



# CASA improvements for VLBI Des Small, JIVE

# JIVE and the EVN


**JIVE**  
 Joint Institute for VLBI  
 ERIC

  
**EUROPEAN  
 NETWORK**

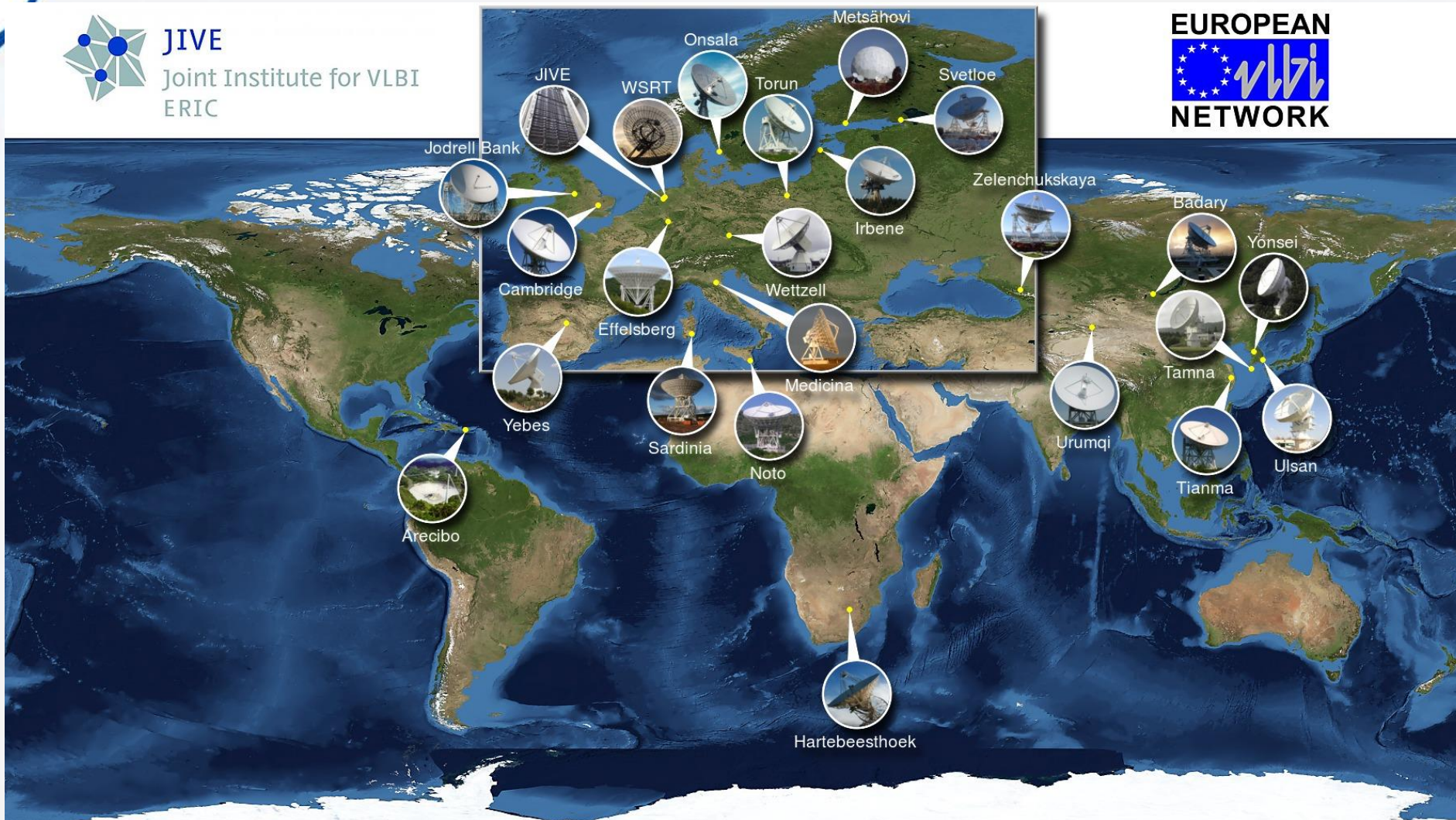


Image by Paul Boven (boven@jive.eu). Satellite image: Blue Marble Next Generation, courtesy of Nasa Visible Earth (visibleearth.nasa.gov).

