

iP2

Status on Ho

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erc | USNAC

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Context | ACDM Works

Only 6 free parameters

Baryon Acoustic Oscillation

Clusters

...

Weak Lensing

Baryon Nucleosynthesis

Scolnic 2018



Context | ACDM Works, **except when it doesn't !**

Only 6 free parameters | *but "Λ" and "CDM"*

*H*_o Tension $\int \sigma$

Universe's expansion is too fast



Planck 2020



Heymans et al. 2020



Two approaches | Hubble-Lemaître Constant

Direct Method

$$H_0 \sim d_l / v_h$$



Redshifts & Distances

Indirect Method

$$H(\underline{z}) = H_0 \times \sqrt{\Omega_r (1 + \underline{z})^4 + \Omega_m (1 + \underline{z})^3 + \Omega_\Lambda (1 + \underline{z})^{3(1 + \underline{z})^4}}$$



Model & High redshift anchoring



Type Ia Supernova Cosmology



Type Ia Supernova Cosmology









Indirect determination of *H*₀









Ho Tension | SHoES vs. Planck



Are Supernovae & CMB in tension ? No!



Inverse Distance Ladder

Get independent distances for SNe Ia Sets the scale of the Universe at z~1 **"BAO."** "CMB" GOODS SCP 3000 CANDELS 2000 +CLASH 13.8 Gyr 5.5 Gyr 3.5 Gyr



BOSS DR12 | Alam et al. 2017

See also e.g.: Aubourg et al. 2015 • Macaulay et al. 2018



12



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Ho Tension | Early vs. Late



H₀ Tension | Change the model ?

H_o Tension | 5 σ

Universe's expansion is too fast



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Multiple images of the background quasar

Background quasar

Looking further into the past



Ho Tension | Systematics in strong lensing



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Ho Tension | +Mega-masers & KiloNova





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H₀ Tension | TRGB vs. Cepheid



Strong Lensing systematics actively studied

> Nearby so Sensitive to peculiar velocity correction

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The Progenitor issue | Astrophysical biases









High fraction of young stars

Rigault et al. 2020 Nicolas et al. 2021 Briday et al. 2022 Ginolin et al. *in prep*

 $lsSFR \propto \frac{\# Young Stars}{\pi}$ # Old Stars



22

Astrophysical Bias affecting H₀



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Rigault et al. 2015

3% bias on H_0

So a 2 km s⁻¹ Mpc⁻¹ shift

Total current SH0ES error budget **1.04 km s⁻¹ Mpc⁻¹**

SH0ES "corrected" \sim 71 ± 1.5 km s⁻¹ Mpc⁻¹

Rigault et al. in prep. | Rigault et al. 2015, 2020

SH0ES rebuttal

"If we **mimic the Cepheids selection** function and only take Hubble flow SNe Ia from *Spiral* hosts, *H*₀ reduces by 0.5%"

Riess et al. 2022 | Riess et al. 2016, 2019



Zwicky Transient Facility (ZTF) is acquiring ~1000 SNeIa per year at z<0.1 since 2018



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ZTF

We will be observing during O4

Localization of GW170817 was smaller than ZTF FoV



Abbot et al. (2017), PRL 119, 161101



ZTF | Changing the scale of SN Cosmology



Smith, Rigault et

al.	in	pr	ер	
			26	

ZTF Sample Toward a self-consistant H₀

Measure "L_{SN}"

Calibrator Sample

Volume limited ZTF-SNeIa < 50 Mpc

Technique TRGB (doable in any galaxy)

Statistics: ~5 per year (~30 by end of ZTF)

No selection function since both volume limited samples Unique photometric system, no absolute photometric calibration issue only relative, which is way easier

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Hubble Flow Sample

Volume limited ZTF-SNeIa z<0.06 Mpc

ZTF detects, follows and classifies *all* SNe Ia in the northern sky up to z~0.06

Statistics: Already >800 acquired

The Hubble Tension

Many more points (e.g. 2022)

- SNeIa->SNII: de Jaeger+2022 | 75±5%
- Geometry+Cepheids: <u>Kenworthy</u>+2022 | 73±4%
- BAO+BBN: Schöneberg+2022 | $68 \pm 0.5\%$
- ...



Sensitive to peculiar velocity correction





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