

# Hyper-Kamiokande experiment

ILANCE, LLR, LPNHE, OMEGA

# IN2P3 physicists and engineers

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OMEGA: S. Callier, P. Dinaucourt, S. Conforti, F. Dulucq, L. Raux, C. de la Taille

APC: C. Volpe (theoretician)

Natural continuation of our participation in the Japanese neutrino program (T2K and SK)  
In close collaboration with CEA-IRFU colleagues



# Hyper-Kamiokande in Japan

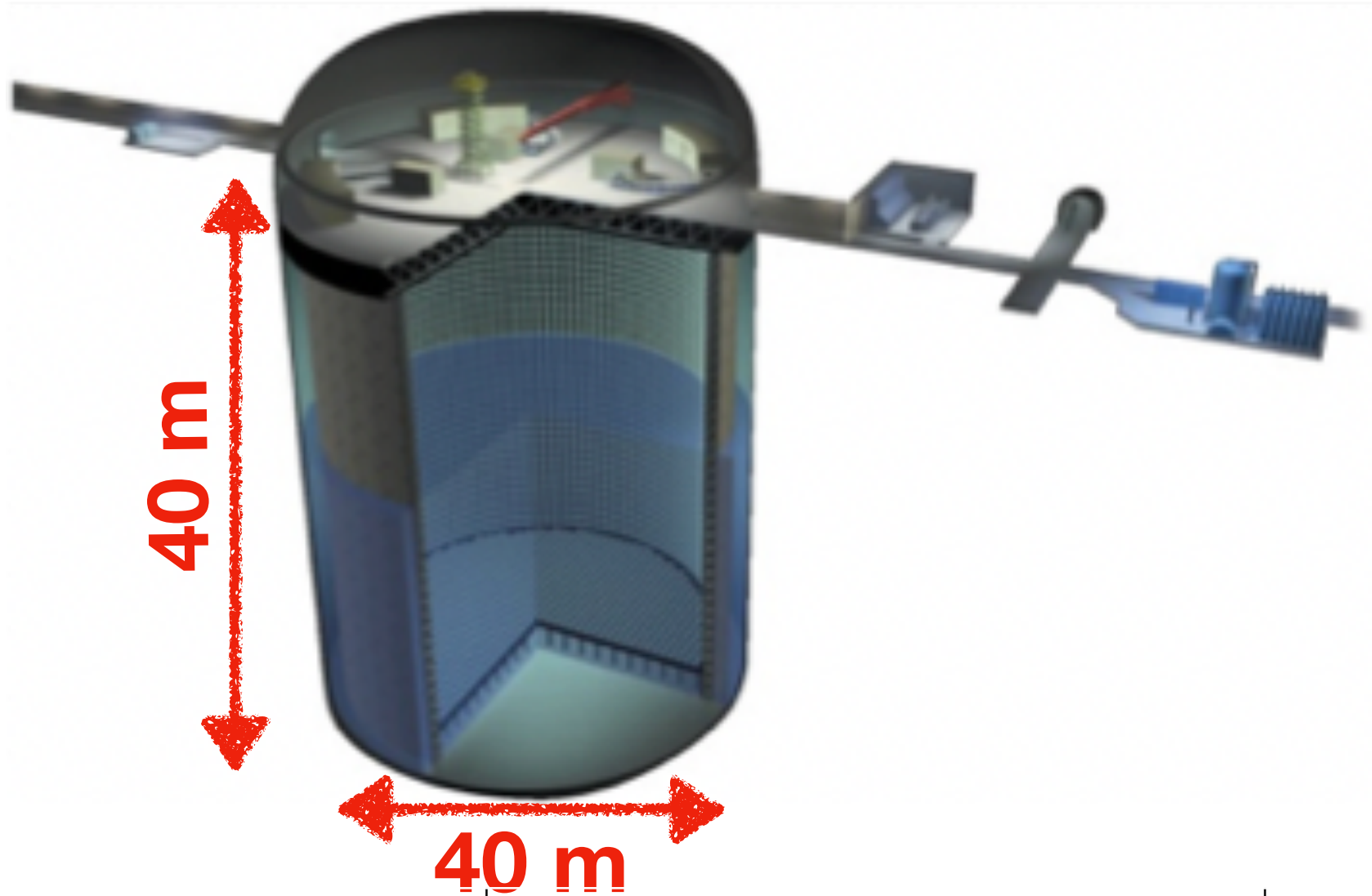


- Third generation Water Cherenkov detector in Japan
- Based on the experience from T2K and Super-Kamiokande
- 295 km and 2.5° off-axis w.r.t. existing neutrino beam ( $\langle E_\nu \rangle \approx 600$  MeV) from J-PARC
- Existing near detector ND280 currently being upgraded for T2K-II
- Vast non-accelerator scientific program

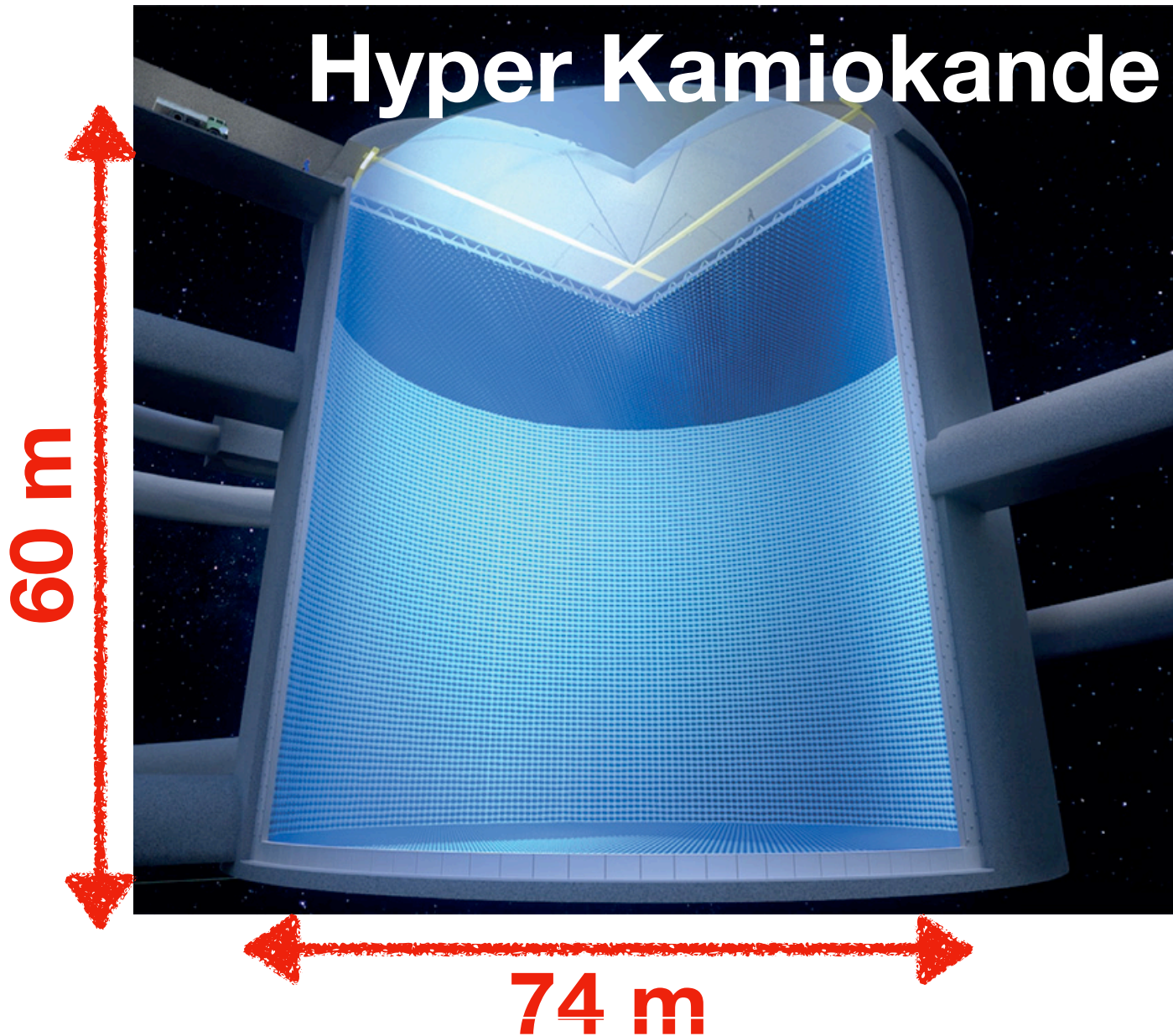


# Hyper-Kamiokande vs Super-Kamiokande

Super Kamiokande



Hyper Kamiokande



	Super Kamiokande	Hyper Kamiokande
Site	Mozumi-yama	Tochibora-yama
Number of ID 20" PMTs	11 129	>20,000
Photo-coverage	40 %	>20%
Single-photon efficiency/PMT	~12%	~24%
Dark rate/PMT	~4 kHz	~4kHz
Time resolution of 1 photon	~3 ns	~1.1 ns
Total/fiducial mass (kton)	50 / <b>22.5</b>	260 / 187

**Fiducial volume x8:**  
→ non-beam  $\nu$  physics

**Beam neutrino  
event rate x 20:**  
→ beam  $\nu$  physics

**Start operations in 2027 with 240 kt.MW and an assumed runtime  $10^7$  s per year**



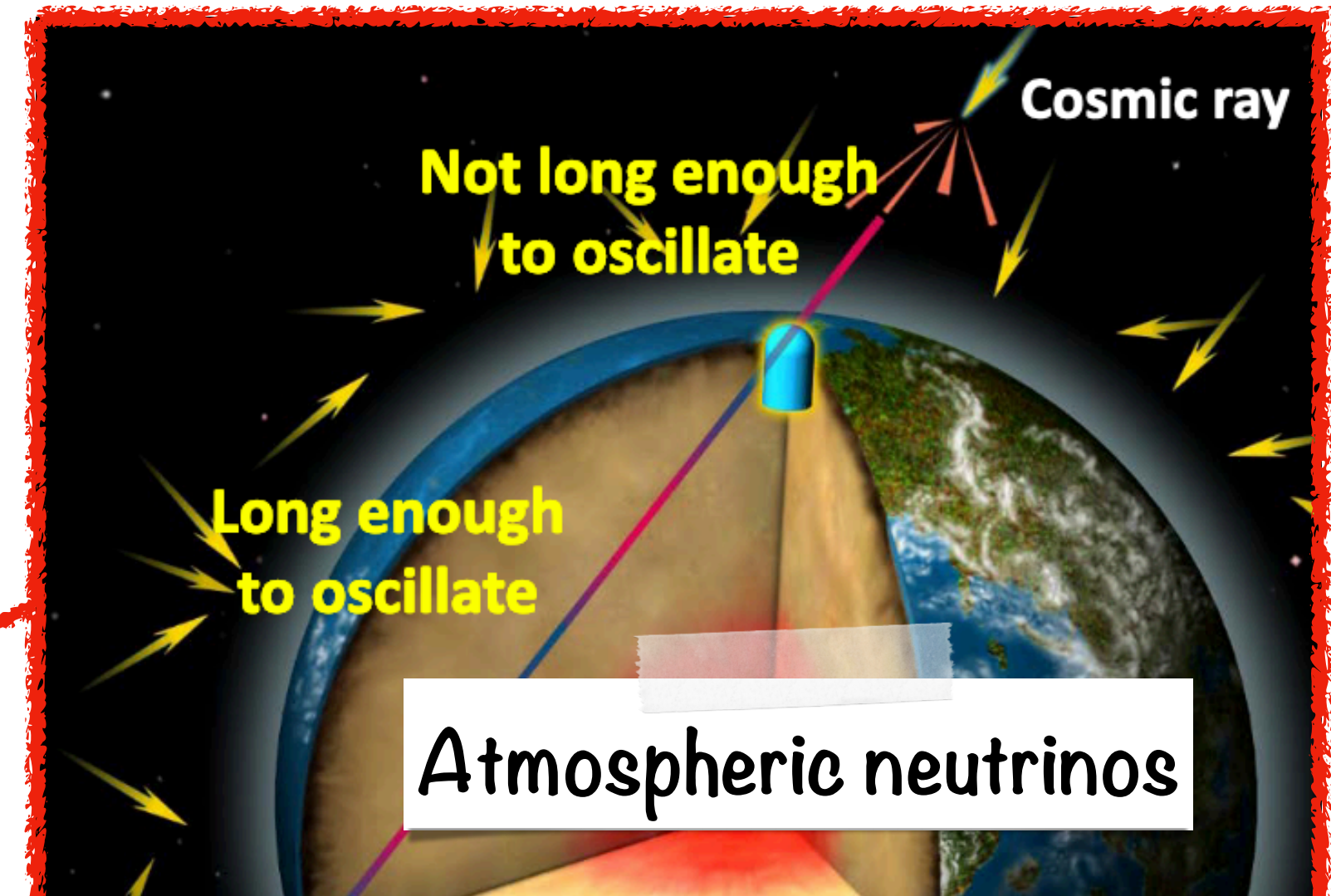
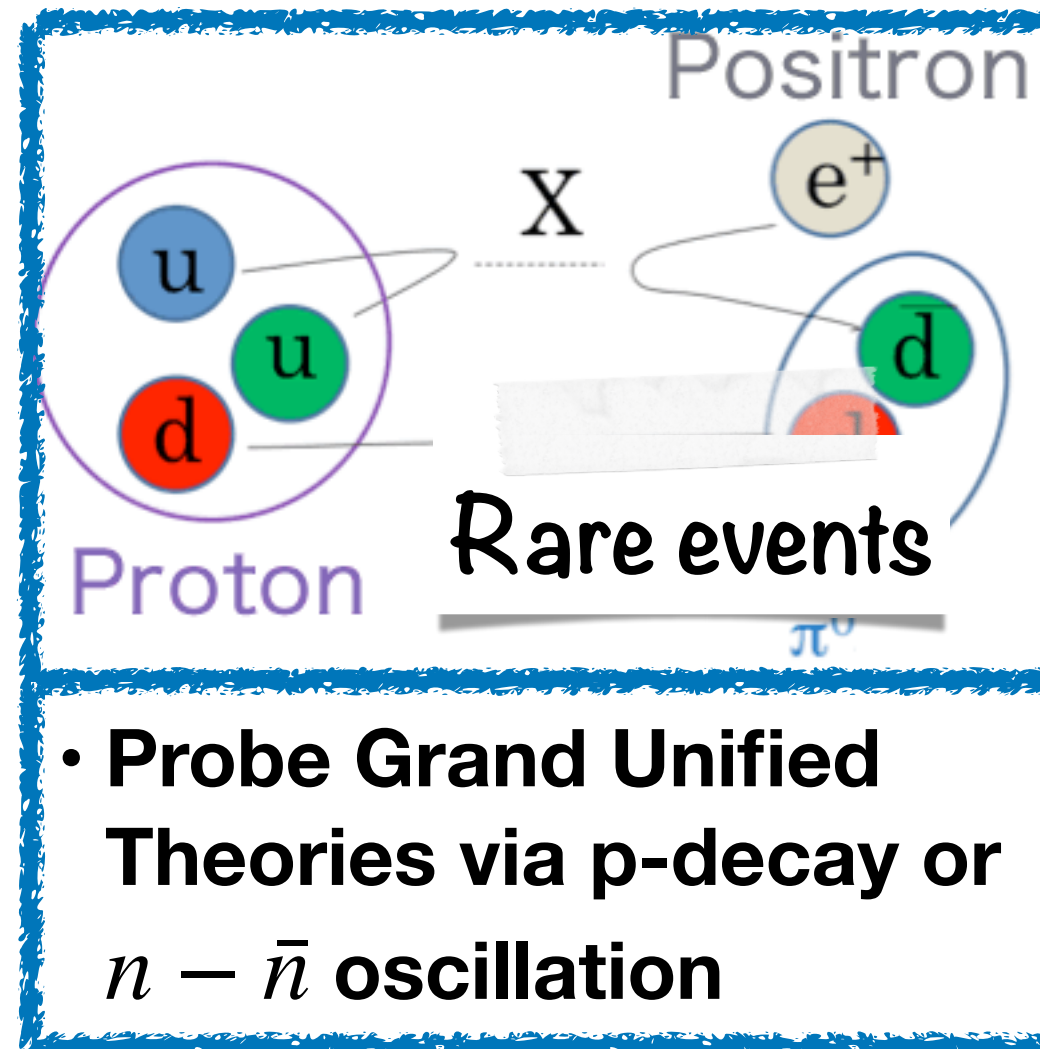
# Hyper-Kamiokande physics program

## Solar neutrinos

- MSW effect
- Non-standard interactions

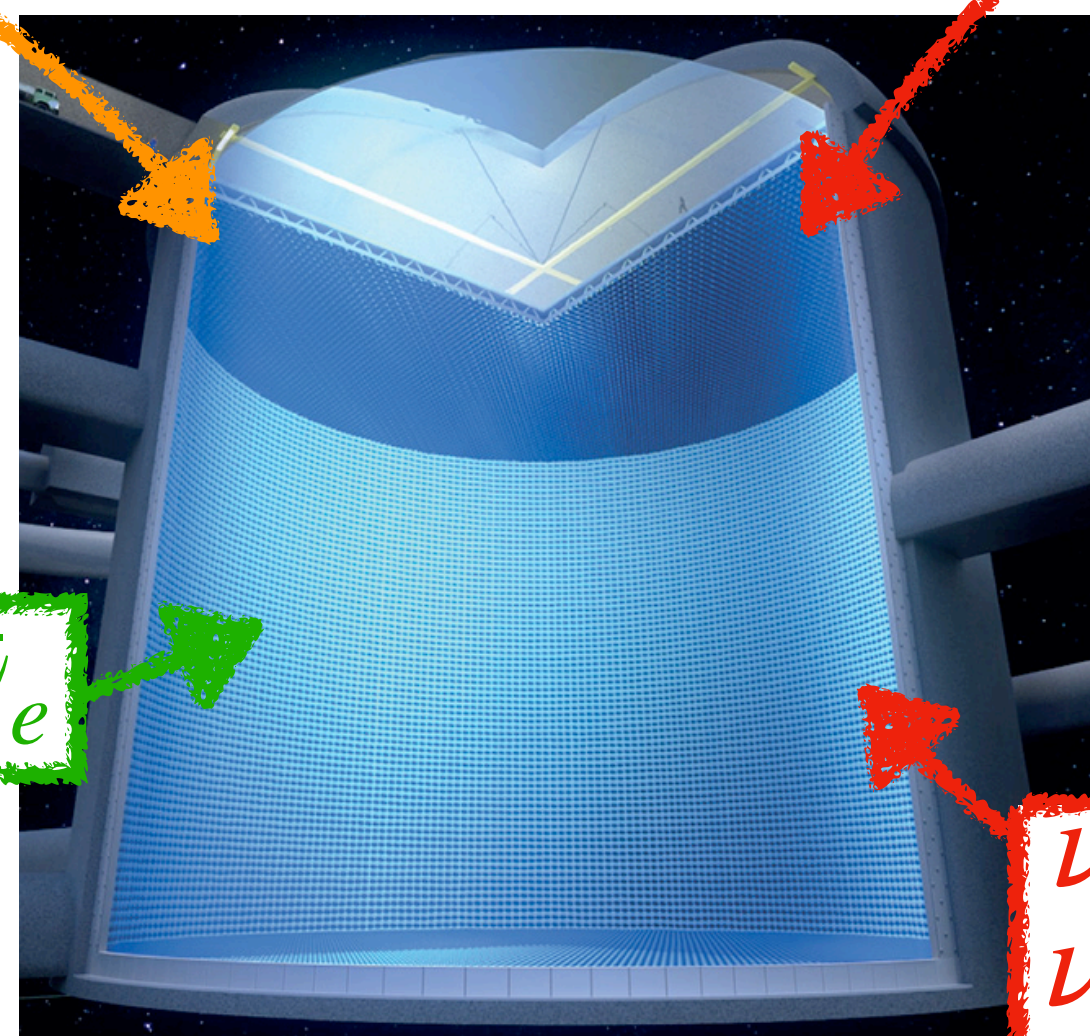
## Supernovae neutrinos

- Transient SN  $\nu$ : constrain SN profile models
- Relic SN  $\nu$ : constrain cosmic star formation



- Observe CP violation for leptons at  $5\sigma$
- Precise measurement of  $\delta_{CP}$
- High sensitivity to  $\nu$  mass ordering

## J-PARC accelerator neutrinos



$\bar{\nu}_e$

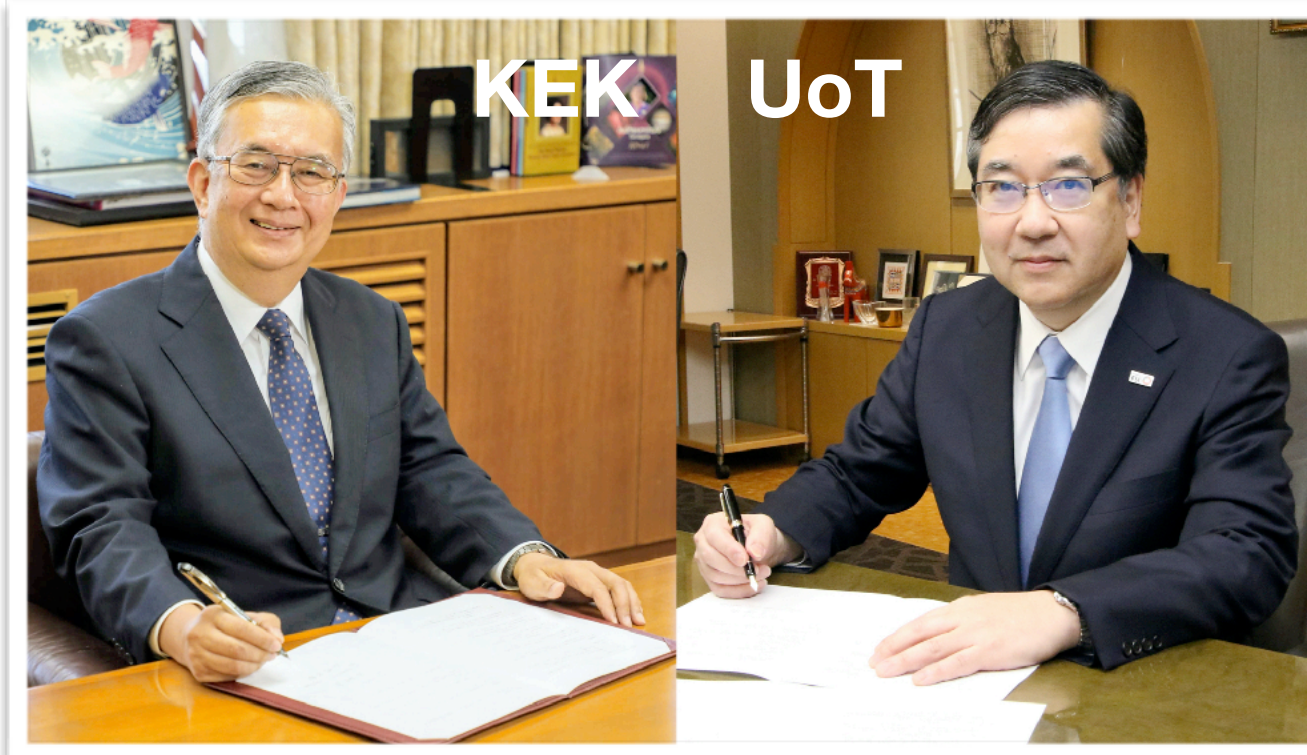
$\nu_e$   $\bar{\nu}_e$   
 $\nu_\mu$   $\bar{\nu}_\mu$

$\nu_e$   $\bar{\nu}_e$   
 $\nu_\mu$   $\bar{\nu}_\mu$



# Experiment approval: key dates

Conclusions of 2018 SC : “Le projet n’est pas actuellement approuvé au Japon et il n’y a pas suffisamment d’informations quant à l’organisation du projet pour envisager et discuter des participations directes à HK.”



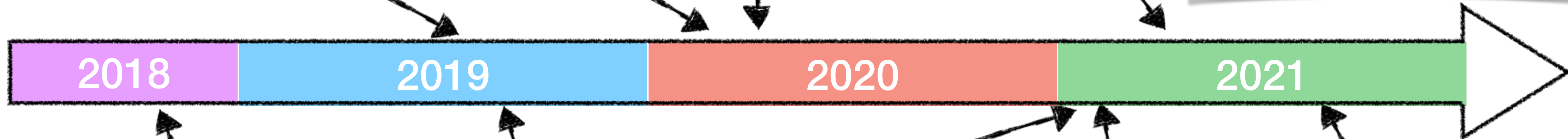
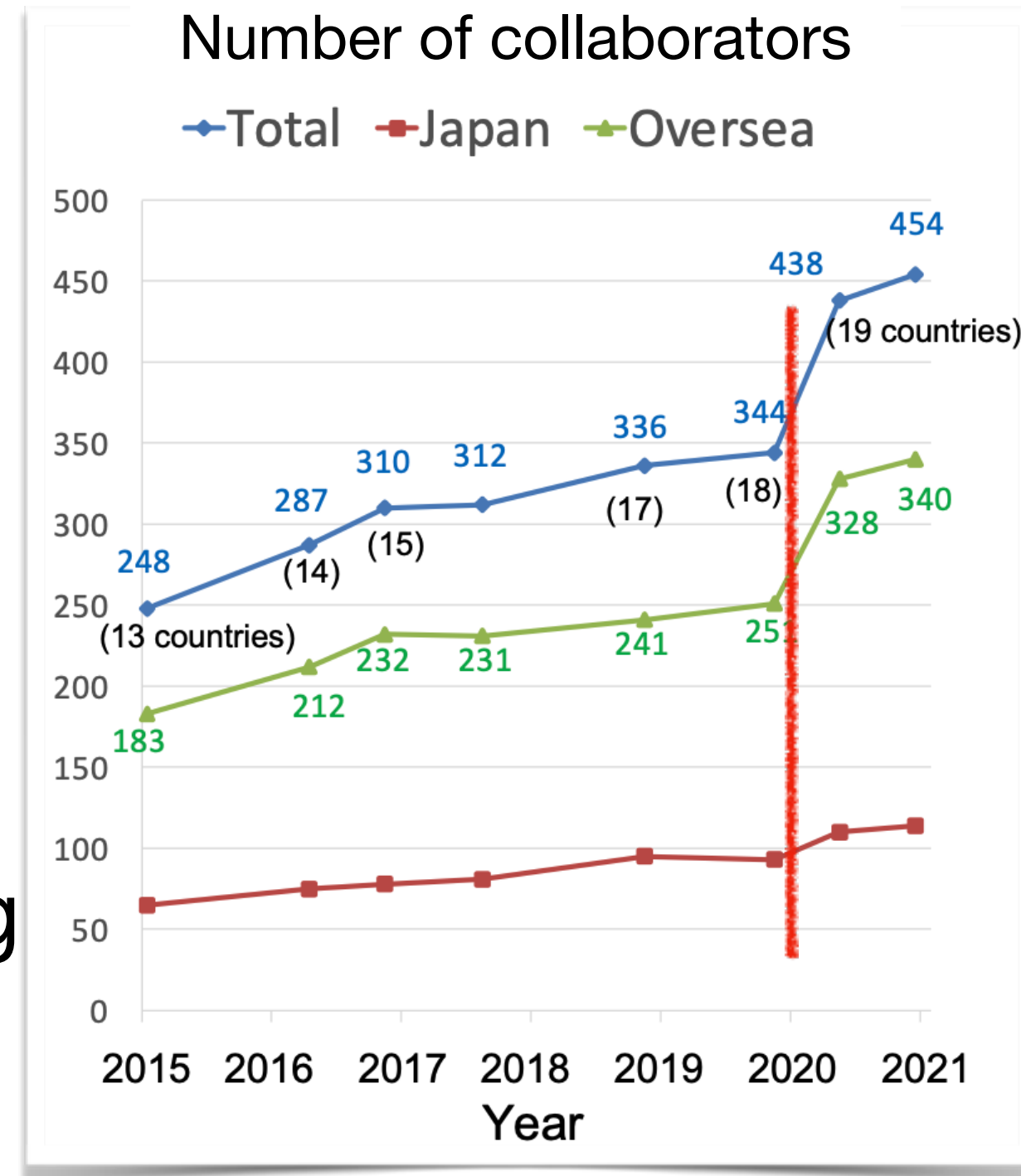
May 2020:

Aug. 2019:  
MEXT approval  
of HK project

Feb. 2020:  
HK budget  
approved

Signature of MOU

May 2021:  
Groundbreaking  
ceremony



Oct. 2018:  
IN2P3 CS  
discussing HK

Oct. 2019:  
LPNHE CS  
approved HK

Jan. 2021:  
CEA CS  
approved HK

March 2021:  
LLR CS  
approved HK

Oct. 2021:  
IN2P3 CS  
discussing HK



# Conclusions of the 2021 SC

## 6.3. Avis du Conseil

La dernière présentation de la contribution de l'IN2P3 à HK en session du Conseil Scientifique de l'IN2P3 est assez récente (octobre 2018). En octobre 2021, le Conseil constate une évolution positive remarquable au cours des deux dernières années :

- Le projet HK est approuvé par le gouvernement japonais en août 2019, le budget (500 M\$) est voté en février 2020, et le début de la prise de données est confirmé pour 2027 ;
- Deux laboratoires de l'IN2P3 (LPNHE, LLR) et le CEA, soutenus par leurs conseils scientifiques respectifs, rejoignent le projet HK, en octobre 2019 pour le LPNHE, en janvier 2021 pour le CEA, et en mars 2021 pour le LLR.

...

Le Conseil souligne cependant qu'un engagement technique direct sur le détecteur lointain de HK est requis pour confirmer et valider le ticket d'entrée de l'IN2P3 dans HK<sup>4</sup>, et considère que la participation centrale proposée par le LPNHE et le LLR en collaboration avec OMEGA sur le digitaliseur HKROC et sur le système de distribution d'horloge répond à cette condition dans une enveloppe budgétaire raisonnable. Le Conseil note également que les équipes de l'IN2P3 proposent une réflexion pour contribuer aux coûts d'opération à travers la participation du CC-IN2P3.



**RD4HK initiated in 2022  
(Benjamin's talk)**



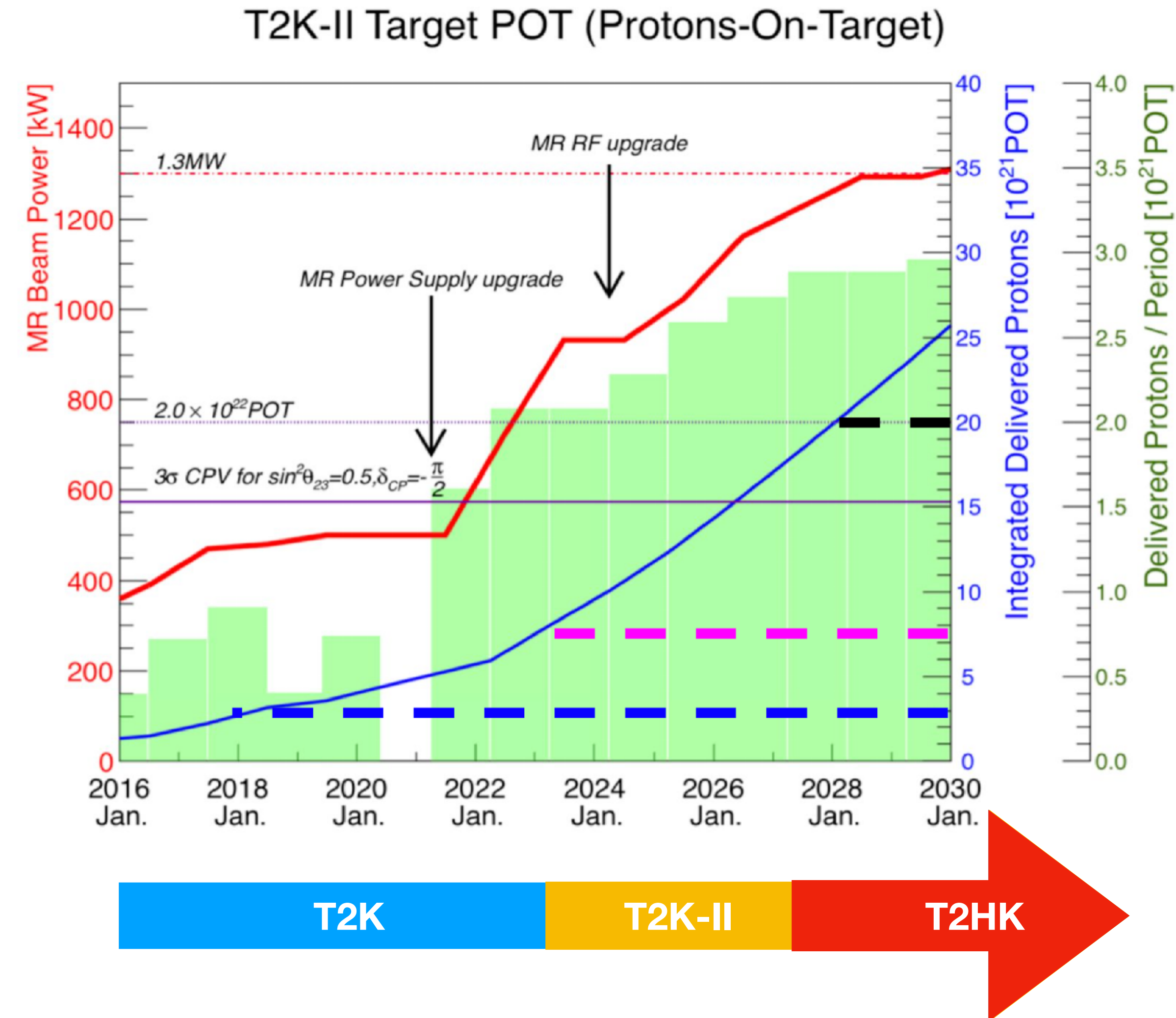
# Tokai to HK: heritage from T2K

## Accelerator upgrade

Power increase (500kW  $\rightarrow$  1.3 MW)

**x2.7 more stats per s (wrt T2K-I)**

$\nu/\bar{\nu}$  flux uncertainty  $< 5\%$  thanks to  
NA61/SHINE hadroproduction measurements





# Tokai to HK: heritage from T2K

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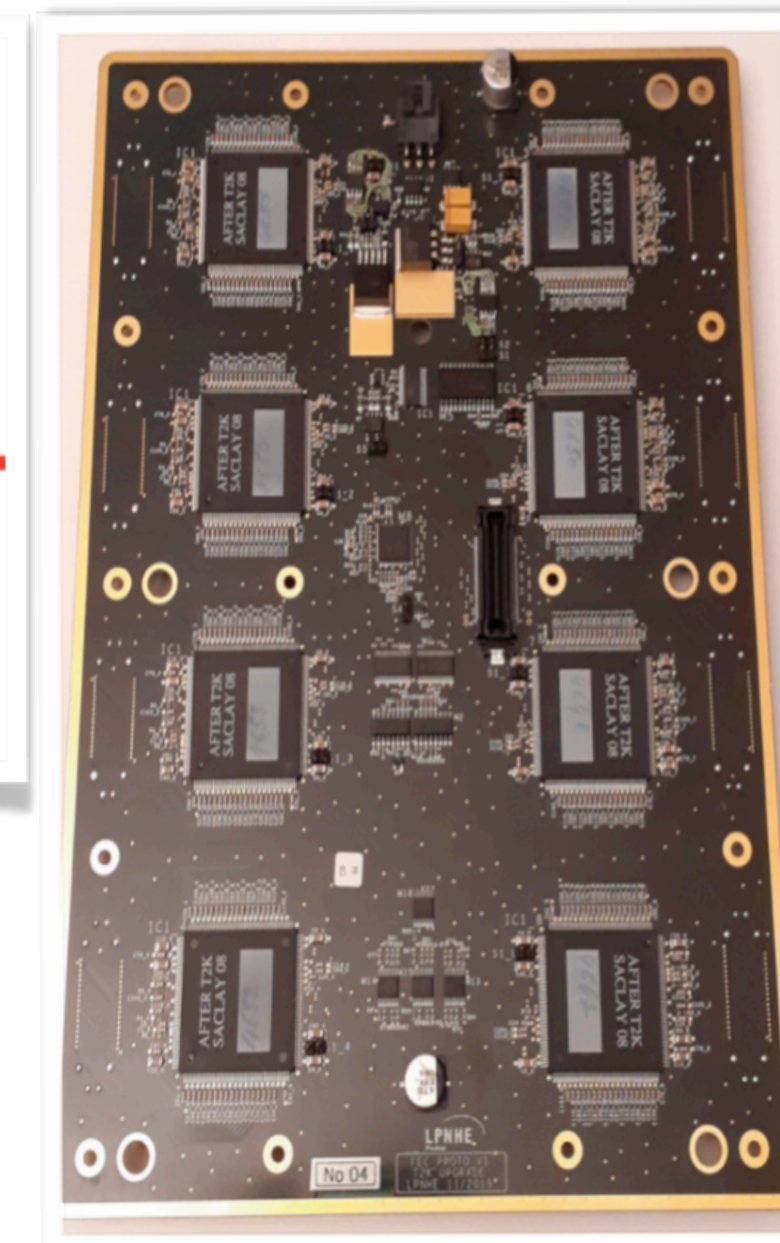
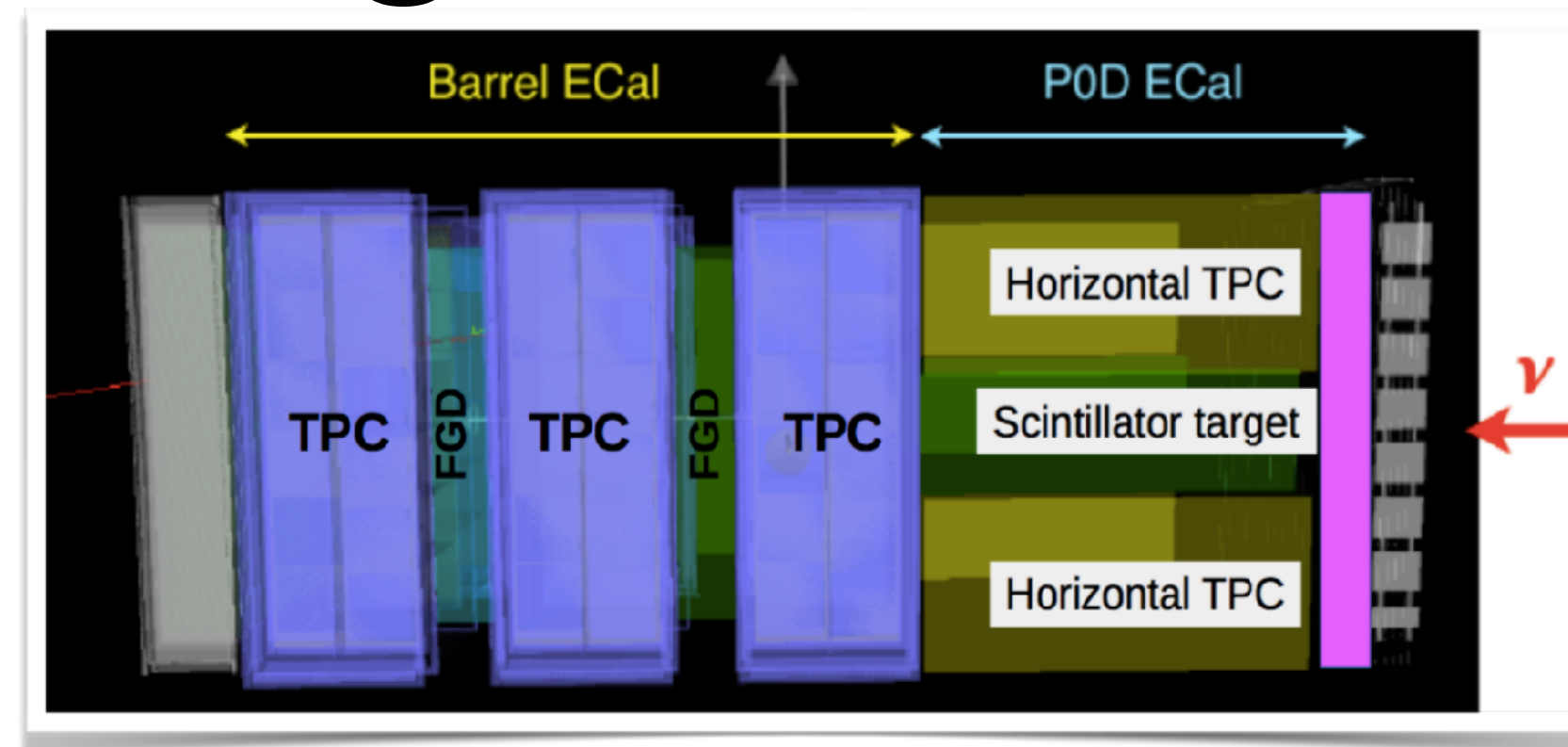
## Magnetized near detector @280 m (ND280) + INGRID

