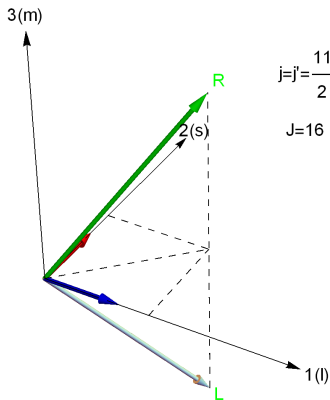
Stationary points



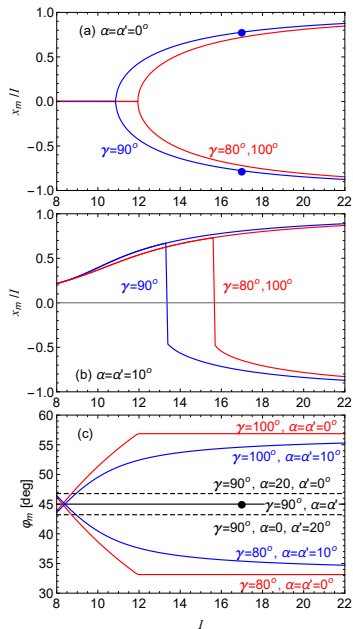
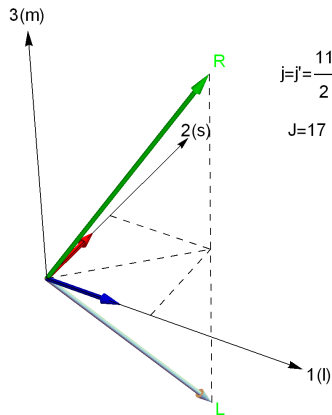
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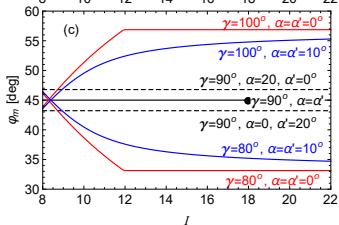
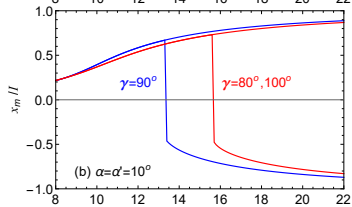
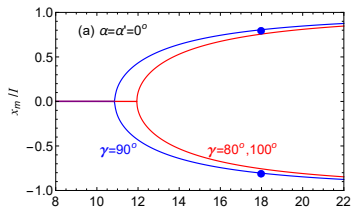
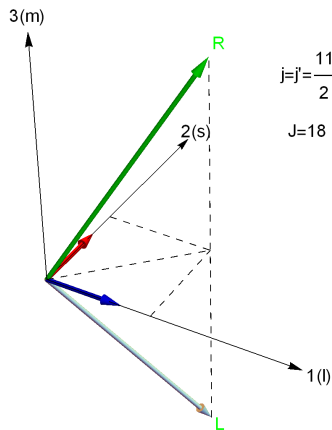
Stationary points



Stationary points



Stationary points



- Energy function is expanded around the corresponding single minimum in $\varphi_0(x)$ for fixed values of x :

$$\tilde{\mathcal{H}}(x, \varphi) \approx \mathcal{H}(r, \varphi_0(x)) + \frac{1}{2} \left(\frac{\partial^2 \mathcal{H}}{\partial \varphi^2} \right)_{\varphi_0(x)} [\varphi - \varphi_0(x)]^2,$$

- Quantization of the properly symmetrized $\tilde{\mathcal{H}}(x, \varphi)$ with $\varphi \rightarrow i \frac{d}{dx}$

$$\hat{H}_c = -\frac{1}{2} \frac{1}{\sqrt{B(x)}} \frac{d}{dx} \frac{1}{\sqrt{B(x)}} \frac{d}{dx} + V(x),$$

Effective mass $B(x) = \left[\frac{\partial^2 \mathcal{H}(x, \varphi)}{\partial \varphi^2} \right]_{\varphi_0(x)}^{-1}$

