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The NEVFAR project:

New Evaluation of  $\nu$  Fluxes At Reactors

# Revisiting the summation calculation of reactor antineutrino spectra

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[International Conference on the Physics of the Two Infinities](#)

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<sup>(a)</sup>Now at ILANCE (CNRS/UTokyo), Japan

<sup>(b)</sup>Now at TUM, Germany

## 1. Introduction & motivations

- a. Reactors as antineutrino sources
- b. Experimental anomalies
- c. Modeling methods

## 2. Revised summation method

- a.  $\beta^-$  spectrum calculation
- b. Nuclear data content
- c. Uncertainty budget

## 3. Comparison to experiments and models

- a. Integral measurements
- b. Spectrum shape

## 4. Conclusion & perspectives

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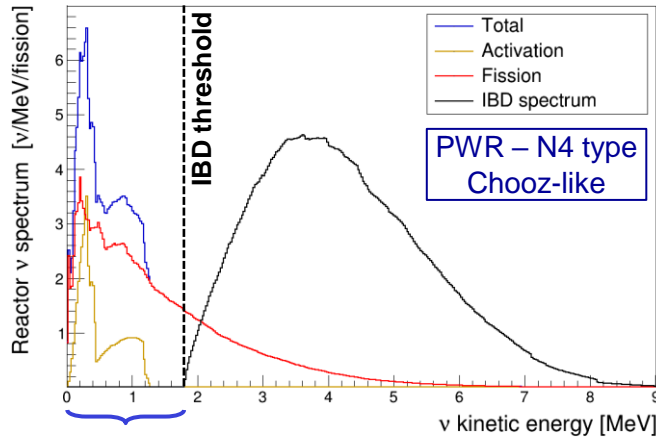
- a. Integral measurements
- b. Spectrum shape

## 4. Conclusion & perspectives

### PRESSURIZED WATER REACTOR (PWR)

- Fuel: lowly enriched uranium,  $^{238}\text{U}$  + 3-5%  $^{235}\text{U}$
- High power:  $\sim 3 - 4 \text{ GW}_{\text{th}}$
- Close reactor design & fuel contents for all PWR

⇒ Similar  $\bar{\nu}_e$  spectra



#### Typical fission fraction

$^{235}\text{U}$	0.559	} ~83% $\Phi_{\bar{\nu}_e}$
$^{239}\text{Pu}$	0.291	
$^{238}\text{U}$	0.088	
$^{241}\text{Pu}$	0.062	

#### Activation [ $\bar{\nu}_e$ /fission]

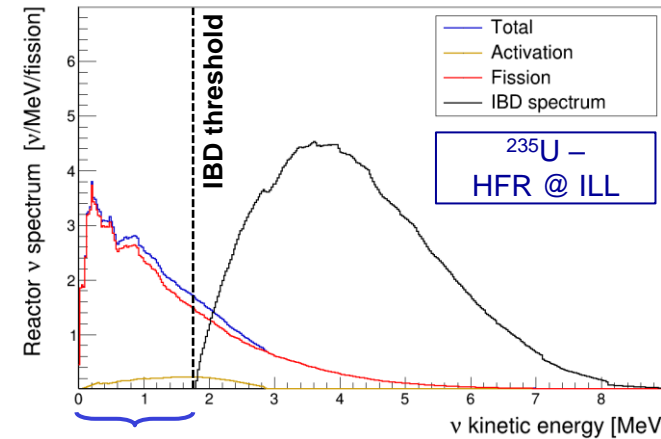
$^{239}\text{U}$	0.6	} ~17% $\Phi_{\bar{\nu}_e}$
$^{239}\text{Np}$	0.6	

Average event/fission for a Chooz type reactor ( $\sim 4\%$   $^{235}\text{U}$ ) over a 12-month core cycle

### RESEARCH REACTOR

- Fuel: highly enriched uranium,  $>20\%$   $^{235}\text{U}$
- Low power :  $\sim 0.1 \text{ kW}_{\text{th}} - 100 \text{ MW}_{\text{th}}$  but very short baseline accessible
- Wide array of designs & fuel contents + reactor specific structural material activation

⇒ Reactor-specific  $\bar{\nu}_e$  spectra



#### Fission fraction

$^{235}\text{U}$	0.993	} ~93% $\Phi_{\bar{\nu}_e}$
$^{239}\text{Pu}$	0.007	

#### Activation [ $\bar{\nu}_e$ /fission]

$^{239}\text{U}$	0.013	} ~7% $\Phi_{\bar{\nu}_e}$
$^{239}\text{Np}$	0.013	
$^{28}\text{Al}$	0.346	
$^{56}\text{Mn}$	0.058	

Average event/fission for the HFR at the ILL ( $\sim 93\%$   $^{235}\text{U}$ ) over a 50-day core cycle

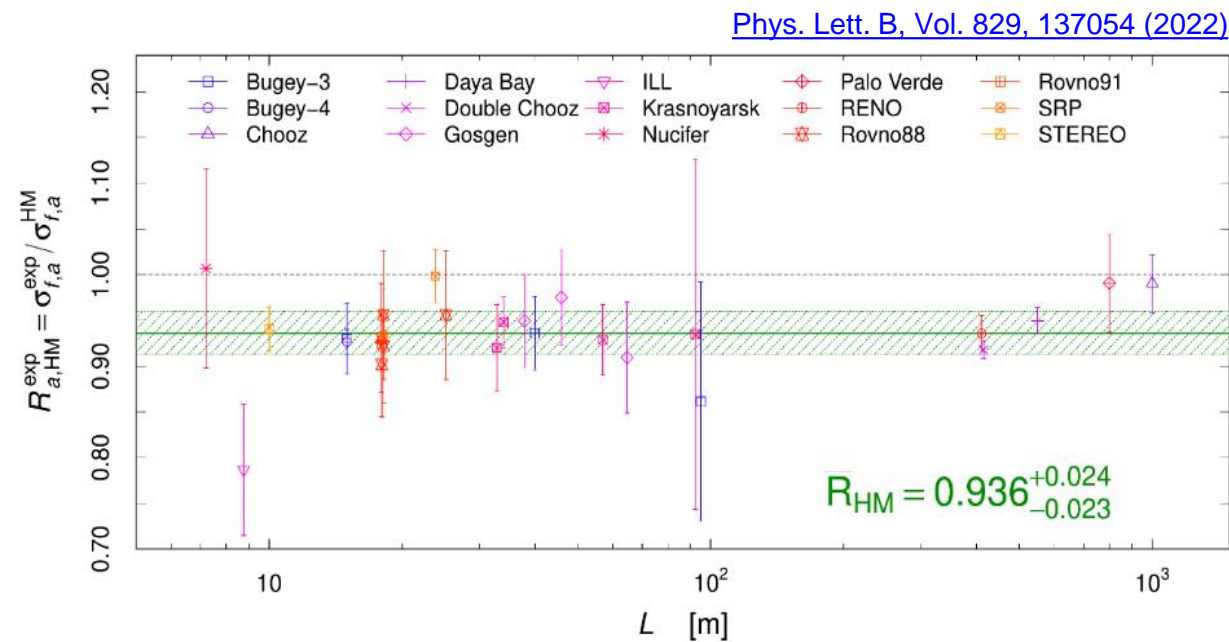
⇒  $\bar{\nu}_e$  contribution depends on reactor type and changes with time

⇒ Prediction needed for fission and activation spectra

## REACTOR ANTINEUTRINO ANOMALY (RAA)

[Phys. Rev. D 83, 073006 \(2011\)](#)

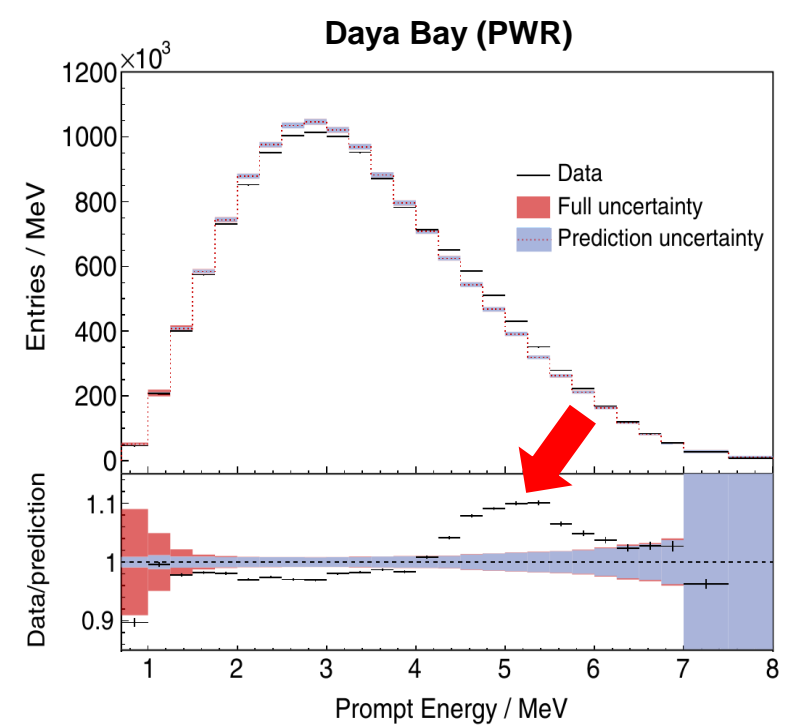
- Systematic measured **IBD rate deficit** compared to Huber-Mueller model
  - ▶ Observed in more than 20 reactor experiments
  - ▶ Confirmed in all recent reactor experiments at PWR and research reactors
- Ratio of measured over predicted IBD rate:  $0.936^{+0.024}_{-0.023}$ 
  - ▶ Statistical significance:  $2.5\sigma$
- RAA possible origins
  - ▶ Experimental bias *Unlikely*
  - ▶ New physics (sterile neutrino)
  - ▶ **Mismodeling / underestimation of  $\bar{\nu}_e$  spectrum uncertainty**



