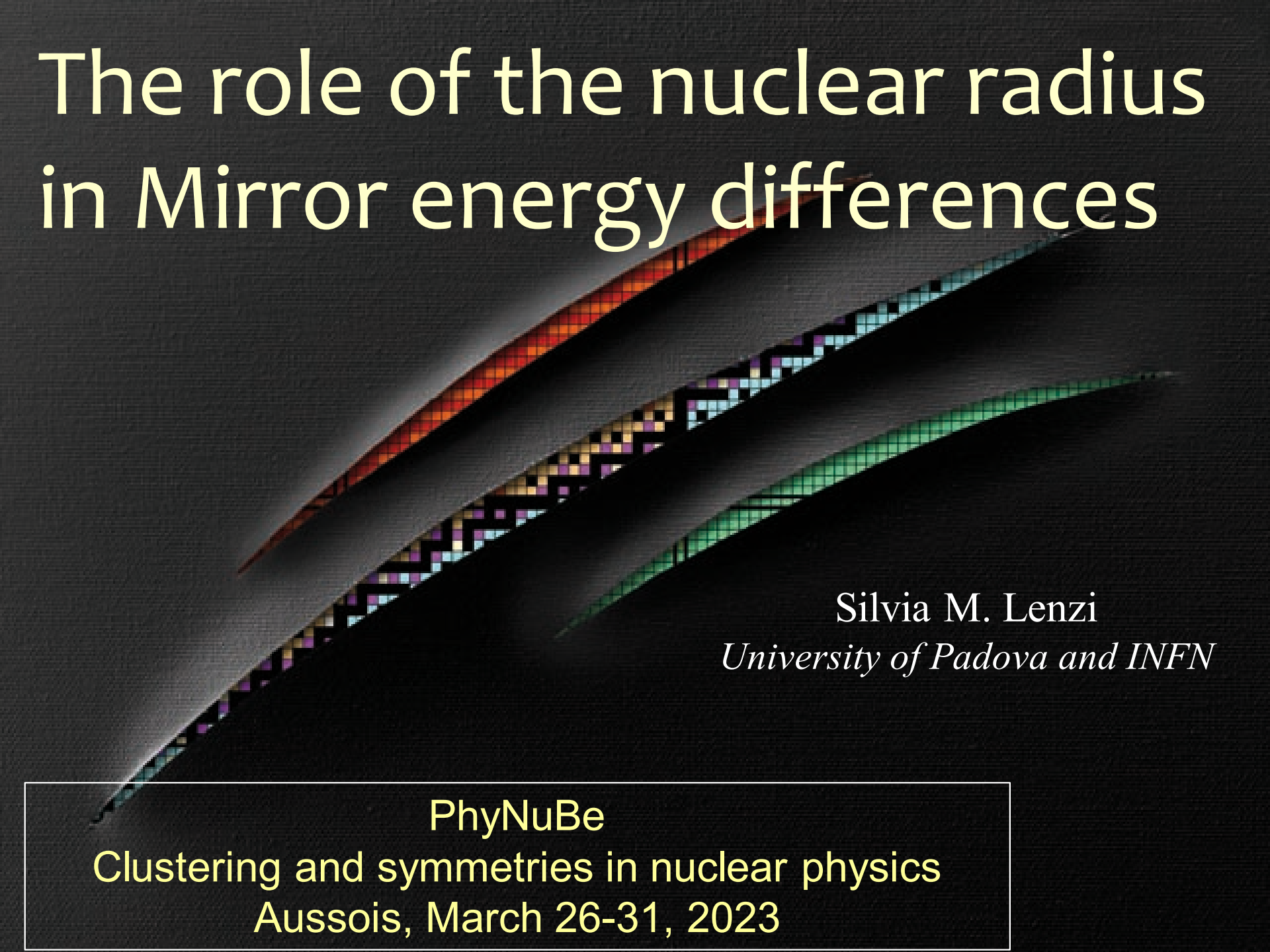


# The role of the nuclear radius in Mirror energy differences



Silvia M. Lenzi

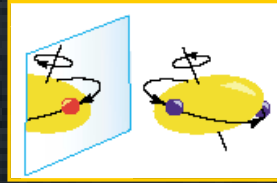
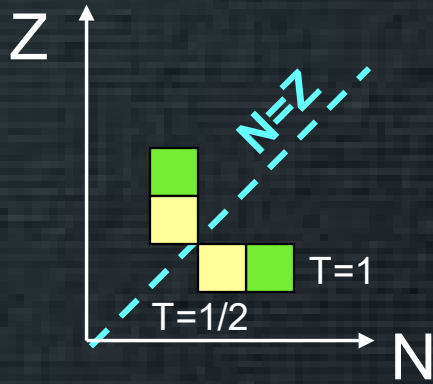
*University of Padova and INFN*

PhyNuBe

Clustering and symmetries in nuclear physics

Aussois, March 26-31, 2023

# Mirror energy differences

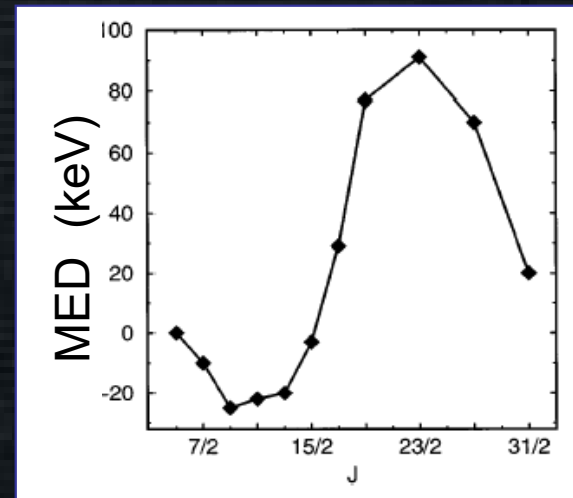
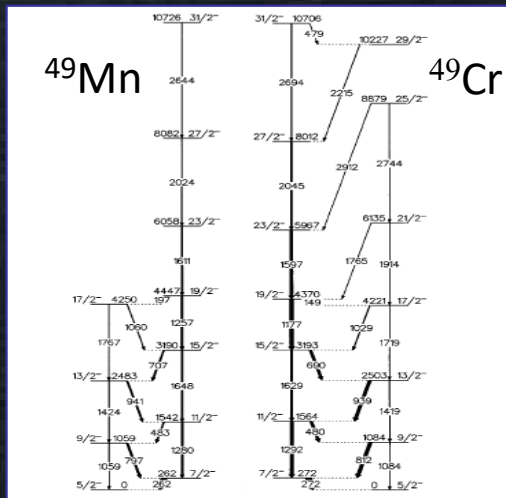


difference in excitation energies

$$MED_J = E^*_{J,-|T_z|} - E^*_{J,|T_z|}$$

Test the charge symmetry of the interaction

$$V_{pp} = V_{nn}$$



These (small) differences are mainly due to the Coulomb interaction

# What can we learn from MED

They contain a richness of information about spin-dependent structural phenomena

We can compute these energy differences with the shell model and learn about **nuclear** structure features:



- How the nucleus generates its angular momentum
- Isospin non-conserving terms of the interaction
- Evolution of radii (deformation) along a rotational band
- Estimate the neutron skin
- Learn about the configuration of the states

# Outline

**First part:** show the effect of the nuclear radius changes as a function of the angular momentum in predicting the MED

