

Participation of IN2P3 physicists in the **Hyper-Kamiokande** experiment

ILANCE, LLR, LPNHE, OMEGA

IN2P3 physicists and engineers

ILANCE: M. Gonin, G.Pronost (postdoc)

LLR: A.Afiri, M. Buizza-Avanzini, O. Drapier, F. Gastaldi, M.Louzir, T. Mueller, J. Nanni, P. Paganini, B. Quilain, A. Beauchene (PhD student)

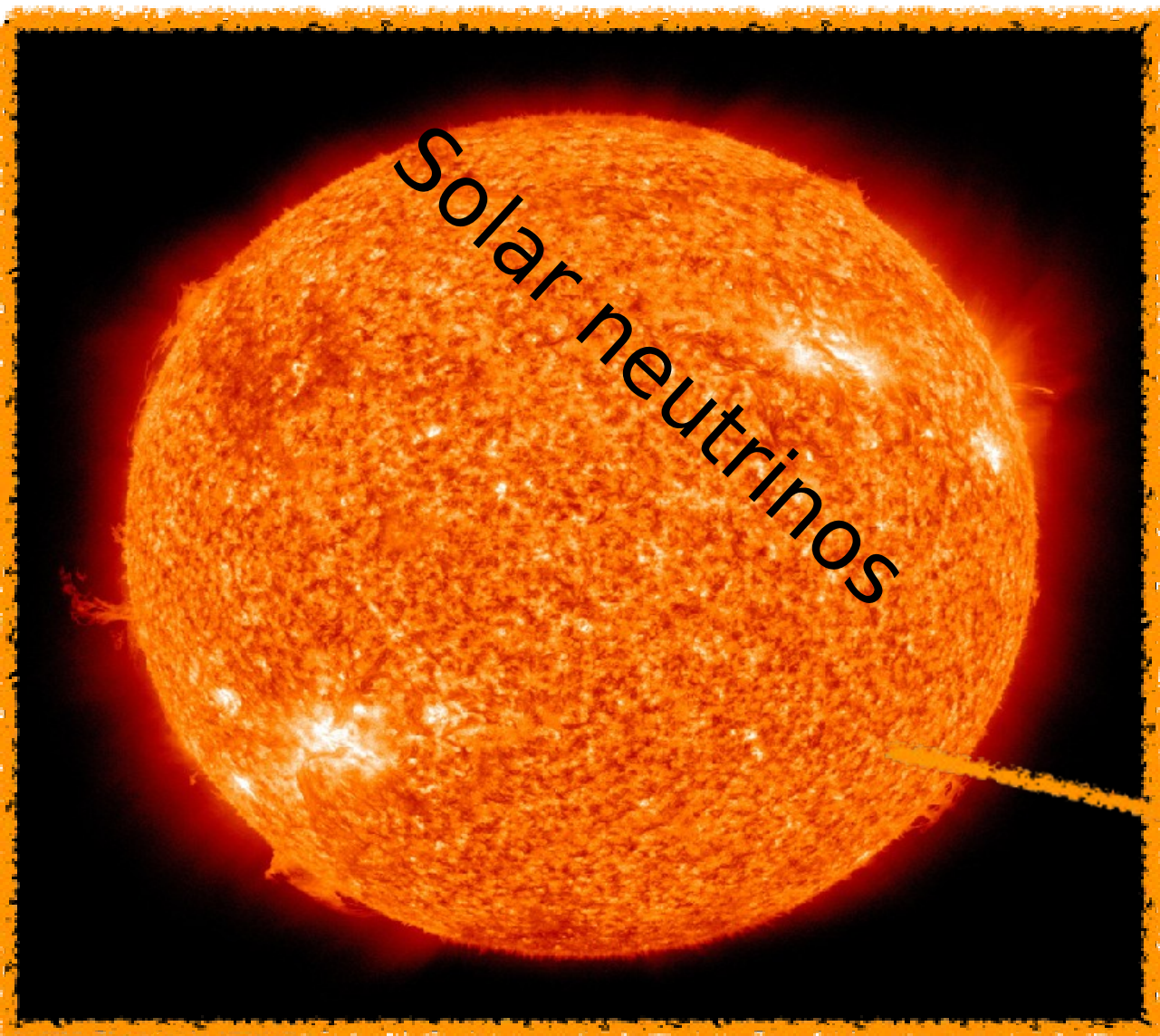
LPNHE: A. Blondel, J. Dumarchez, C. Giganti, M. Guigue, M.Martini, B. Popov, S. Russo, V. Voisin, M. Zito, W.Saenz (postdoc ANR), L. Mellet (PhD student), C.Dalmazzone (PhD student) + one more ANR postdoc starting from 2023

OMEGA: S. Callier, P.Dinaucourt, S. Conforti, F. Dulucq, L. Raux, C. de la Taille

APC: C.Volpe

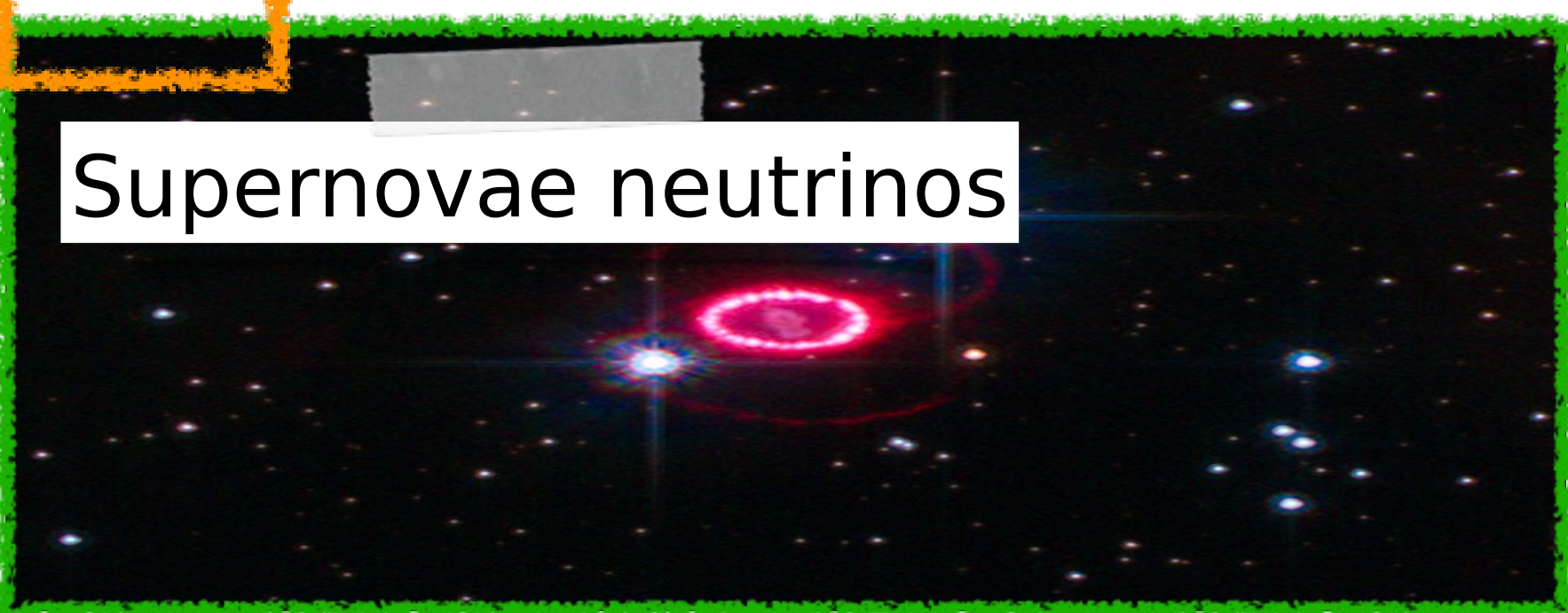
Natural continuation of our participation in the Japanese neutrino program (T2K and SK)

Hyper-Kamiokande physics program

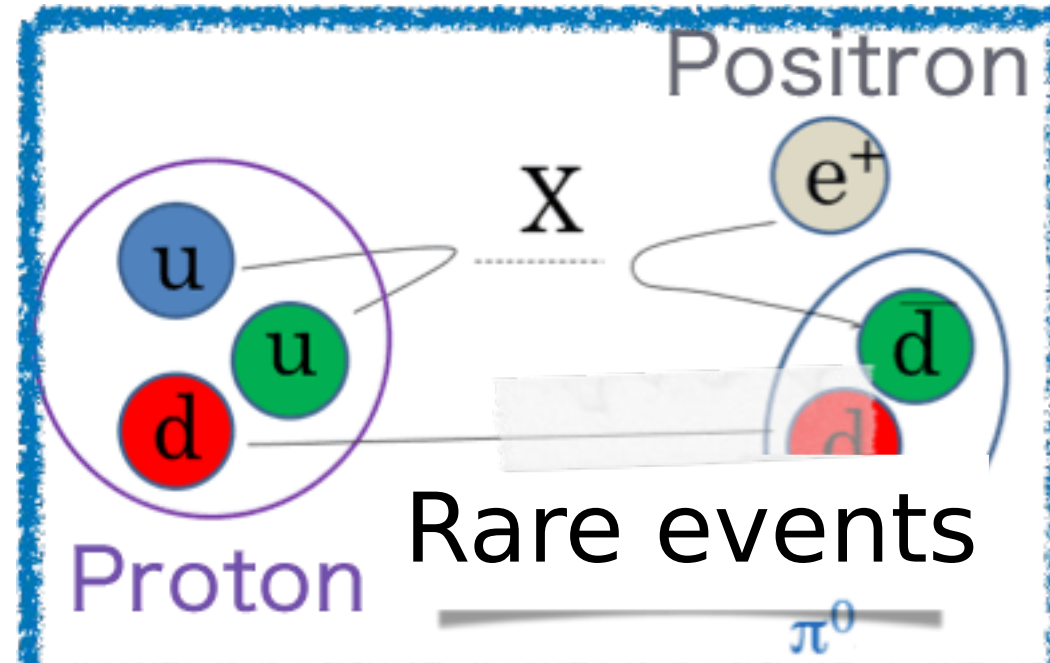


ν_e

MSW effect
Non-standard interactions

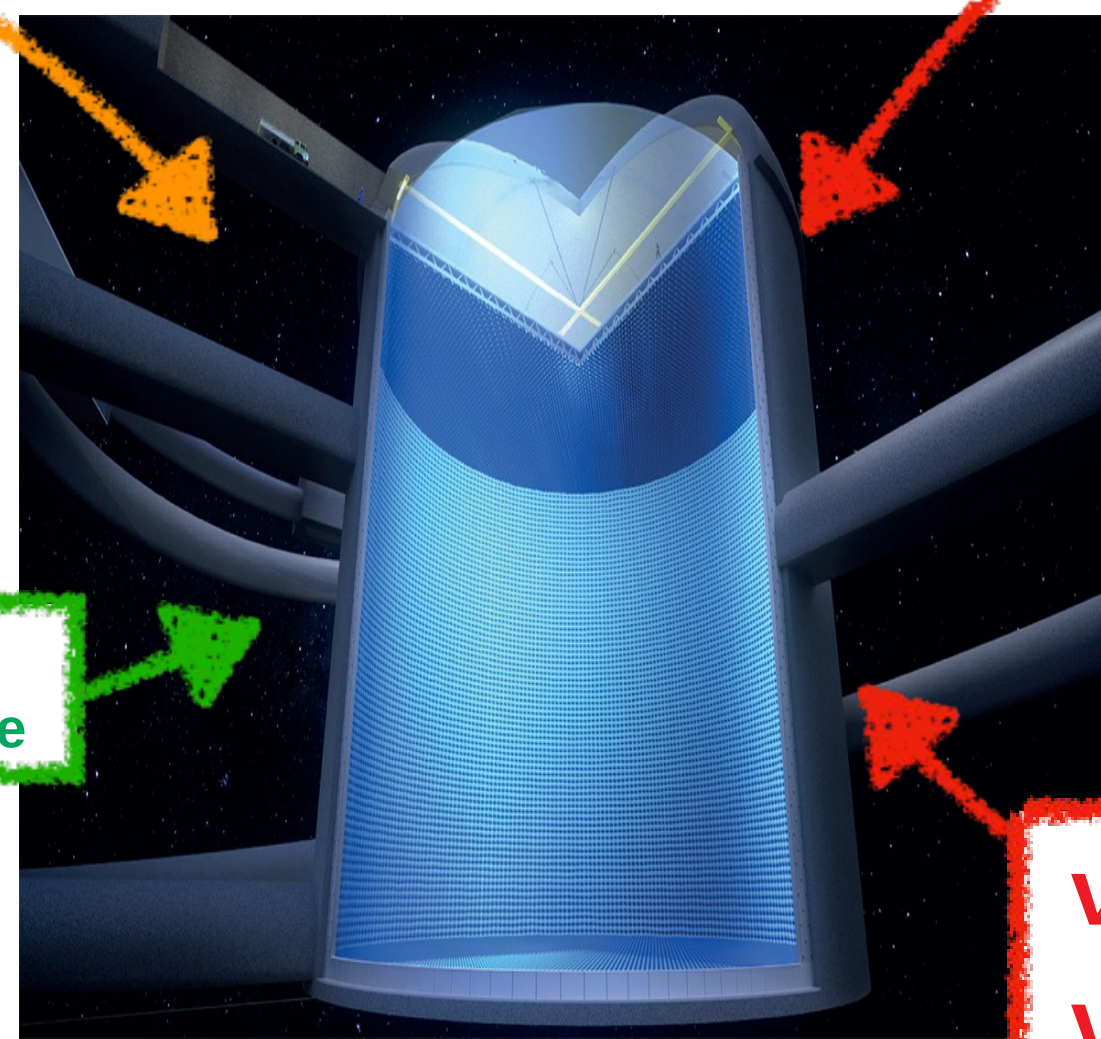


- Transient SNv : constrain SN profile models
- Relic SNv : constrain cosmic star formation



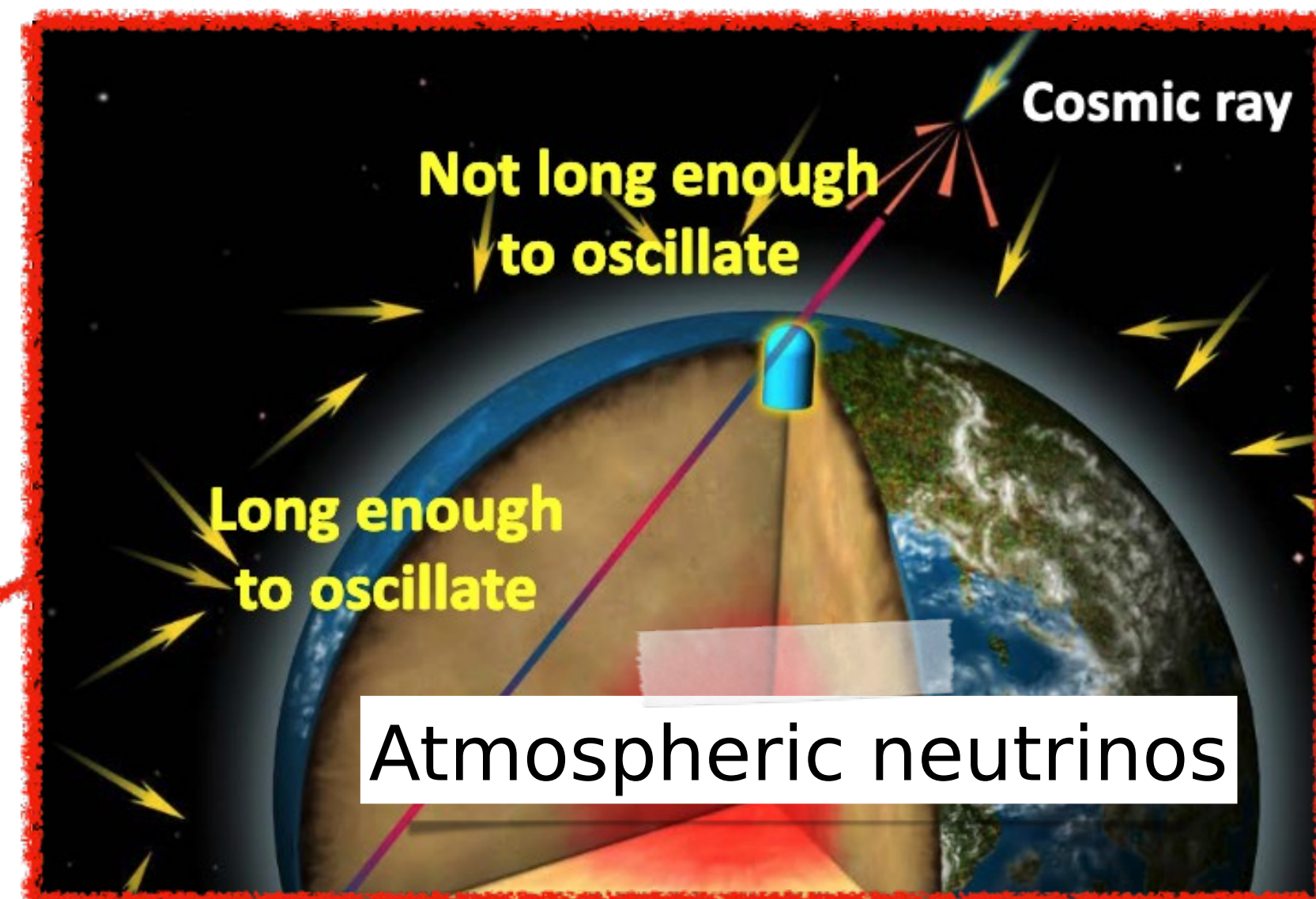
- **Probe Grand Unified Theories via p-decay or n- \bar{n} oscillation**

$\nu_e \bar{\nu}_e$
 $\nu_\mu \bar{\nu}_\mu$

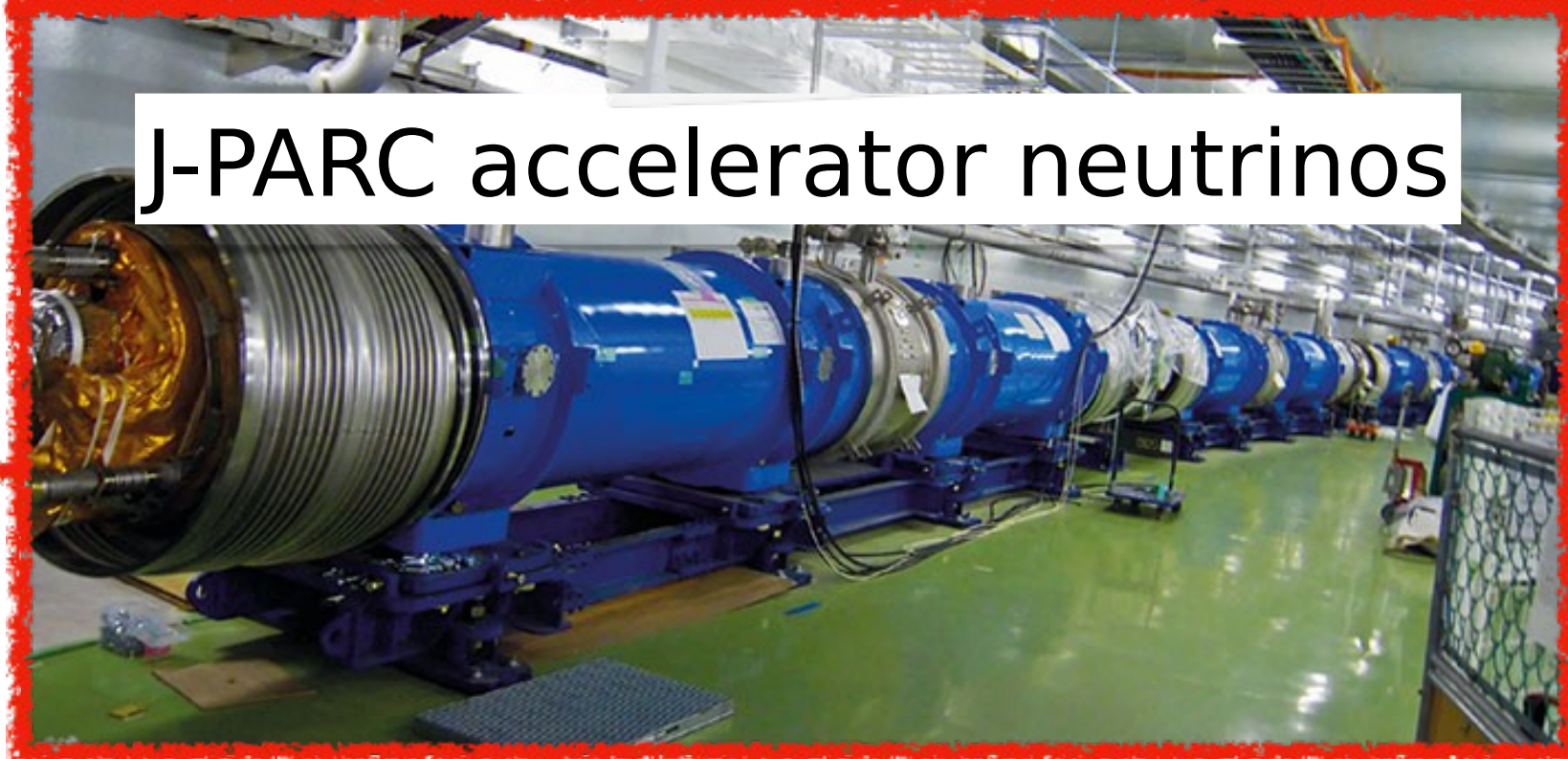


$\bar{\nu}_e$

$\nu_e \bar{\nu}_e$
 $\nu_\mu \bar{\nu}_\mu$



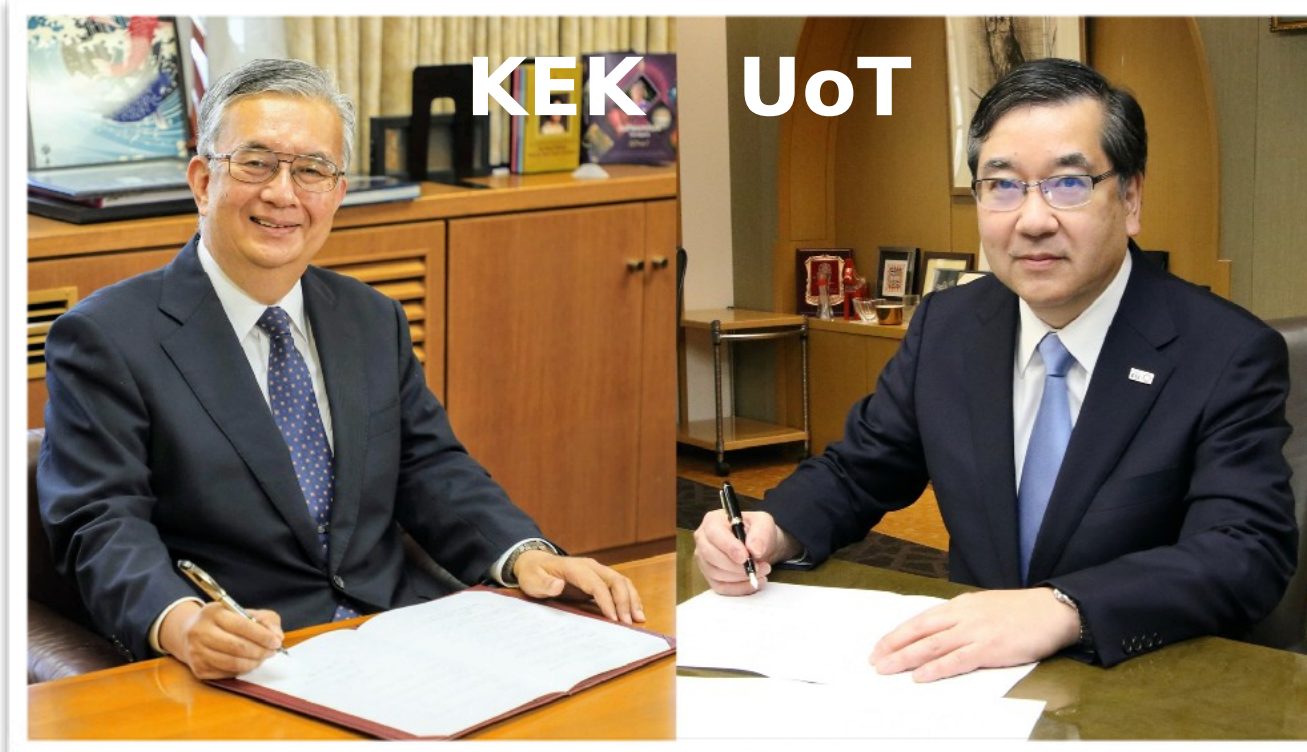
- **Observe CP violation for leptons at 5σ**
- **Precise measurement of δ_{CP}**
- **High sensitivity to ν mass ordering**



J-PARC accelerator neutrinos

Experiment approval: key dates

Conclusions of 2018 SC : “Le projet n’est pas actuellement approuvé au Japon et il n’y a pas suffisamment d’informations quant à l’organisation du projet pour envisager et discuter des participations directes à HK.”

