# FIRST WIMP SEARCH RESULTS FROM THE LUX-ZEPLIN EXPERIMENT Amy Cottle, University of Oxford



#### COLLABORATION







#### Science and Technology Facilities Council







#### <u>@lzdarkmatter</u> https://lz.lbl.gov/





**LZ Collaboration Meeting University Of Maryland** 5<sup>th</sup>-7<sup>th</sup> January 2023



36 Institutions: ~250 scientists, engineers, and technical staff



#### **INTRODUCTION TO LZ**



#### <u>NIM A, 163047 (2019)</u>



- Based at the Sanford Underground
- Dual-phase xenon time projection •

#### **TPC DETECTION PRINCIPLE**

- Interactions in the xenon create
  - Light prompt scintillation S1
  - Charge electrons drifted and extracted into gas -> proportional scintillation - S2
- Excellent 3D position reconstruction (~mm)
- S2:S1 ratio discriminate electronic recoils (ERs) from potential WIMP nuclear recoils (NRs)
- Distinguish between single scatter (SS) and multiple scatter (MS) interactions



# **VETO DETECTOR ANTI-COINCIDENCE**

- 17 tonnes Gd-loaded scintillator in OD



#### **TPC & SKIN ASSEMBLY**







## **OD CONSTRUCTION & UNDERGROUND INSTALLATION**

#### Water Tank Panoramic

#### **Cryostat Insertion**

#### **Instrumented OD**



## FIRST SCIENCE RUN (SR1)

- 116 calendar days -> 89 live days
- Stable detector conditions
  - Temperature of 174.1 K
  - Gas pressure of 1.791 bar
  - Drift field of 193 V/cm
  - Extraction field of 7.3 kV/cm (in gas)
- Continuous purification at 3.3 t/day through hot getter system
- Demonstration run, no explicit bias mitigation



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# **TPC CALIBRATIONS**

- Backgrounds predominantly ERs; WIMPs produce NRs
- Tritiated methane (CH<sub>3</sub>T) injection to calibrate ER band
  - Spatially homogenous β source
- DD neutron generator (NR band) •
  - Monoenergetic 2.45 MeV neutrons
- 99.9% discrimination of beta backgrounds under NR band median achieved

4.50 4.25 4.00[[phd]] 3.75 S2( 3.50  $\log_{10}$ 3.25 3.00 2.75 2.50







# **ACCIDENTAL COINCIDENCE BACKGROUNDS**

- Unrelated S1s & S2s can accidentally combine to produce single scatter events
- Rate: population of definite accidental events with drift time >1 ms
- Distribution: fake events constructed from lone S1 & S2 pulse waveforms
- Analysis cuts developed to combat observed pulse/event pathologies
  - >99.5% efficiency in removing accidentals
  - SR1 WIMP search counts:  $1.2 \pm 0.3$



# counts/tonne/year

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## WIMP ANALYSIS - ROI & FV

- Region of Interest definition •
  - 3 < S1c < 80 photons detected (phd); three-fold PMT coincidence
  - Uncorrected S2 > 600 phd; log10 (S2c) < 5
- Fiducial volume (FV) definition •
  - 86 us < drift time < 936.5 us cut to avoid higher background rates at TPC edges
  - Radial cut chosen to ensure <0.01 wall \_\_\_\_ background counts in the FV
- Calculated fiducial mass of  $5.5 \pm 0.2$  t •



# WIMP ANALYSIS - CUTS & DATA QUALITY

1.0

- Event selection criteria
  - FV, ROI, single scatter cuts
  - Veto detector anti-coincidence
  - S1/S2 shape cuts
- Cuts developed on non-WIMP ROI background & calibration data
- Rejection of live time with detector instabilities, high TPC pulse rates
  - $60 \pm 1$  live days



# WIMP ANALYSIS - DATA & STATISTICAL INFERENCE

• 335 events after all cuts

- PDFs created with energy deposit + detector response simulations\*
- Profile likelihood ratio (PLR) analysis



- 1 & 2-Sigma Contours
  Post-fit total background distribution
- <sup>37</sup>Ar
- <sup>8</sup>B
- 30 GeV/c<sup>2</sup> WIMP
- NR band from DD

\* j.astropartphys.2020.102480

4.50 4.25 4.00 [[phd]] 3.75  $\log_{10}(S2c$ 3.50 3.25 3.00 2.75



# WIMP ANALYSIS - BACKGROUNDS & STATISTICAL INFERENCE

Component	Expected Events	Best Fit Events
β decays & detector γs	215 ± 36	222 ± 16
37 <b>Ar</b>	[0, 288]	$52.5^{+9.6}_{-8.9}$
<sup>127</sup> Xe	9.2 ± 0.8	$9.3 \pm 0.8$
<sup>124</sup> Xe	5.0 ± 1.4	5.2 ± 1.4
<sup>136</sup> Xe	$15.1 \pm 2.4$	$15.2 \pm 2.4$
Solar v ERs	27.1 ± 1.6	27.2 ± 1.6
<sup>8</sup> B CEvNS	0.14 ± 0.01	0.15 ± 0.01
Det. Neutrons	$0.00^{+0.02}$	$0.00^{+0.02}$
Accidentals	$1.2 \pm 0.3$	$1.2 \pm 0.3$
Total w/o <sup>37</sup> Ar	$273 \pm 36$	280 ± 16
Total w/ 37Ar		333 ± 17