**Second part:** use the measured MED to estimate the nucleon skin for every excited state

**Third part:** the MED of non-natural parity states as a test of the type of p-h excitations across the gap

# Calculating the MED

We start from diagonalizing a nuclear hamiltonian that conserves isospin and treat Coulomb and other eventual isospin symmetry breaking (ISB) contributions **perturbatively**

$$H = H_{nuc} + V_C + V_B$$

$$V_C + V_B = V_{CM} + V_{Cm} + V_B$$

Multipole  
Coulomb

$$V_{CM}$$

- correlations
- energy gains

$$V_B$$

Isospin symmetry  
breaking term of  
non-Coulomb origin

monopole  
Coulomb

$$V_{Cm}$$

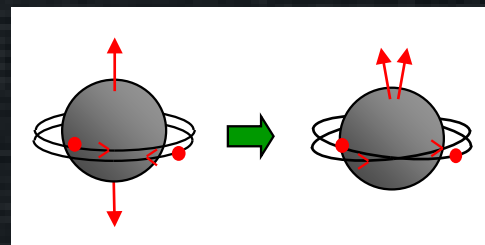
- represents a spherical mean field extracted from the interacting shell model
- determines the single-particle energies and the shell evolution

# The different terms in the MED

$$MED_J^{theo} = \Delta \langle V_{CM} \rangle_J + \Delta \langle V_{Cm} \rangle_J + \Delta \langle V_B \rangle_J$$

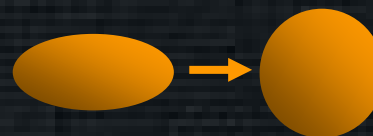
$V_{CM}$  Multipole part of the Coulomb interaction

Between valence protons only



$V_{Cm}$  monopole part of the Coulomb interaction

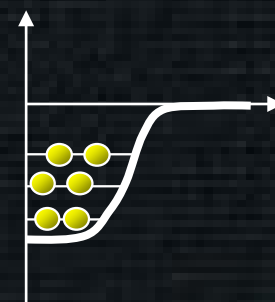
radial effect:  
radius changes with J



$\ell \cdot \ell$  term to account for shell effects

$\ell \cdot s$  electromagnetic spin-orbit term

change the single-particle energies



$V_B$  isospin symmetry breaking term

$$V_{\pi\pi}^{J=0} - V_{\nu\nu}^{J=0} = -100 \text{ keV}$$

for all orbits

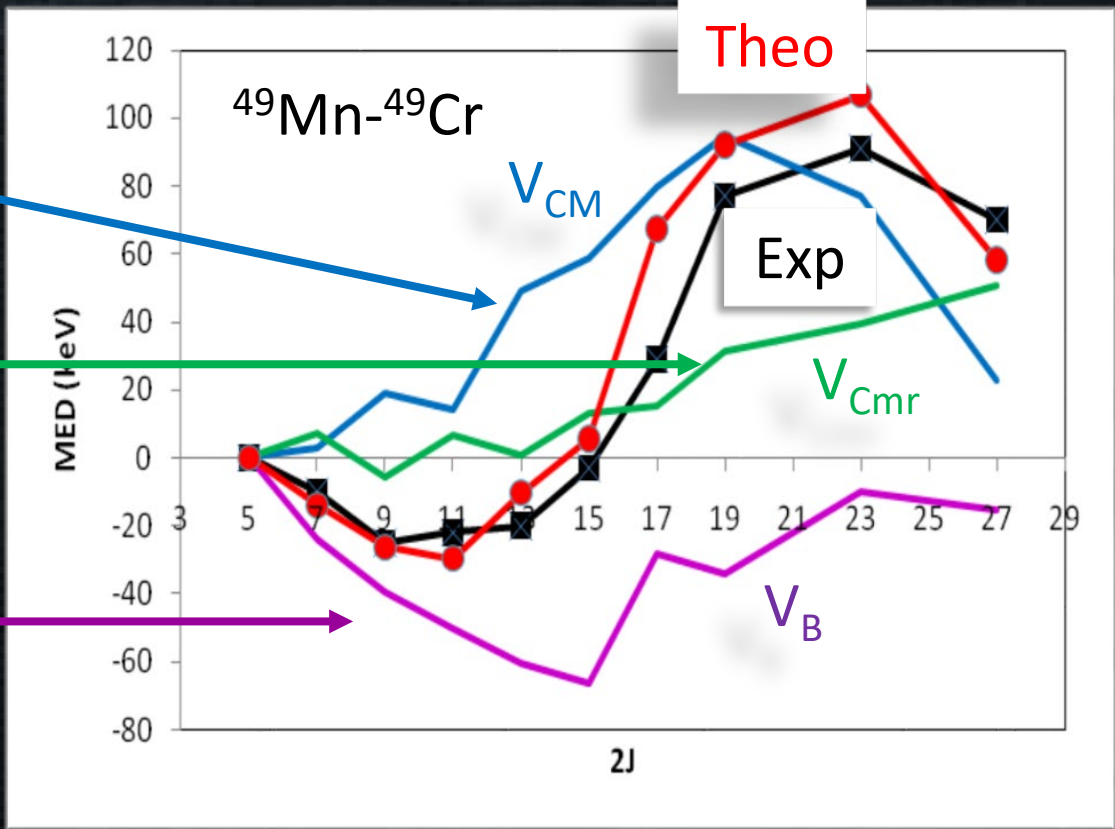
# Calculating the MED with SM

$$MED_J^{theo} = \Delta \langle V_{CM} \rangle_J + \Delta \langle V_{Cm} \rangle_J + \Delta \langle V_B \rangle_J$$

**VCM:** gives information on the nucleon alignment or recoupling

**VCmr:** gives information on changes in the nuclear radius

Important contribution from the ISB **VB** term: of the same order as the Coulomb contributions



A. P. Zuker et al., PRL 89, 142502 (2002)

# Evidence of the Coulomb radial effect

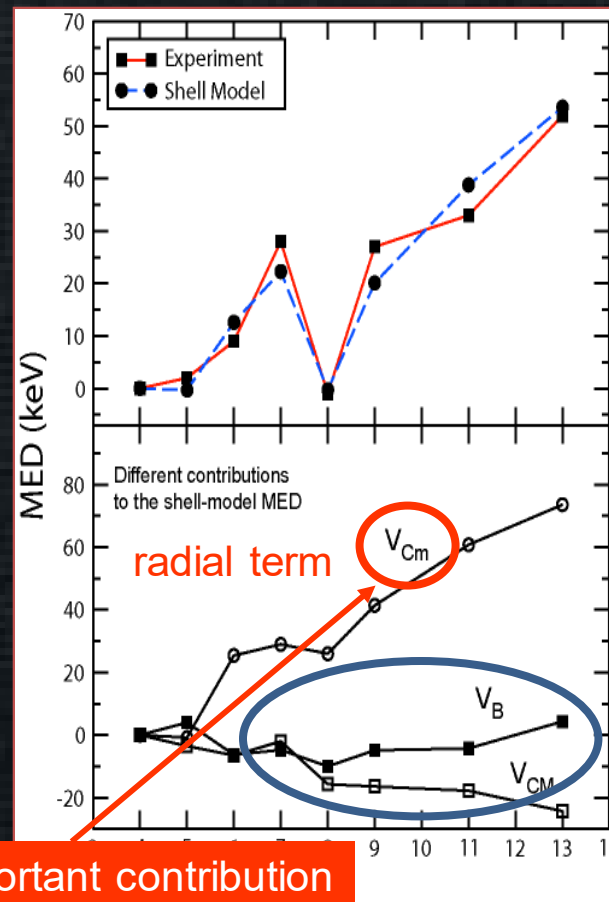
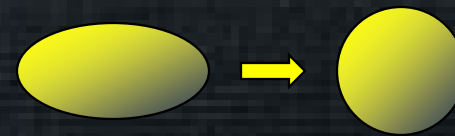
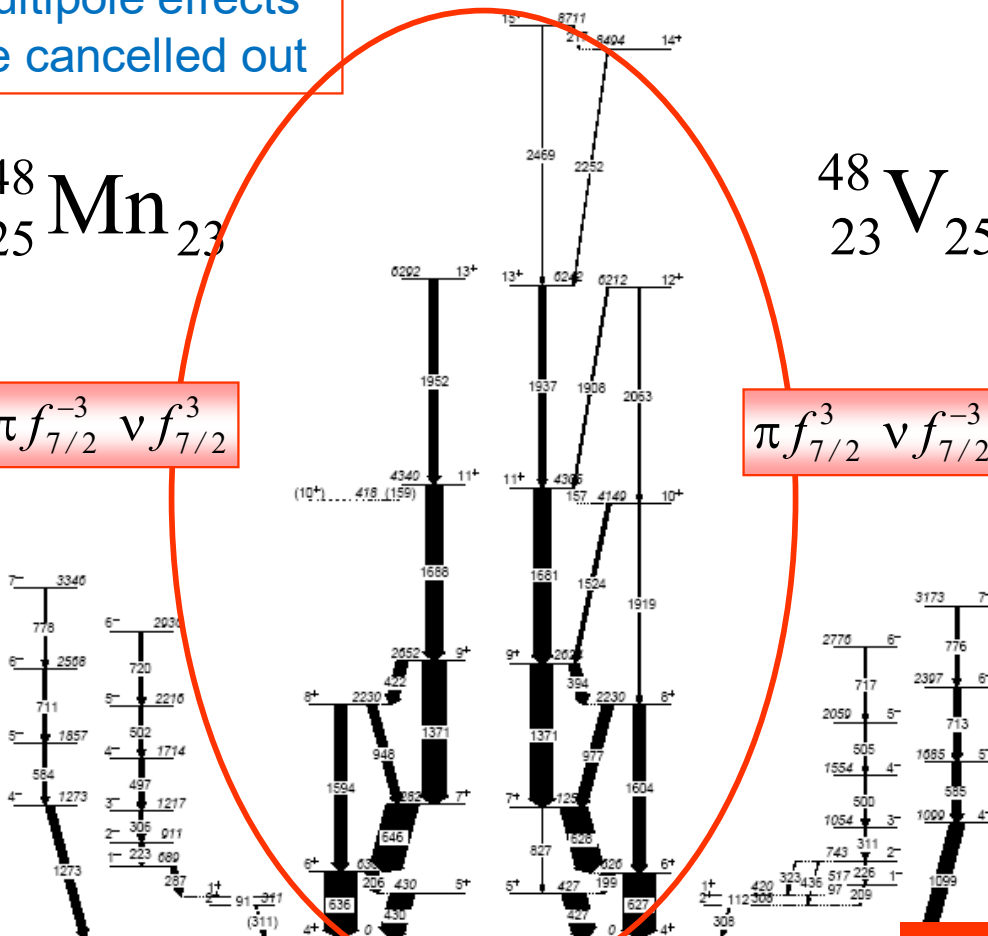
Multipole effects are cancelled out