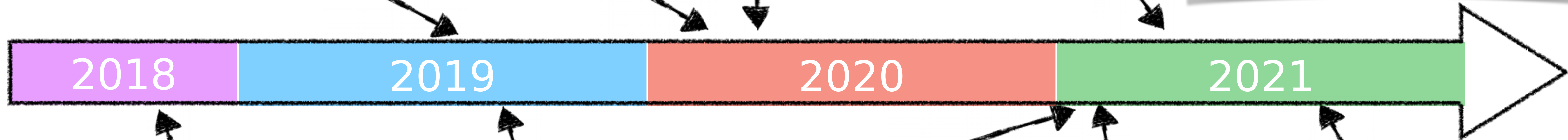
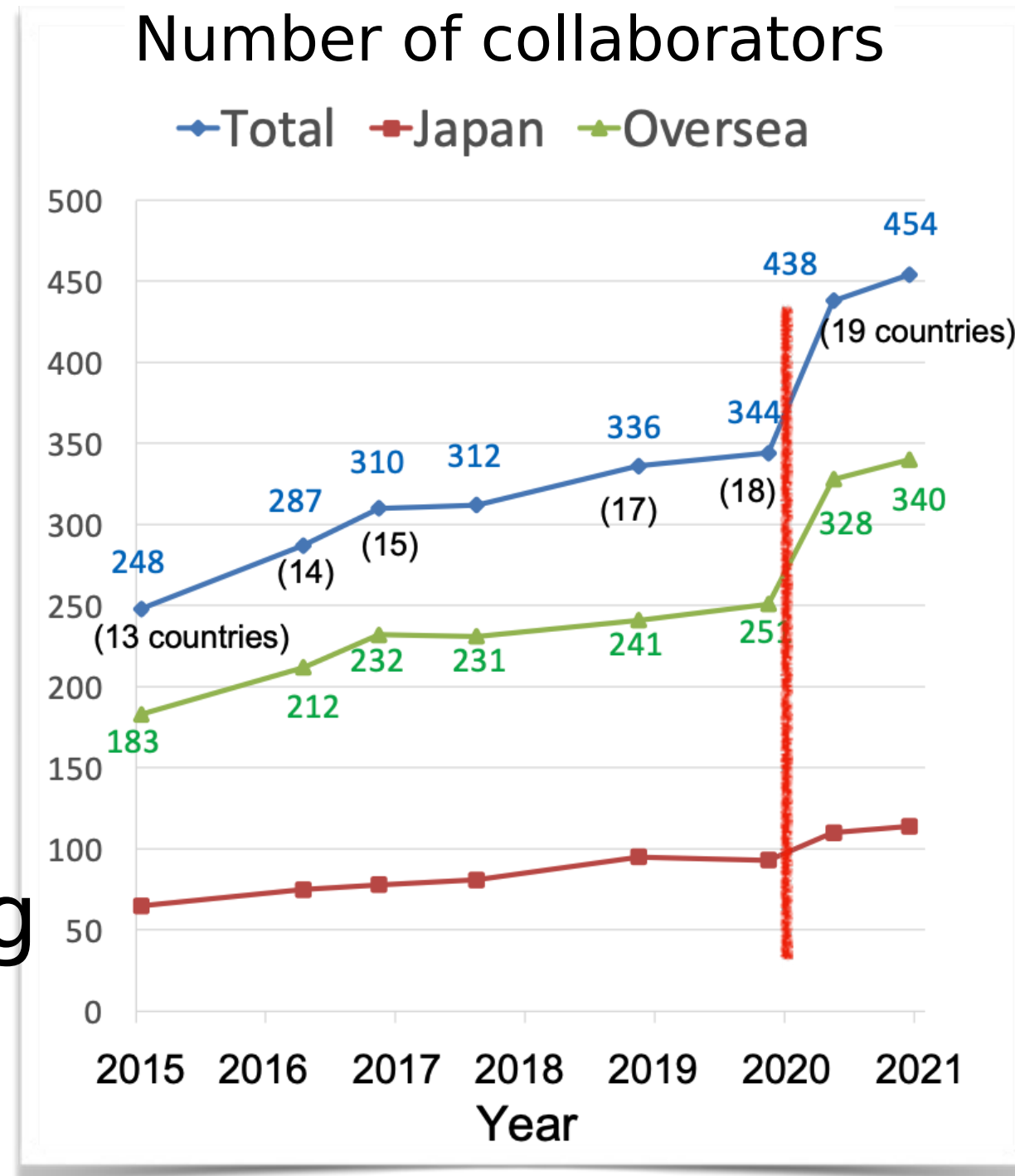


May 2020:

Aug. 2019: MEXT approval of HK project

Feb. 2020: HK budget approved

May 2021: Groundbreaking ceremony



Oct. 2018: IN2P3 CS discussing HK

Oct. 2019: LPNHE CS approved HK

Jan. 2021: CEA CS approved HK

March 2021: LLR CS approved HK

Oct. 2021: IN2P3 CS discussing HK

Conclusions of the 2021 SC

6.3. Avis du Conseil

La dernière présentation de la contribution de l'IN2P3 à HK en session du Conseil Scientifique de l'IN2P3 est assez récente (octobre 2018). En octobre 2021, le Conseil constate une évolution positive remarquable au cours des deux dernières années :

- Le projet HK est approuvé par le gouvernement japonais en août 2019, le budget (500 M\$) est voté en février 2020, et le début de la prise de données est confirmé pour 2027 ;
- Deux laboratoires de l'IN2P3 (LPNHE, LLR) et le CEA, soutenus par leurs conseils scientifiques respectifs, rejoignent le projet HK, en octobre 2019 pour le LPNHE, en janvier 2021 pour le CEA, et en mars 2021 pour le LLR.

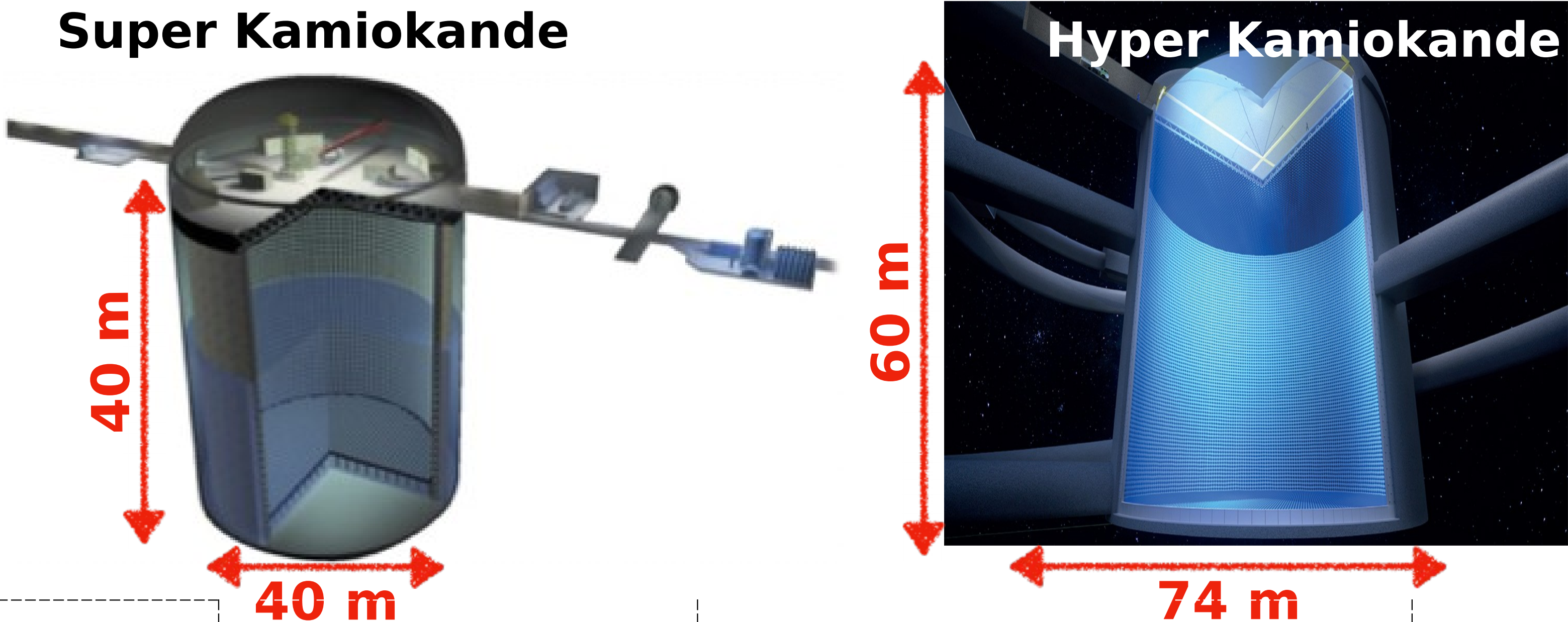
...

Le Conseil souligne cependant qu'un engagement technique direct sur le détecteur lointain de HK est requis pour confirmer et valider le ticket d'entrée de l'IN2P3 dans HK⁴, et considère que la participation centrale proposée par le LPNHE et le LLR en collaboration avec OMEGA sur le digitaliseur HKROC et sur le système de distribution d'horloge répond à cette condition dans une enveloppe budgétaire raisonnable. Le Conseil note également que les équipes de l'IN2P3 proposent une réflexion pour contribuer aux coûts d'opération à travers la participation du CC-IN2P3.



RD4HK starting from 2022

Hyper-Kamiokande vs Super-Kamiokande



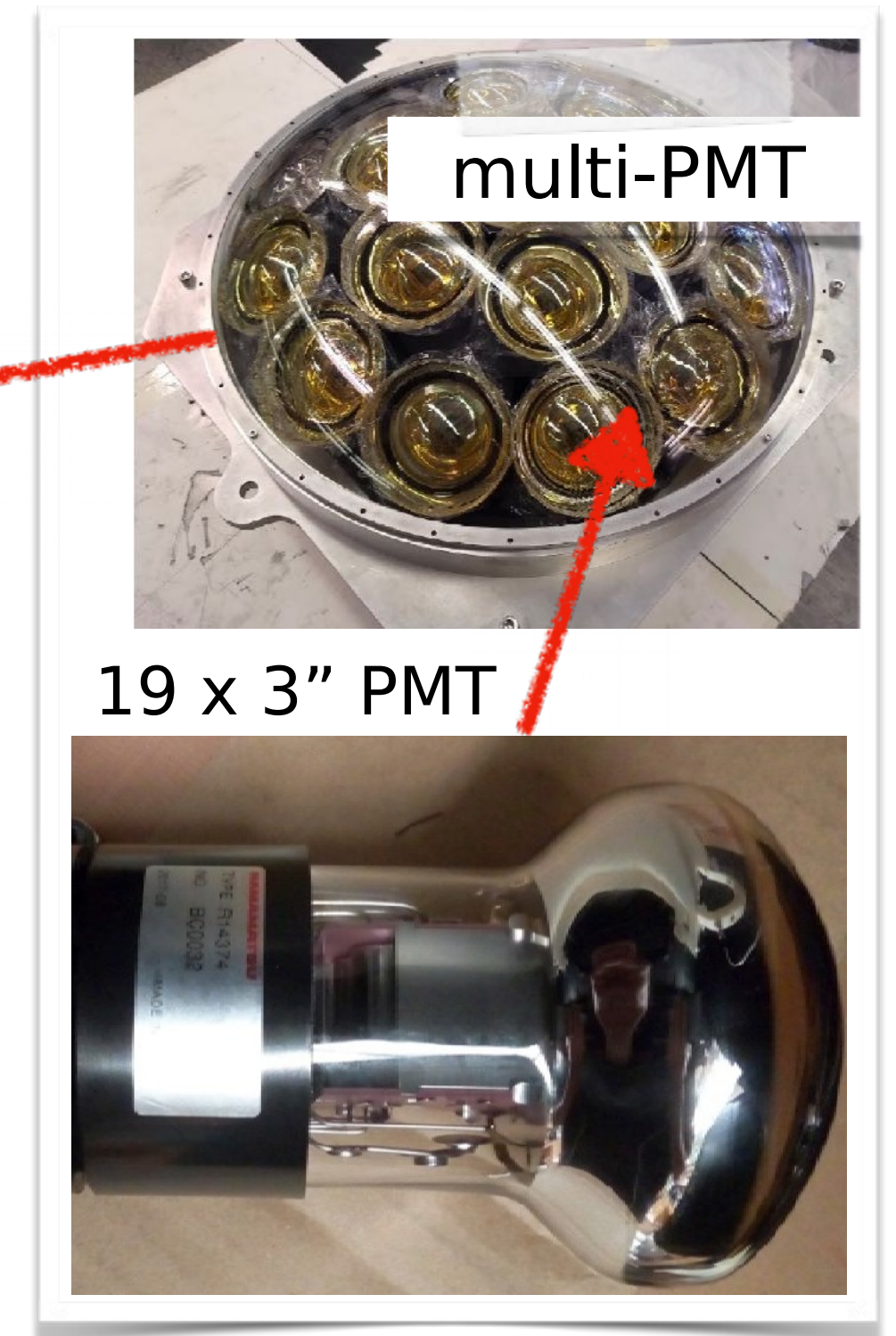
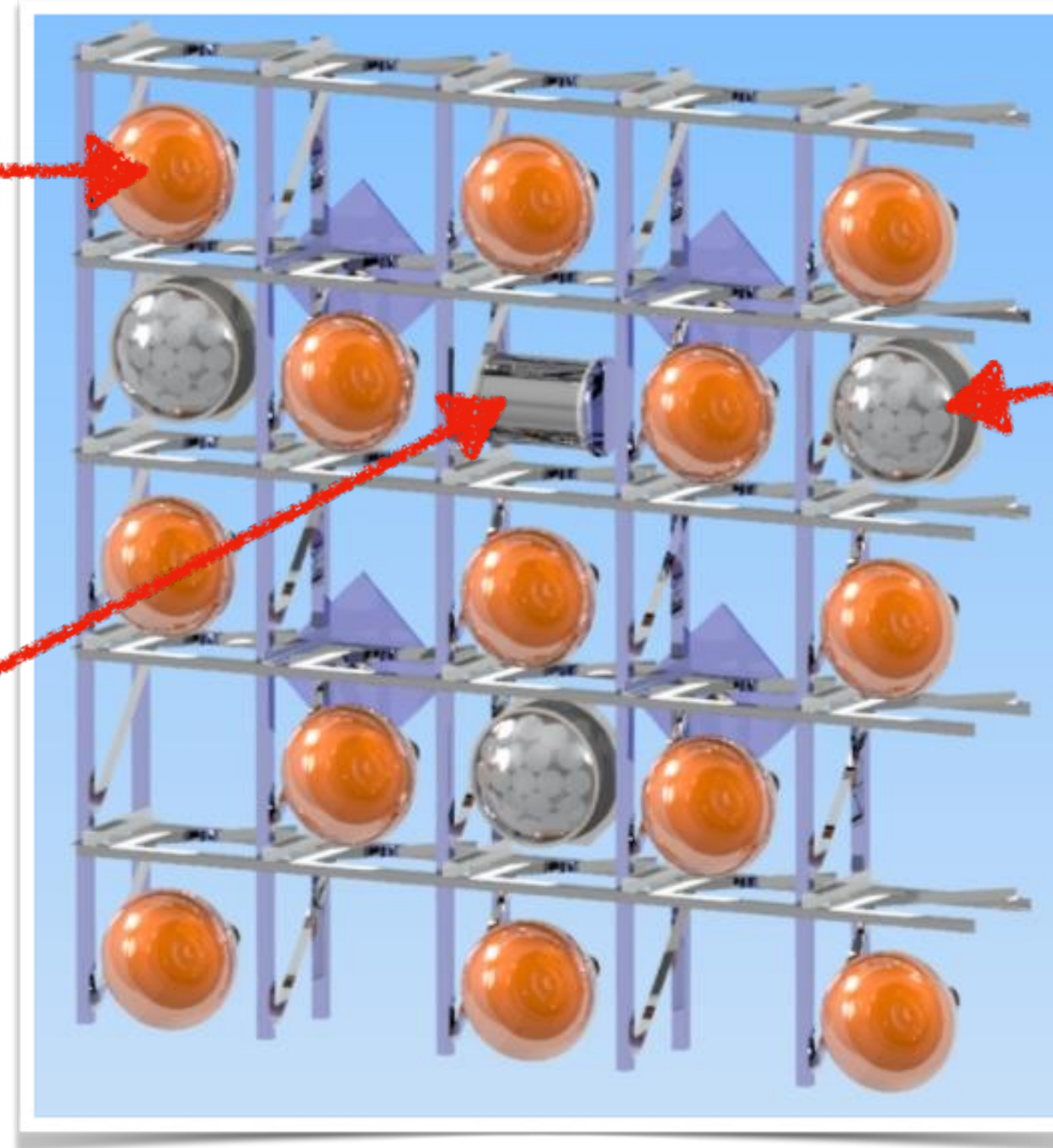
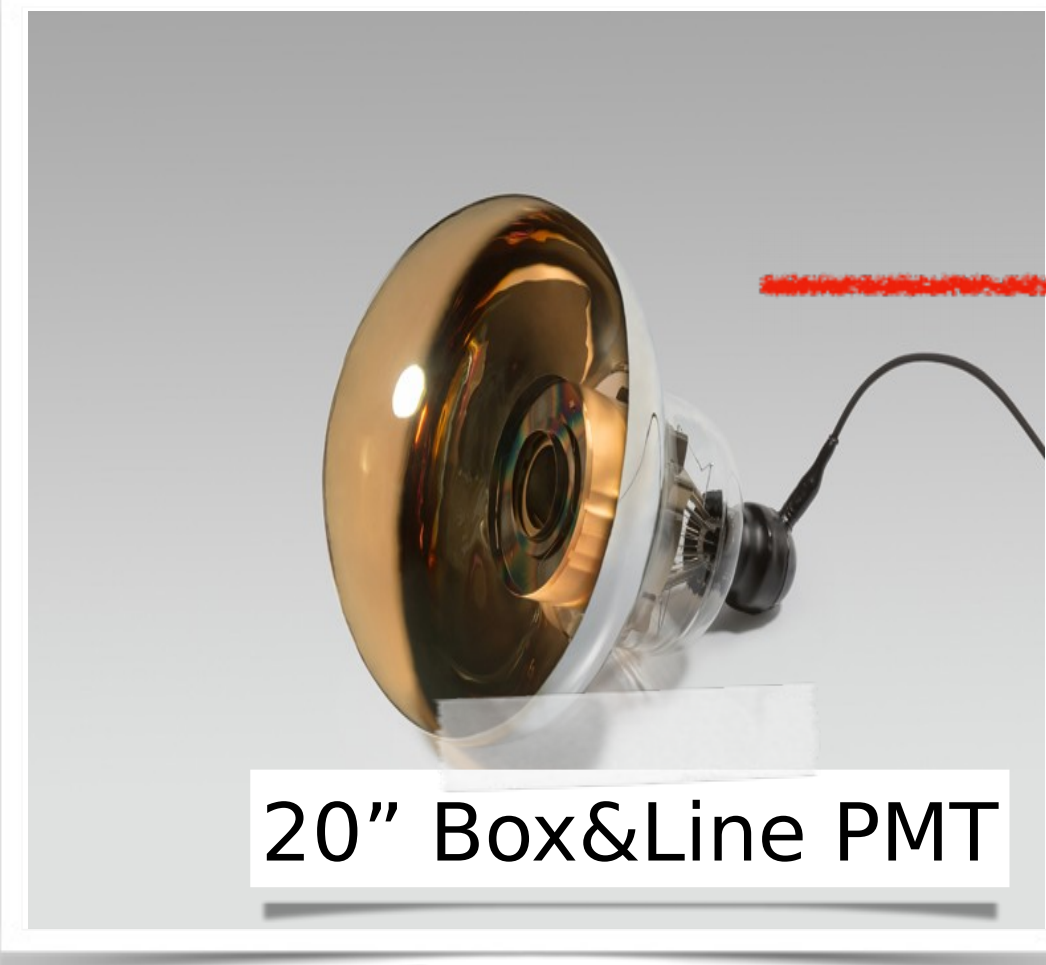
| | Super Kamiokande | Hyper Kamiokande |
|------------------------------|------------------|------------------|
| Site | Mozumi-yama | Tochibora-yama |
| Number of ID 20'' PMTs | 11 129 | >20,000 |
| Photo-coverage | 40 % | >20% |
| Single-photon efficiency/PMT | ~12% | ~24% |
| Dark rate/PMT | ~4 kHz | ~4kHz |
| Time resolution of 1 photon | ~3 ns | ~1.5 ns |
| Total/fiducial mass (kton) | 50 / 22.5 | 260 / 187 |