Chiral(wobbling) potential $V(x) = \mathcal{H}(x, \varphi_0(x)) + \frac{B''(x)}{8 [B(x)]^2} - \frac{9 [B'(x)]^2}{32 [B(x)]^3}$.

$$\varphi_0(x) \begin{cases} \text{wobbling configuration (axial\&planar rotation)} \\ = 0 (\text{Mod } \pi/2) = \text{const.} \quad [\text{RB, C. M. Petrache, Phys. Rev. C } \mathbf{106}, 014313 (2022)] \\ \text{chiral configuration (planar\&aplanar rotation)} \\ = \frac{j' \cos \alpha'}{j \cos \alpha} = \text{const.} \quad \gamma = 90^\circ, \quad A_1 = A_2 \quad [\text{RB, Phys. Lett. B } \mathbf{817}, 136308 (2021)] \\ = f(x), \quad \gamma \neq 90^\circ \quad f(x) - \text{numerically determined} \quad [\text{RB, Phys. Lett. B } \mathbf{797}, 134853 (2019)] \end{cases}$$

Eigensystem is bounded by $|x| \leq I$



Particle in the box basis states

$$f_I^s(x) = \frac{[B(x)]^{-\frac{1}{4}}}{\sqrt{I}} \left\{ \sum_{n=1}^{n_{Max}} A_n^s \cos \left[\frac{(2n-1)\pi x}{2I} \right] + \sum_{n=1}^{n_{Max}} B_n^s \sin \left[\frac{2n\pi x}{2I} \right] \right\},$$



Solutions $F_{Is}(x) = f_I^s(x) [B(x)]^{\frac{1}{4}}$

Quantum number of the bands $n = s - 1$, s - solution's order.

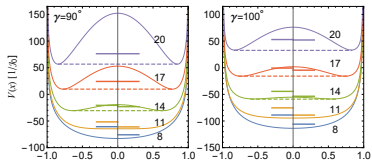
- Total energy $E(I, n) = E_{diag}(\mathcal{J}_0, \gamma, \alpha, \alpha'; Is) + CI(I + 1) + E_0$

- Probability distribution from the chiral Hamiltonian $\rho_s^I(x) = |F_{Is}(x)|^2$

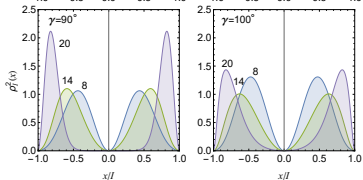
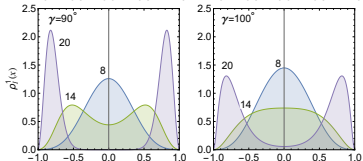
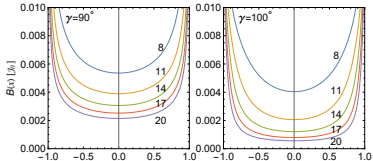
- Total wave function $|\Psi_{IMs}\rangle = \sum_{K=-I}^I \left[\sum_{K=-I}^I F_{Is}(K)^2 \right]^{-1/2} F_{Is}(K) |IKM\rangle$

- For $\alpha = \alpha' = 0$, s plays the role of a parity-like quantum number associated to the $\pm x$ symmetry of the problem.

Potential →



Effective mass →



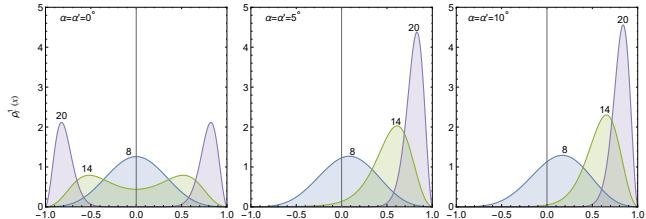
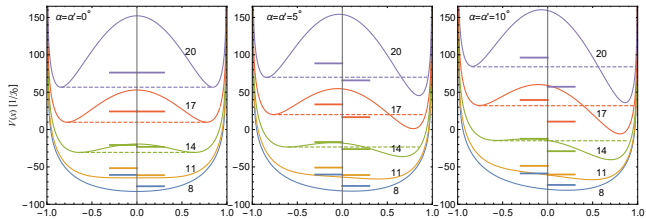
$j = j' = 11/2$
 $\alpha = \alpha' = 0^\circ$ No tilting
 Principal axis alignment