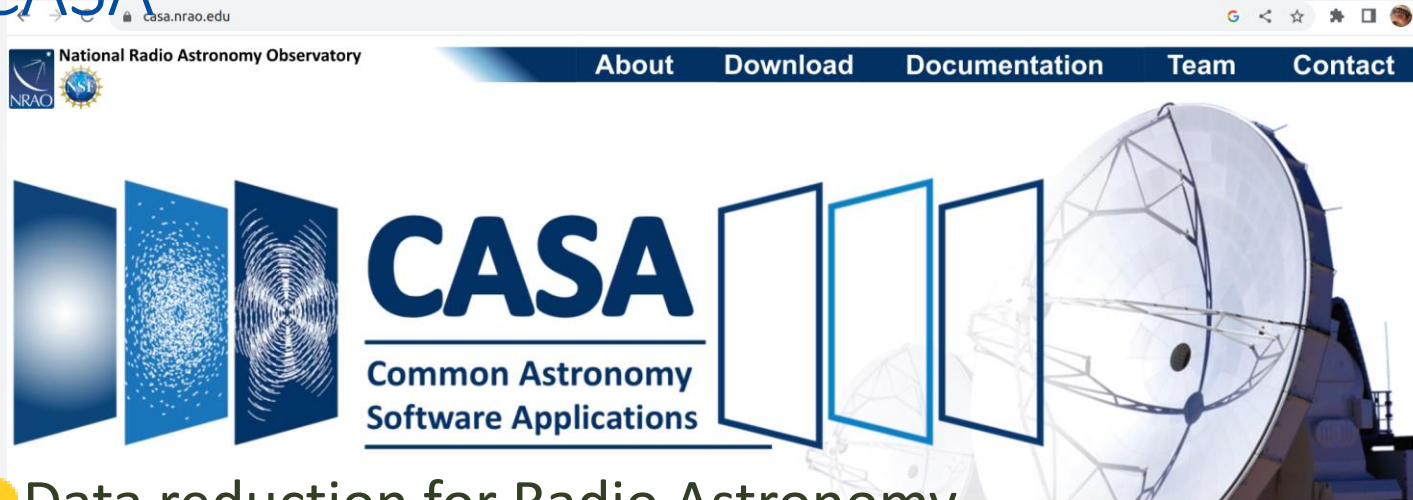
# JIVE and the EVN



- JIVE is a European Research Infrastructure
- Works for EVN
- Correlates most EVN experiments
- Maintains software correlator (SFXC)
- Supports users in data reduction
  - Originally AIPS
  - Increasingly CASA





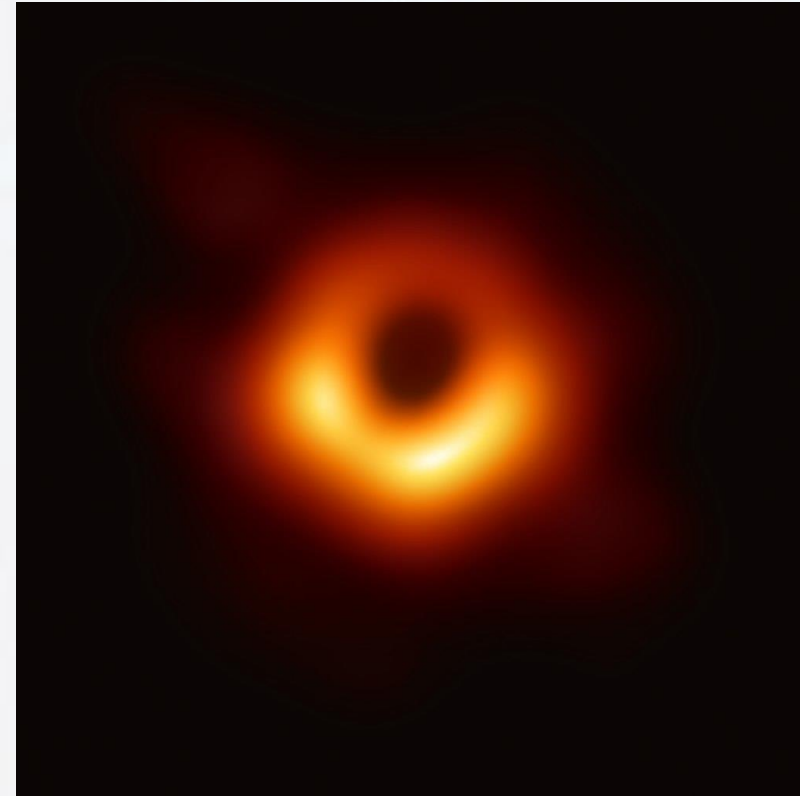


- Data reduction for Radio Astronomy
- Developed by NRAO, ESO, NAOJ & JIVE
  - C++ and Python
- Handles full data reduction
- Built for connected element arrays
- Used by VLA and ALMA pipelines
- Development process well defined

