



GRB 221009A: The BOAT

Antonio de Ugarte Postigo (CNRS - Observatoire de la Côte d'Azur)

7 September 2023

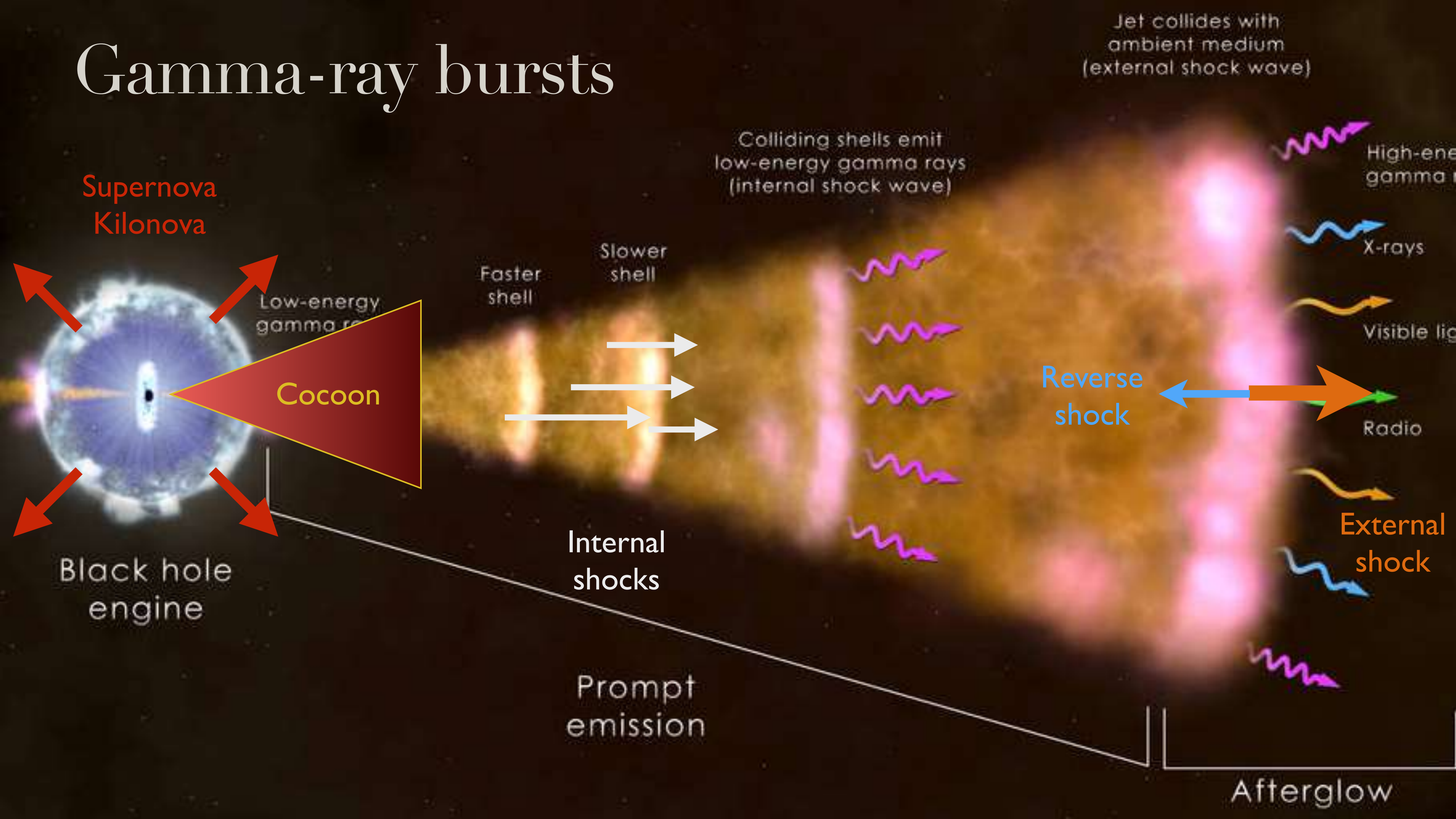


GRB 221009A: The Brightest Of All Time

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Gamma-ray bursts



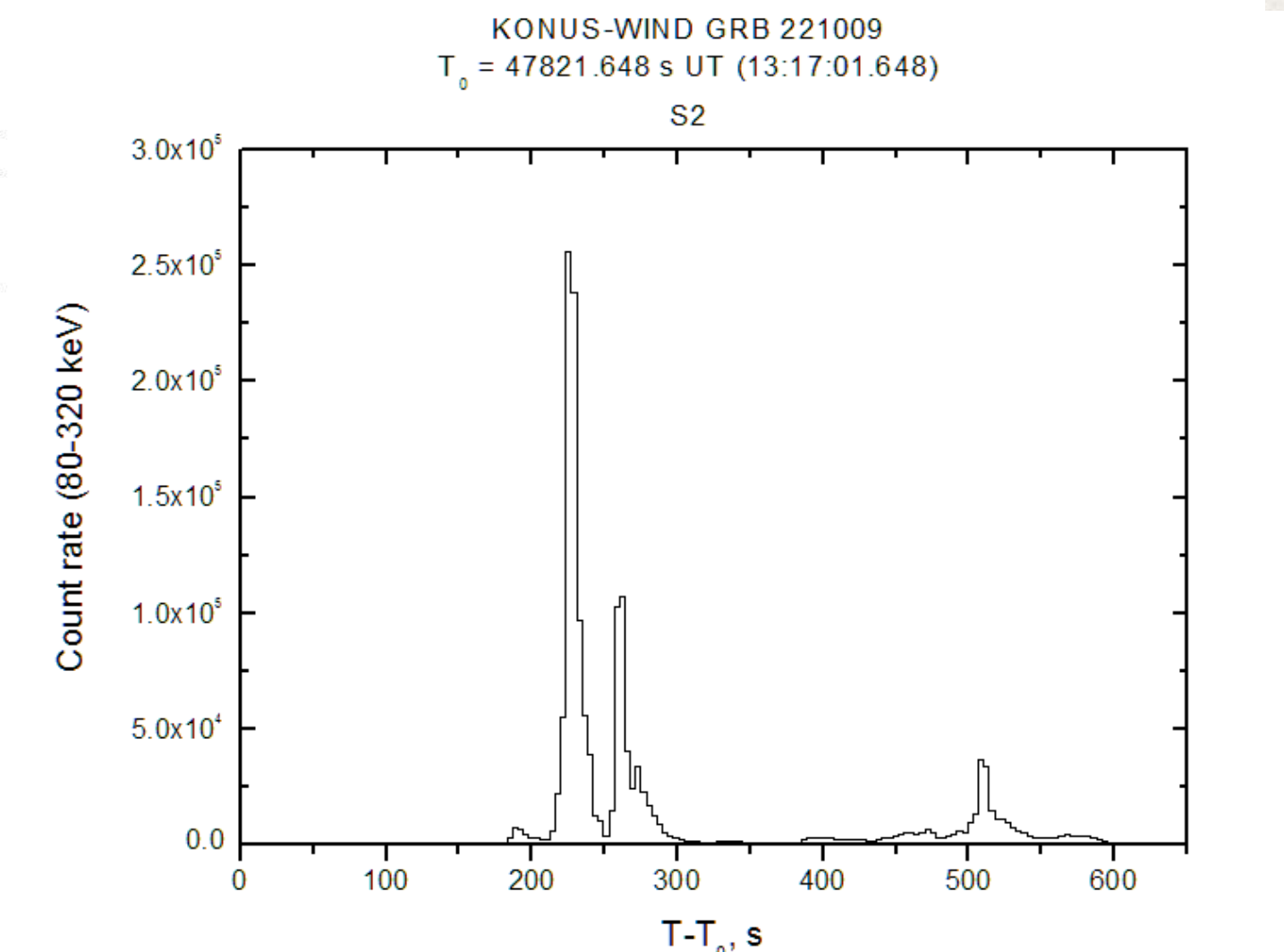
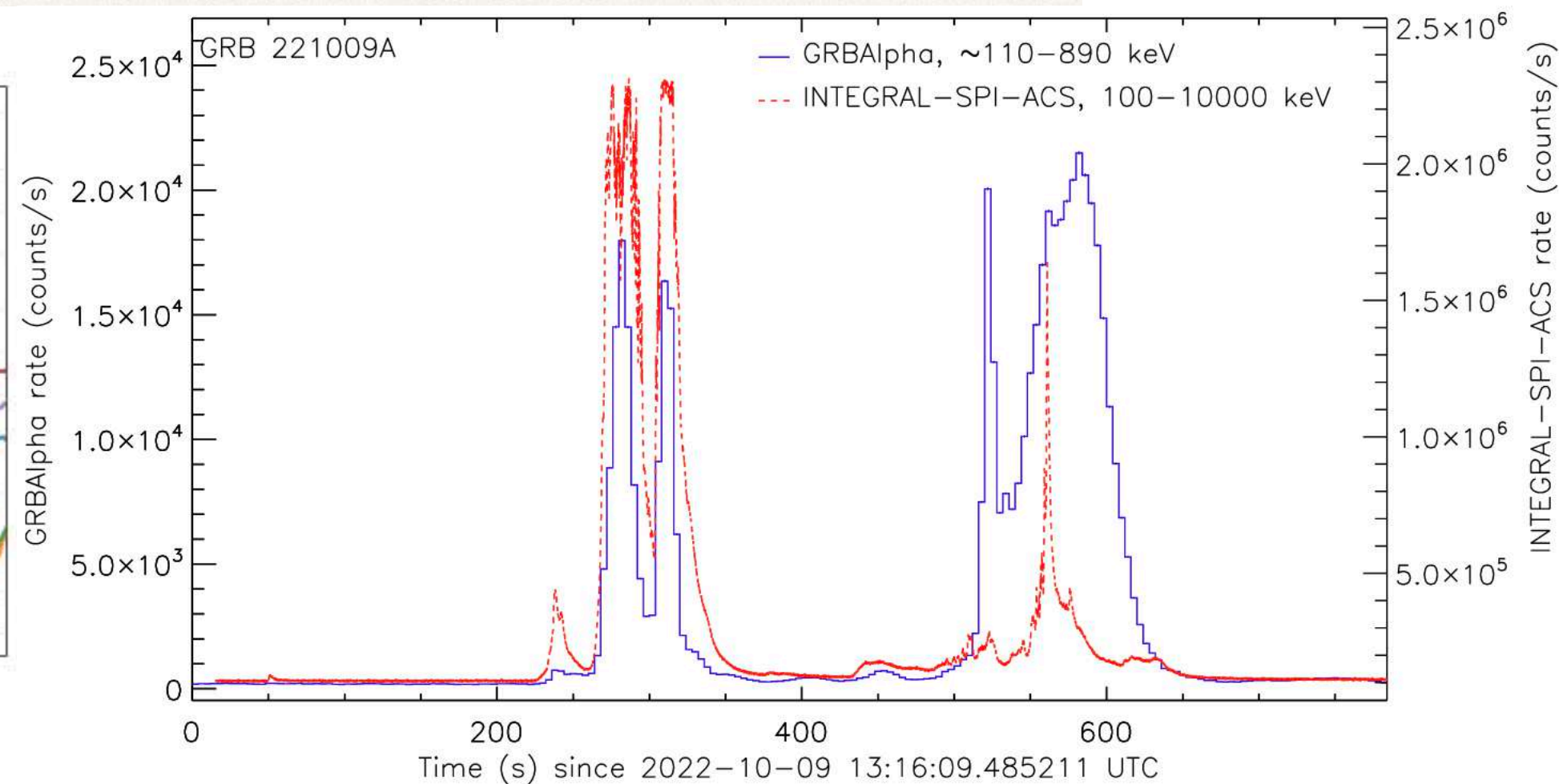
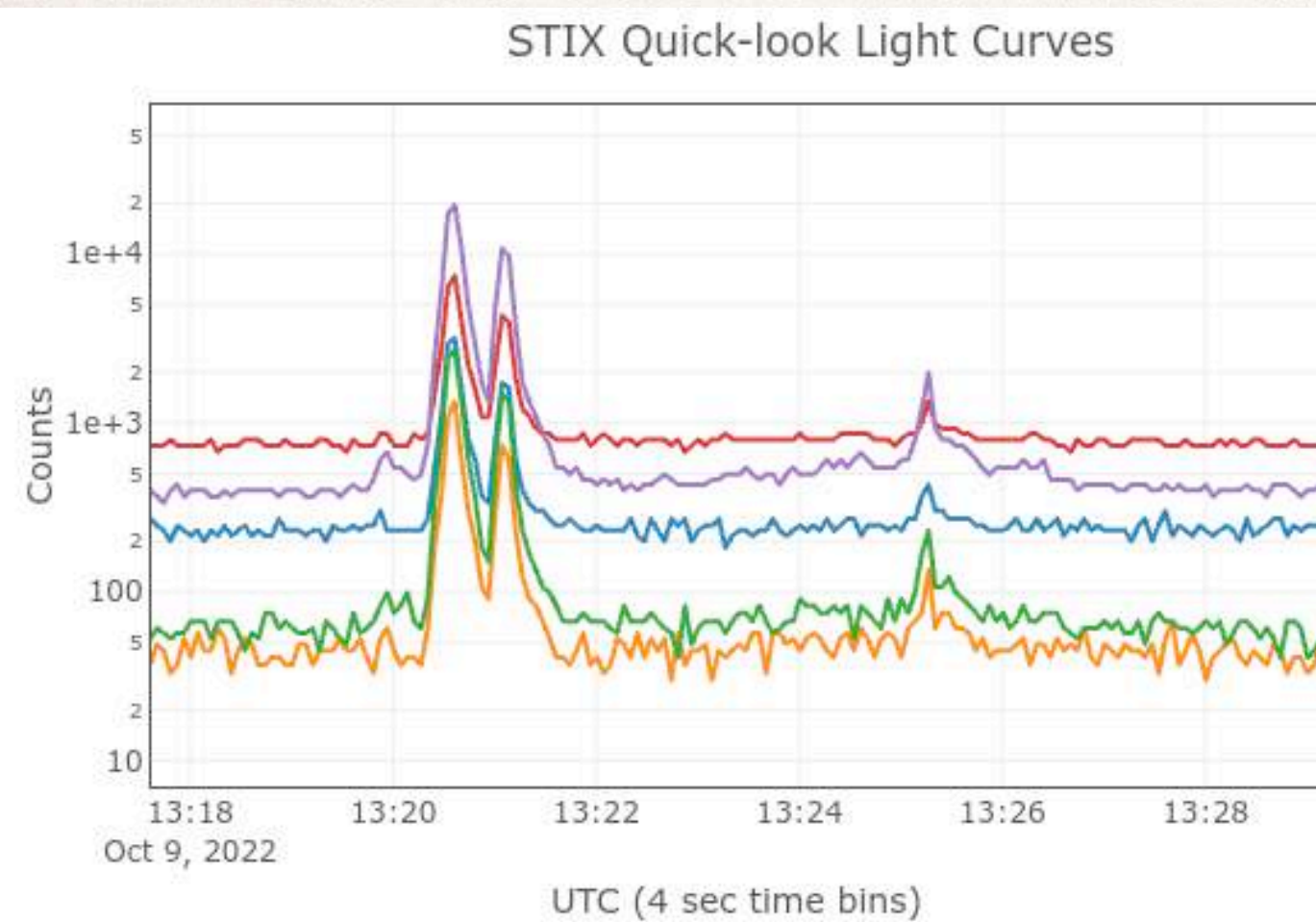
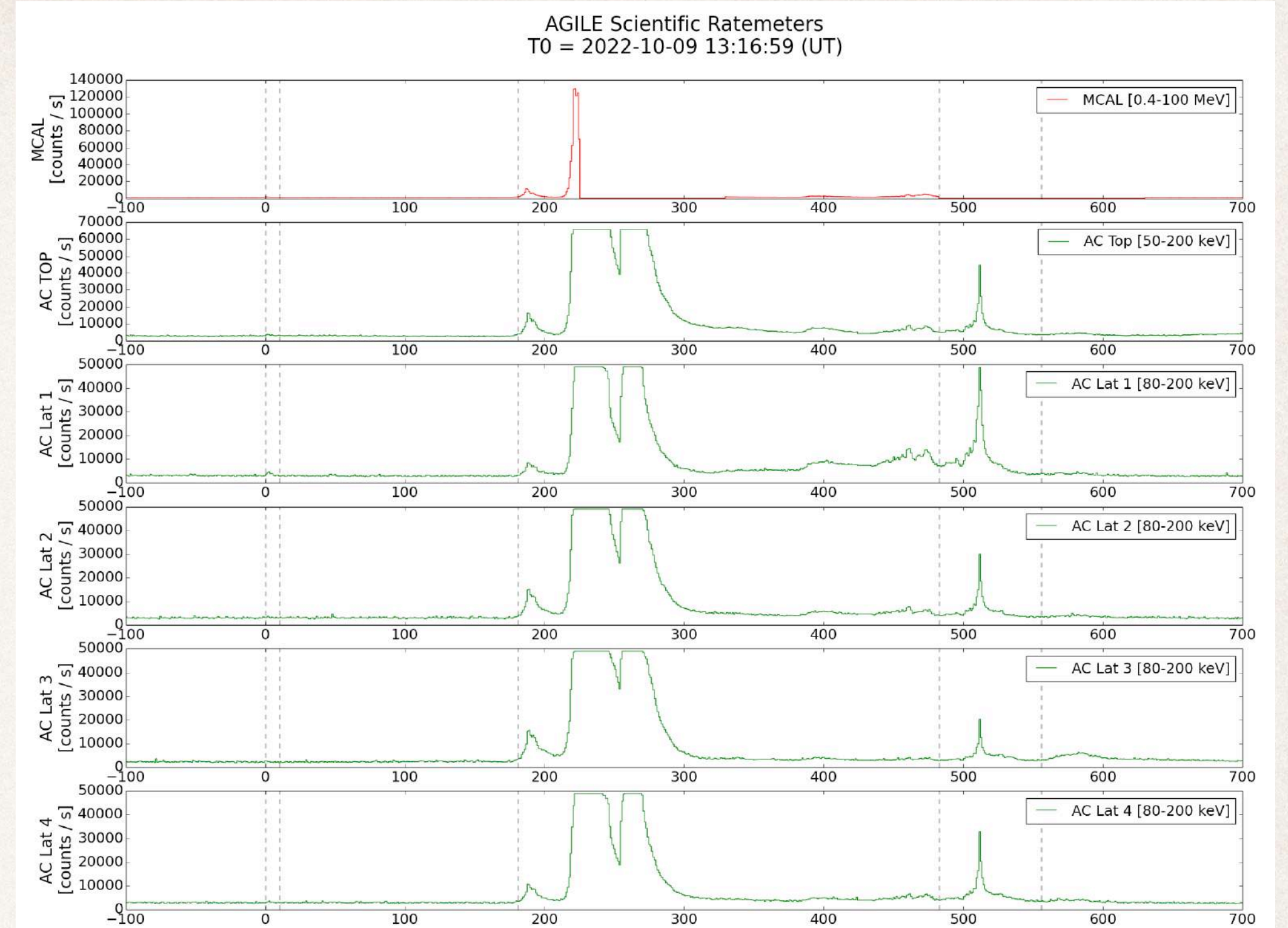
Discovery of GRB 221009A