← ground band

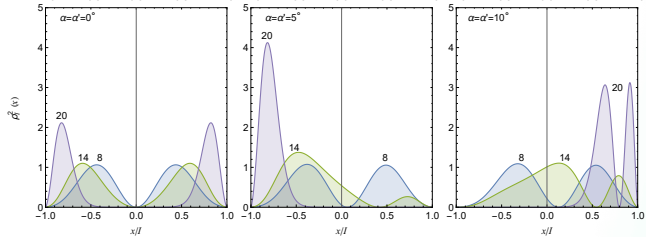
← excited band

$$j = j' = 11/2$$

$$\gamma = 90^\circ$$



← ground band



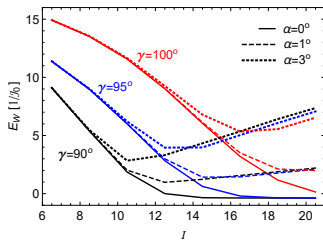
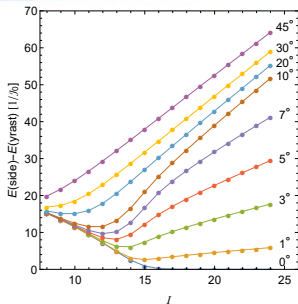
← excited band

Chiral Geometry

$$j = j' = 11/2$$

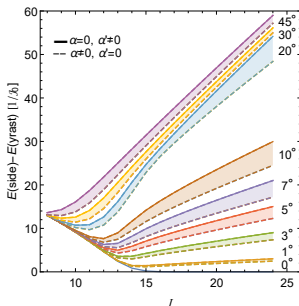
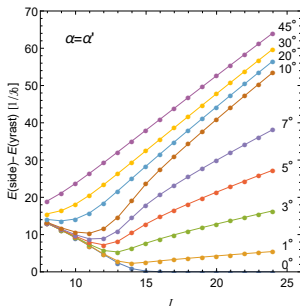
$$\gamma = 90^\circ \rightarrow$$

$$\alpha = \alpha'$$



Transverse wobbling $j' = 11/2, j = 0 \uparrow$

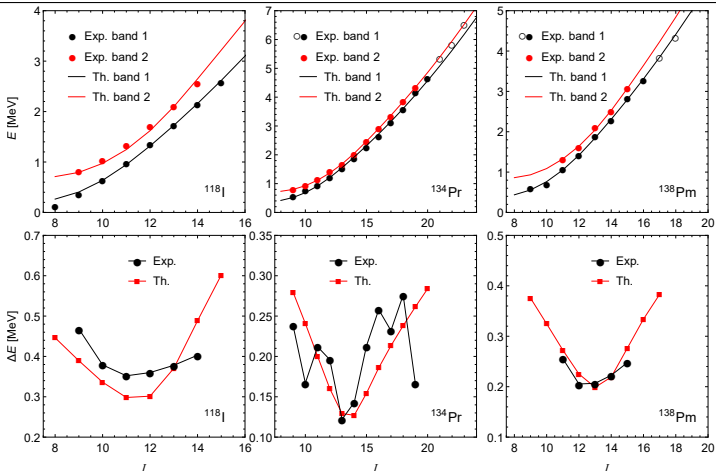
$$E_W(I) = E(I) - [E(I+1) + E(I-1)]/2$$



Chiral geometry $j = 9/2, j' = 11/2, \gamma = 90^\circ \rightarrow$

Applications - Chiral configuration with tilted alignments and $\gamma = 90^\circ$

Nucl.	j (hole)	j' (particle)	α	α'	E_0 [MeV]	\mathcal{J}_0 [MeV $^{-1}$]	C [keV]	rms [keV]
^{118}I	$9/2 (\pi)$	$11/2 (\nu)$	3°	10°	2.293	29.294	0.19	61.4
^{134}Pr	$11/2 (\nu)$	$11/2(\pi)$	3°	3°	1.759	48.076	3.24	43.9
^{138}Pm	$11/2 (\nu)$	$11/2(\pi)$	0°	8°	2.303	35.681	3.34	31.4