- Each feature is a JIRA ticket and git branch
- Scope of the issue defined
- Implementation, verification, validation
  - Verification requires unit tests and regression tests
  - Validation done by someone else
  - Centralised package building/testing on commit (can be turned on)
- Software engineering best practices!
- (*We really don't want to break ALMA pipelines!*)

# CASA, VLBI and JIVE

- CASA didn't have VLBI features
- NRAO always wanted them
- JIVE added them
- Continued work on new features







# New CASA papers for added FAIRness!



Cornell University

arXiv > astro-ph > arXiv:2210.02276

Astrophysics > Instrumentation and Methods for Astrophysics

[Submitted on 5 Oct 2022]

## CASA, the Common Astronomy Software Applications for Radio Astronomy

THE CASA TEAM, Ben Bean (1), Sanjay Bhatnagar (2), Sandra Castro (3), Jennifer Donovan Meyer (4), Bjorn Emonts (4), Enrique Garcia (3), Robert Garwood (4), Kumar Golap (2), Justo Gonzalez Villalba (3), Pamela Harris (2), Yohei Hayashi (4), Jagannathan (2), Wataru Kawasaki (5), Aard Keimpema (6), Mark Kettenis (6), Jorge Lopez (4), McNichols (4), David Mehringer (4), Renaud Miel (5), George Moellenbrock (2), Federico Petry (3), Martin Pokorny (2), Ryan Raba (4), Urvashi Rau (2), Darrell Schiebel (4), Neal Steeb (6), Des Small (6), Jan-Willem Steeb (4), Kanako Sugimoto (5), Ville Suoranta (4), Takahiro Ueda (6), Akeem Wells (4), Wei Xiong (1), Arpad Szomoru (6), Morgan Griffith (4), Brian Glenn Foster (3), ESO, (4) NRAO Charlottesville, (5) NAOJ, (6) JIVE, (7) IDIA

CASA, the Common Astronomy Software Applications, is the primary data processing software for the Karl G. Jansky Very Large Array (VLA), and is frequently used also for other radio telescopes. The CASA synthesis, and Very Long Baseline Interferometry (VLBI) telescopes. One of its core functionalities is the processing of VLA Sky Survey (VLASS), and the Nobeyama 45m telescope. This paper presents a high-level overview of the procedures for calibrating and imaging astronomical radio data in CASA. CASA is being developed and maintained based at the National Radio Astronomical Observatory (NRAO), the European Southern Observatory (ESO), and the Joint Institute for VLBI European Research Infrastructure Consortium (JIVE-ERIC), under the leadership of the

Cornell University

arXiv > astro-ph > arXiv:2210.02275

Astrophysics > Instrumentation and Methods for Astrophysics

[Submitted on 5 Oct 2022]

## CASA on the fringe -- Development of VLBI processing capabilities for CASA

Michael Janssen, George A. Moellenbrock, Dirk Petry, Ciriaco Goddi, Justin D. Linford, Kazi L.J. Rygl, Akeem Wells, Marina Medina, Neal Schweigart, Marjolein Verkouter, Aard Keimpema, Arpad Szomoru, Huib Jan van Grooten

Very Long Baseline Interferometry (VLBI) data has been implemented in the CASA package. This includes two new tasks to handle fringe finding tasks have been adjusted to handle VLBI visibility data and calibration meta-data properly. With these changes, CASA can now process VLBI data and spectral line observations in CASA. This article describes the development and implementation, and the use of CASA for European VLBI Network or Very Long Baseline Array data in CASA. Though the CASA VLBI functionality has been used for processing of the Effelsberg 100m Telescope data processing, in this paper we compare results for the same dataset processed in CASA and conclude that CASA in some cases performs better, though it cannot match AIPS for single-core processing. The new capabilities for easy development of pipelines or Jupyter notebooks, and thus contributes to raising VLBI data processing to a higher level of FAIRness and reusability.

zenodo

October 3, 2022

### Jupyter-CASA

Aard Keimpema; Mark Kettenis; Des Small; Ilse van Bemmel

A Jupyter kernel for CASA, a popular data processing suite for radio astronomy. The software is packaged together with CASA as a Docker container.

3 views 2 downloads

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codemeta.json	2.0 kB	Download
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jupyter-casa-docker-6.5.1.tar.gz	3.1 GB	Download
md5:88b6b6bc2d3d2c455ebc853662704ea0		Download

Public Citations

Show only: Literature (0) Dataset (0) Software (0) Unknown (0)

No citations.