Used for T2K Oscillation Analysis for >10 years

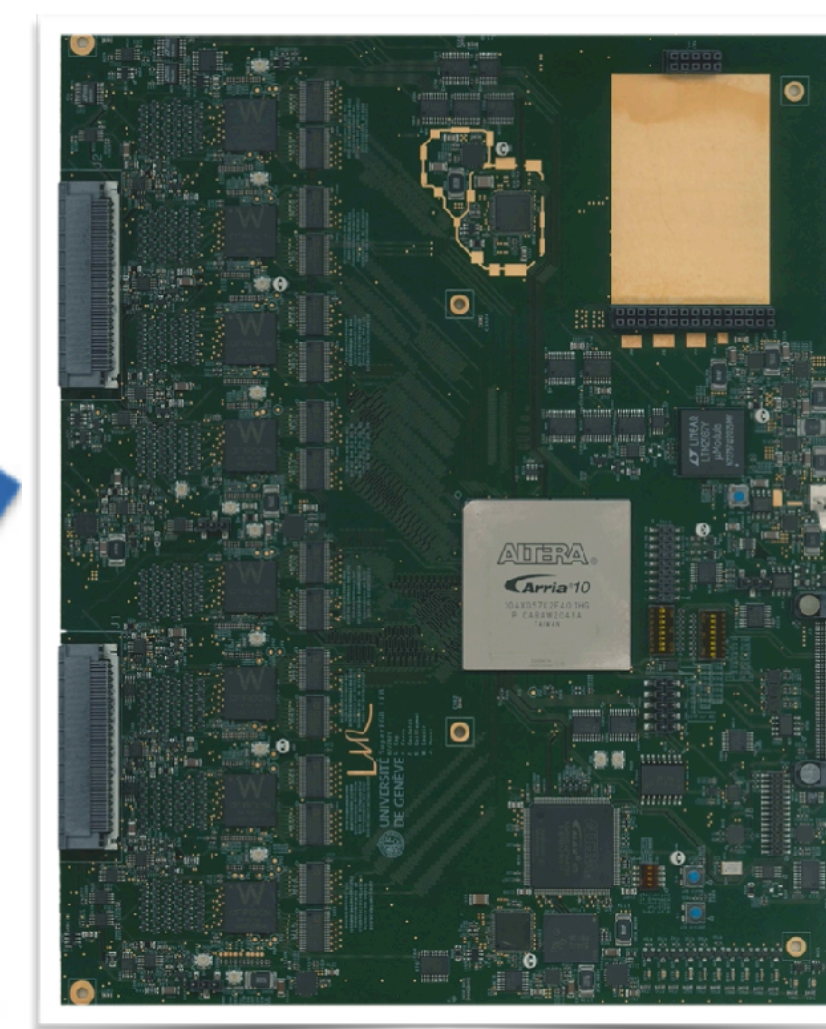
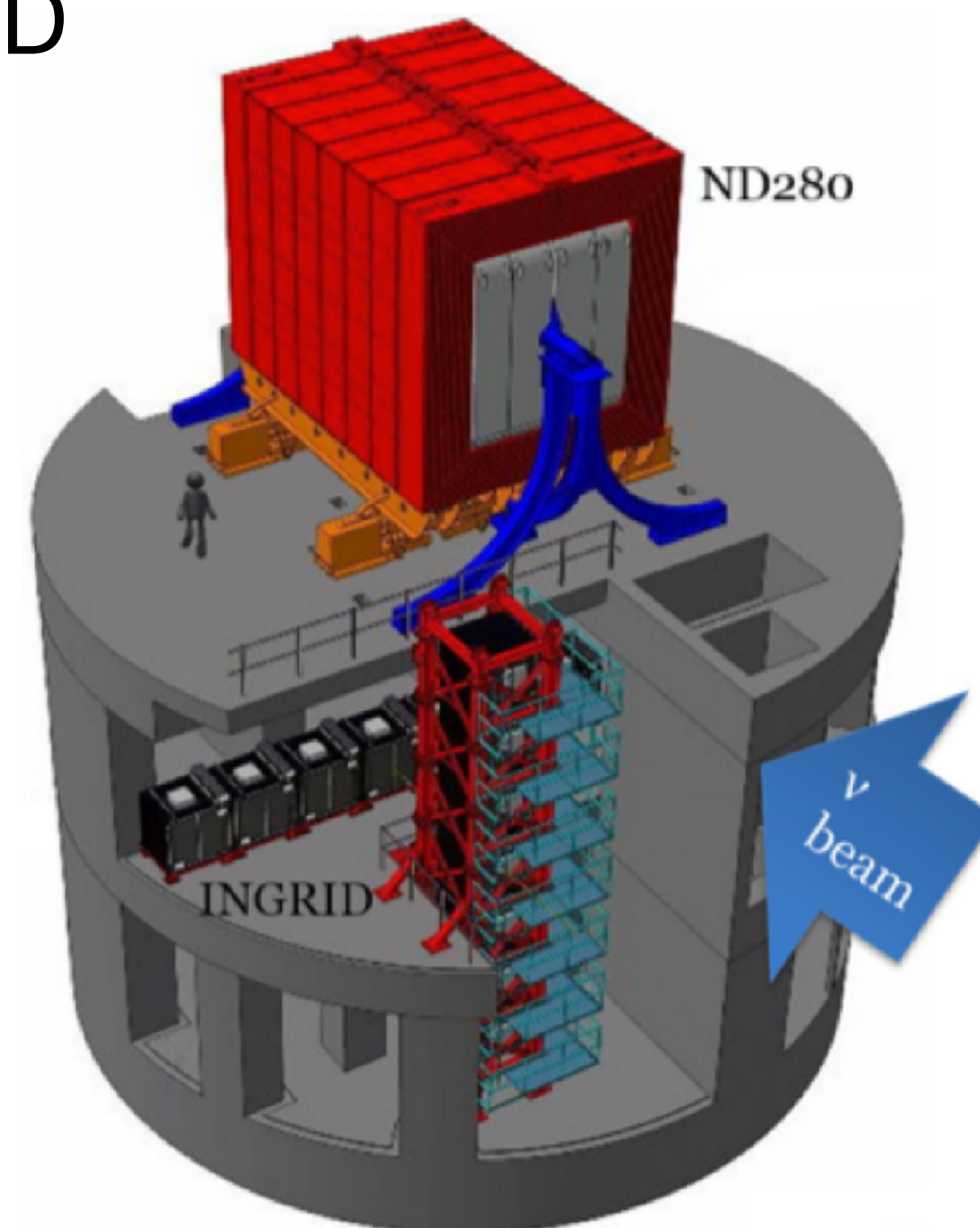
Being upgraded now for T2K-II

(crucial contributions and strong support from IN2P3)

**Systematics uncertainties under control  
from Day-1 of HK**



FEC for HA-TPC readout



FEB for SuperFGD readout



# Tokai to HK: what will be new

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Magnetized near detector @280 m (ND280) + INGRID

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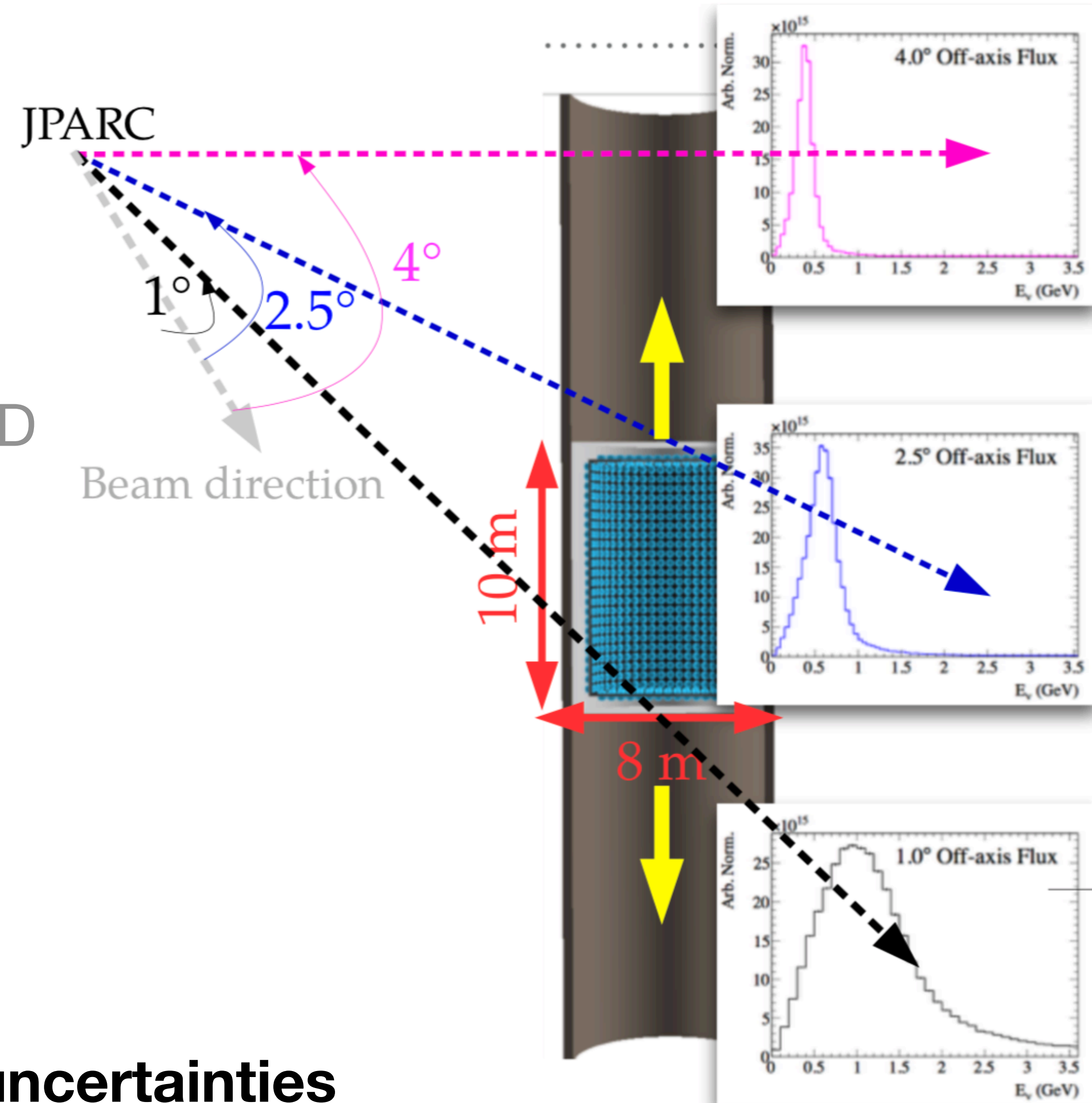
Intermediate Water Cherenkov Detector (IWCD/E61)

Measure  $\nu$  interactions on Water

High stats. sample of  $\nu_e$  interactions

**Needed to reach final HK goal for systematics uncertainties**

*(if IWCD is delayed, ND280 will be the only near detector on day 1 of HK)*

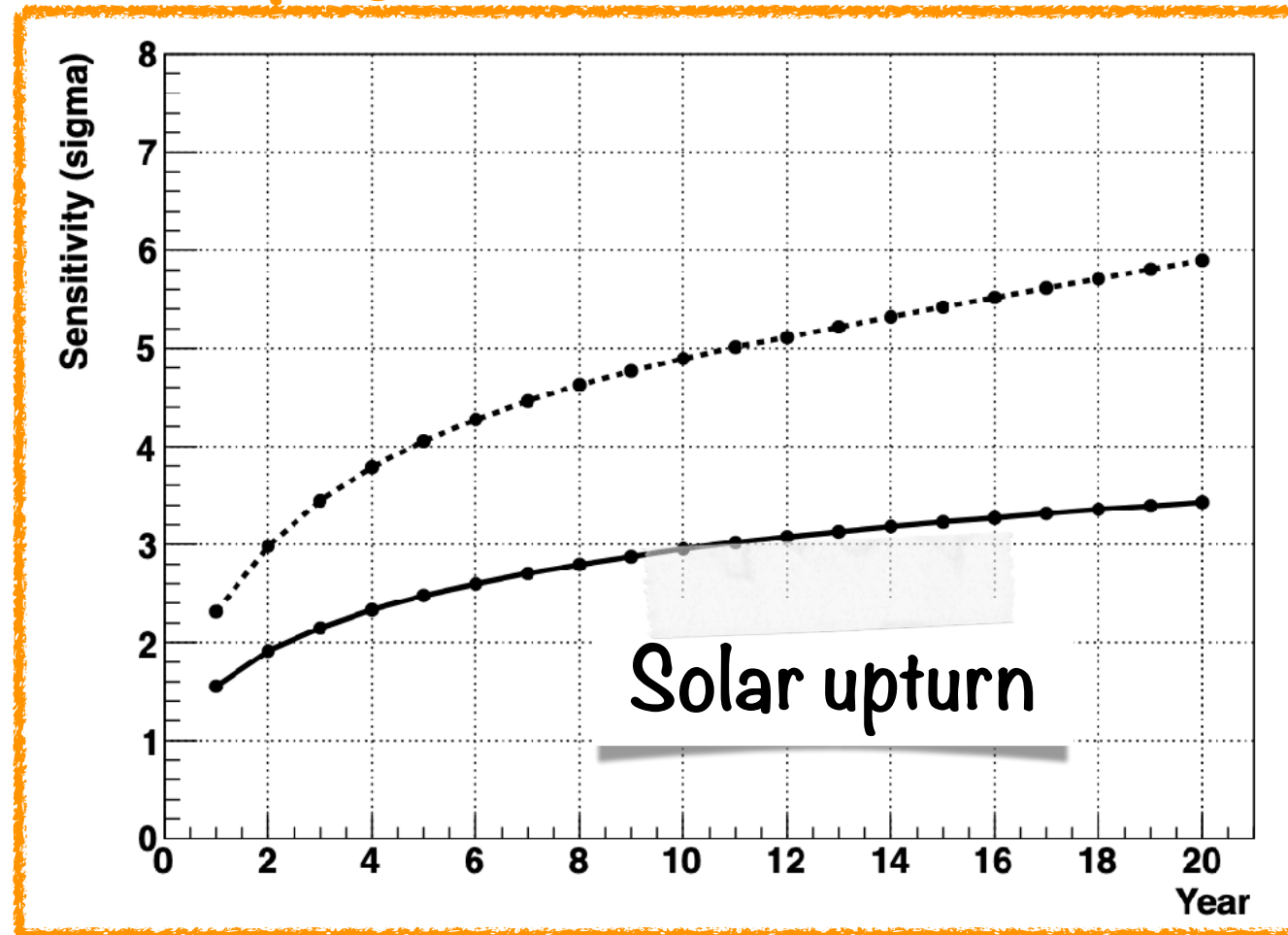


NEW

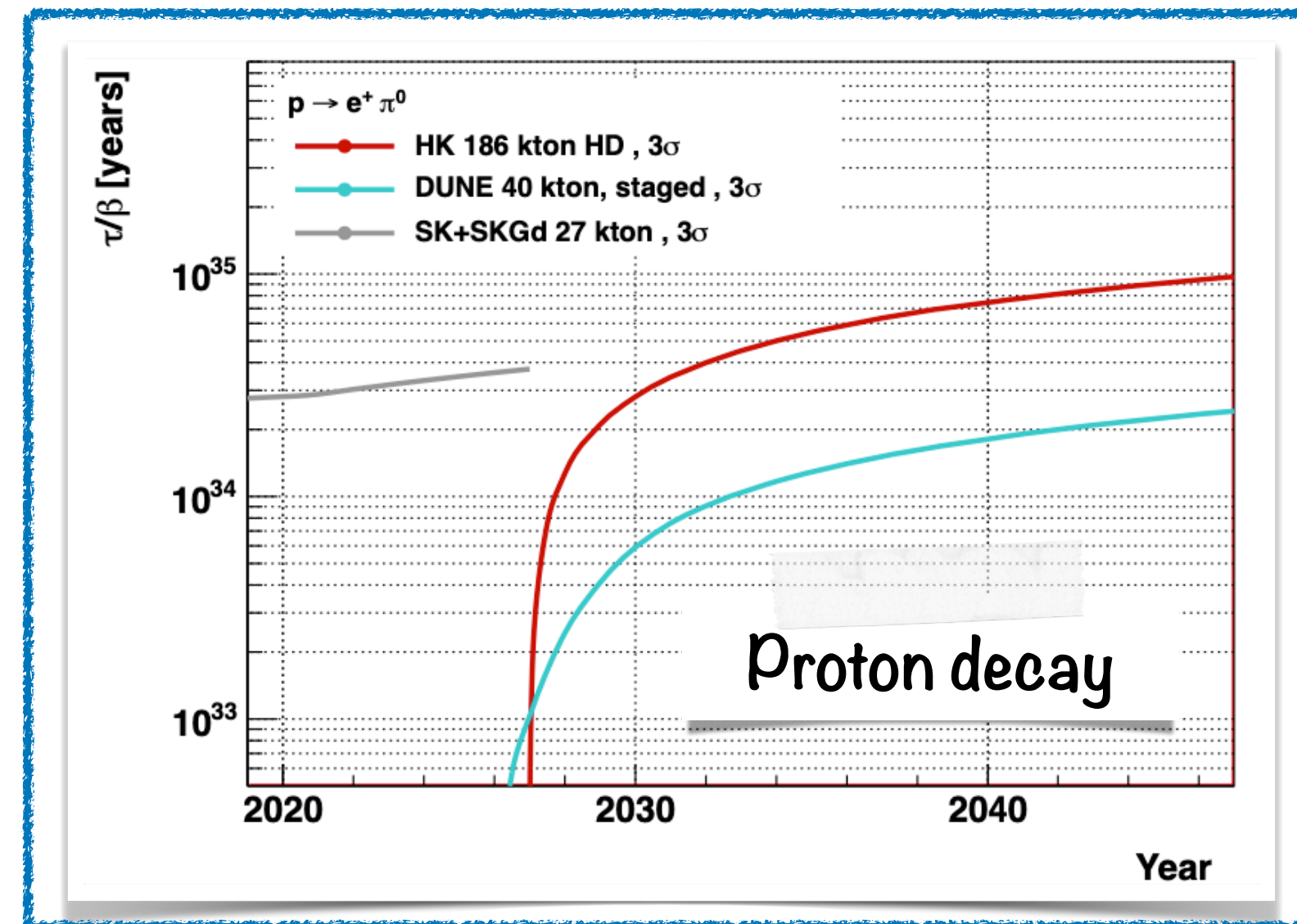


# Hyper-Kamiokande Physics

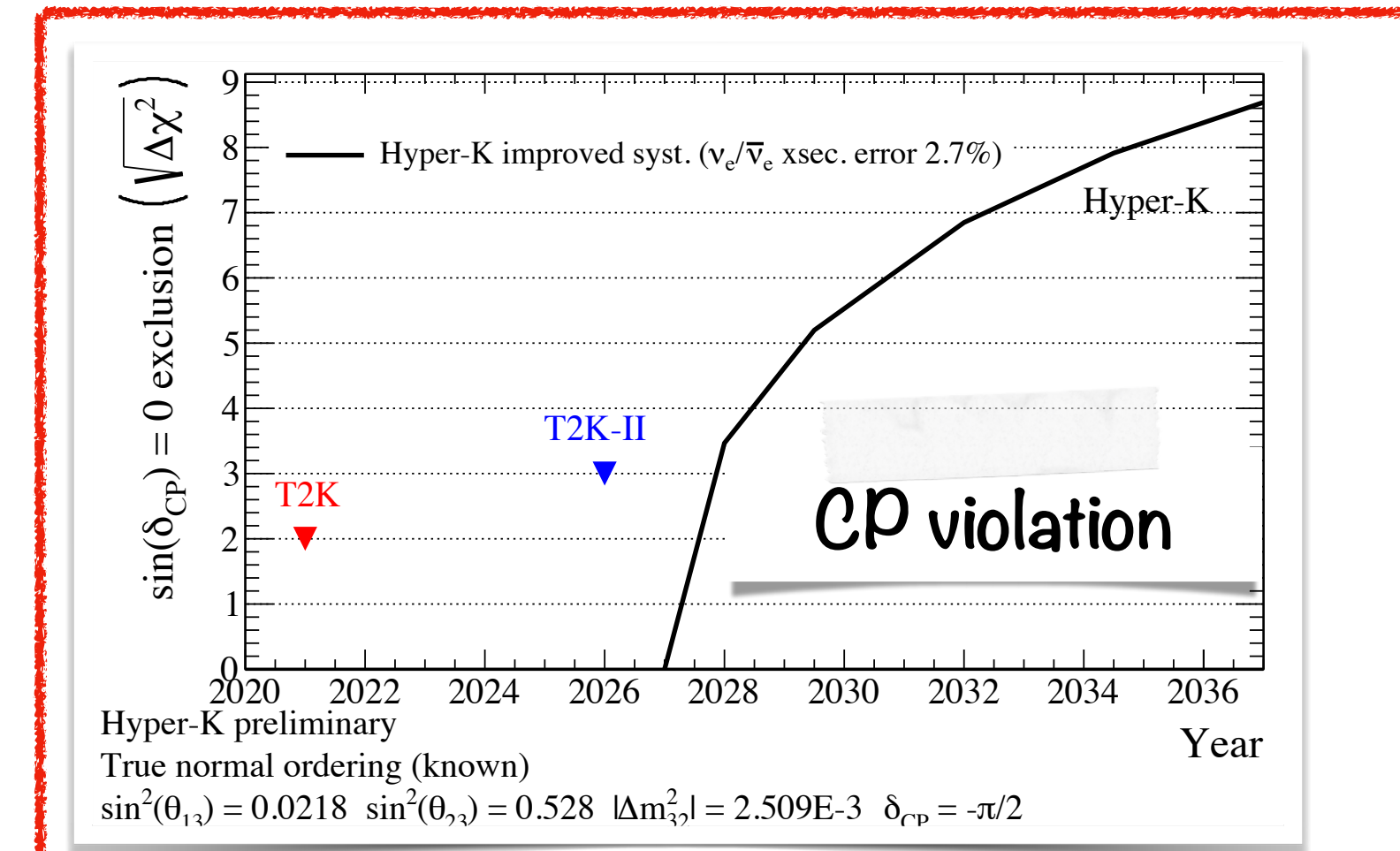
## Solar physics



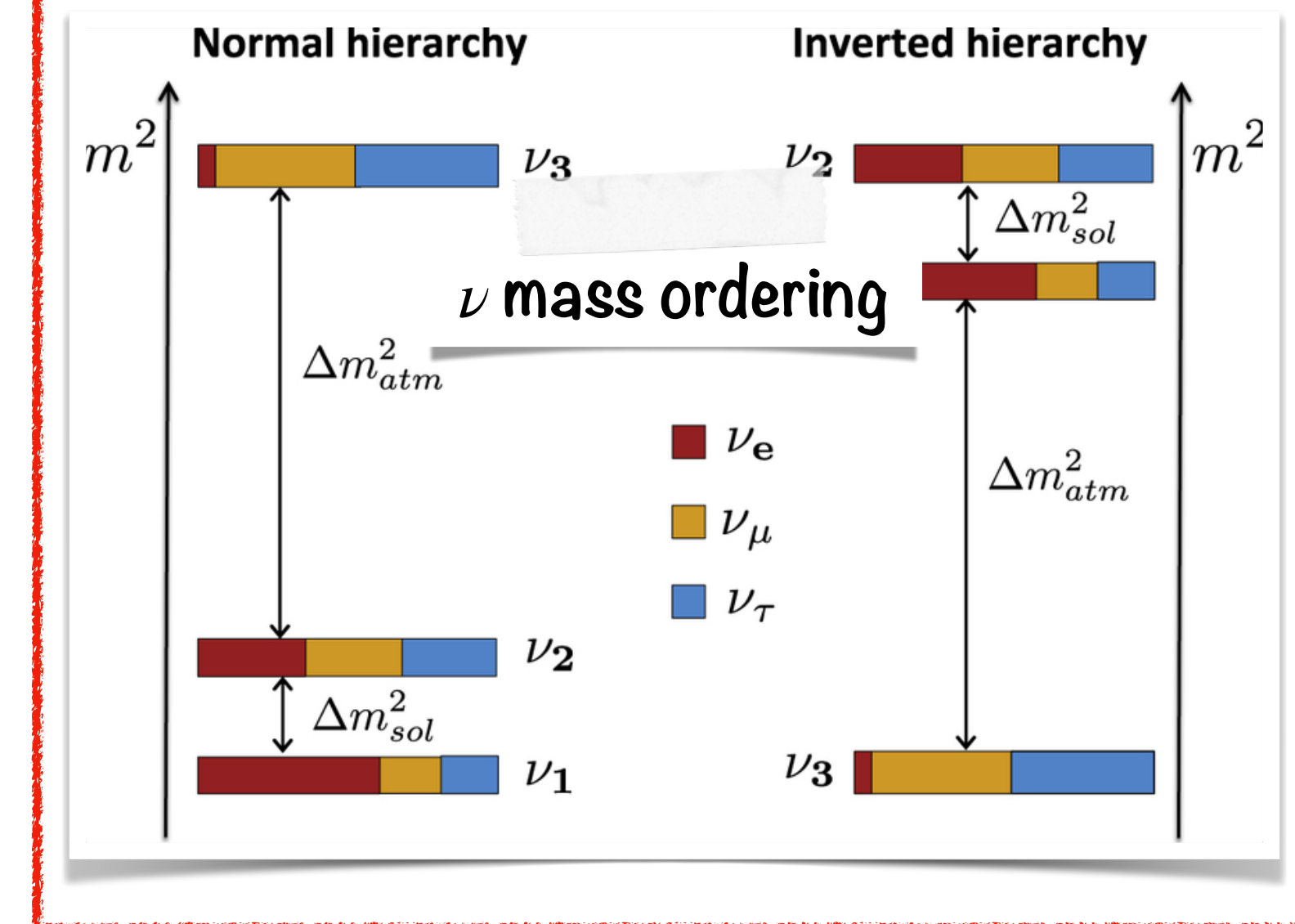
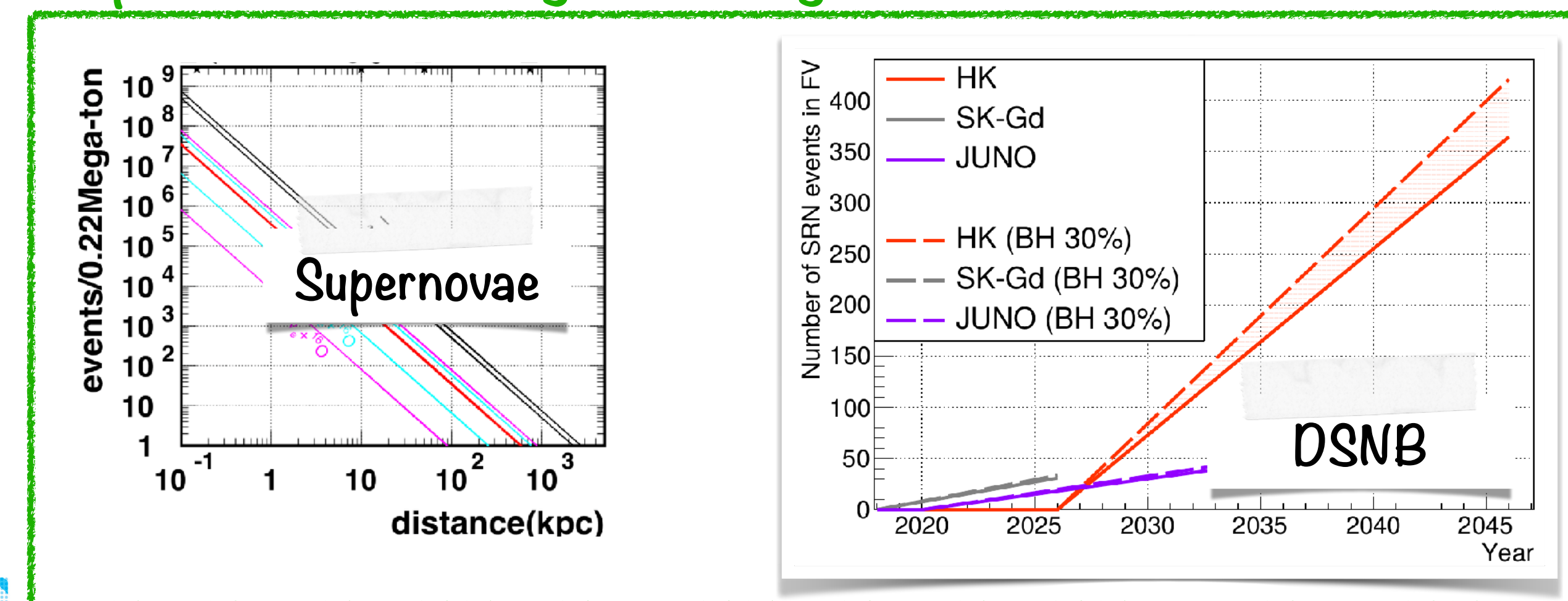
## Rare events



## Neutrinos oscillation



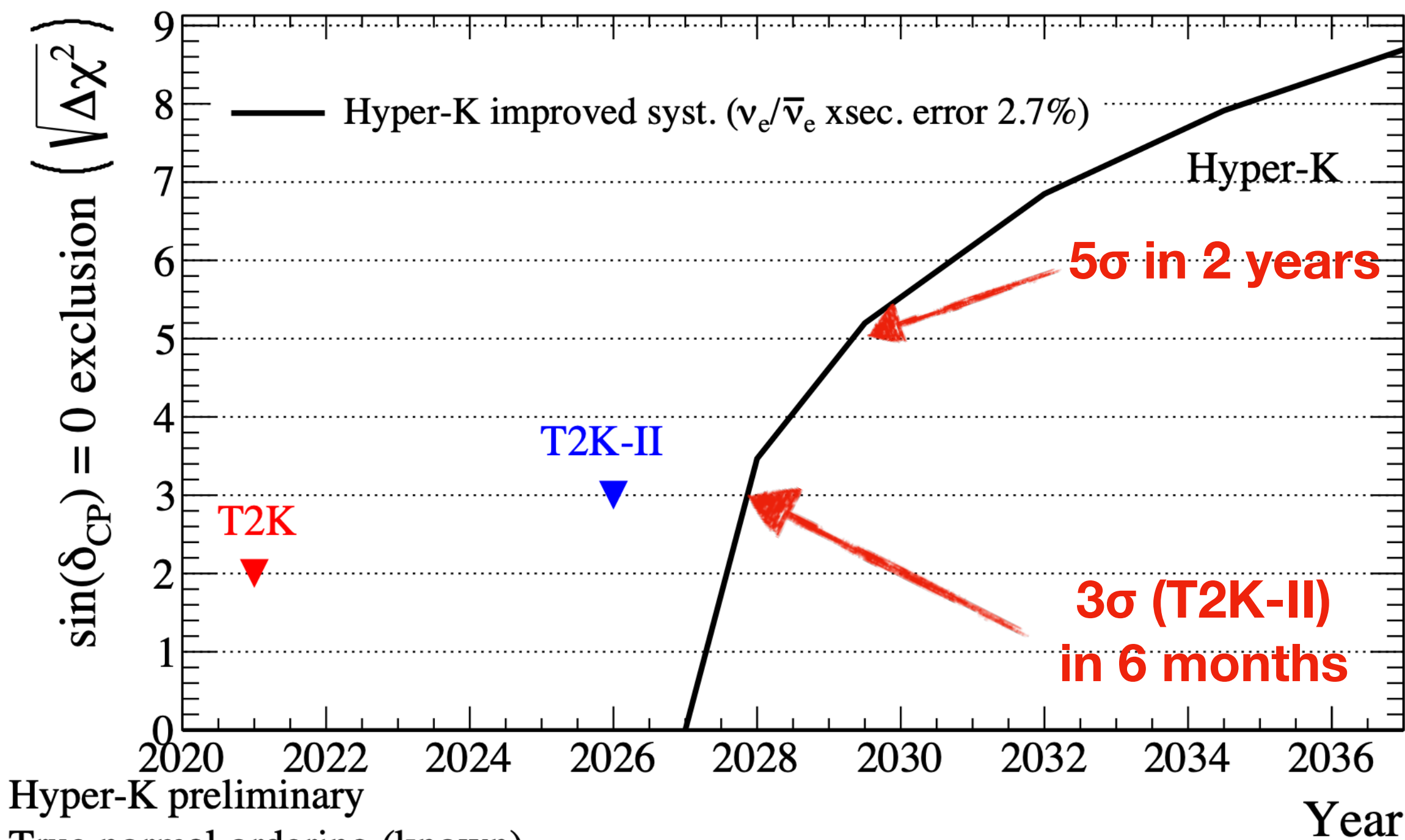
## Supernovae modeling and Early Universe