$^{48}_{25}\text{Mn}_{23}$

$^{48}_{23}\text{V}_{25}$

$$\pi f_{7/2}^{-3} \nu f_{7/2}^3$$

$$\pi f_{7/2}^3 \nu f_{7/2}^{-3}$$



M.A. Bentley et al.,  
PRL 97, 132501 (2006)

The nucleus changes shape  
towards band termination



# The evolution of the radius

Coulomb energy of a charged sphere:

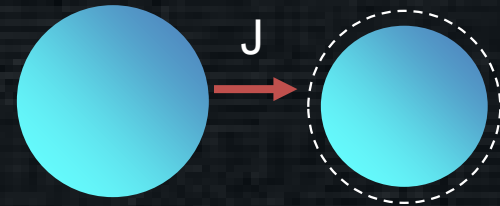
$$E_C = \frac{3Z(Z-1)e^2}{5R_C}$$

The difference between the energy of the ground states:

$$\Delta E_C(J=0) = E_C(Z_>) - E_C(Z_<) = \frac{3n(2Z_> - n)e^2}{5R_C}$$

$$T_z = \pm \frac{n}{2}$$

If  $R_C$  changes as a function of the angular momentum...



$$\Delta E_{Cr}(J) = \Delta E_C(J) - \Delta E_C(0) = \frac{3}{5}n(2Z_> - n)e^2 \left( \frac{R_C(0) - R_C(J)}{R_C^2} \right)$$

$$= -\frac{3}{5}n(2Z_> - n)e^2 \frac{\Delta R_C(J)}{R_C^2} = nC \cdot \Delta R_C(J)$$

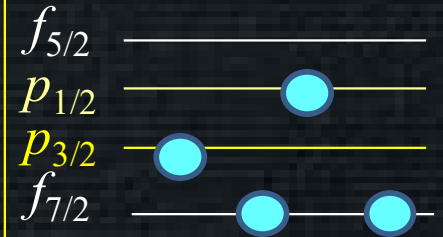
Radial contribution to the MED

# The radial effect with the shell model

The radius of a nucleus depends on the occupation of the different orbitals.

In the  $fp$  shell  $p$  orbitals have larger radius than the  $f$ .

The radial term will depend on the change of occupation of the  $p$  orbitals as a function of  $J$



$$V_{Cm,r}(J) = 2|T_z|\alpha \frac{(z_p + n_p)_J}{2}$$

$z_p$  and  $n_p$  are the number of protons and neutrons in the  $p$  orbitals, relative to the g.s. ( $J=0$ )

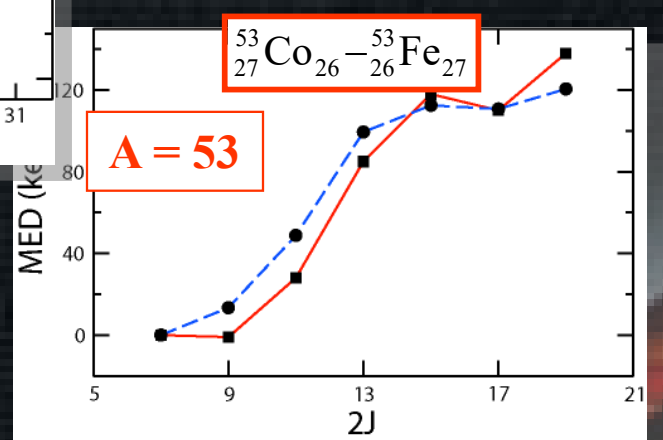
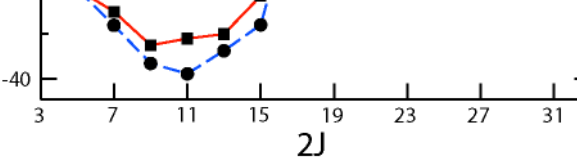
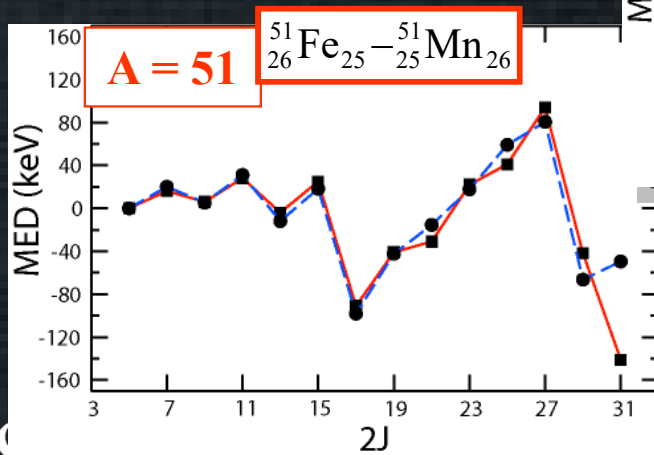
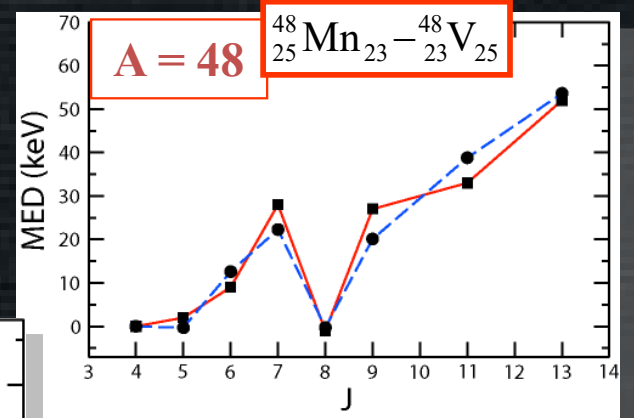
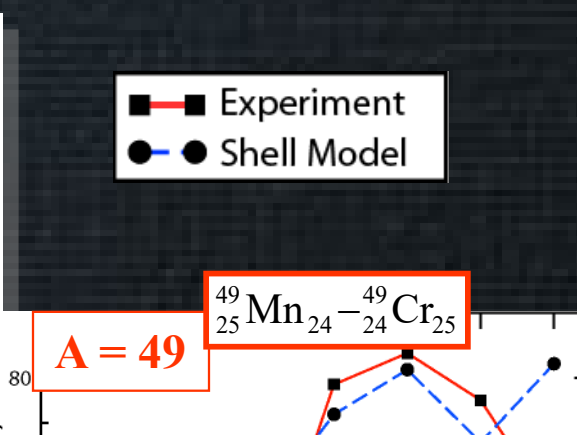
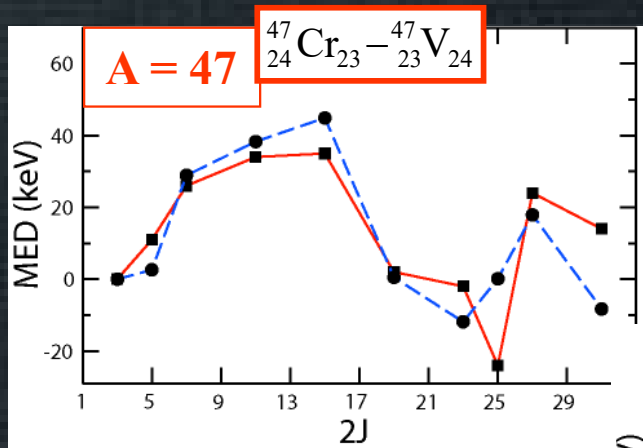
The radial monopole term depends on the occupation of the  $p$  orbitals