# WIMP ANALYSIS - SR1 SPIN-INDEPENDENT LIMIT

- Two-sided PLR search with power-constrained limit defined using rejection power
- Minimum cross-section of  $\sigma_{SI}$ = 9.2 × 10<sup>-48</sup> cm<sup>2</sup> for WIMP mass of 36 GeV/c<sup>2</sup>
- No evidence for WIMPs

#### <u>Key</u>

- Observed limit
- Median expected sensitivity

EPJC 81, 907 (2021), arXiv:1105.3166



#### WIMP ANALYSIS - SR1 SPIN-DEPENDENT LIMITS **WIMP-Neutron Scattering**



## WIMP SEARCH PROSPECTS

- SR1 covers just 6% of planned full exposure of 1000 live days
  - Still a lot of parameter space explorable with LZ
  - 1000-day projected sensitivity:

90% CL minimum: 1.4 x 10<sup>-48</sup> cm<sup>2</sup> at 40 GeV/c<sup>2</sup>

PRD 101, 052002 (2020)

 $10^{-42}$ 10-43  $[cm^2]$  $10^{-44}$ **G**SI 10-45 10-46 10-47 10-48 10-49

-nucleon

WIMP-





## CONCLUSIONS

- World-leading spin-independent WIMP search limit achieved with just 6% of planned exposure
  - Now entering discovery parameter space
- Background sources well examined & documented in <u>dedicated paper</u>
- Multiple other physics channels to explore
  - Papers in preparation on SR1 data
- <u>XLZD consortium</u> formed, looking towards the ultimate xenon rare physics observatory



# BACKUP SLIDES



## TPC ENERGY RESPONSE

- S1s & S2s position-corrected using <sup>131m</sup>Xe background, <sup>83m</sup>Kr calibration
- Doke plot constructed with monoenergetic electron recoil peaks





# **VETO DETECTOR RESPONSE**

- Skin & OD response and inter-detector timings calibrated
  - OD optical calibration system
  - External γ-ray & neutron sources (e.g. <sup>22</sup>Na; DD, AmLi, <sup>252</sup>Cf)
- <sup>127</sup>Xe Skin tagging efficiency of 78 ± 5% based on K-shell analysis
- OD tagging efficiency of TPC-interacting neutrons of  $89 \pm 1\%$  (AmLi calibrations)
  - TPC-OD coincidence window: 1200 µs; threshold equivalent to ~200 keV





## **AR37 ESTIMATE**

- Ar37 a significant background in early LZ data/SR1 WIMP search
  - K-shell e<sup>-</sup> capture -> 2.8 keV
  - $\tau_{1/2} = 35$  days
- Can be produced via cosmic spallation on xenon
  - Calculated using the ACTIVIA package & estimated exposure of the xenon during transport\*
  - Large uncertainties in spallation cross-section

\* PRD 105, 082004 (2022)







#### NEUTRONS

- OD Gd-loaded scintillator high thermal neutron capture cross-section
  - Measured OD neutron tagging
     efficiency of 89 ± 1%
- Likelihood analysis of sideband of events passing all WIMP search cuts except OD anti-coincidence
  - Constraint in sideband of 0<sup>+0.8</sup>
     events
  - Constraint on SR1 WIMP search neutron background of 0<sup>+0.2</sup> events



# LIVE TIME VETOES

- thus contributing to accidental coincidence backgrounds
- Removal of periods • after S2s (e-/ph trains) excludes ~30% of our live time
- Working on optimising this live time veto for future runs



## LIMIT SHAPE

#### Downward fluctuation in limit caused by deficiency of events under Ar37 contour







Calibrations and Xe127 M-shell counts as expected under signal acceptance model -> background under-fluctuation





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## LZ LIMIT UPDATE





#### **BACKGROUNDS PAPER**



#### <sup>60</sup>Co Dat Cavern gamma <sup>232</sup>Th-early Fit Internals <sup>136</sup>Xe $2\nu\beta\beta$ <sup>232</sup>Th-late 2000 1500 2500

Reconstructed Energy [keVee]

accessful background model built for SR1 underpinning WIMP search result

odel extends beyond the WIMP search region of interest to other energy ranges



#### WHAT'S NEXT FOR LZ?



• LZ plans to take 1000 live days of data = x17 more exposure than SR1

Broad physics programme available e.g. neutrinoless double beta decay, solar axions • PRC 102, 014602 (2020), PRD 104, 092009 (2021)

## **BEYOND LZ - XLZD CONSORTIUM**

- XLZD consortium formed from the LZ, XENON and DARWIN collaborations
- Coming together to build the ultimate dual-phase multi-ten tonne xenon dark matter detector
- Observatory for other rare physics
- See <u>https://xlzd.org</u> and our joint white paper (<u>arXiv:2203.02309</u>)