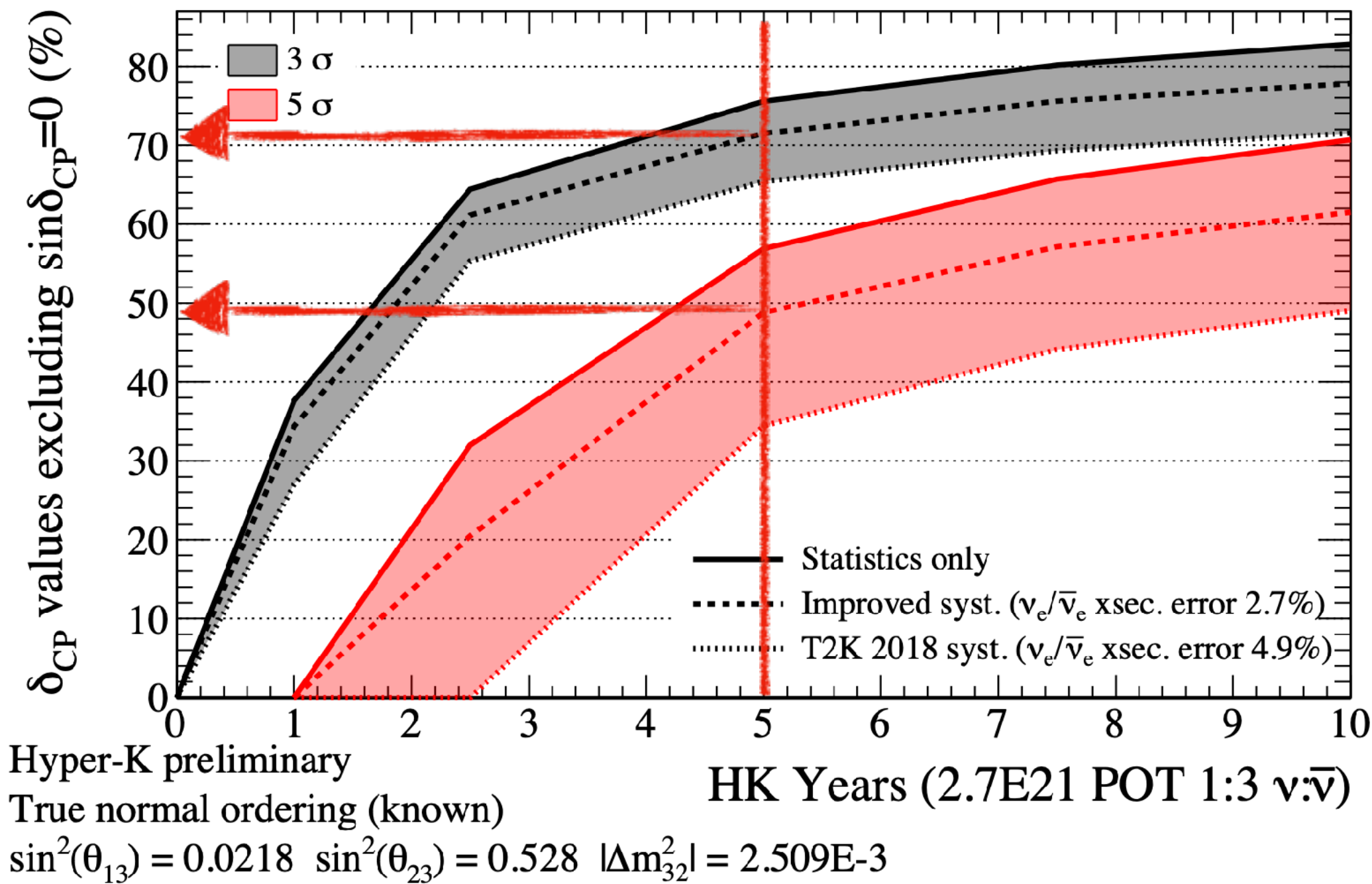
# Fast CP-violation discovery

Known mass ordering



Hyper-K preliminary  
True normal ordering (known)  
 $\sin^2(\theta_{13}) = 0.0218$   $\sin^2(\theta_{23}) = 0.528$   $|\Delta m_{32}^2| = 2.509\text{E-}3$   $\delta_{CP} = -\pi/2$

If  $\delta_{CP}=-\pi/2$ , CP violation discovered before any other LBL- $\nu$  experiment  
Fastest experiment to survey possible  $\delta_{CP}$  values



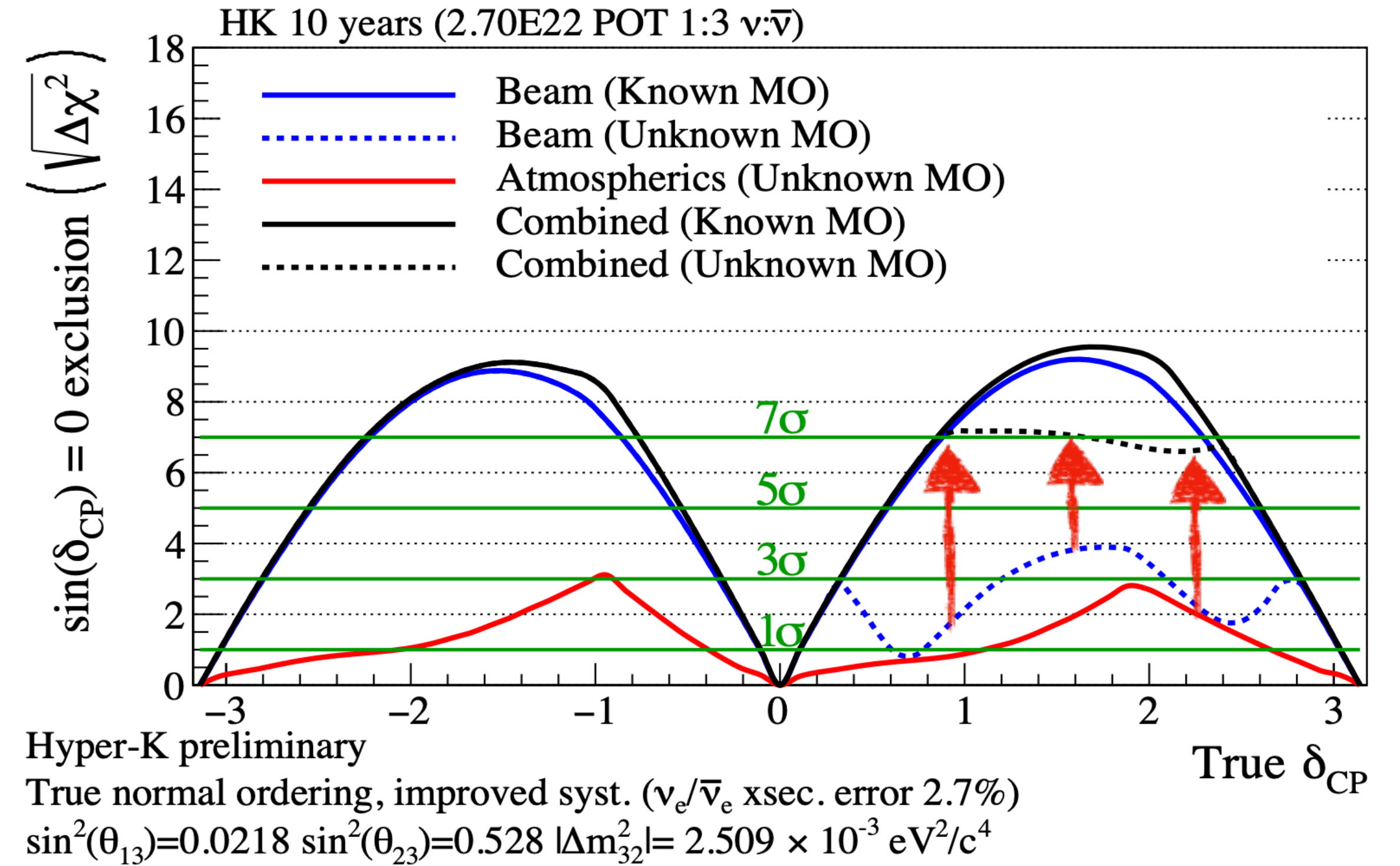
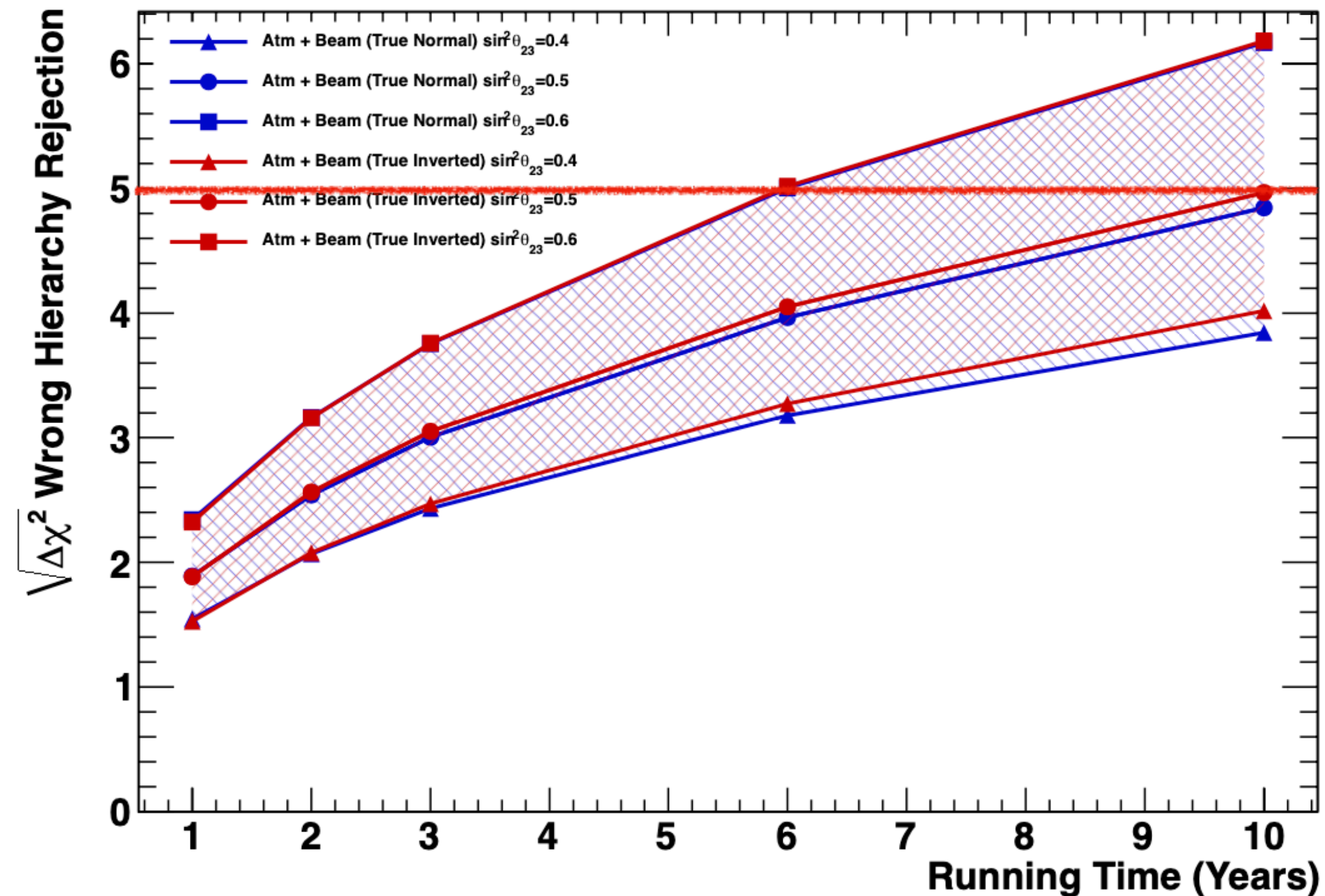
	$\delta_{CP} = -\pi/2$		All $\delta_{CP}$	
	3 $\sigma$	5 $\sigma$	50 % 5 $\sigma$	70 % 3 $\sigma$
Hyper Kamiokande	0.5 y	2 y	5 y	5 y
DUNE (staged*)	4 y	8 y	10 y	13 y

\* 2 modules@1.2 MW y1; 3 modules y2;  
4 modules y4; @2.4 MW y7

DUNE CDR [arXiv:2002.03005](https://arxiv.org/abs/2002.03005)  
IUPAP Neutrino panel [report](#)



# Mass ordering sensitivity with atmospherics

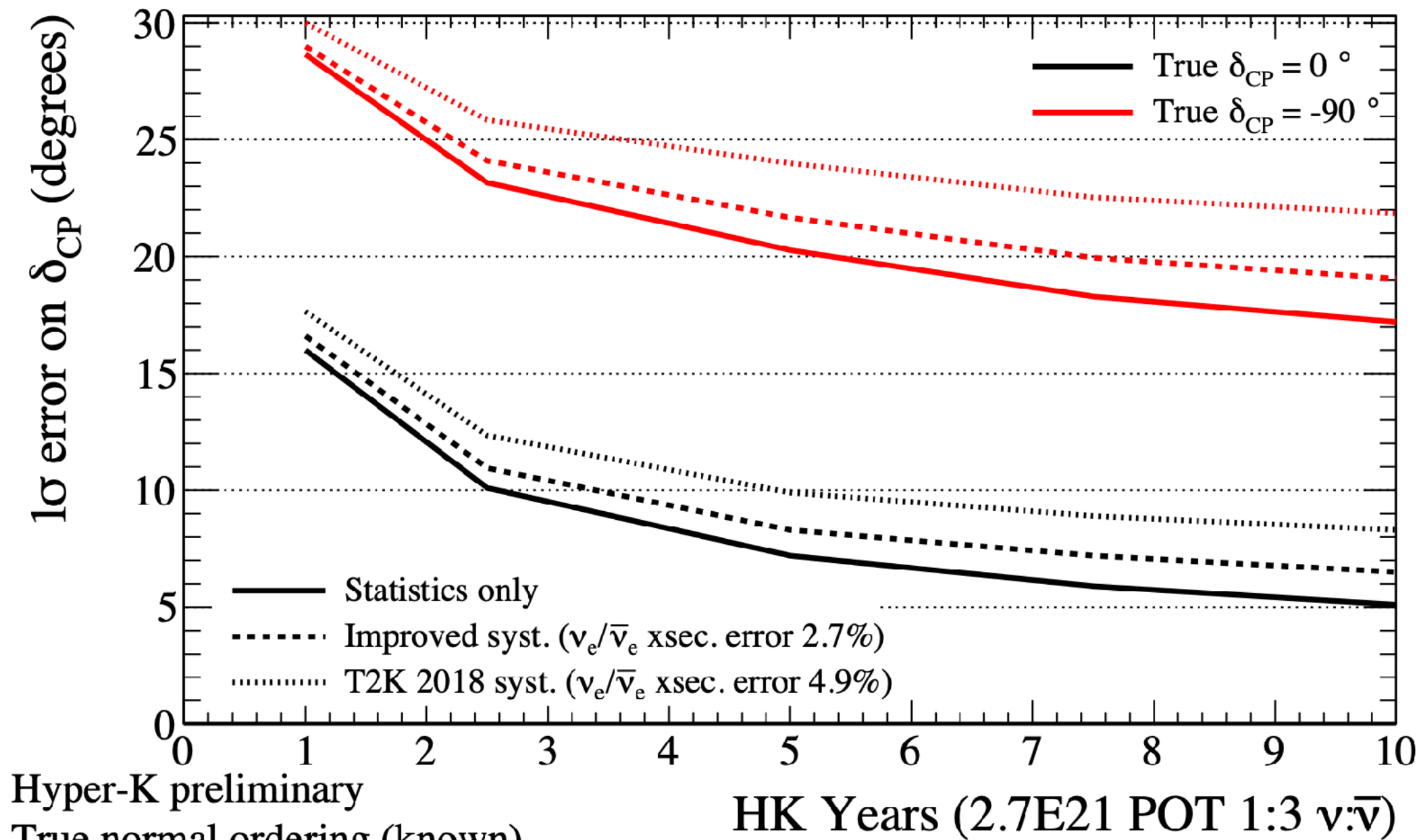


If not discovered by T2K/SK, NOvA, ORCA or JUNO before 2027, HK can determine MO after 6-10 years via atmospheric  $\nu$

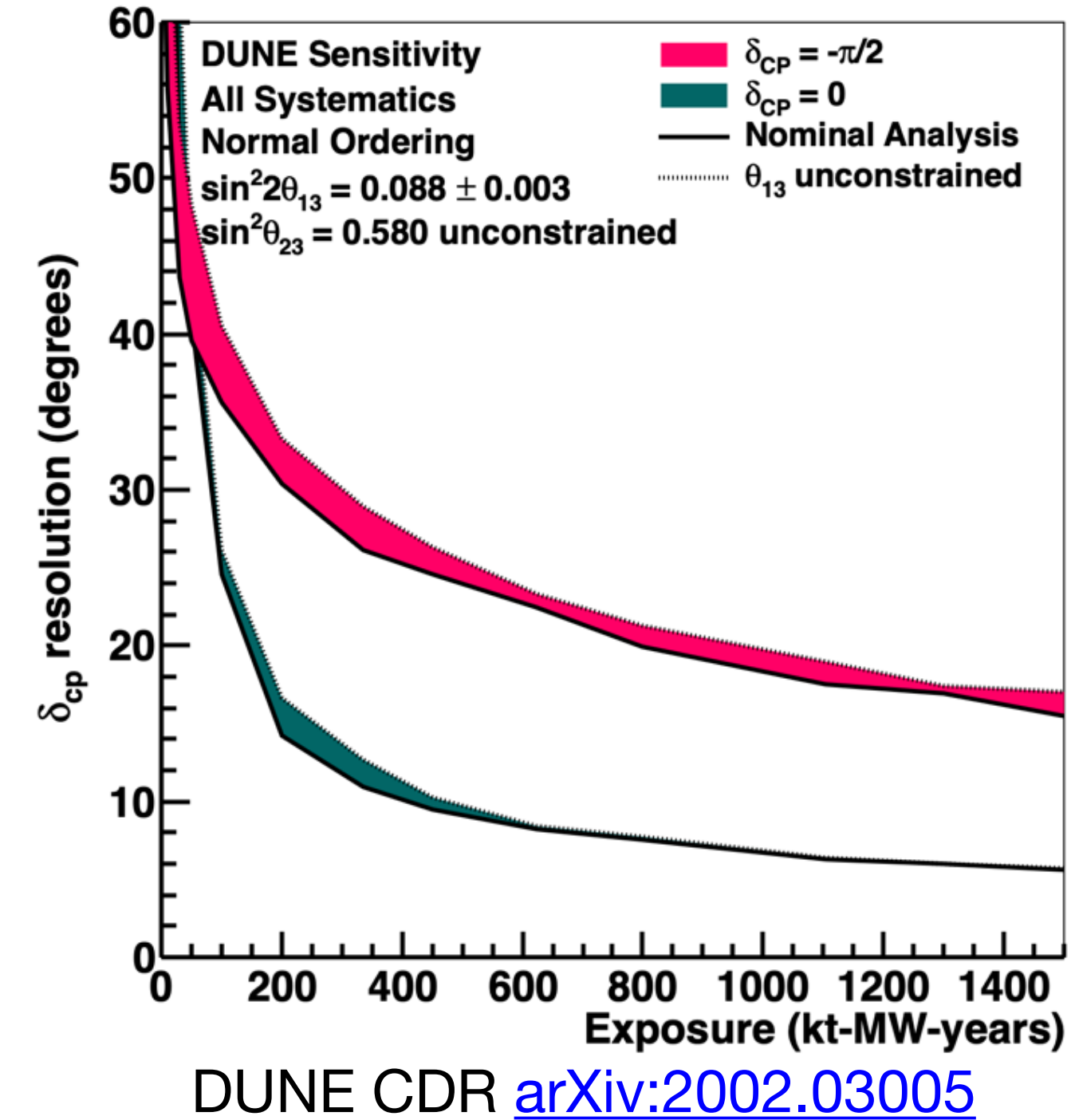
	$\sin^2 \theta_{23}$	Atmospheric neutrino	Atm + Beam
Mass ordering	0.40	2.2 $\sigma$	→ 3.8 $\sigma$
	0.60	4.9 $\sigma$	→ 6.2 $\sigma$

Sensitivity to CPV is little affected if we add atmospheric  $\nu$   
 → MO prior knowledge not really required to explore  $\delta_{CP}$

# $\delta_{CP}$ measurement resolution



	$\delta_{CP} = -\pi/2$		$\delta_{CP} = 0$	
	30°	20°	15°	10°
Hyper Kamiokande	1 y	7 y	1 y	3 y
DUNE	5 y	12 y	5 y	8 y



Precision = sensitivity to matter-antimatter models

→ HK will quickly reach precision on  $\delta_{CP}$  of 30°(15°) for  $\delta_{CP} = -\pi/2$  (0)

For the ultimate precision on  $\delta_{CP}$  it will be important to further reduce systematics uncertainties w.r.t. T2K (ND280 Upgrade + IWCD)



# Proton decay

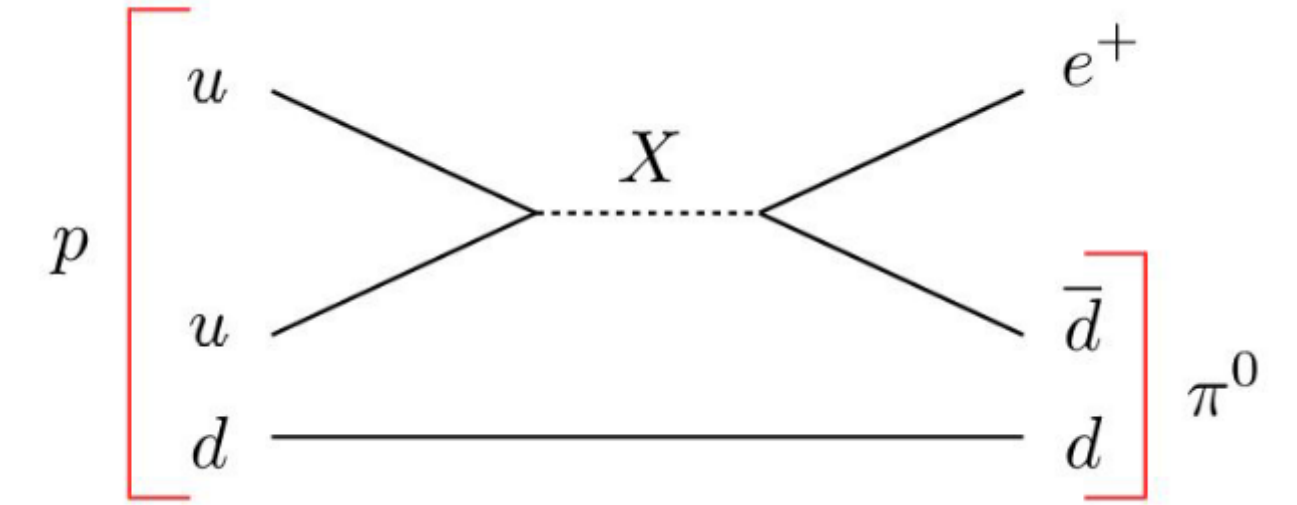
Motivated by Grand-Unification Theories

HK will have the best limit on  $p \rightarrow e^+ \pi^0$  for bound protons

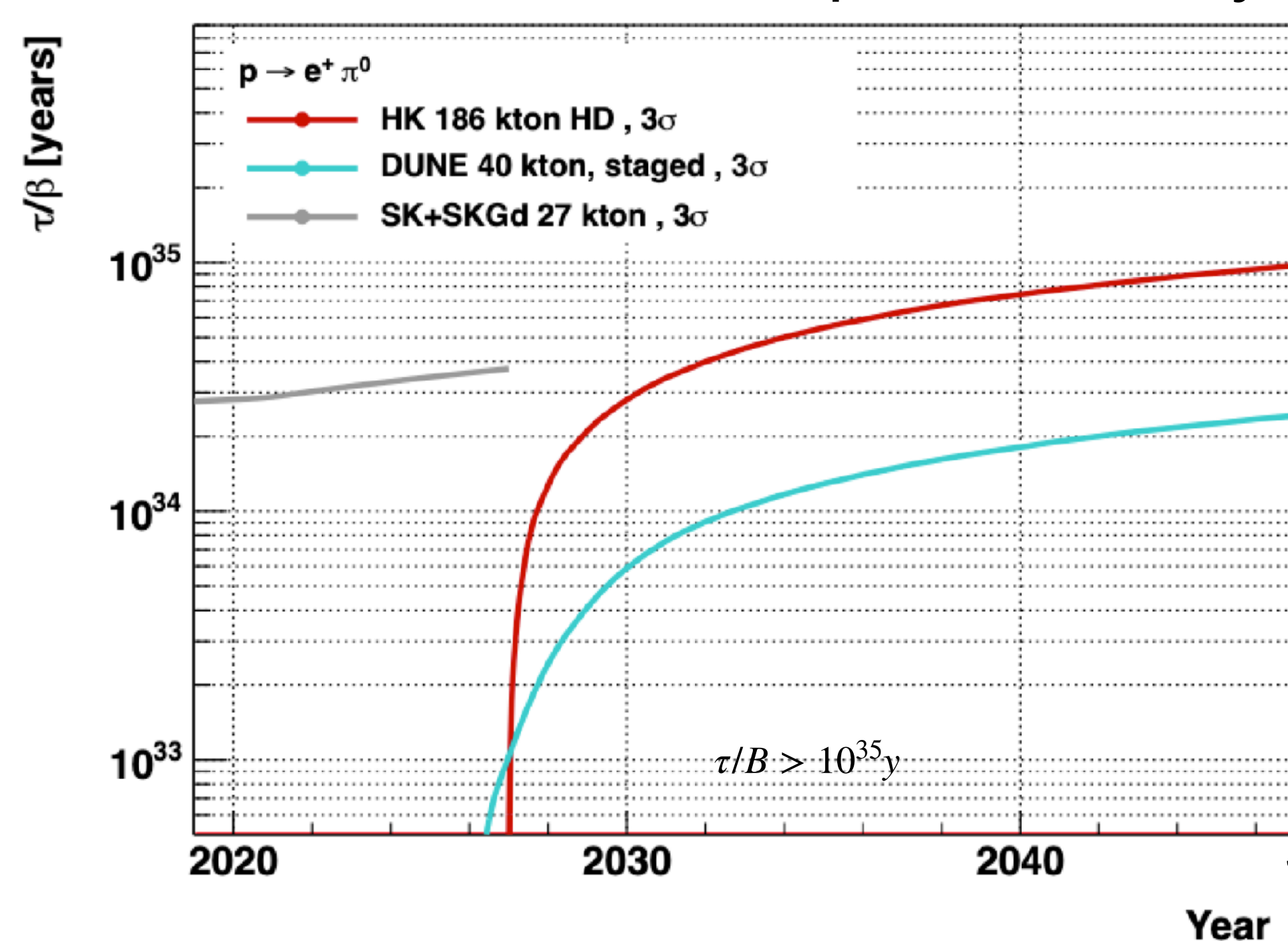
→ about 1 order of magnitude better than current limits

Thanks to its huge mass, HK will also have leading sensitivity to channels with invisible particles ( $p \rightarrow \nu K^+$ )

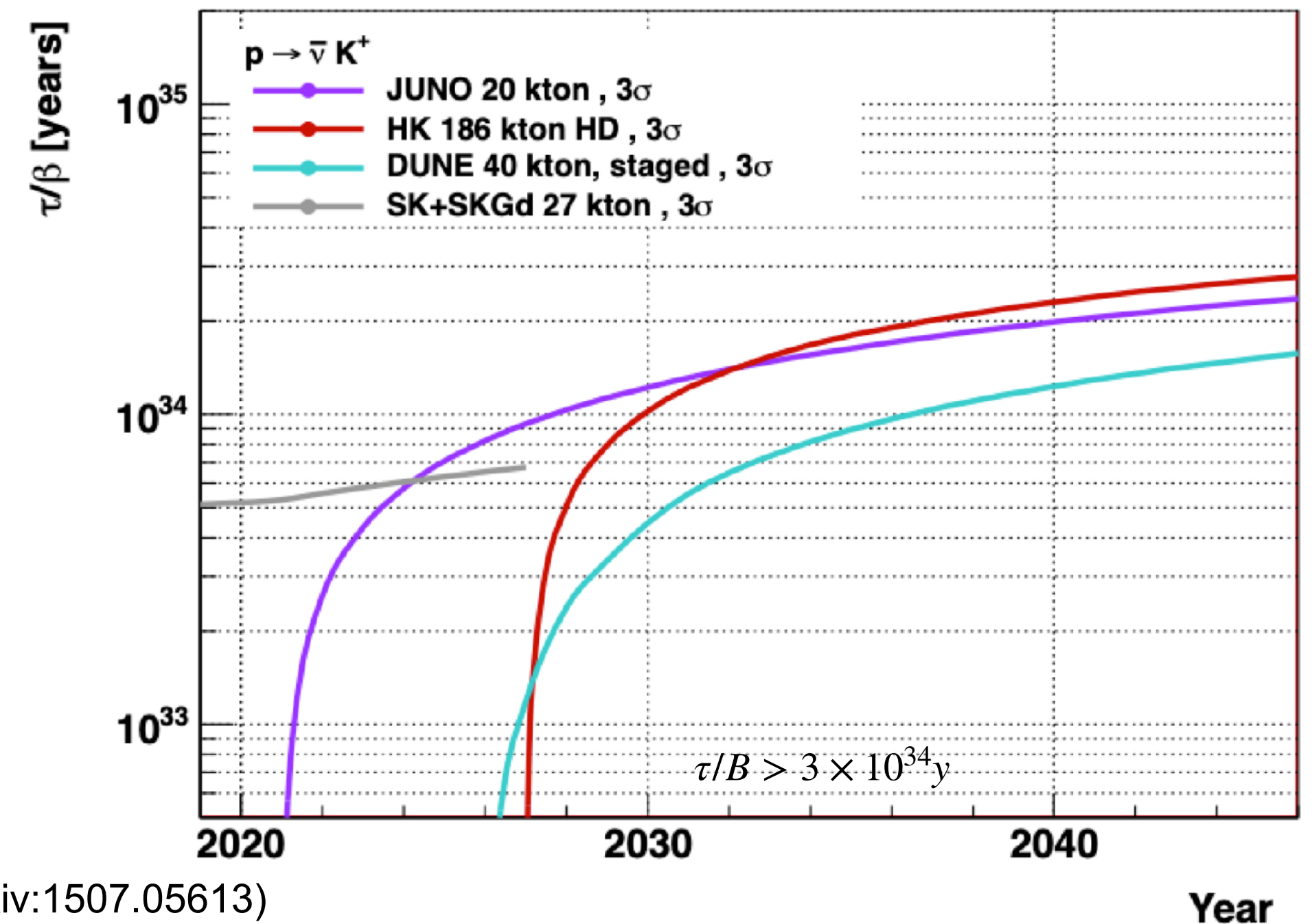
HK is sensitive to free proton decay



Phys. Lett. B 233 (1-2) 178-182



JUNO: J. Phys. G 43 (2016) 030401 (arXiv:1507.05613)  
DUNE: FERMILAB-PUB-20-025-ND (arXiv:2002.03005)



# Astrophysical neutrinos

## Supernova neutrinos

[arXiv:2101.05269](https://arxiv.org/abs/2101.05269)

Increase by  $\sim 10$  in stat sensitivity w.r.t. SK

SN1987A type  $\sim 2500$  events

Galactic center:  $\sim 50000+$  events

Direction ( $1^\circ$ @10kpc)  $\rightarrow$  triangulation

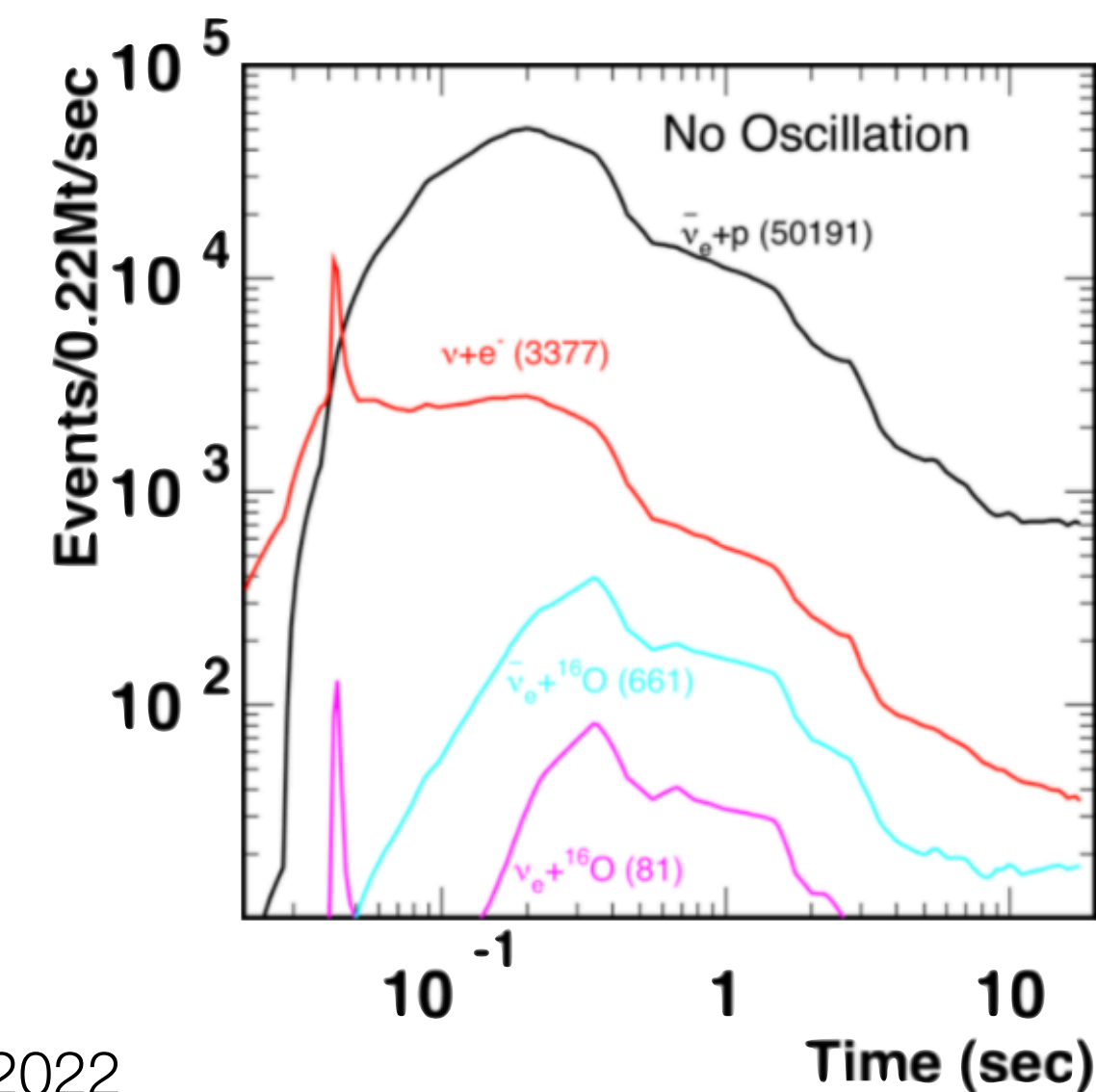
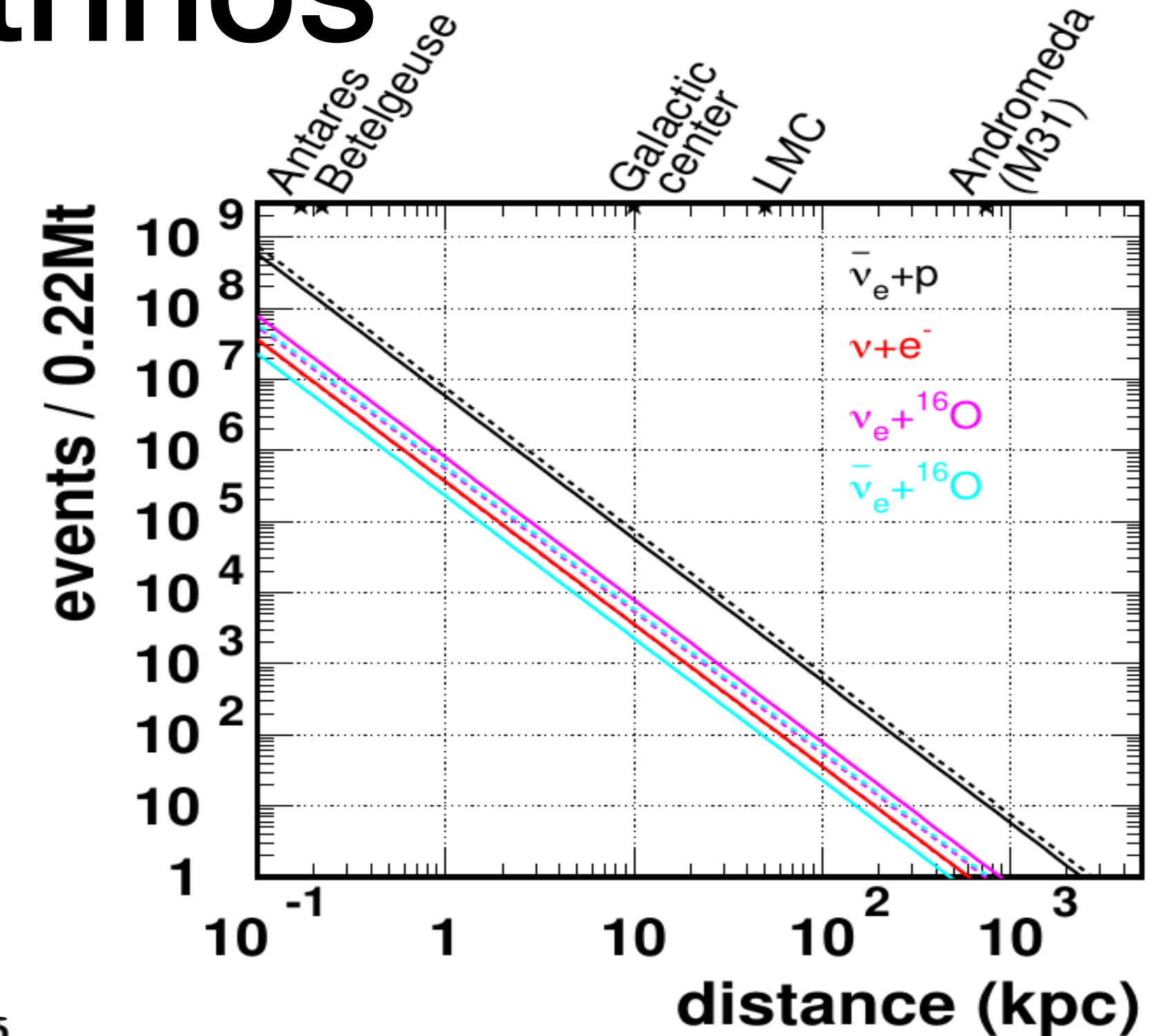
Time profile: collapse models

Since HK is sensitive to about 1 Mpc SN,  
detection of SN explosion expected every 10 y