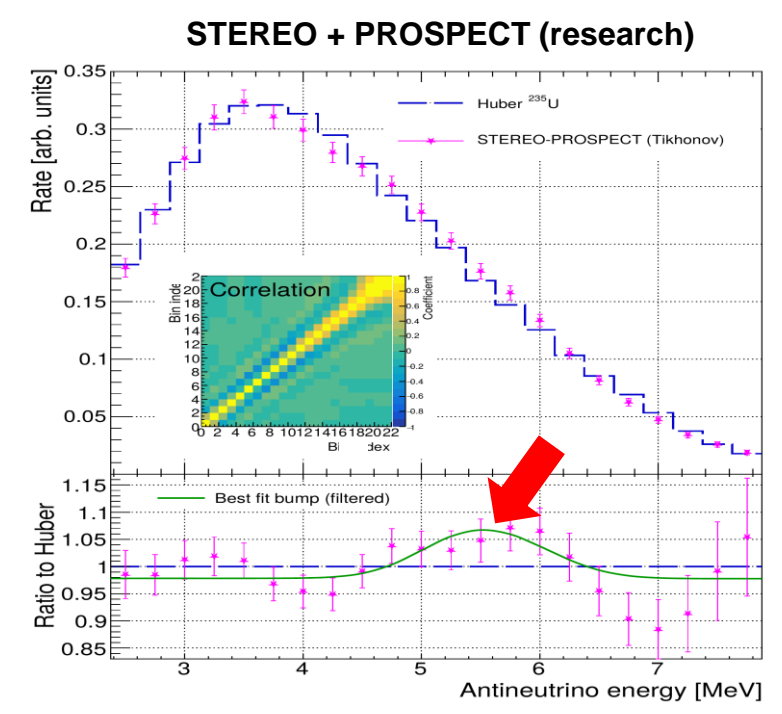
### SHAPE ANOMALY

- ~8-10% difference between predictions and data at 5-6 MeV
- First observed by Double Chooz, Daya Bay and RENO
  - ▷ Confirmed by recent very-short baseline reactor experiments (NEOS, STEREO, PROSPECT, DANSS)

- Possible origins
  - ▷ Detector energy scale calibration Checked
  - ▷ Fuel composition
  - ▷ **Prediction issue, single / multiple actinide(s) ?**



[Phys. Rev. Lett. 123, 111801 \(2019\)](https://arxiv.org/abs/1808.07409)



[arXiv:2107.03371](https://arxiv.org/abs/2107.03371)

## CONVERSION METHOD

- Measure **exp.  $\beta$  fission spectra**
- Convert virtual  $\beta$  branch fit to  $\bar{\nu}_e$  branches



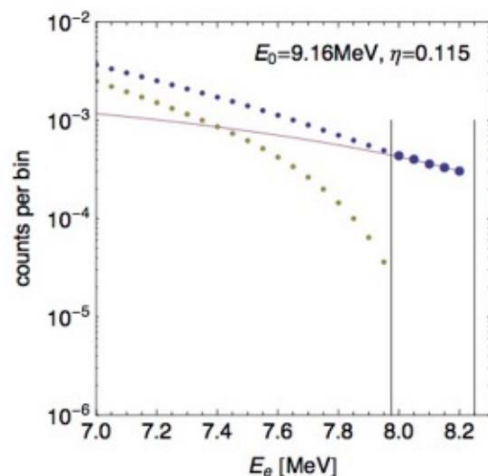
### PROS

- Small uncertainties ~2-3%
- Access total  $\bar{\nu}_e$  fission spectrum



### CONS

- Limited to exp. range, 2-8 MeV
- No activation spectrum
- **HM subject to the anomalies**
- BILL data questioned → Kurchatov Inst.
- Impact of forbidden branches on fit



## SUMMATION METHOD

- Fission spectrum prediction = **sum of all  $\beta$  branches** listed in nuclear databases
- +900  $\beta^-$  emitters ~ 10 000  $\beta^-$  transitions



### PROS

- Prediction  $\forall$  energy,  $\forall$   $\beta$  emitter
  - ▶ CEvNS
- Convenient to understand physics
- **Mandatory for activation spectra**



### CONS

- Uncomplete/biased nuclear database
- Modeling approximations
- **Uncertainties very complex to estimate**

⇒ **Summation method requires many input data**

*W. Bühring and H. Behrens formalism (1982)*

### Branch spectrum

$$S_\beta = KpW(W_0 - W)^2 F_0 C(1 + \delta_{WM} + \delta_{RC})$$

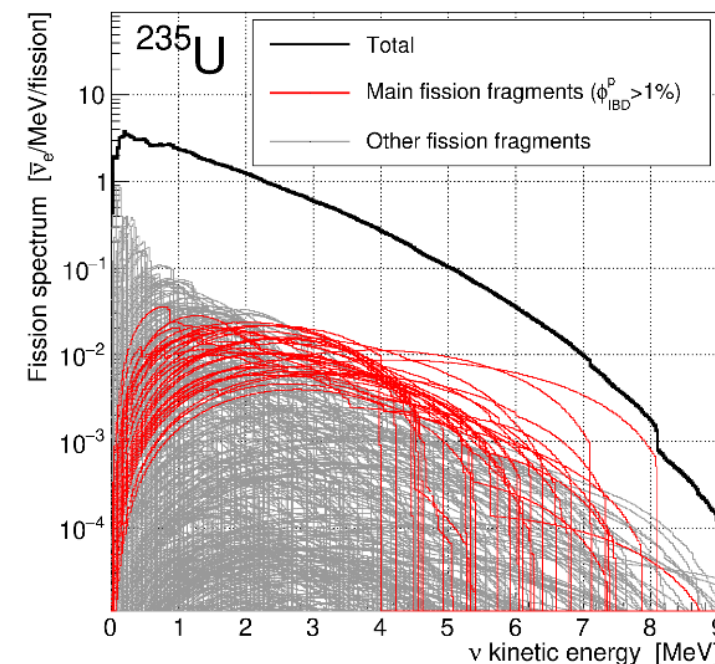
$$S_\nu(E) = S_\beta(E_0 - E)$$

### Isotope spectrum

$$S_i = \sum_b BR_i^b S_\nu^b$$

### Fission spectrum

$$S_f = \sum_i FY_f^i S_i$$



## CONVERSION METHOD

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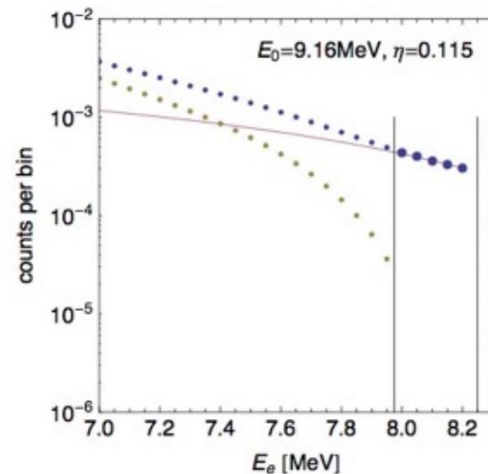
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### CONS

- Uncomplete/biased nuclear database
- Modeling approximations
- **Uncertainties very complex to estimate**

$\Rightarrow$  **Reliable summation method required for multiple purposes**

## THE NE $\nu$ FAR PROJECT

(New Evaluation of  $\nu$  Fluxes At Reactor)



Revise summation method with BESTIOLE code

- Improve  $\beta$ -decay modeling
  - ▶ Refine **non-unique forbidden transition** modeling
- Impact of database incompleteness and quality
  - ▶ Update nuclear database with **Pandemonium-free data**
  - ▶ Impact of **nuclides with no data** modeling
- Build a **comprehensive uncertainty** budget
  - ▶ Nuclear data and modeling uncertainties



**All results are preliminary**



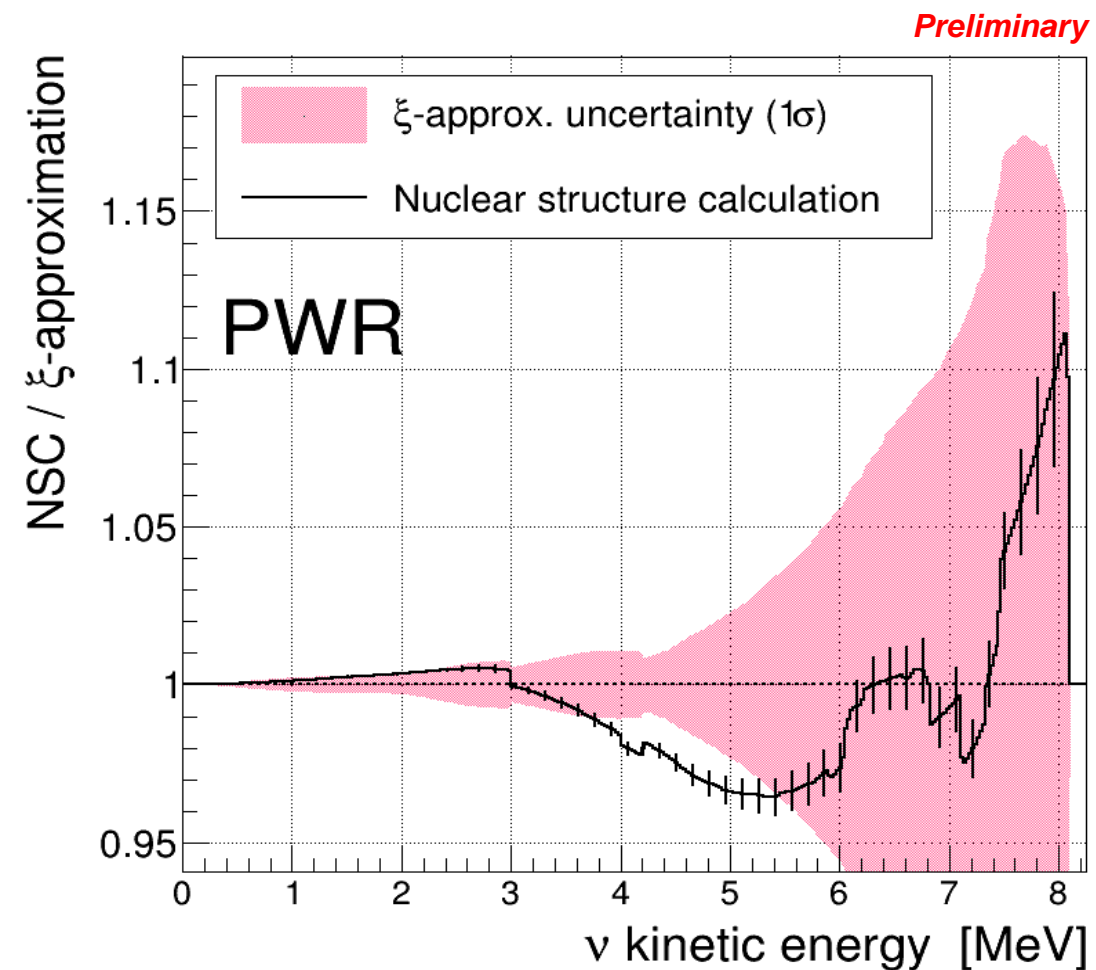
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## MODELING OF NON-UNIQUE TRANSITIONS