- ❖ Initially identified by *Swift* as a Galactic source Swift J1913.1+1946
 - ❖ Was in the galactic plane
 - ❖ Slow evolution
 - ❖ Triggered on the afterglow
- ❖ *Fermi*/GBM reported it as an extremely bright GRB almost an hour before *Swift*



An extremely bright event

- ❖ Detected by: *Fermi* (GBM & LAT), *Swift*, *MAXI*, *INTEGRAL*, *Insight-HXMT* / HE, *Konus-Wind*, *IPN*, *AGILE*, *STIX*, *ART-XC*, *GRBA* Alpha, *SIRI-2*, *HEBS*, ...
- ❖ Saturated many of the detectors

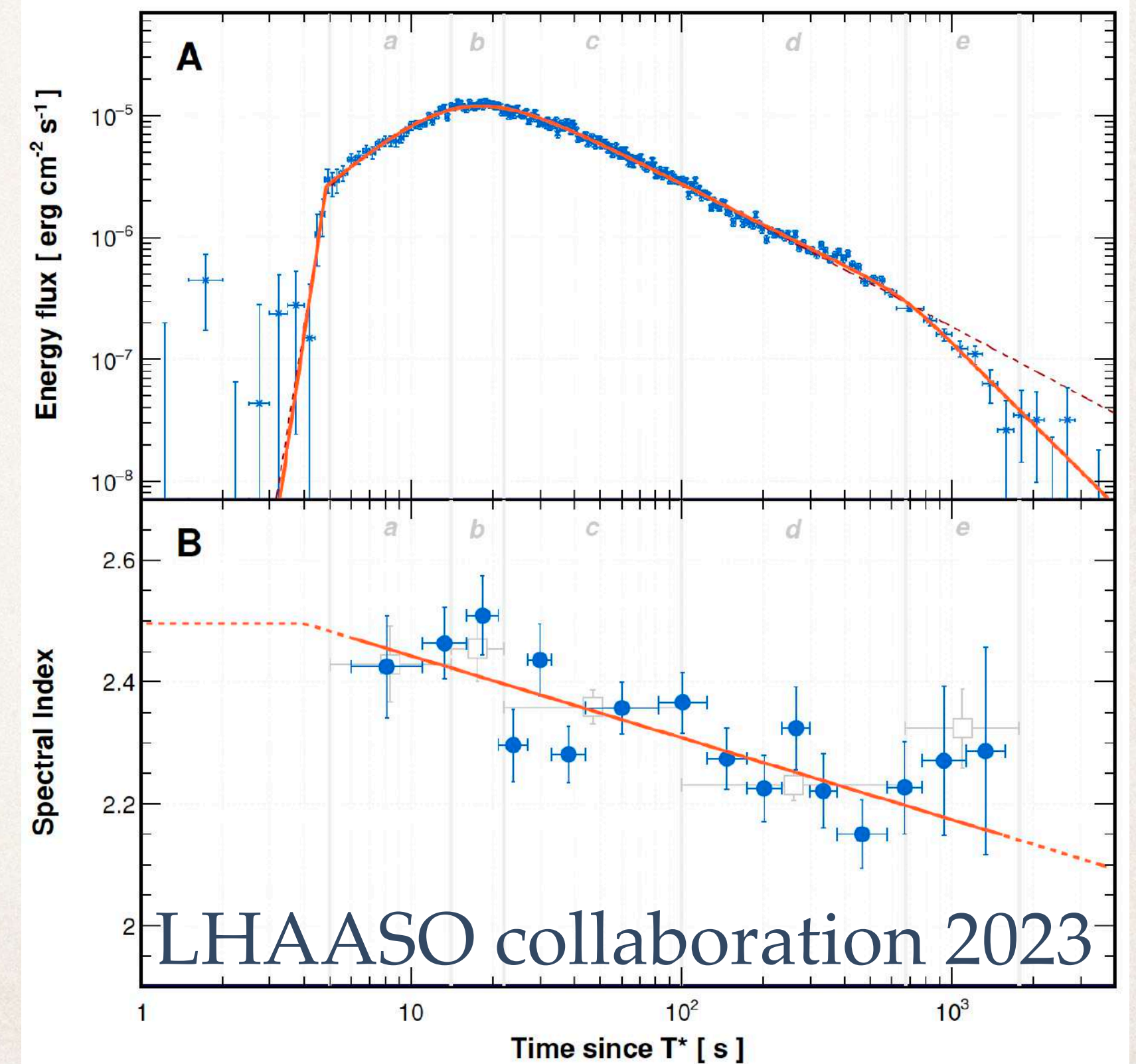


LHAASO

detection at 18TeV



- ❖ Large High Altitude Air Shower Observatory (LHAASO)
- ❖ Sensitive to 10^{11} and 10^{15} eV
- ❖ Located in Tibet
- ❖ More than 64 000 photons above 0.2 TeV
- ❖ Starting minutes after the GRB



Disturbances in the Earth's ionosphere

- ❖ Detected by radio transmitters at < 100 kHz

- ❖ Schnoor et al. GCN 32744, Guha et al. GCN 32745

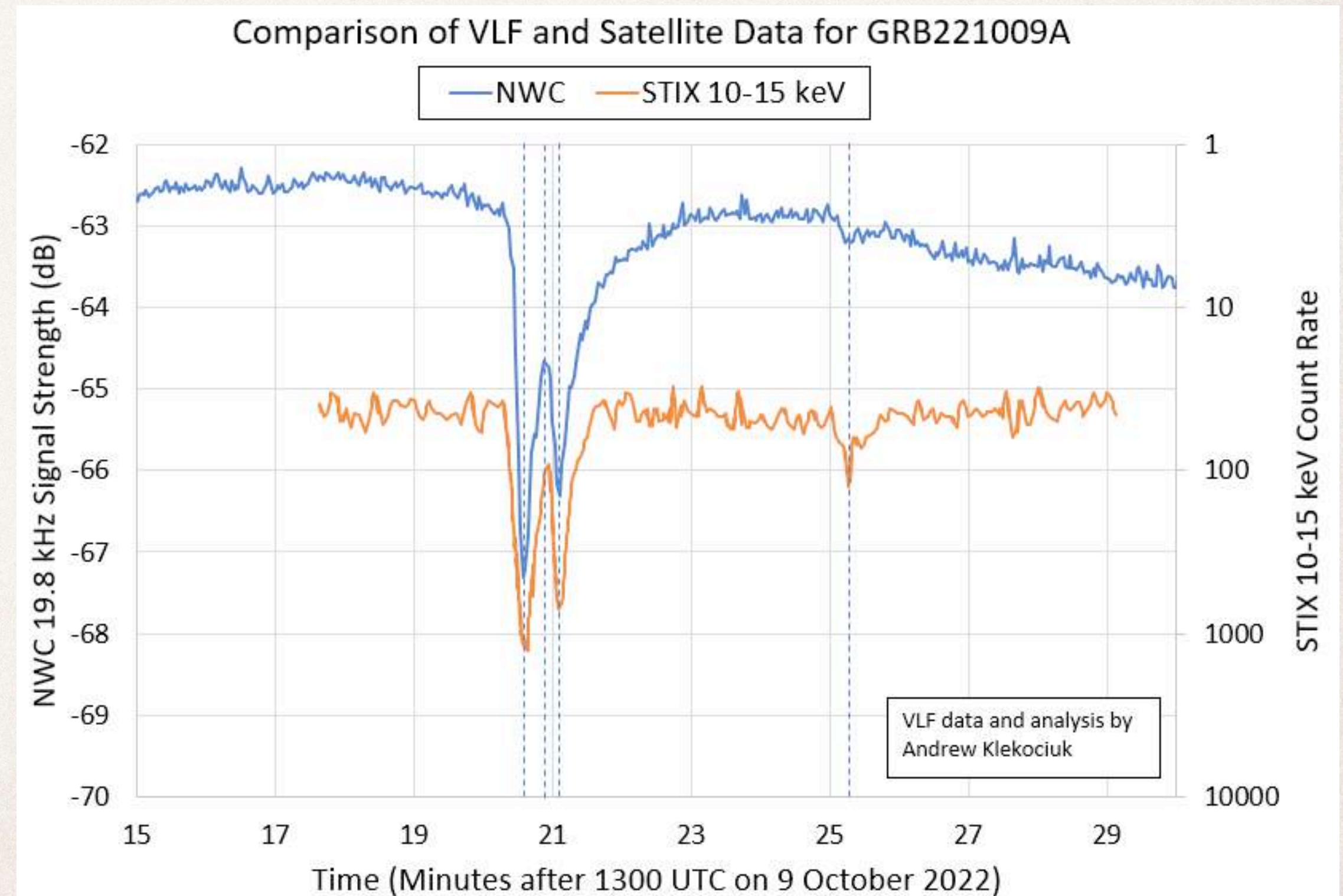
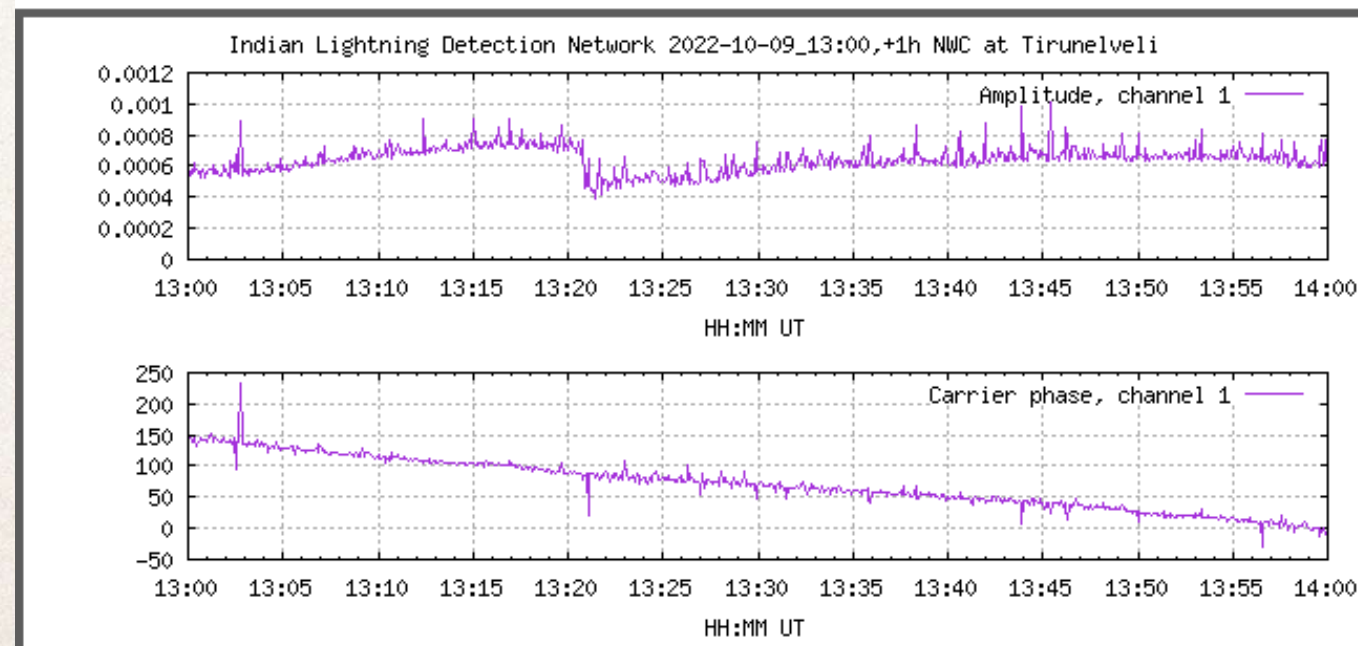
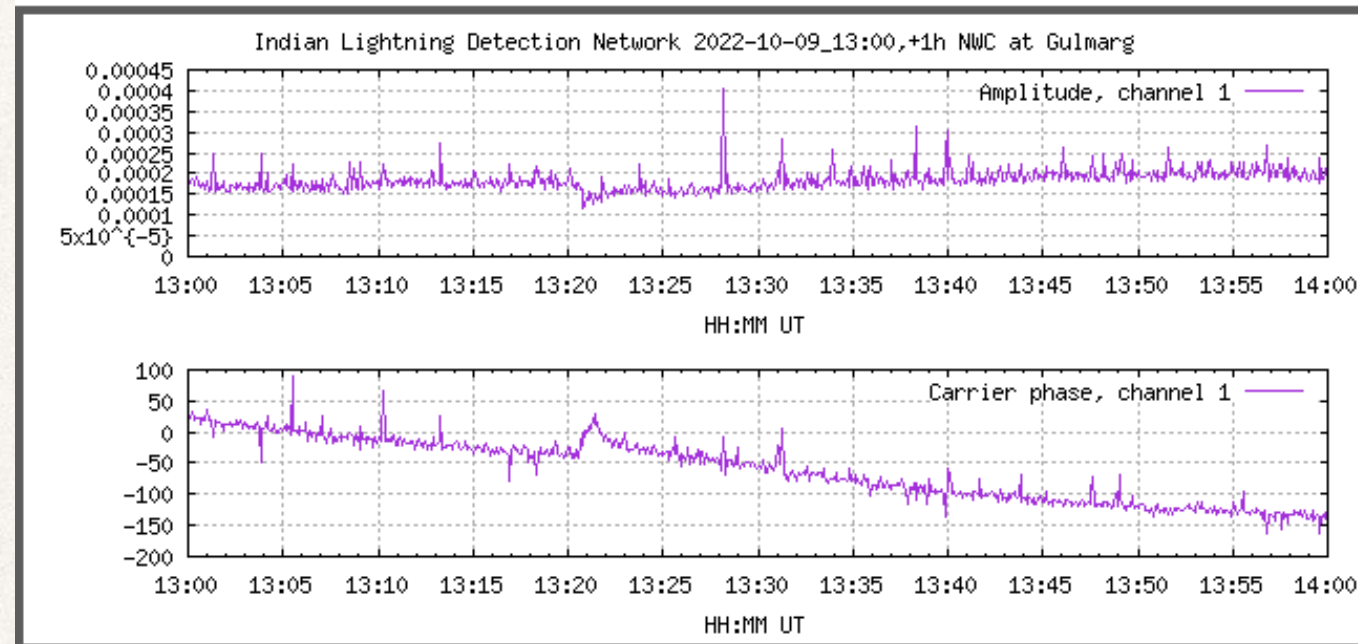
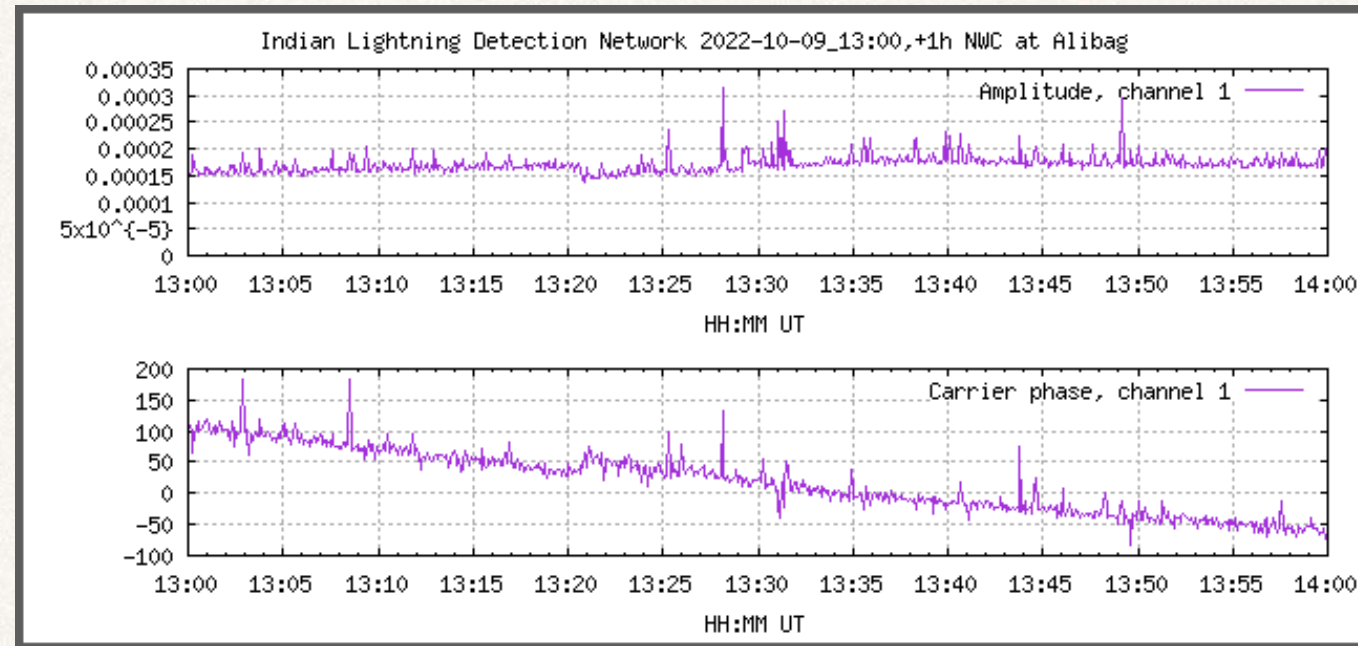