## Gravitational waves sources

Nearby (10 Mpc) neutron star mergers

$\rightarrow$  Unique multi-messengers observatory





# SN-relic neutrinos (DSNB) in HK

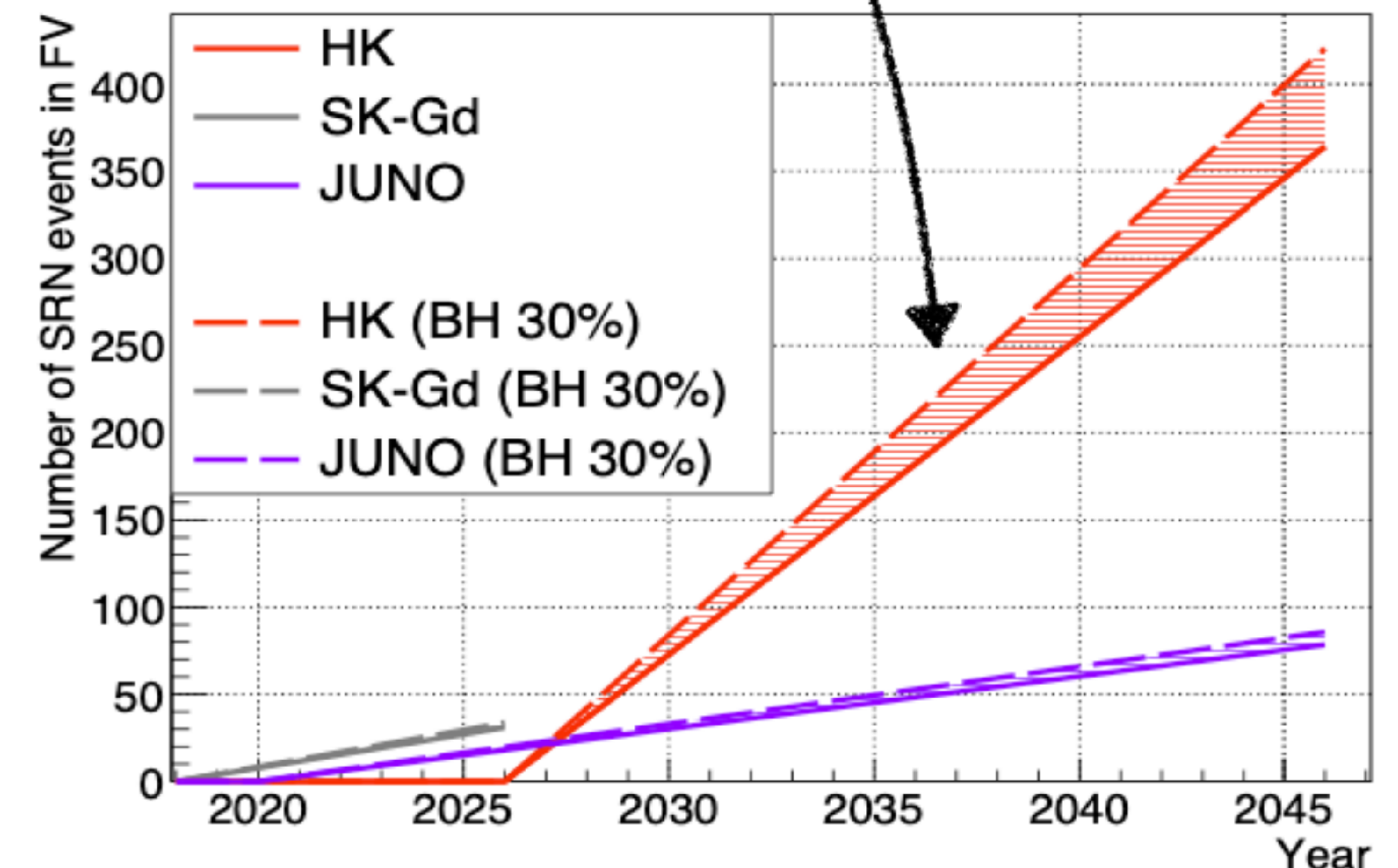
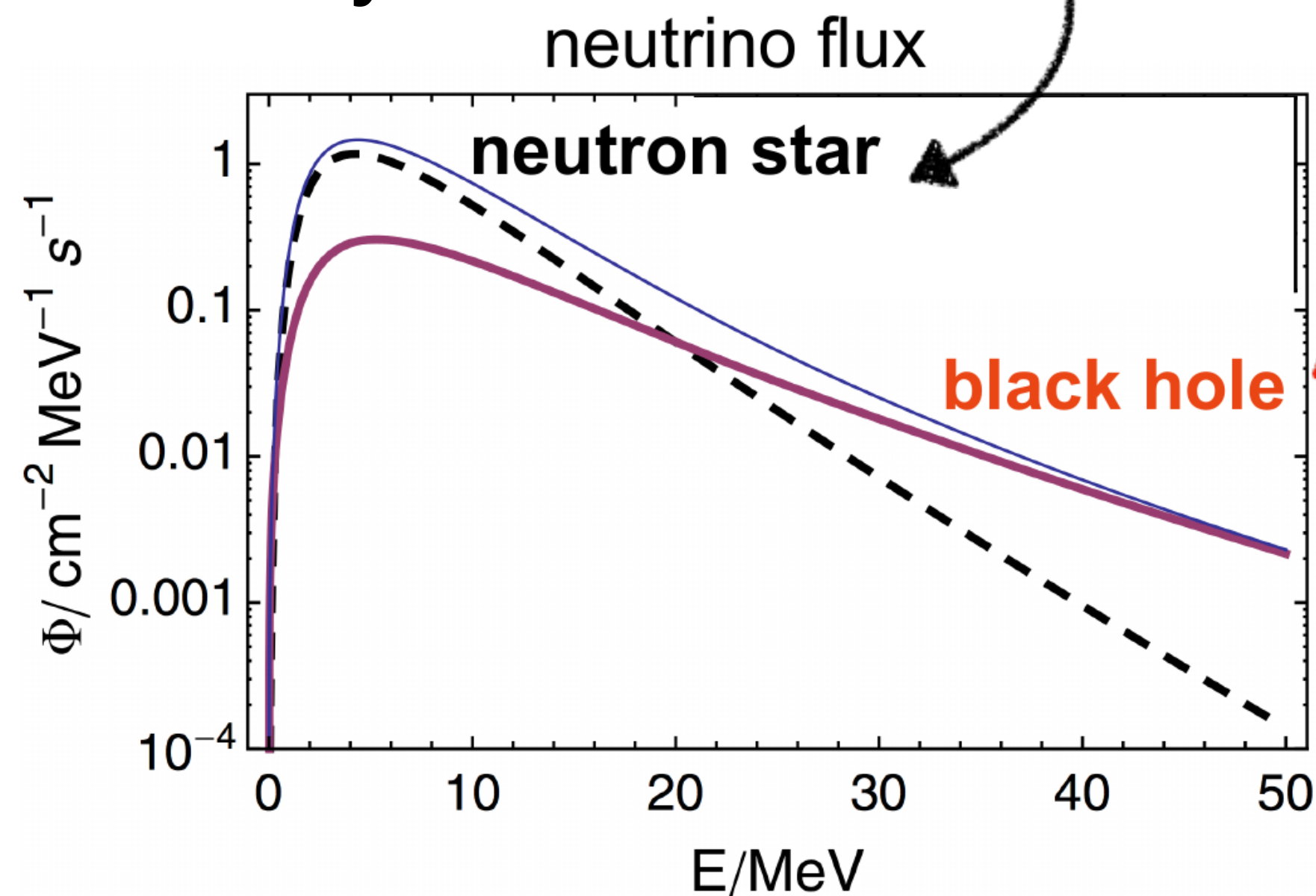
SN-relic neutrinos (SNRv) offer new constraints on cosmic star history

→ Could be first detected by SK-Gd

→ The spectrum will be determined by HK

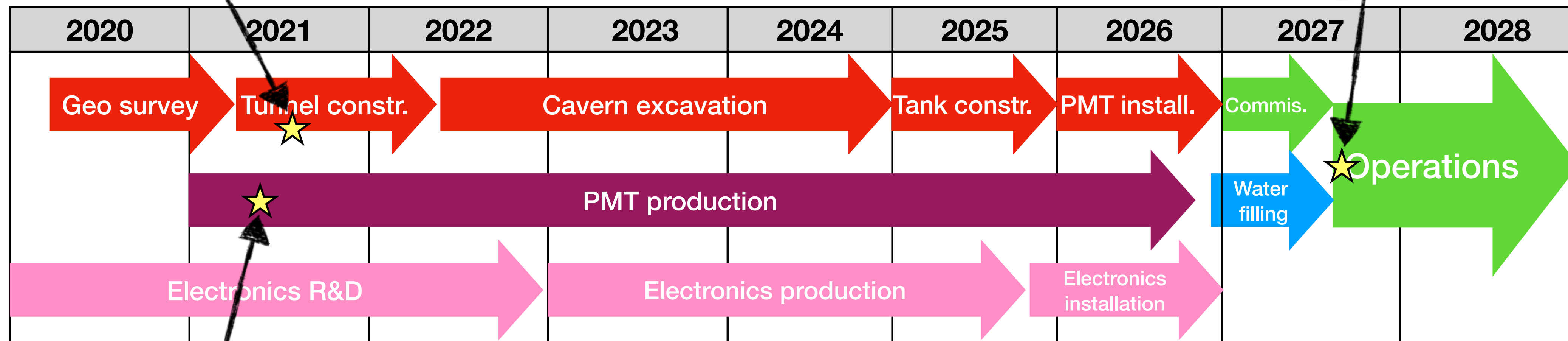
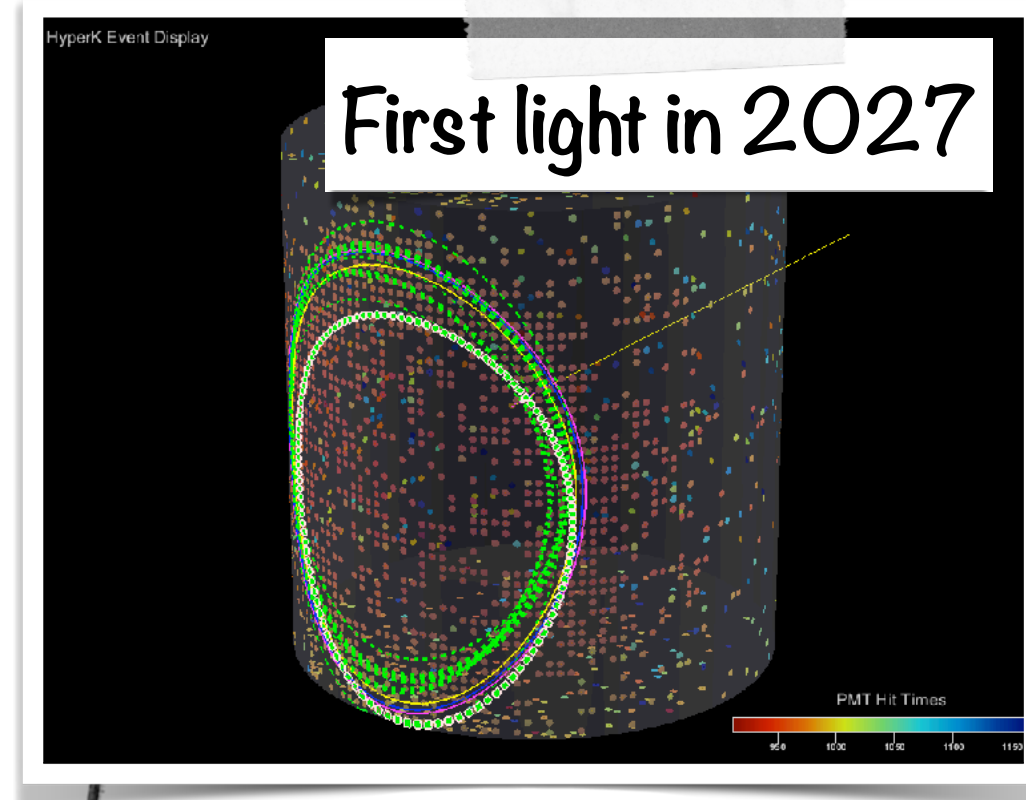
Impact of redshift: low energy  $\leftrightarrow$  probe older stars

Sensitivity to neutron star vs **black hole** formation





# HK schedule



**Strong engagement of Japan:** ~500 M\$ for construction  
Expected from other countries: ~100 M\$

International contributions are nearly formalized  
(proposed IN2P3 contributions presented by Benjamin)



# Construction is on schedule

Entrance yard finished (Aug. 2020)



New main building finished (Spring 2022)



Groundbreaking ceremony  
(May 2021)



Tunnel entrance (June 2021)



Survey completed ✓

Excavation on-going 

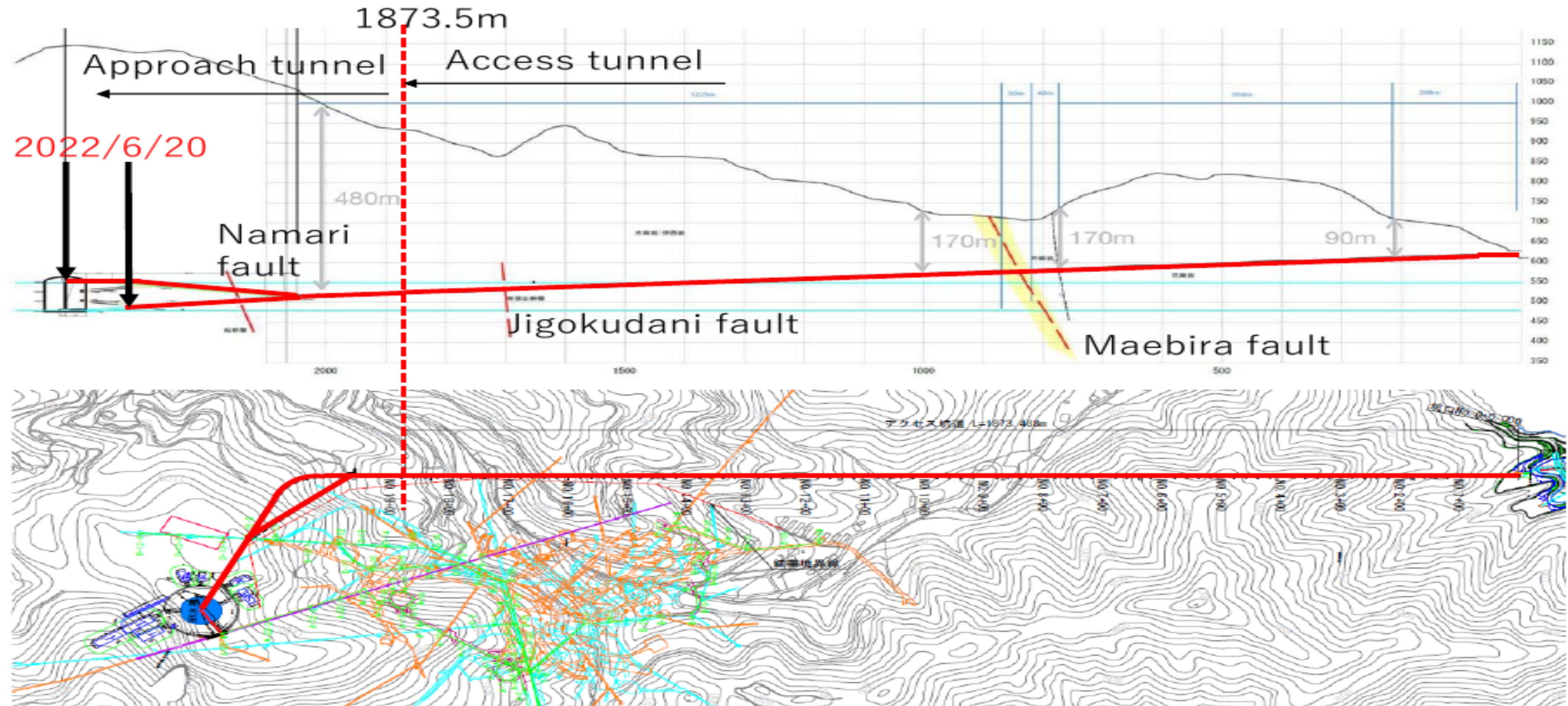
New main building ✓

Everything on track! ✓



# Recent progress: access tunnel excavation

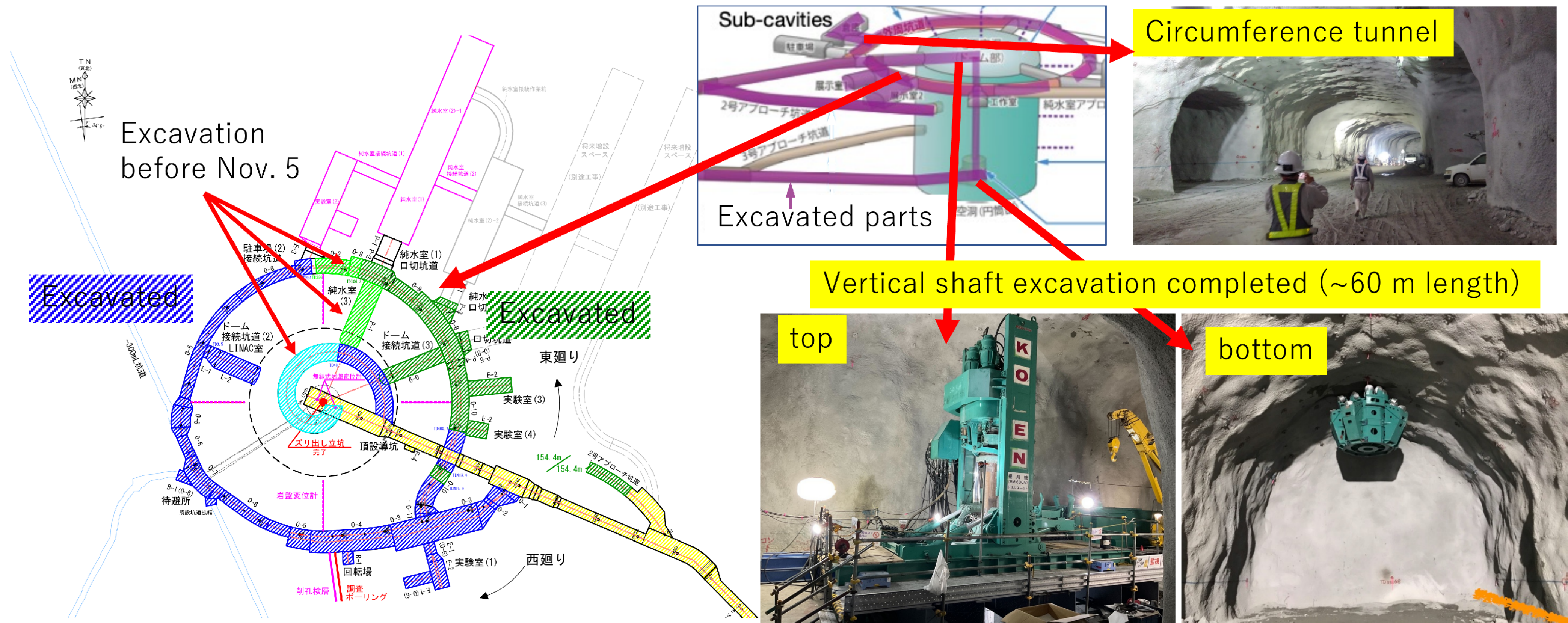
## Tunnel excavation overview



The dome center was reached on 24<sup>th</sup> of June 2022!



# Recent progress: cavern excavation

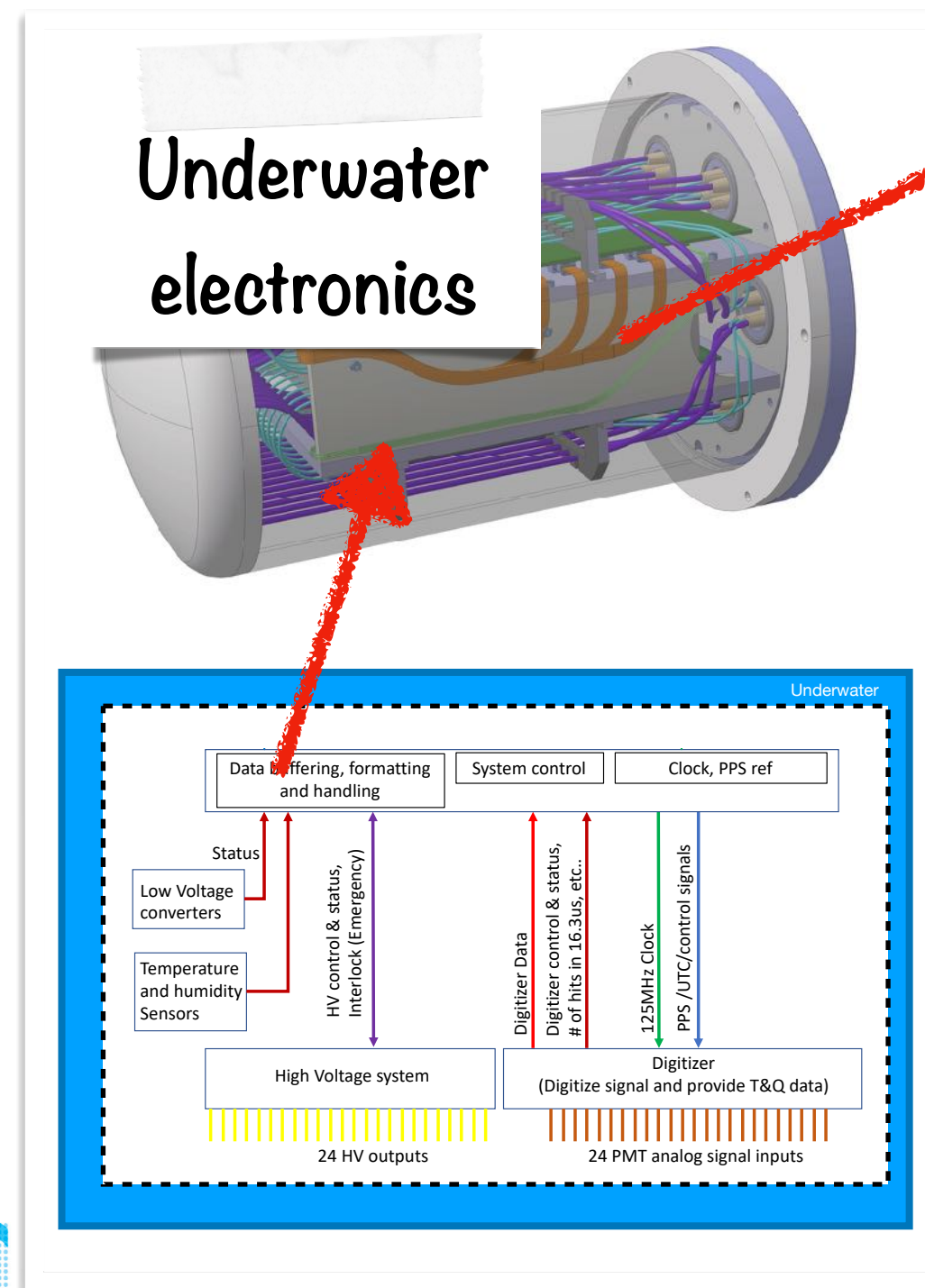
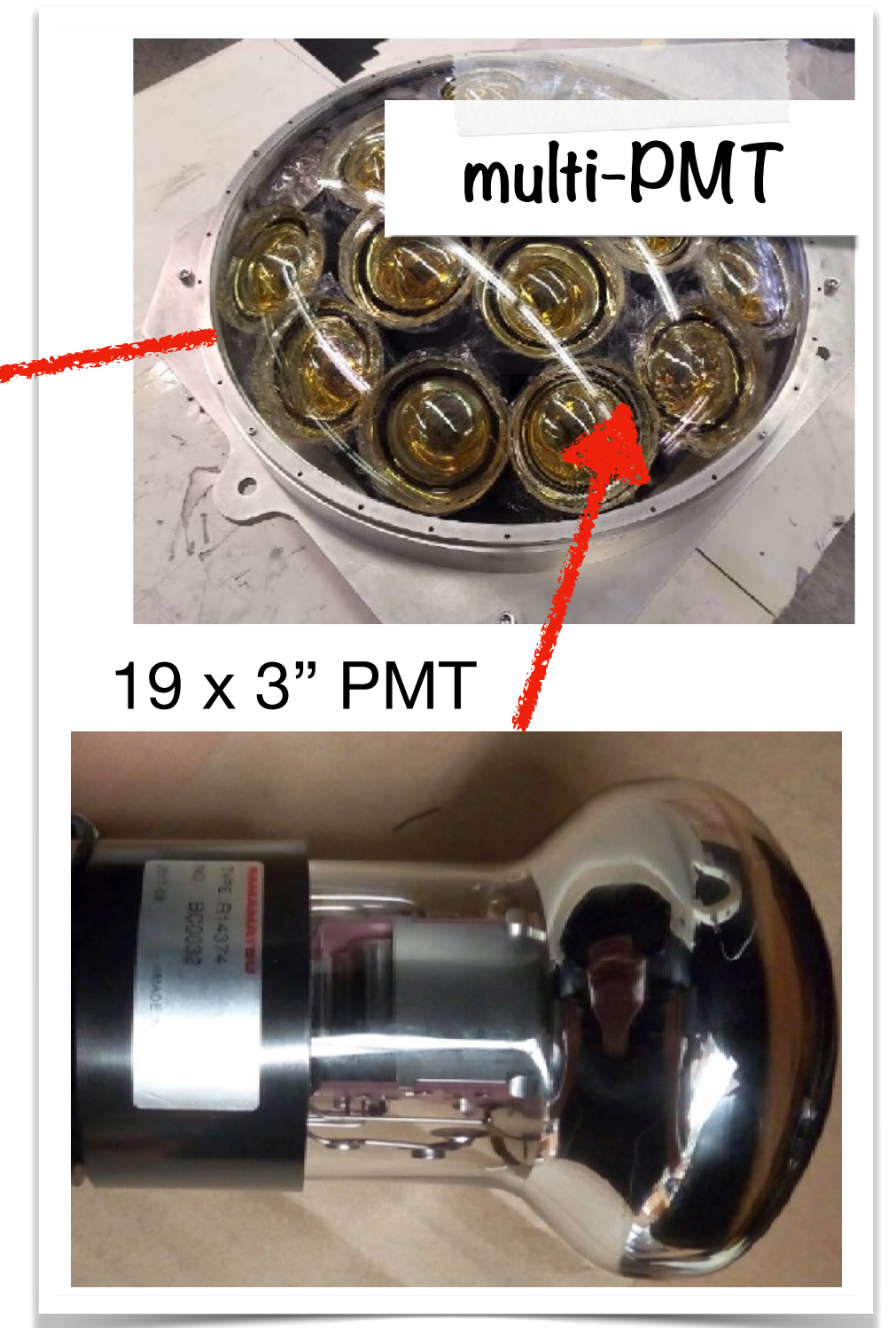
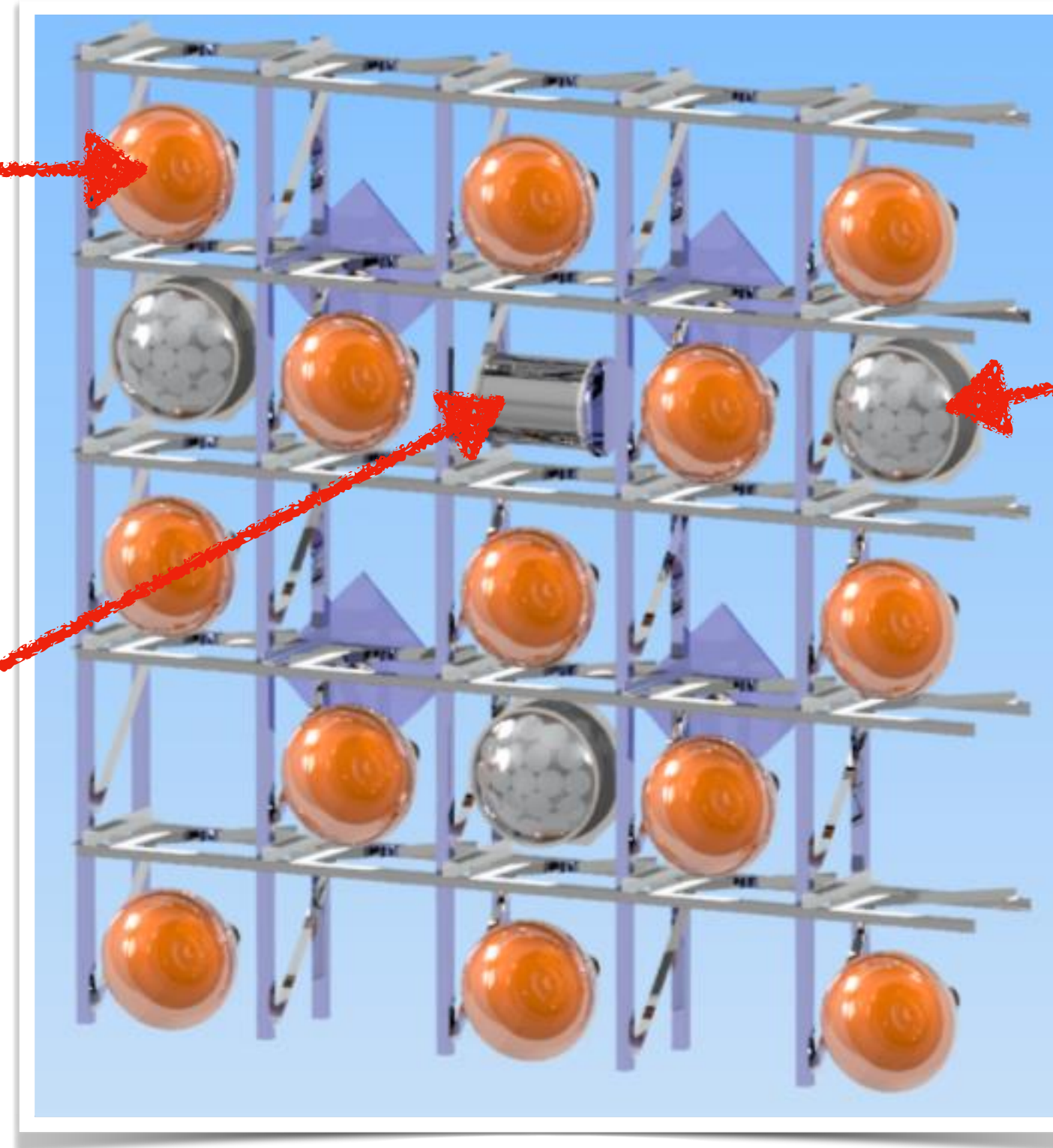
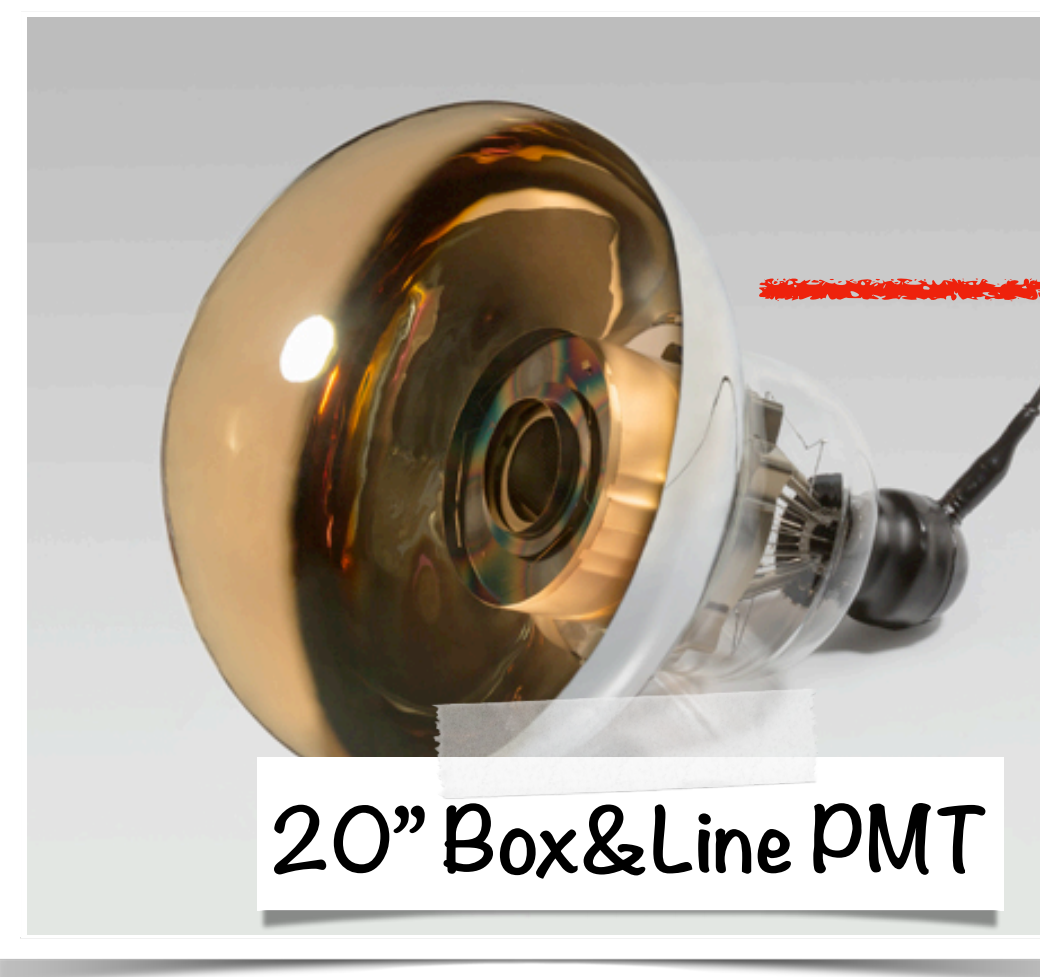


- The excavation work is on schedule. In Oct., a spiral tunnel towards the top of the dome is excavated.
- Ready for excavating the top of the dome in Nov.

**Vertical shaft excavation completed on schedule!**



# 20" PMTs, mPMTs and readout



Inner detector composed of

- 20k+ 20" PMTs (Hamamatsu R12860)
- ~1k mPMTs (19 3" R12199-02 PMTs)

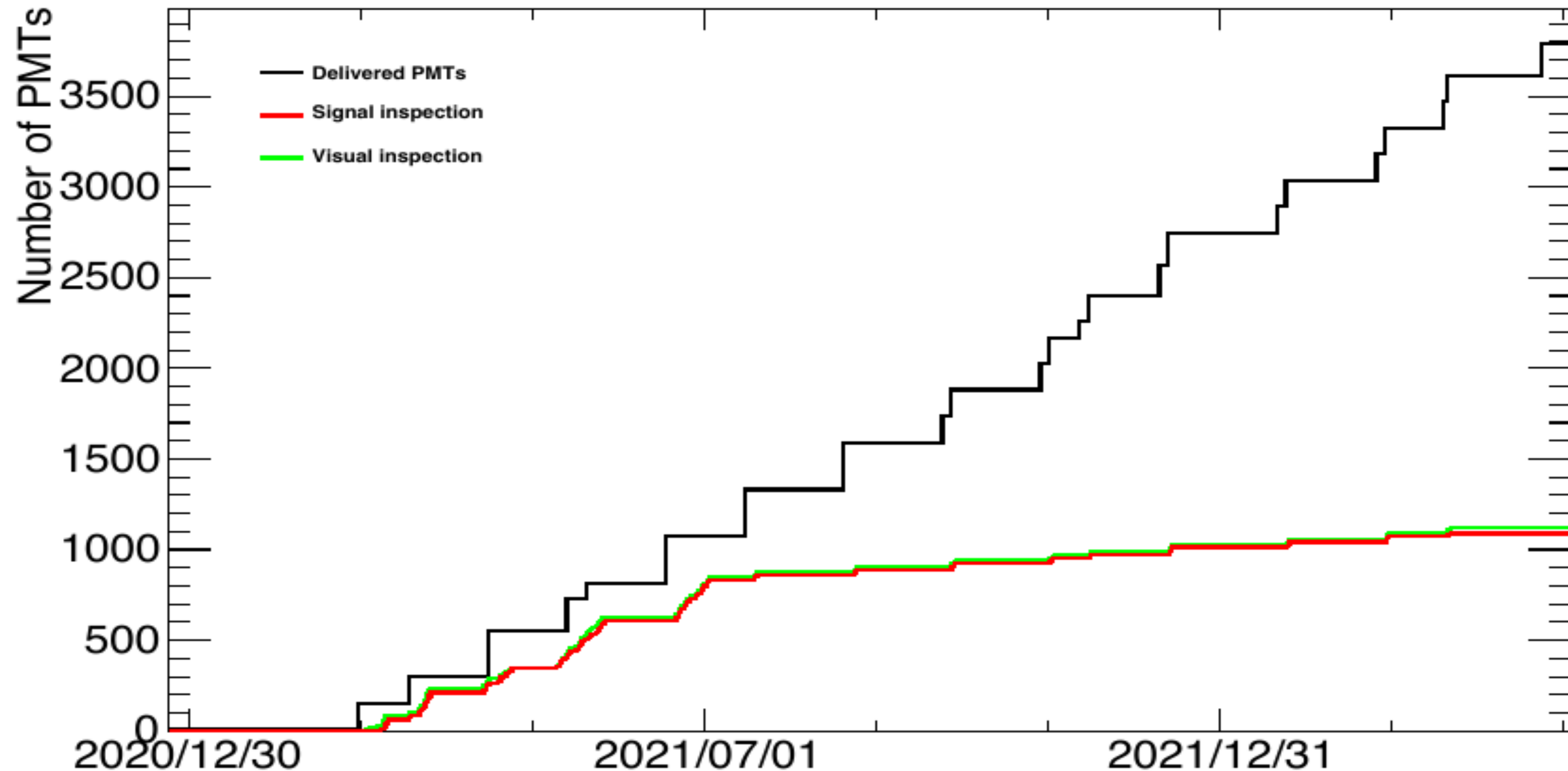
**NEW**

→ Better signal-to-noise ratio, directionality, timing

**Size of detector requires an in-water electronics**



# Recent progress: 20" PMT production



> 3700 20" PMTs already produced (at the beginning the signal and visual inspection was done for almost all PMTs). Since July 2021 the inspection rate is 10%.