$\alpha$  is not a free parameter but can be estimated from experimental data:

The radial parameter amounts to  $\alpha \sim 200$  keV for nuclei in the  $f_{7/2}$  shell

# MED in the $f_{7/2}$ shell

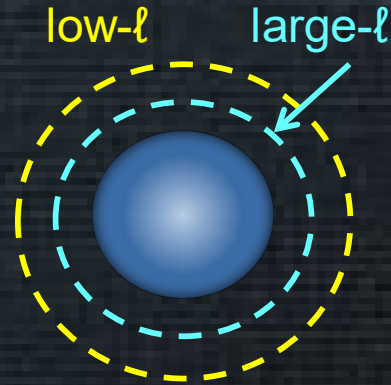
Very good quantitative description of data without free parameters



M.A. Bentley and SML,  
Prog. Part. Nucl. Phys. 59,  
497-561 (2007)

# The size of the orbital radius

In a main shell, the radial extension of low- $\ell$  orbits is much larger than the others



In the sd shell Bonnard and Zuker have found a very peculiar behaviour of the  $1s_{1/2}$  orbit:

$$r_s - r_d \approx 1.6 \text{ fm} \quad Z, N \leq 14$$

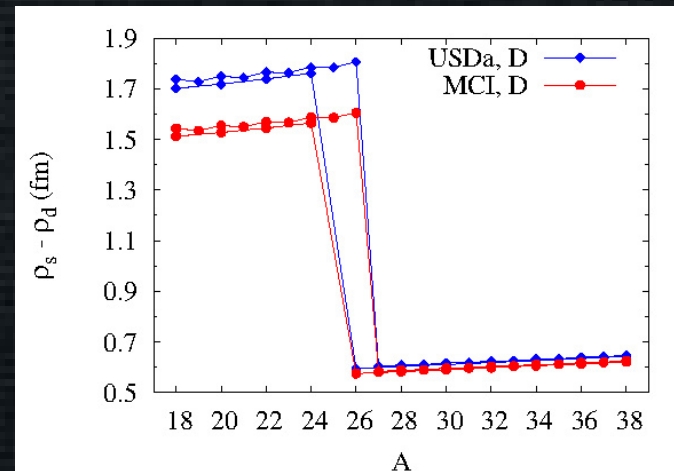
$$r_s - r_d \approx 0.6 \text{ fm} \quad Z, N > 14$$

A similar behavior is predicted in the  $fp$  shell: when the  $p$  orbits are occupied by one or more nucleons, the orbital radius decreases.

$$V_{Cm,r}(J) = |T_z| \alpha (z_p + n_p)_J$$

$$\alpha \approx 200 \text{ keV for } (z_p + n_p) < 1$$

$$\alpha \ll 200 \text{ keV for } (z_p + n_p) \geq 1$$



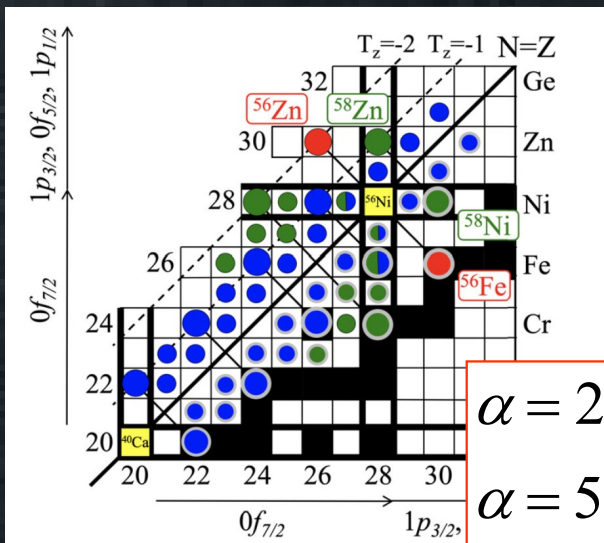
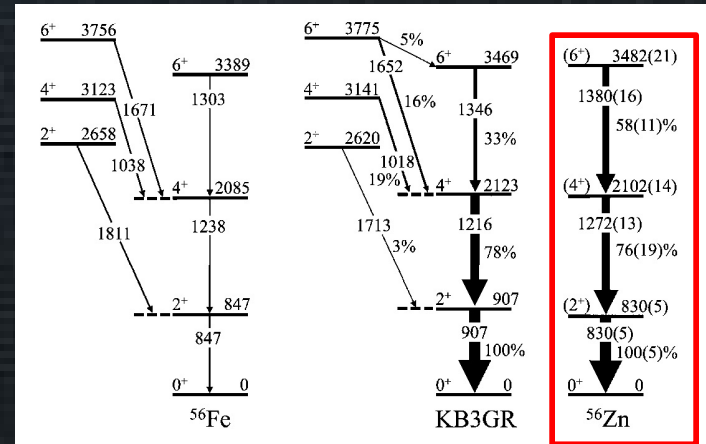
J. Bonnard and A. P. Zuker,  
JoP Conf. Series 1023 (2018) 012016