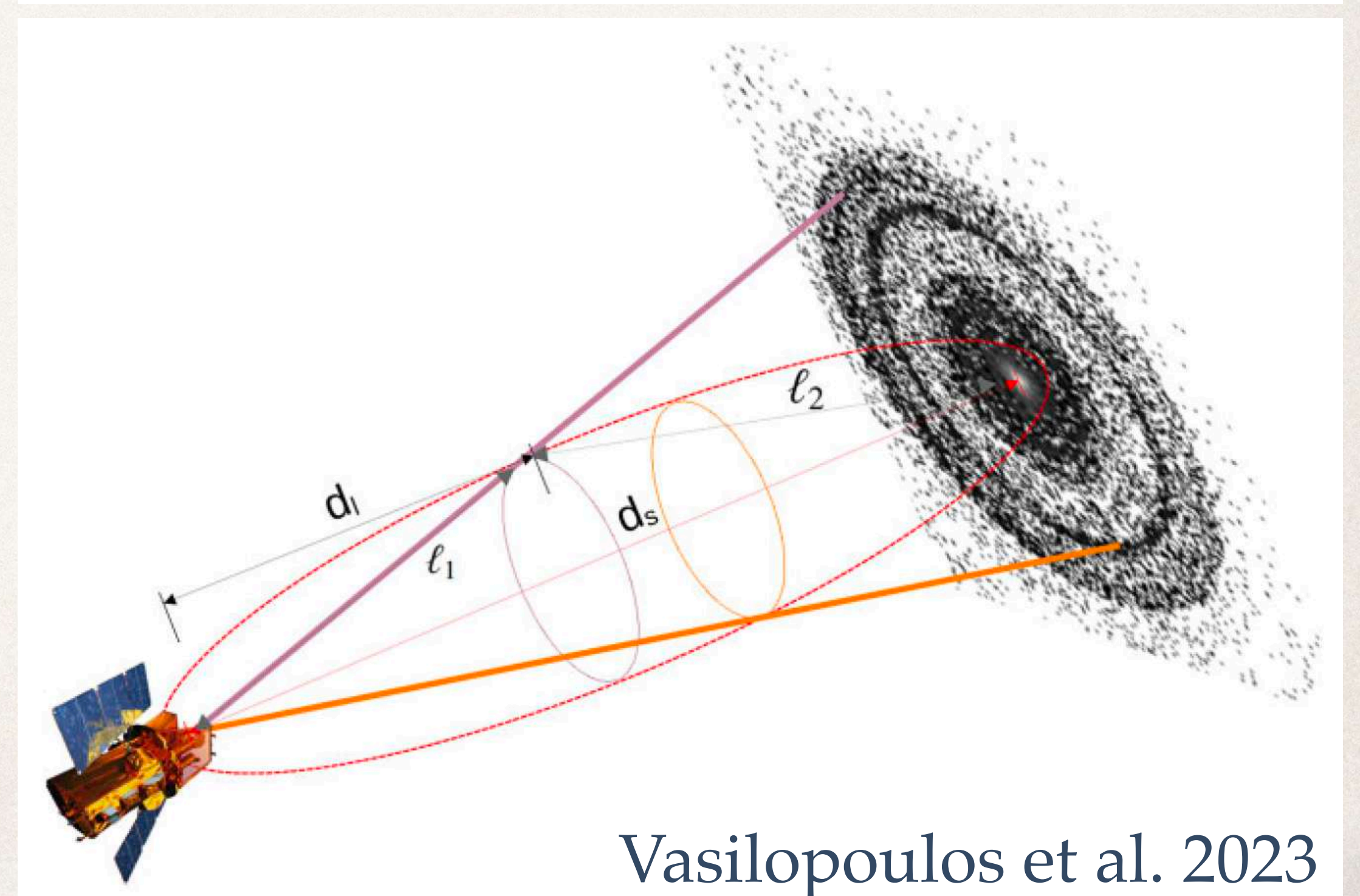
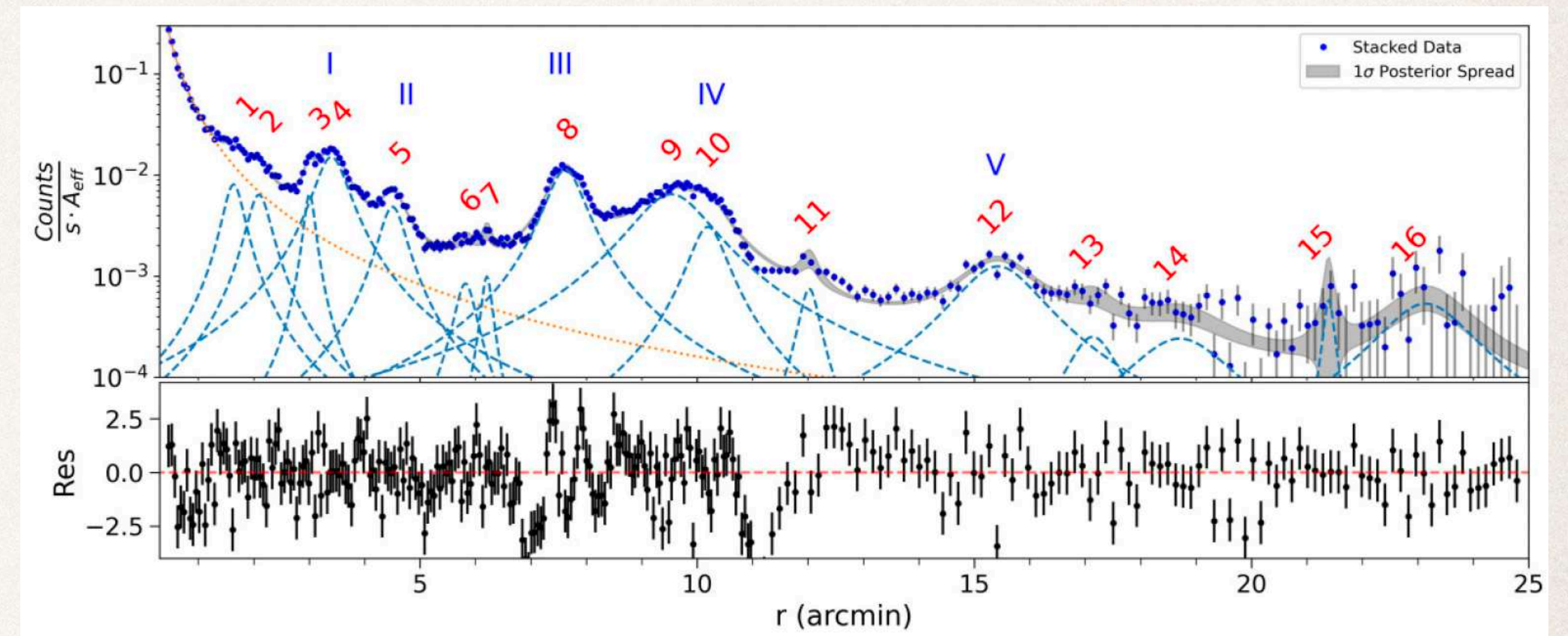
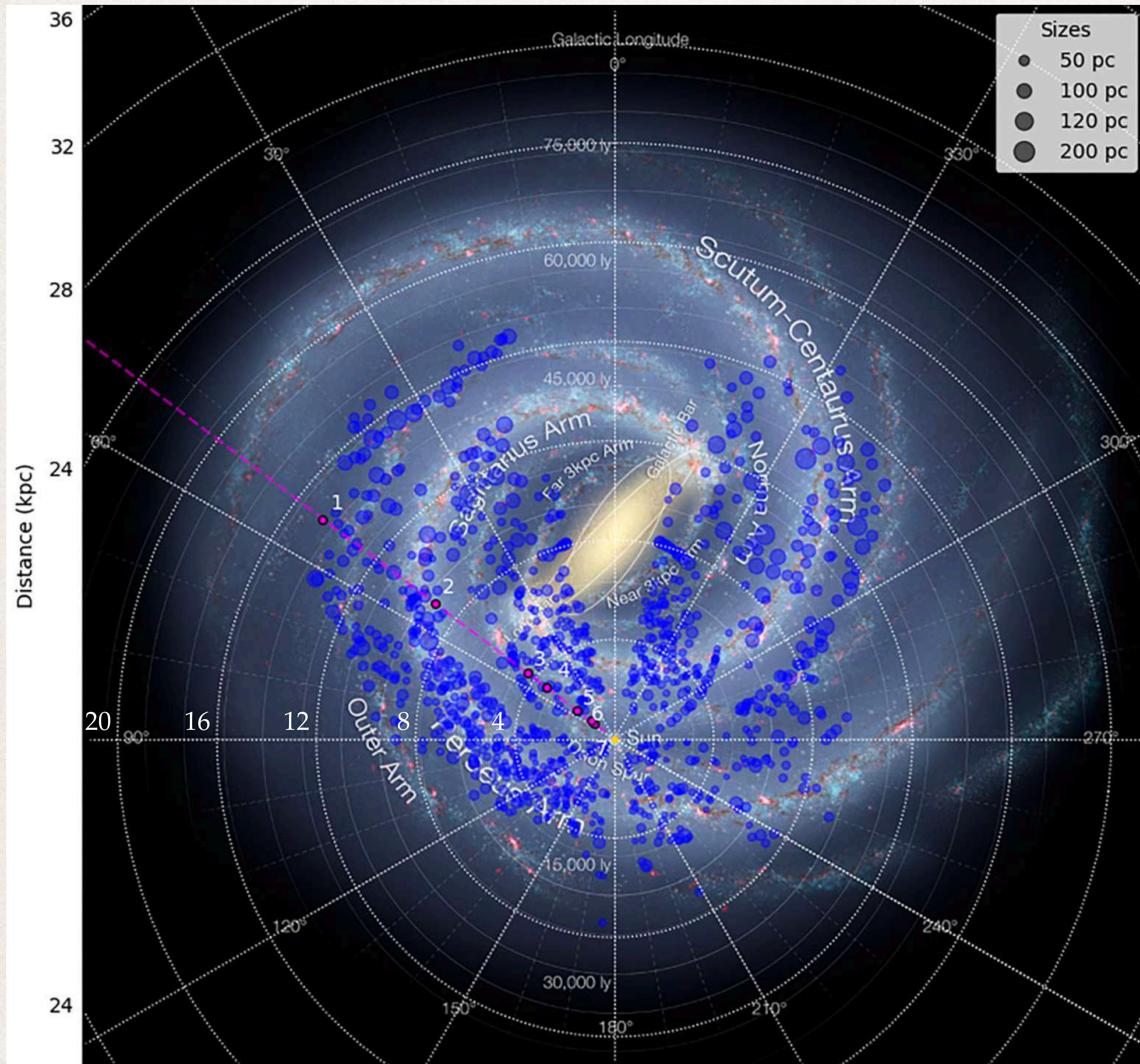
Photo-bombed by the Milky Way

GRB221009A



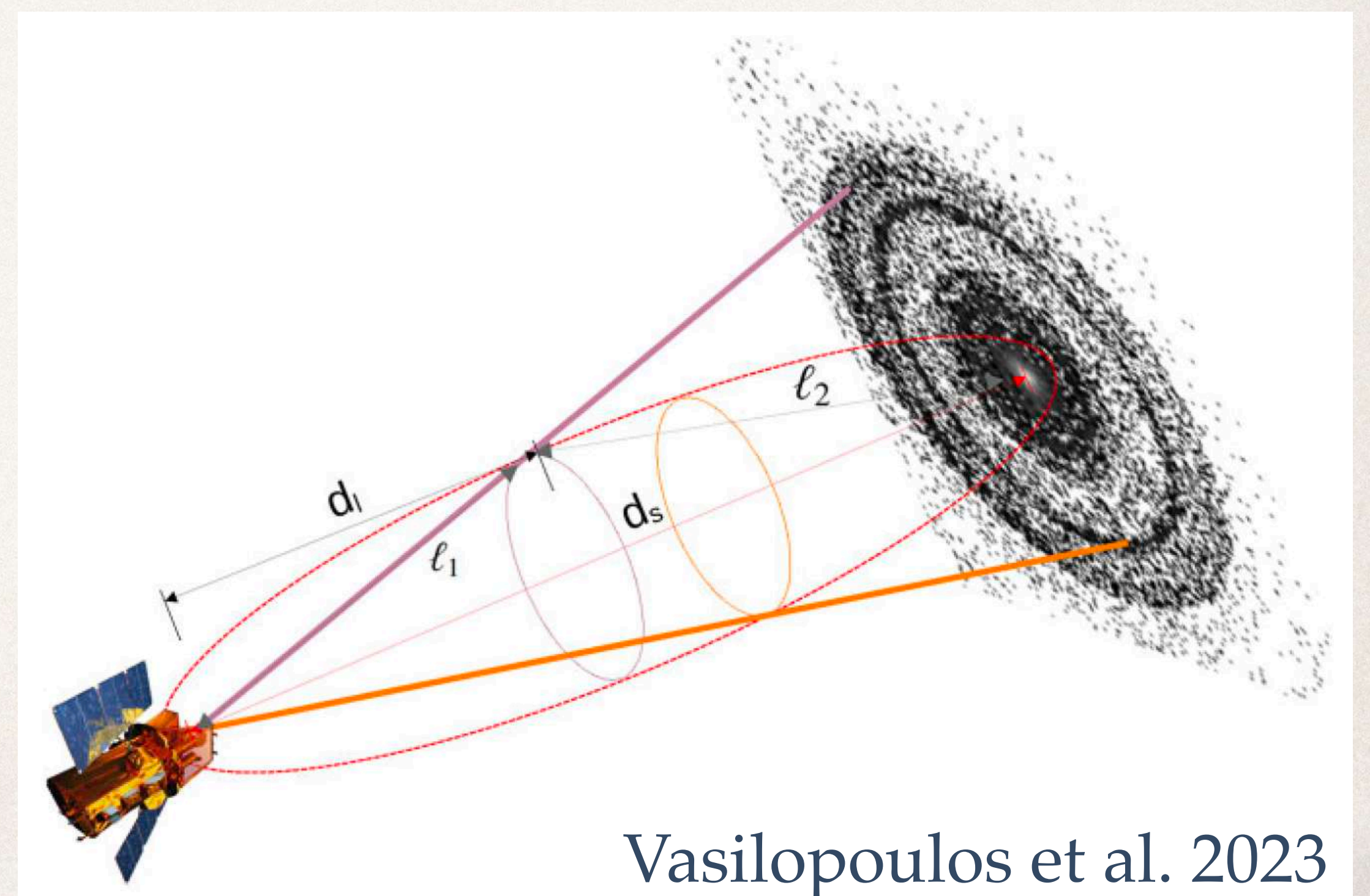
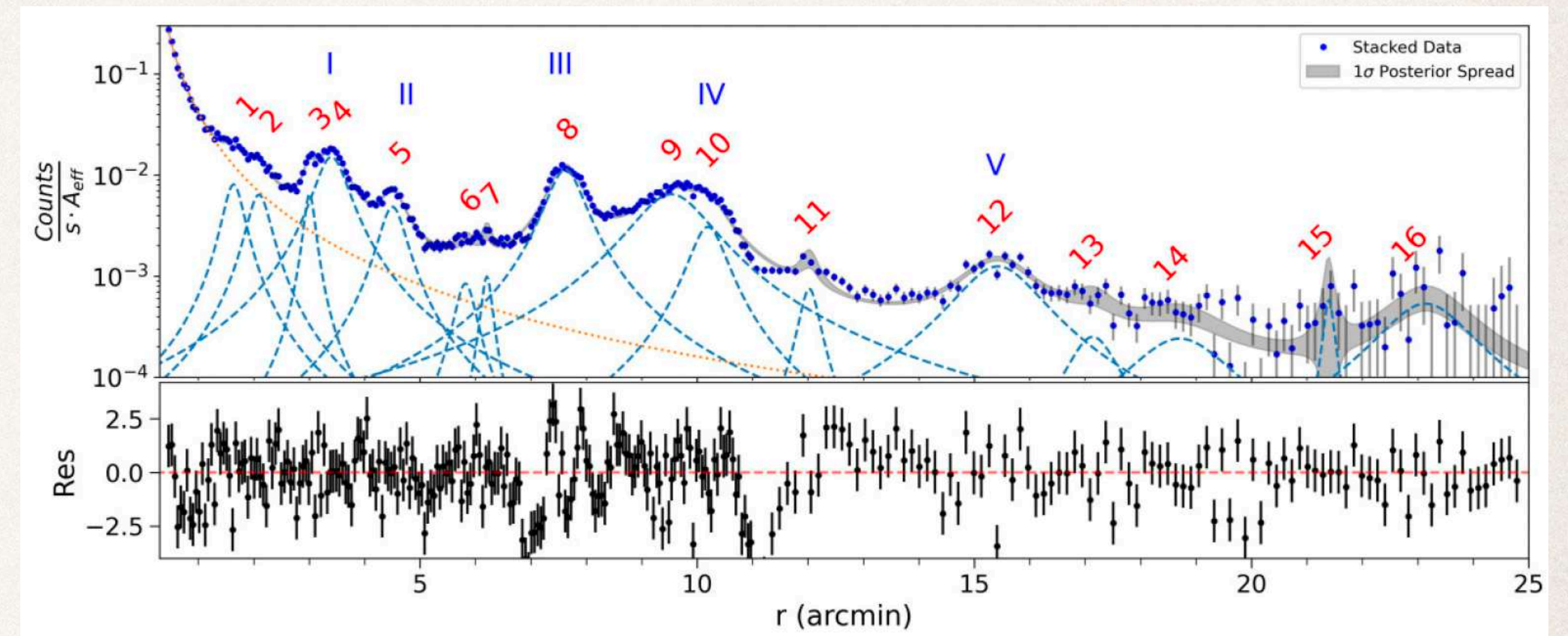
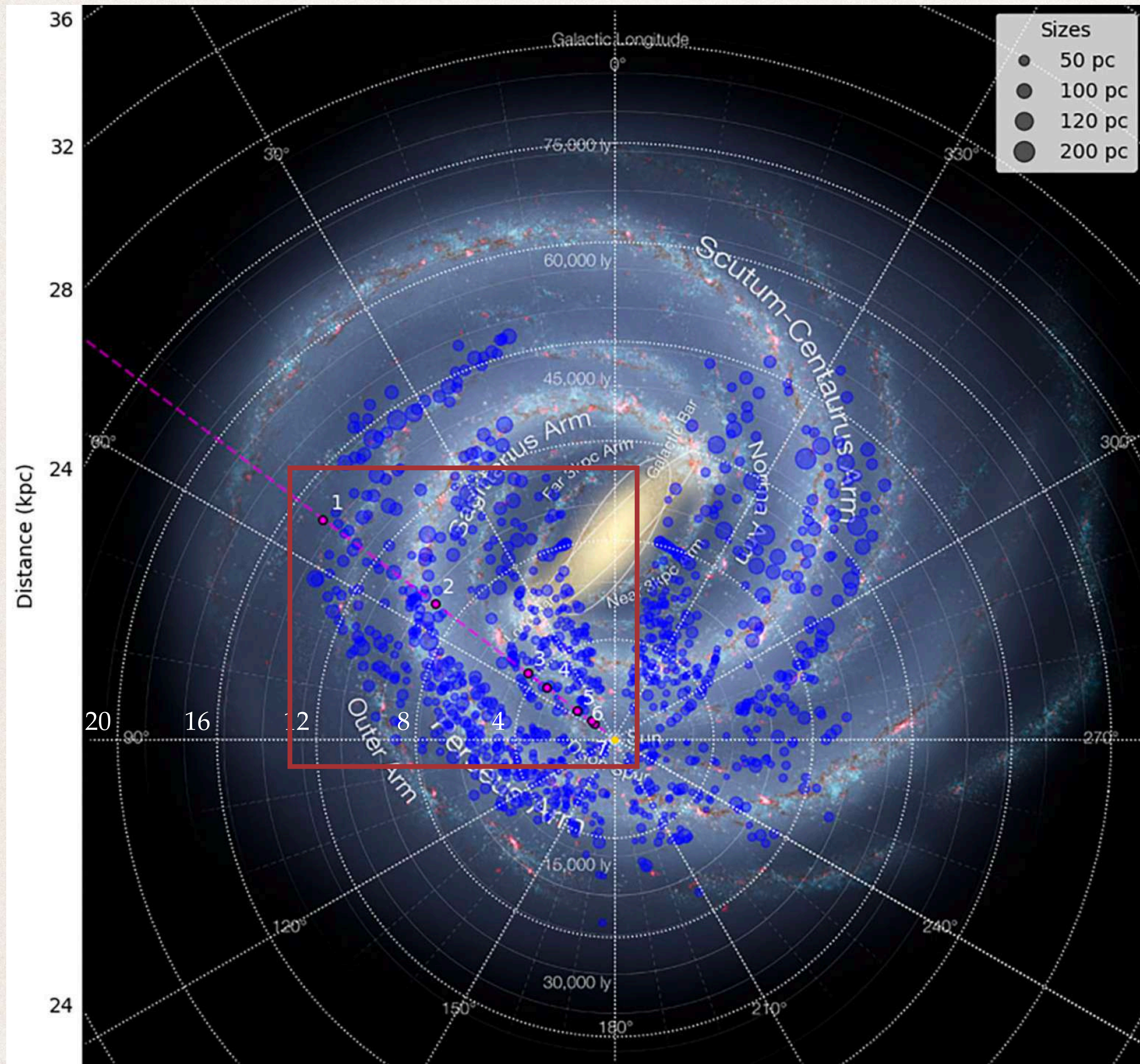
- ❖ In the Galactic plane
- ❖ High optical extinction $A_V \sim 4.1$ mag, $E(B-V) \sim 1.32$ mag