- Dismissed in previous modeling (modeled as allowed or unique forbidden)
- Hayes *et al.* (2014) + Hayen *et al.* (2019): modelings of non-unique transitions in conversion predictions → partial explanation of shape anomaly
- Nuclear structure calculation with NuShellX
  - Very time consuming (man & cpu)
  - No general nor systematic trend
- 23 non-unique forbidden transitions contribute to
 

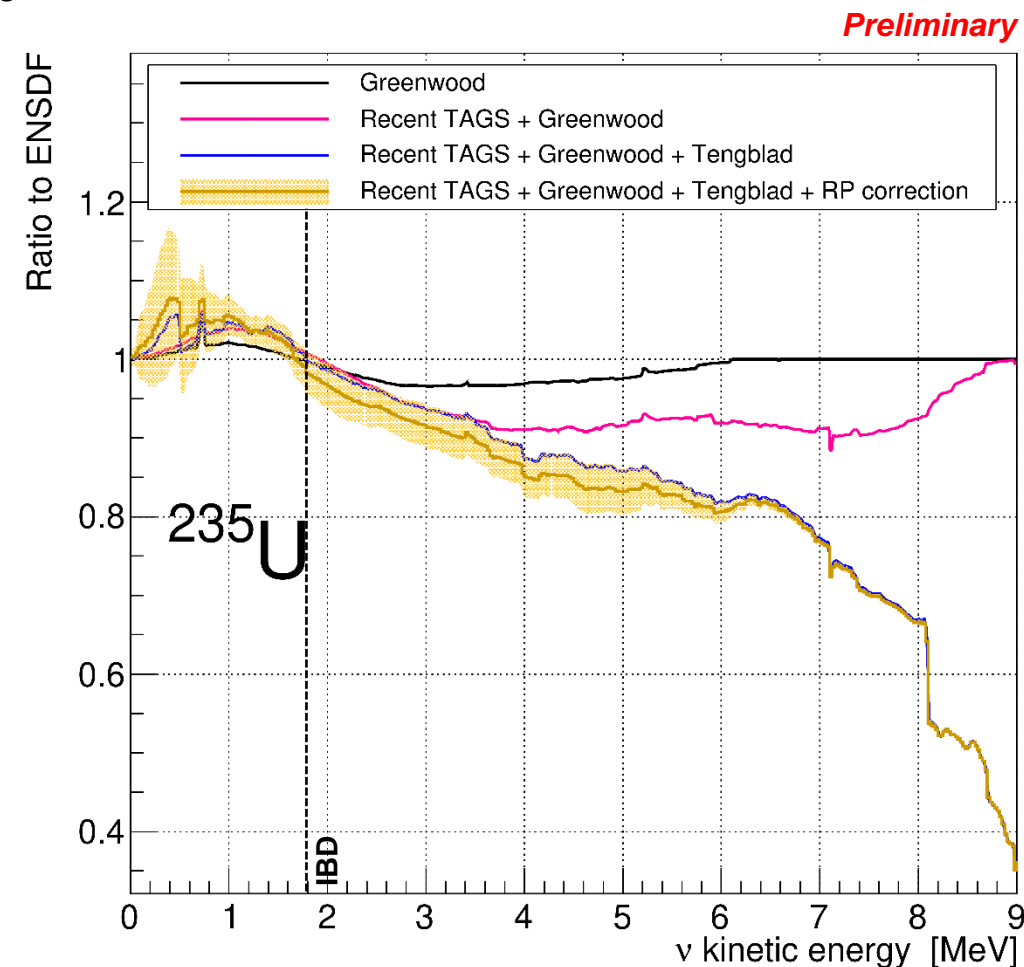
~27% of IBD yield
~22% of CE $\nu$ NS yield

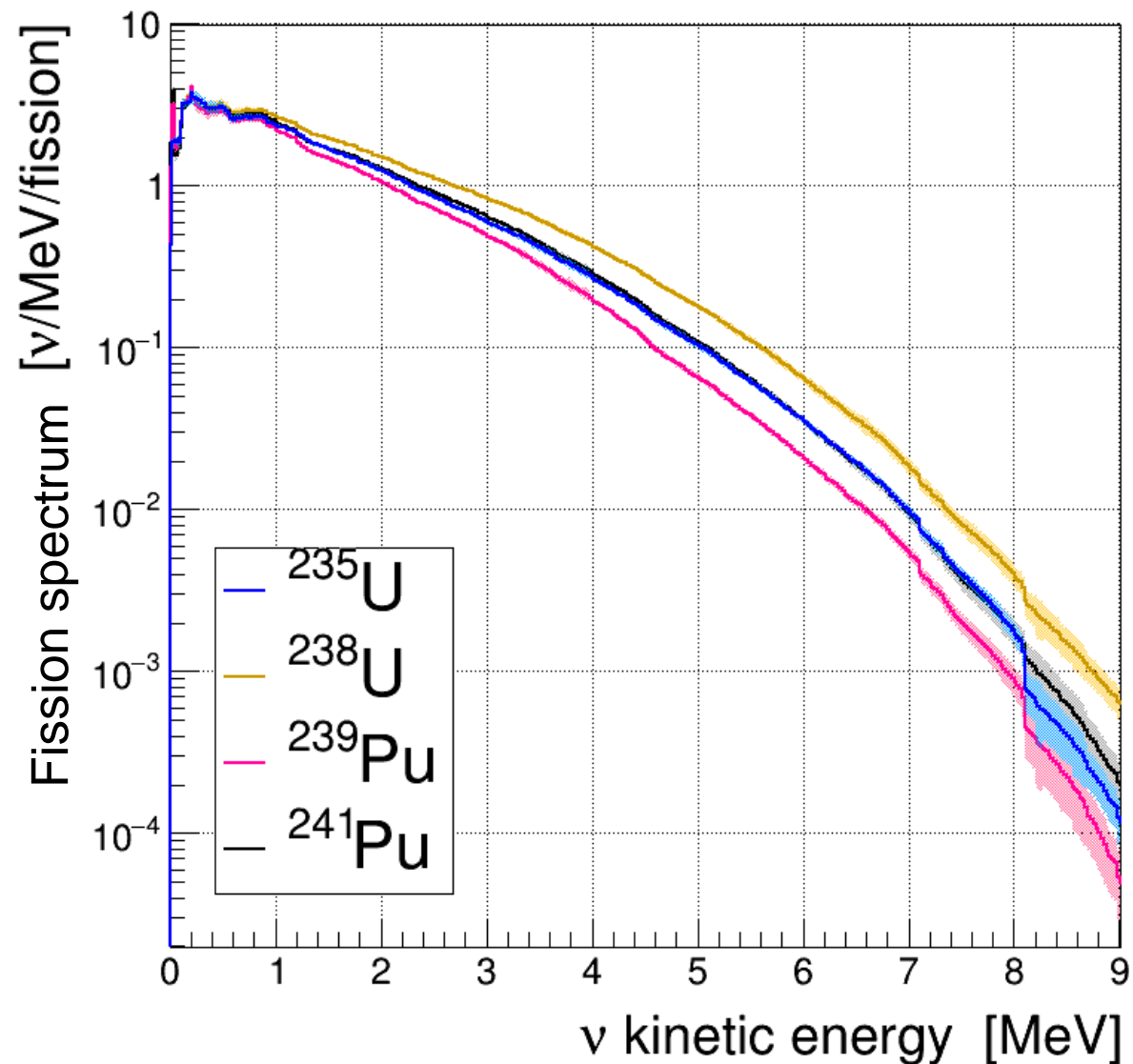
⇒ Using NSC decreases IBD yield by  $(1.3 \pm 0.2)\%$



## TACKLING THE PANDEMONIUM EFFECT IN SUMMATION SPECTRA

- HPGe detector, high energy resolution + decreasing efficiency for increasing energies
  - $\beta$  feedings to low (high) energy levels are overestimated (underestimated)
- Nuclear database are biased by the Pandemonium effect
  - Estienne et al. (2019): including Pandemonium-free TAGS data decreases IBD yields and shape differences
- Including up-to-date Pandemonium-free data (TAGS + Direct  $\beta$  measurements)
  - ⇒ **IBD yield decreased by  $(12.8 \pm 1.5) \%$**
  - ⇒ **~ 65% of IBD and CE $\nu$ NS yields**
- Remaining isotopes potentially impacted by Pandemonium in nuclear database
  - 29 isotopes identified by IAEA
  - Apply correction for residual Pandemonium effect
    - ⇒ **IBD yield decreased by  $(2.2 \pm 2.4) \%$**
    - ⇒ **~ 12% of IBD and CE $\nu$ NS yields**





### IBD yields ( $10^{-43}$ cm<sup>2</sup>/fission)

$^{235}\text{U}$ :	$6.25 \pm 0.21$
$^{238}\text{U}$ :	$10.01 \pm 0.32$
$^{239}\text{Pu}$ :	$4.48 \pm 0.15$
$^{241}\text{Pu}$ :	$6.58 \pm 0.21$

⇒ IBD yield uncertainty ~3%

### CEνNS yields\* ( $10^{-43}$ cm<sup>2</sup>/fission)

$^{235}\text{U}$ :	$1113 \pm 34$
$^{238}\text{U}$ :	$1669 \pm 48$
$^{239}\text{Pu}$ :	$882 \pm 25$
$^{241}\text{Pu}$ :	$1169 \pm 33$