# MED in T=2 A=56 mirrors

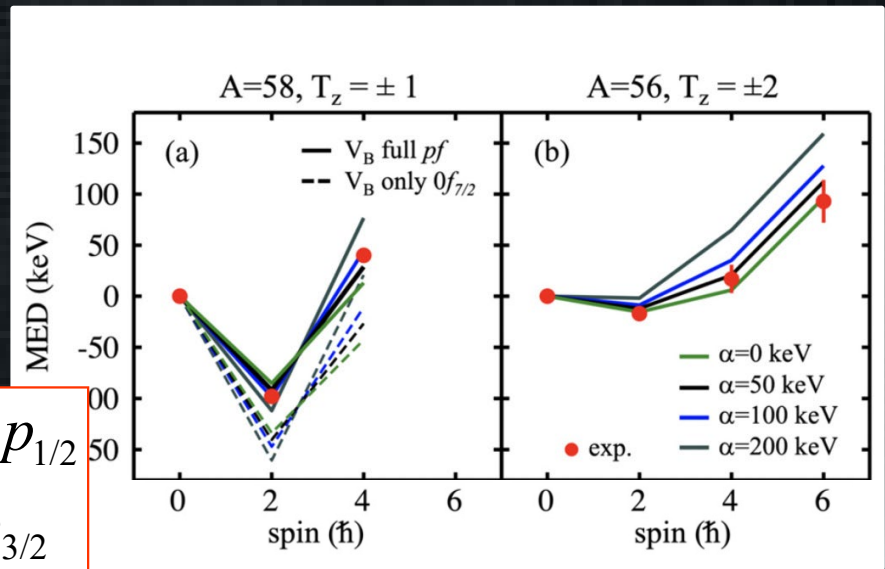
Recently, excited states in  $^{56}\text{Zn}$  ( $T=2$ ) have been observed for the first time in RIKEN

$$V_{Cm,r}(J) = |T_z| \alpha (z_p + n_p)_J$$

The role of the radial term increases with T



$\alpha = 200$  keV for  $p_{1/2}$   
 $\alpha = 50$  keV for  $p_{3/2}$

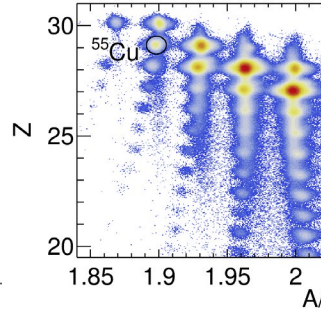


A. Fernandez et al., Phys. Lett. B 823 (2021) 136784

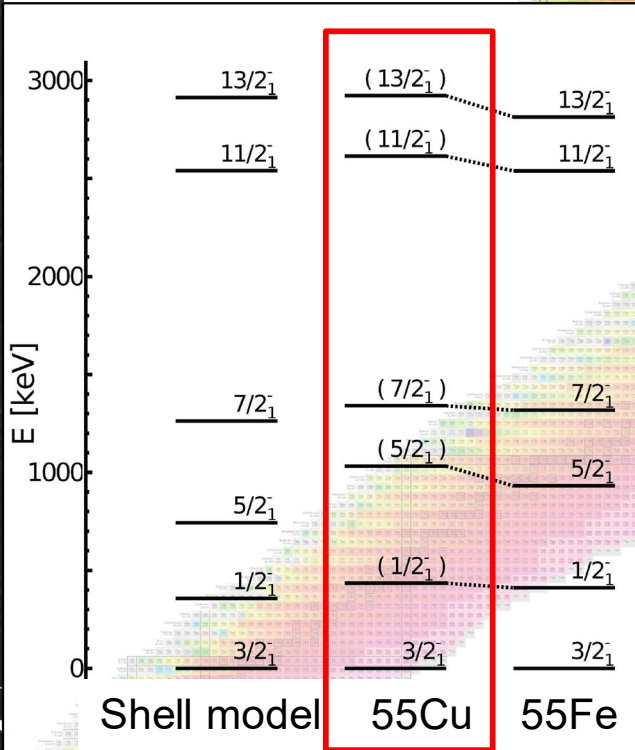
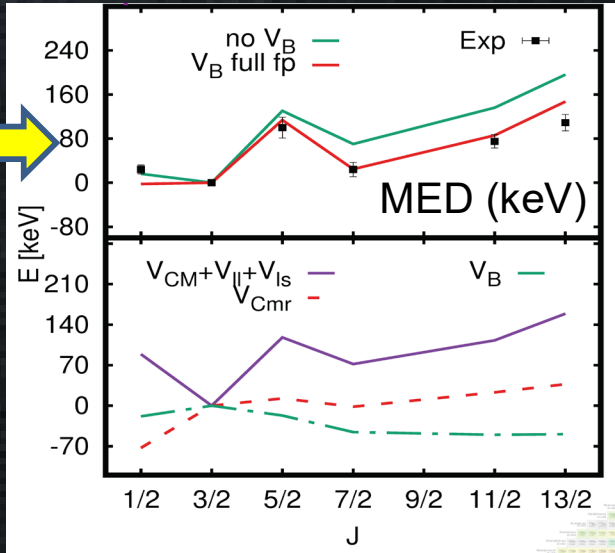
# MED in $T=3/2$ $A=55$ mirrors

$3/2^-$   
 $^{55}\text{Cu}$

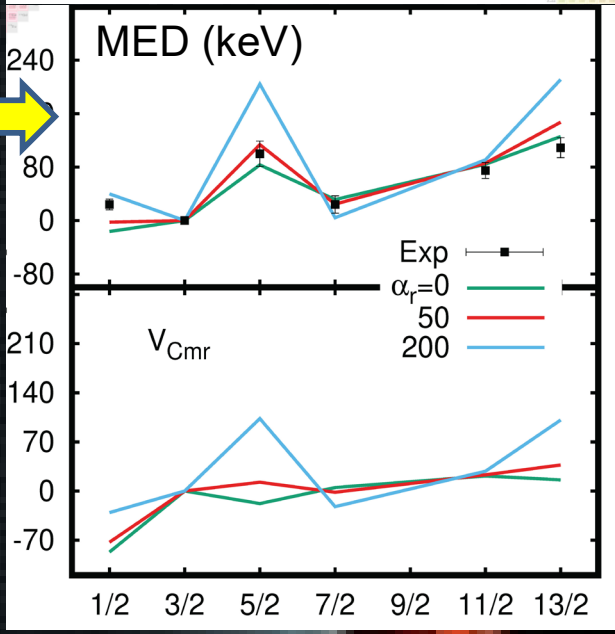
First spectroscopy of  $^{55}\text{Cu}$



Importance of the contribution of the **ISB  $V_B$**  term for **all orbits**



Effect of the **reduced radius of the  $p_{3/2}$  orbit** that is occupied by more than one nucleon



Sara Pigliapoco *et al.*, *tbp* (2023)

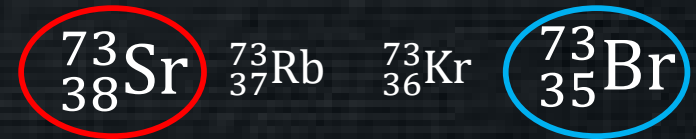
# Take-home message

The radius of the low- $\ell$  orbit in a main shell decreases when is occupied by one or more nucleons

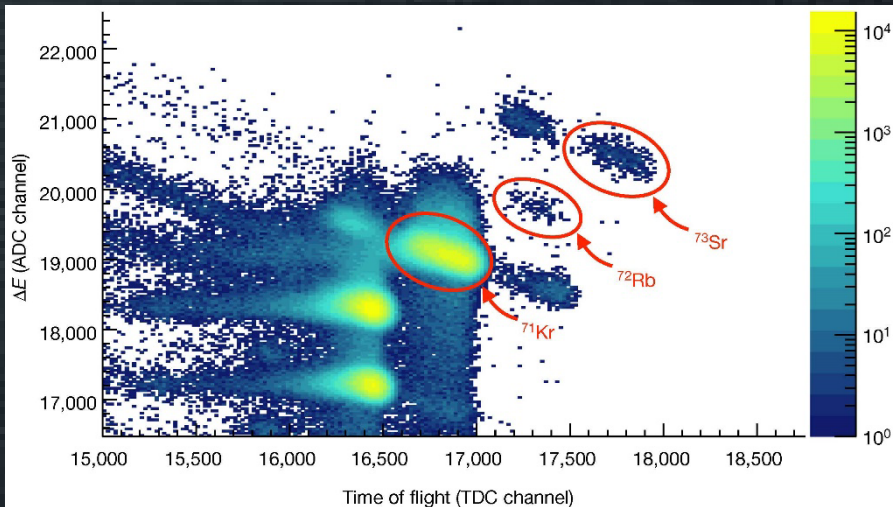
Recent MED data for  $T = 3/2$   $A = 61, 73$  and  $T=1$   $A = 62, 70$  confirm these conclusions