Dust scattering of X-rays



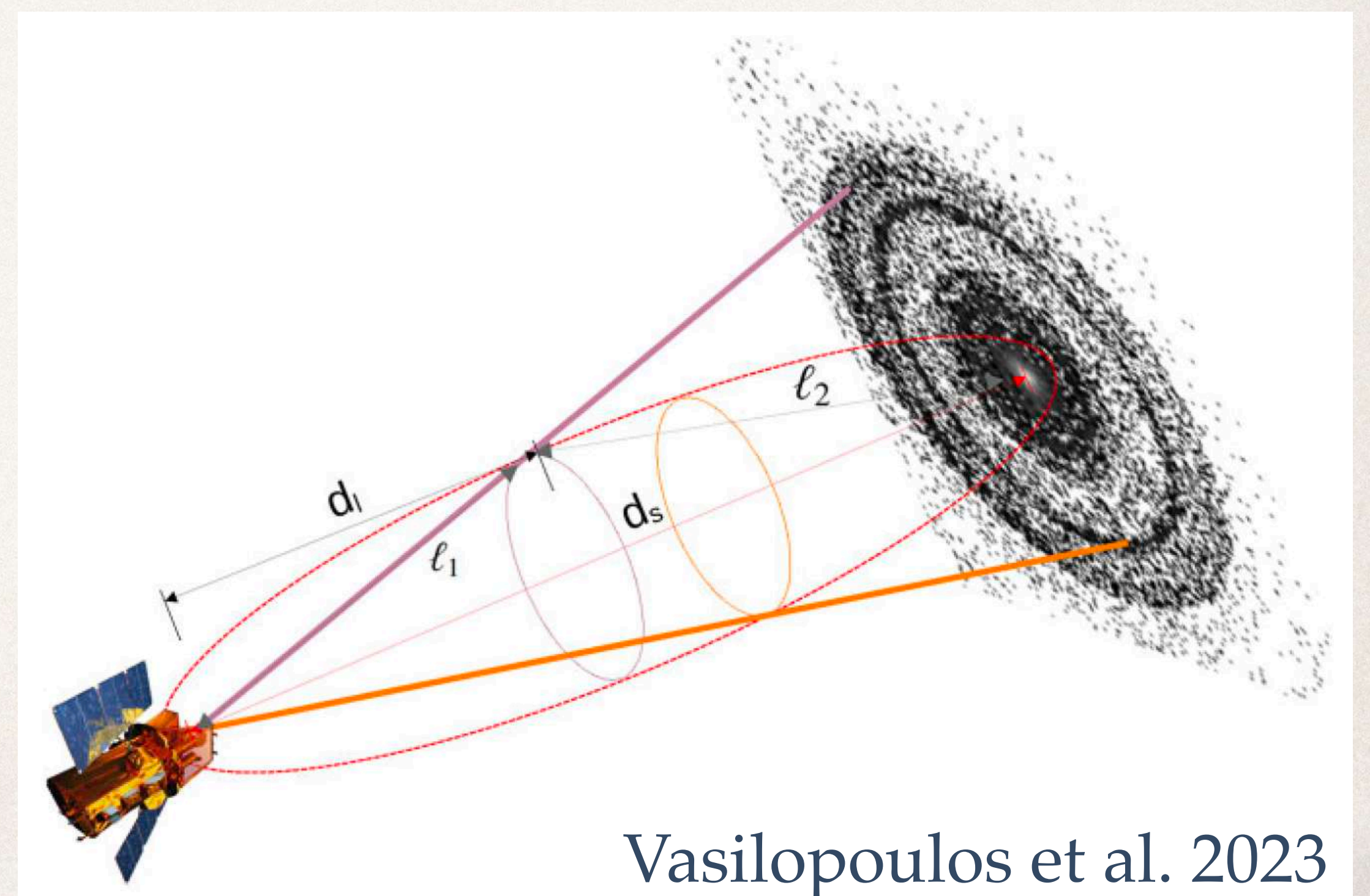
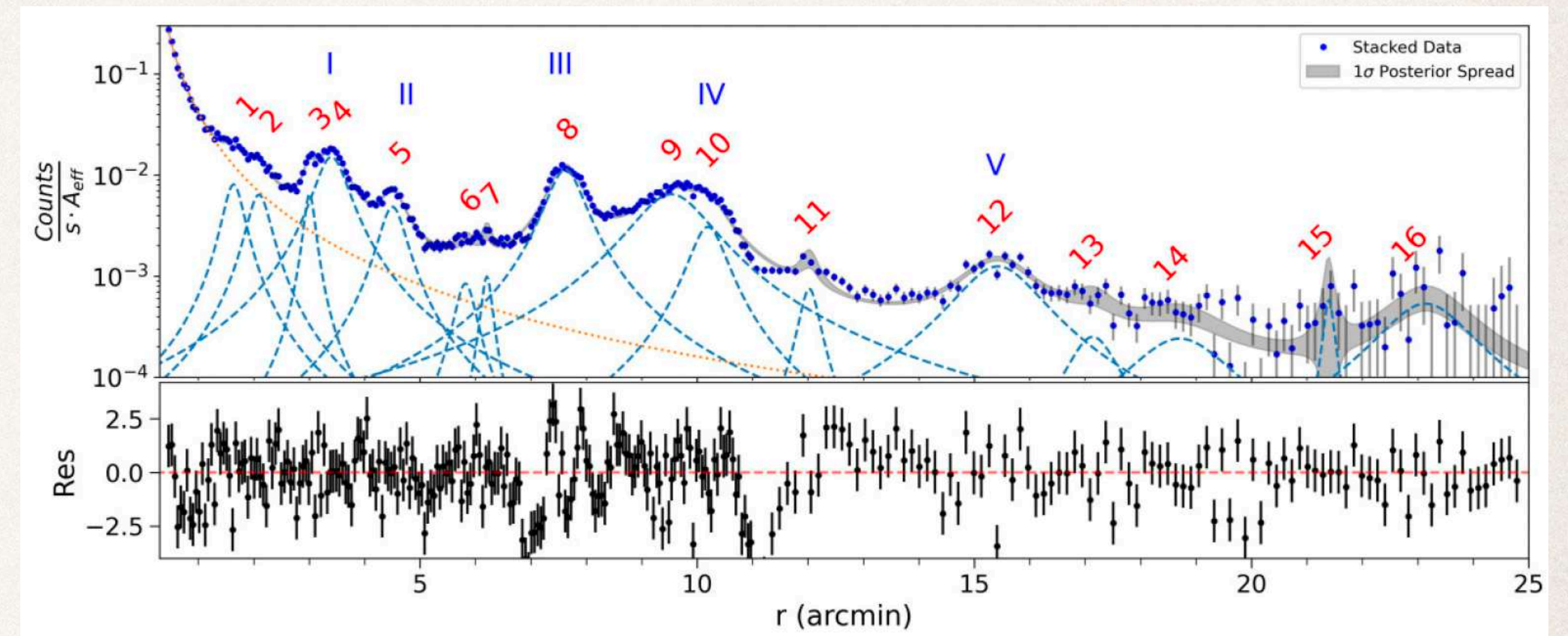
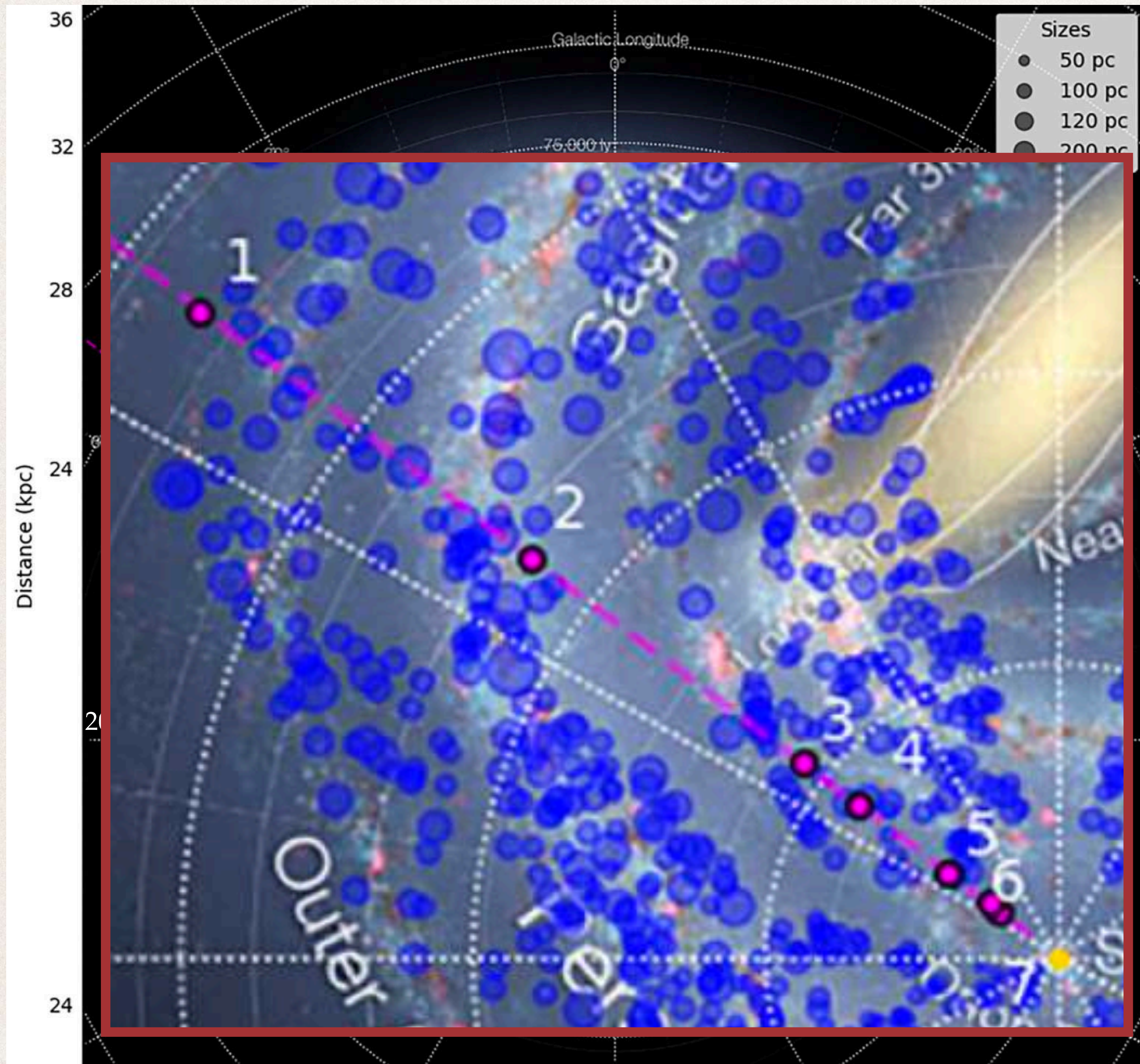
Vasilopoulos et al. 2023

Dust scattering of X-rays



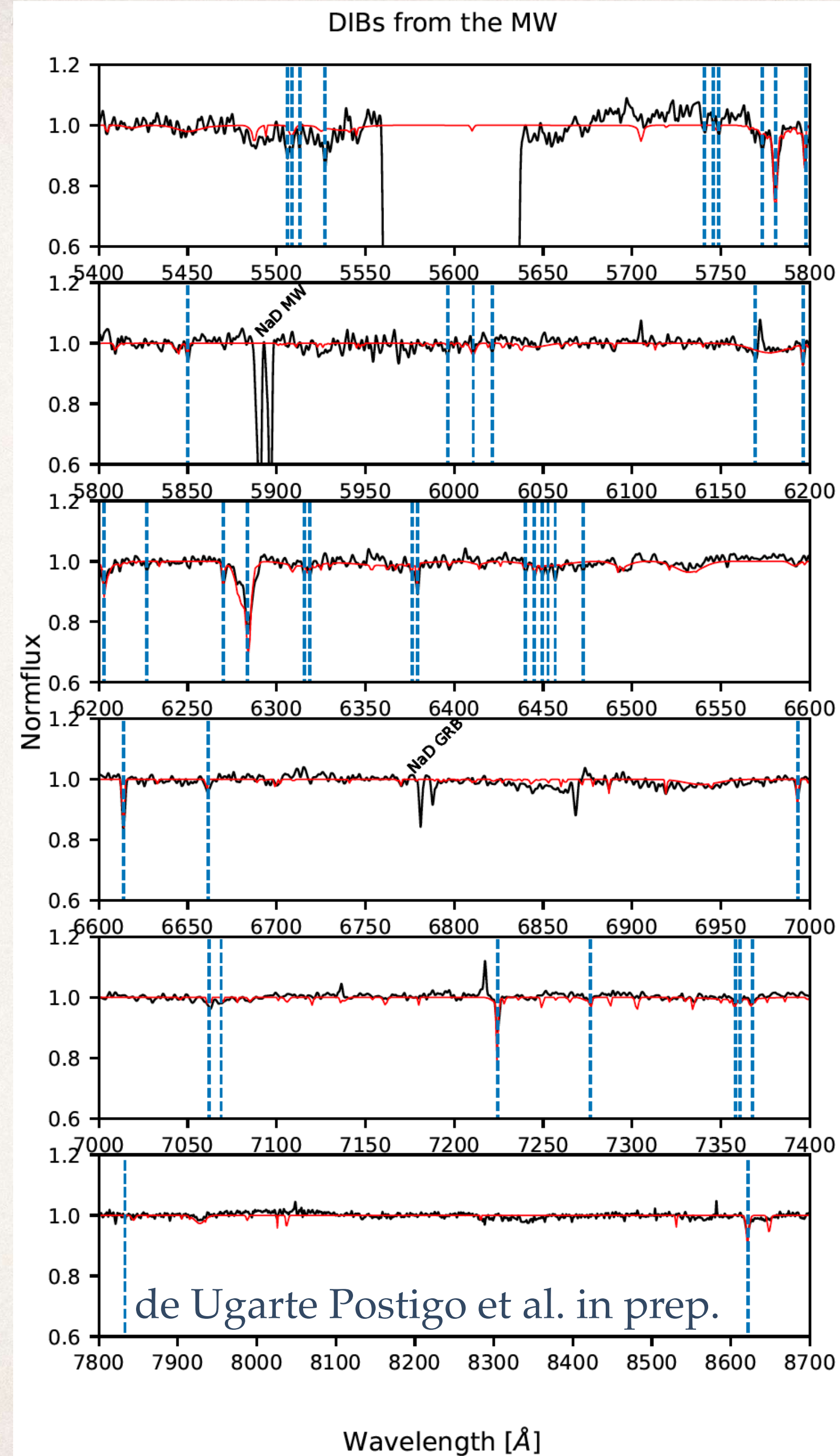
Vasilopoulos et al. 2023

Dust scattering of X-rays



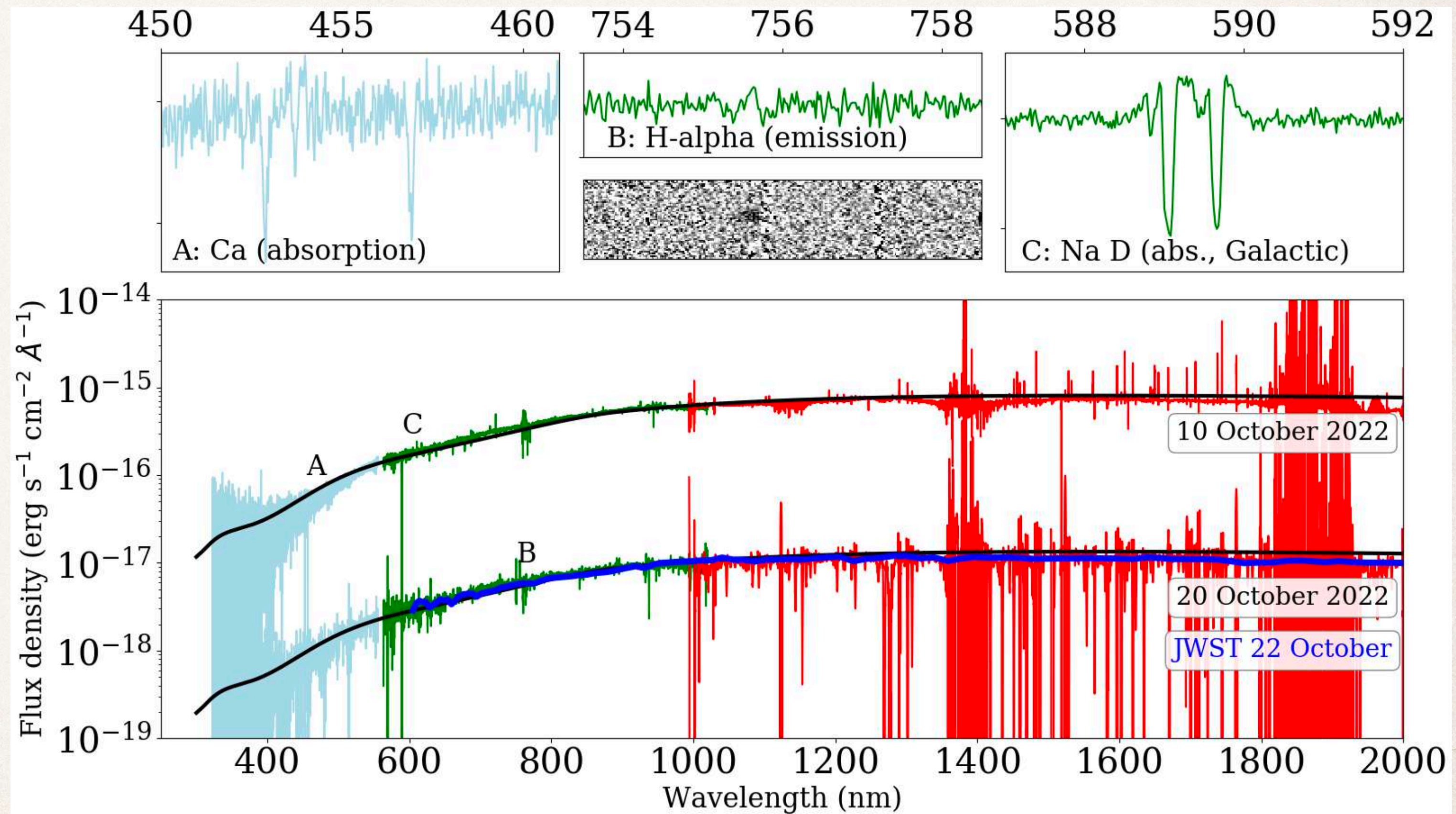
Diffuse interstellar bands

- ❖ Imprints of interstellar material in high-density Galactic sight lines
- ❖ Most carriers still unknown
- ❖ Related to extinction (Friedman et al. 2011):
 - ❖ $E(B-V) [5780.50] = 1.46 \pm 0.04$
 - ❖ $E(B-V) [6283.80] = 1.33 \pm 0.04$
 - ❖ $E(B-V) [6204.50] = 1.72 \pm 0.12$
 - ❖ $E(B-V) [6196.00] = 2.69 \pm 0.21$
 - ❖ $E(B-V) [6613.60] = 1.50 \pm 0.05$
 - ❖ $E(B-V) [5797.10] = 2.38 \pm 0.12$
- ❖ And a weighted mean of $\langle E(B-V) \rangle = 1.48 \pm 0.02$



Distance and energetics

- ❖ $z = 0.151$ (1.9 Gly) from X-shooter (de Ugarte Postigo et al. GCN 32648; Malesani et al. 2023)
- ❖ Absorption and emission features
- ❖ Very energetic in gamma-rays
 $E_{\text{iso}} \approx 1\text{-}1.5 \cdot 10^{55}$ erg
(Kann & Agüí-Fernández 2022; Frederiks et al. 2023; S. Lesage et al. 2023)
- ❖ Follow-up with VLT, NOEMA, GRANDMA, GTC, HST, JWST, ...