# Intermediate conclusions

Hyper-Kamiokande has a vast and rich physics program including:

- Precision study of  $\nu$  oscillations (fast CP-violation discovery,  $\delta_{CP}$  measurement, etc)

- Rare events observatory (e.g. proton decay)

- Multi-messenger astrophysics

→ **Impressive and quick discovery potential!**

**Construction started and on schedule → Operation will start in 2027**

Natural continuation of our involvement in Japanese  $\nu$  program (T2K/T2K-II and SK)

**Upgraded ND280 will be used as the HK near detector starting from day 1**

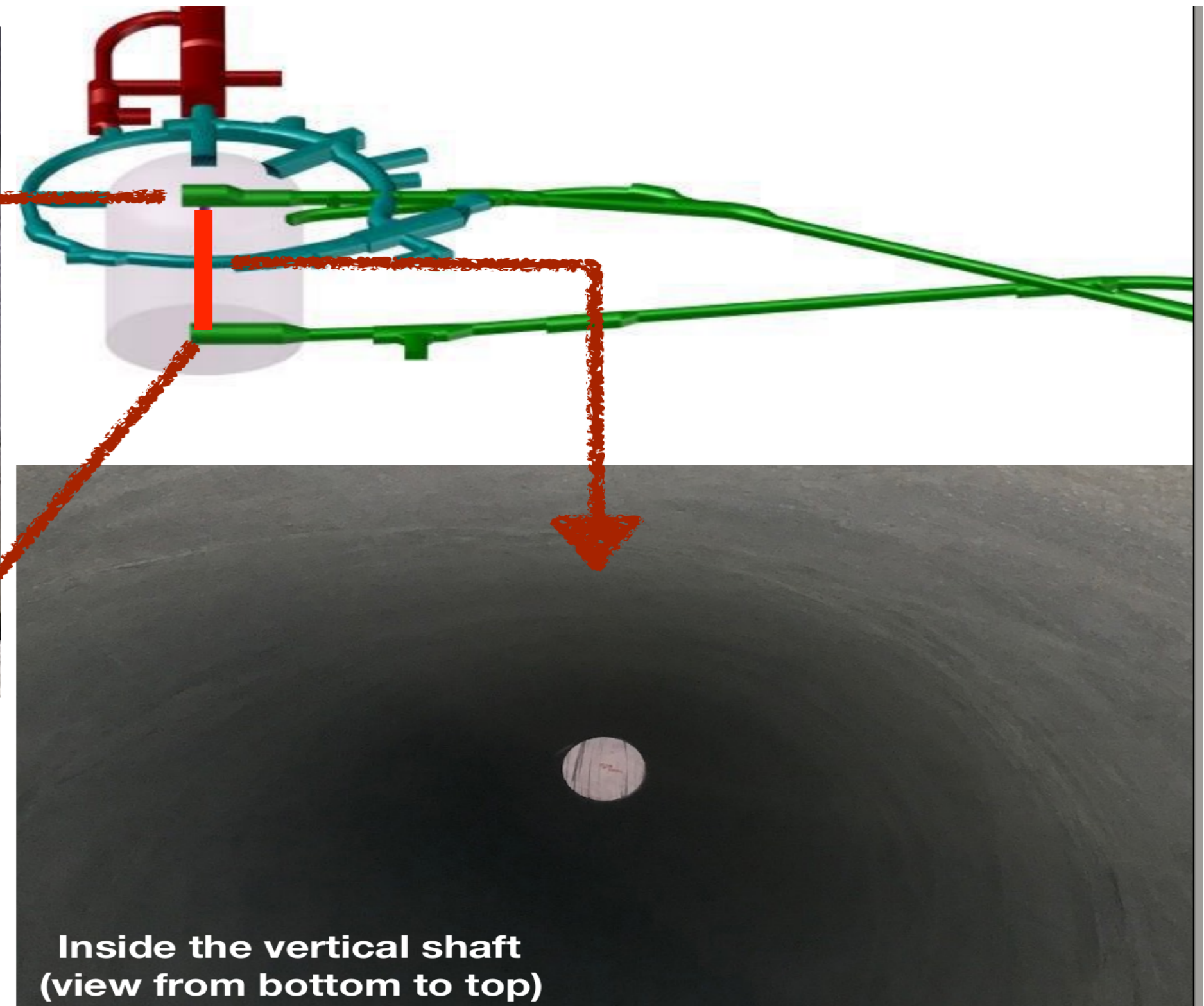
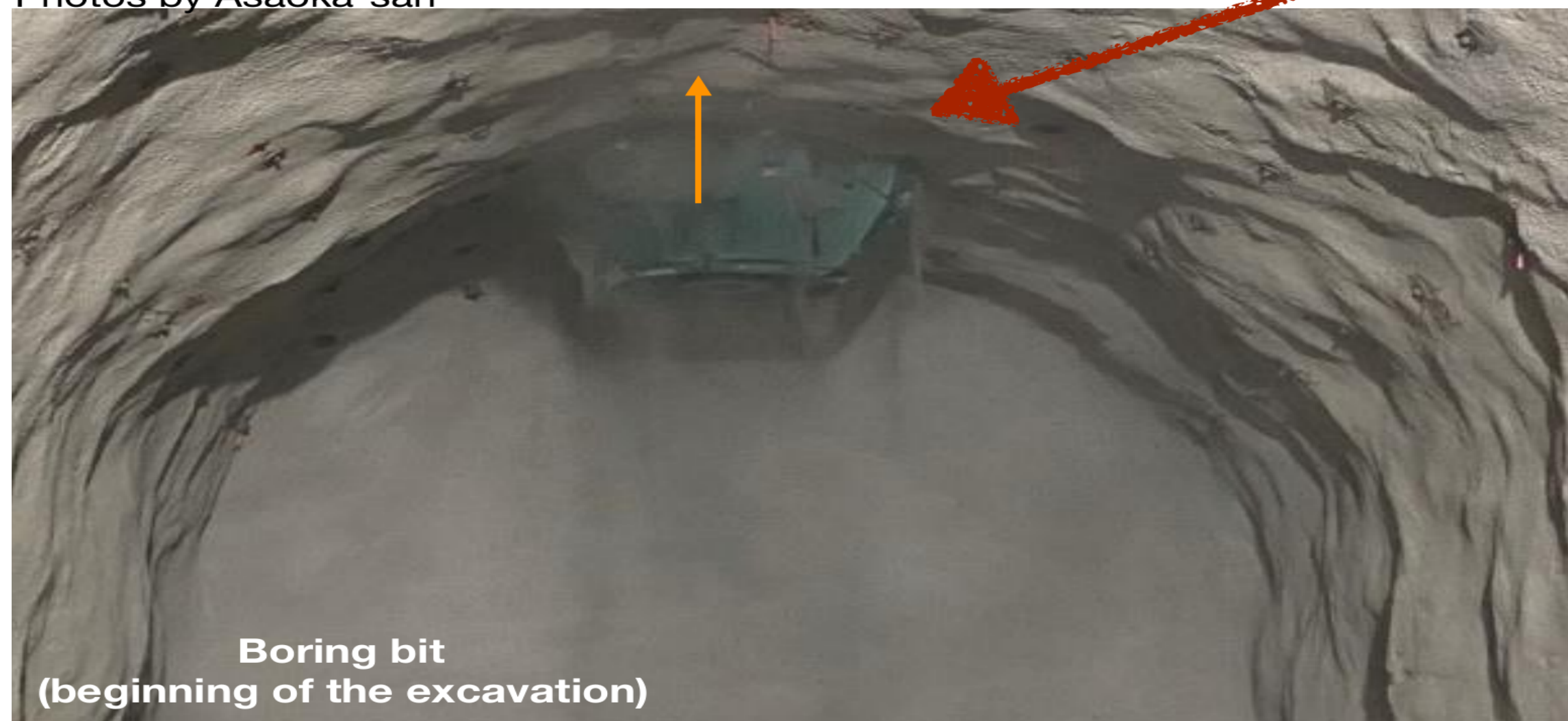
**Unique program of world-leading measurements and discoveries up to ~2040**



# Backup slides



# Recent progress: cavern excavation

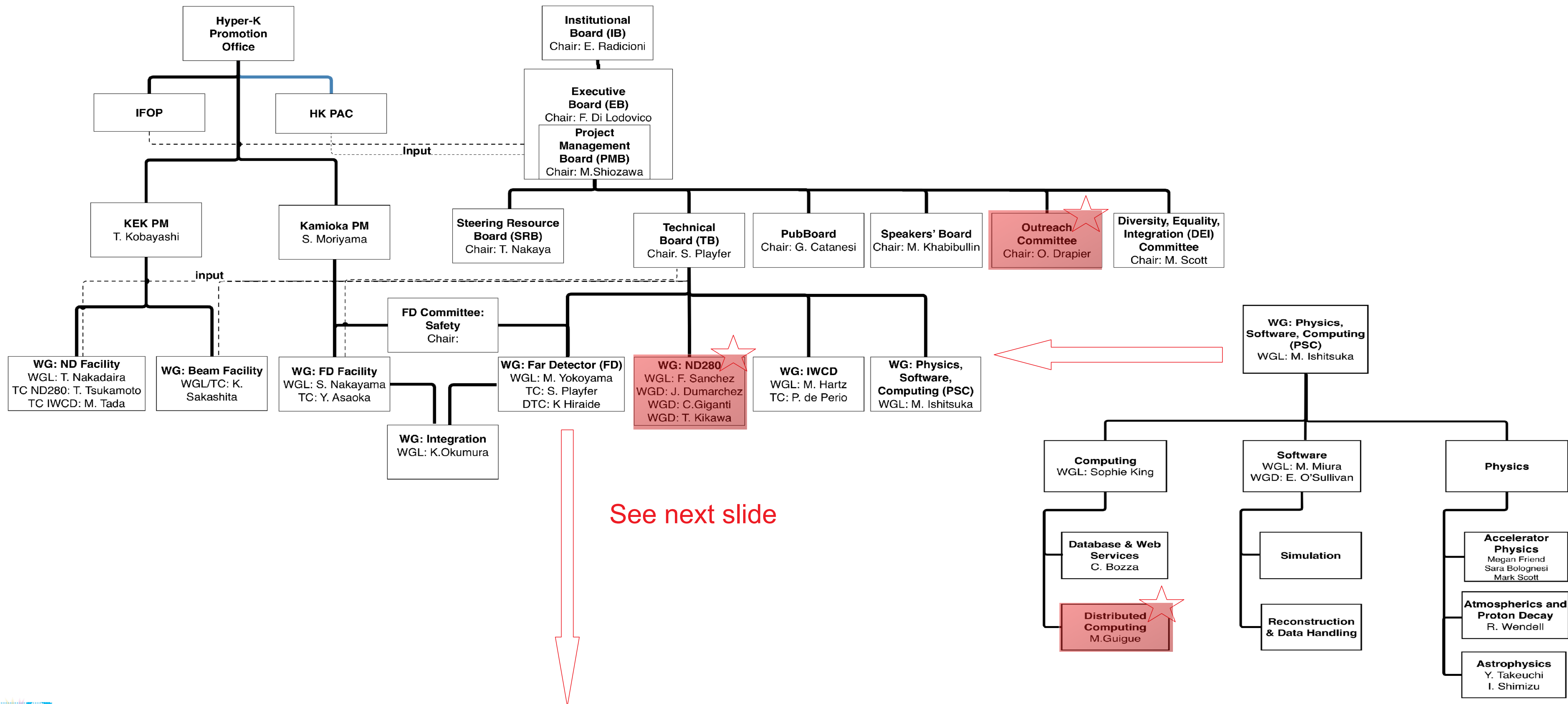


- The excavation of the vertical shaft completed on schedule
  - Vertical shaft:  $\Phi 3.4\text{m}$ , 63m-long
- Collect geological data with borehole camera
  - No obvious geological defective throughout the vertical shaft
- No heavy metal content in the excavated rock

**Vertical shaft excavation completed on schedule!**



# HK organizational chart

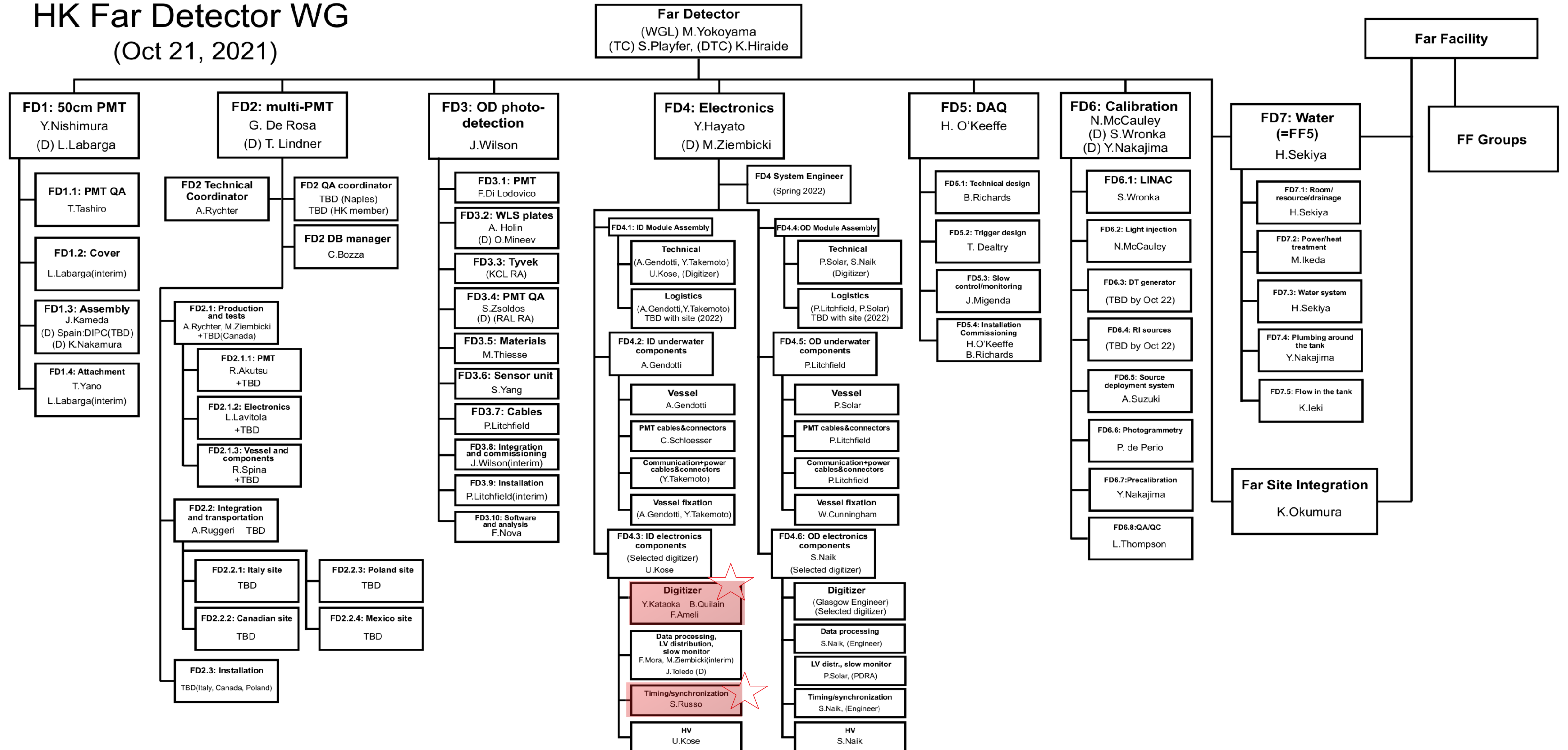


See next slide



# Far detector Working Groups

## HK Far Detector WG (Oct 21, 2021)

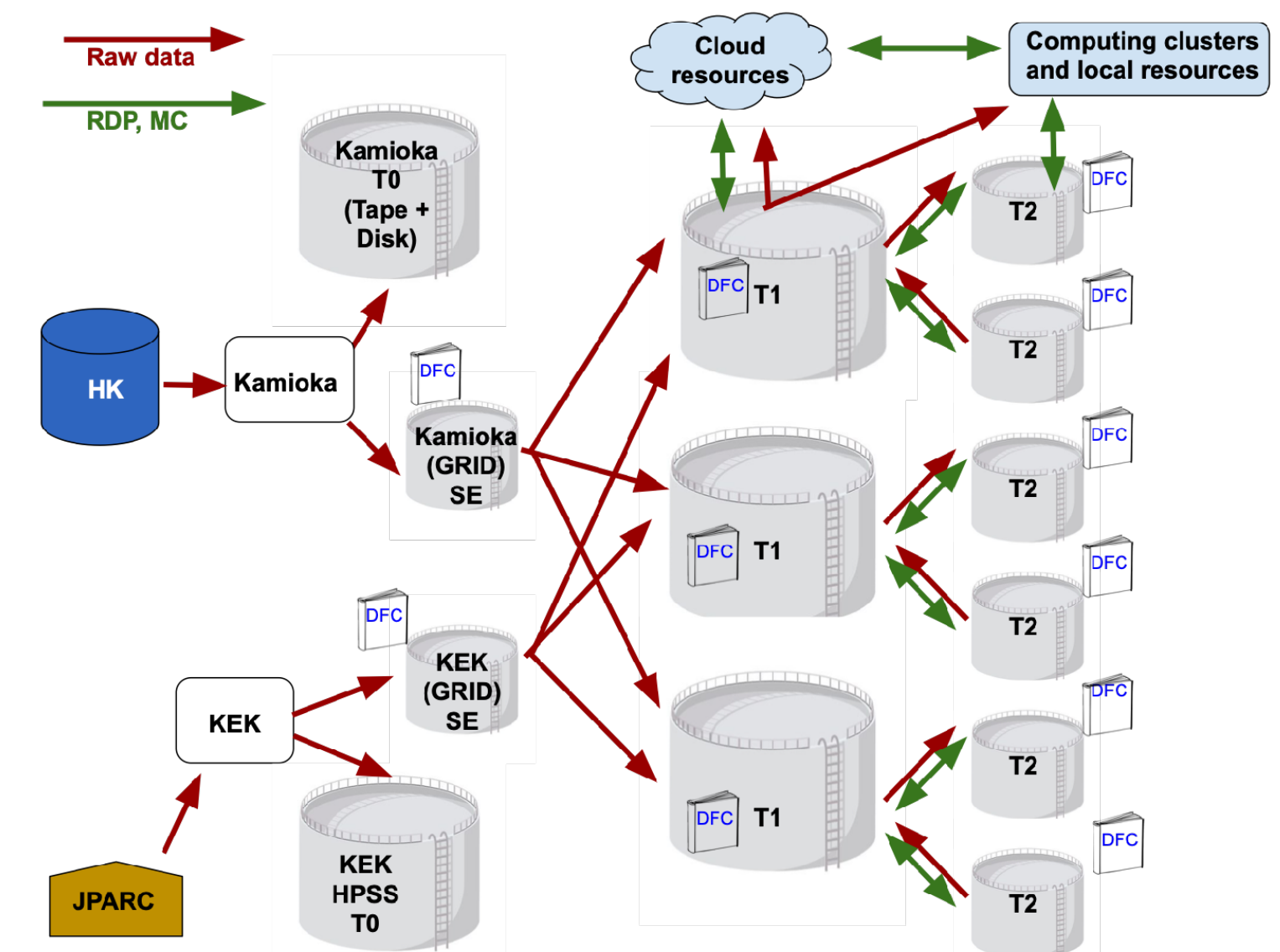




# IN2P3-CEA technical contributions to HK



- **ND280 Upgrade and maintenance**
  - Super-FGD electronics
  - High-Angle TPCs electronics & readout
- **Construction of HK far detector**
  - Front-end electronics & timing system
- **International computing effort**
  - CC-IN2P3 as T1 for HK





# SWOT

## **Strengths:**

- Well-known exceptional Water Cherenkov technology
- Use of existing neutrino beam and near detector complex built for T2K, thus saving large amount of money for the long-baseline program and reducing systematics uncertainties from the first day of the experiment.
- Construction budget for Hyper-Kamiokande have been allocated by Japanese government in 2019 with a budget profile that will allow to start the experiment in 2027.
- Leading roles of French groups and acquired expertise in the ongoing T2K experiment

## **Weaknesses:**

- Small groups at LLR and LPNHE. Mitigated by the large overlap in terms of physics case, technologies and tools between T2K, SK and HK.
- Hyper-Kamiokande is an IN2P3 R&D project since 2022, but not yet an IN2P3 master project, undermining our visibility within the collaboration.

## **Opportunities:**

- Fast measurement of CP violation, before any other experiments.
- Huge target mass, making HK the most sensitive observatory for rare events in the MeV–GeV energy region.
- IN2P3 groups can build on their long standing expertise in the T2K experiment to propose strong contributions in Hyper-Kamiokande. There are still many possibilities to contribute, e.g. electronics of the outer detector, further ND280-upgrade, IWCD.
- Deployment of a similar timing system for the near and intermediate detectors without any additional R&D.

## **Risks:**

- Approval by CS-IN2P3 and allocation of IN2P3 funding is needed to capitalize on the R&D on the digitizer to make a strong contribution to the HK detector in addition to the timing system and computing.
- Since the HKROC digitizer was not selected as the primary option for the HK inner detector, there is still an uncertainty about the outcomes of this contribution.



# Financial resources

Item	Cost (M)	Partially covered with external fundings	Funding approval	Construction period	Requested fundings (M)
ND280 Upgrade	6	T2K Collaboration	2019	2019 – 2022	0.6 (obtained)
Far detector timing	0.6	ANR - INFN - CEA	2022	2023 – 2026	0.4
Communication cables	2	European countries	2022	2023 – 2026	0.2-0.4
Chip and Front-end	2.5	CEA - INFN	2022	2023 – 2026	1-2.5
Computing (CC-IN2P3)	3.8	CEA	2021	2021 – 2037	3.8
Note that costs for computing are spread over a much longer period of time (15 years).					
Total	14.9	-	-	-	~6-7.7

External and internal investments:

- 70k€ for R&D towards HK (Sorbonne Université)
- 400k€ (X) + 90k€ (IN2P3) for R&D on HKROC
- 300k€ (ANR) + 150 k€ (IN2P3) for R&D on time generation and clock distribution

**Significant efforts to acquire external fundings before asking for IN2P3 investments**



# Conclusions and outlook

Hyper-Kamiokande has a vast and rich physics program including:

- Precision study of  $\nu$  oscillations (fast CP-violation discovery,  $\delta_{\text{CP}}$  measurement, etc...)

- Rare events observatory (e.g. proton decay)

- Multi-messenger astrophysics

→ **Impressive and quick discovery potential!**

**Construction started and on schedule → Operation will start in 2027**

Natural continuation of our involvement in Japanese  $\nu$  program (T2K/T2K-II and SK)

- Upgraded ND280 will be used as the HK near detector starting from day 1

**Unique program of world-leading measurements and discoveries up to ~2040**

Identified contributions on electronics (RD4HK during 2022) and computing

Used funding from external sources (Sorbonne University, Ecole Polytechnique, ANR)

Technological decisions taken in September 2022

→ **approval by CS-IN2P3 & recognition of HK as IN2P3 master project are needed for full-scale participation in the HK experiment**



# Time Distribution costs estimations

<b>Second stage distribution</b>	Total number of PMTs	PMT per FE	Total number of FE (Time distribution nodes)	Number of nodes per distributor	Number of distributors	Distributor cost per unit	Total distributor's cost
Outer detector	13 300	72	185	16	12	3 000	36 000
Inner detector	20 000	24	833	16	53	3 000	159 000
<b>MultiPMT</b>	2 000	12	167	16	11	3 000	<b>33 000</b>

Total distributor's cost	228 000
Atomic Clock cost	54 000
GNSS recevier	13 000
Antenna + cable	4 000
Clock multiplication board	2 000
<b>Total clock ref cost</b>	<b>73 000</b>
<b>Redundant system</b>	<b>73 000</b>
<b>Grand total clock reference</b>	<b>146 000</b>
Slow Control switch	2 000
<b>Total cost</b>	<b>376 000</b>
Contingency	20%
<b>Gran total</b>	<b>451 200</b>

Time generation estimated cost: 145 k€ (LPNHE)  
Time distribution (first and second stages) estimated cost: 250 k€ (LPNHE +CEA)  
Time distribution endpoint cost: 80 k€ (INFN)

Total = ~480 Keuros



# Computing efforts in Hyper Kamiokande

Tier model similar to CERN's  
Resources and data management using DIRAC  
Software containerized and shared via CVMFS

System similar in other IN2P3 exp. (Belle-II,  
CTA, LSST...)

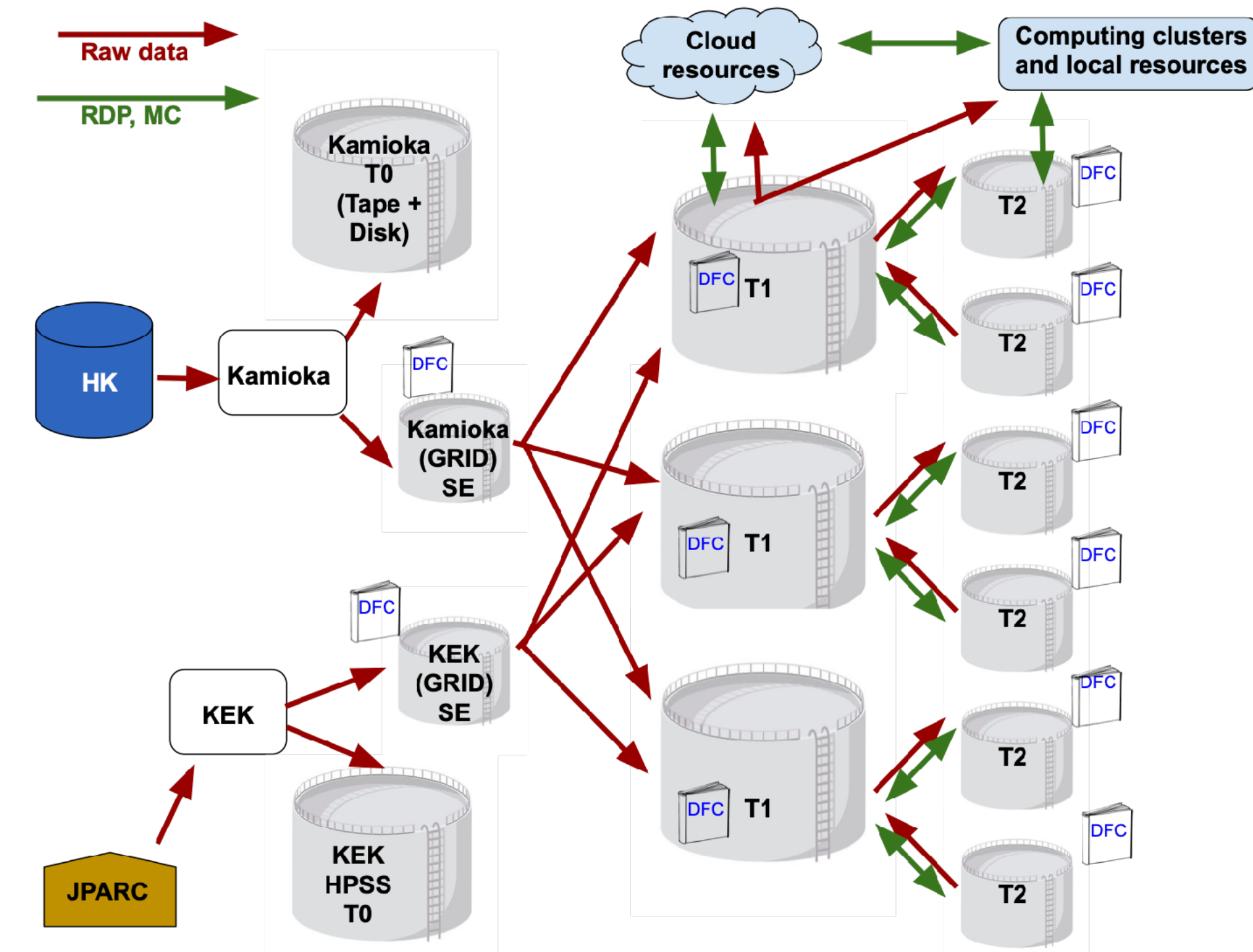
obvious synergies

First 10 years of operations:

25 PB (data + MC = mostly Far Detector)

880 MCPU.hours

(minimal with one copy of each file)





# Proposed Contribution: CC-IN2P3 as T1 site

CC-IN2P3 Tier1 for LHC (WLCG)  
infrastructure, expertise available

Low-rate data stored on tapes

Disk and CPU for productions

Database management

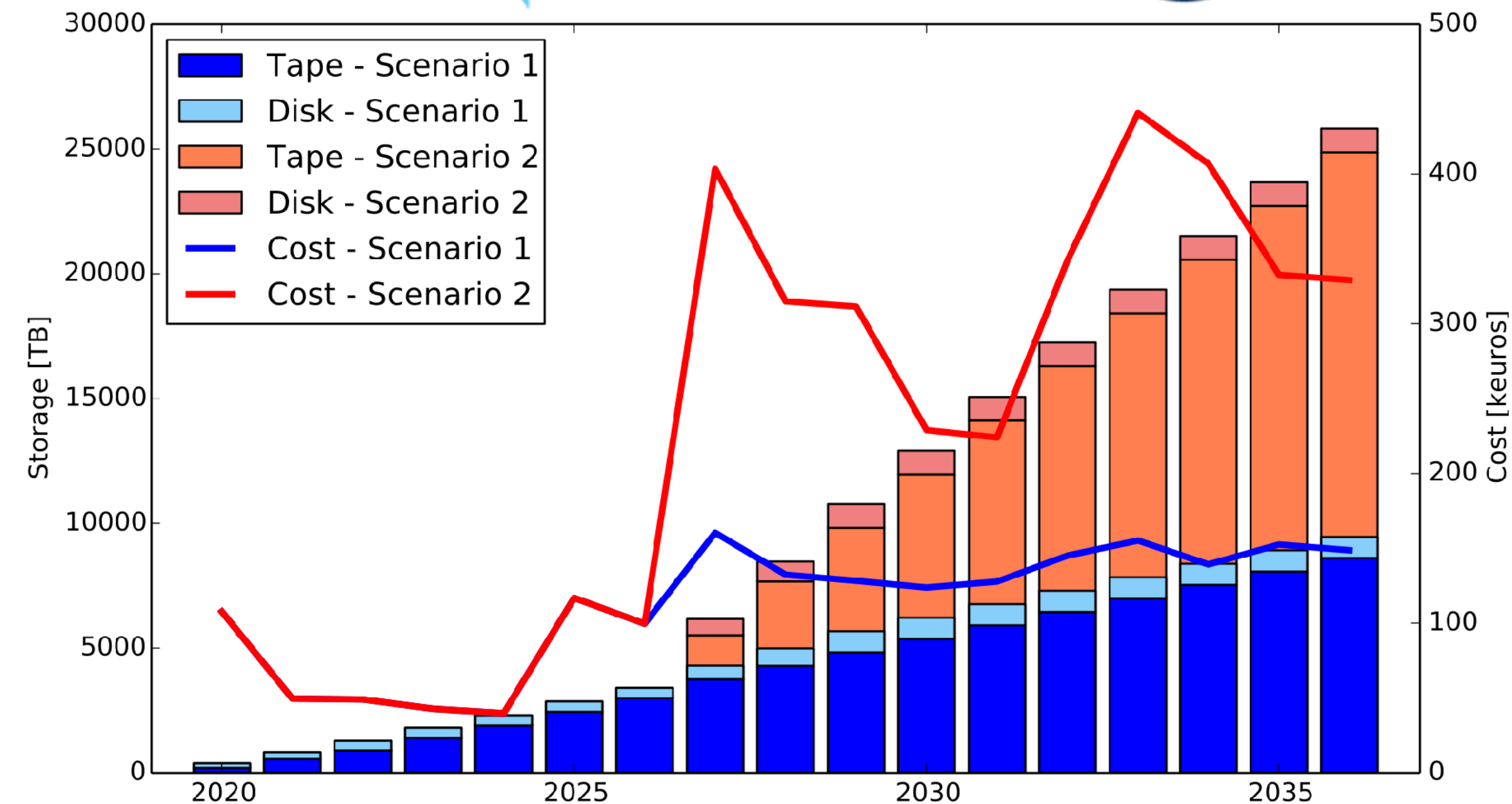
Two proposed scenarii:

1. ND280 data storage

2. Near and Far detectors data storage

First step: CC-IN2P3 as T1 site for T2K  
integration of CC as grid site into GridPP  
disk allocation and data transfer

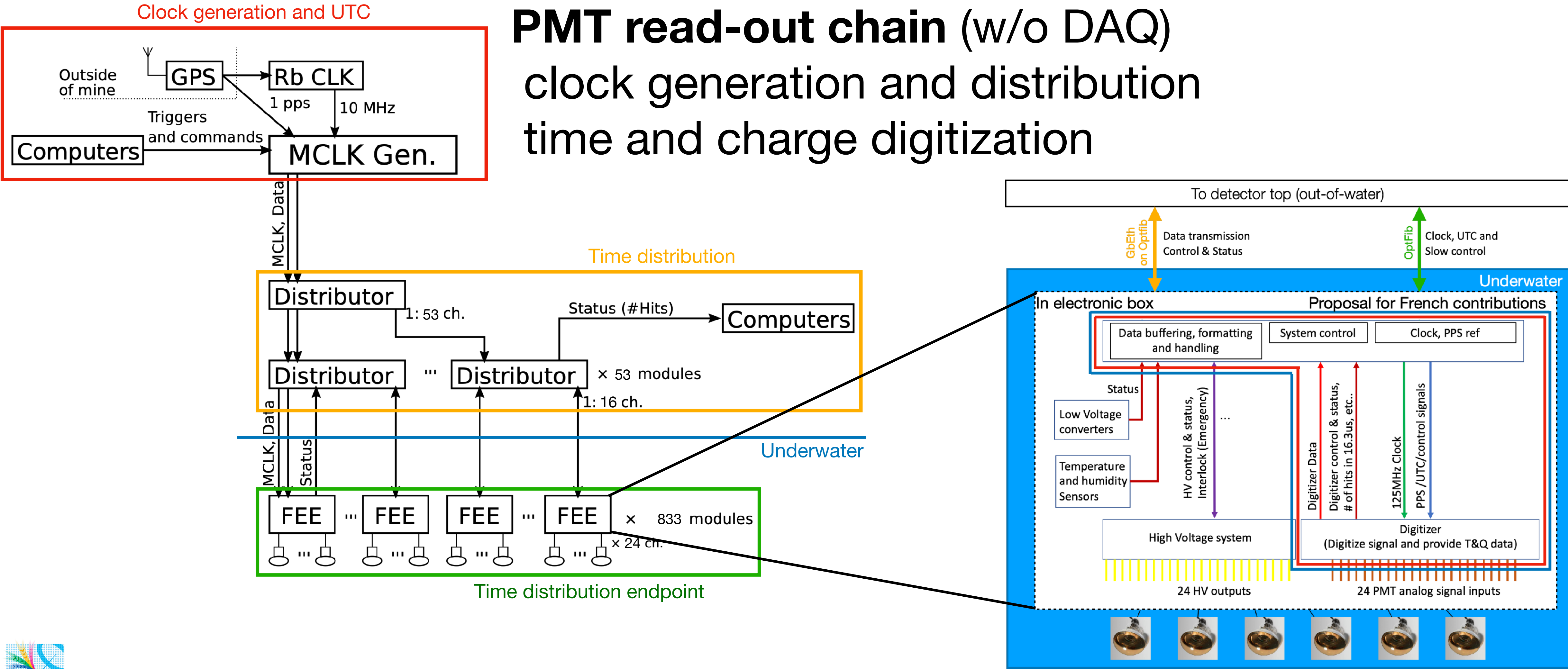
**First production for HK at CC-IN2P3 completed**





# Contribution to HK Far Detector electronics

France would develop and produce the whole **PMT read-out chain (w/o DAQ)** clock generation and distribution time and charge digitization





# HKROC: new chip and readout

## HK requirements fulfilled

# Chip and readout board for PMT inspired by HGCR0C chip

1. Large dynamic range: 3 gains / ch. → up to 2500 pC
2. Excellent charge & time resolution: (<200 ps)
3. TSMC CMOS 130nm etching
4. Reduced dead-time (<50 ns): SAR ADC sampling waveform at 40 MHz