# The case of $^{73}\text{Sr}$ - $^{73}\text{Br}$

Study the decay of the  
 $T = 3/2, T_z = -3/2, ^{73}\text{Sr}$

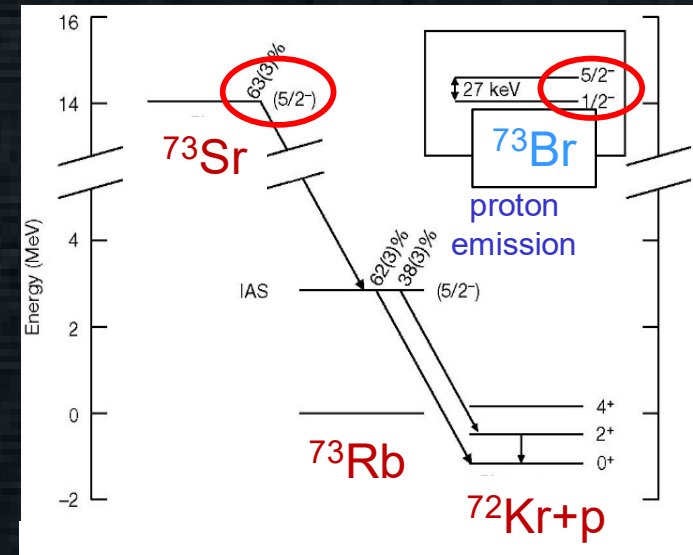


Isobaric multiplet



A  $J^\pi = 5/2^-$  spin assignment to the g.s. of  $^{73}\text{Sr}$  is needed to explain the proton-emission pattern observed from the  $T = 3/2$  IAS in  $^{73}\text{Rb}$

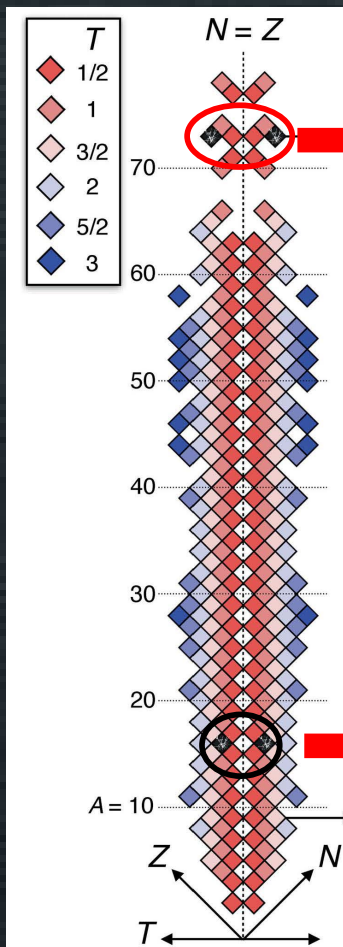
→ the ground state of  $^{73}\text{Sr}$  differs from that of its mirror  $^{73}\text{Br}$



E. M. Ho, A. M. Rogers *et al.*,  
 Nature 580 52 (2020)

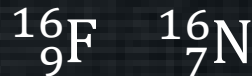


# Interpretation of the data



Prolate and oblate shapes coexist in this mass region

Thomas-Ehrman shift



Using the effective interaction PFSDG-U (F. Nowacki et al. PRL 117, 272501 (2016), we compute the MED thus predicting the excitation energy of the 1/2- state at 16 keV in  $^{73}\text{Sr}$

S.M. Lenzi, A. Poves and A. O. Macchiavelli, PRC 031302(R) (2020)



MED and nuclear skin

# Charge radii and nuclear skin

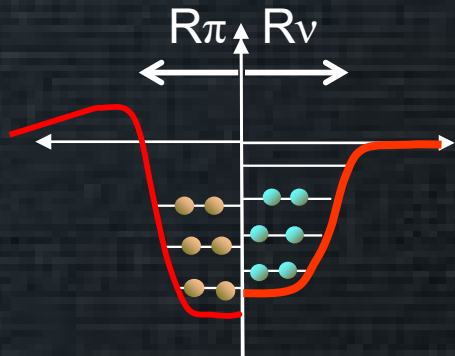
Charge radii can be measured via electron scattering

These measurements are limited to stable nuclei

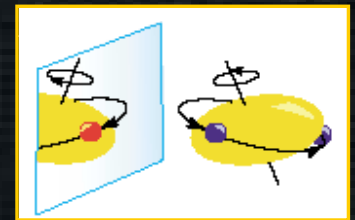
Neutron skin is still more difficult to measure

Laser spectroscopy allows to measure radial shifts along isotopic chains

This applies to ground states or isomeric states



Can we get any information on the evolution of radii in excited states and on the neutron skin?

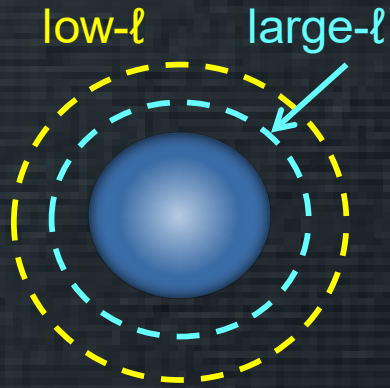


# We will now use the MED data to deduce the nuclear skin

We will calculate the excitation energy of the levels in both mirror nuclei within the shell model framework using a **realistic interaction in the *sd* shell that naturally includes all ISB terms.**

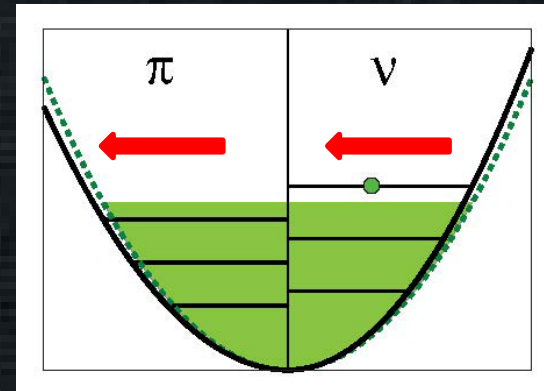
We will compute the MED and compare with data, varying the neutron skin in an iterative way, until we fit the experimental MED

# Proton and neutron radii



Studying mirror energies in doubly-magic nuclei + 1 nucleon

## Isovector monopole polarizability



A	$J^\pi$	$\Delta r_{\nu\pi}$ (fm)	
17	$5/2^+$	0.056	$d_{5/2}$
	<b><math>1/2^+</math></b>	<b>0.147</b>	$s_{1/2}$
41	$7/2^-$	0.015	$f_{7/2}$
	$5/2^-$	0.018	$f_{5/2}$
	<b><math>3/2^-</math></b>	<b>0.038</b>	$p_{3/2}$
	<b><math>1/2^-</math></b>	<b>0.037</b>	$p_{1/2}$

The addition of a nucleon induces changes in the potential wells of both protons and neutrons and tends **to equalize the radii**

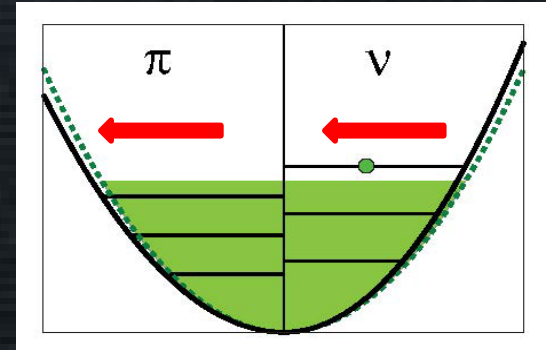
J. Bonnard, S.M.L. and A.P. Zuker,  
PRL 116, 212501 (2016)

# Radii and MED

The size parameters are determined using:

- the (measured) charge radius of the neutron-rich partner
- the (measured) MED
- Isospin-symmetry arguments

$$\langle r_{\pi,v}^2 \rangle \propto \frac{1}{\hbar\omega_{\pi,v}}$$



measured!