\* For a Ge target nucleus and 20 eV detector threshold

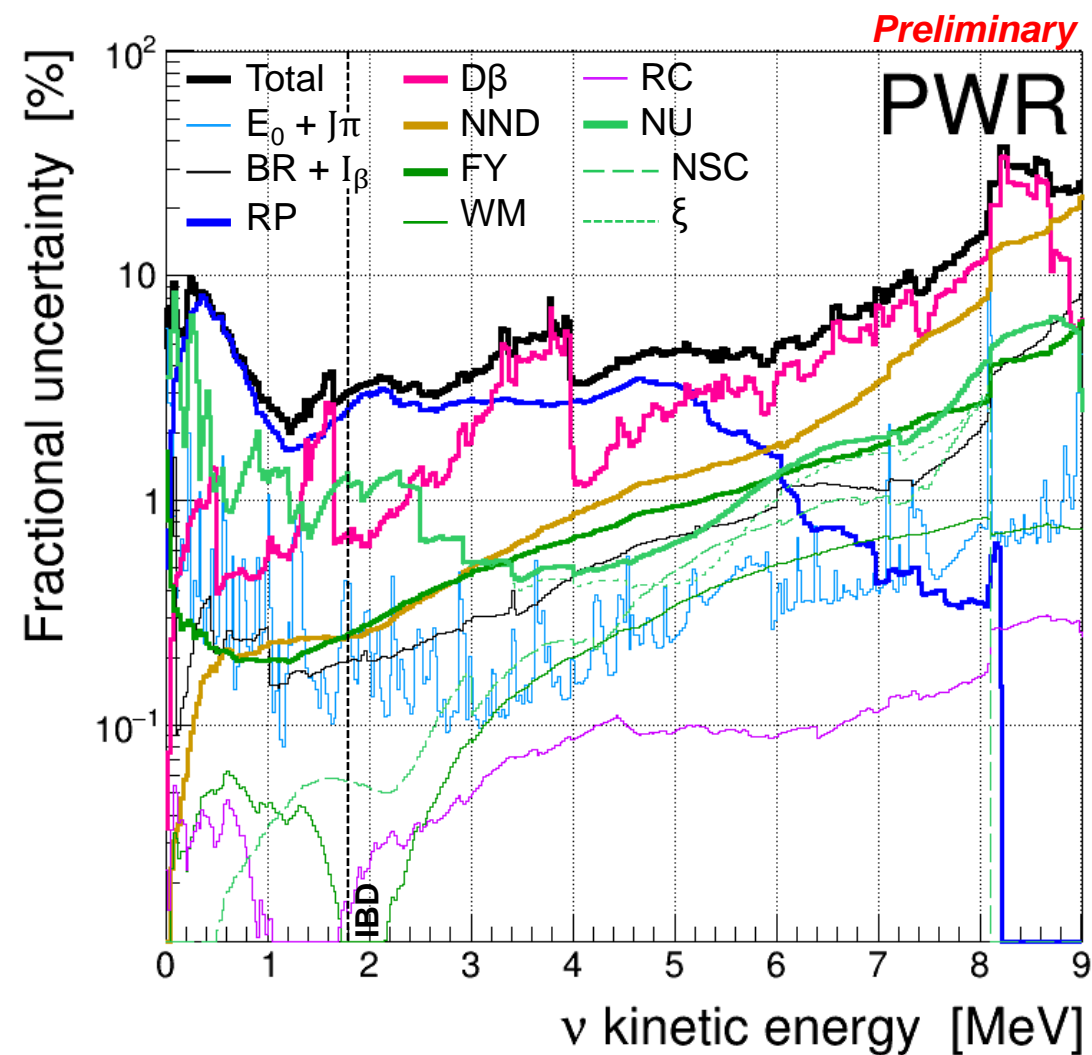
⇒ CEνNS yield uncertainty ~3%

### NORMALIZATION UNCERTAINTY

PWR			$\langle\sigma_{IBD}\rangle$	$\langle\sigma_{CEvNS}\rangle^*$	
[ $10^{-43}$ cm <sup>2</sup> /fission]			6.08	1090	
Uncertainty	Abbrev.	Method	[%]	[%]	
DATA	Endpoint + Spin-parity	$E_0 + J\pi$	MC	0.1	0.1
	Branching ratio + $\beta^-$ intensity	$BR + I_\beta$	MC + Analytic	0.4	0.3
	Residual Pandemonium	RP	Analytic	2.5	2.4
	Direct $\beta$ measurement	$D\beta$	Analytic	1.5	1.2
	Nuclides with no data	NND	Pool modeling	0.8	0.5
	Fission yield	FY	Analytic	~0.7	~0.6
	Fission fraction		Analytic	~0.7	~0.7
MODELING	Weak magnetism	WM	Model comparison	0.3	0.2
	Radiative corrections	RC	Model comparison	0.1	0.1
	Non-unique transitions	NU	Model comparison	0.4	0.4
	• Nuclear struct. calcul.	NSC		0.2	0.1
	• $\xi$ -approximation	$\xi$		0.3	0.3
	Cross-section		Analytic	0.1	0.5
<b>TOTAL</b>			<b>3.1</b>	<b>2.9</b>	

\* For a Ge target nucleus and 20 eV detector threshold

### FRACTIONAL UNCERTAINTY



⇒ Uncertainty budget dominated by RP and  $D\beta$

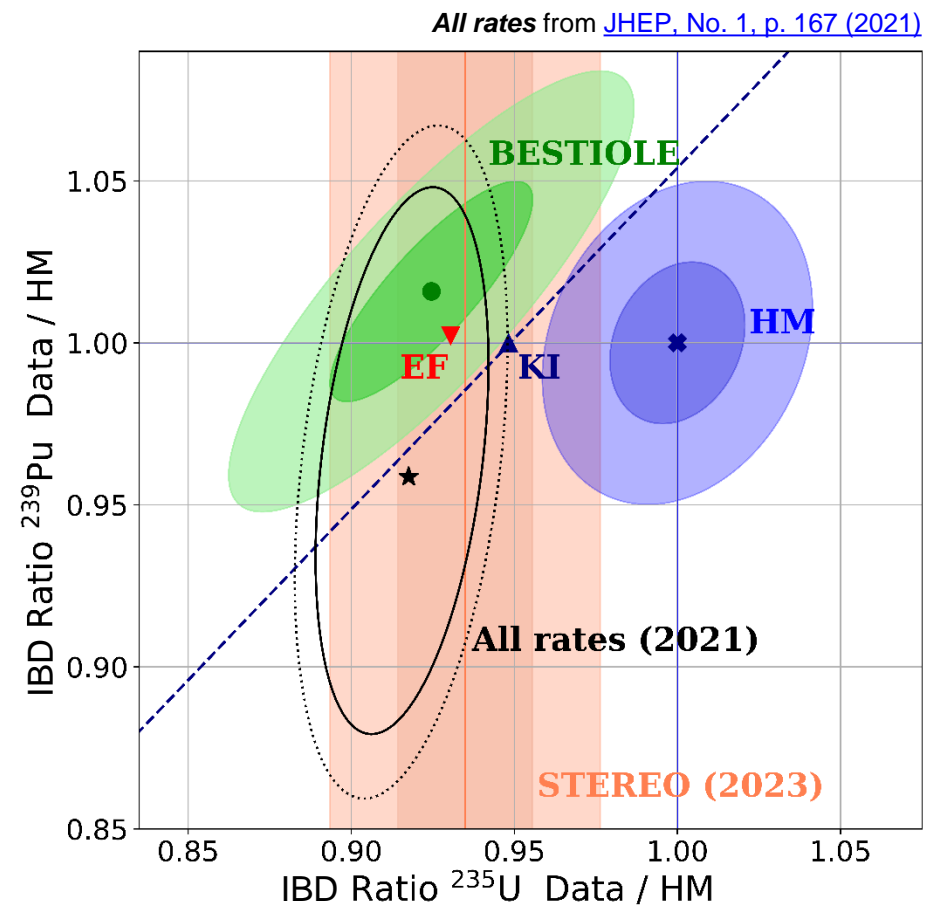
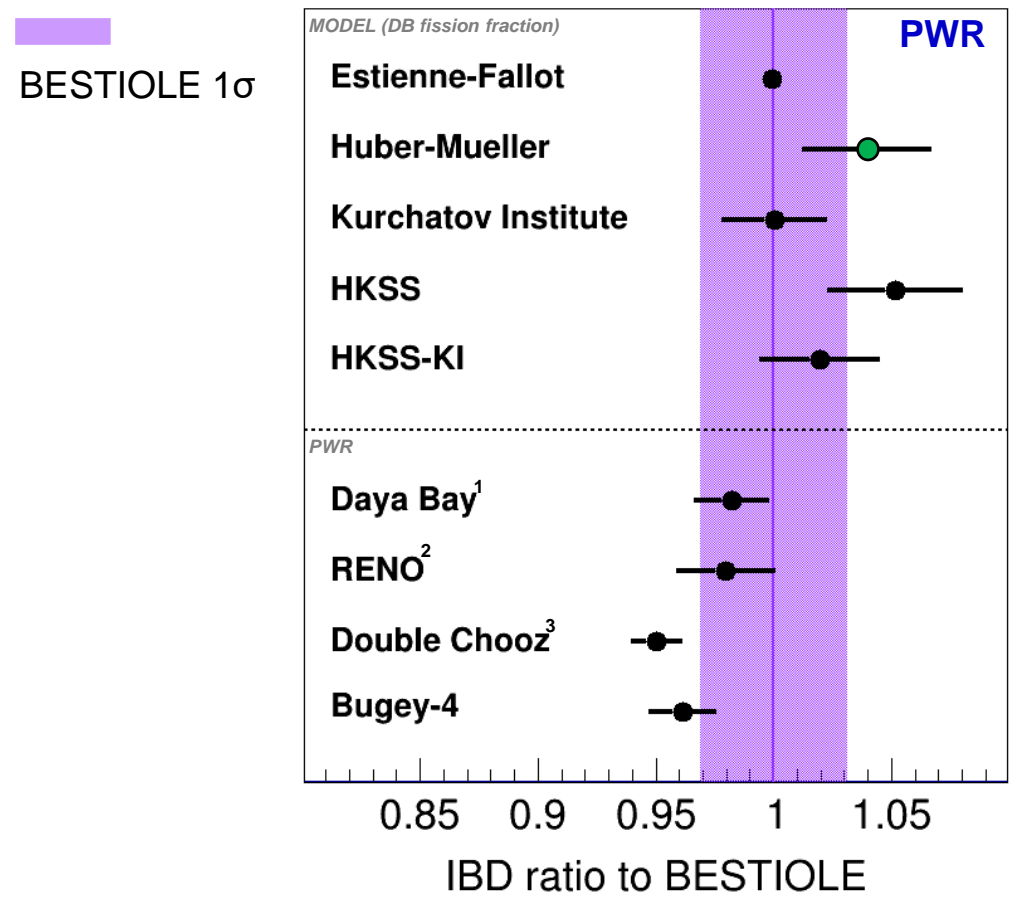
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### 3. Comparison to experiments and models