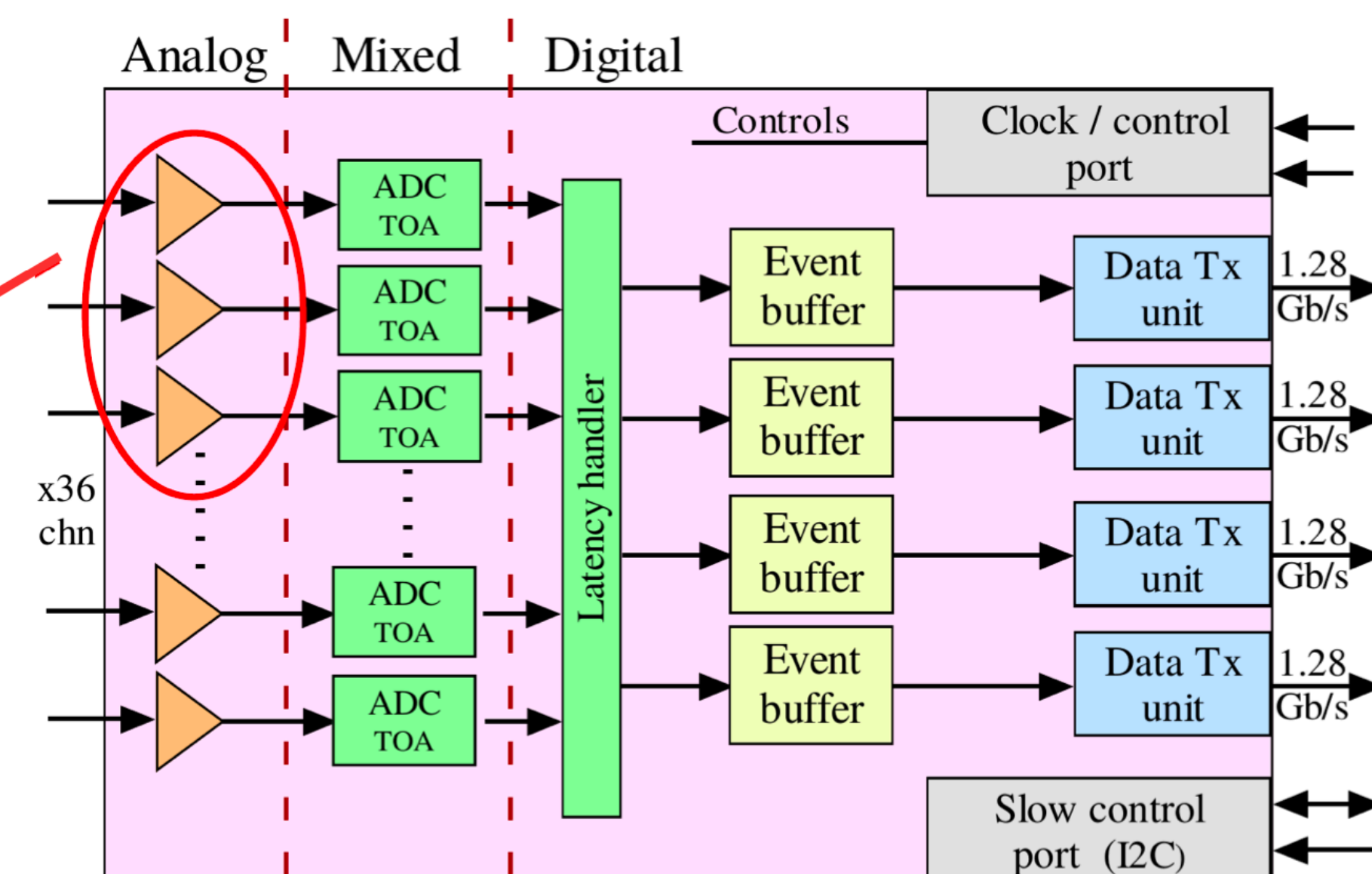
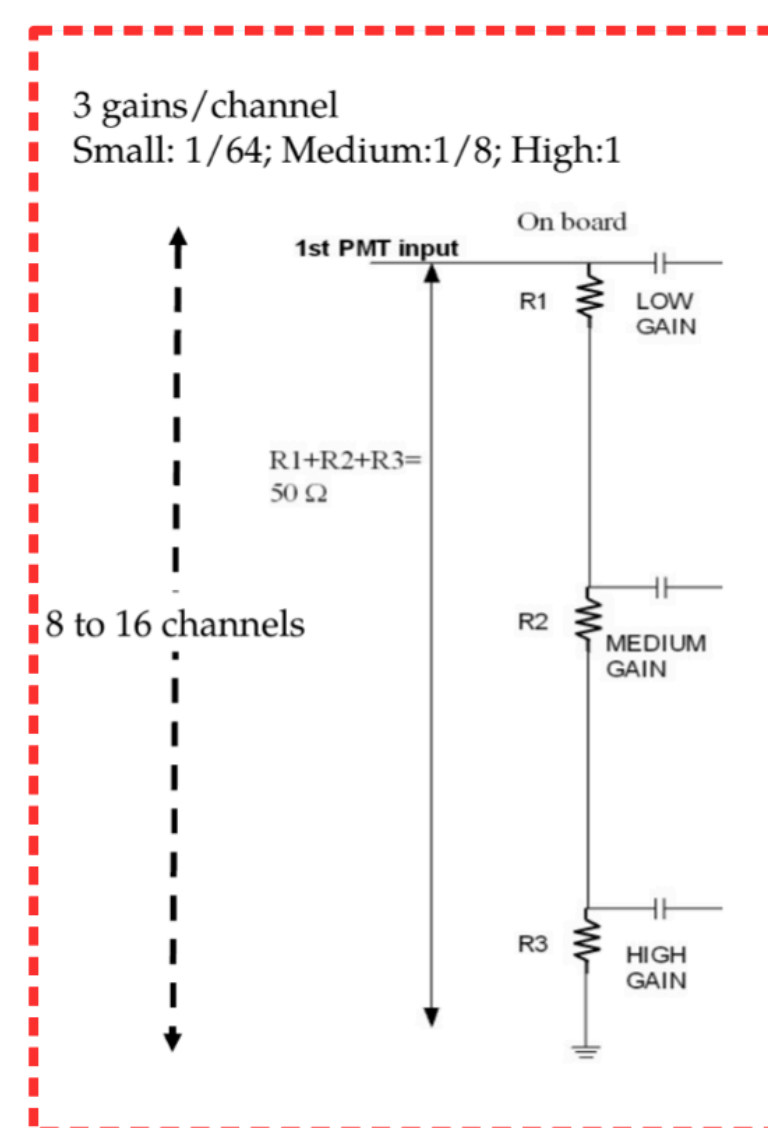
## R&D funded by X (400 k€)

# HKROC delivered in Nov 2021

# First tests in December 2021

# Synergy OMEGA/LLR/IRFU

# Use in future WC detectors!



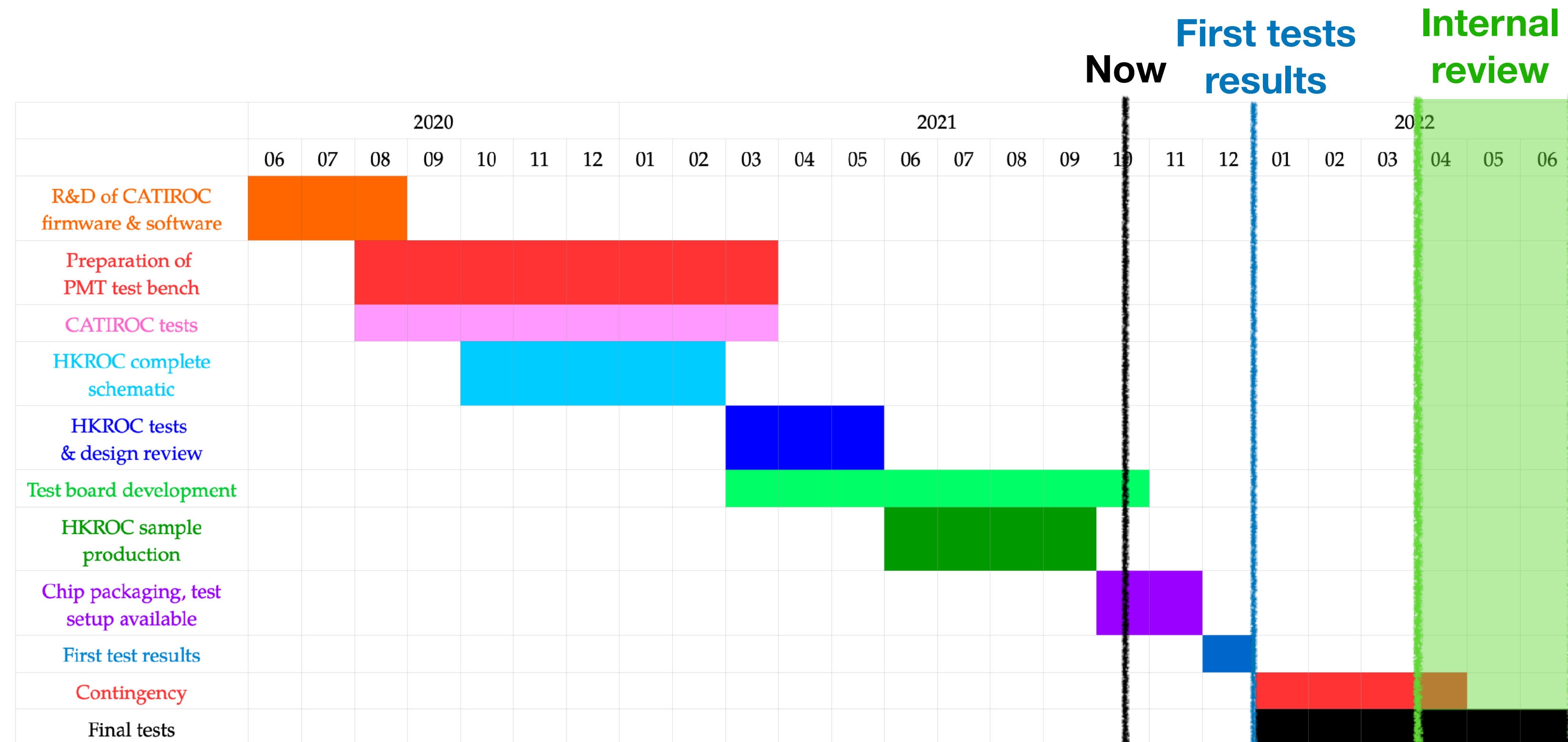


# Competitors and time constraints

Two other competitors : Japan QTC (SK chip) & Italian discrete solution  
 HK internal review Apr-June 2022 **selection of digitizer solution**

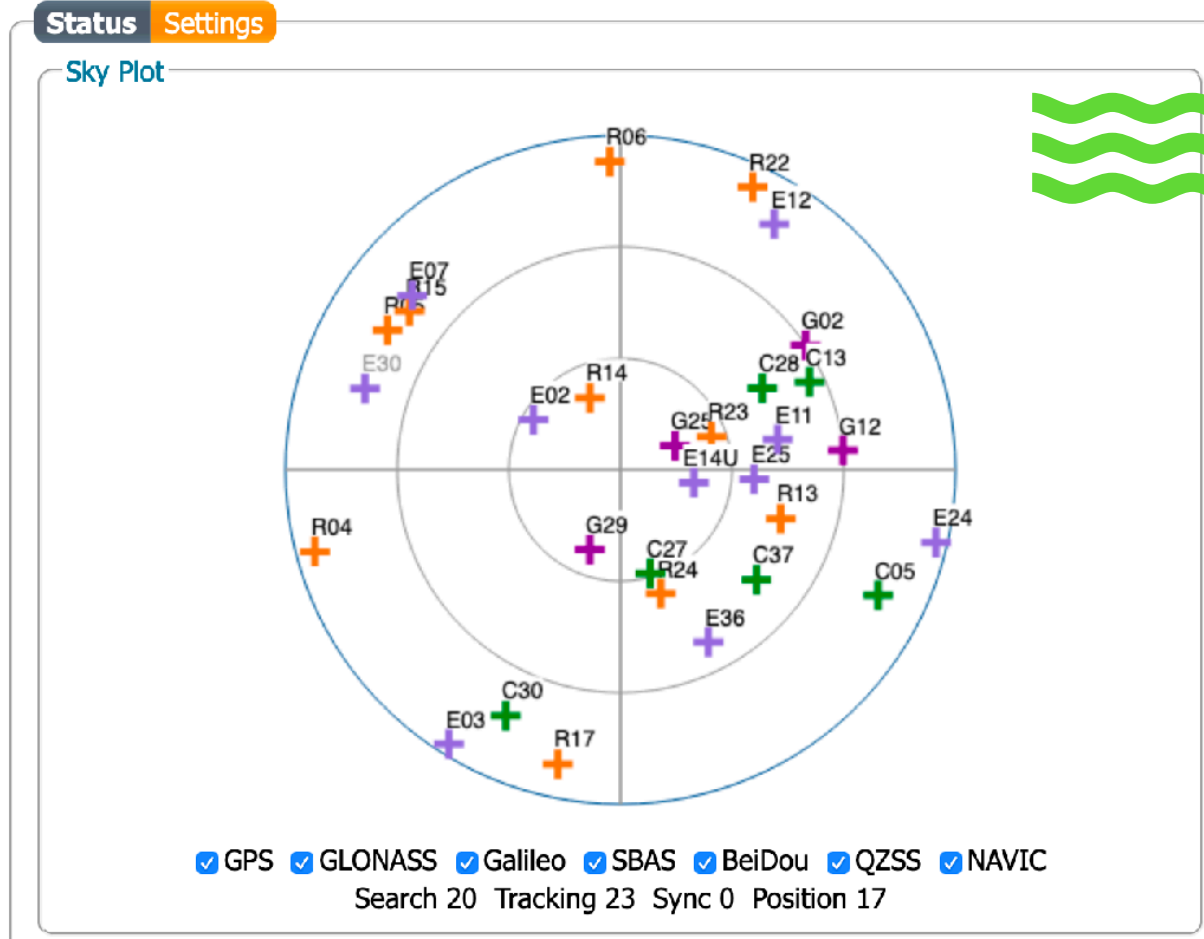
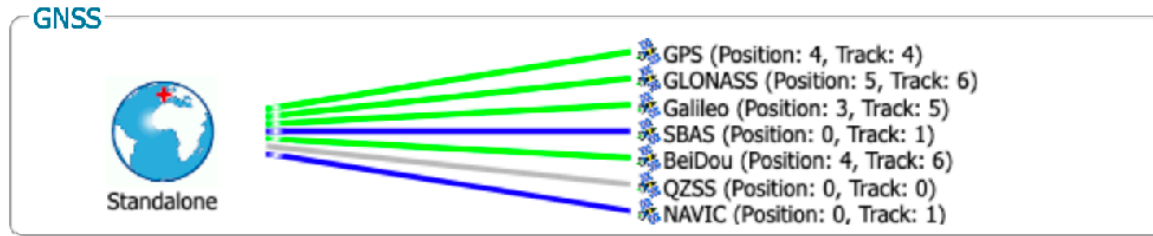
## Selection criteria:

- Performances
- Group expertise
- **Official engagement on fundings in June 2022**

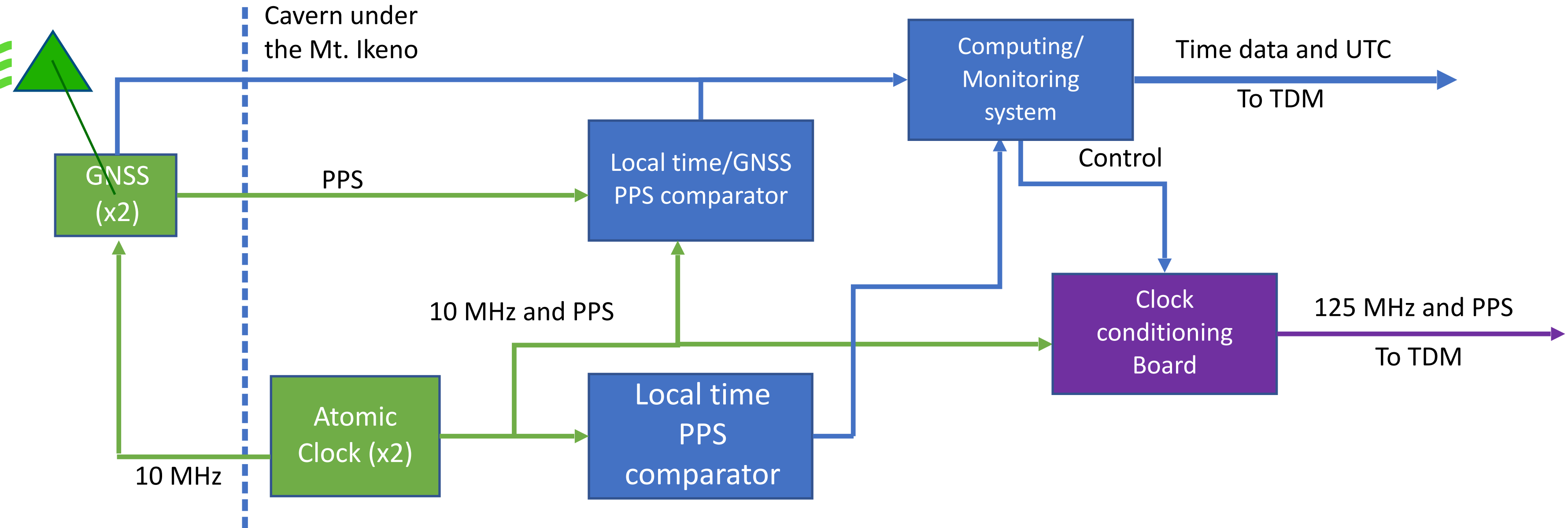




# Time Generation & UTC



Septentrio PolaRx5 satellites detection



Creation of reliable Universal Time for global synchronization and stable 125 MHz frequency for front-end digitizers

**Strong collaboration between LPNHE and SYRTE (Observatoire de Paris)**

Calibration of clocks and GNSS antennas

Creation of a dedicated lab to study clocks and GNSS at LPNHE

Long-term studies and comparison using atomic clocks, antennas and PPS-SYRTE

**R&D program supported by SU Emergence and ANR “Bertha”**

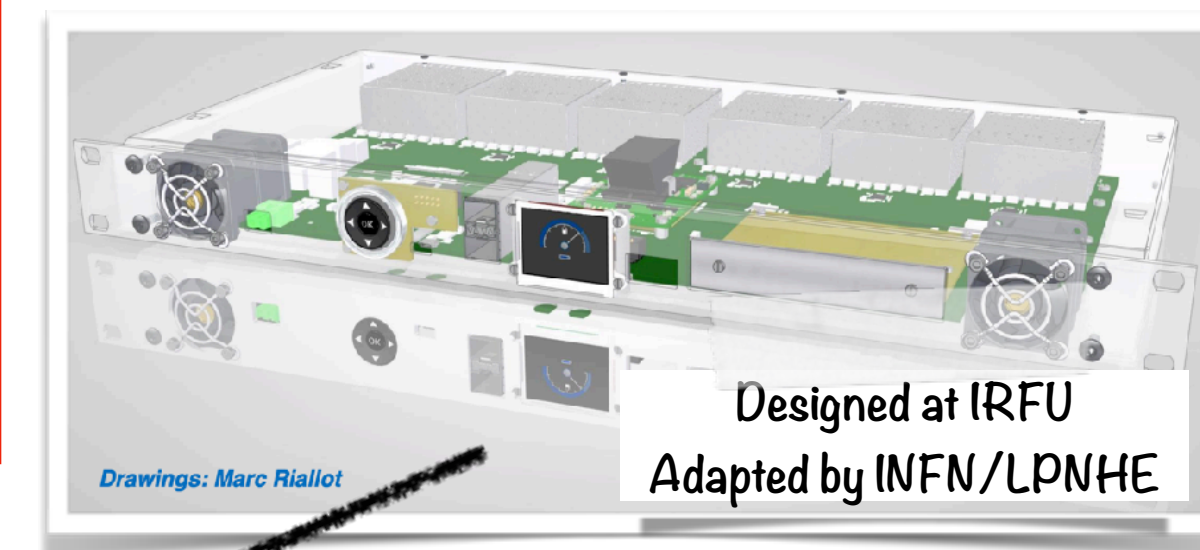
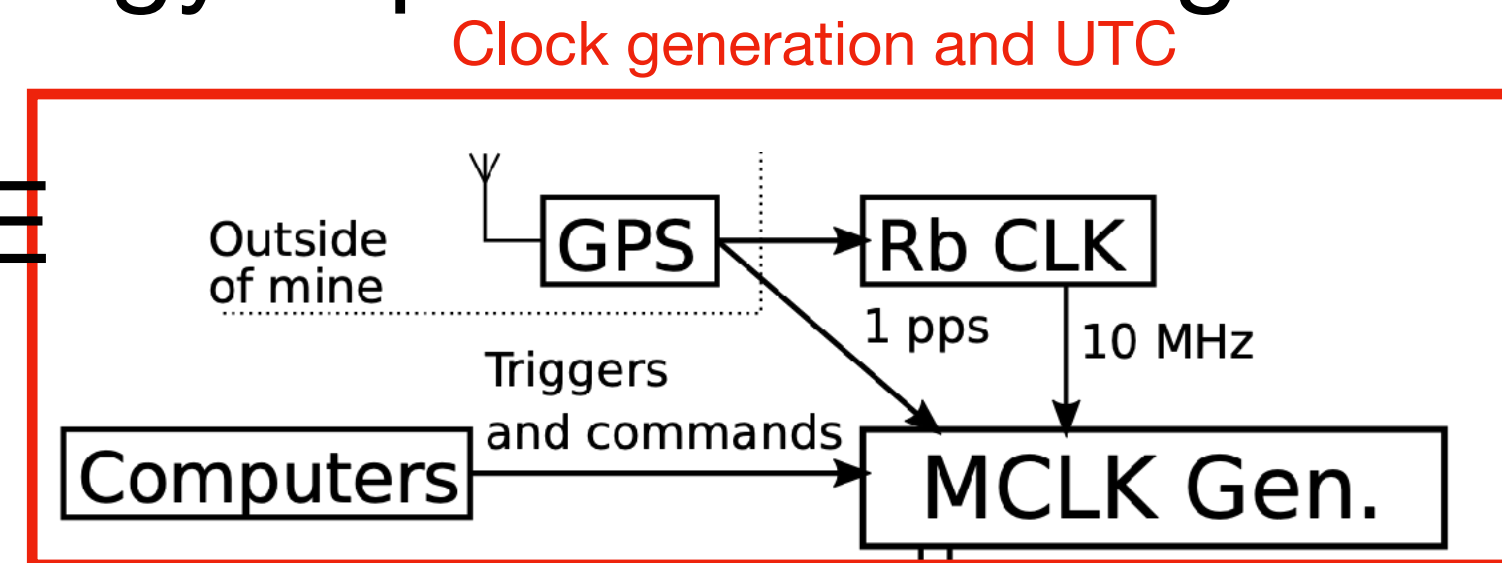


# Clock distribution

# Strong collaboration between LPNHE/INFN/IRFU

# Baseline proposal of time distributor modules (TDM) was finalized

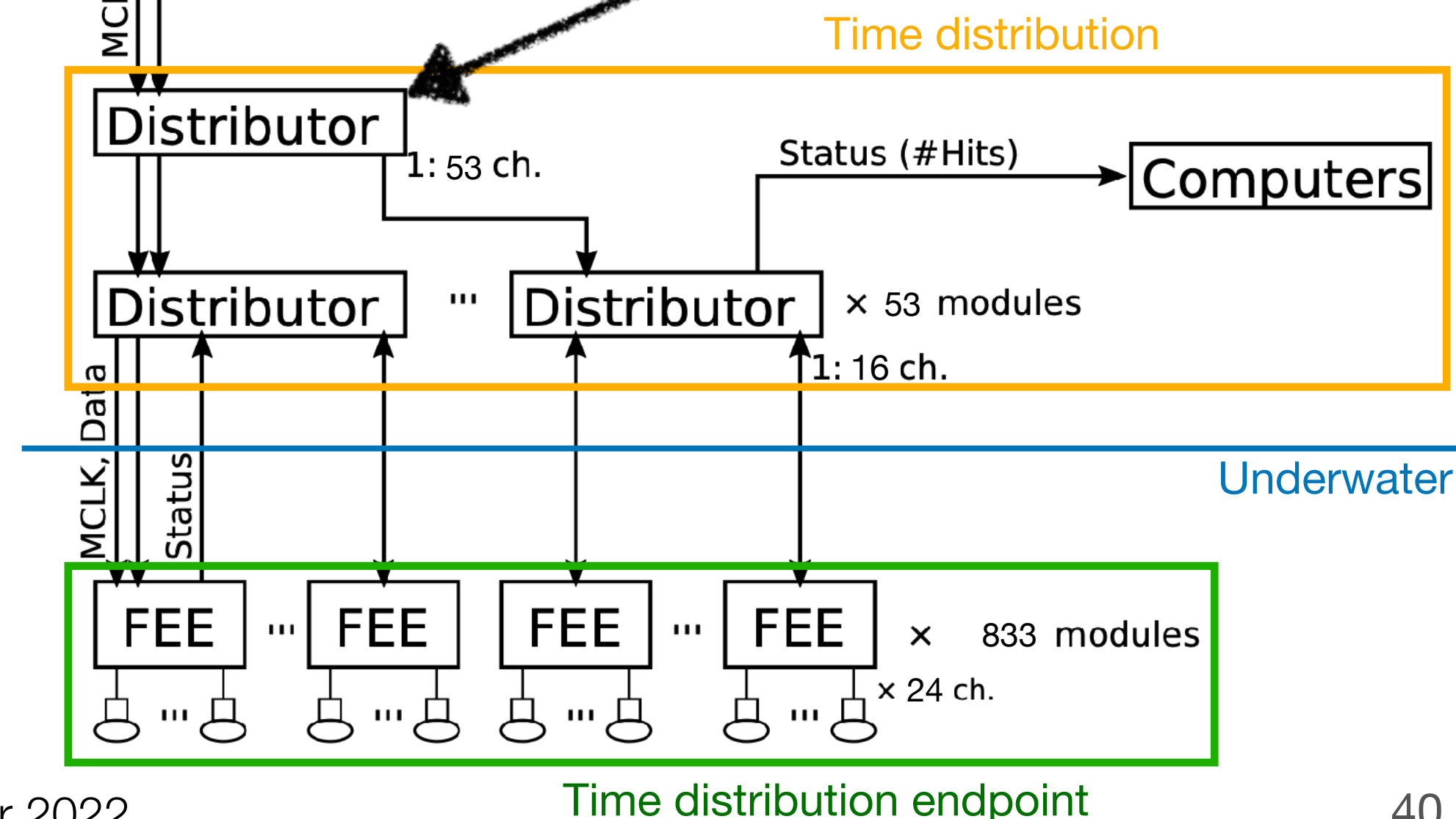
- Clock and Data Recovery (CDR) technology implemented using ser-des
- “Tree-like” structure
- Board design shared btw CEA & LPNHE
- Main distributor (53 ports)
- Second distributors (16 ports)
- Slow control and asynchronous signals distribution



# FE board development in collaboration with HKROC's

Other competitor: SK-based clock distribution system  
selection of the solution in June 2022

## Official engagement on fundings for approval





# Upcoming milestones

Technological decision  
for HKROC and clocks

	S2 2020	S1 2021	S2 2021	S1 2022	S2 2022	S1 2023	S2 2023
HKROC R&D							
HKROC production							
Front-end board R&D							
Front-end board production							
Time generation R&D							
Clock distribution R&D							
Clock distribution production							
CC-IN2P3 integration							
Productions at CC-IN2P3							
Electronics internal review							
MOU signature							

R&D almost complete

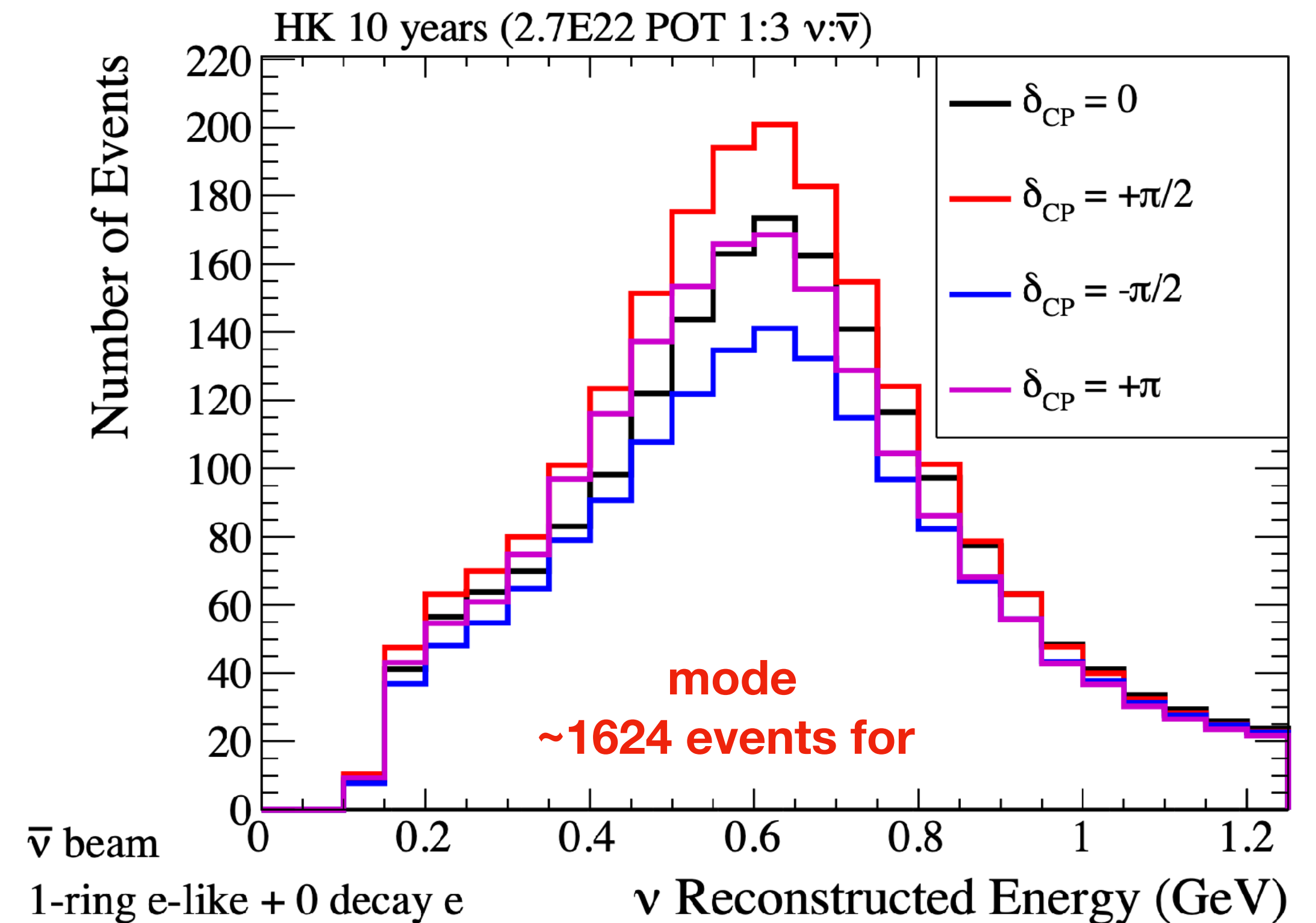
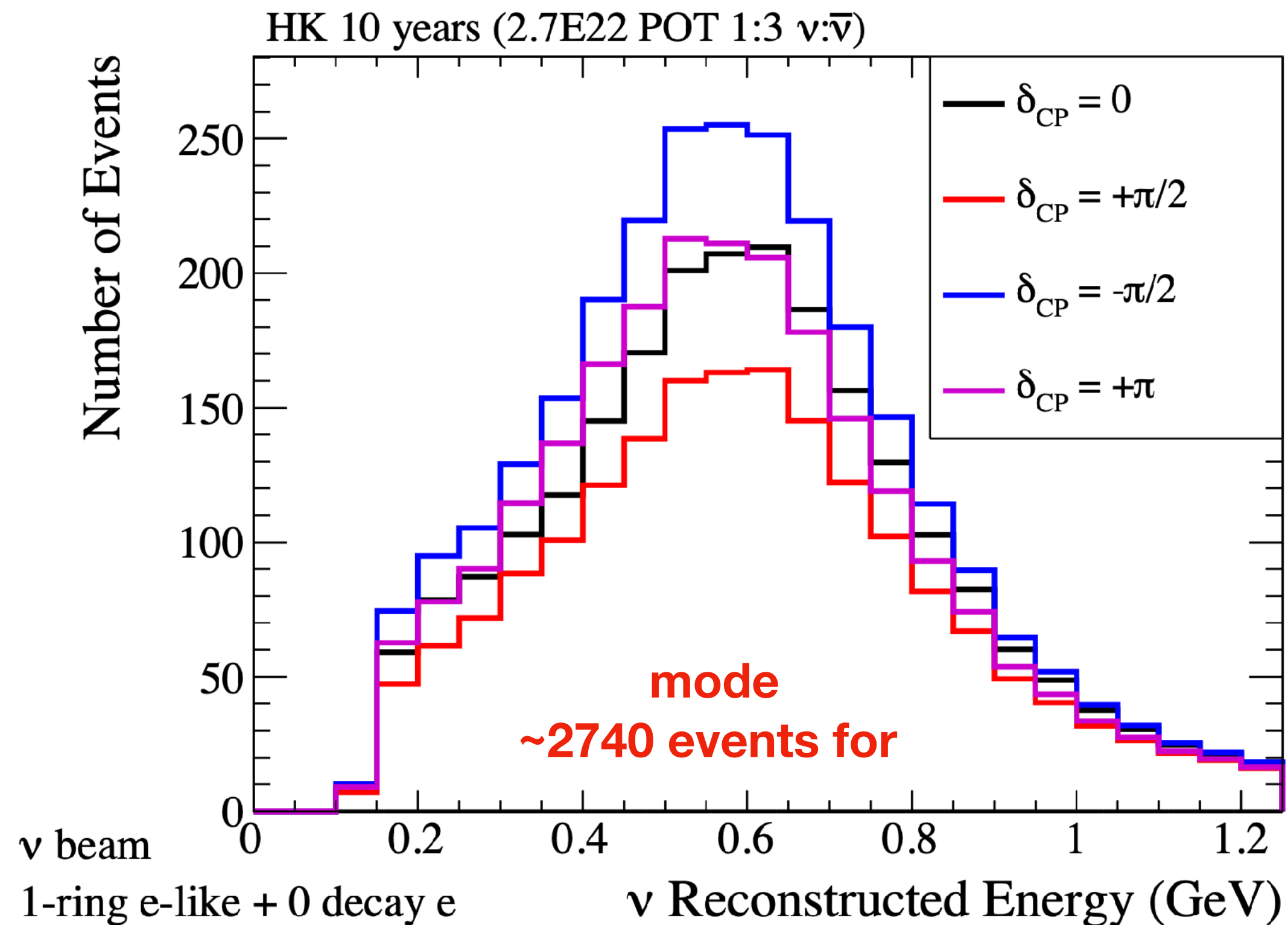
HK internal review beginning of 2022 **technological decision June 2022**

MOU signature in 2022



# HK LBL- oscillations measurements

Data taking plan: **1.3 MW** **6 cycles/year** **10 years** ( )