Once in a life time?

- ❖ Atteia GCN 32793:
 - ❖ one in 130-520 years 2023 (energetics)
- ❖ Malesani et al. 2023:
 - ❖ One in > 68 years (count rate)
 - ❖ One in > 112 years (fluence)
 - ❖ Up to one in a millennia
- ❖ Burns et al. 2023:
 - ❖ One in 10 000 years (energetics)
- ❖ Kann et al. 2023:

It is unlikely that a chance like this will come again in the coming decades or even centuries, making this an event to be remembered through the ages.

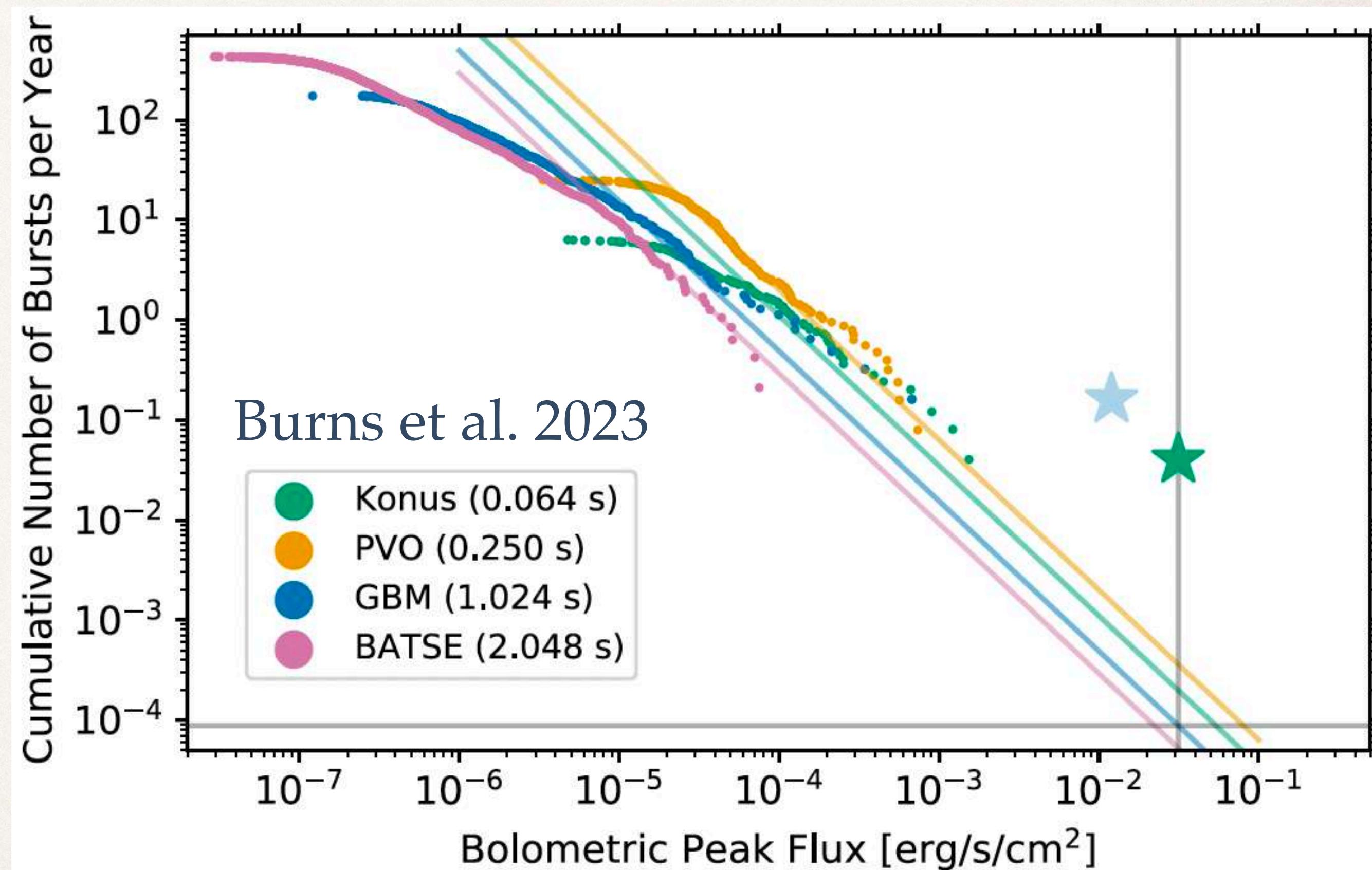


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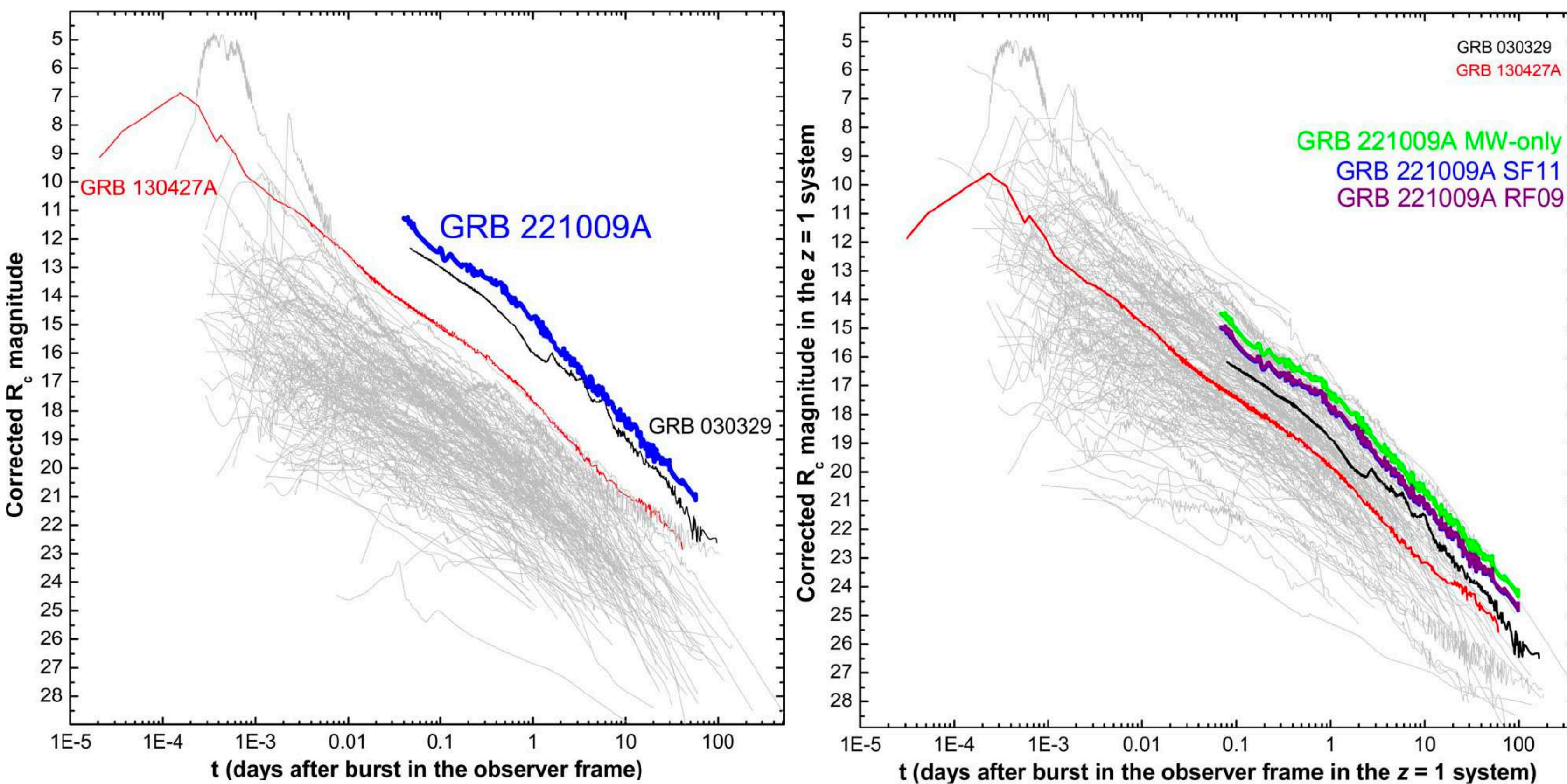
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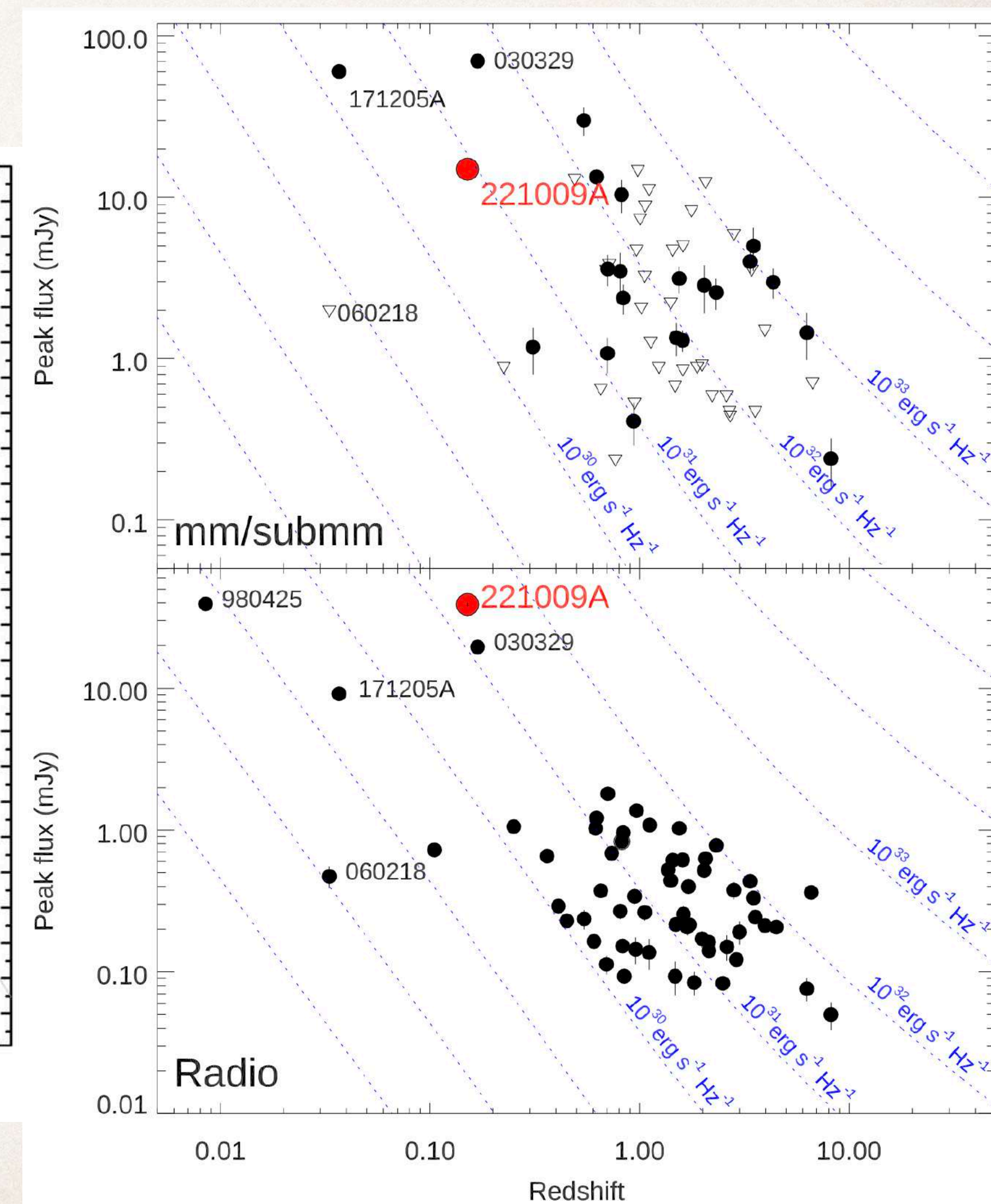
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Afterglow luminosity



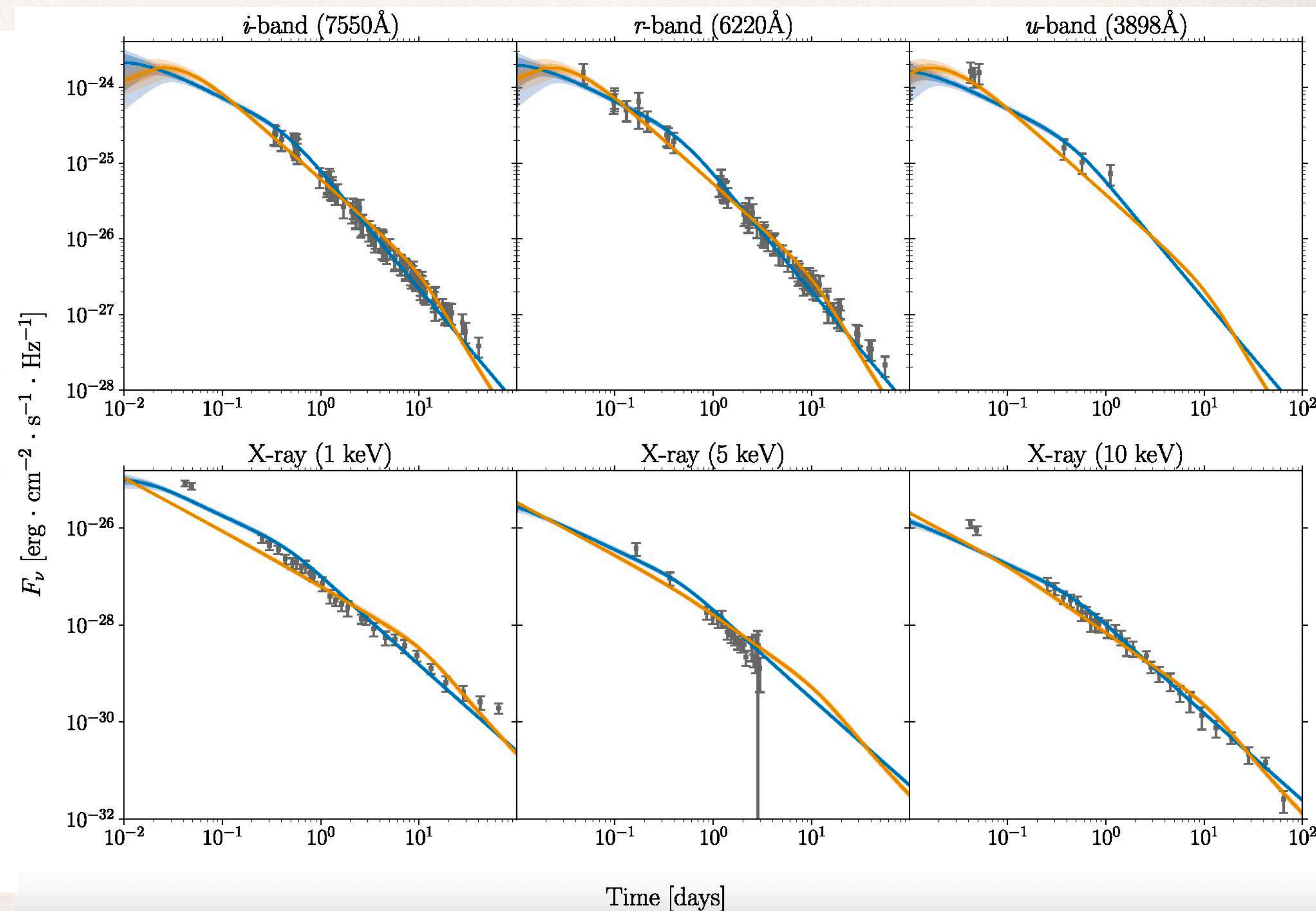
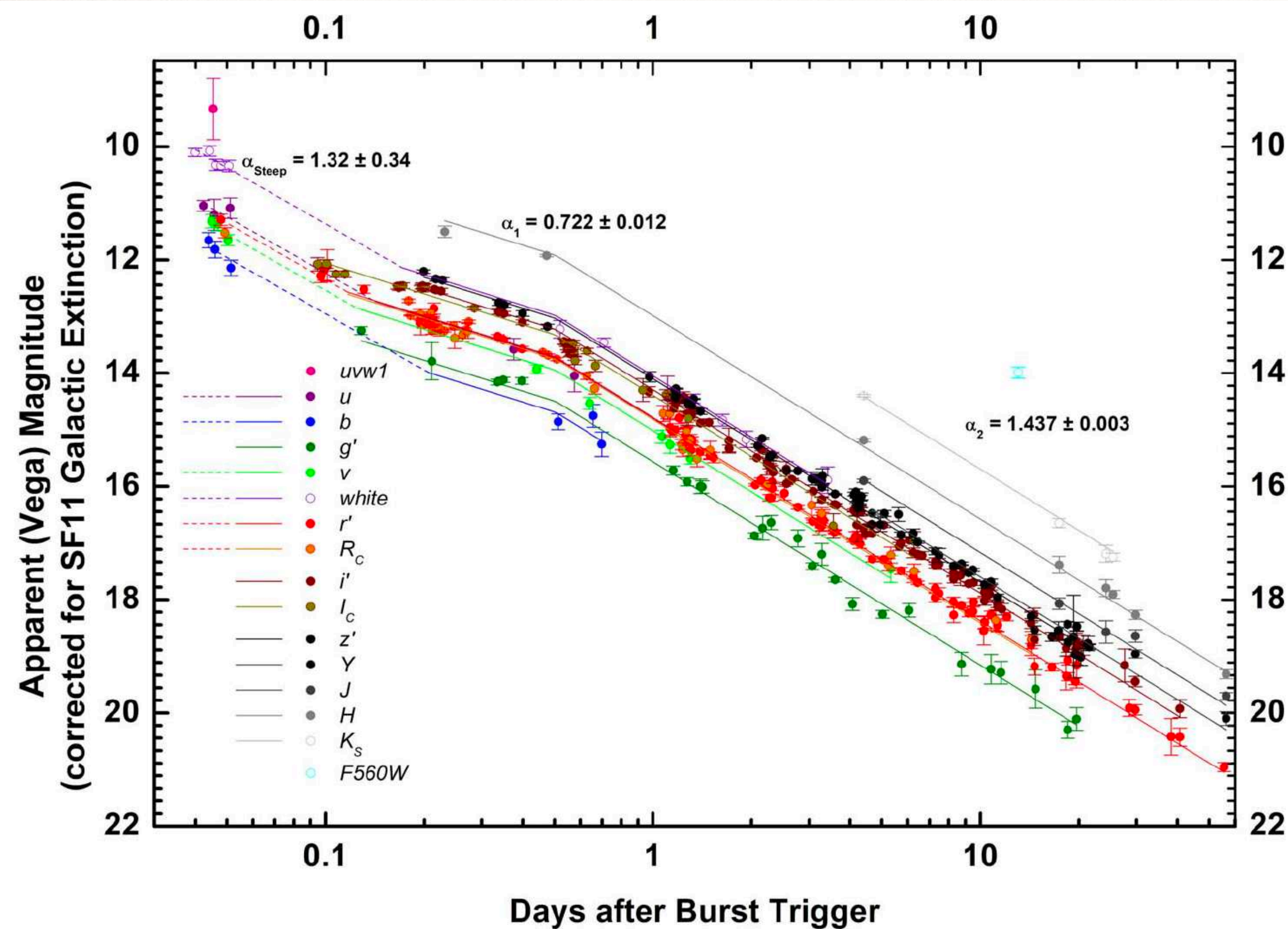
Kann et al. 2023



de Ugarte Postigo et al. in prep.