Applications - Chiral configuration with tilted alignments and $\gamma = 90^\circ$

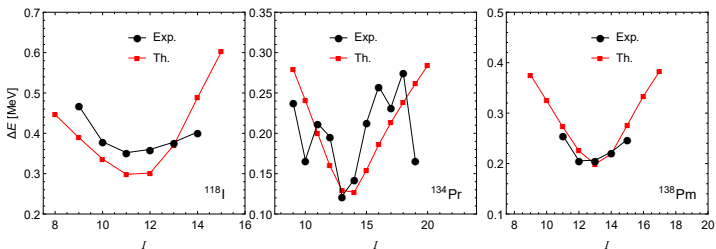
Nucl.	j (hole)	j' (particle)	α	α'	E_0 [MeV]	\mathcal{J}_0 [MeV $^{-1}$]	C [keV]	rms [keV]
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^{134}Pr Small tilting \rightarrow still the best candidate for chiral symmetry.

^{138}Pm PRM calculations

[P. Siwach, P. Arumugam, L. S. Ferreira, E. Maglione, Phys. Lett. B **811**, 135937 (2020).]

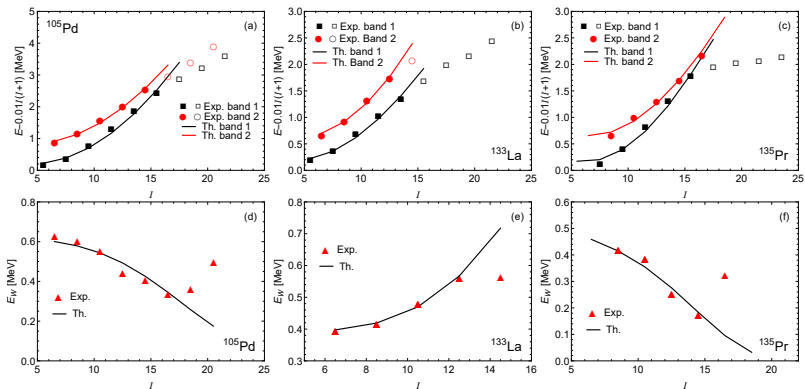
Only the dominant proton quasiparticle configuration has a sizable medium axis component.



[RB, Phys. Lett. B **817**, 136308 (2021); Front. Phys., submitted (2023)]

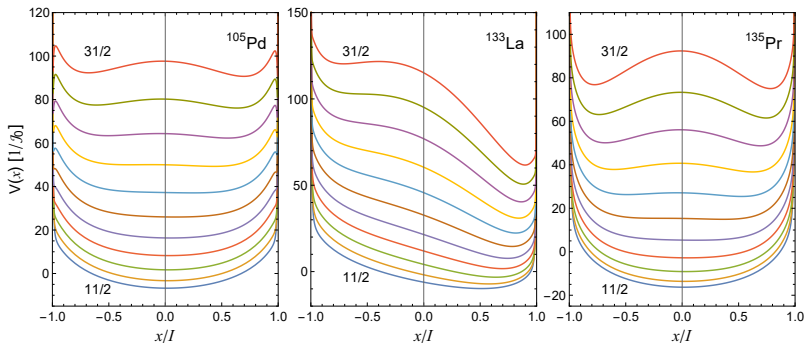
Applications - Wobbling configuration with tilted alignments

Nucl.	γ	α	$1/\mathcal{J}_0$ [keV]	E_0 [keV]	C [keV]	I_c
^{105}Pd	104°	1°	30.49	272.28	6.891	12.52
^{133}La	101°	32°	24.40	325.89	8.049	14.88
^{135}Pr	100°	1°	30.99	679.23	3.822	10.64



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^{133}La

- Tilted axis wobbling.
- Still small tilting, to be considered as a hole quasiparticle.