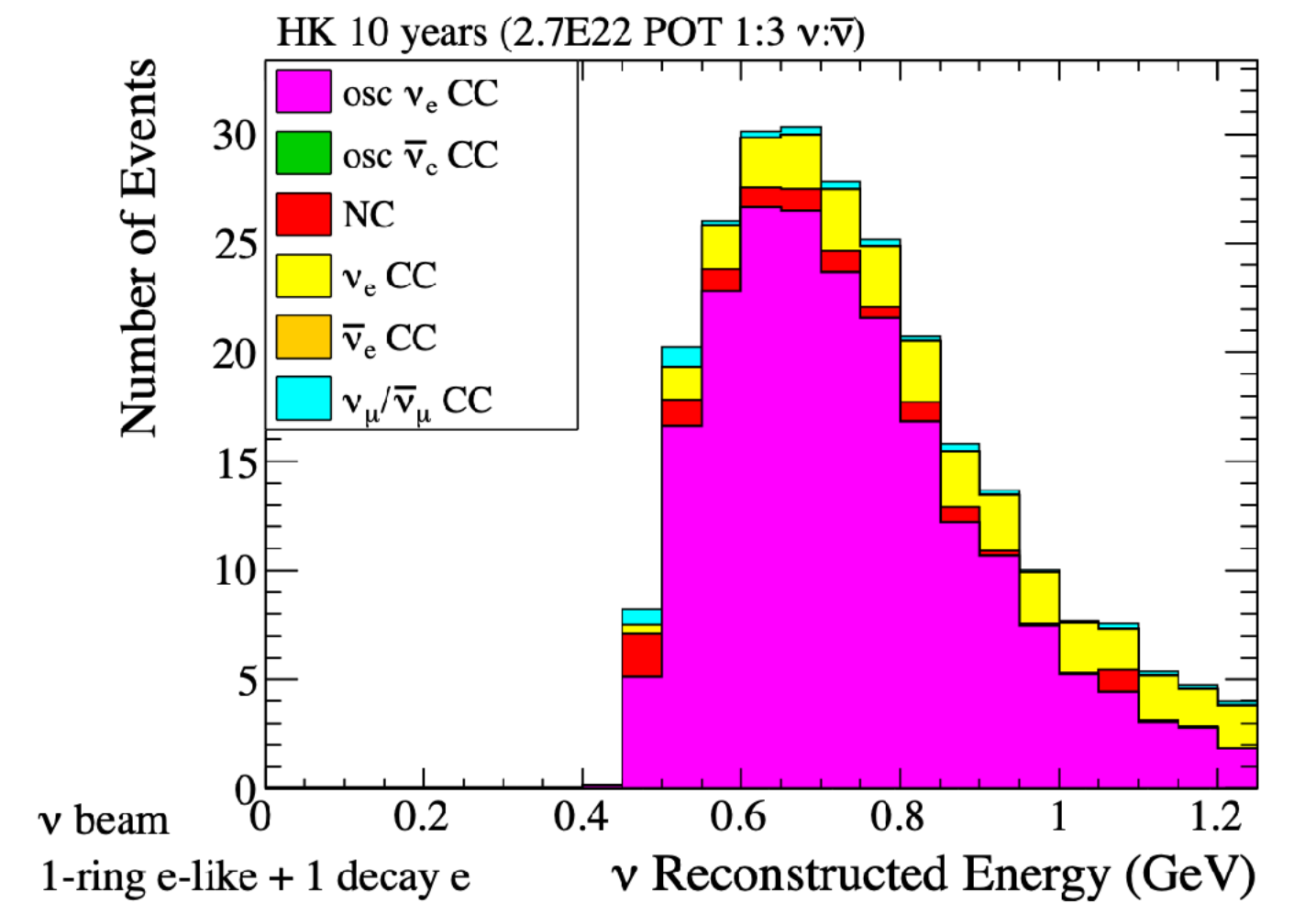
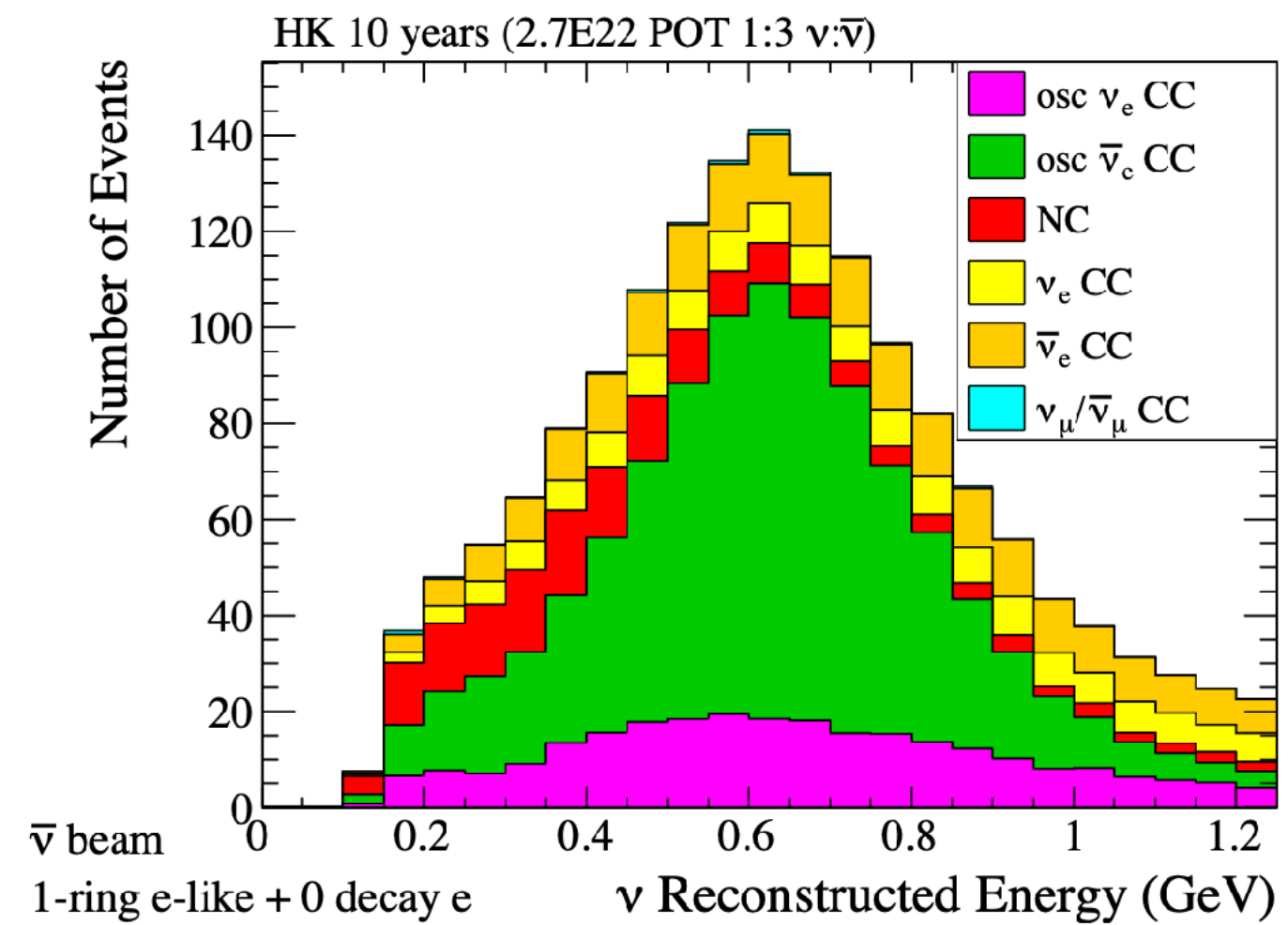
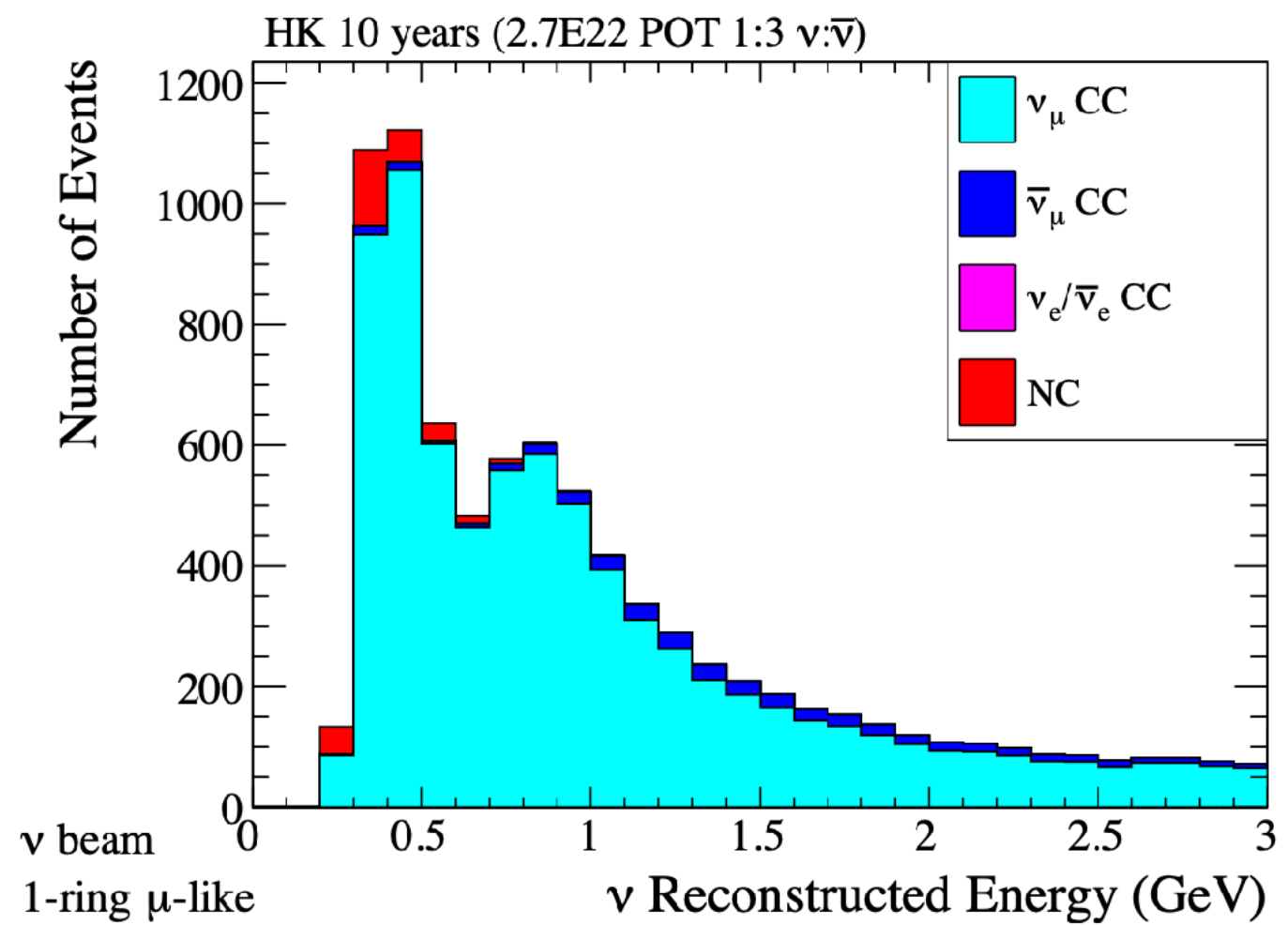
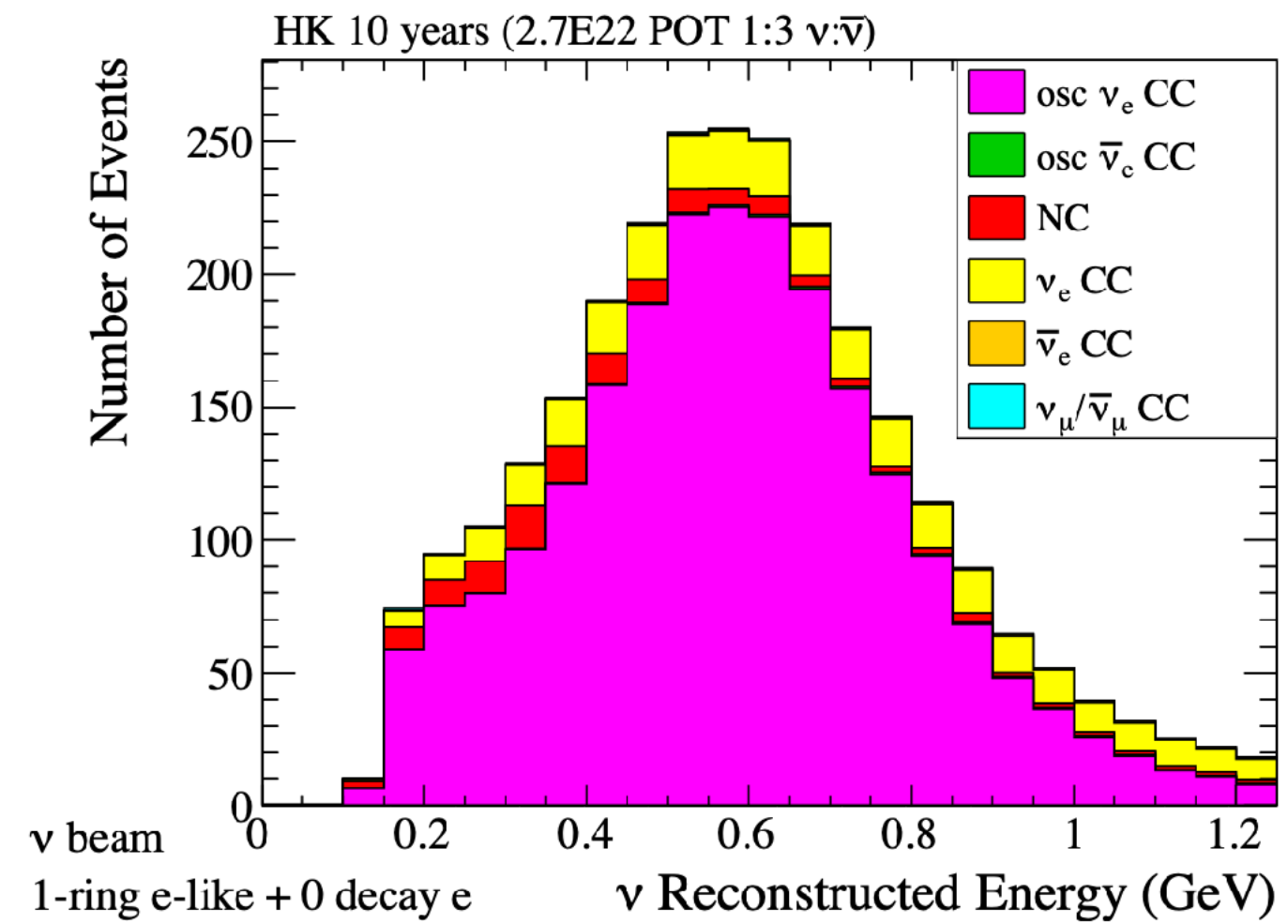
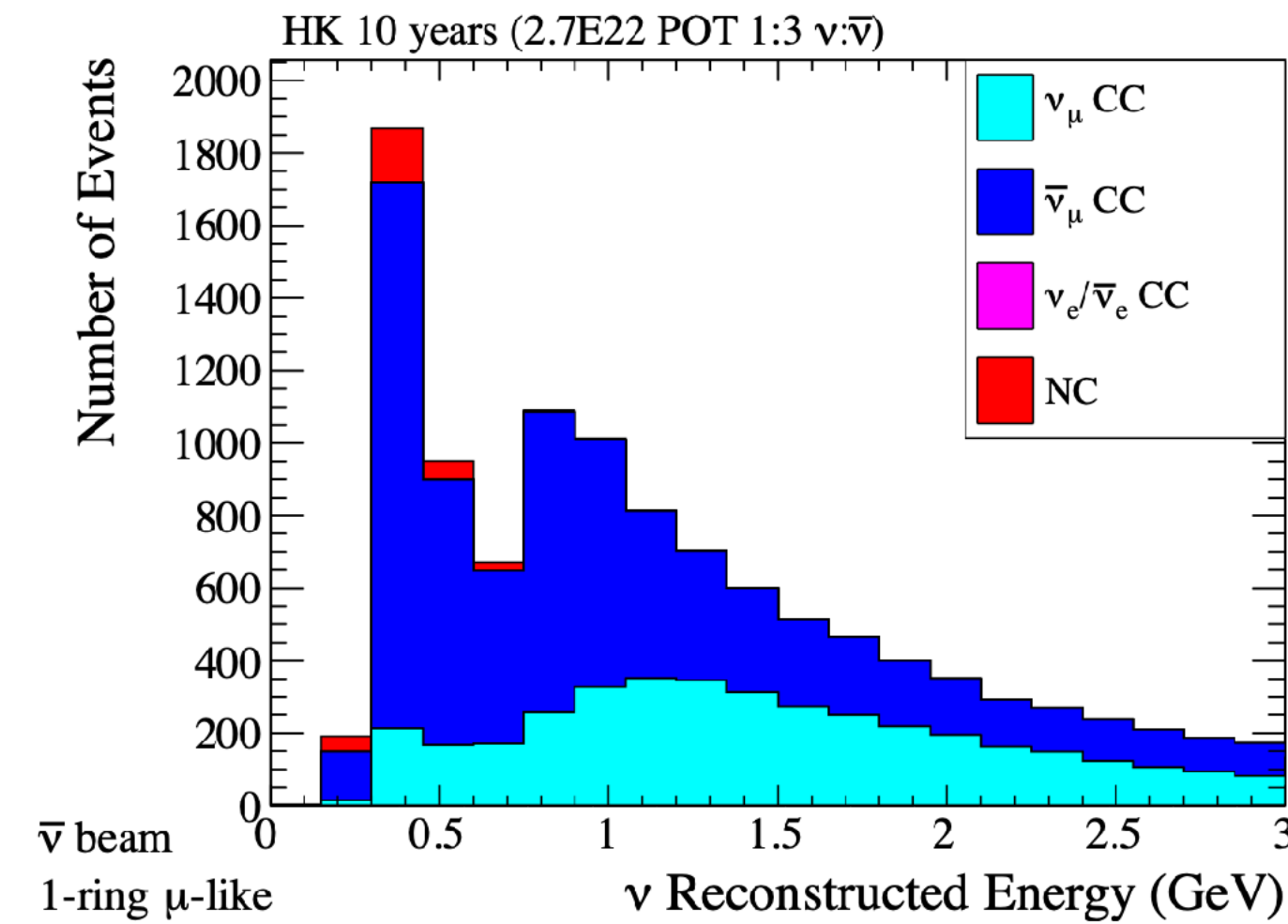


Appearance measurements: increased by 2 orders of magnitude (75:15 for T2K-I Nature)

**Higher statistics measurements of the oscillation parameters**

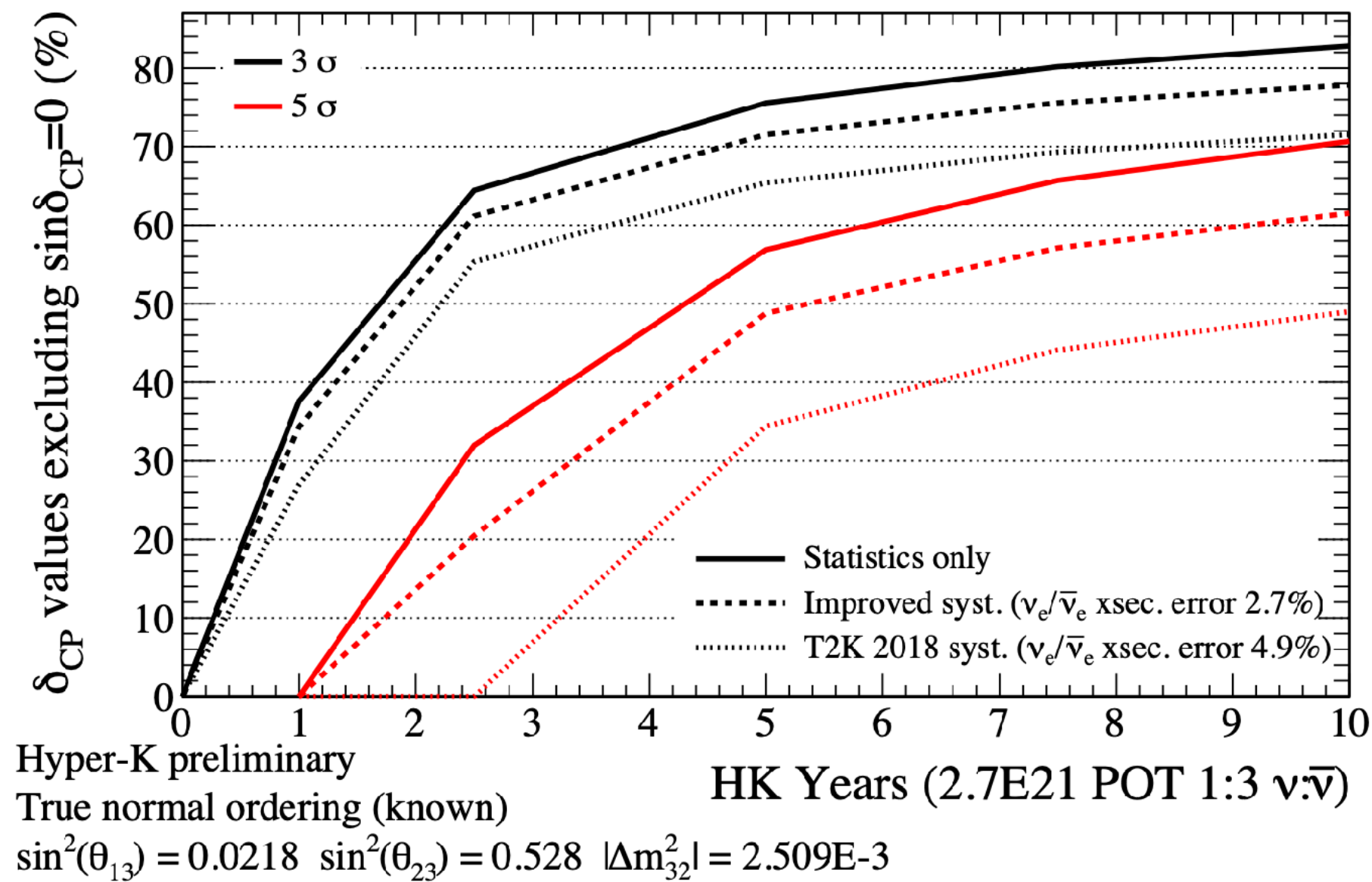
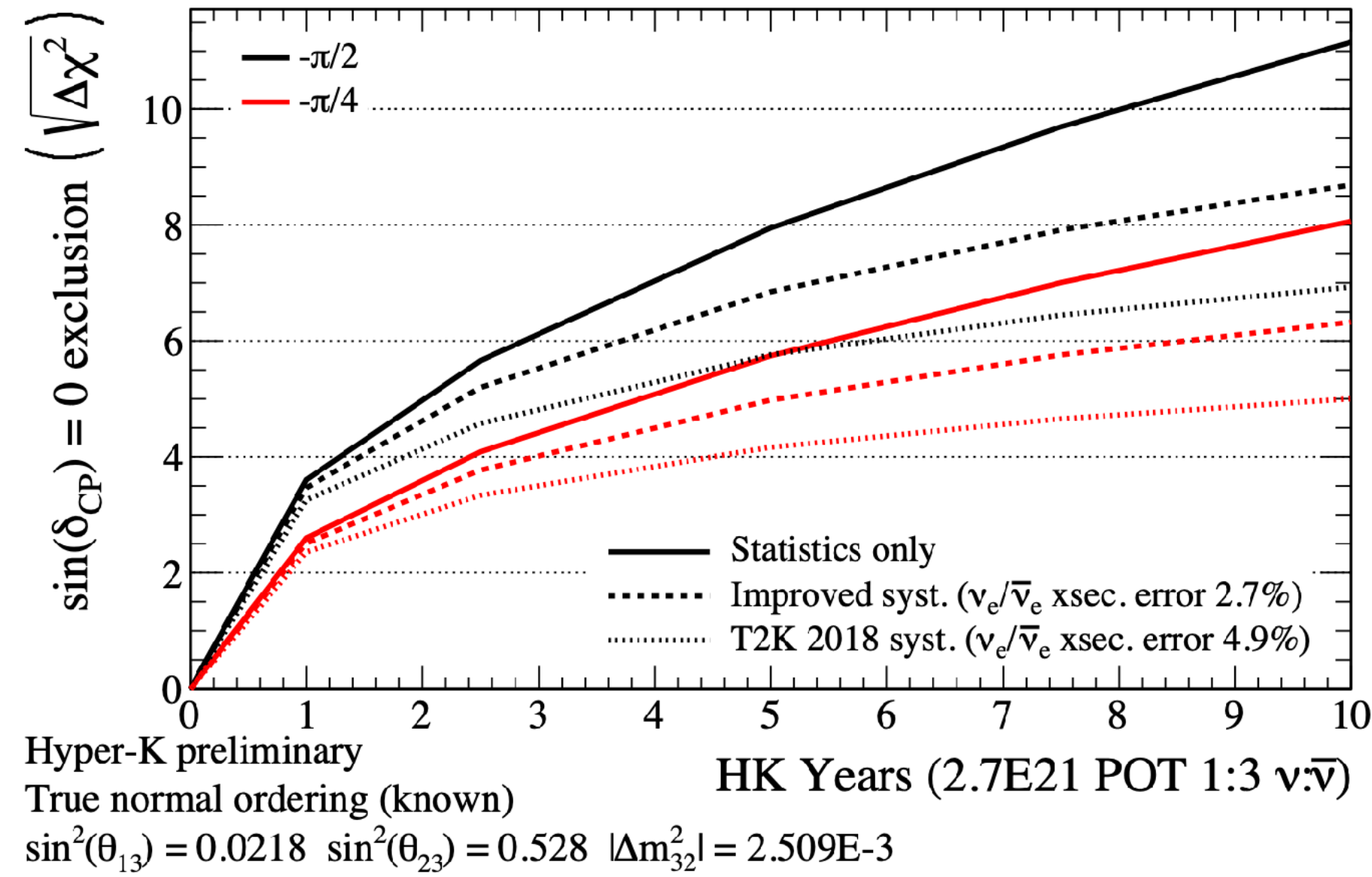


# HK OA samples





# sensitivity





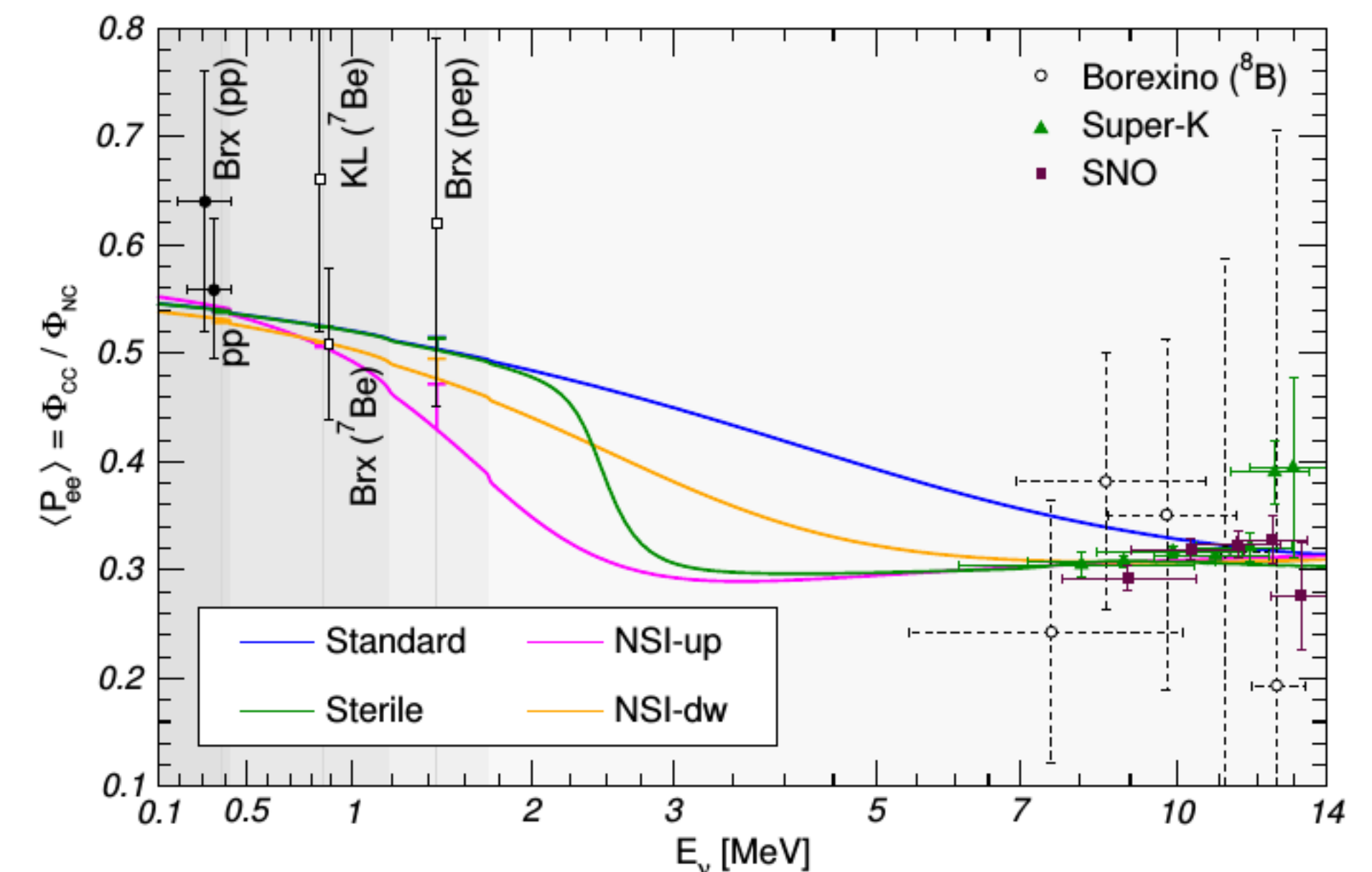
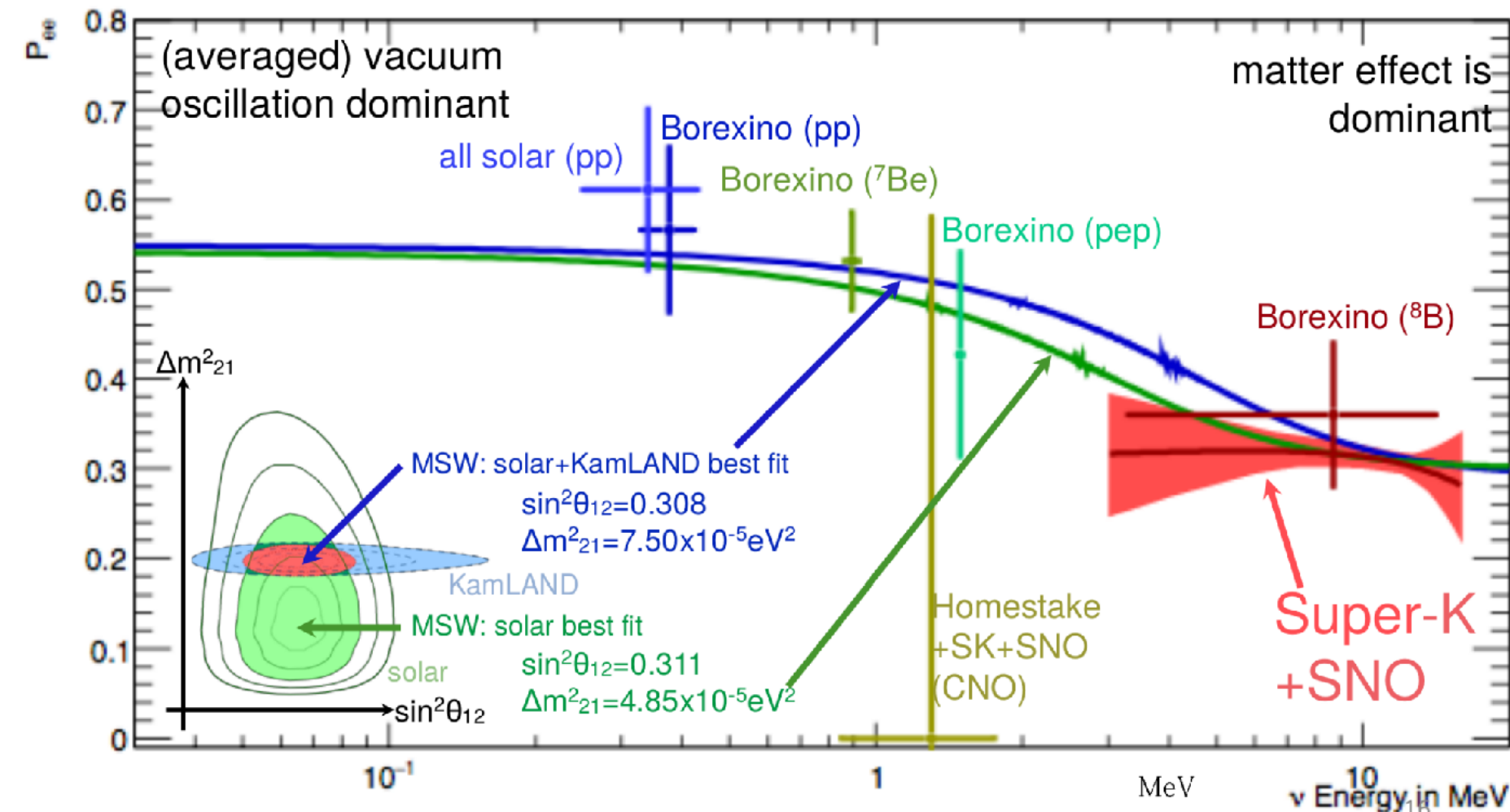
# Solar neutrinos: upturn

Matter effect in the Sun dominant at high energy  
transition in probability around a few MeV

Displacement towards lower energy could be explained by

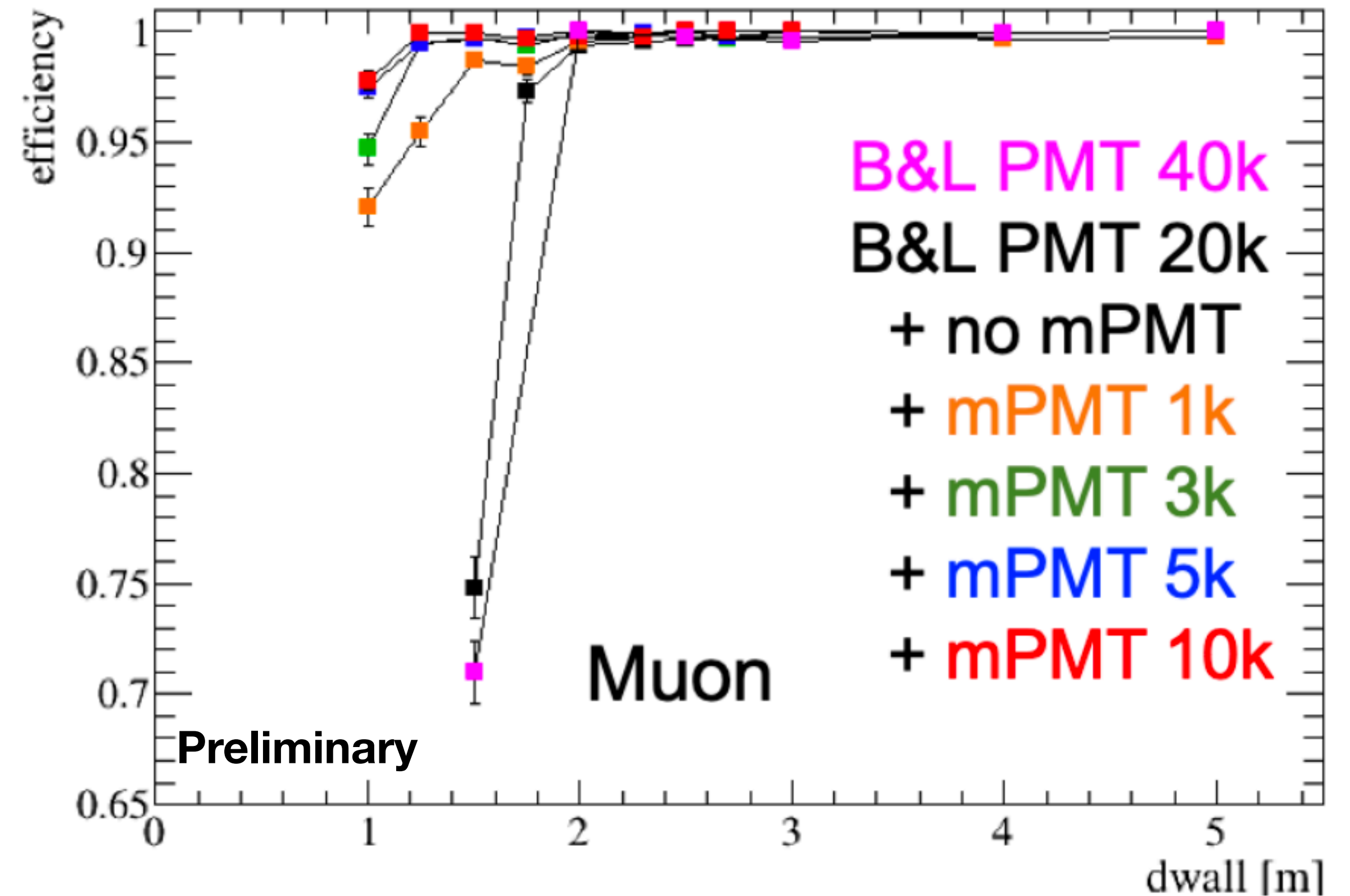
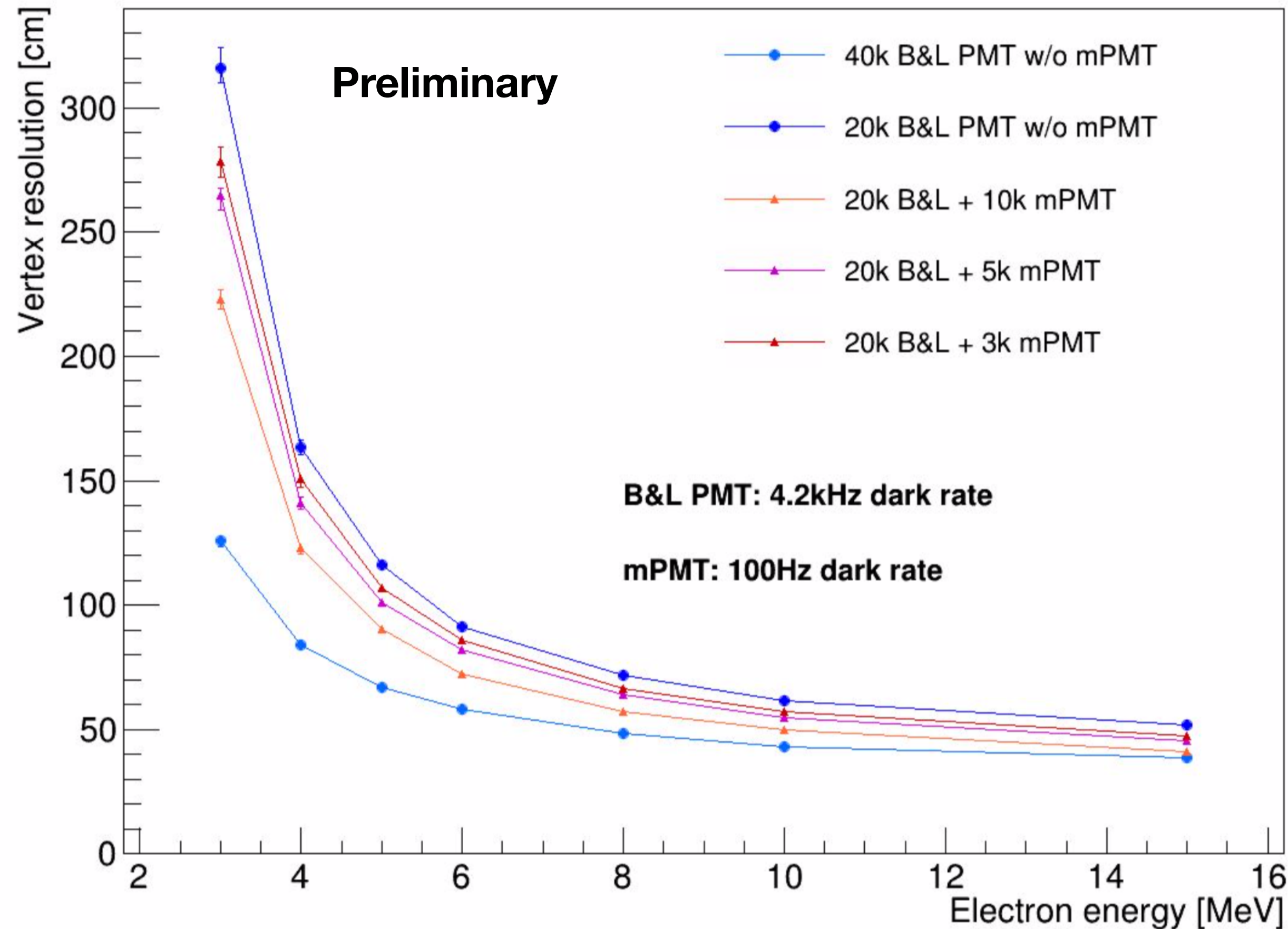
- Statistical fluctuation ?
- Light sterile neutrino ?
- Non Standard Interaction in the dense Sun ?