GRANDMA follow-up

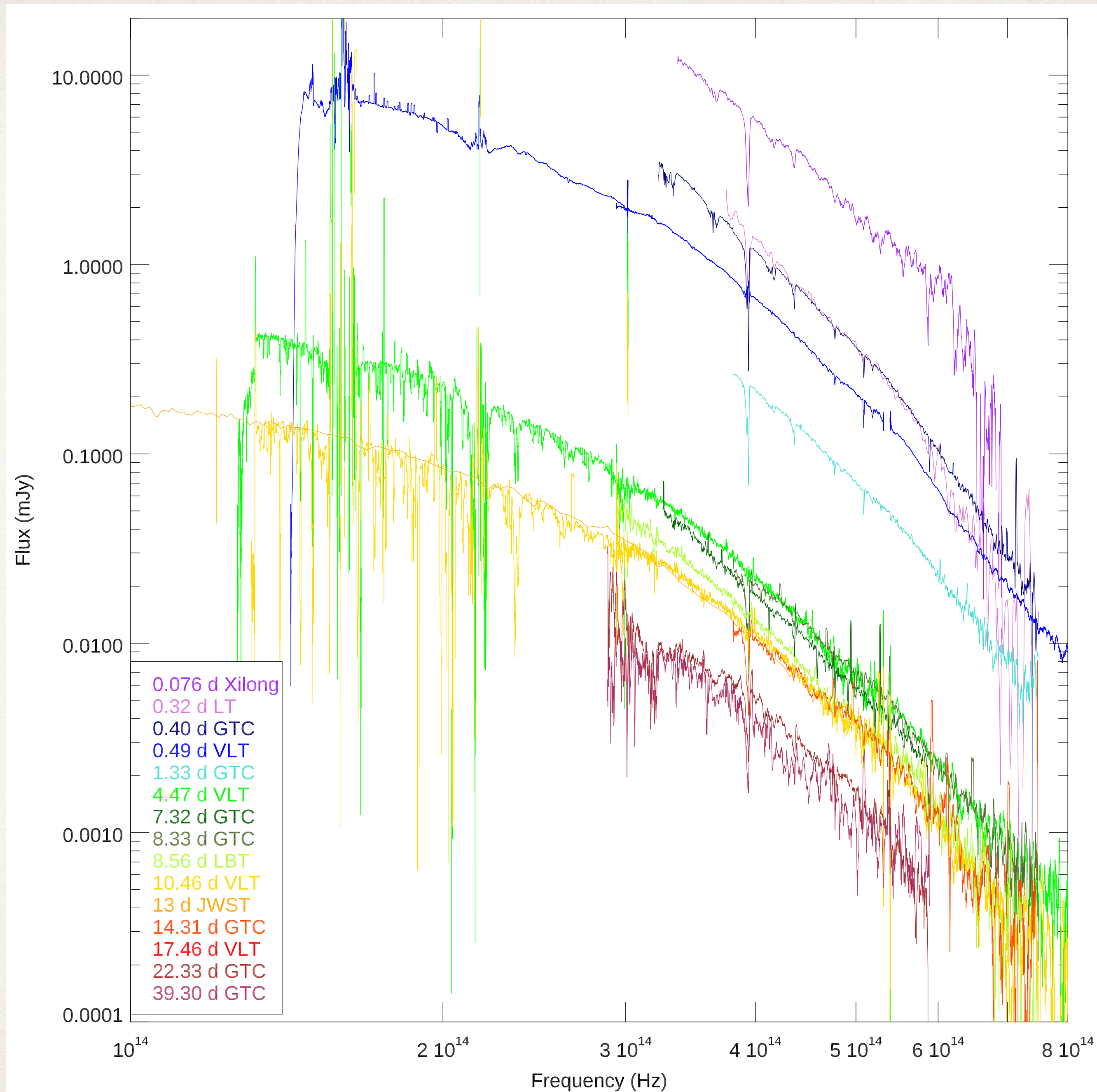
- ❖ Over 300 observations from:
 - ❖ 15 GRANDMA telescopes
 - ❖ 19 amateur telescopes



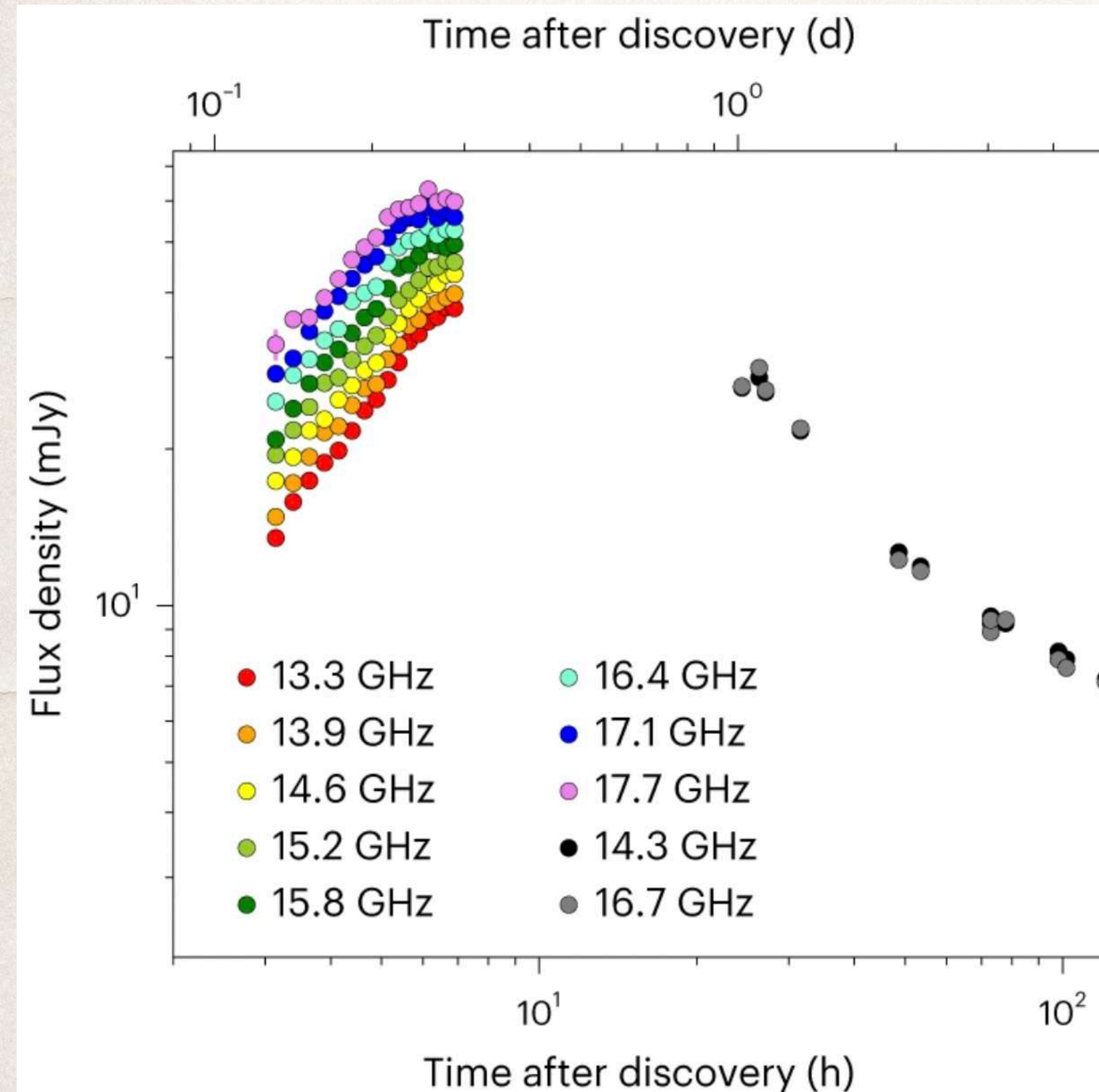
Spectral evolution

- ❖ 15 spectral epochs from GTC, VLT, JWST, LT and Xilong
- ❖ From 1.8 hrs to 39 days
- ❖ Extreme extinction complicates disentangling afterglow and supernova component

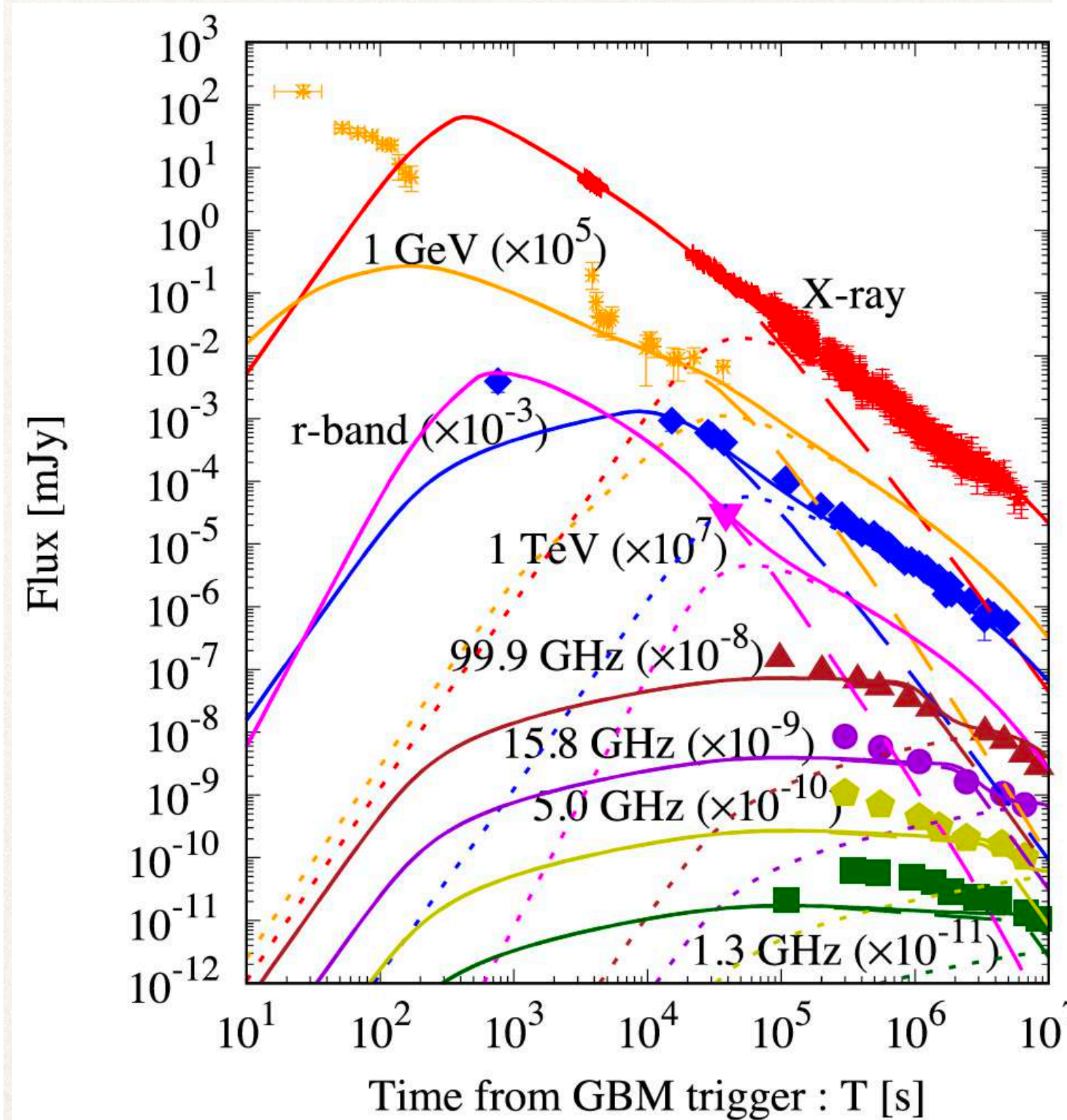
de Ugarte Postigo et al. in prep.



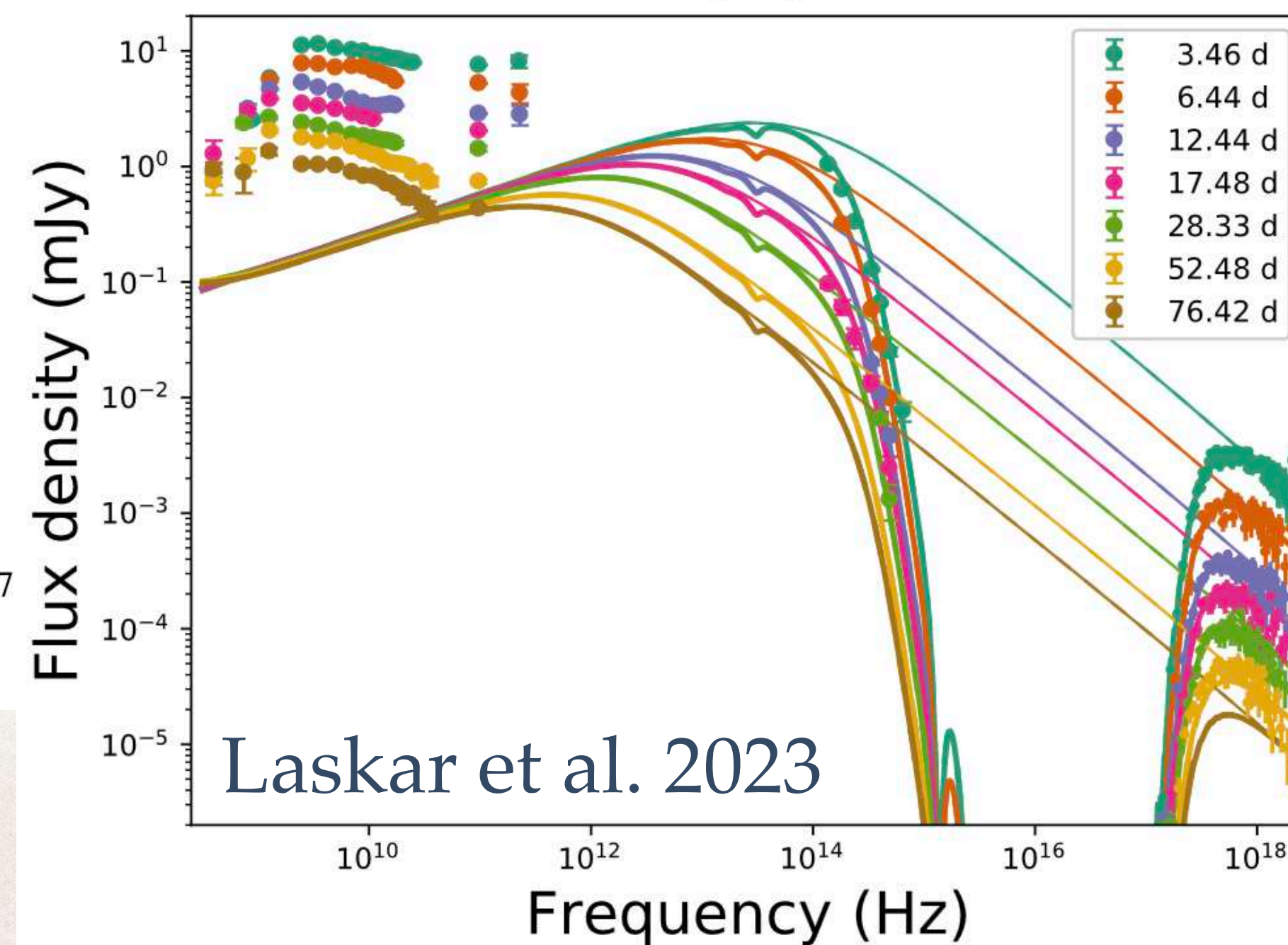
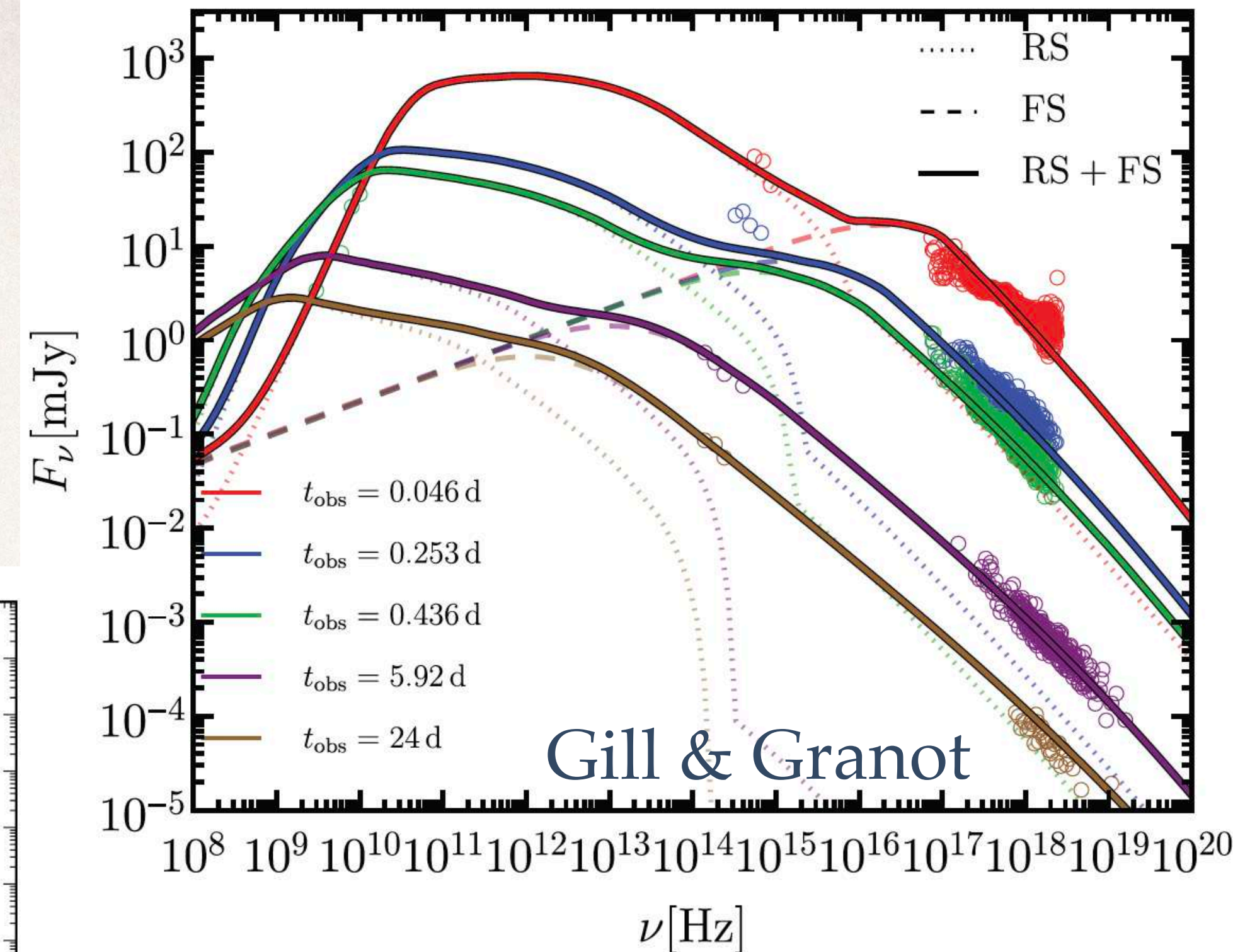
Broad-band emission