Publication date: October 3, 2022

DOI: 10.5072/zenodo.1111745

Keyword(s): JIVE, Astronomy, jupyter-notebook

Grants: European Commission

- ESCAPE - European Science Cluster of Astronomy & Particle physics ESFRI research infrastructures (824064)

Related identifiers: https://arxiv.org/abs/2210.02275 (Preprint), https://arxiv.org/abs/2210.02276 (Preprint)



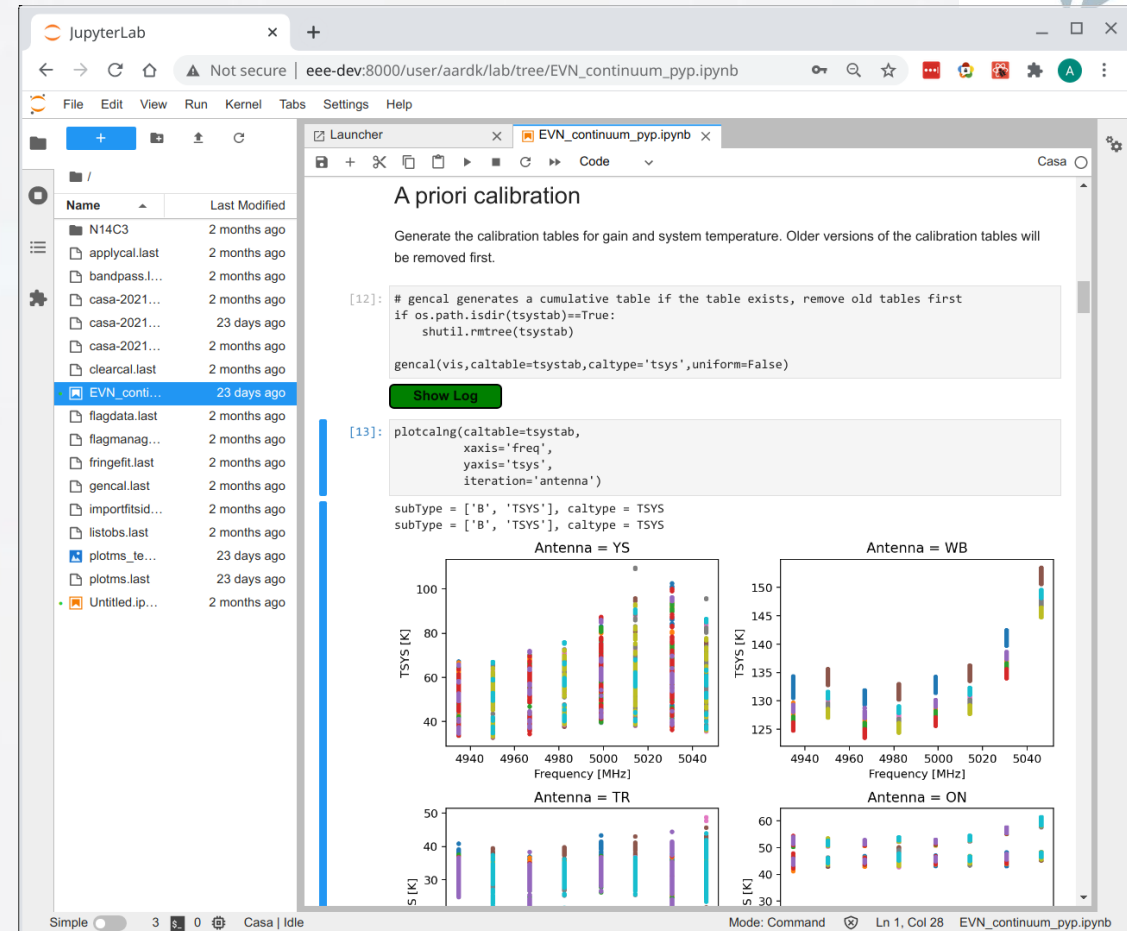


# Our OSSR contribution



JIVE

- Wrapped CASA in Jupyter notebooks
- We can offer “CASA in the cloud”
- Trial run at ERIS 2022
- Will be covered by A Keimpema



**A priori calibration**

Generate the calibration tables for gain and system temperature. Older versions of the calibration tables will be removed first.

```
[12]: # gencal generates a cumulative table if the table exists, remove old tables first
if os.path.isdir(tsystab)==True:
    shutil.rmtree(tsystab)

gencal(vis, caltable=tsystab, caltype='tsys', uniform=False)
```

**Show Log**

```
[13]: plotcalg(caltable=tsystab,
            xaxis='freq',
            yaxis='tsys',
            iteration='antenna')
```

subType = ['B', 'TSYS'], caltype = TSYS  
subType = ['B', 'TSYS'], caltype = TSYS

**Antenna = YS**

**Antenna = WB**

**Antenna = TR**

**Antenna = ON**

Mode: Command Ln 1, Col 28 EVN\_continuum\_pyp.ipynb



Thanks!