#### a. Integral measurements

*All plots are preliminary*

Predictions and Bugey-4 taken from [Giunti et al., Phys. Lett. B, 829, 137054 \(2022\)](#)  
 1: [PRL 123, 111801 \(2019\)](#) 2: [PRD 104, L111301 \(2021\)](#) 3: [PRL 125, 201801 \(2020\)](#)



**DB / BESTIOLE =  $0.982 \pm 0.015$  (exp)  $\pm 0.031$  (model)**

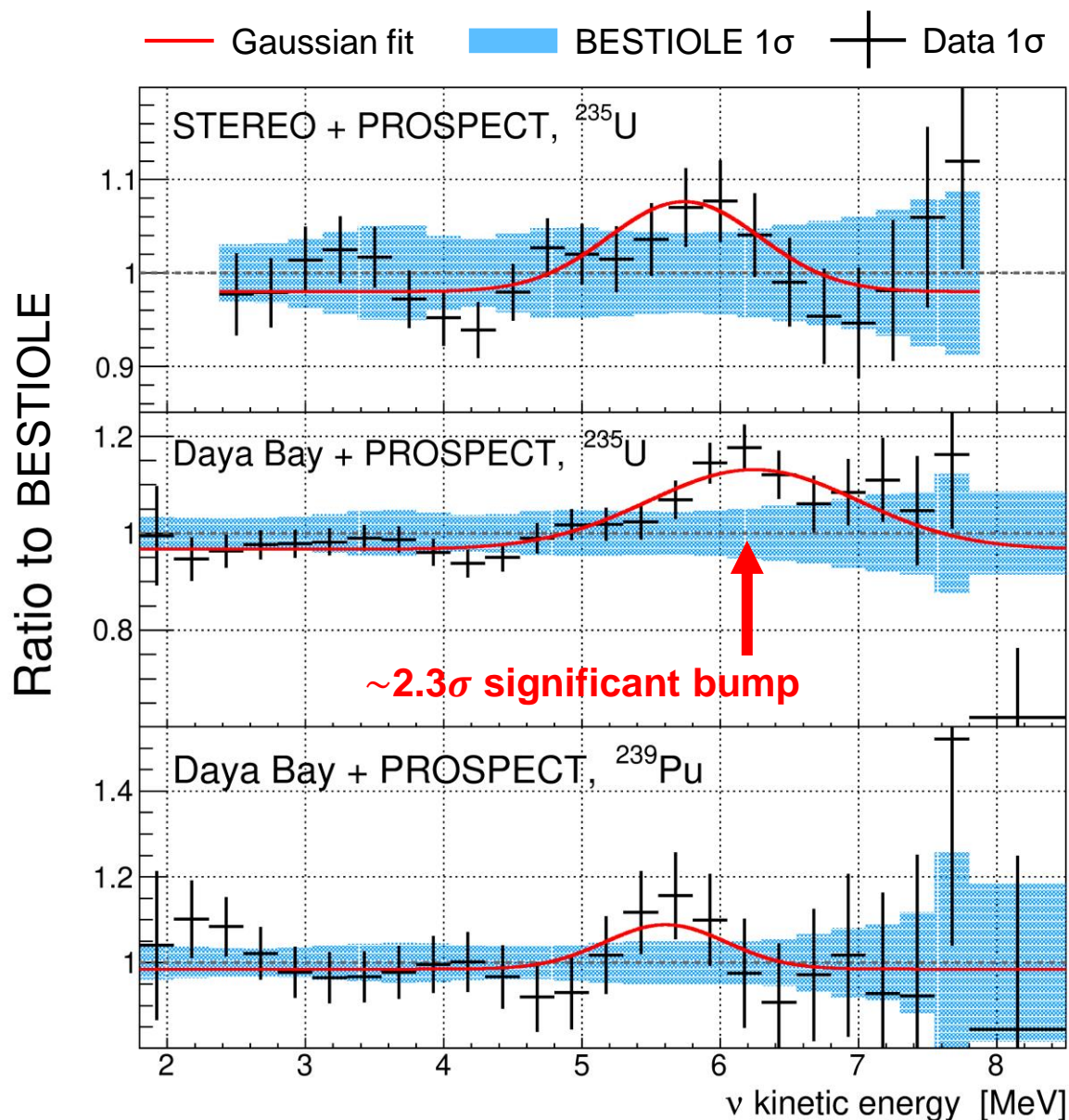
**DB / HM =  $0.945 \pm 0.014$  (exp)  $\pm 0.024$  (model)**

**⇒ Significance at 0.5σ for BESTIOLE and 1.9σ for HM**

**⇒ BESTIOLE consistent within ~2σ with recent data**

**⇒ Discrepancy with HM favors RAA caused by <sup>235</sup>U HM flux**

## RATIO OF IBD SPECTRA



- Shape only comparison, predictions normalized to data
- Gaussian distortion not significantly favored in 5-7 MeV
  - Gaussian bump hypothesis favored by  $\leq 2.3\sigma$

**$\Rightarrow$  Good agreement with experimental IBD spectra within uncertainty**



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### KEY POINTS OF BESTIOLE SUMMATION PREDICTION

#### All modeling impacts considered and quantified

- **Nuclear structure calculation** for 23 non-unique branches
  - ▶ IBD yield **decreased by  $(1.3 \pm 0.2)\%$**

#### Quality of data checked for all data sources

- Correction for Residual Pandemonium
  - ▶ IBD yield **decreased by  $(2.2 \pm 2.4)\%$**
  - ▶ Measurement needed to validate RP correction

#### Comprehensive uncertainty budget

- Uncertainty budget of summation model for the **first time ever**

#### Complete revision of summation method

- **Good overall agreement with data**
- Results favors RAA caused by  $^{235}\text{U}$  HM flux

**Final IBD and  $\text{CE}_{\nu\text{NS}}$  yield  
uncertainty budget  $\sim 3\%$**

**Led by RP correction**

**$\Rightarrow$  more Pandemonium-free data needed**

**$\Rightarrow$  Article coming soon with  
supplementary materials**

# BACKUP SLIDES

## ALLOWED TRANSITION SPECTRUM CORRECTIONS

Effect	Magnitude
Phase space factor Traditional Fermi function	Unity or larger
Finite size of the nucleus Radiative corrections Shape factor Atomic exchange Atomic mismatch	$10^{-1}-10^{-2}$
Atomic screening Shakeup Shakeoff Isovector correction Distorted Coulomb potential due to recoil Diffuse nuclear surface Nuclear deformation Recoiling nucleus Molecular screening Molecular exchange	$10^{-3}-10^{-4}$
Bound state $\beta$ decay Neutrino mass Forbidden decays	Smaller than $1 \times 10^{-4}$

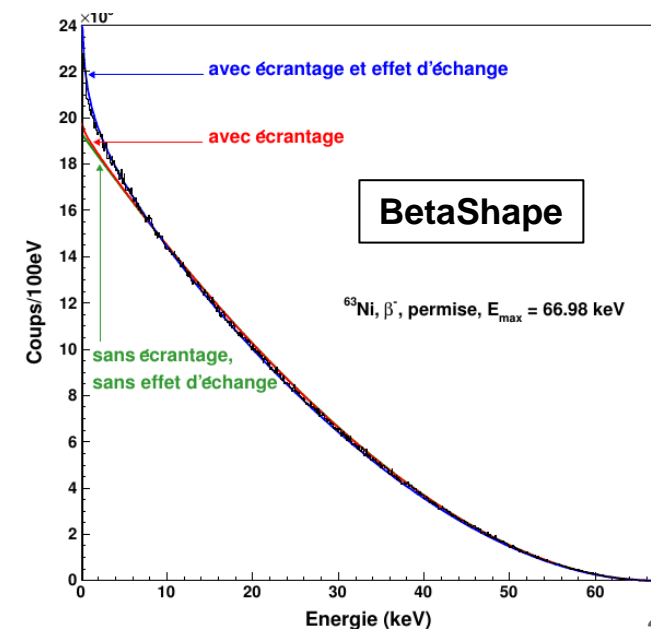
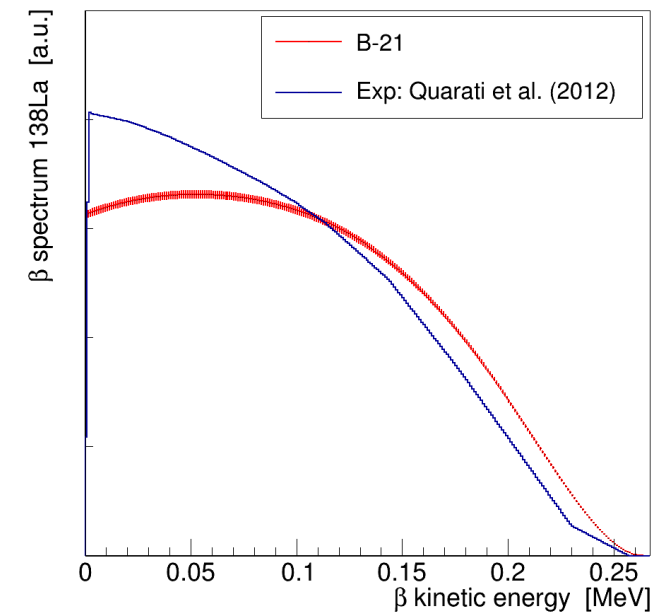
Electron emitted into a bound state

- Rearrange the atomic electrons
- Impact the  $\beta$  spectrum at low energy

Different Hamiltonian for initial and final nuclear states

- Impact the  $\beta$  spectrum near the endpoint energy
- Not expected to be significant for  $\nu$  spectrum