Bright et al. 2023

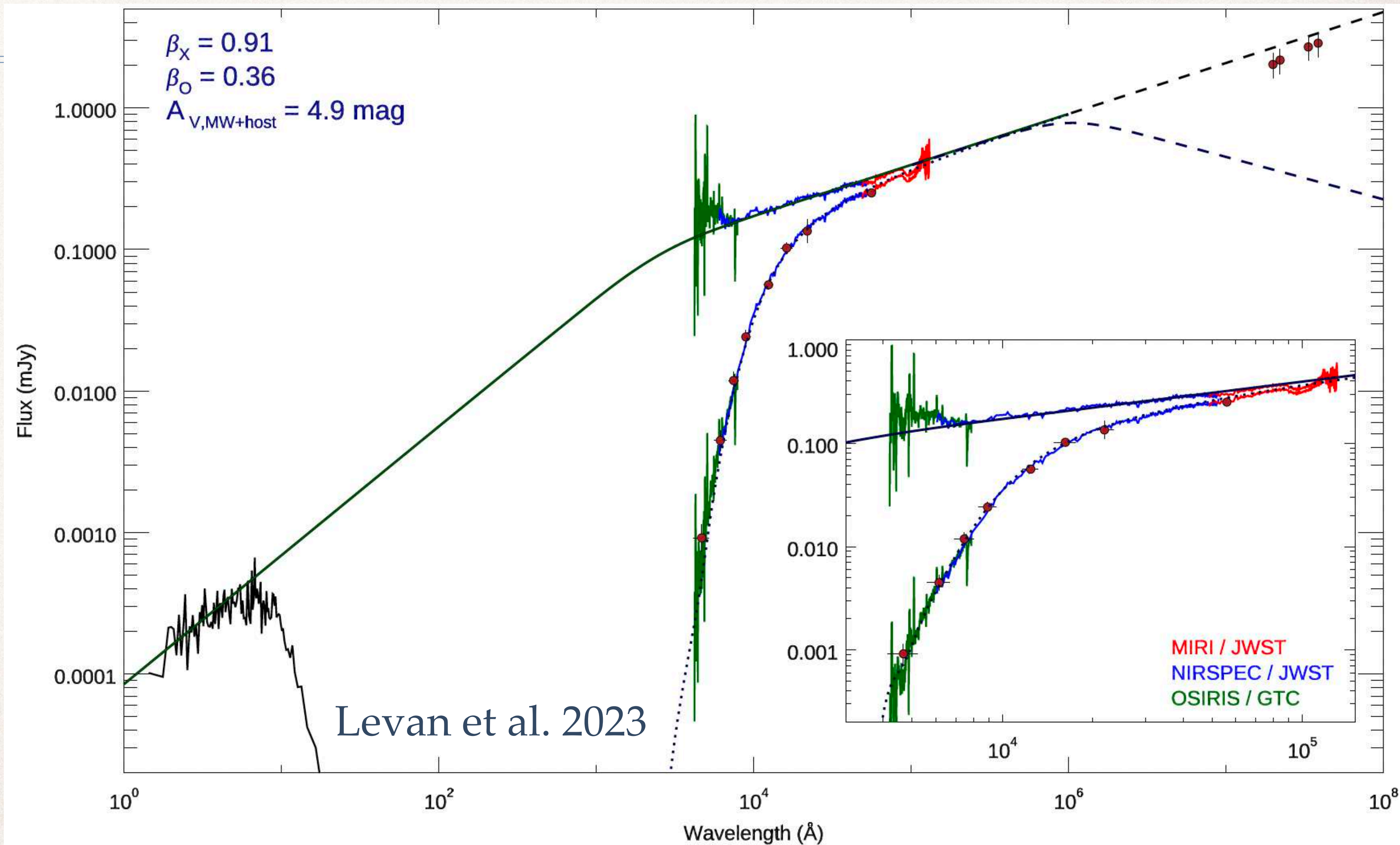
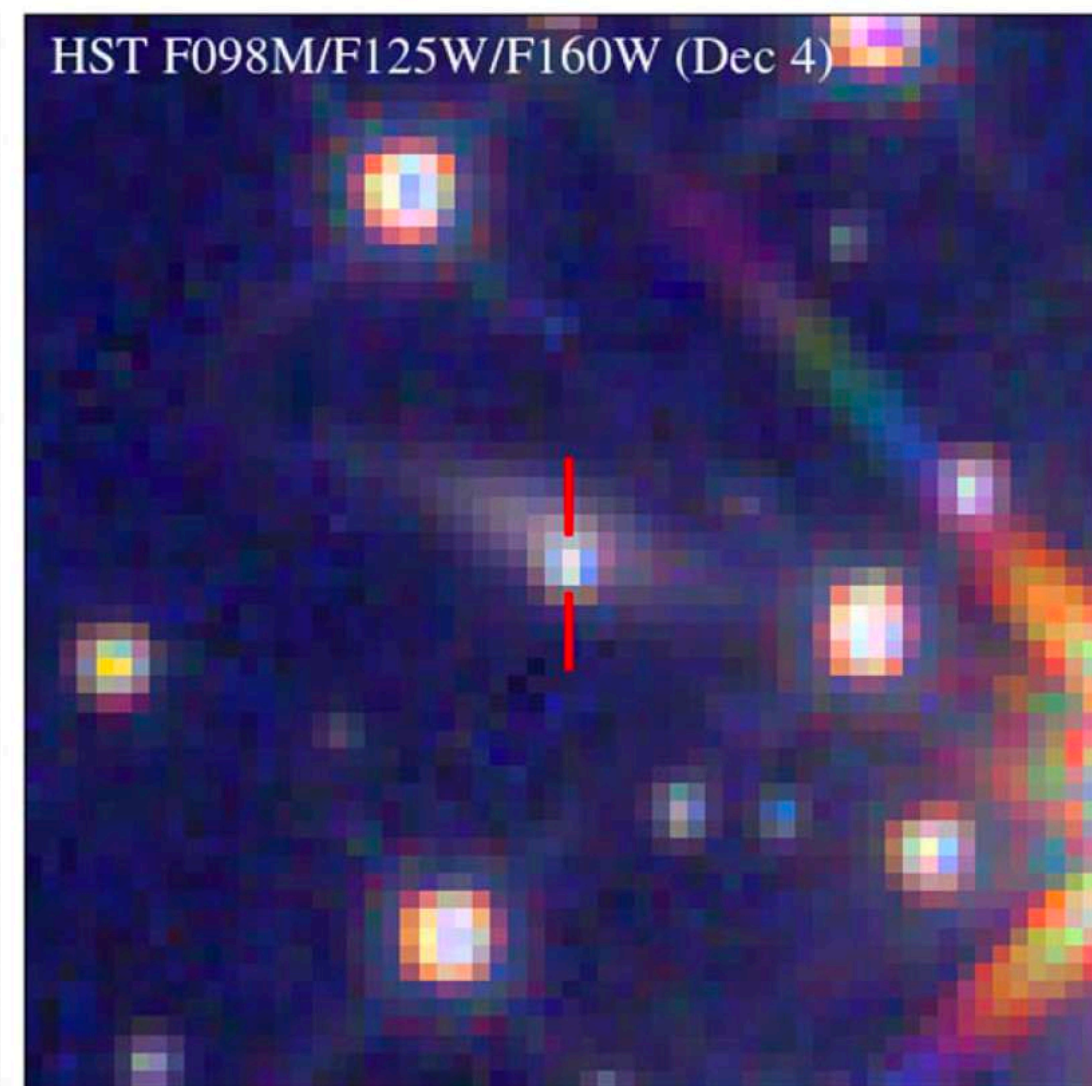
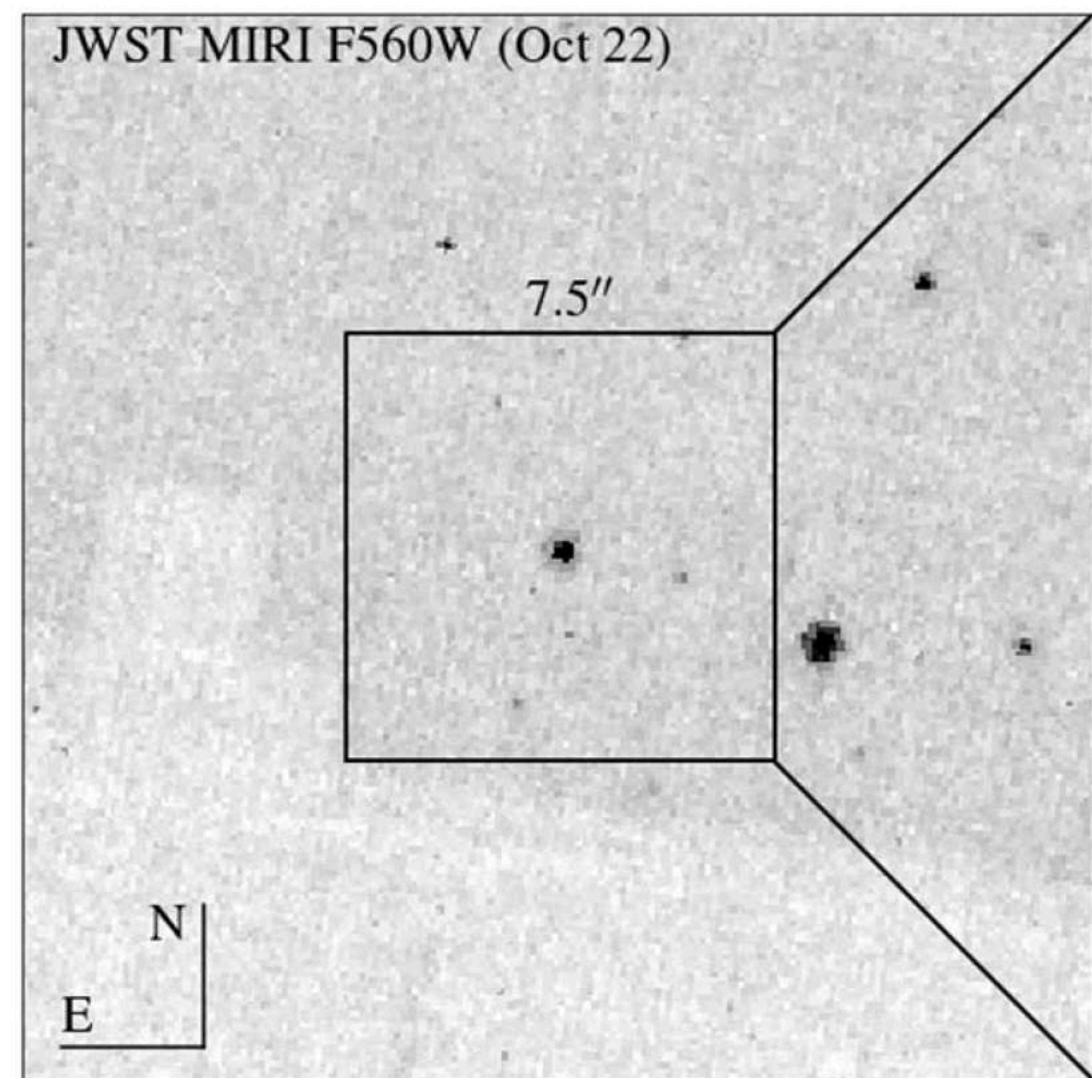


Sato et al. 2023

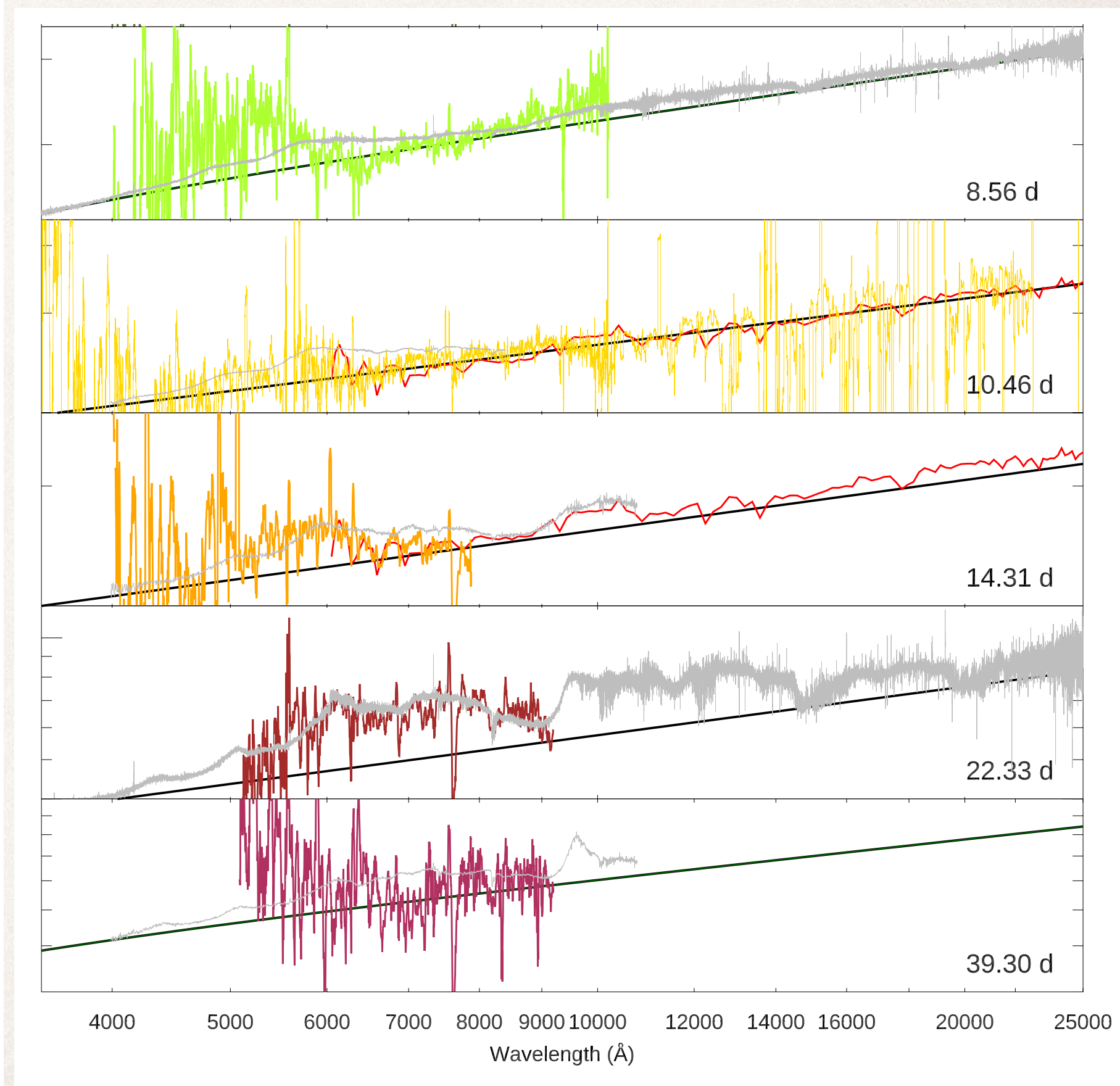
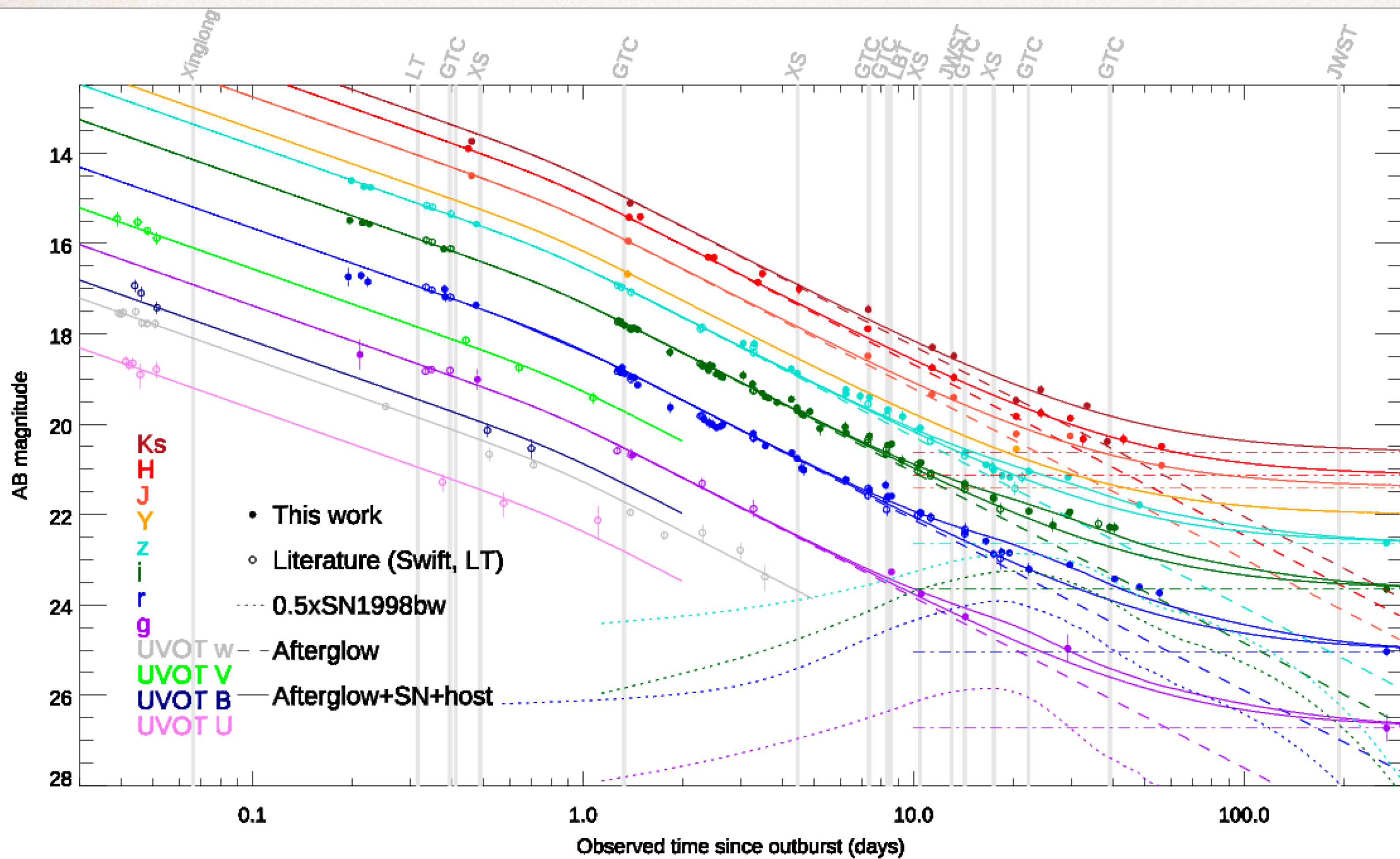