$$r_{\pi}(N > Z) = r_{\nu}(N < Z) \quad \text{isospin symmetry}$$

The charge radius of the proton-rich partner is **obtained from the MED**

$$r_{\pi}(N < Z) = r_{\nu}(N > Z)$$

J. Bonnard et al., PRL 116, 212501 (2016)

Due to the isovector monopole polarization, we need to determine the size of both potential wells to calculate the matrix elements of the effective interaction to obtain the MED. **They are different for protons and neutrons!**

# MED and neutron skin in A=23

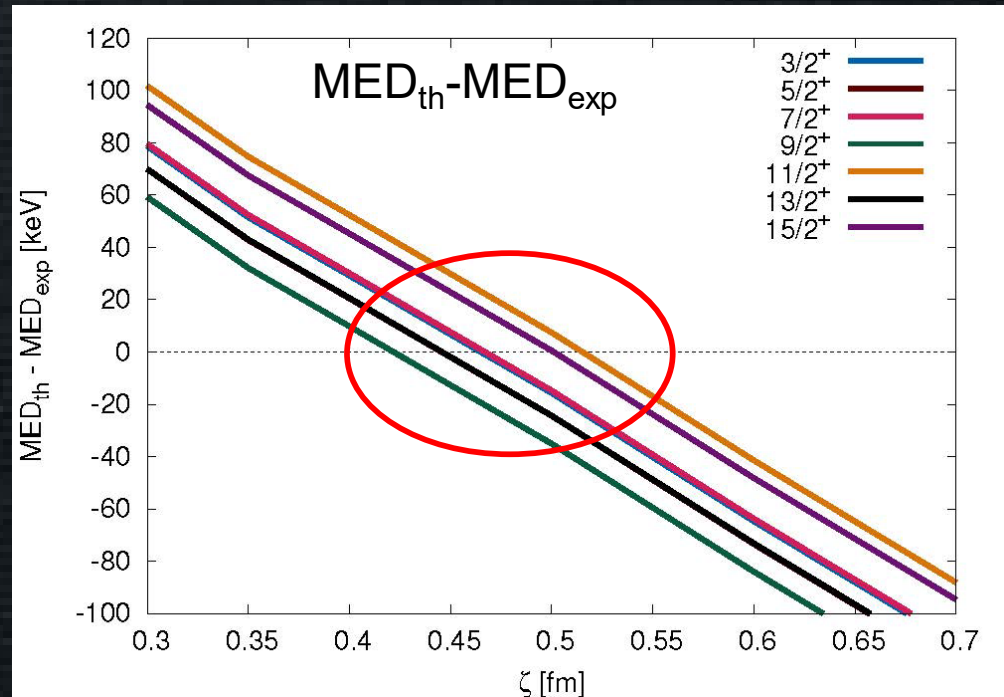
The MED depend linearly on the value of the neutron skin

neutron skin

$$\Delta r_{v\pi} = \sqrt{\langle r_v^2 \rangle} - \sqrt{\langle r_\pi^2 \rangle} = \zeta \cdot f(A, T)$$

J. Duflo, A. P. Zuker, PRC 66, 051304 (2002)

We vary  $\zeta$  to match the experimental MED

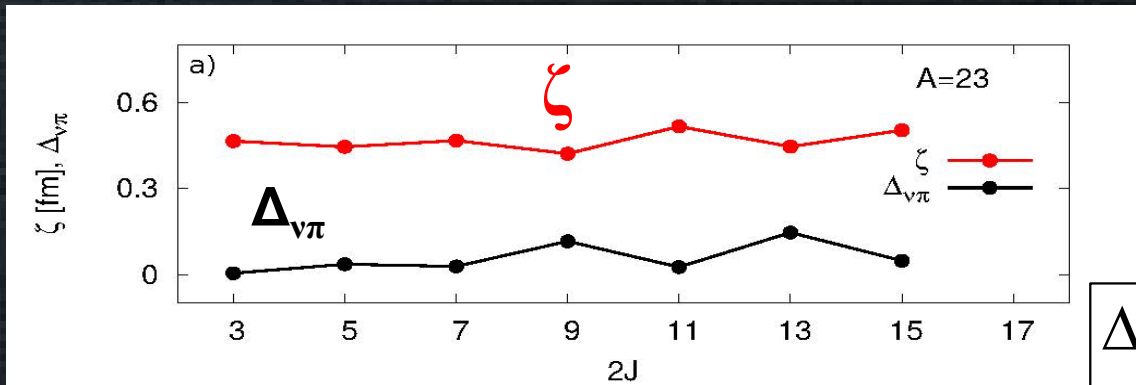


A. Boso *et al.*, Phys. Rev. Lett. 121, 032502 (2018)

# MED and neutron skin

For  $^{23}\text{Na}$  we obtain the neutron skin (in fm):

$J$	3/2	5/2	7/2	9/2	11/2	13/2	15/2
$\Delta r_{\nu\pi}$	0.0211	0.0202	0.0211	0.0192	0.0233	0.0202	0.0226



$$\Delta_{\nu\pi} = n_{s_{1/2}} - Z_{s_{1/2}}$$

Interestingly, the skin is correlated with the difference of occupation number of neutrons and protons  $\Delta_{\nu\pi}$  in the low- $\ell$  orbit  $s_{1/2}$ !

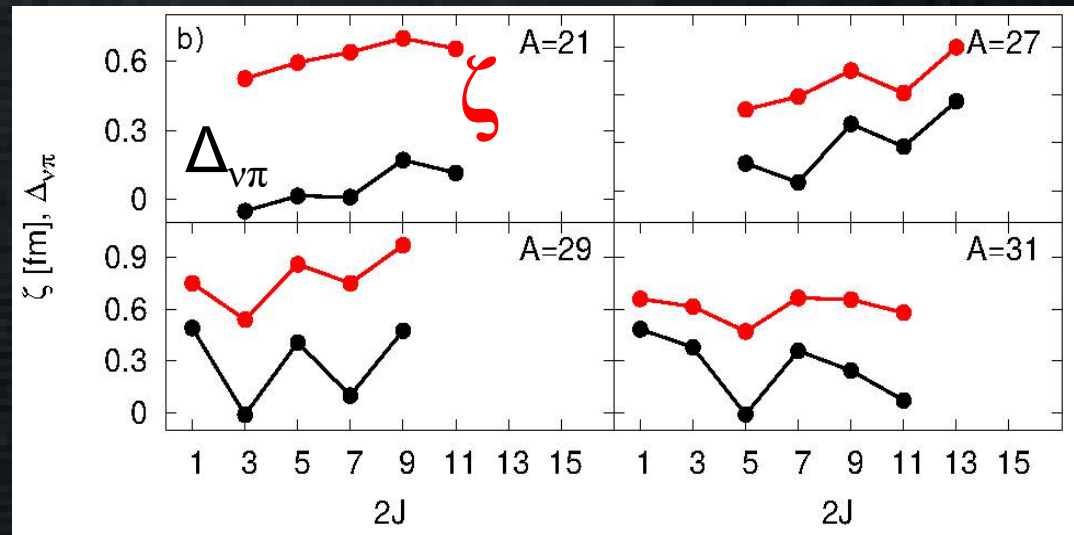
A. Boso *et al.*, Phys. Rev. Lett. 121, 032502 (2018)



# Correlation between skin and difference of occupation numbers

We apply this procedure to the MED for nuclei in the  $sd$  shell and deduce the value of the skin for each excited state

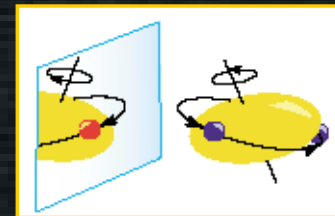
In all cases, the skin is correlated with the difference of occupation number of neutrons minus protons ( $\Delta_{\nu\pi}$ ) in the  $s_{1/2}$  orbit!