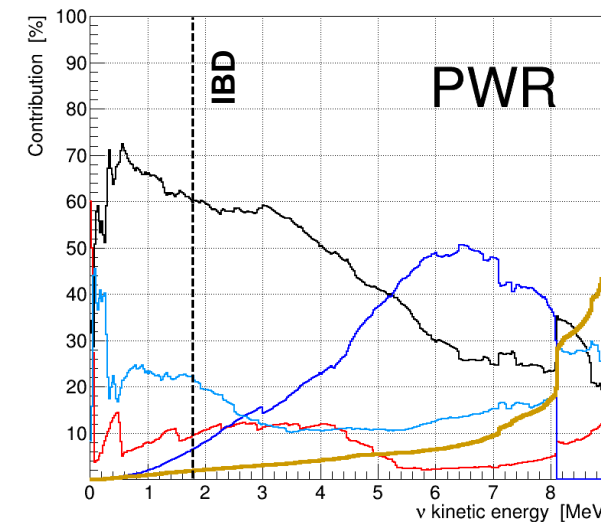
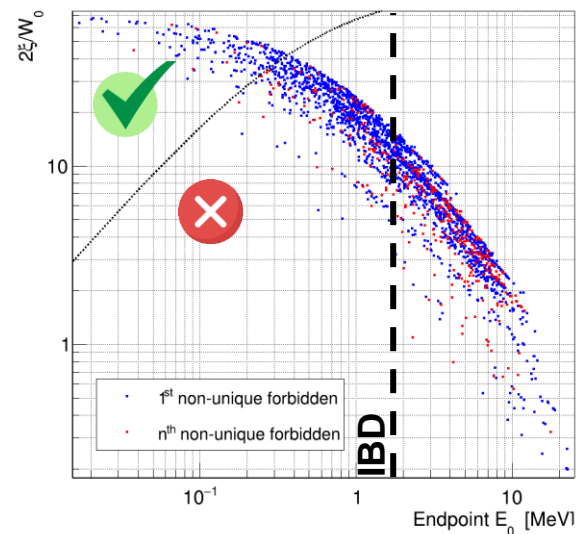
<sup>138</sup>La, 2<sup>nd</sup> unique forbidden



# NON-UNIQUE FORBIDDEN TRANSITION MODELING WITH NUCLEAR STRUCTURE CALCULATION

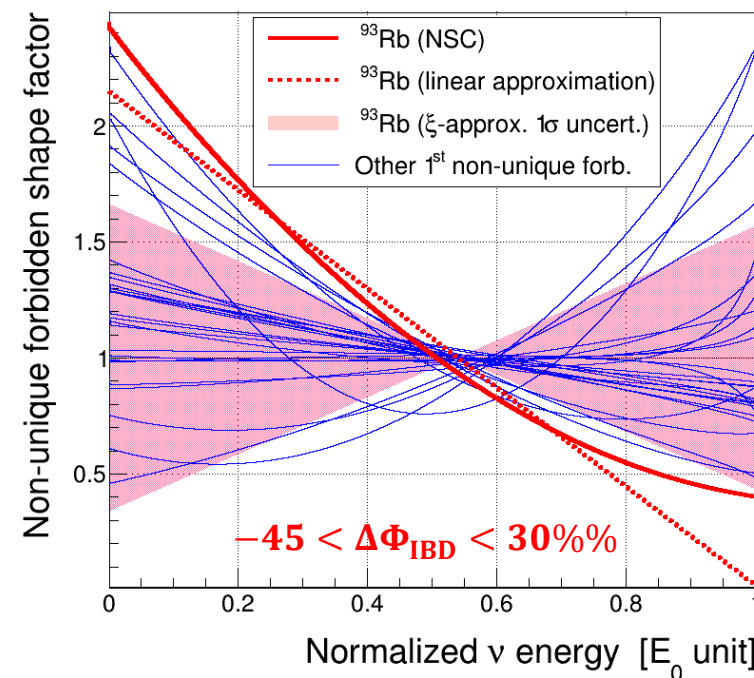
$$\langle \xi_f J_f || T_\lambda || \xi_i J_i \rangle = \hat{\lambda}^{-1} \sum_{a,b} \langle a || T_\lambda || b \rangle \langle \xi_f J_f || [c_a^\dagger \tilde{c}_b]_\lambda || \xi_i J_i \rangle$$

transition matrix element      tensor rank      single particle matrix element      one-body transition density



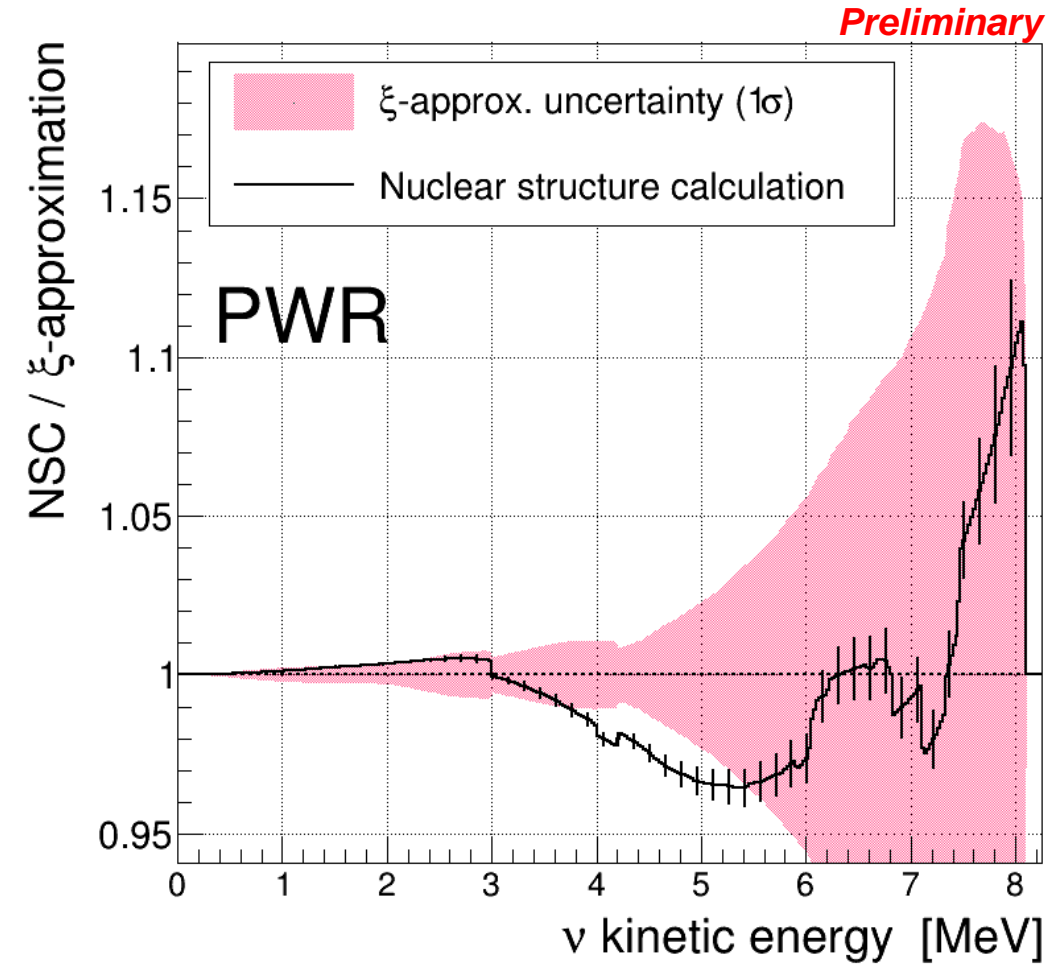
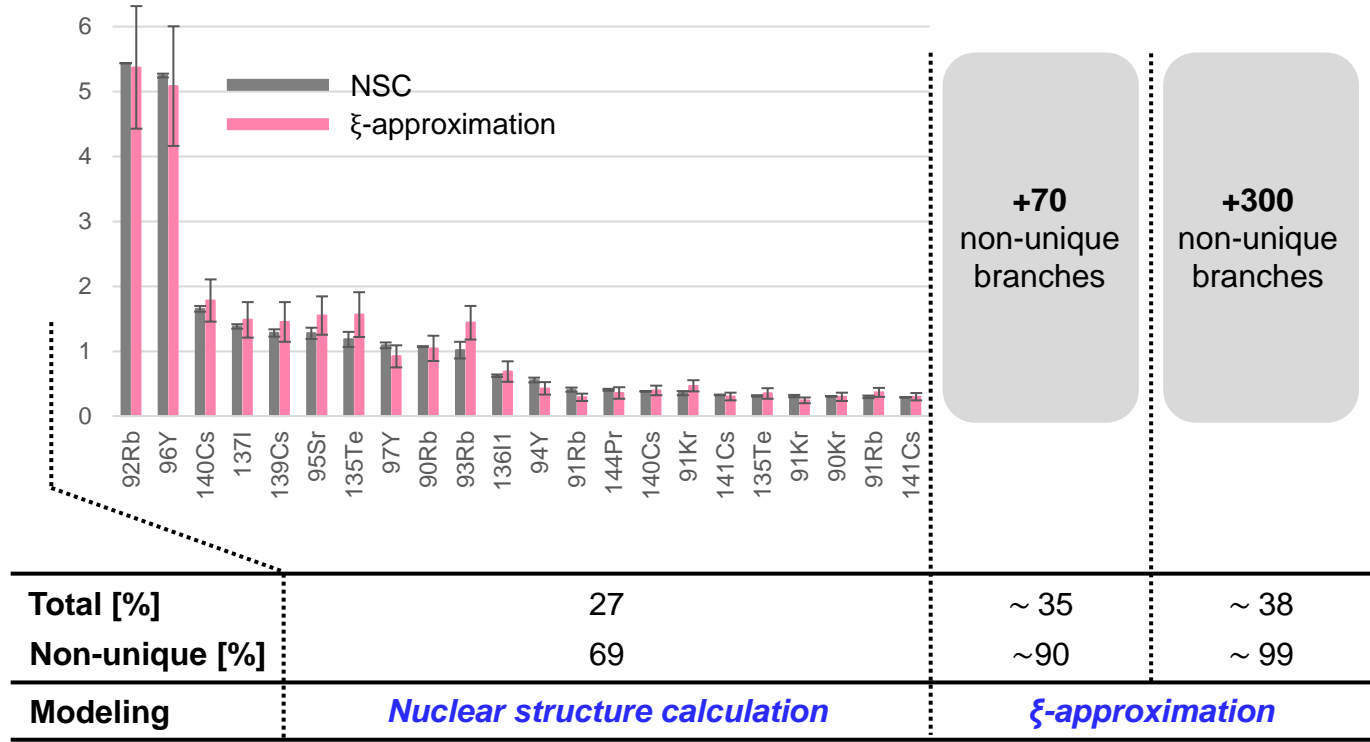
## Advanced modeling:

1. Compute transition matrix element
  - One-body transition density computed with nuclear shell model
  - Sum over the superposition of all possible nucleon states
2. Use of Conserved Vector Current (CVC) hypothesis to determine the vector relativistic matrix elements
  - Coulomb displacement energy includes mismatch between initial and final nucleon wave functions
3. Free-nucleon value of the axial-vector coupling constant  $g_A$  assumed
4. Full numerical lepton current considered, without any expansion of lepton wave functions
5. Control of nuclear current expansion: all dominant contributions included over several orders of magnitude



**NUCLEAR STRUCTURE CALCULATION (NSC)**

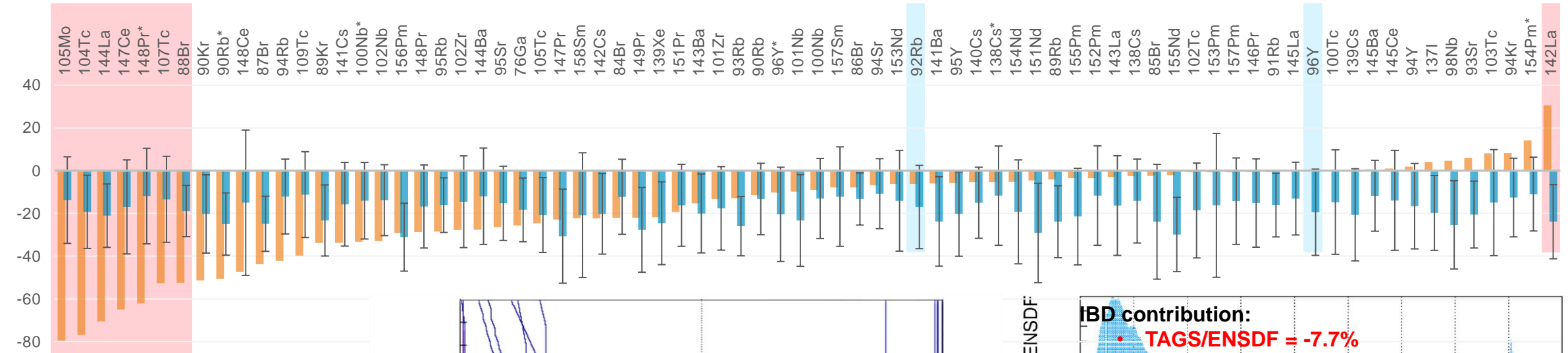
**IBD contributions of non-unique transitions for a Chooz-like PWR**



**⇒ IBD yield decreased by  $(1.3 \pm 0.2)\%$  compared to full  $\xi$ -approximation**

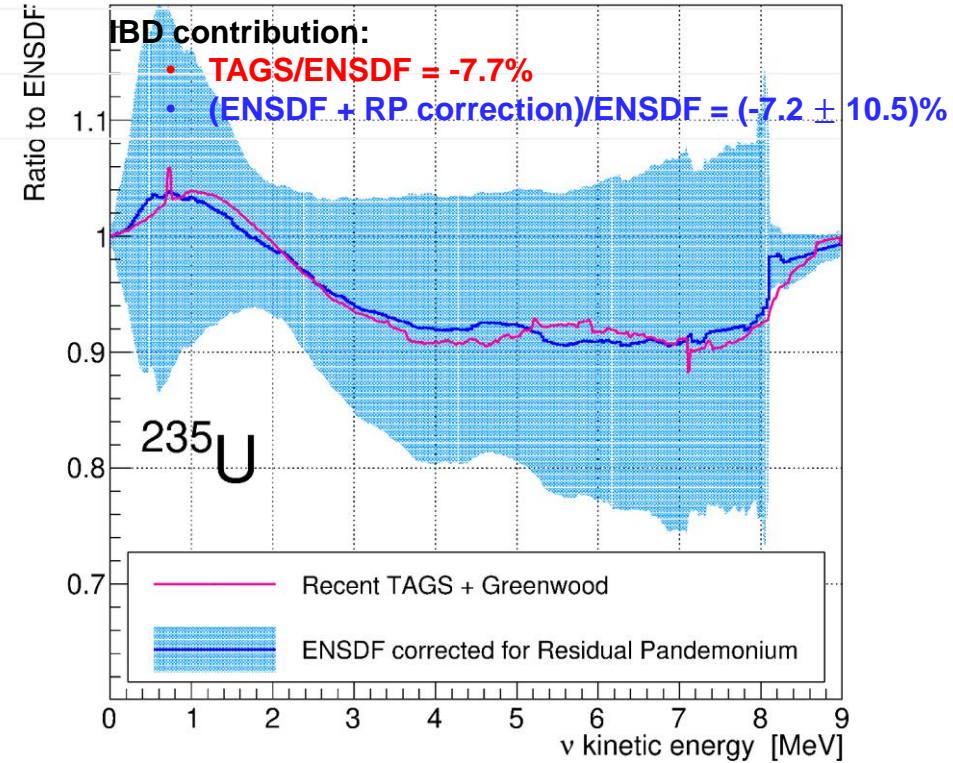
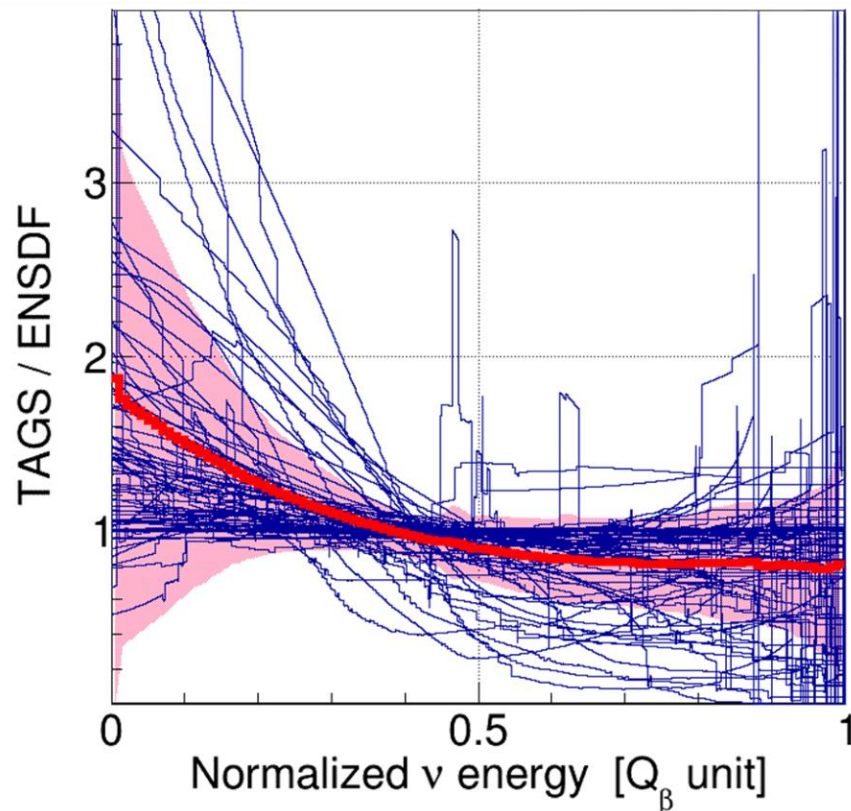
# RESIDUAL PANDEMONIUM EFFECT

Diff > 2σ



TAGS/ENSDF - 1

Correction/ENSDF - 1



## RESIDUAL PANDEMONIUM CORRECTION FOR A PWR

- Apply RP correction on the 29 isotopes from WPEC-25
- IBD yield decreased by  $(2.2 \pm 2.4)\%$  with RP correction

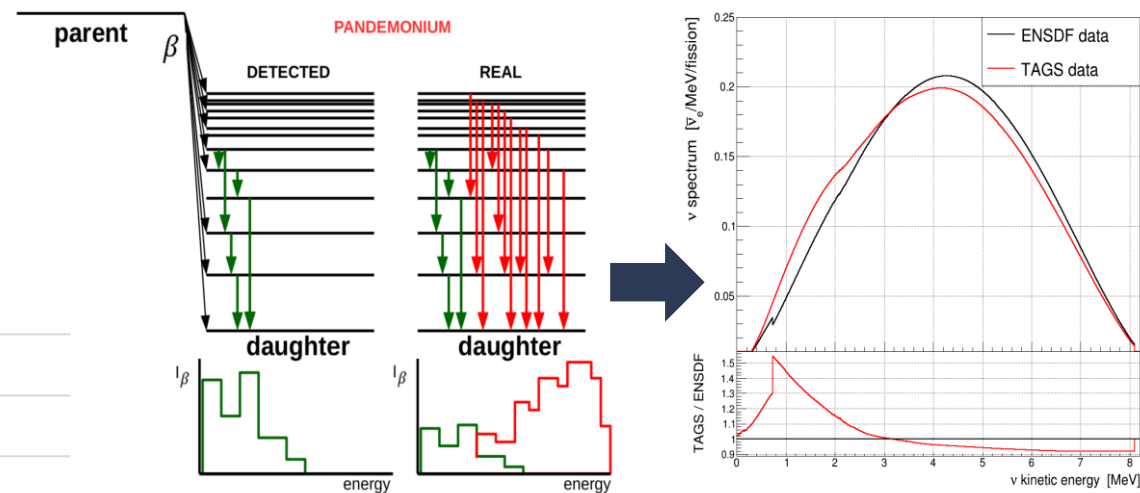
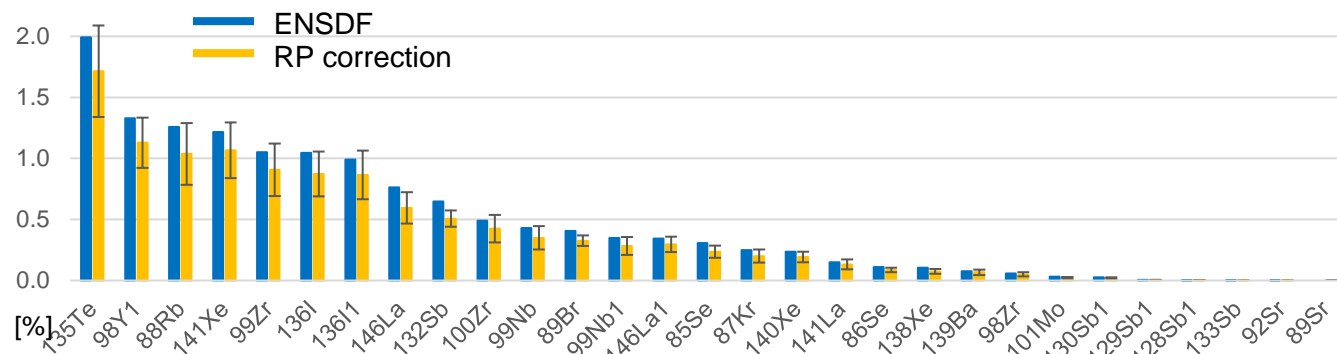
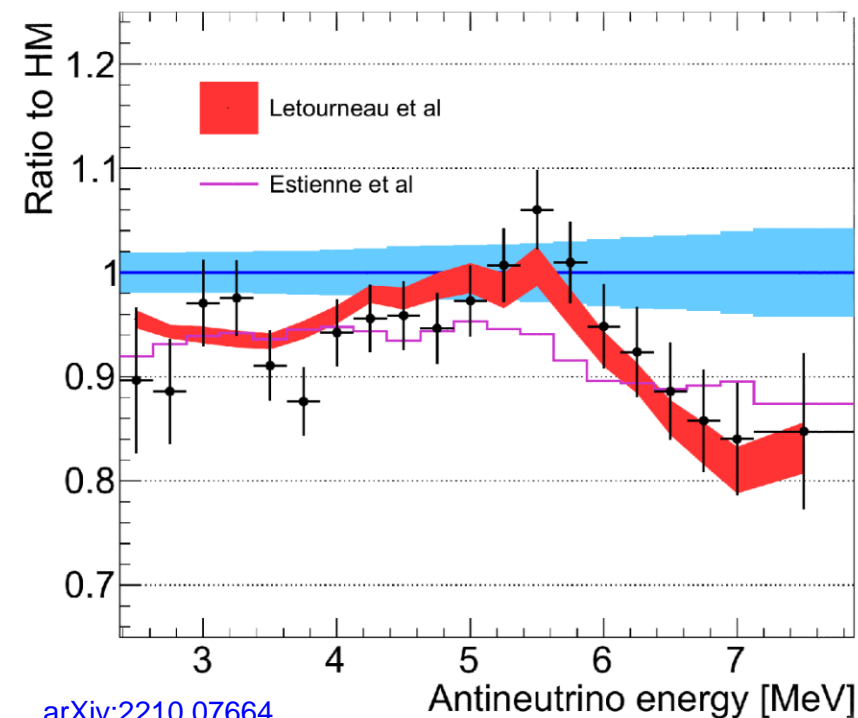