JWST & HST

- ❖ First JWST observation of a GRB
- ❖ Hint of a faint SN

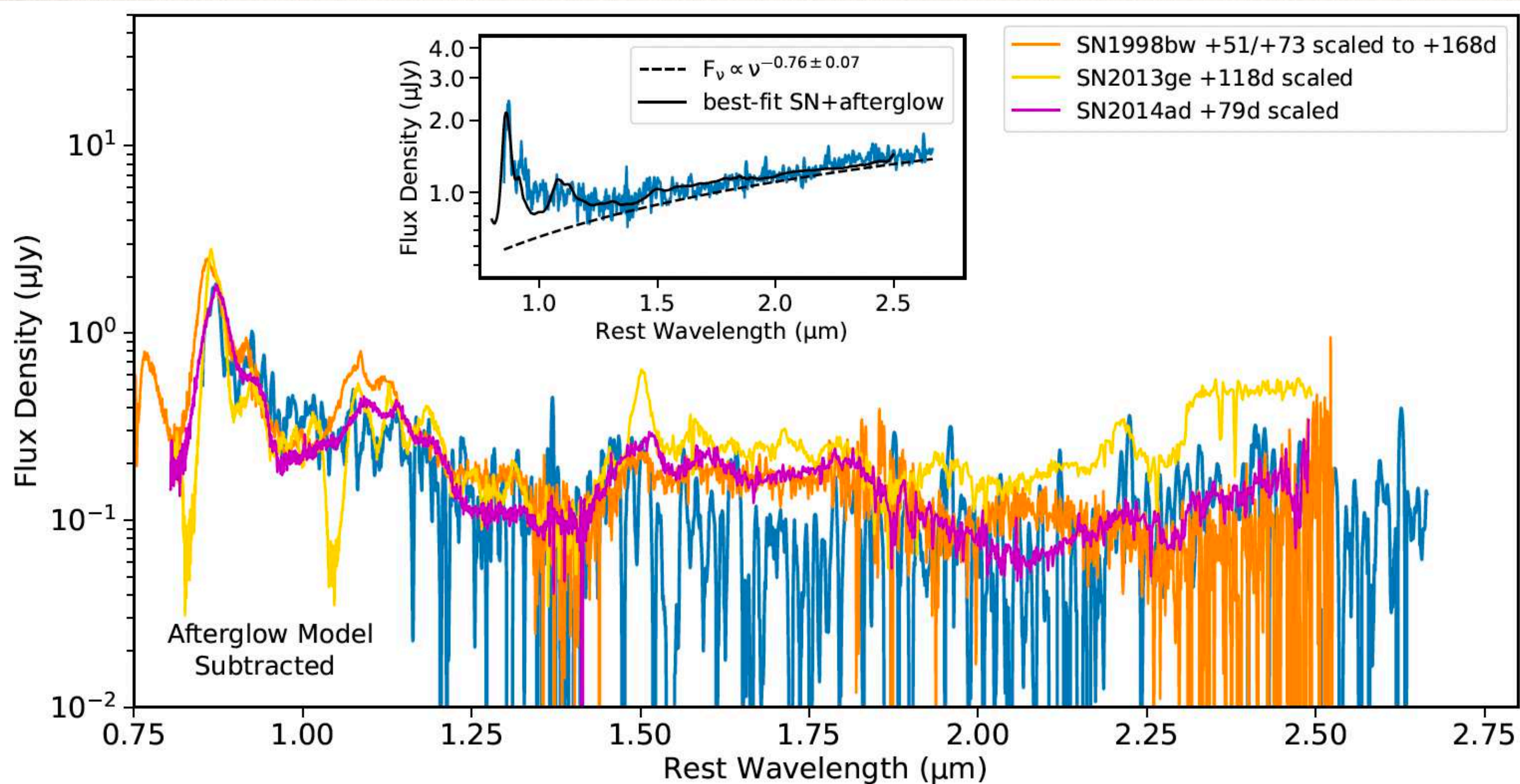


Supernova component

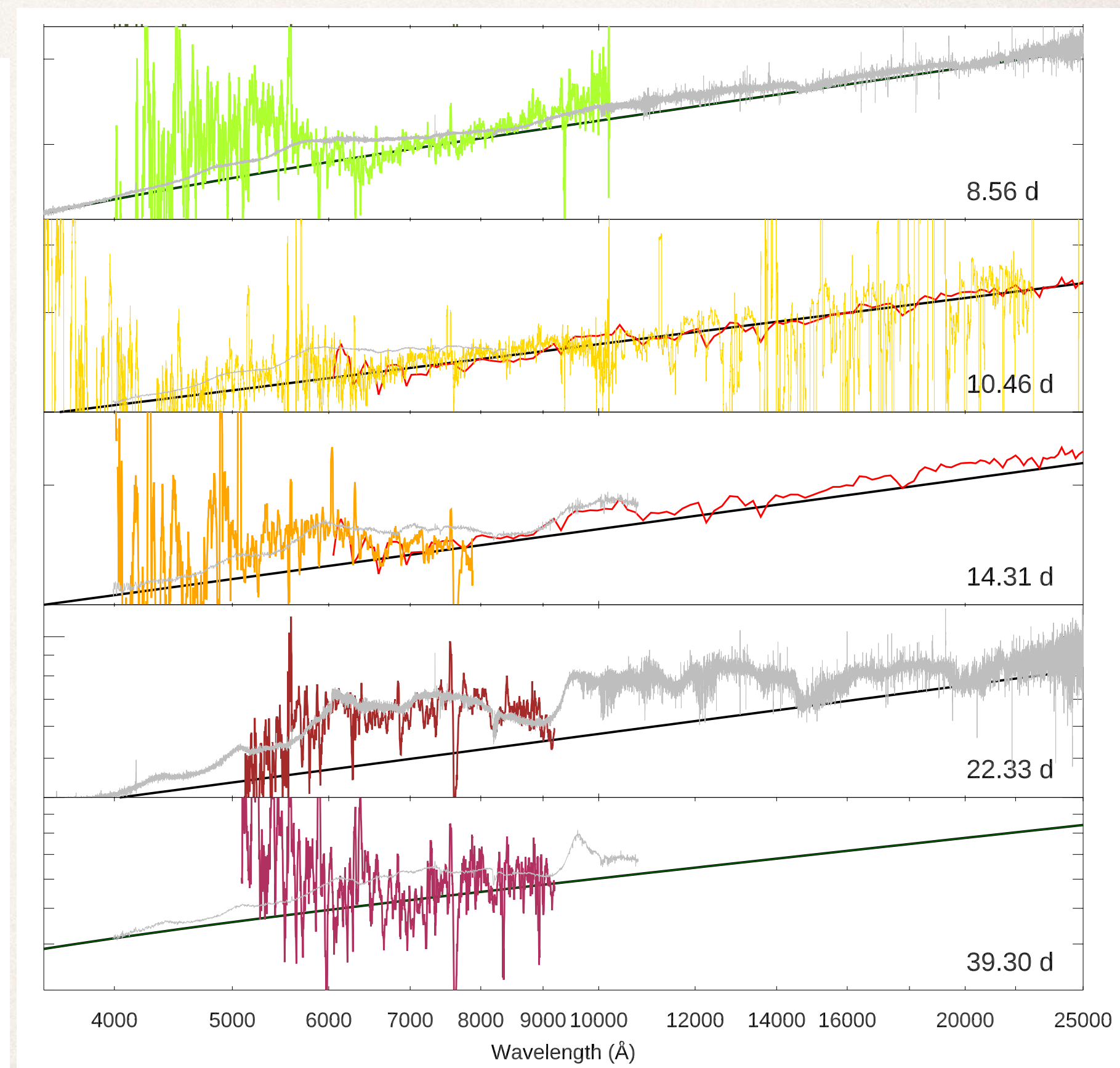


de Ugarte Postigo et al. in prep.

Supernova component



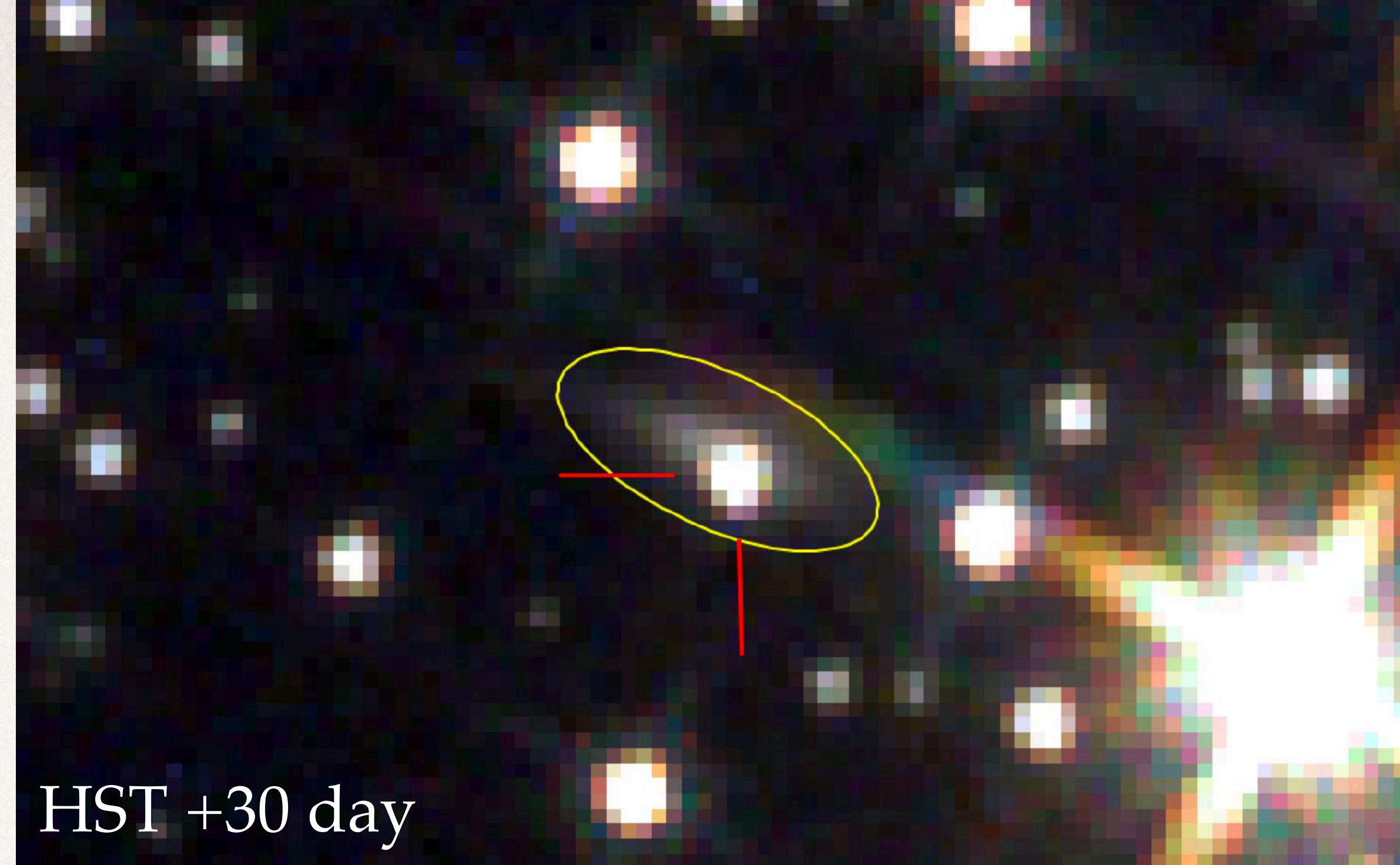
Blanchard et al. 2023



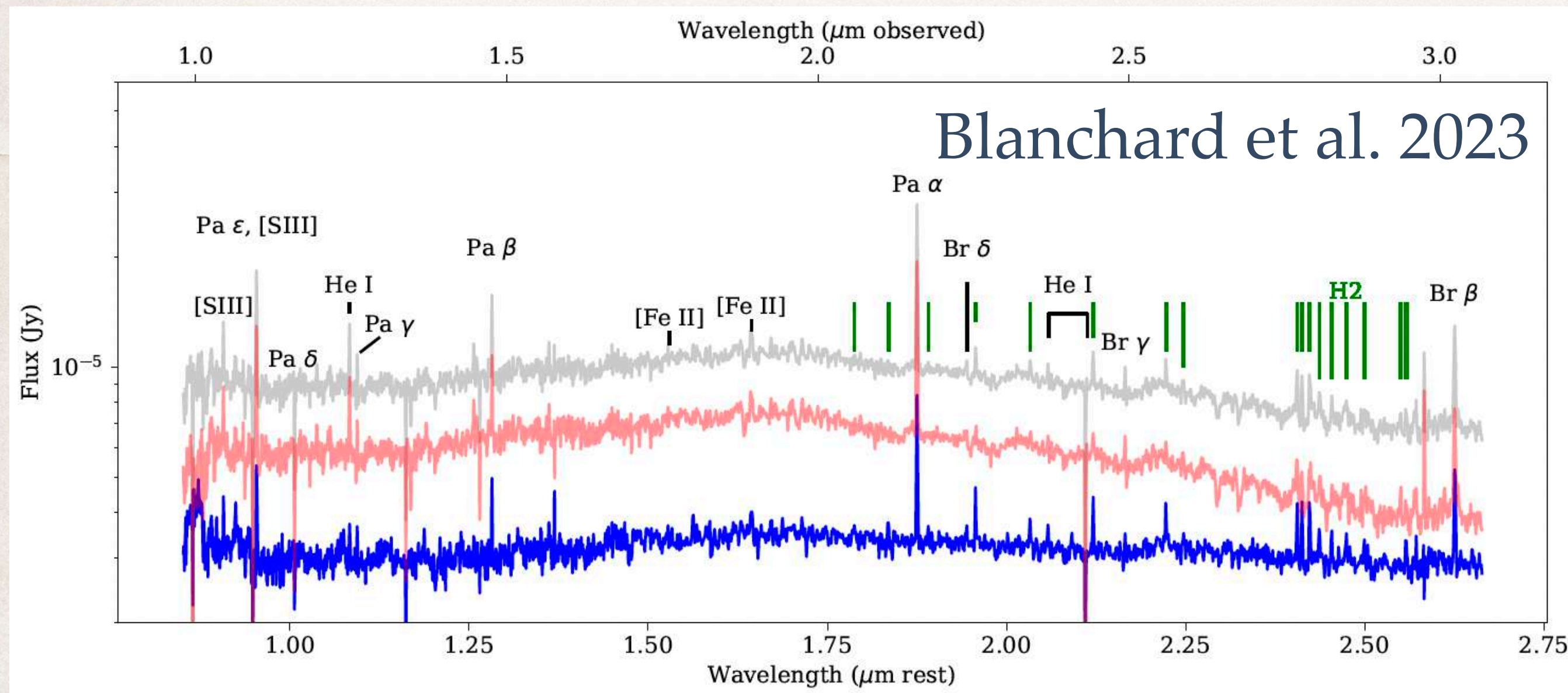
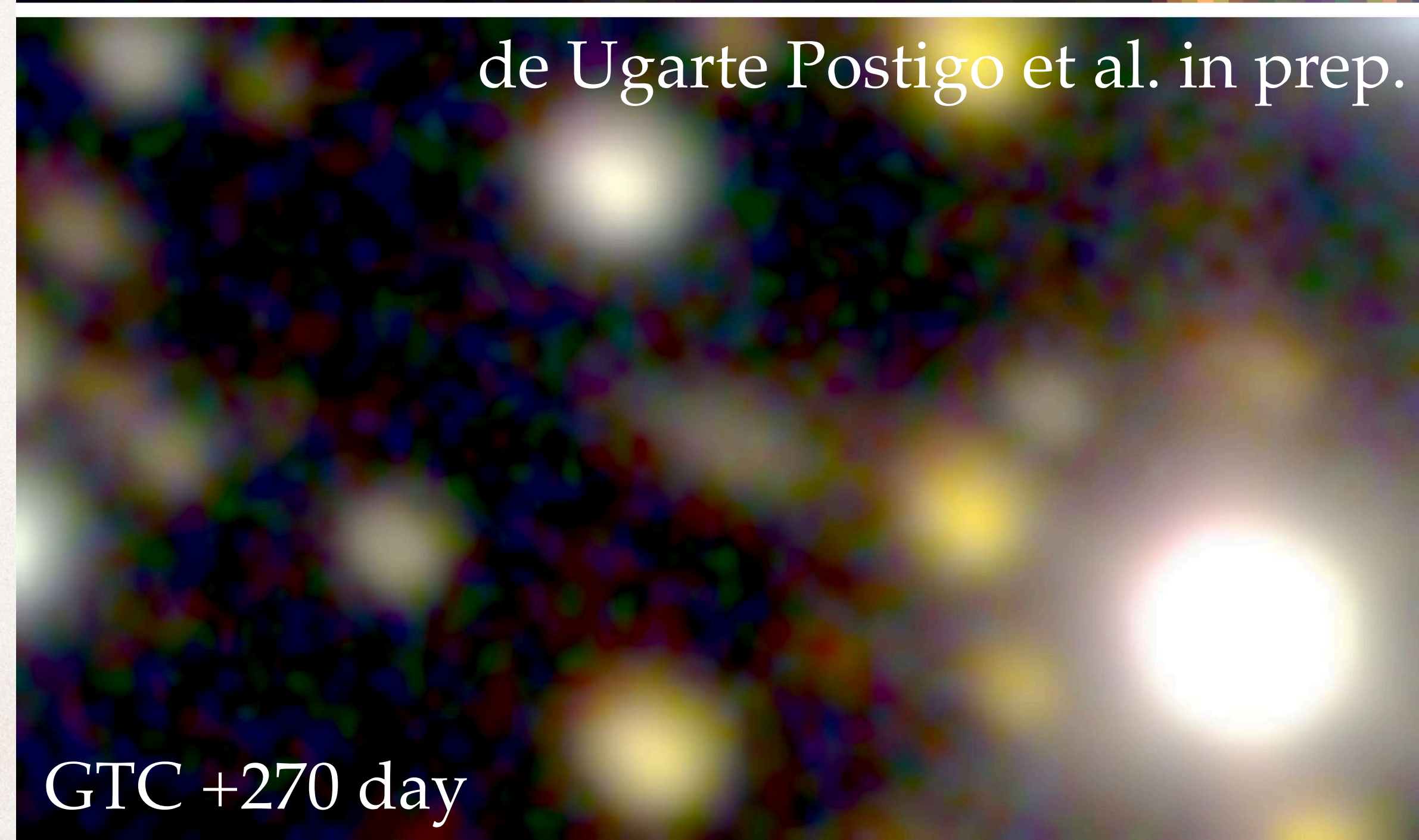
de Ugarte Postigo et al. in prep.

Host galaxy

- ❖ Near edge-on galaxy
- ❖ Star forming $SFR = 0.17 M_{\odot}/\text{yr}$
- ❖ Low metallicity $\log Z/Z_{\odot} = -0.9$
- ❖ The GRB location shows H_2 emission



de Ugarte Postigo et al. in prep.





In memory of
Alex

Conclusions



- ❖ The Brightest Of All Time
- ❖ Probably the most energetic in gamma-rays ($E_{\text{iso}} \sim 10^{55}$ erg)
- ❖ Nearby, $z = 0.151$
- ❖ Luminous but not extreme optical afterglow
- ❖ Hidden behind strong Galactic extinction
- ❖ Associated with a supernova
- ❖ In a low-metallicity galaxy