Fiducial volume x8:
→ non-beam ν physics

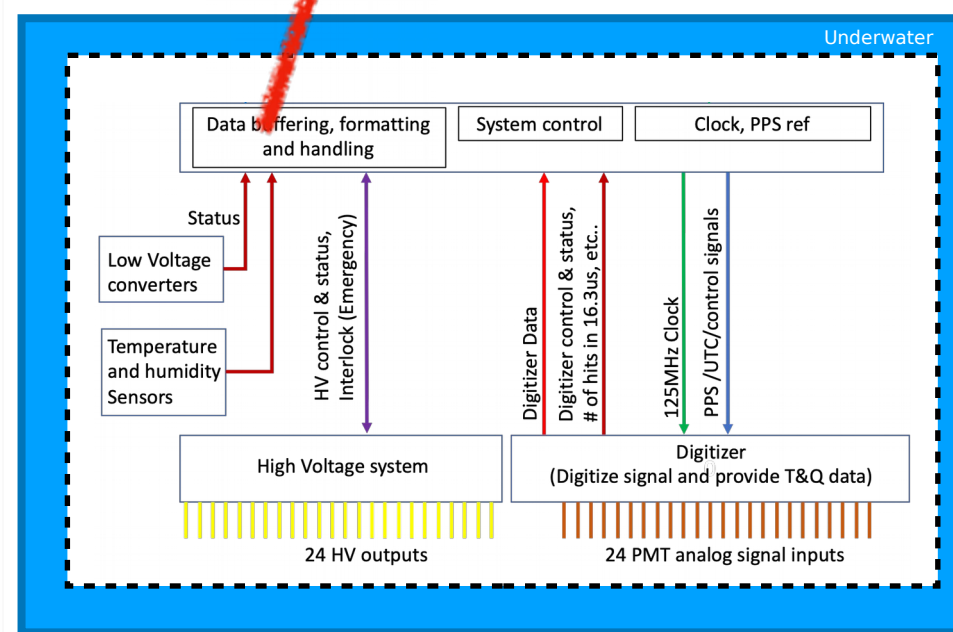
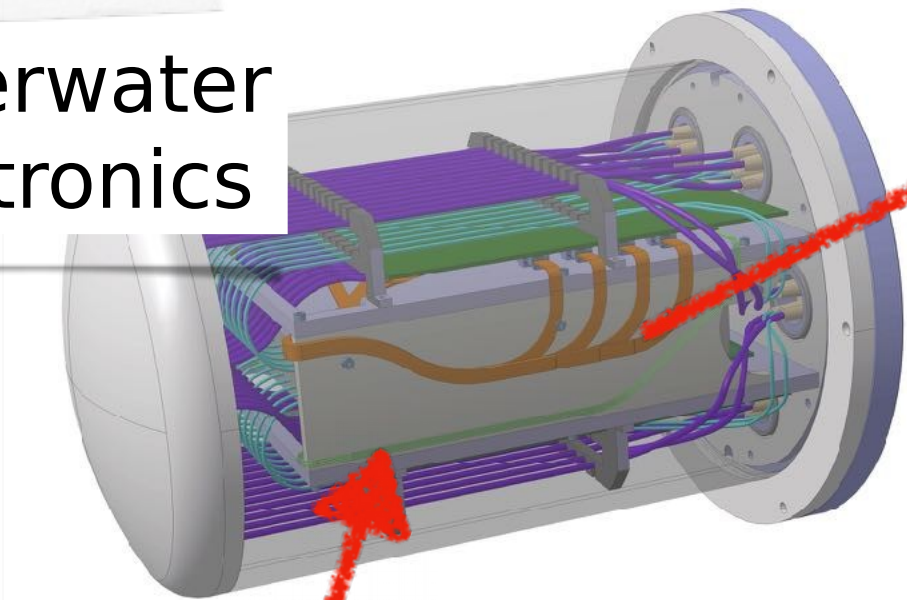
Beam neutrino event rate x 20:
→ beam ν physics

Start operations in 2027 with 240 kt.MW and an assumed runtime 10^7 s per year

20" PMTs, mPMTs and readout



Underwater
electronics



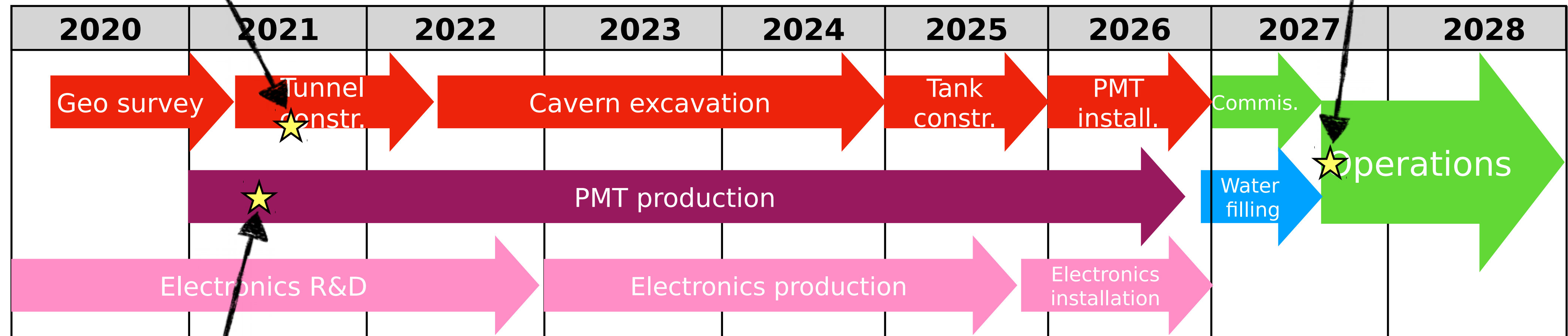
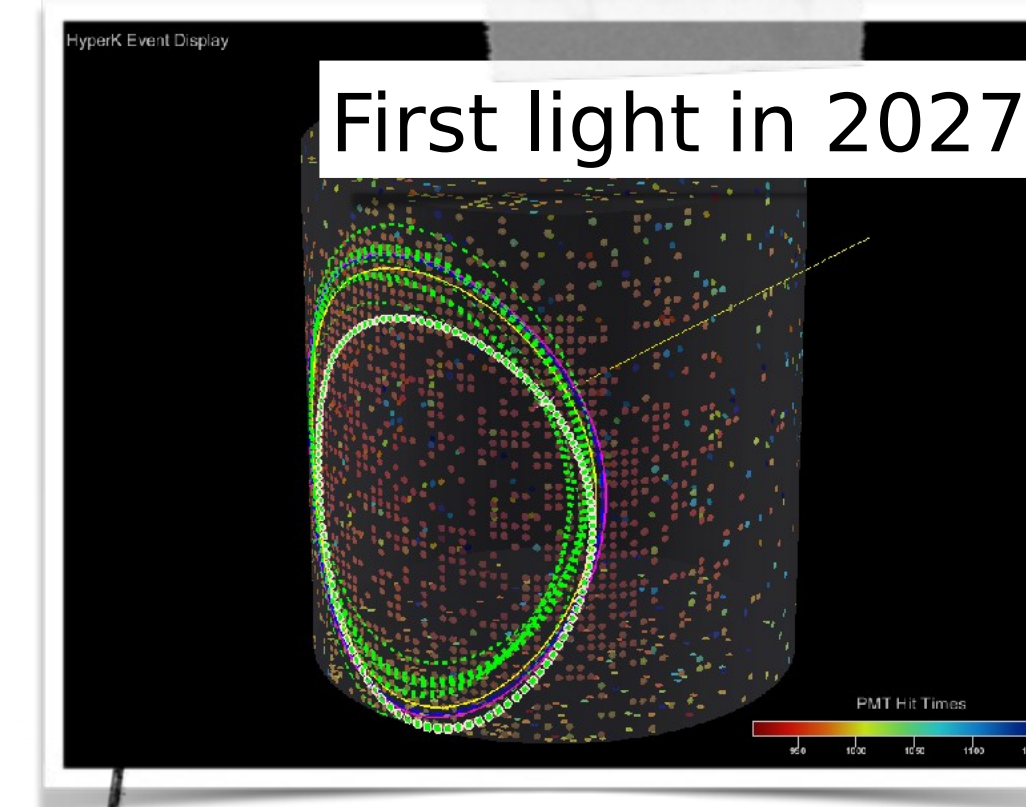
- Inner detector composed of
- 20k+ 20" PMTs (Hamamatsu R12860)
 - ~1k mPMTs (19 3" R12199-02 PMTs)

Better SNR, directionality, timing

NEW

Size of detector requires an in-water electronics

HK schedule



Strong engagement of Japan: ~500 M\$ for construction
Expected from other countries: ~100 M\$

International contributions are being formalized

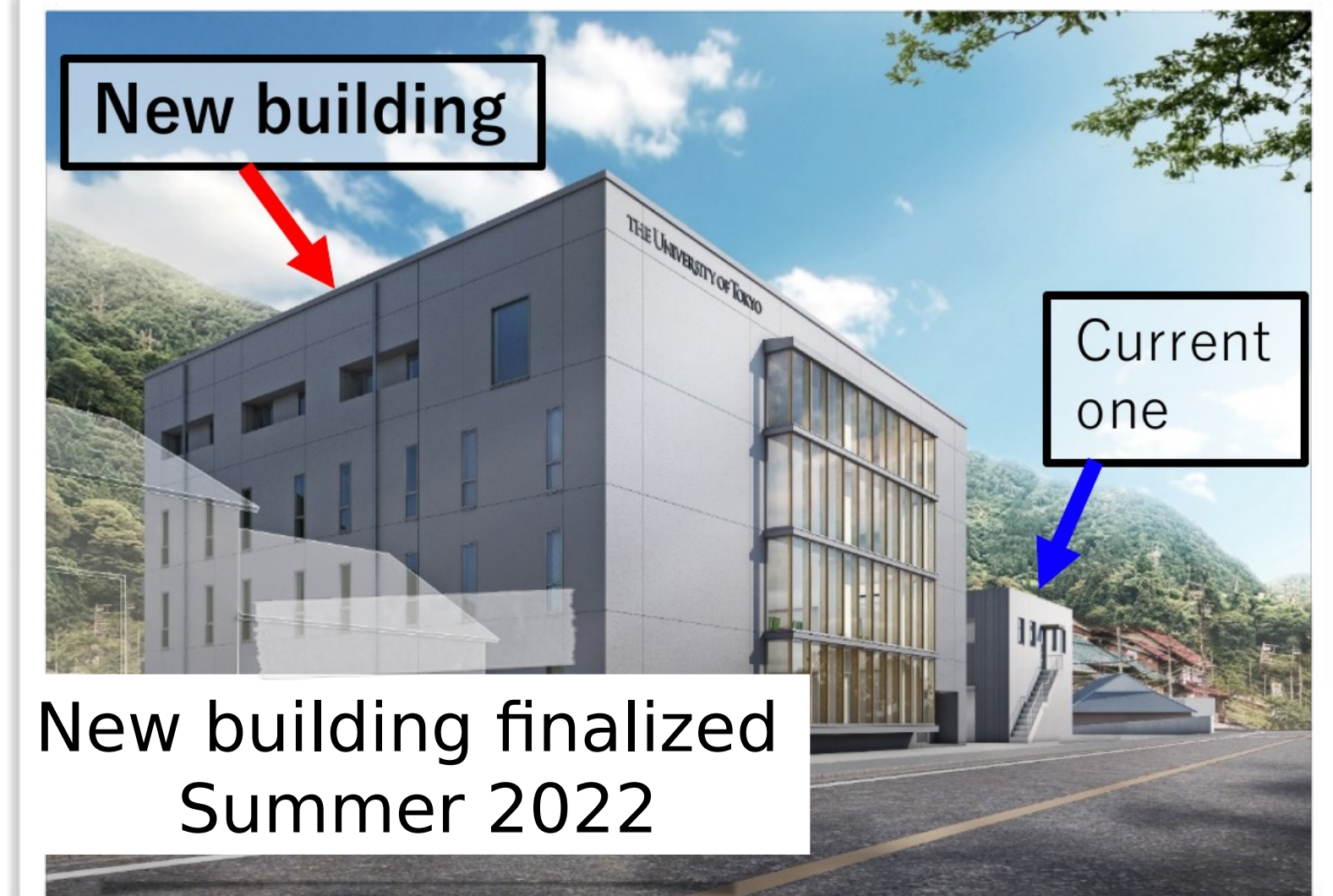
Construction is on schedule

Entrance yard finished (Aug. 2020)



New building

Current one



New building finalized
Summer 2022

Groundbreaking
ceremony
(May 2021)



Tunnel entrance (June 2021)



Survey completed ✓

Excavation on-going

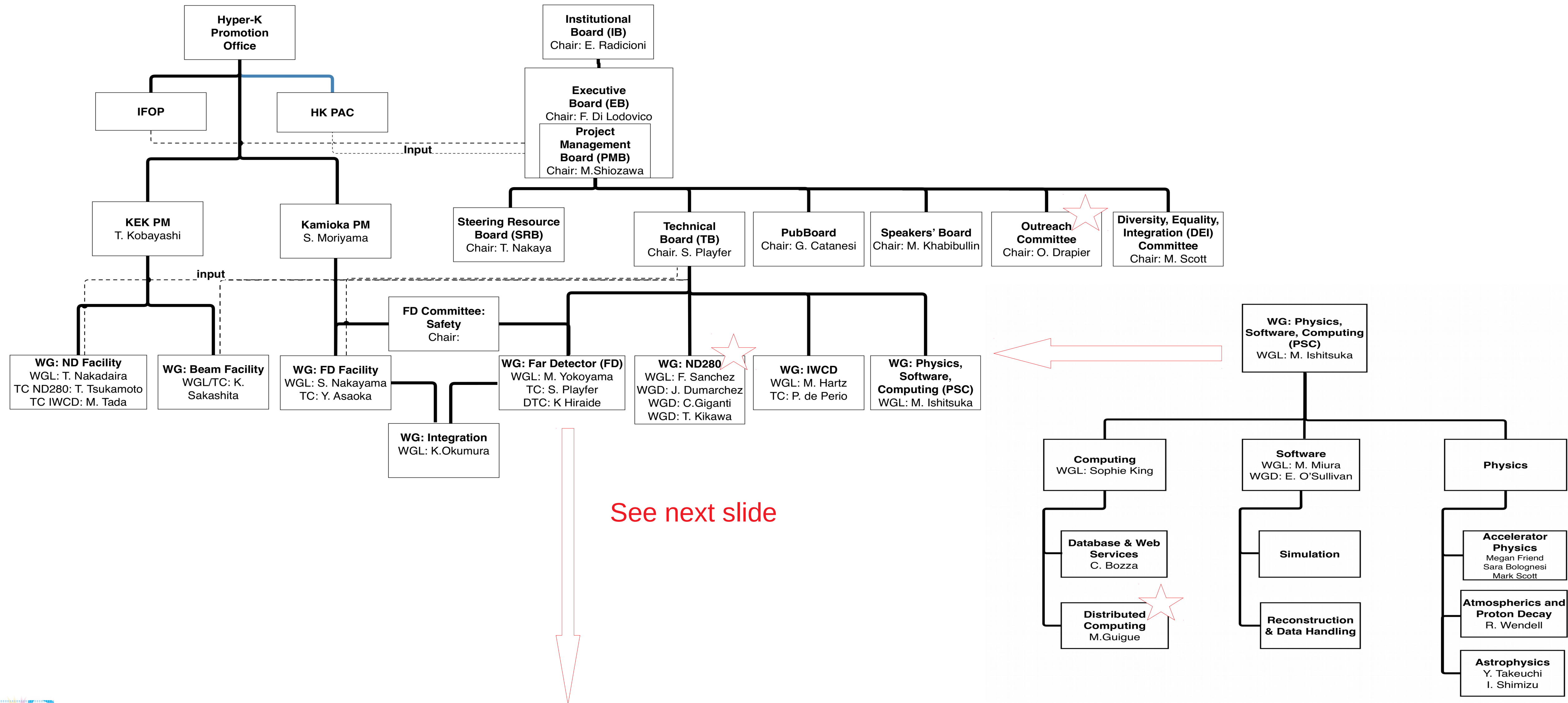


New main building



Everything on track! ✓

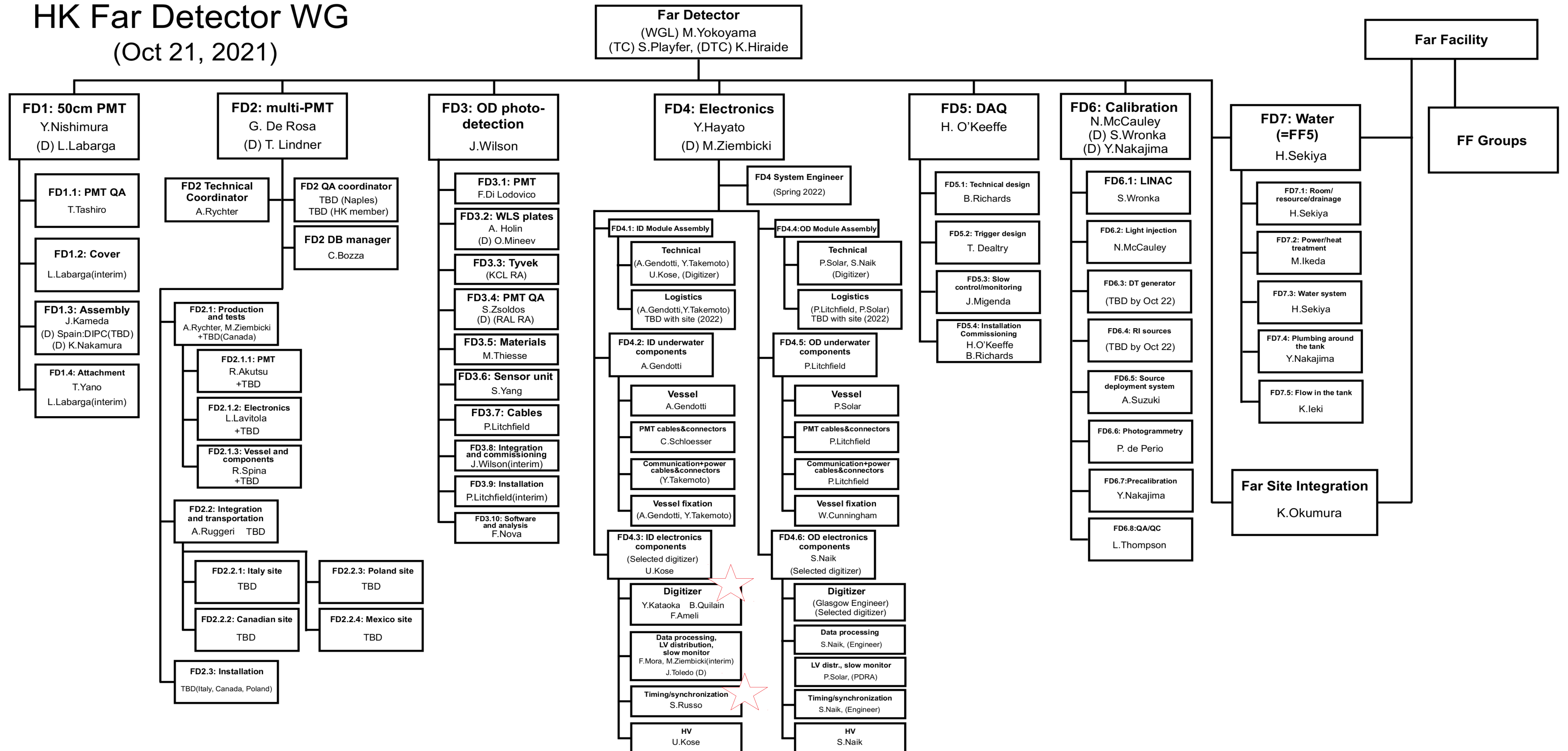
HK organizational chart



See next slide

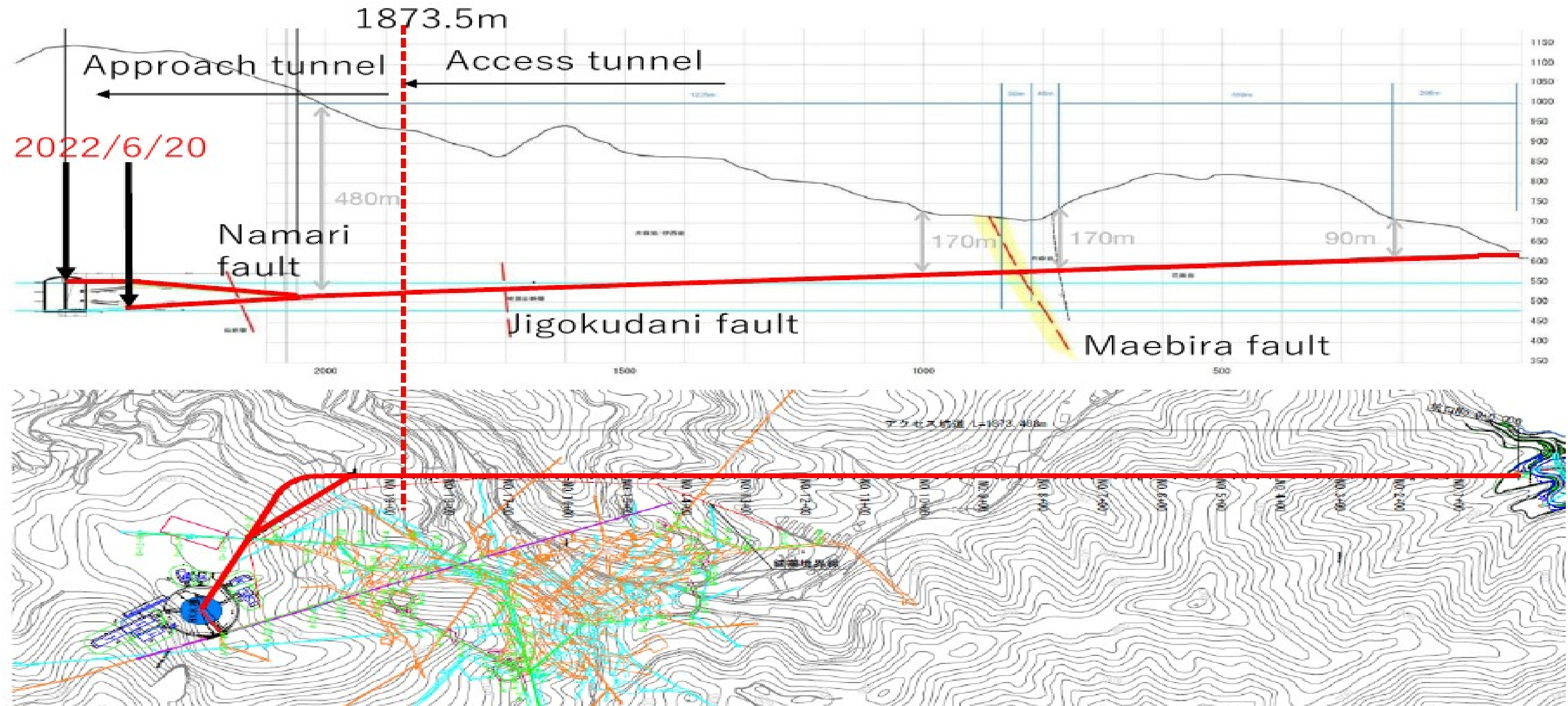
Far detector Working Groups

HK Far Detector WG (Oct 21, 2021)