A. Boso *et al.*,  
Phys. Rev. Lett. 121, 032502 (2018)

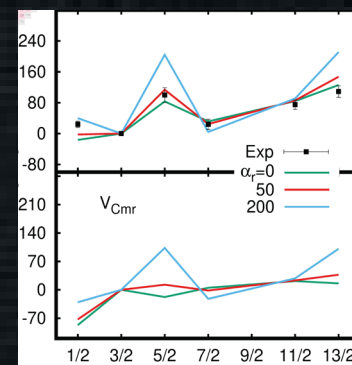
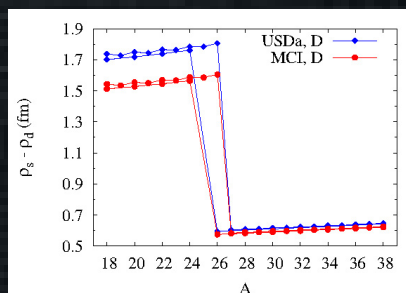
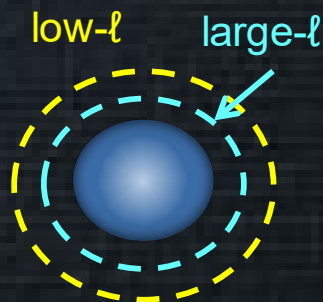
# What do we learn from MED?

The MED are sensitive to the nuclear structure and therefore constitute a very powerful tool to understand several nuclear properties provided we use a non-free parameter, unique method for all mass regions



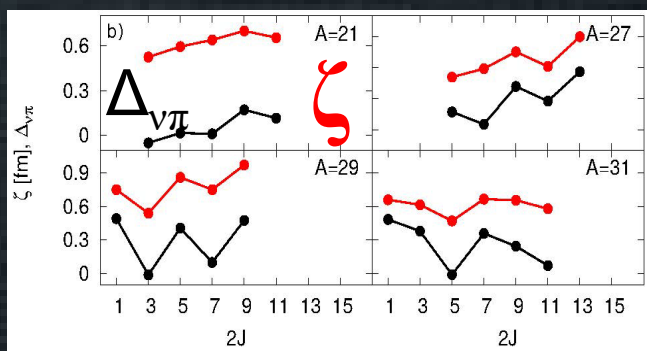
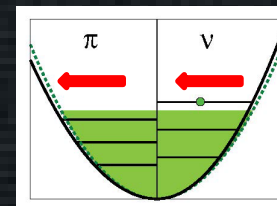
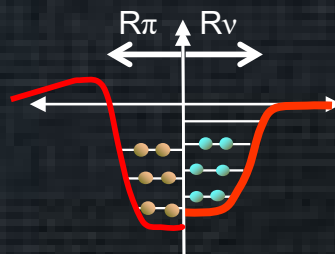
In particular, MED depend on the nuclear radius (average of protons and neutrons).

Low- $\ell$  orbits reduce their radius with occupancy.



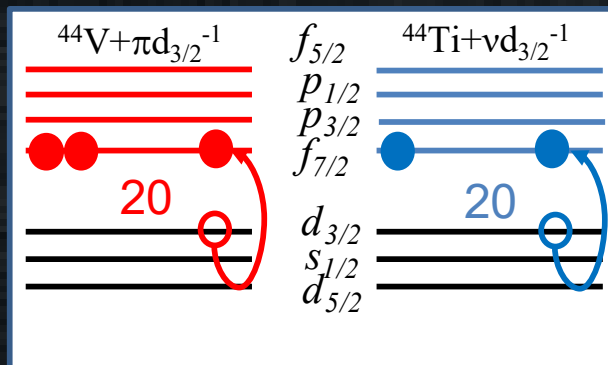
# What do we learn from MED?

MED can give us information on the nuclear skin and its evolution with the spin.



There is a clear correlation between the skin and the difference of occupation of neutrons and protons of the low- $l$  orbit.

The MED of non-natural parity give direct information on the character of p-h excitations in the wavefunction



# Thank you for your attention

1222-2022  
800  
ANNI



UNIVERSITÀ  
DEGLI STUDI  
DI PADOVA



Silvia M. Lenzi – March 29, 2023

# Backup slides

# MED and neutron skin

Using a 5-parameter fit of measured charge radii for  $A < 60$

J. Duflo, A. P. Zuker, PRC 66, 051304 (2002)

$$\sqrt{\langle r_{\pi}^2 \rangle} = A^{1/3} \left( \rho_0 + \frac{\zeta}{2} \frac{t_z}{A^{4/3}} - \frac{\nu}{8} t_z^2 \right) e^{g/A} + \lambda D_{\pi\nu}$$

Obtain  $\hbar\omega$

compute the matrix elements of a realistic CD interaction

Calculate the MED

Vary  $\zeta$  to match the experimental MED

$$\Delta r_{\nu\pi} = \sqrt{\langle r_{\nu}^2 \rangle} - \sqrt{\langle r_{\pi}^2 \rangle} = \frac{\zeta t}{A} e^{g/A}$$

Obtain the neutron skin

Example: for nuclei with one nucleon over a doubly-closed shell nucleus, the neutron skin varies linearly with the  $\zeta$  parameter

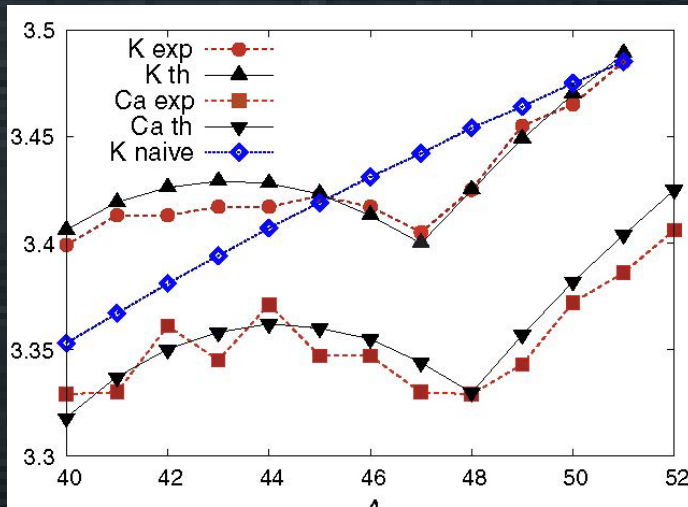
A	J $\pi$	$\Delta r_{\nu\pi}$ (fm)	$\zeta$ (fm)
17	5/2 <sup>+</sup>	0.056	0.90
	<b>1/2<sup>+</sup></b>	<b>0.147</b>	<b>2.37</b>
41	7/2 <sup>-</sup>	0.015	0.61
	5/2 <sup>-</sup>	0.018	0.71
	<b>3/2<sup>-</sup></b>	<b>0.038</b>	<b>1.50</b>
	<b>1/2<sup>-</sup></b>	<b>0.037</b>	<b>1.48</b>

*d*  
*s*  
*f*  
*f*  
*p*  
*p*

J. Bonnard et al., PRL 116, 212501 (2016)

# Understanding ISB effects

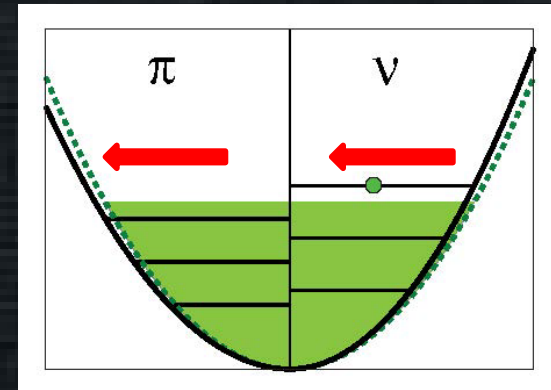
Charge radii in Ca and K isotopes



$\leftarrow$   $\nu f_{7/2}$   $\leftarrow$   $\nu p_{3/2}$   $\rightarrow$

Proton radii increases when filling with neutrons the  $p_{3/2}$  shell!

An effect to consider is the isovector polarizability



the addition of a nucleon induces changes in the potential wells of both protons and neutrons

An important conclusion from this work is that low- $\ell$  orbits have much larger radii than their partners in a main shell