Illustration of the Pandemonium effect impact on the  $\bar{\nu}_e$  spectrum of <sup>92</sup>Rb

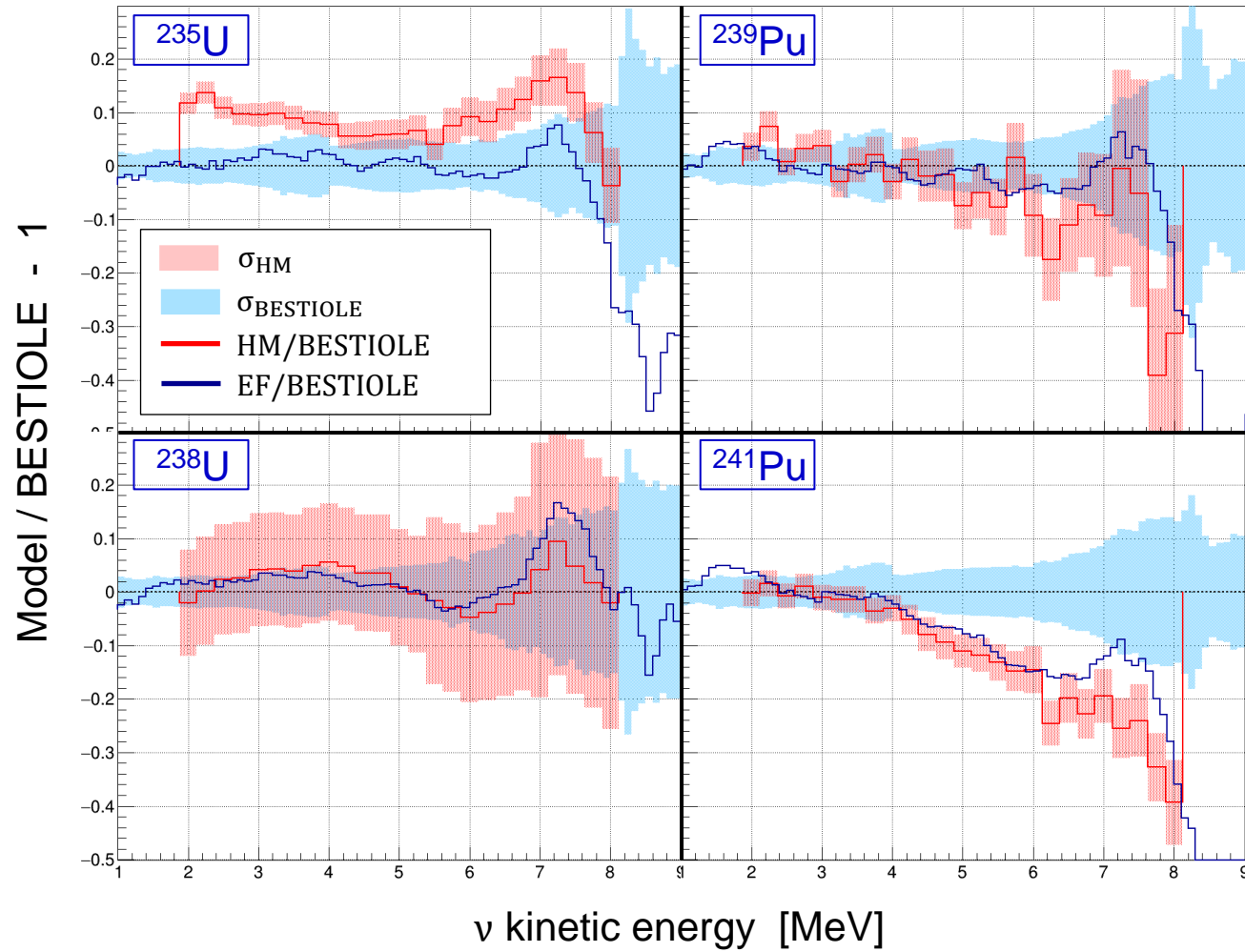
## PHENOMENOLOGICAL DECAY STRENGTH MODEL

- 1 parameter Gamow-Teller strength model
- Generate missing transitions for all fission products from ENSDF
  - ▶ Correct Pandemonium effect and missing transitions
- [Letourneau et al., PRL 130, 021801 \(2023\)](#)
- Very good agreement with STEREO



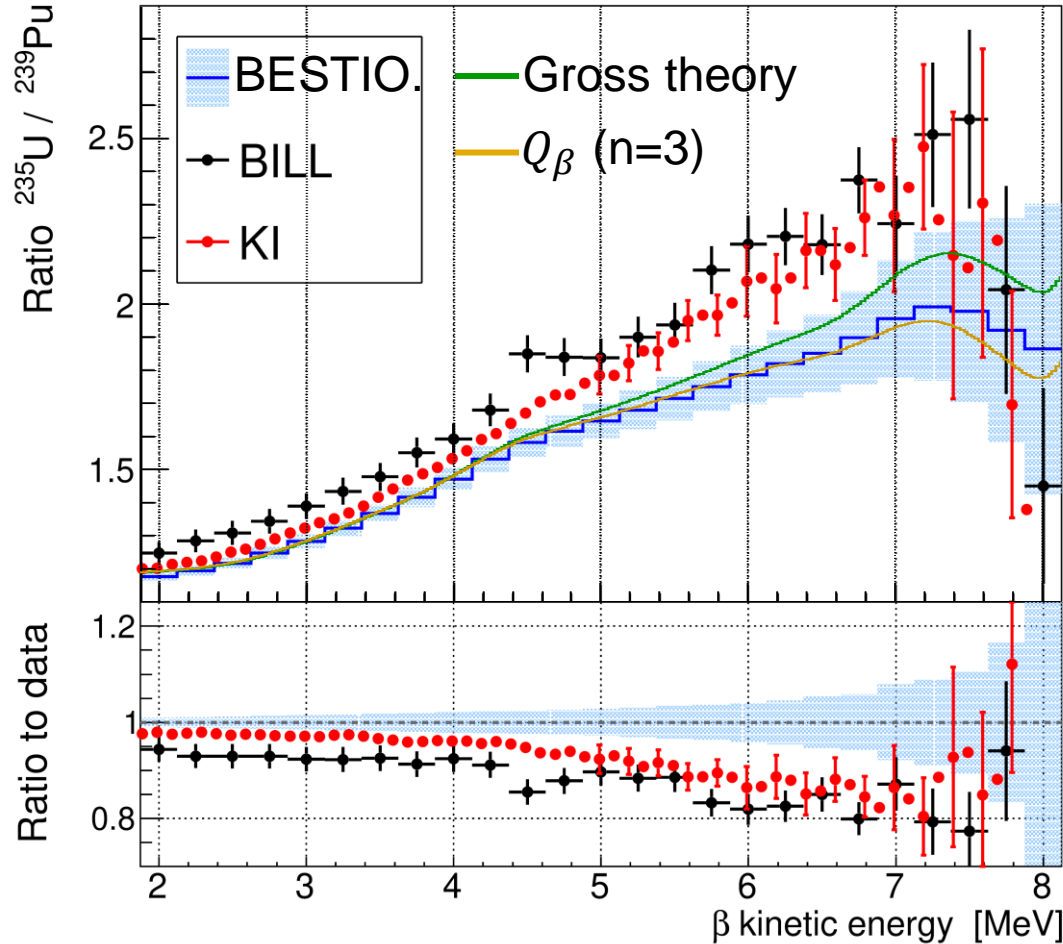


## COMPARISON TO OTHER MODELS

Differences with Estienne-Fallot (EF):

- FY data (JEFF-3.1.1 vs JEFF-3.3)
- NND modeling (>7 MeV)

### IMPACT OF NND MODEL



⇒ Impact of NND modeling observed  
>6 MeV, important >7 MeV

### $^{235}\text{U}/^{239}\text{Pu}$ RATIO OF IBD SPECTRA