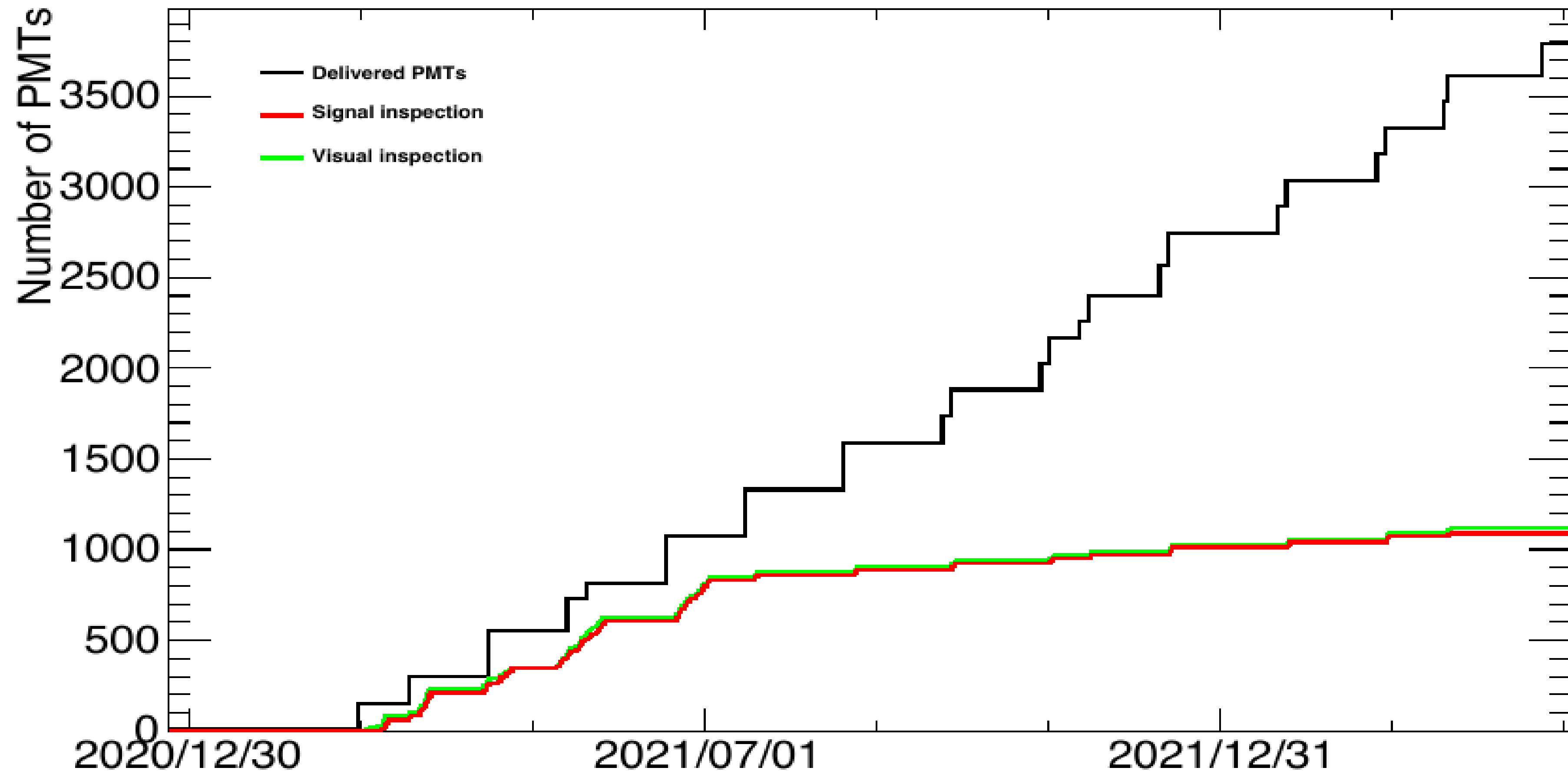
Recent progress: excavation

Tunnel excavation overview



The dome center was reached on 24th of June 2022!

Recent progress: 20'' PMT production



> 3700 20'' PMTs already produced (at the beginning the signal and visual inspection was done for almost all PMTs). Since July, 2021 the inspection rate is 10%.

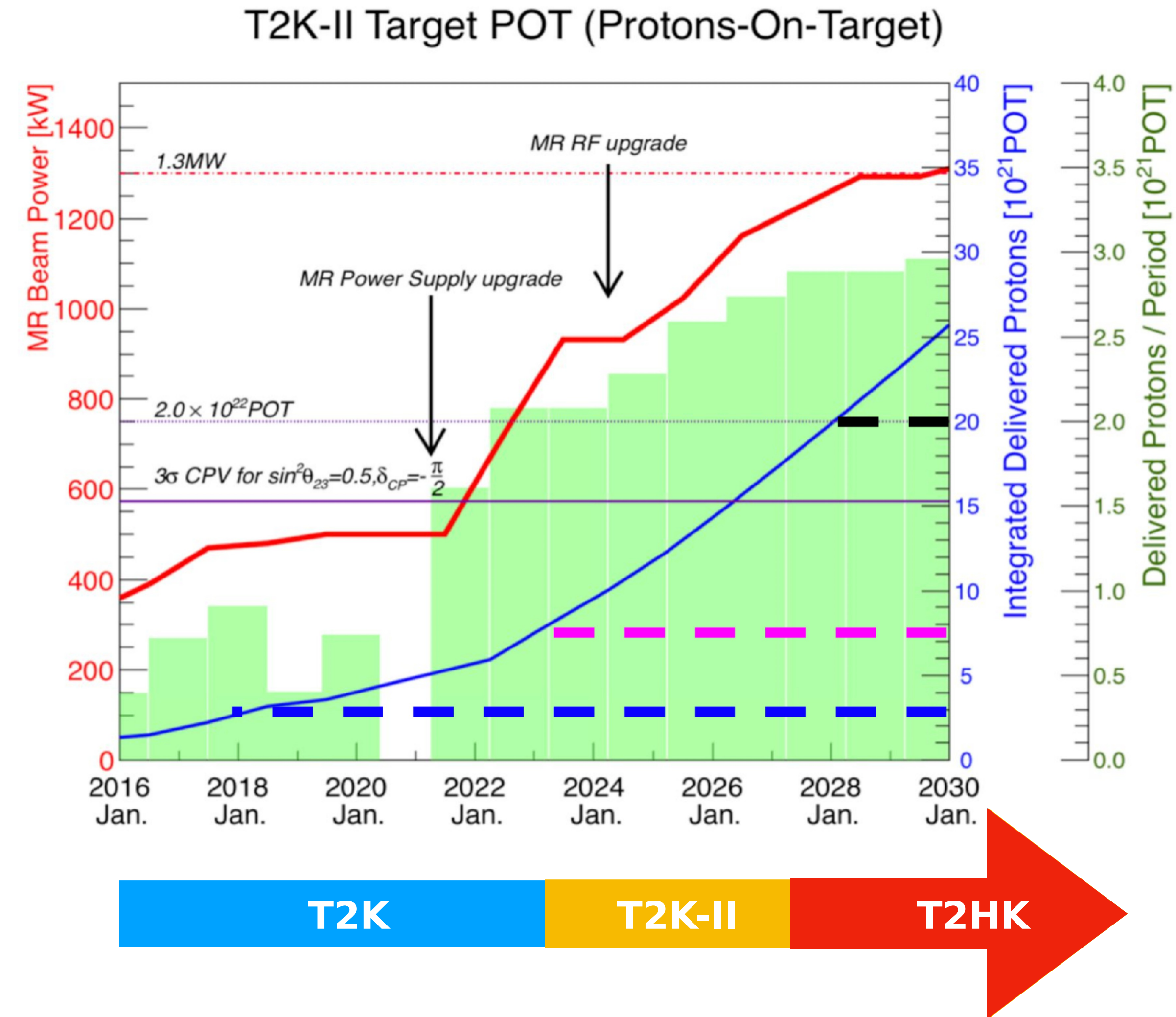
Tokai to HK: heritage from T2K

Accelerator upgrade

Power increase (500kW --> 1.3 MW)

x2.7 more stats per s (wrt T2K-I)

$\nu / \bar{\nu}$ flux uncertainty < 5% thanks to NA61



Tokai to HK: heritage from T2K

Accelerator upgrade

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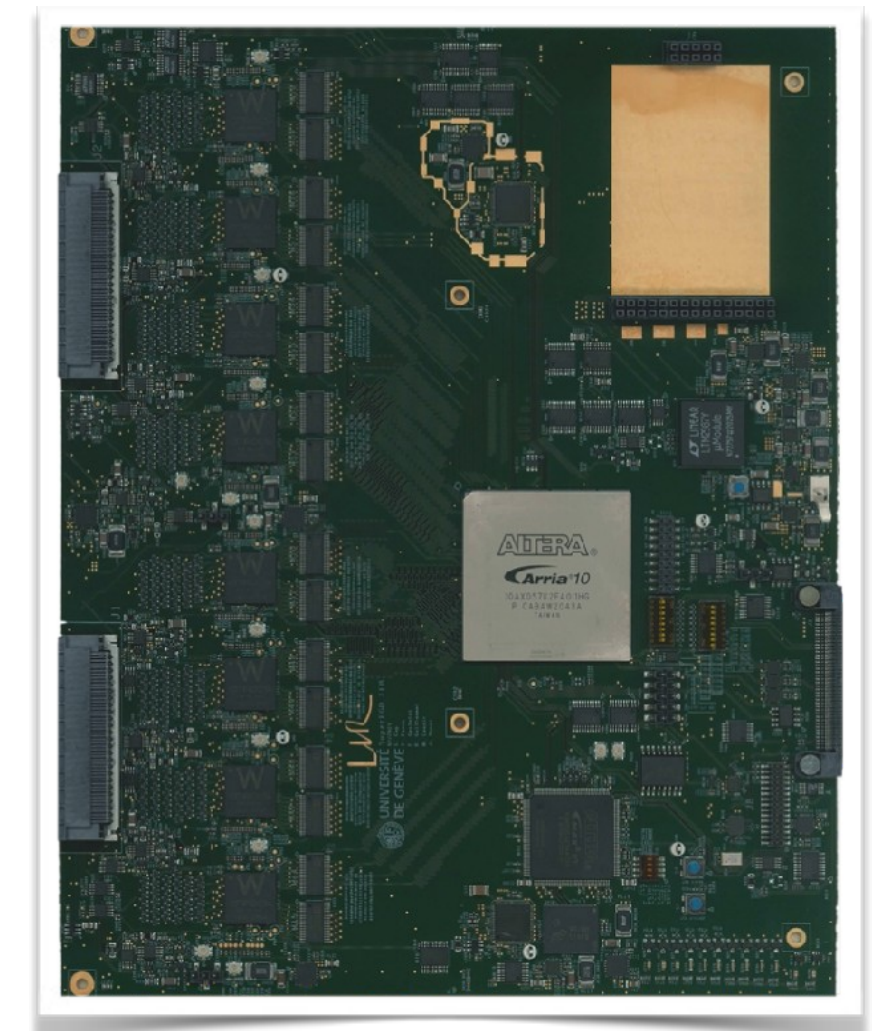
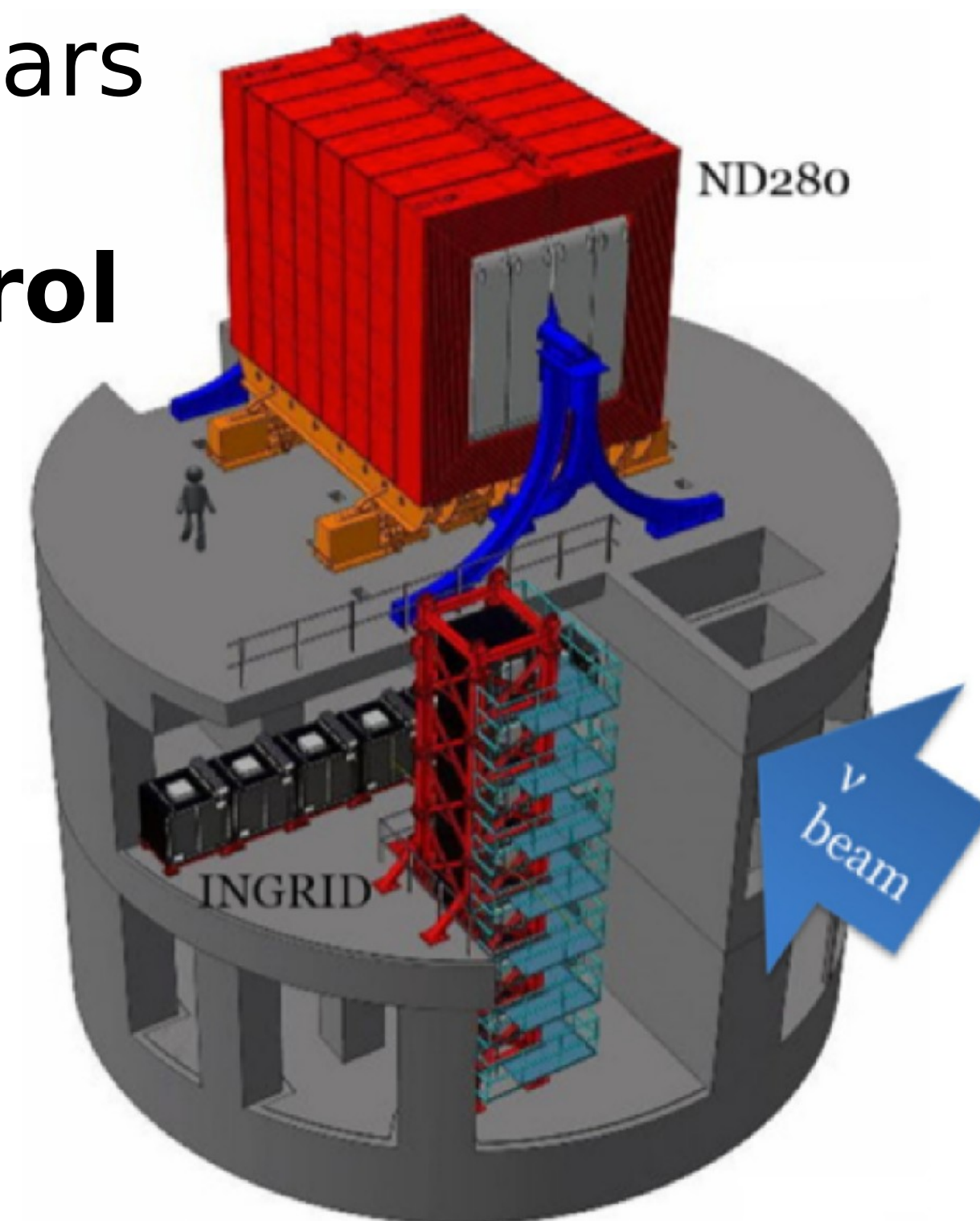
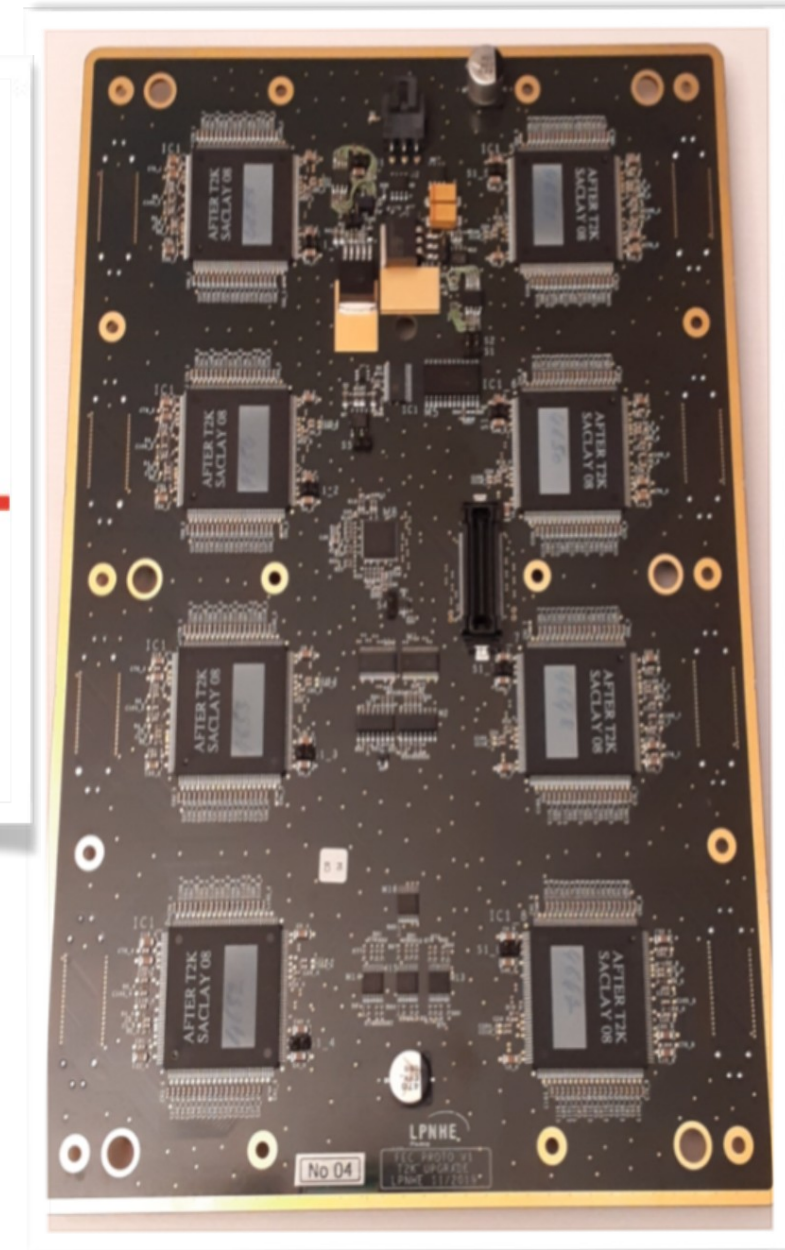
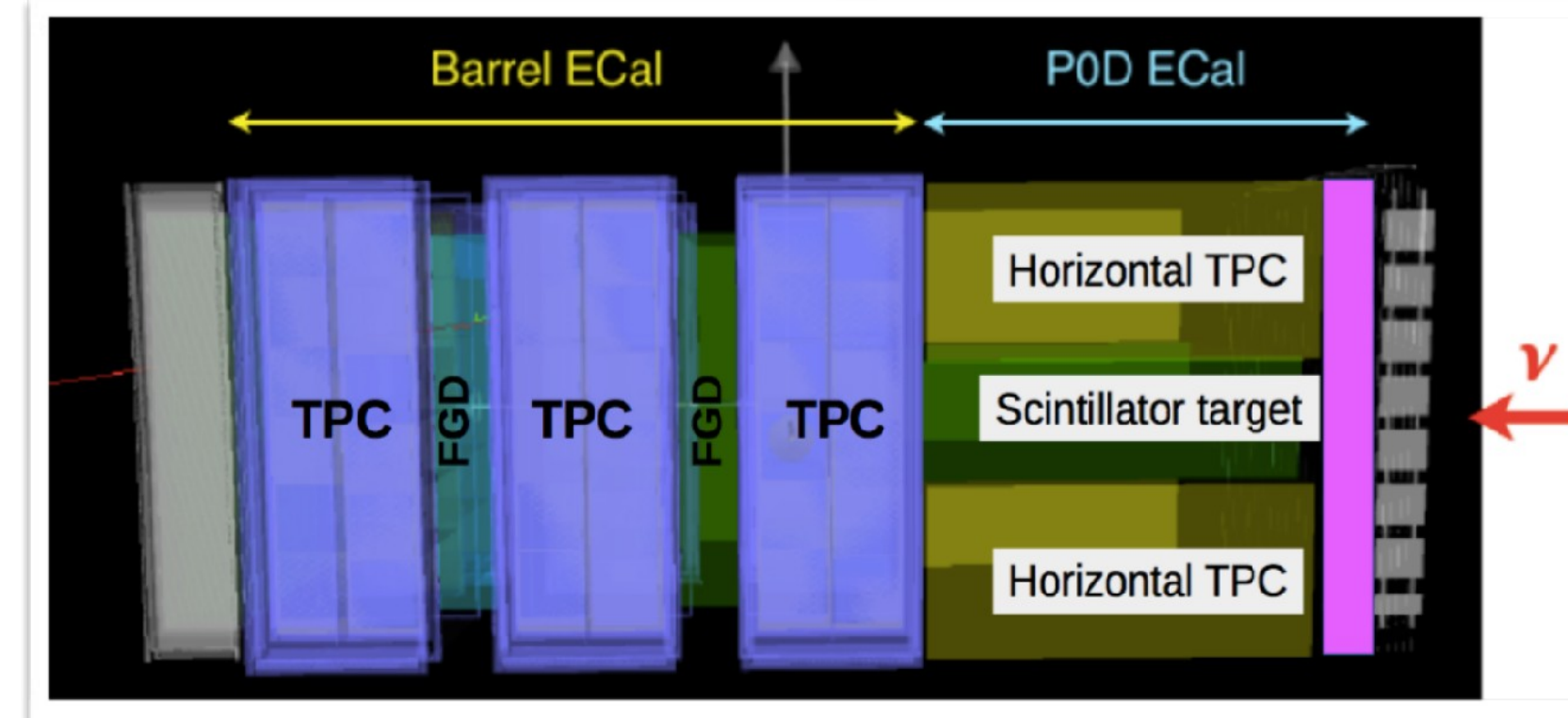
$\nu / \bar{\nu}$ flux uncertainty < 5% thanks to NA61