Measurement accessible using mPMTs (lower energy threshold)





# mPMT impact on physics



Vertex resolution improved w/ mPMT for  $dWall \leq 8m \rightarrow$  FV expansion  
mPMT allows to explore  $< 5$  MeV region.



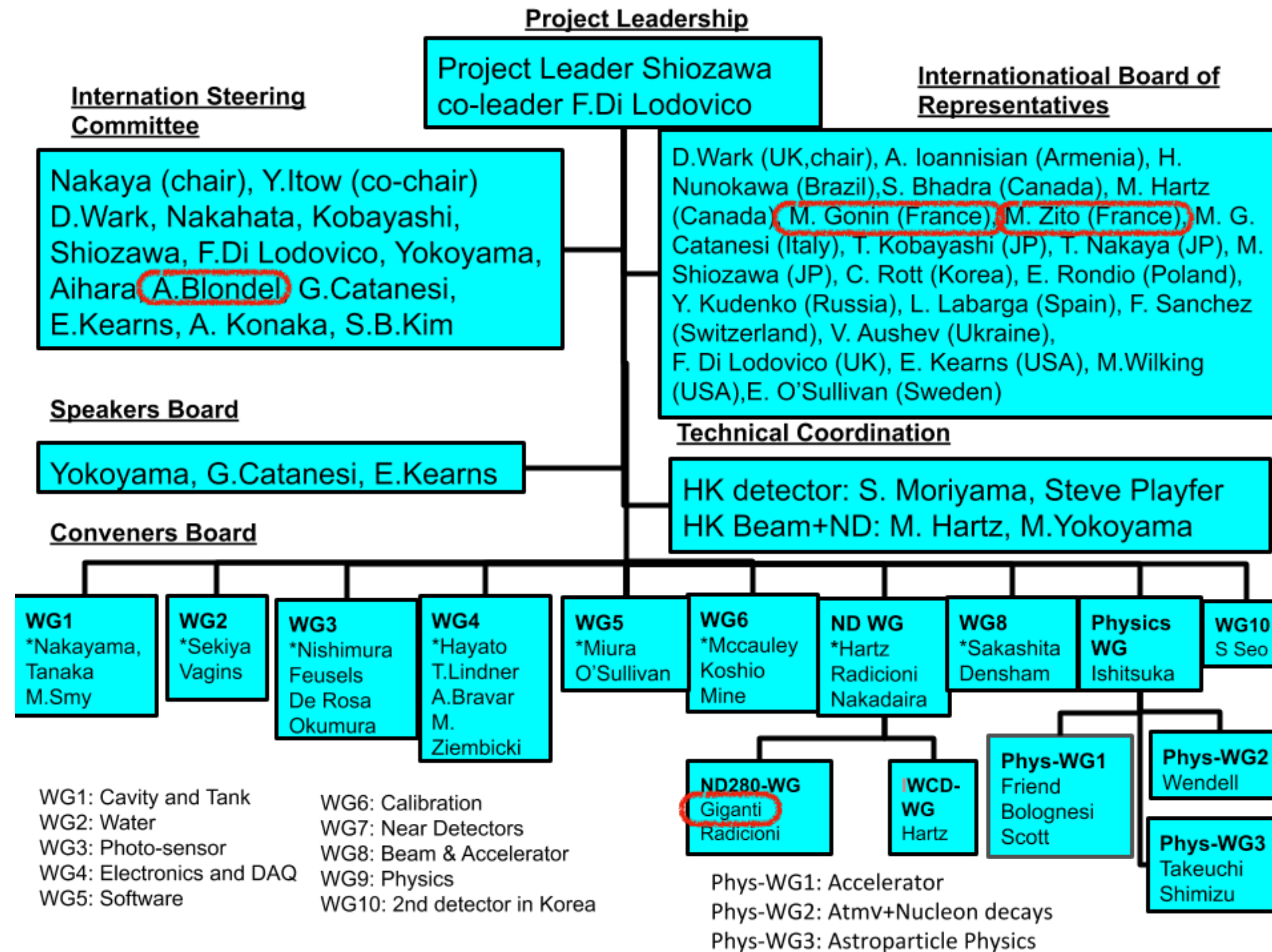
# Digitizer costs

1.9 M€ for ID to be shared btw IN2P3 and CEA

0.7 M€ for OD to be shared btw IN2P3, CEA and UK



# Collaboration Organigram

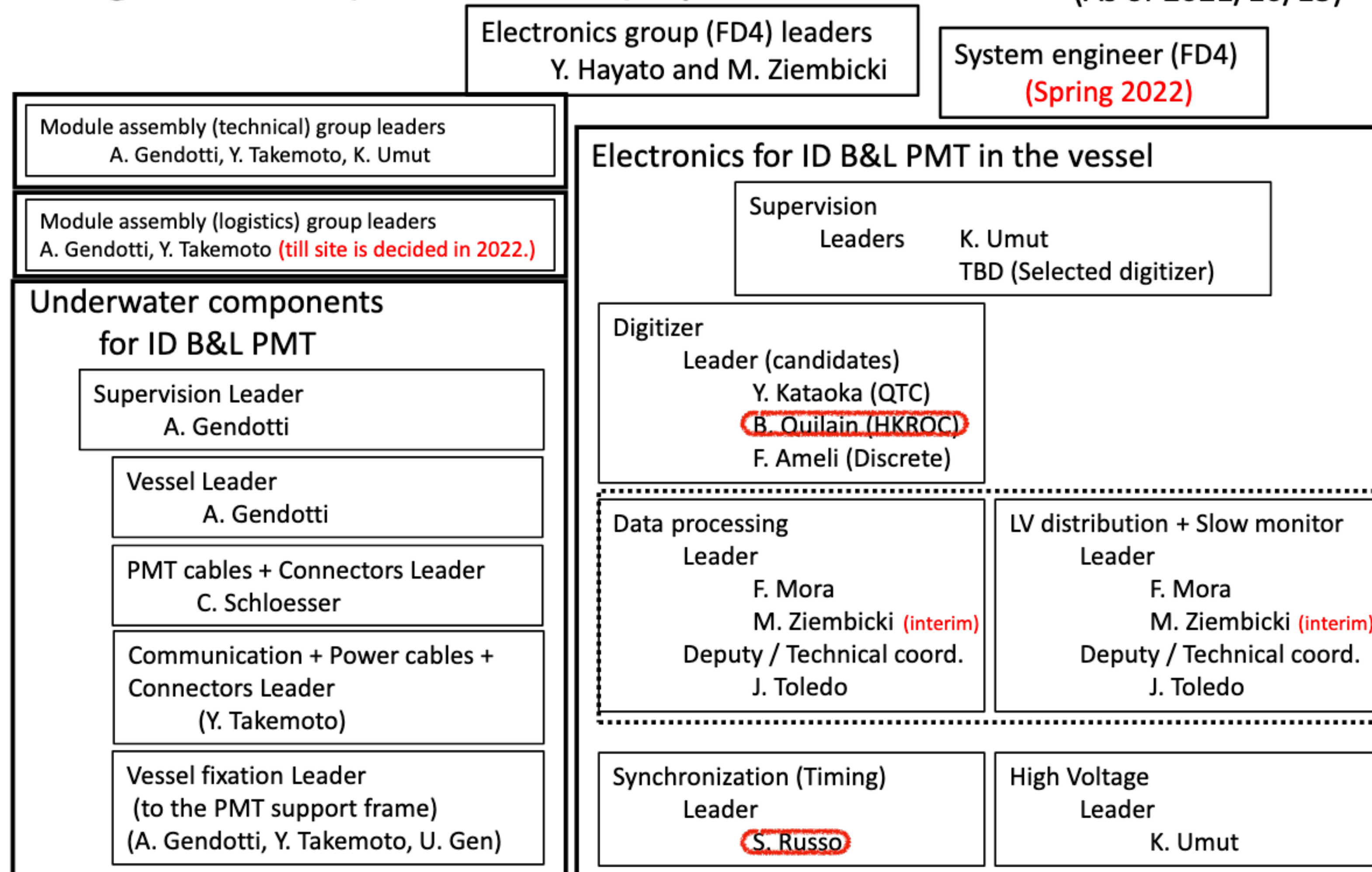




# Electronics organization for ID B&L

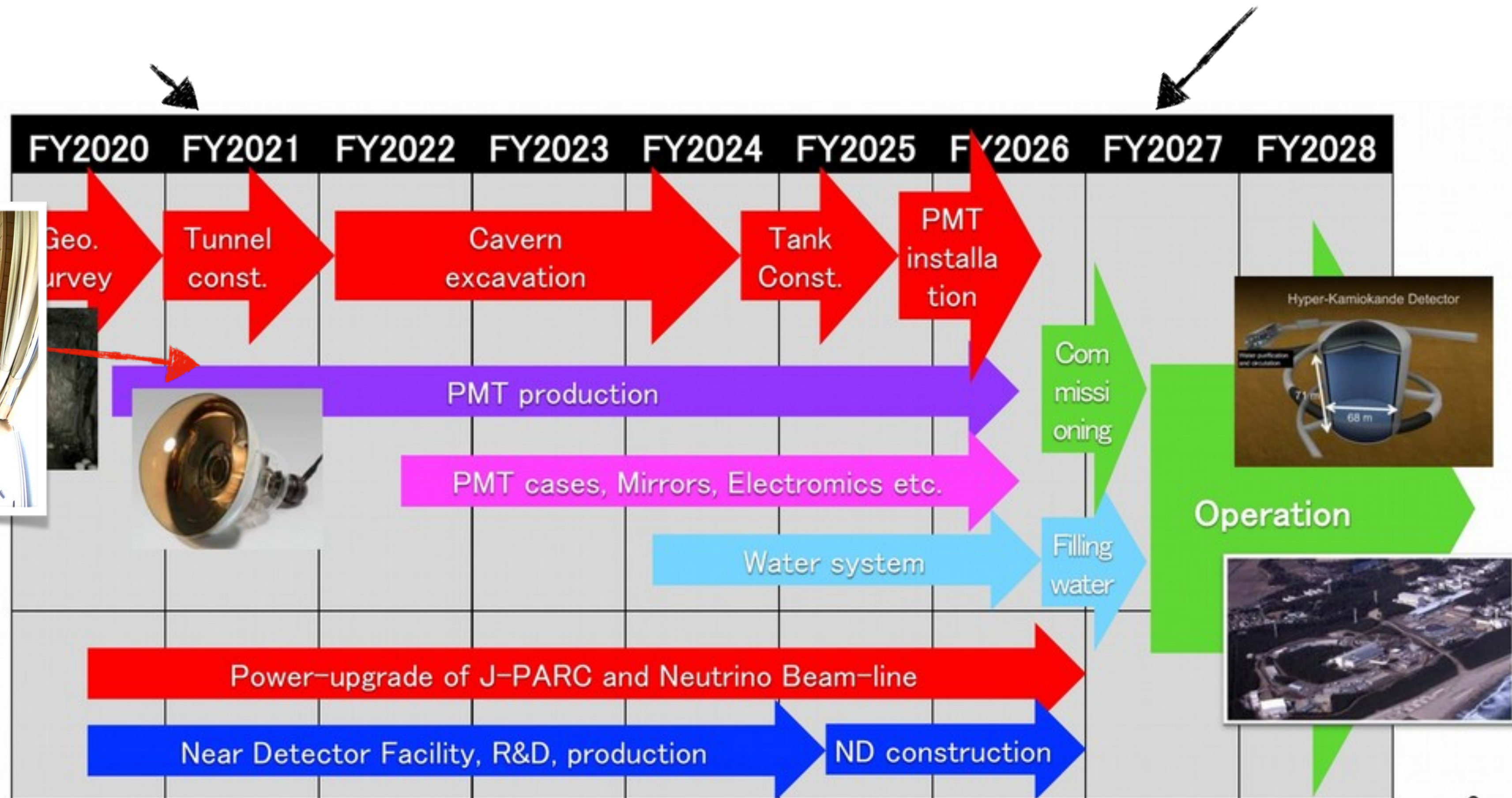
## Organization (ID B&L PMT) update

(As of 2021/10/13)





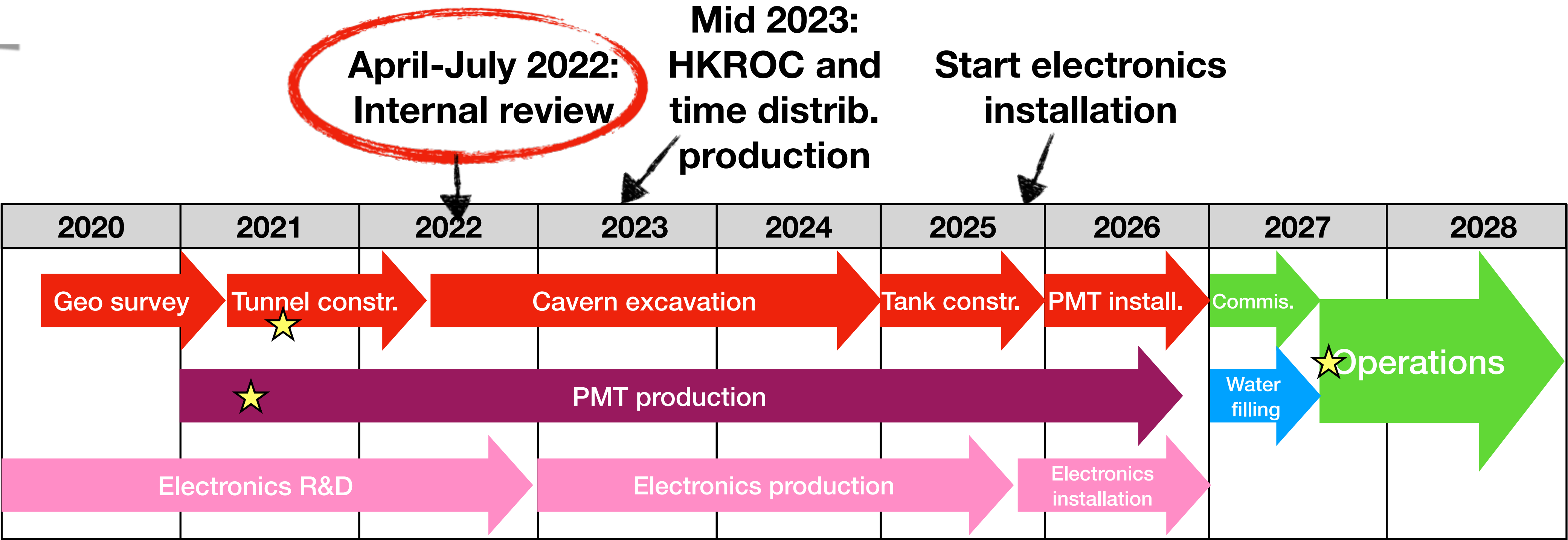
# Hyper-Kamiokande schedule



3



# HK schedule





# Tunnel construction





# Computing timeline

## Computing resources timeline



**2021** \* Update **computing predictions/forecast**

**2021 → 2023** \* **Work with countries/institutes** that will contribute computing resources  
→ Intent to contribute, **application procedures**, allocations  
→ Acquire small scale **resources for integrating/testing/using**

**2023** \* Confirmed/**secured resources** to cover **MC validation and final prod (2024-2027)**

**2024 → 2025** \* **Computing forecasts** for data era recomputed based on validation samples and updated information from DAQ group.

\* Work with countries/institutions to use new predictions to make **computing applications** for the **data taking phase** of HK

**2025** \* Confirmed/**secured resource** allocation for the first 5+ years of data taking with plans in place to expand throughout HK lifetime

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\* Computing predictions will be revised every couple of years

\* Addition of new resources, and development of integrating different types of new resources can be on-going. But the largest computing resources, which cover the bulk of our needs, will follow above timeline.

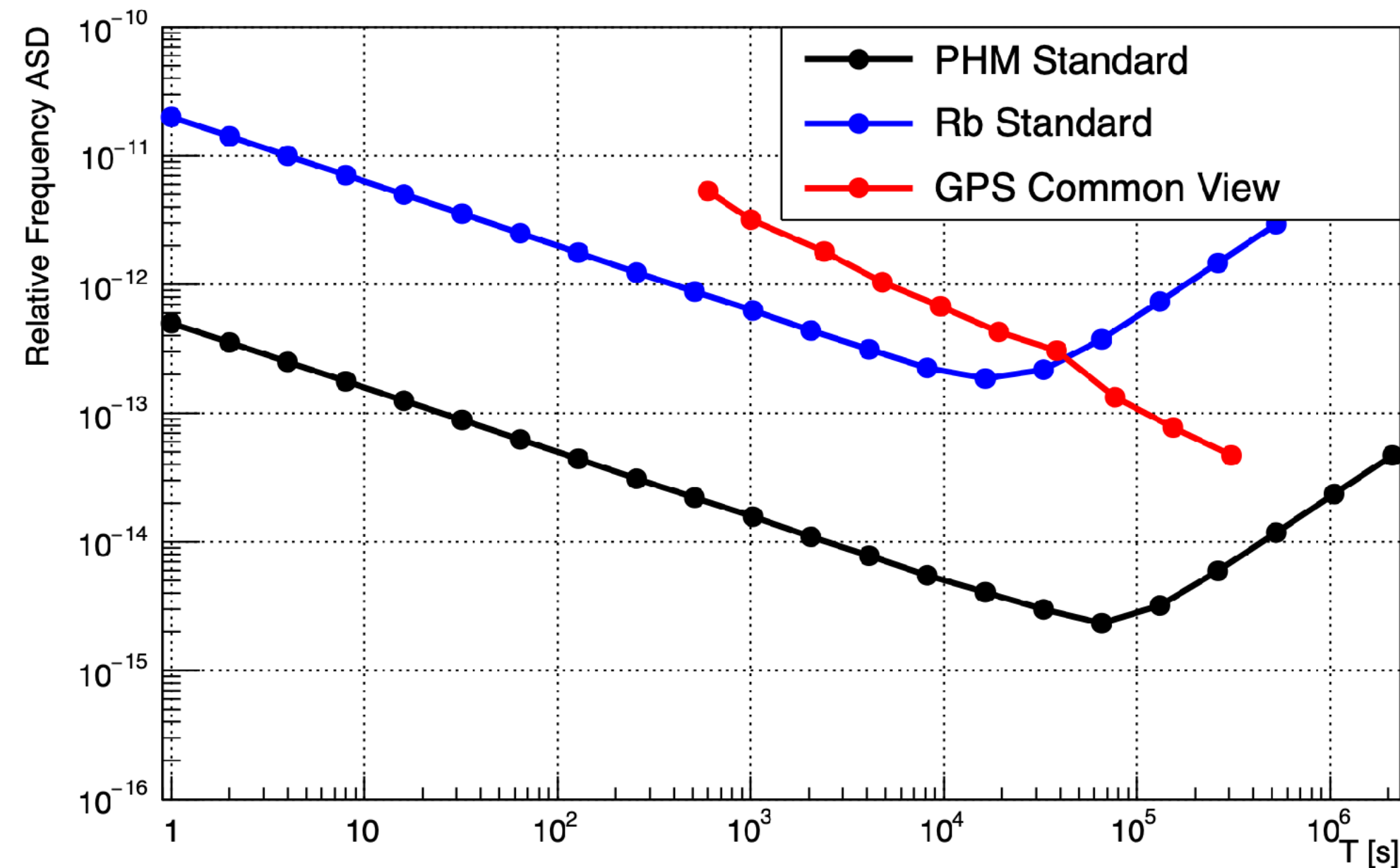
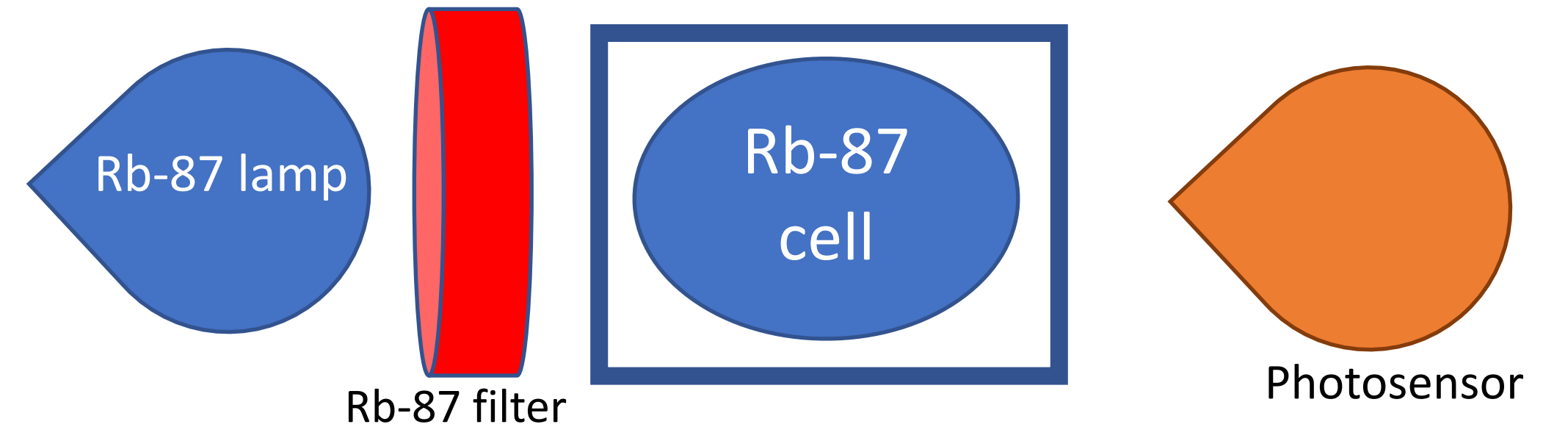
14



# Rb atomic clock



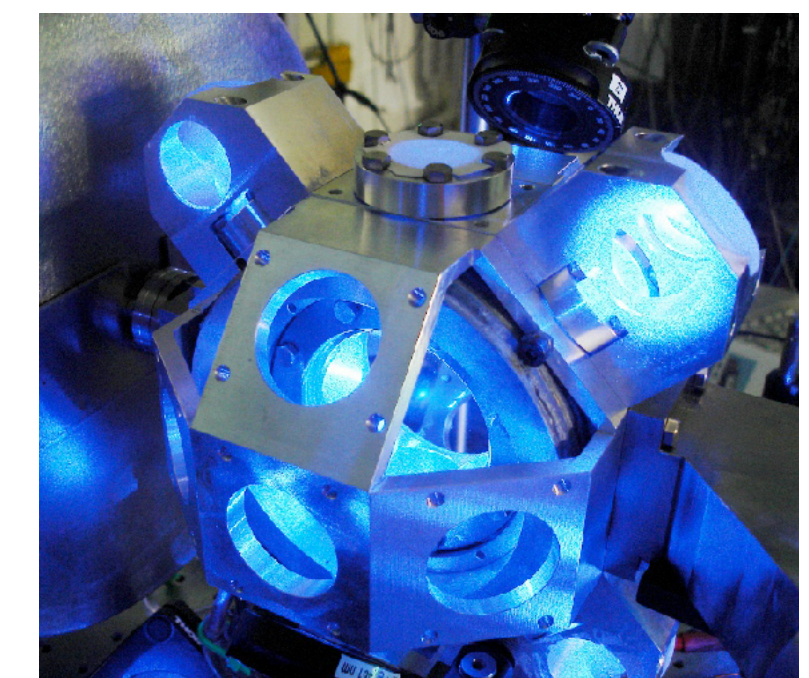
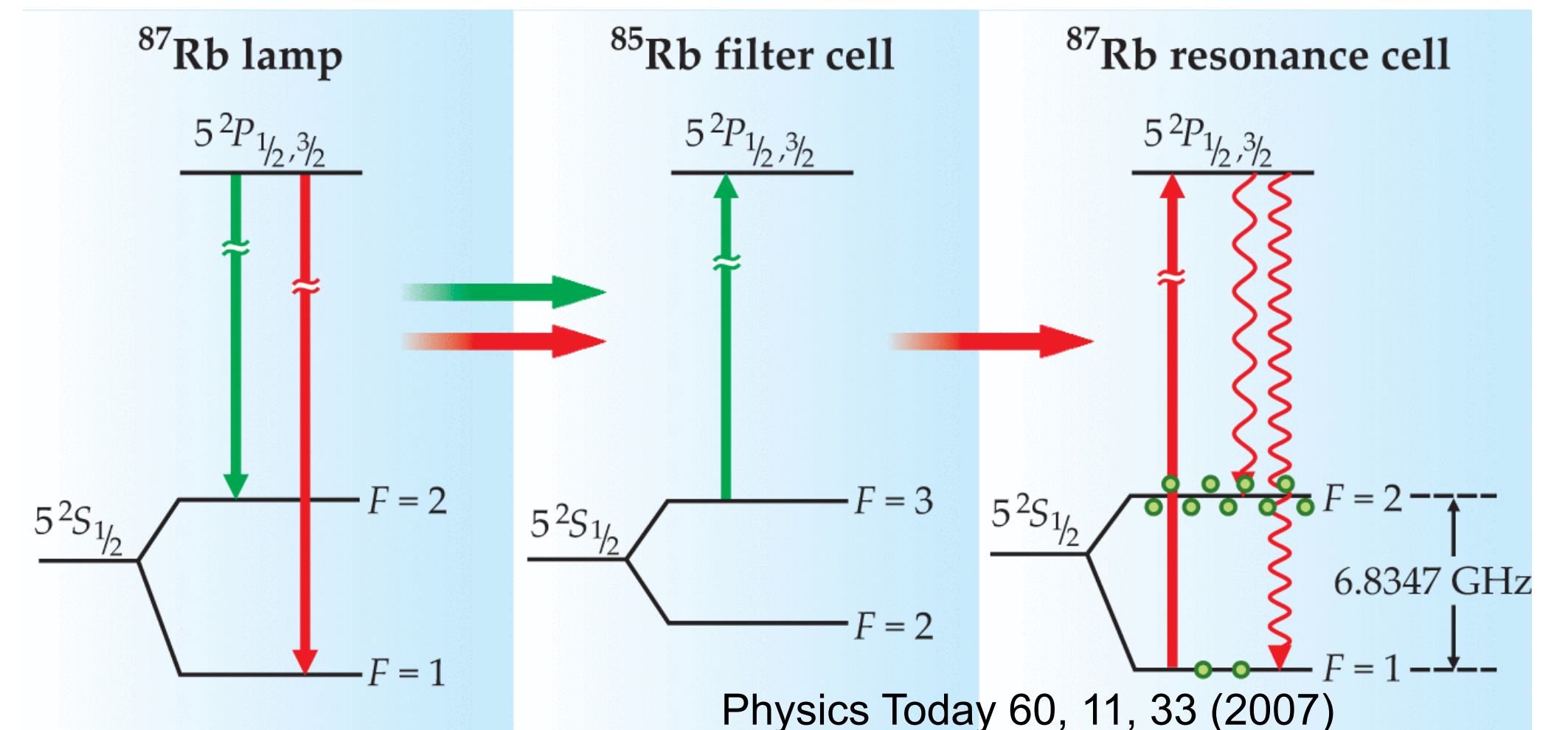
Microwaves oven  
@ 6.8347 GHz



Allan standard deviation: [paper](#)

Comparison with time reference at SYRTE **performances studies**

Test using more stable clocks e.g. Passive Hydrogen Maser (PHM1008)



Strontium  
optical  
atomic

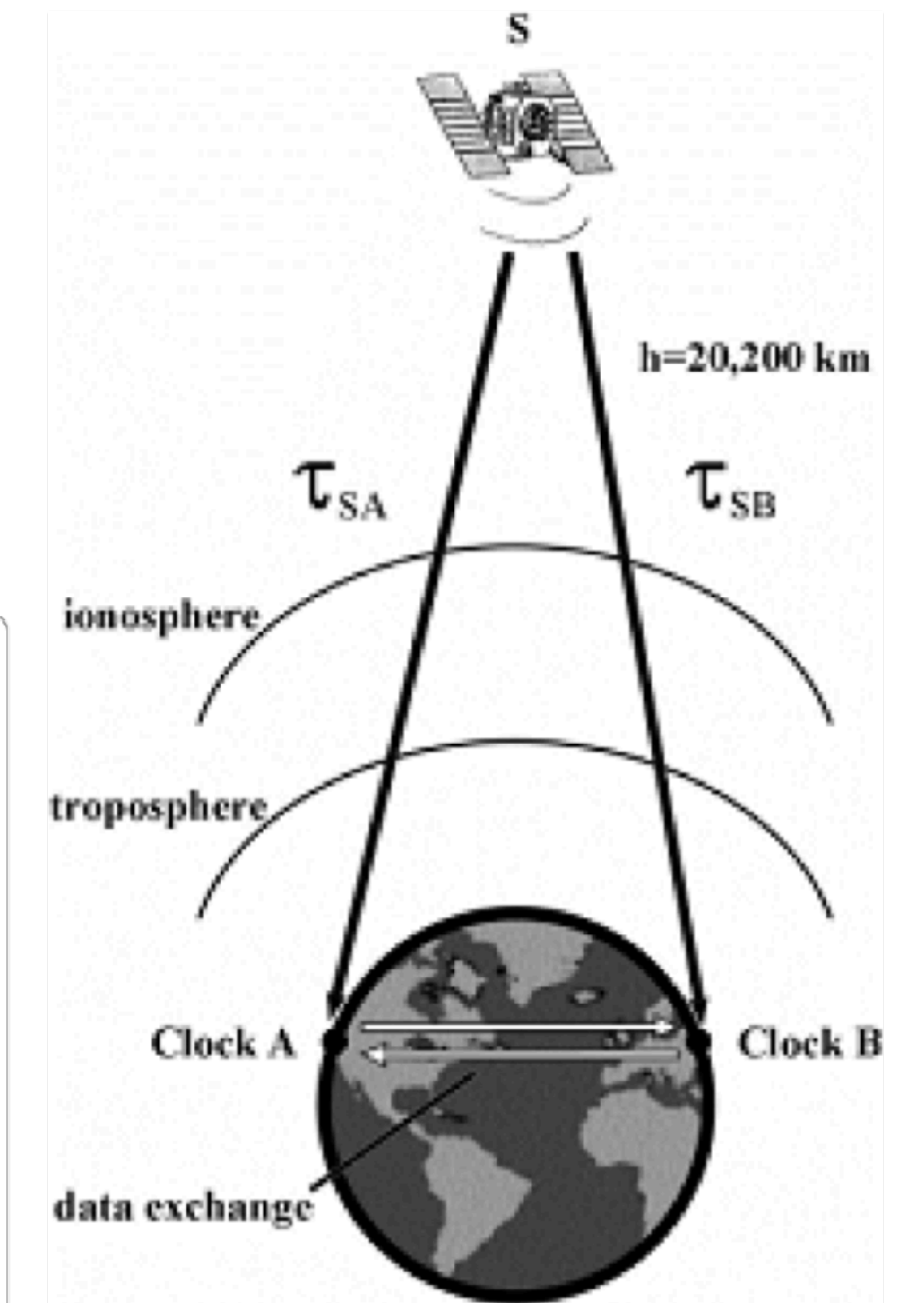
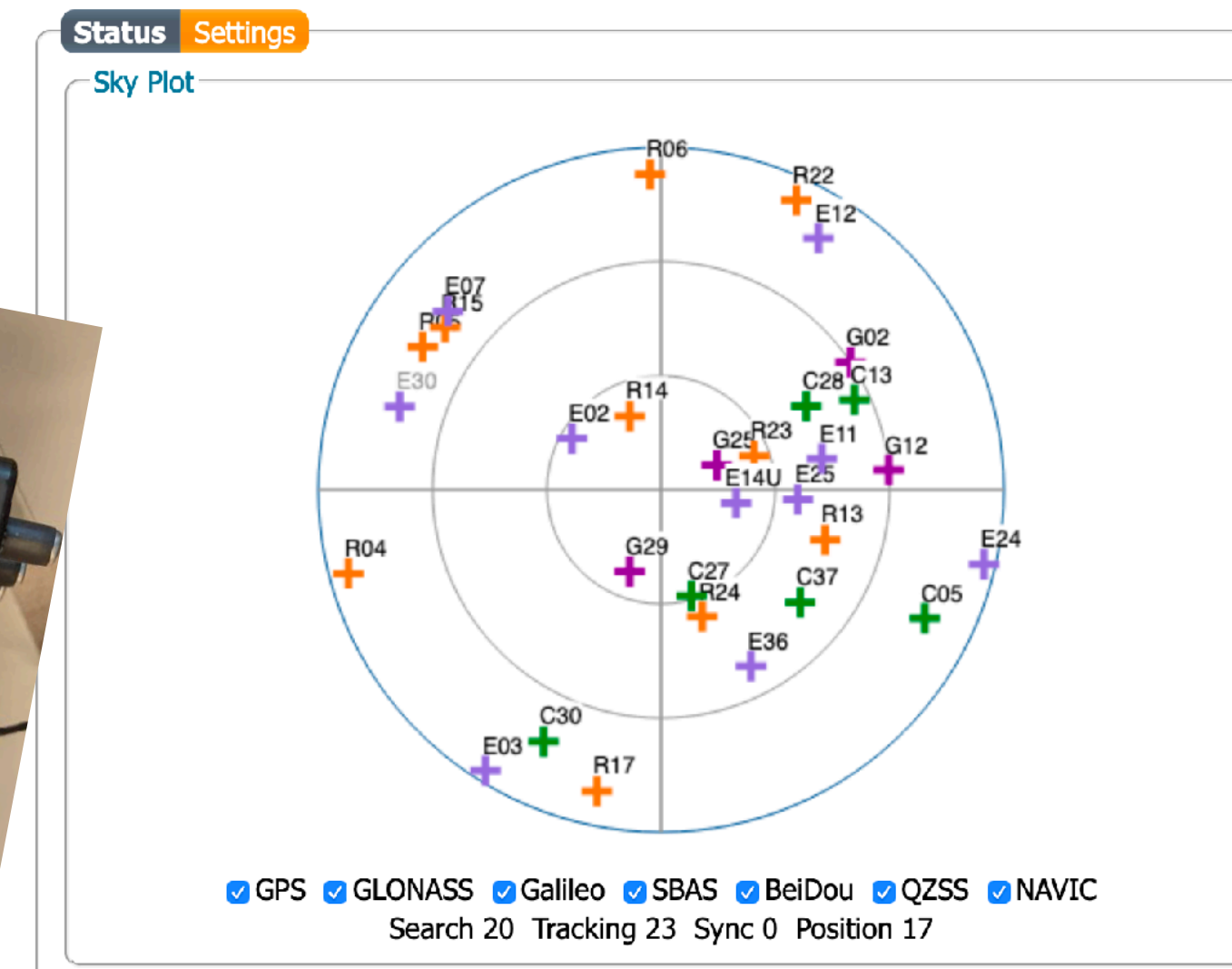
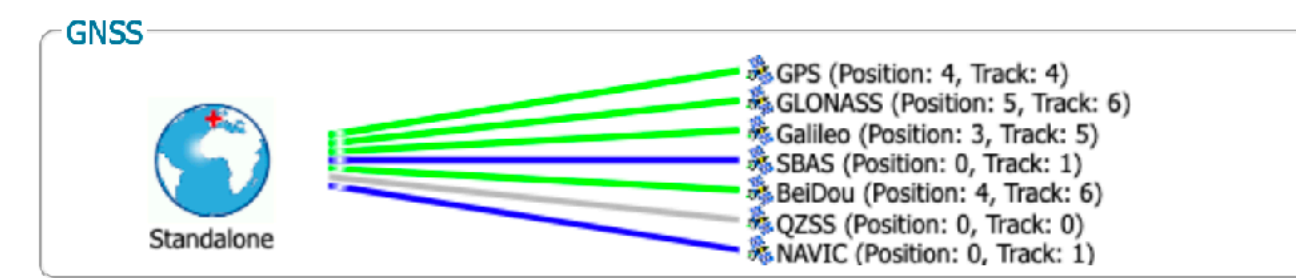


# GNSS antenna and receiver

Receive signals from various satellites e.g. Galileo

Produce 1 PPS using received signals

Synchronization with UTC (after calibration & correction)





# Tests at Jussieu and SYRTE

Dedicated tests and performances at SYRTE

Clock and GNSS calibration

Creation of a dedicated lab to study clocks and GNSS at LPNHE

Room for long-term tests (5th floor, Tower 13)

Time transfer (White Rabbit) between Syrte-Obs.Paris and Jussieu

PPS with UTC(OP) available in lab soon

Installation of GNSS antenna on the Jussieu roof (Tower 13)

Long-term studies and comparison using  
atomic clocks and PPS-Syrte

First implementation of GNSS signals corrections

Build prototypes for HK

**R&D towards HK and other IN2P3 projects**