[T. M. Semkov *et al.*, Phys. Rev. C **34**, 523 (1986)]

[L. Hildingson, W. Klamra, Th. Lindblad, C. G. Lindén, G. Sletten, G. Székely, Z. Phys. A **338**, 125 (1991)]

[C. M. Petrache *et al.*, Phys. Rev. C **94**, 064309 (2016)]

- PRM calculations

[Q. B. Chen, S. Frauendorf, N. Kaiser, Ulf-G. Meißner, J. Meng, Phys. Lett. B **807**, 135596 (2020)]

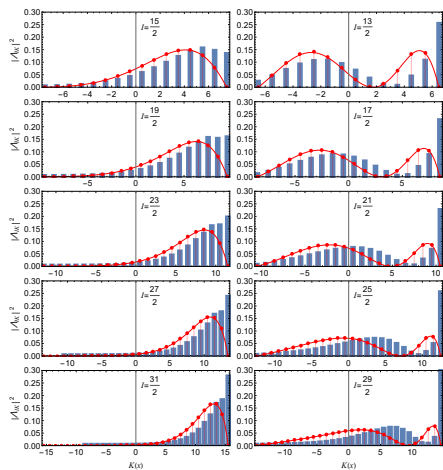
$$\frac{\langle j_{3(m)} \rangle}{\langle j_{2(s)} \rangle} \approx 0.5 \approx \sin \alpha$$

- A particle-rotor system with rigid single-particle spin alignments is studied in a semiclassical approach.
- The existence conditions were identified for distinct dynamic phases.
- A Schrödinger equation is obtained from the classical picture for a continuous projection variable (x), which allows the phenomenological interpretation of the quantum states in terms of oscillations and tilted axis rotations.

References:

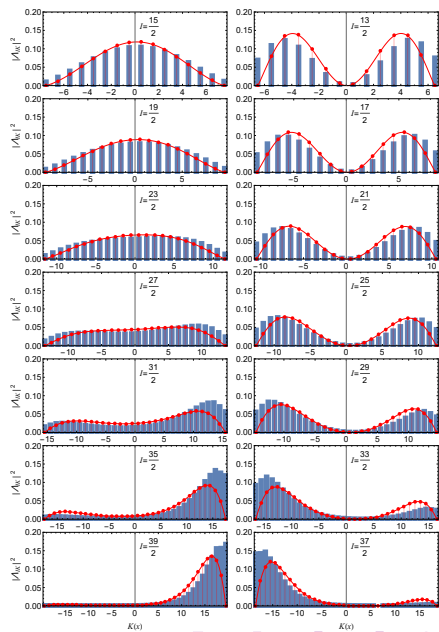
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- RB, A. I. Budaca, Eur. Phys. J. A, submitted.

Wobbling K -plots

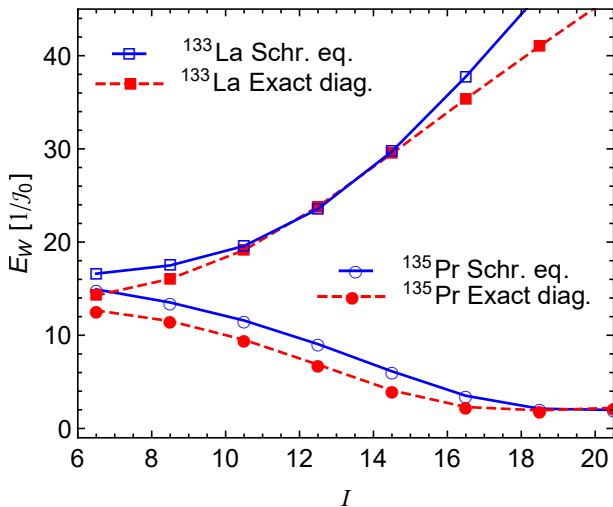


↑
 ^{133}La

$^{135}\text{Pr} \rightarrow$



Schrödinger equation VS exact diagonalization



Second order expansion

$$T_{2\mu}(E2) = t_1 q_{2\mu} + t_2 [q \times q]_{2\mu}$$

in quadrupole moments

$$q_{2\mu} = \beta \left[\cos \gamma D_{\mu 0}^2 + \frac{\sin \gamma}{\sqrt{2}} (D_{\mu 2}^2 + D_{\mu -2}^2) \right]$$