Magnetized near detector @280 m

Used for T2K Oscillation Analysis for >10 years

Being upgraded now for T2K-II

**Systematics uncertainties under control
from Day-1 of HK**



Tokai to HK: what will be new

Accelerator upgrade

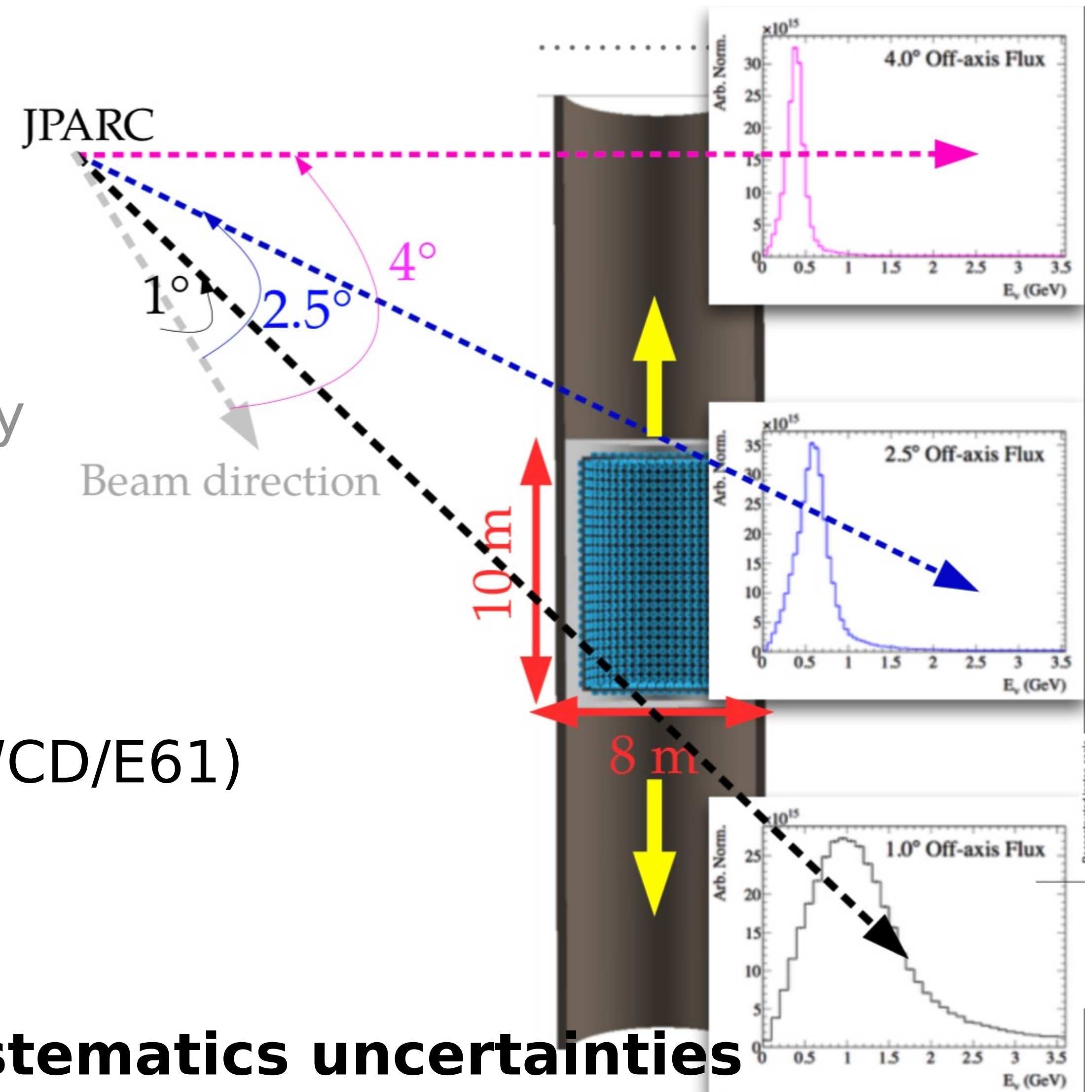
- Power increase (500kW 1.3 MW)
- x2.7 more stats per s (wrt T2K-I)
- $\nu / \bar{\nu}$ flux uncertainty < 5% thanks to NA61
- Magnetized near detector @280 m
- Used for T2K Oscillation Analysis for >10 y
- Being upgraded now for T2K-II
- Systematics uncertainties under control from day 1 of HK

Intermediate Water Cherenkov Detector (IWCD/E61)

- Measure ν interactions on Water
- High stats. sample of ν_e interactions

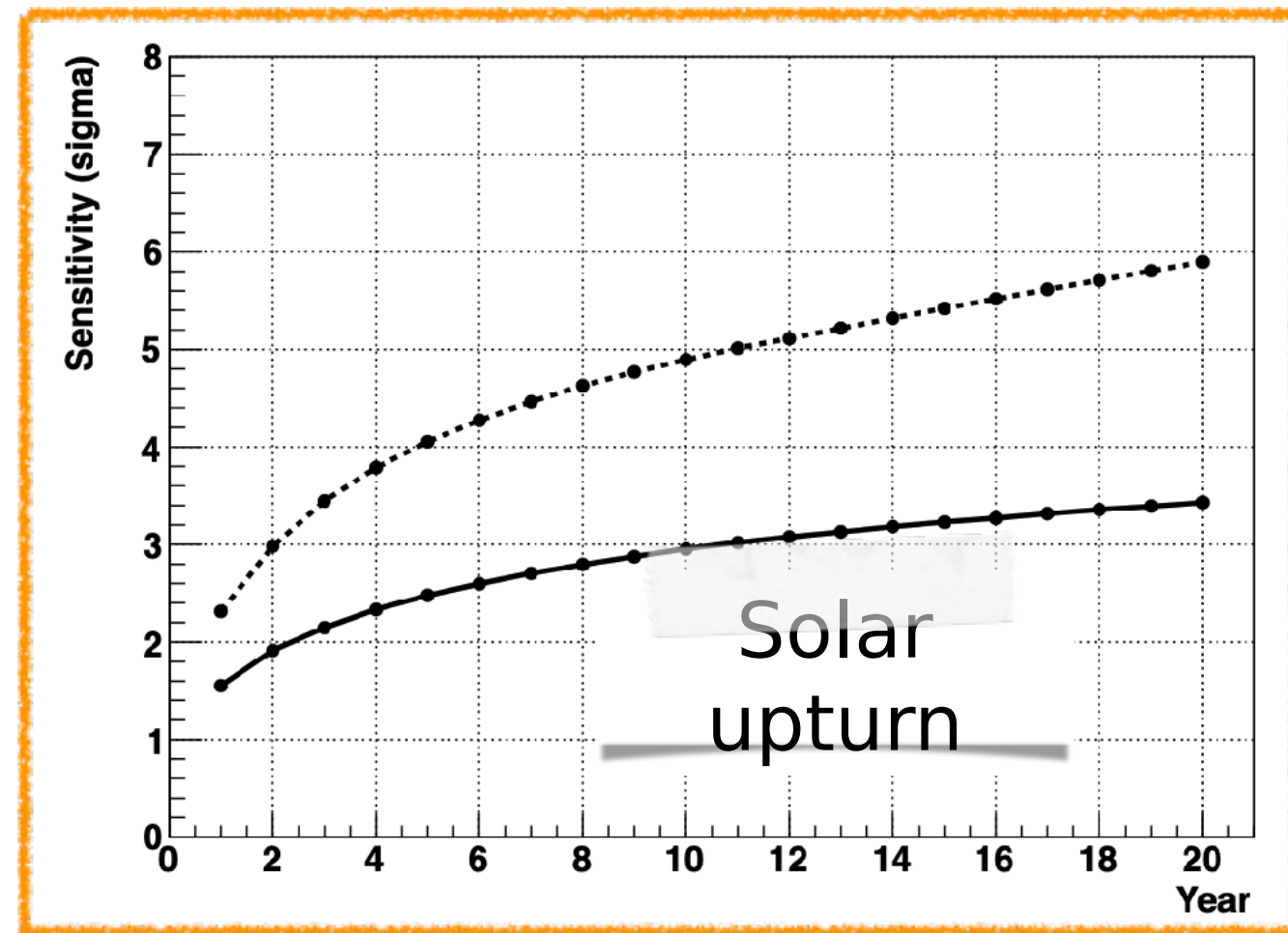
NEW

Needed to reach final HK goal for systematics uncertainties

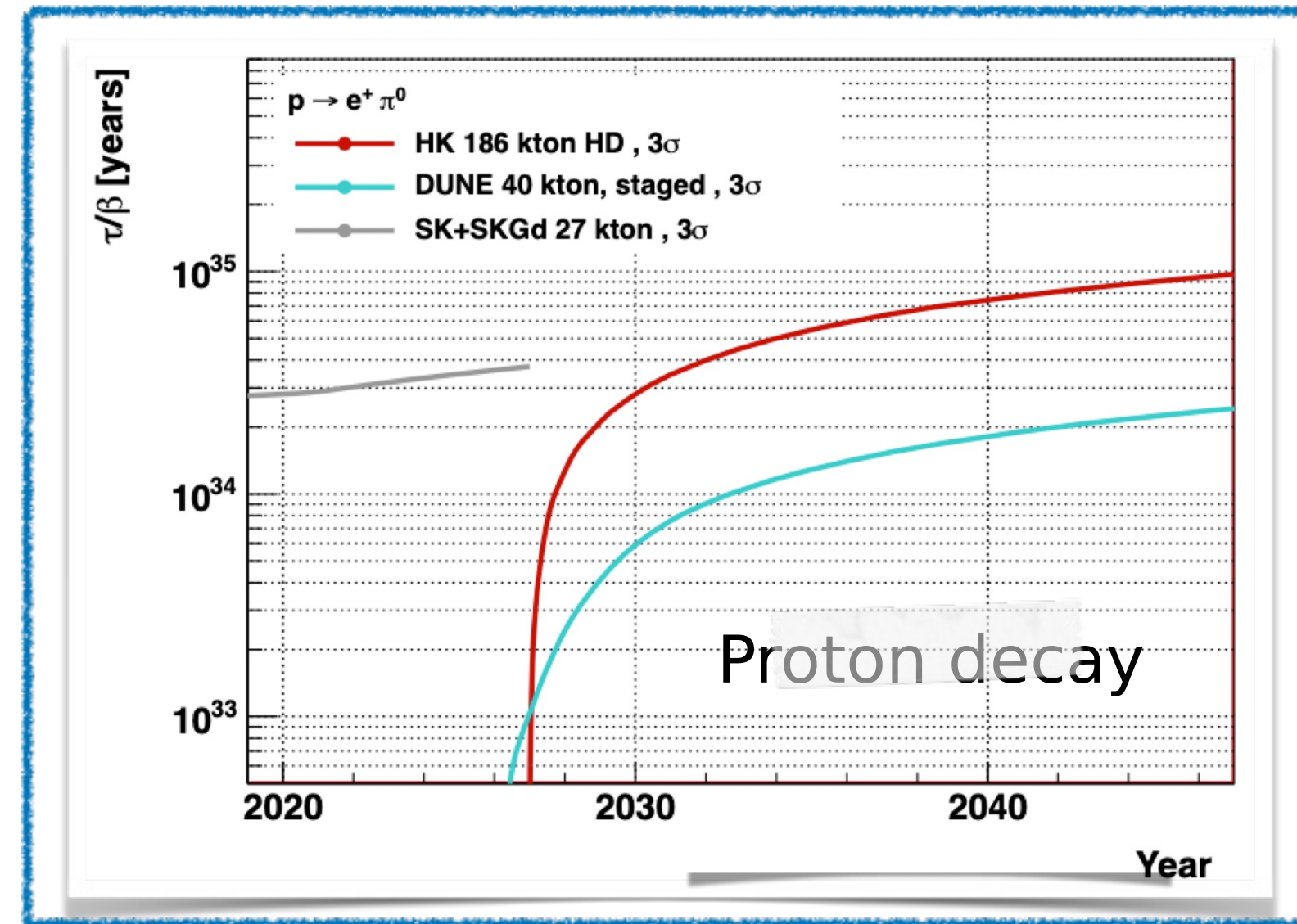


Hyper-Kamiokande Physics

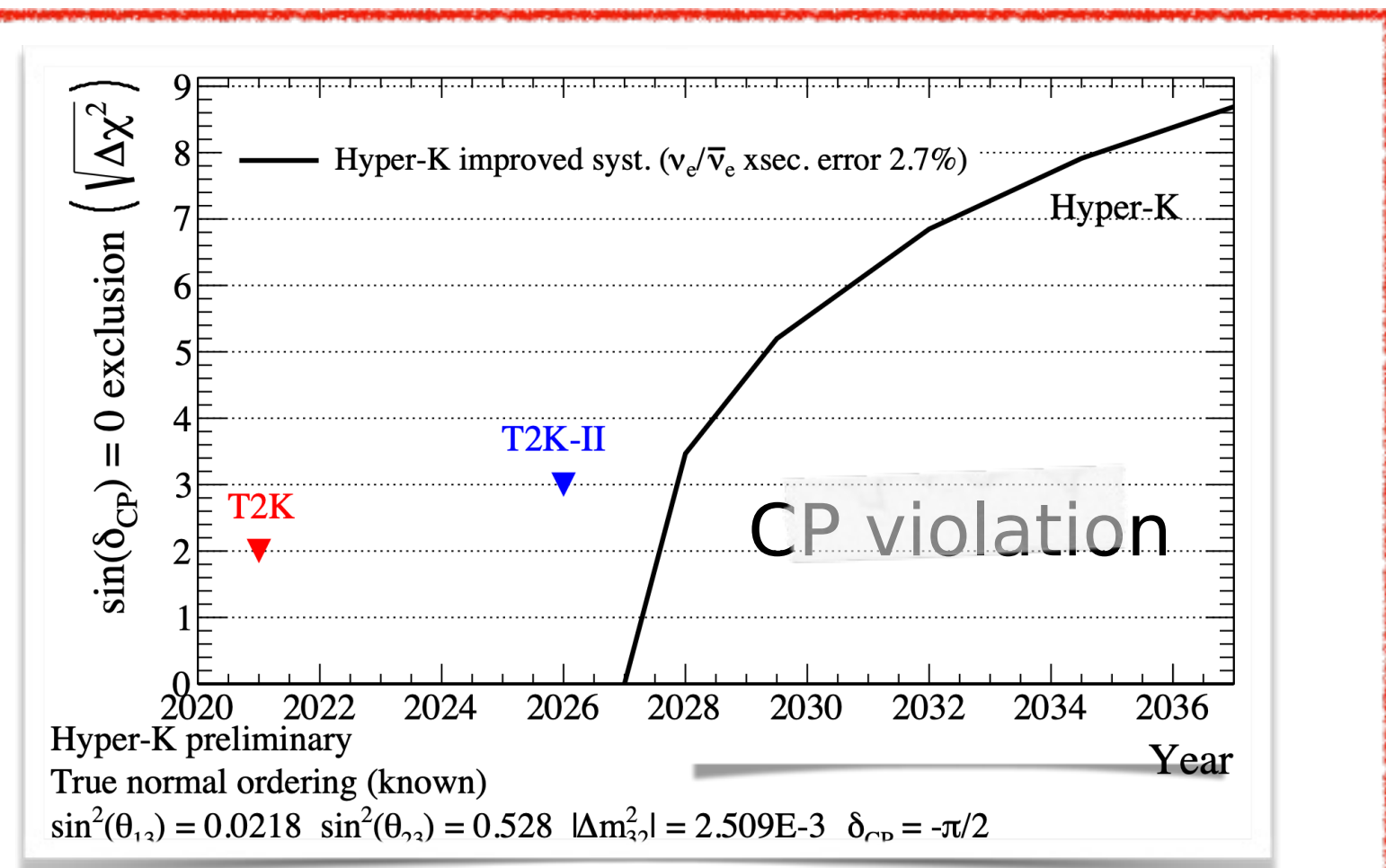
Solar physics



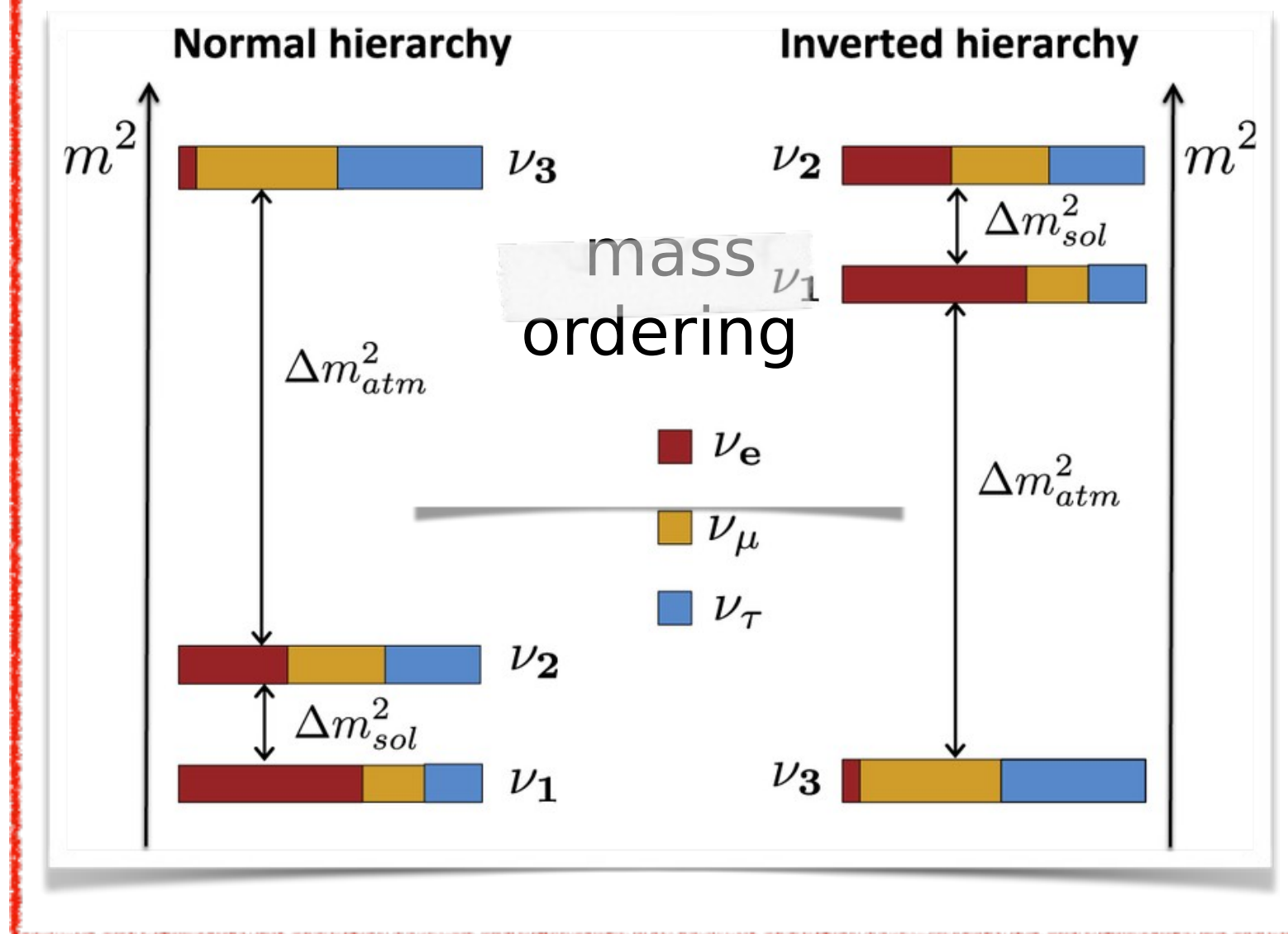
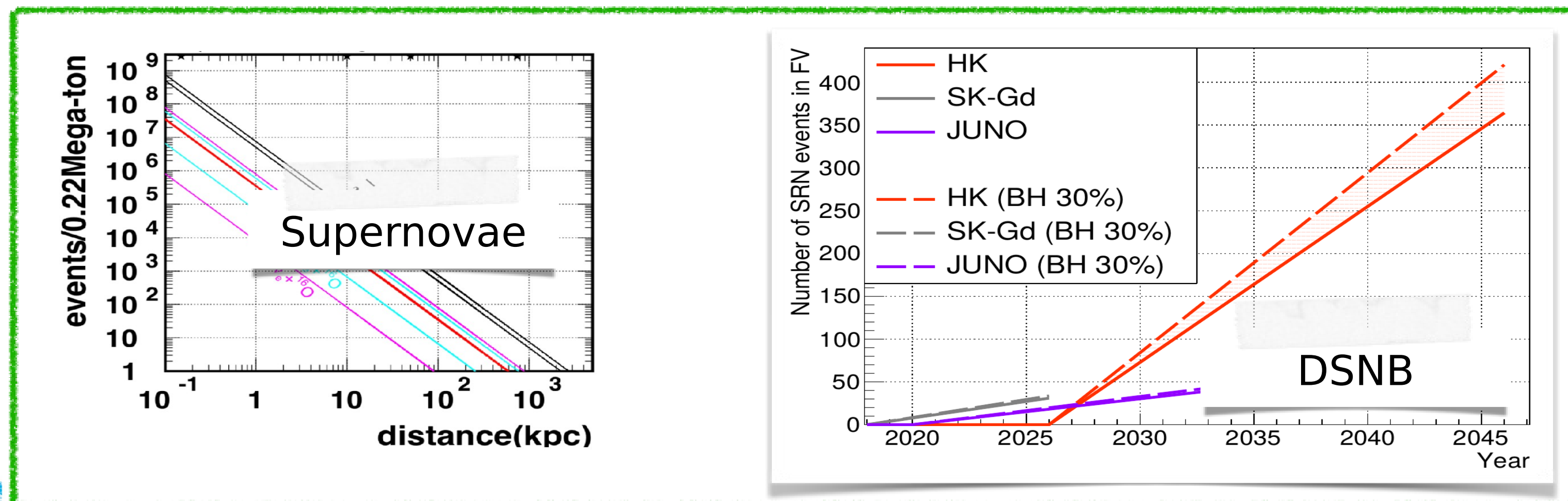
Rare events



Neutrino oscillations

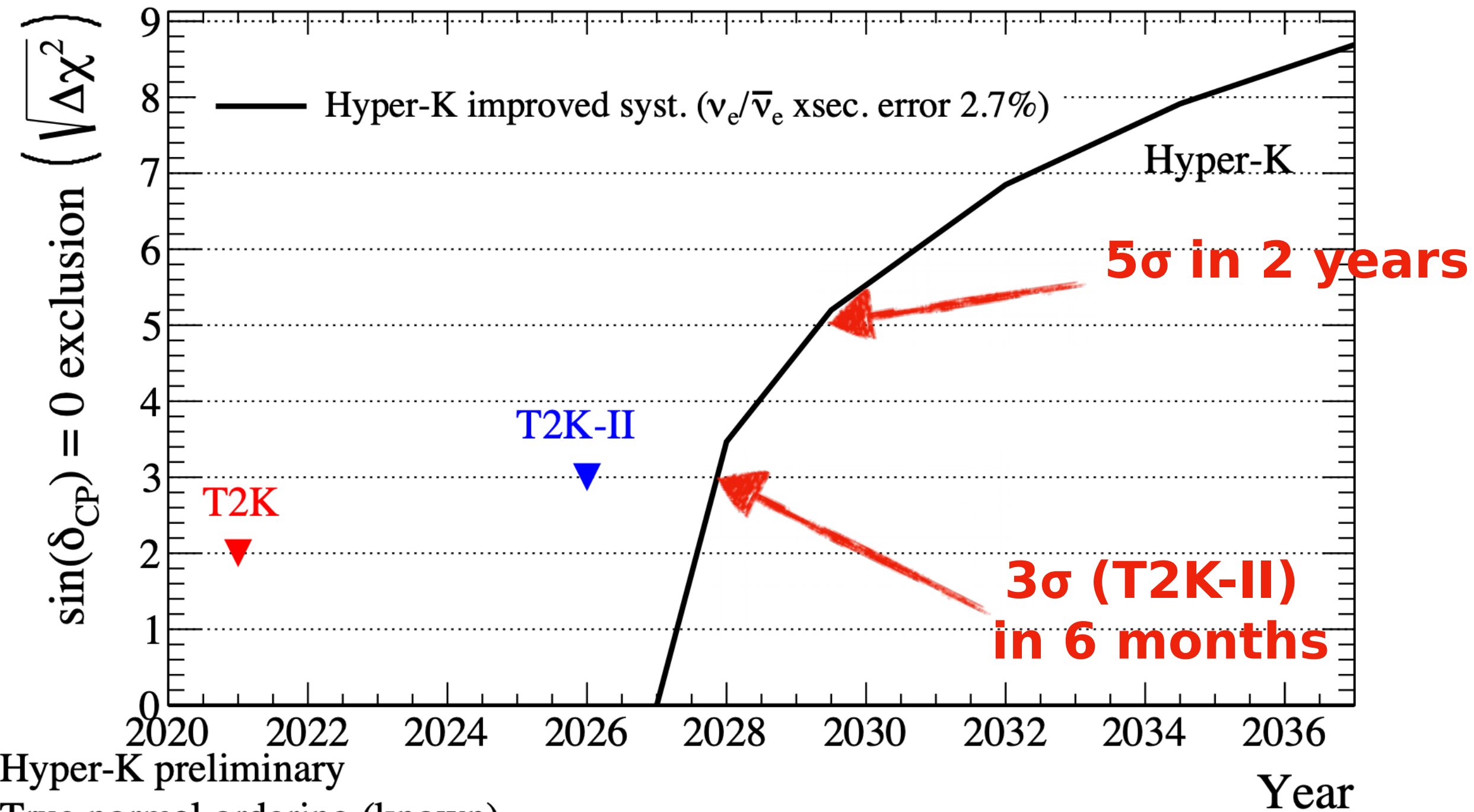


Supernovae modeling and Early Universe



Fast CP-violation discovery

Known mass ordering (KM3NET)



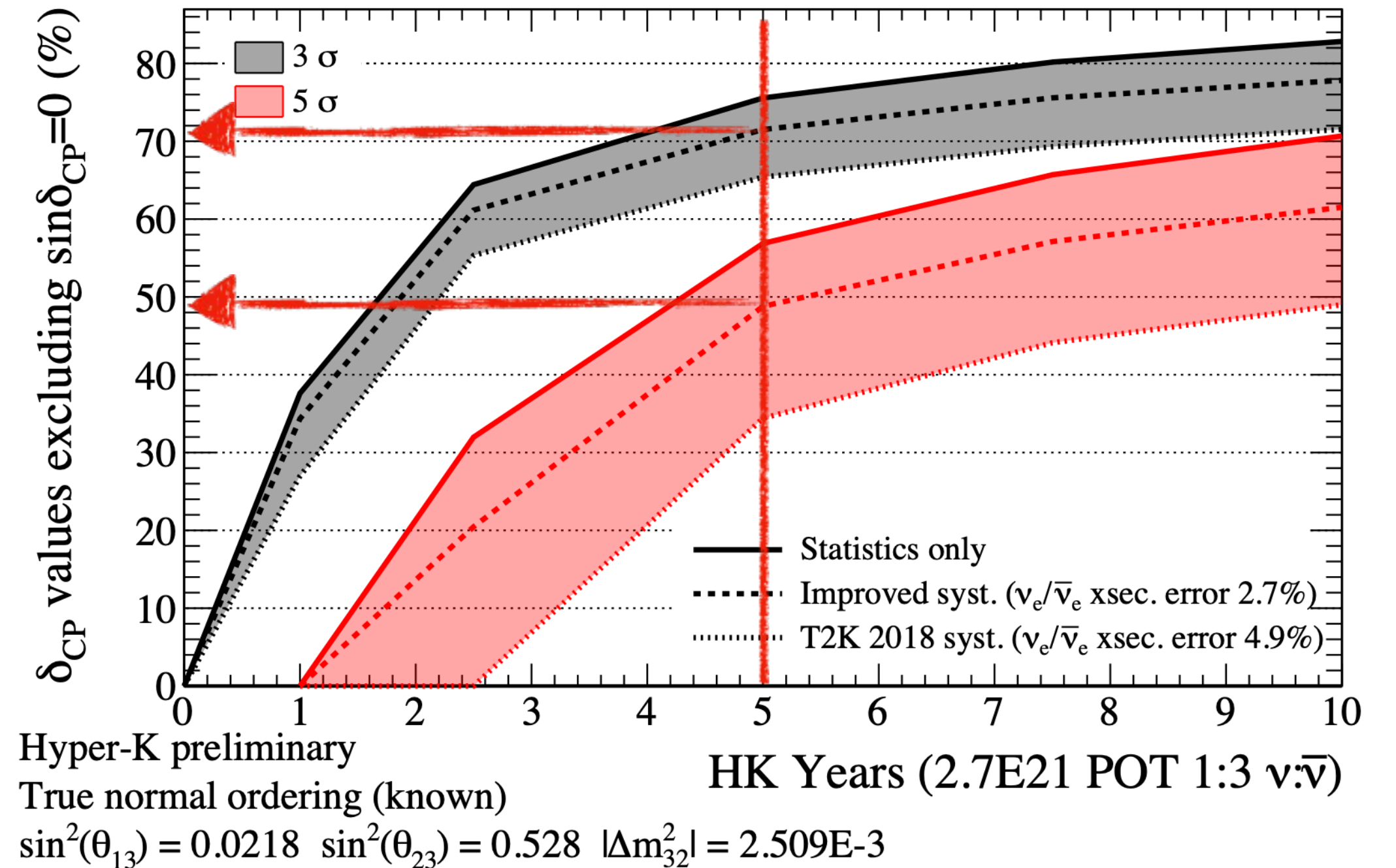
Hyper-K preliminary

True normal ordering (known)

$$\sin^2(\theta_{13}) = 0.0218 \quad \sin^2(\theta_{23}) = 0.528 \quad |\Delta m_{32}^2| = 2.509\text{E-}3 \quad \delta_{CP} = -\pi/2$$

If $\delta_{CP} = -\pi/2$, CP violation discovered before any other LBL- ν experiment

Fastest experiment to survey possible δ_{CP} values

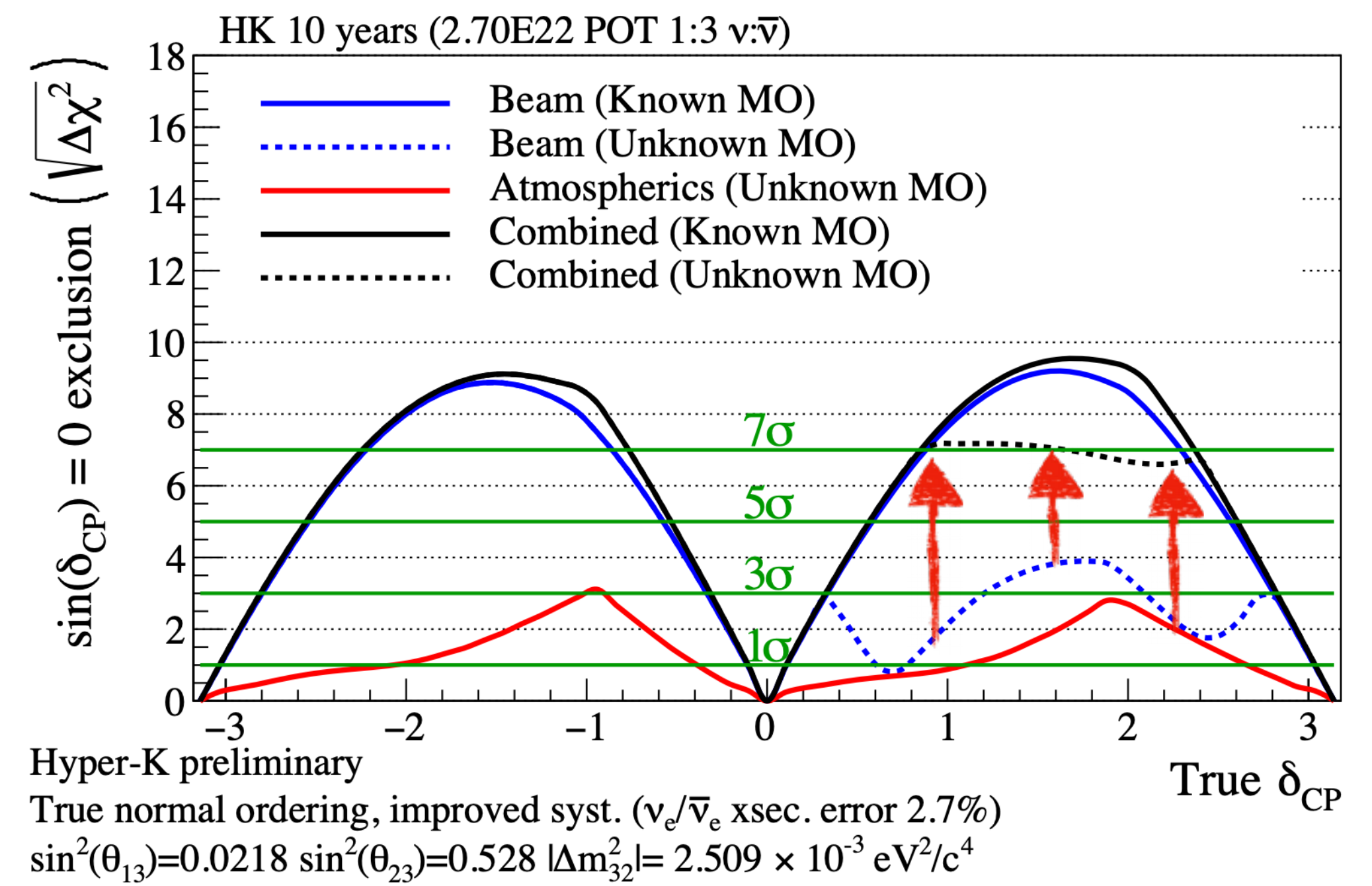
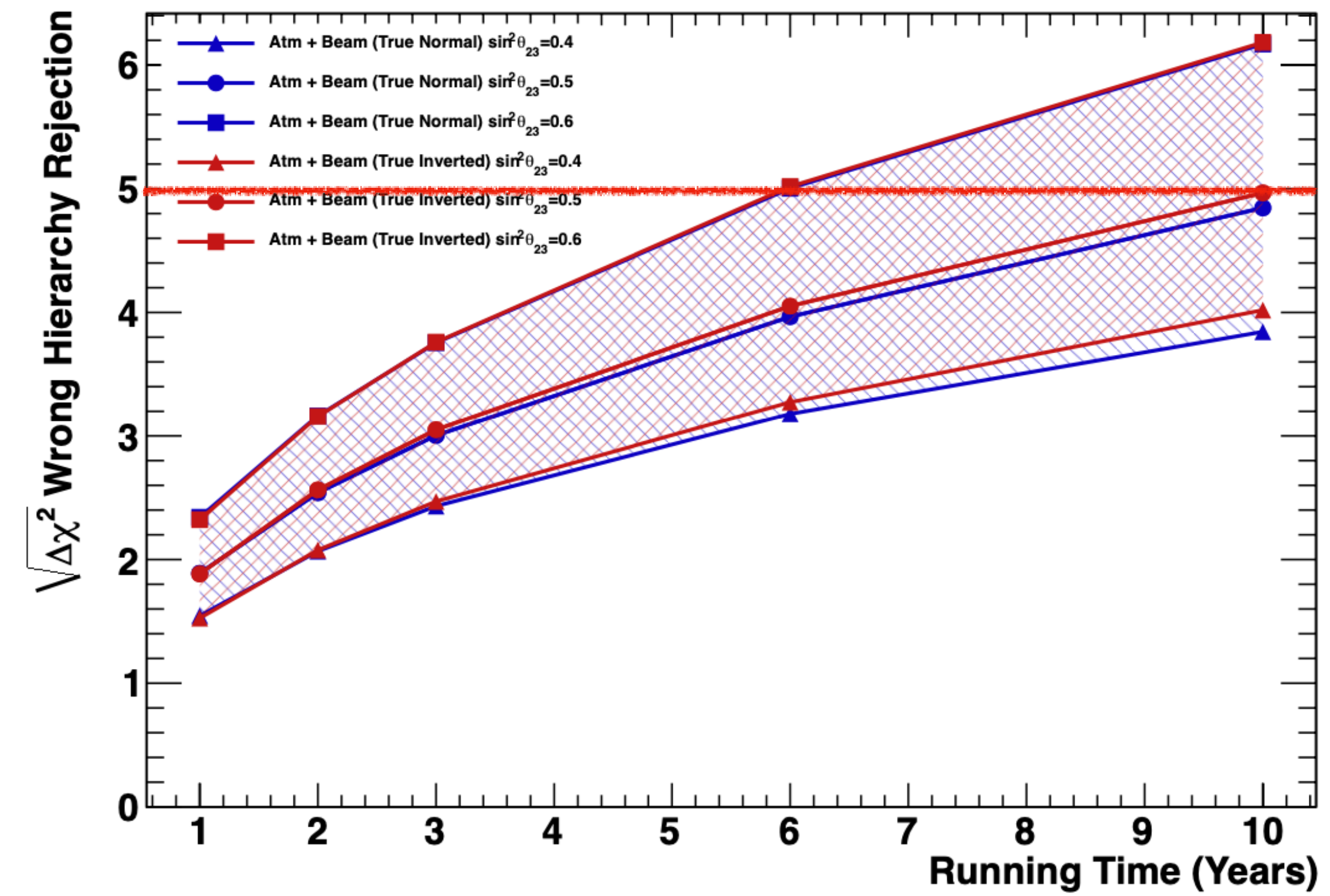


| | $\delta_{CP} = -\pi/2$ | | All δ_{CP} | |
|------------------|------------------------|------------|-------------------|-----------------|
| | 3 σ | 5 σ | 50 % 5 σ | 70 % 3 σ |
| Hyper Kamiokande | 0.5 y | 2 y | 5 y | 5 y |
| DUNE (staged*) | 4 y | 8 y | 10 y | 13 y |

* 2 modules@1.2 MW y1; 3 modules y2; DUNE CDR [arXiv:2002.03005](https://arxiv.org/abs/2002.03005)
4 modules y4; @2.4 MW y7

IUPAP Neutrino panel [report](#)

Mass ordering sensitivity with atmospheric neutrinos

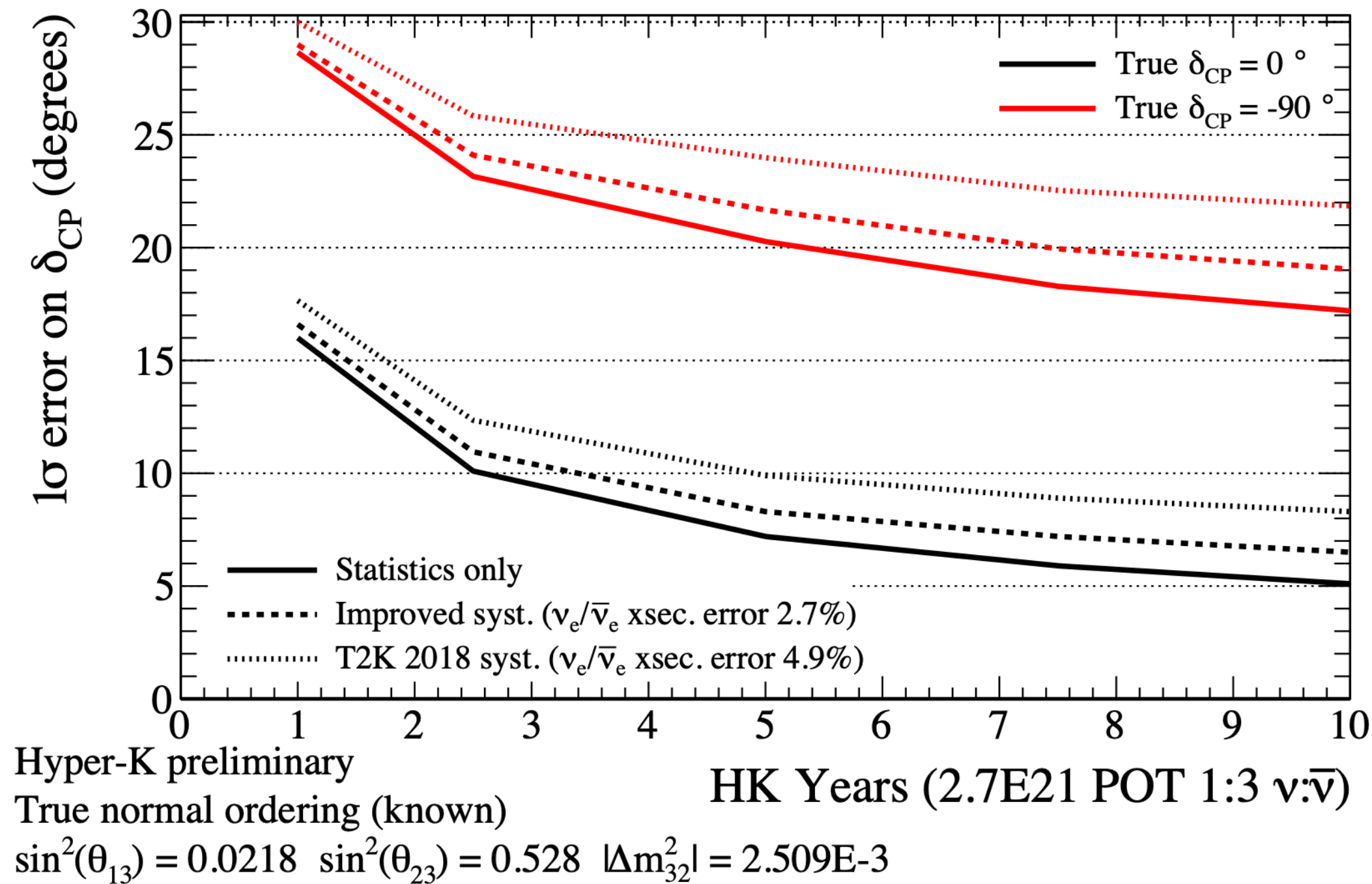


If not discovered by T2K/SK, NOvA, ORCA or JUNO before 2027, HK can determine MO after 6-10 years via atmospheric ν

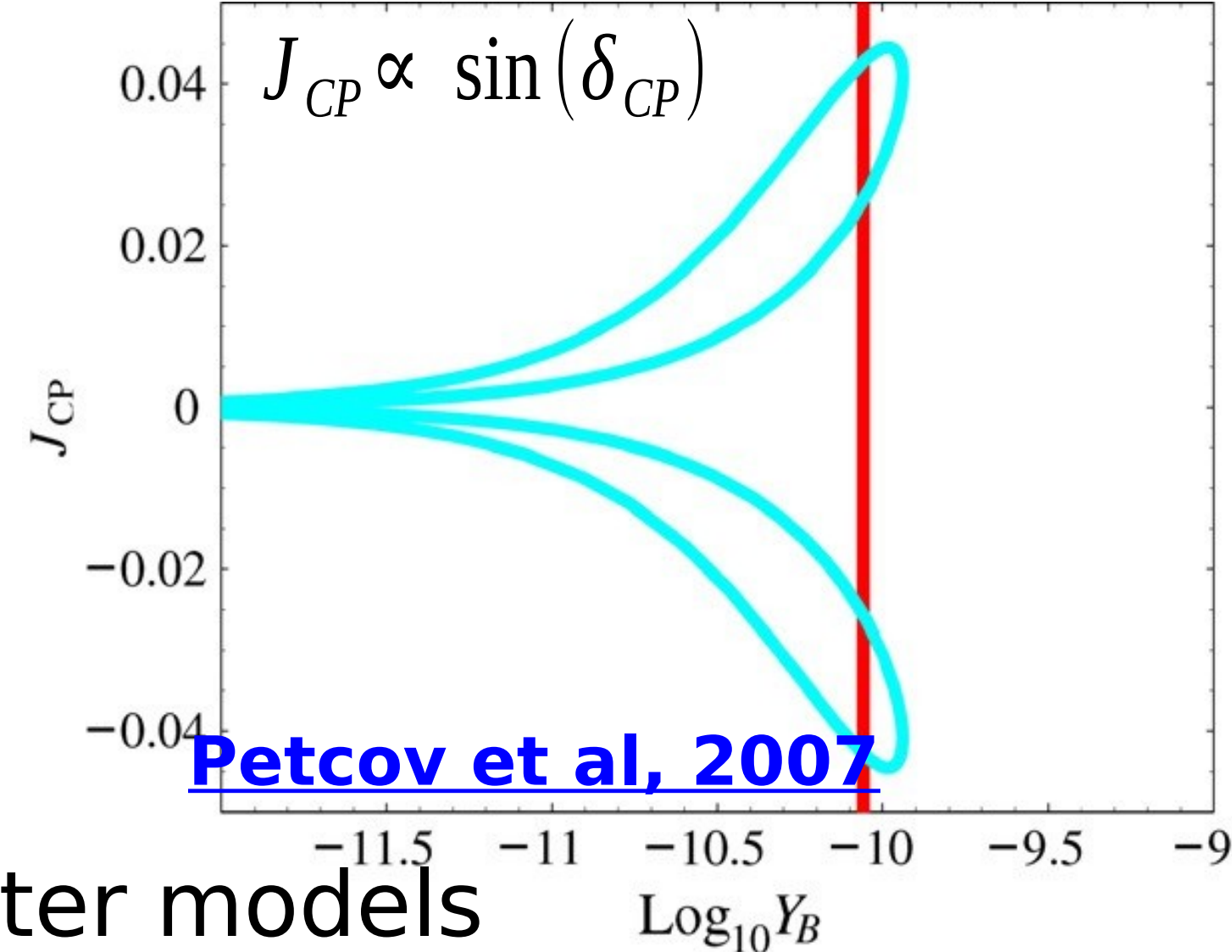
| | $\sin^2 \theta_{23}$ | Atmospheric neutrino | Atm + Beam |
|---------------|----------------------|----------------------|----------------|
| Mass ordering | 0.40 | 2.2 σ | → 3.8 σ |
| | 0.60 | 4.9 σ | → 6.2 σ |

Sensitivity to CPV is little affected if we add atmospheric ν
→ MO prior knowledge not really required to explore δ_{CP}

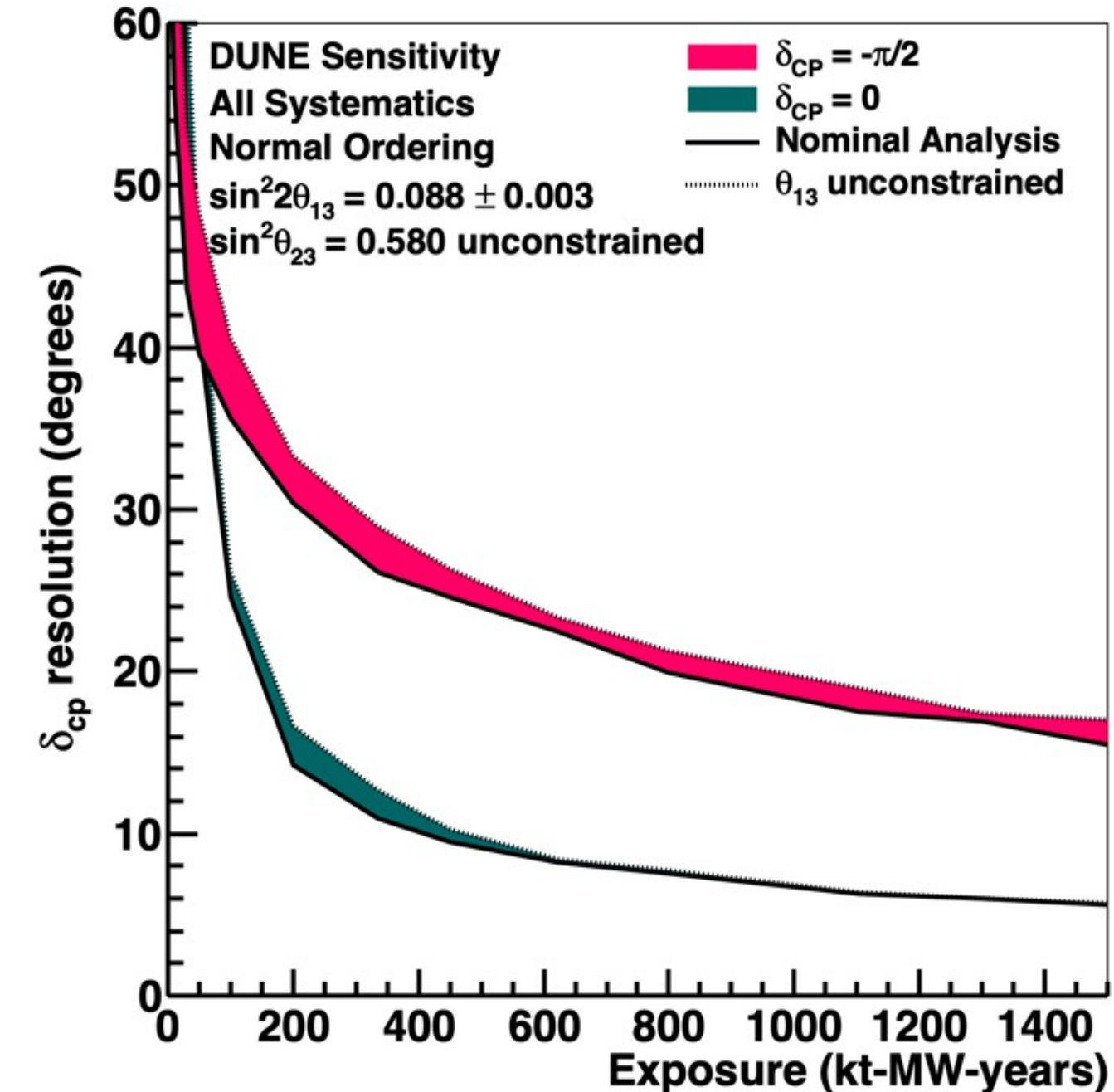
δ_{CP} measurement resolution



| | $\delta_{CP} = -\pi/2$ | | $\delta_{CP} = 0$ | |
|------------------|------------------------|------|-------------------|-----|
| | 30° | 20° | 15° | 10° |
| Hyper Kamiokande | 1 y | 7 y | 1 y | 3 y |
| DUNE | 5 y | 12 y | 5 y | 8 y |



Petcov et al, 2007



DUNE CDR [arXiv:2002.03005](https://arxiv.org/abs/2002.03005)

Precision = sensitivity to matter-antimatter models

→ **HK will quickly reach precision on δ_{CP} of 30°(15°) for $\delta_{CP} = -\pi/2$ (0)**

For the ultimate precision on δ_{CP} it will be important to further

reduce systematics uncertainties w.r.t. T2K (ND280 Upgrade + IWCD)

Proton decay

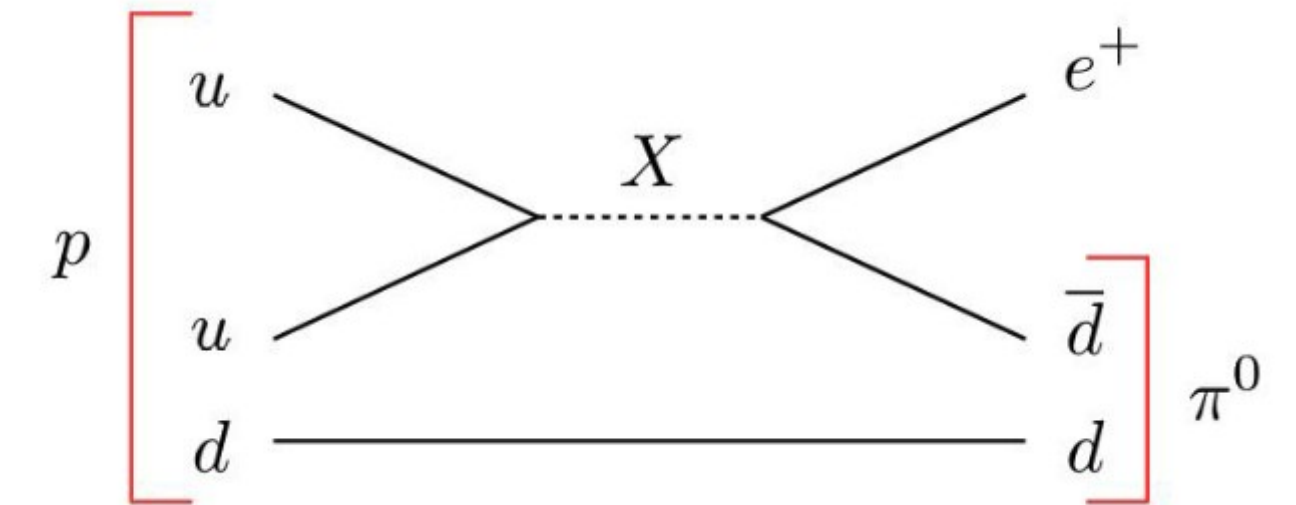
Motivated by Grand-Unification Theories

HK will have the best limit on $p \rightarrow e^+ \pi^0$ for bound protons

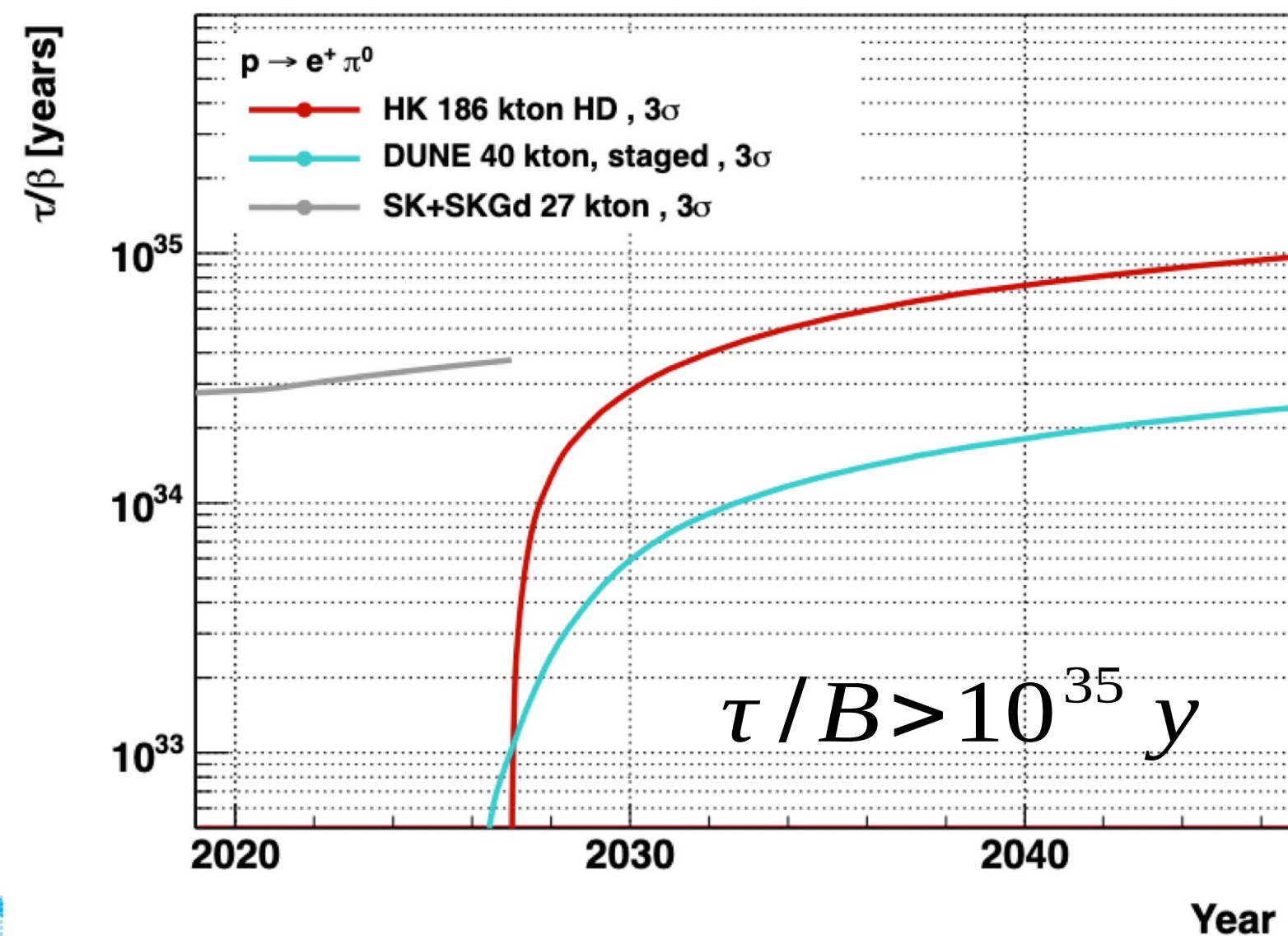
→ about 1 order of magnitude better than current limits

Thanks to its huge mass, HK will also have leading sensitivity to channels with invisible particles ($p \rightarrow \bar{\nu} K^+$)

HK is sensitive to free proton decay



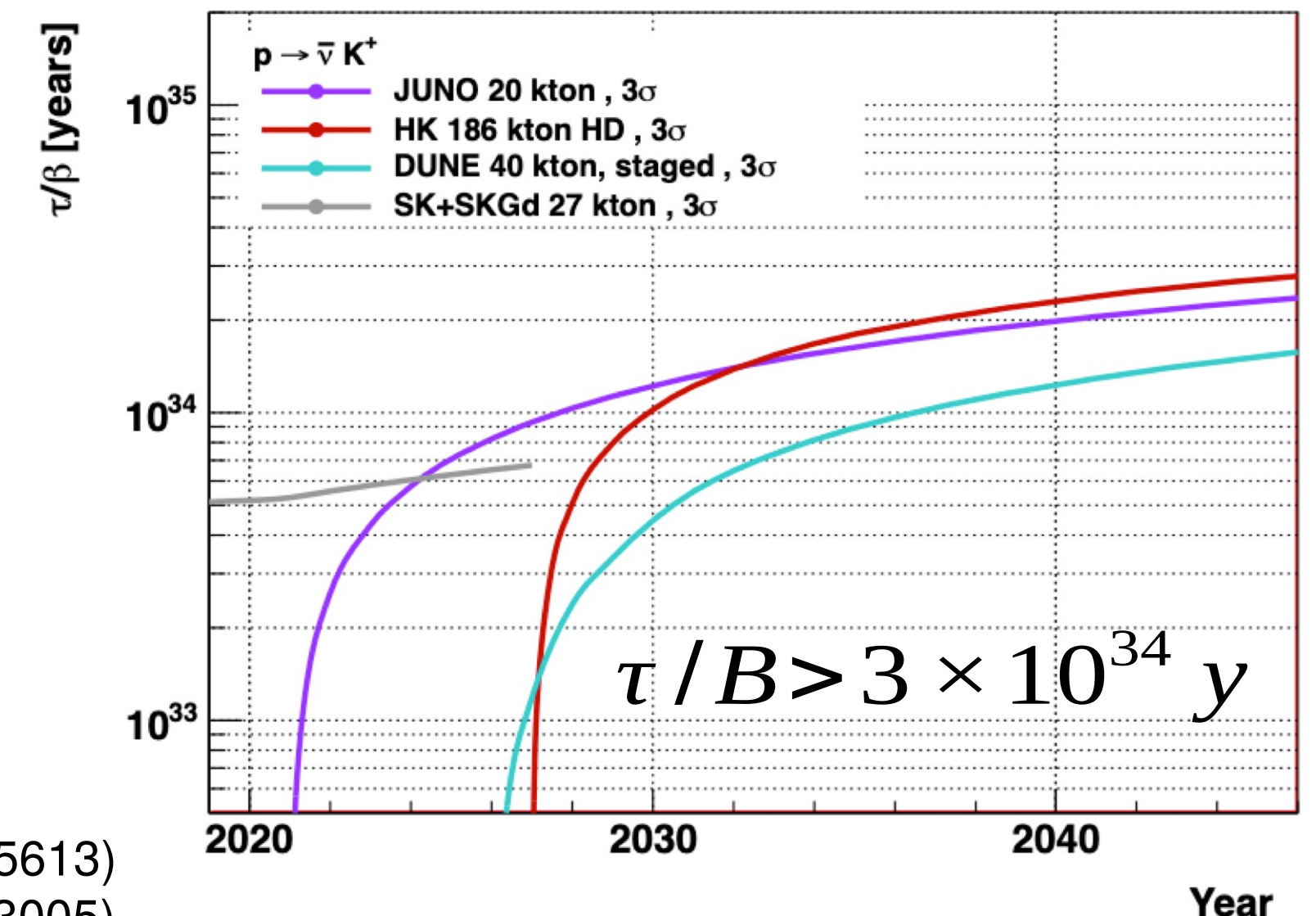
Phys. Lett. B 233 (1-2) 178-182



JUNO: J. Phys. G 43 (2016) 030401 (arXiv:1507.05613)

DUNE: FERMILAB-PUB-20-025-ND (arXiv:2002.03005)

Hyper-Kamiokande — CS IN2P3 October 2022



Astrophysical neutrinos

Supernova neutrinos [arXiv:2101.05269](https://arxiv.org/abs/2101.05269)

Increase by ~ 10 in stats sensitivity wrt. SK

SN1987A type ~ 2500 events

Galactic center: $\sim 50000+$ events

Direction (1° @ 10kpc) \rightarrow triangulation

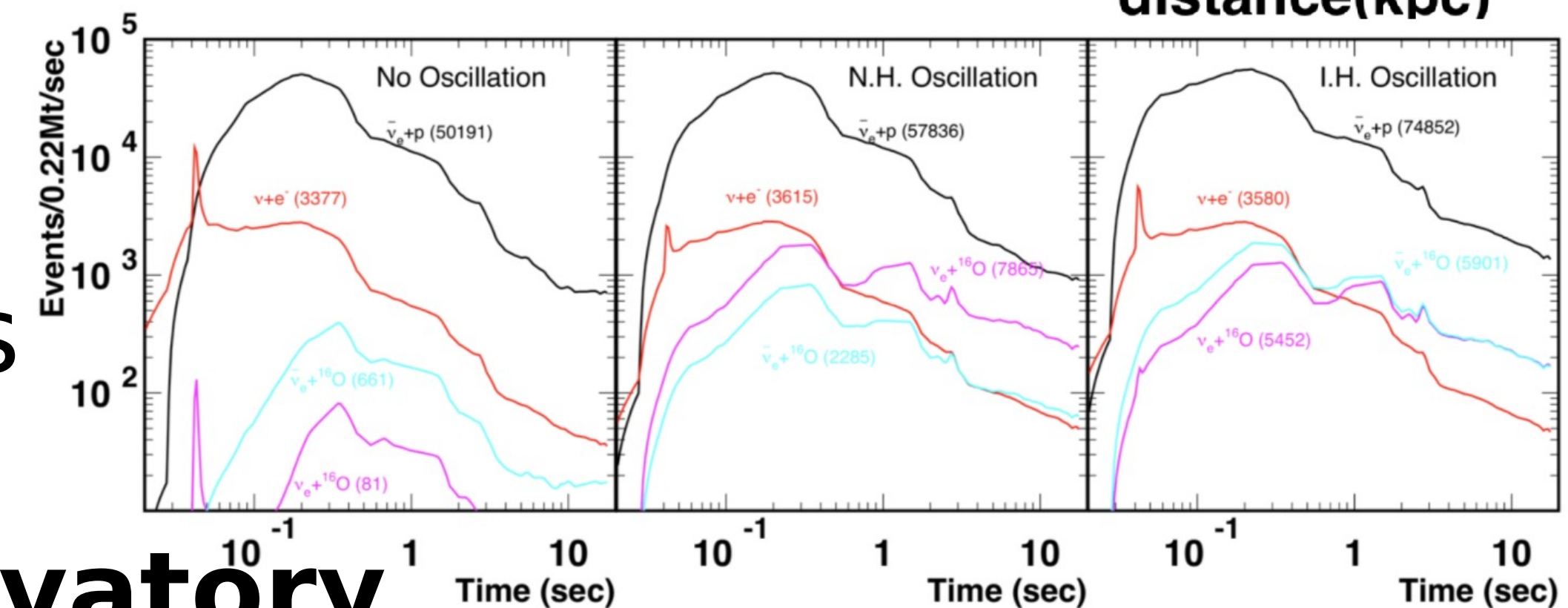
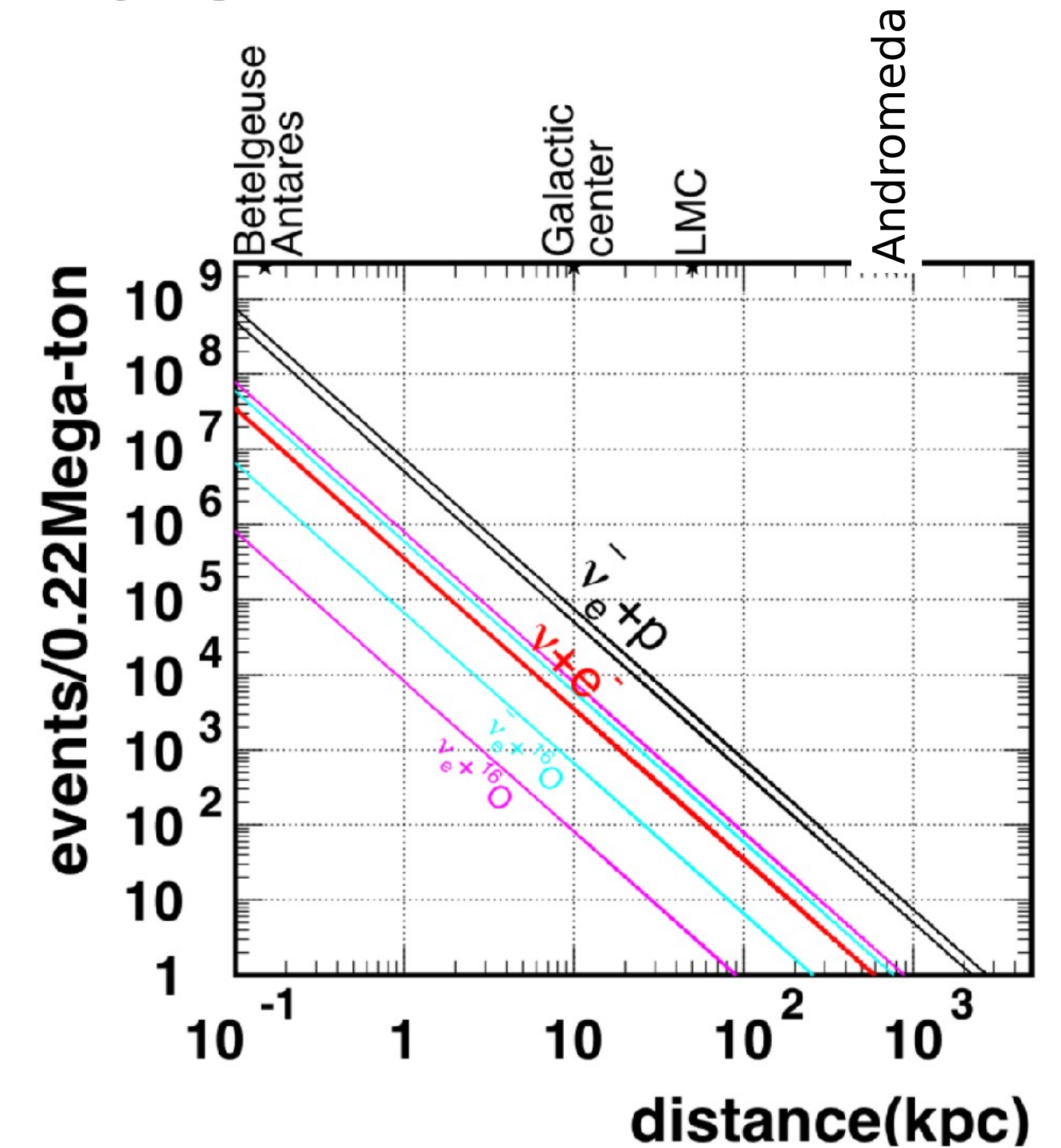
Time profile: collapse models

SN ν detected every 3 years in HK!

Gravitational waves sources

Nearby (10Mpc) neutron star mergers

\rightarrow Unique multi-messengers observatory



DSNB discovery by HK

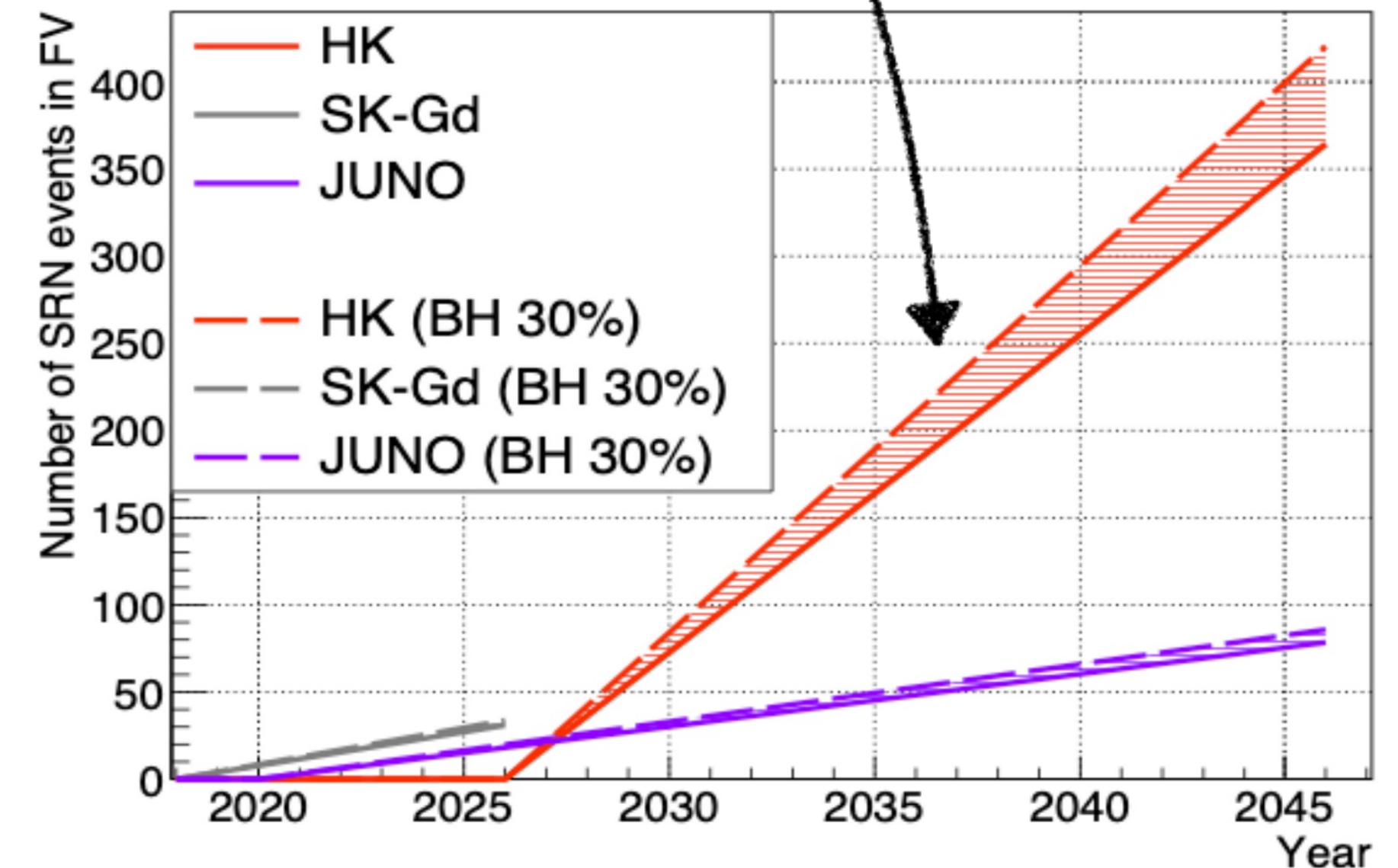
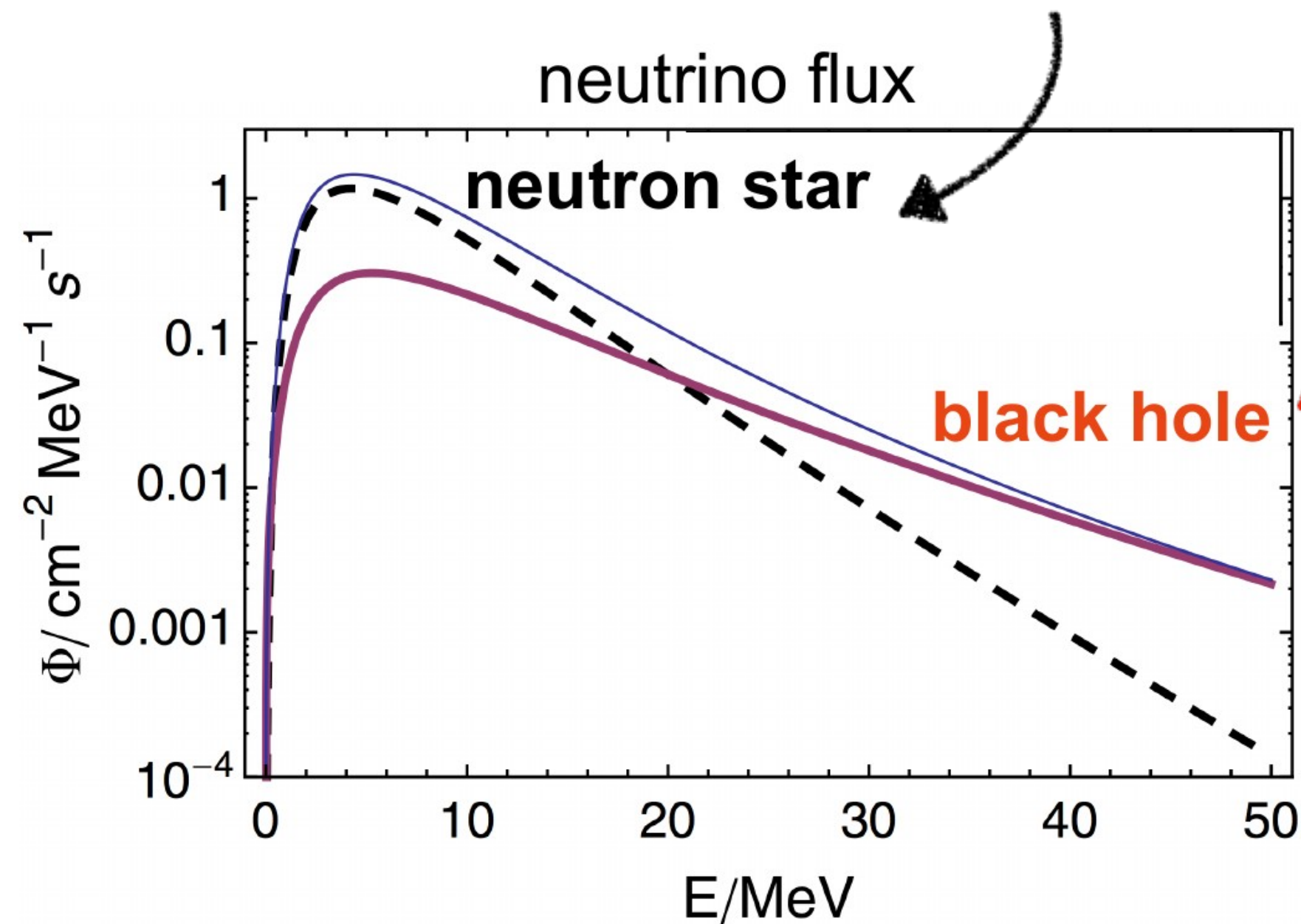
SN-relic neutrinos (SNR ν) offer new constraints on cosmic star history

→ Could be first detected by SK-Gd

→ The spectrum will be determined by HK

Impact of redshift: low energy \leftrightarrow probe older stars

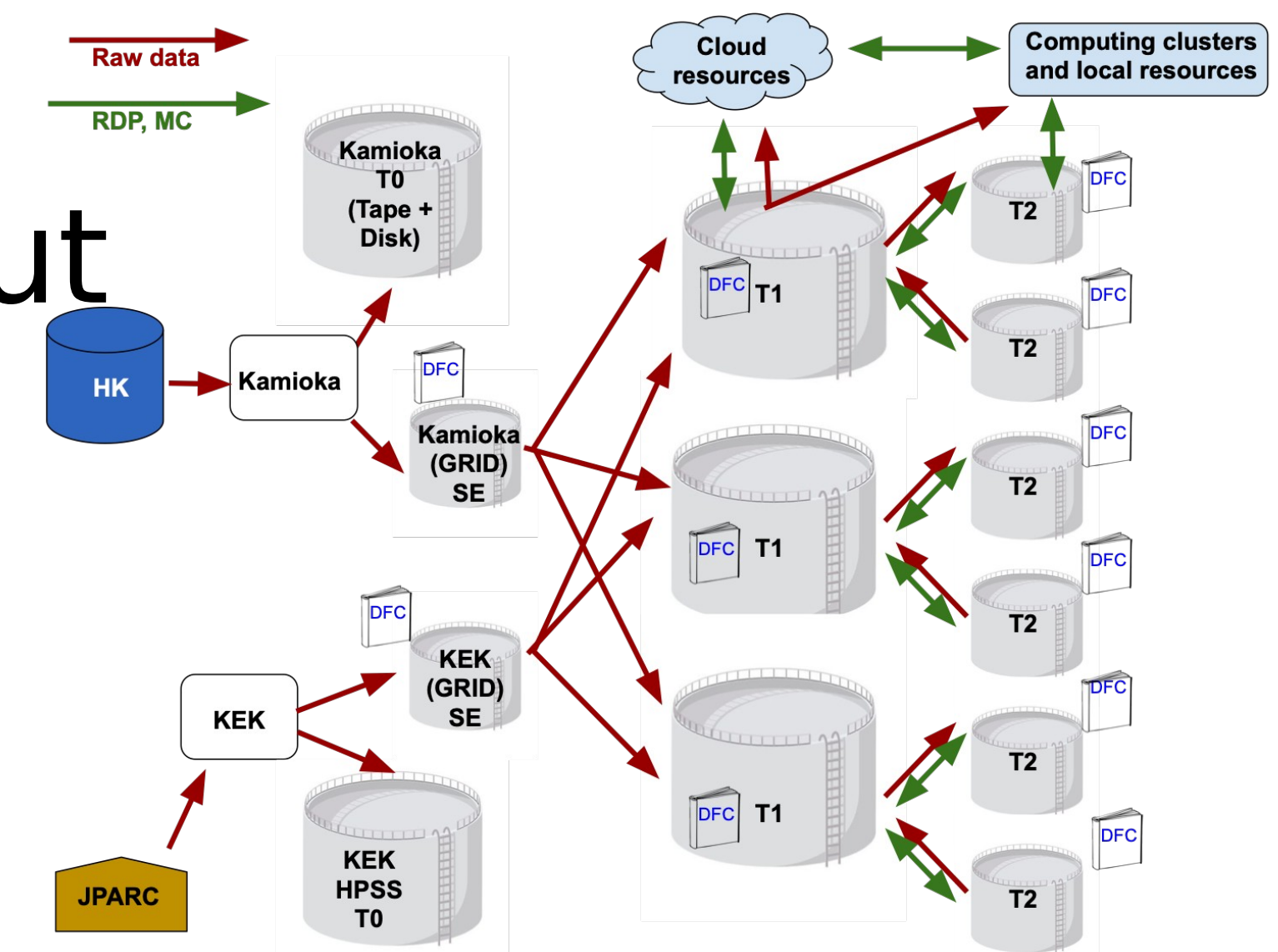
Sensitivity to neutron star vs **black hole** formation



IN2P3-CEA technical contributions to HK



- **ND280 Upgrade and maintenance**
 - Super FGD electronics
 - High-Angle TPCs electronics & readout
- **Construction of HK far detector**
 - Front-end electronics
- **International computing effort**
 - CC-IN2P3 as T1 for HK



Backup slides

SWOT

Strengths:

- Well-known exceptional Water Cherenkov technology
- Use of existing neutrino beam and near detector complex built for T2K, thus saving large amount of money for the long-baseline program and reducing systematics uncertainties from the first day of the experiment.
- Construction budget for Hyper-Kamiokande have been allocated by Japanese government in 2019 with a budget profile that will allow to start the experiment in 2027.

Opportunities:

- Fast measurement of CP violation, before any other experiments.
- Huge target mass, making HK the most sensitive observatory for rare events in the MeV–GeV region.
- IN2P3 groups can build on their long standing expertise in the T2K experiment to propose strong contributions in Hyper-Kamiokande. The possibility of using chips developed by OMEGA for the Hyper-Kamiokande far detector is particularly attractive from this point of view.

Weaknesses:

- Small groups at LLR and LPNHE. Mitigated by the large overlap in terms of physics case, technologies and tools between T2K, SK and HK.
- So far no other IN2P3 groups decided to join the Hyper-Kamiokande experiment.
- Hyper-Kamiokande not an IN2P3 project, undermining our visibility within the collaboration.

Risks:

- Dates for the approval by CS-IN2P3 and fundings of our proposed contributions to the HK electronics before collaboration review in Summer 2022. **Missing the Summer 2022 deadline would compromise the French contributions to the Hyper-Kamiokande far detector.**
- If the solutions not selected or not funded, no other planned French contributions to the far detector construction, putting our participation to Hyper-Kamiokande in jeopardy.

Conclusions and outlook

Hyper-Kamiokande has a vast and rich physics program including:

- ν oscillations (fast CP-violation discovery, δ_{CP} precision...)

- Rare events observatory e.g. proton decay

- Multi-messenger astrophysics (transient and diffuse SN detection)

→ **High and quick discovery potential!**

Construction started and on schedule → Operation will start in 2027

Continuation of T2K and T2K-II

- Upgraded ND280 will be used as the HK near detector

- Unique program of world-leading measurements and discoveries from 2010 to ~2040

Almost done: identified contributions on electronics and computing

Used fundings from external sources (SU, Ecole Polytechnique, ANR)

Technological decisions taken in Summer 2022

→ **approval by CS-IN2P3 & recognition of HK as IN2P3 project are needed for full-scale participation in the HK experiment**

Financial resources

| Item | Cost (M€) | Partially covered with external fundings | Funding approval | Construction period | Requested fundings (M€) |
|----------------------|-----------|--|------------------|---------------------|-------------------------|
| ND280 Upgrade | 6 | T2K Collaboration | 2019 | 2019 – 2022 | 0.6 (obtained) |
| Far detector timing | 1 | ANR - INFN - CEA | 2022 | 2023 – 2026 | TBD |
| Communication cables | 2 | European countries | 2022 | 2023 – 2026 | TBD (~0.5) |
| Chip and Front-end | 1.5-3 | CEA | 2022 | 2023 – 2026 | TBD |
| Computing (CC-IN2P3) | 1.9-3.8 | - | 2021 | 2021 – 2037 | 1.9-3.8 |
| Total | 12.4-15.8 | - | - | - | ~5-7 |

Table 1: Summary of the costs of the various items for Hyper-Kamiokande. Note that it includes only the production costs (the R&D costs has been already funded). See section 5 for more details about the over investments.

External and internal investments:

- 70k€ for R&D towards HK (Sorbonne Université)
- 400k€ (X) + 90k€ (IN2P3) for R&D on HKROC
- 300k€ (ANR) + 50 k€ (IN2P3) for R&D on time generation and clock distribution

Significant efforts to acquire external fundings before IN2P3 investments kicks in

Time Distribution costs estimations

| Second stage distribution | Total number of PMTs | PMT per FE | Total number of FE (Time distribution nodes) | Number of nodes per distributor | Number of distributors | Distributor cost per unit | Total distributor's cost |
|----------------------------------|----------------------|------------|--|---------------------------------|------------------------|---------------------------|--------------------------|
| Outer detector | 13 300 | 72 | 185 | 16 | 12 | 3 000 | 36 000 |
| Inner detector | 20 000 | 24 | 833 | 16 | 53 | 3 000 | 159 000 |
| MultiPMT | 2 000 | 12 | 167 | 16 | 11 | 3 000 | 33 000 |

| | |
|------------------------------------|----------------|
| Total distributor's cost | 228 000 |
| Atomic Clock cost | 54 000 |
| GNSS recevier | 13 000 |
| Antenna + cable | 4 000 |
| Clock multiplication board | 2 000 |
| Total clock ref cost | 73 000 |
| Redundant system | 73 000 |
| Grand total clock reference | 146 000 |
| Slow Control switch | 2 000 |
| Total cost | 376 000 |
| Contingency | 20% |
| Gran total | 451 200 |

Time generation estimated cost: 145 k€ (LPNHE)

Time distribution (first and second stages) estimated cost: 250 k€ (LPNHE +CEA)

Time distribution endpoint cost: 80 k€ (INFN)

Total = ~480 Keuros