



Recent developments around DARWIN

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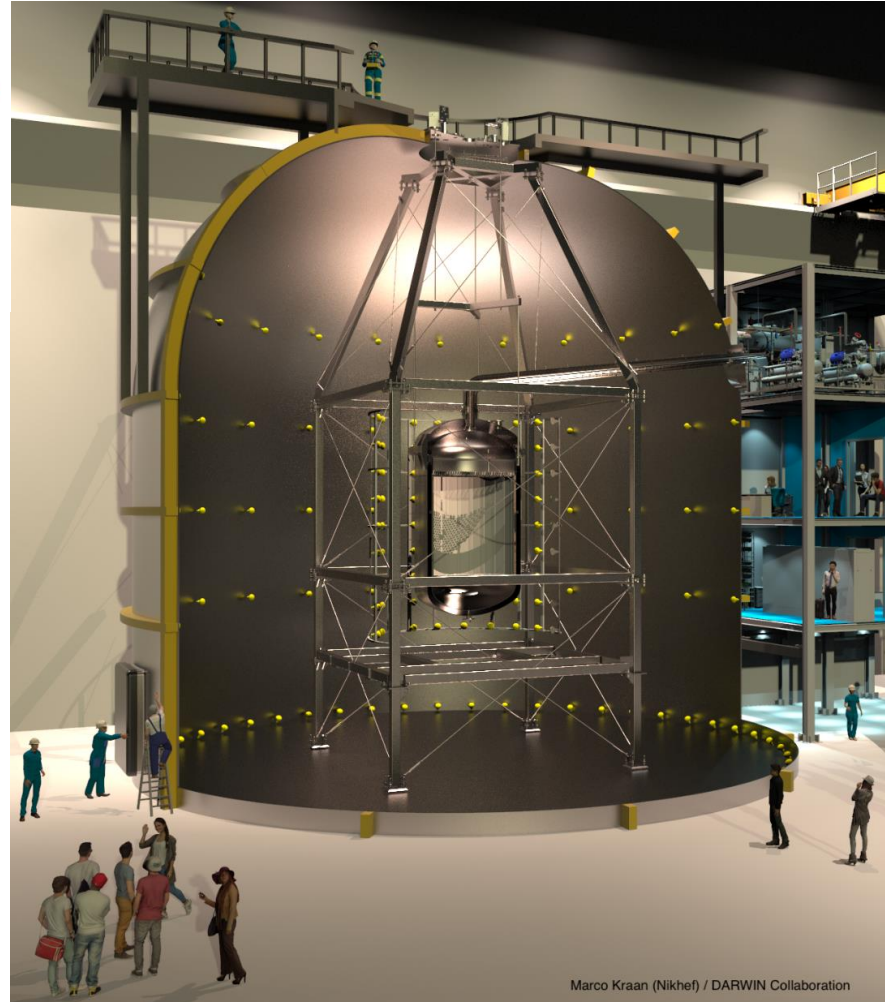
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in cooperation with

Julien Masbou

Luca Scotto Lavina

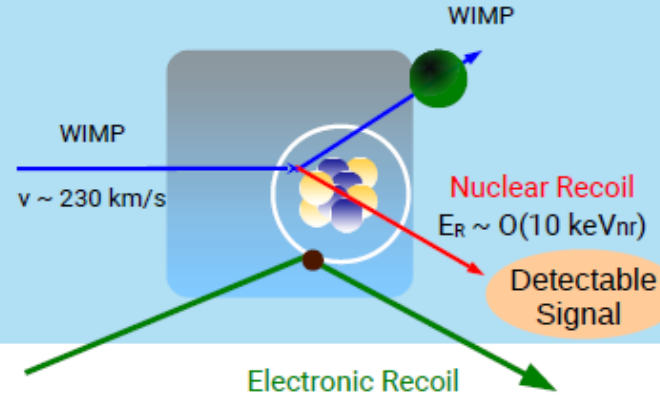


Marco Kraan (Nikhef) / DARWIN Collaboration

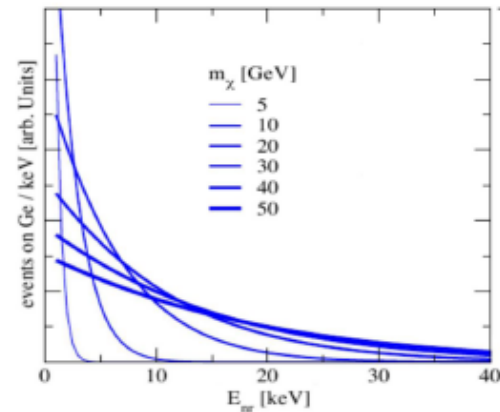
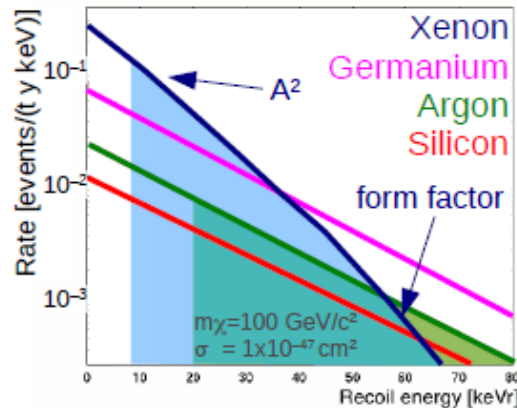
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WIMP scattering principles

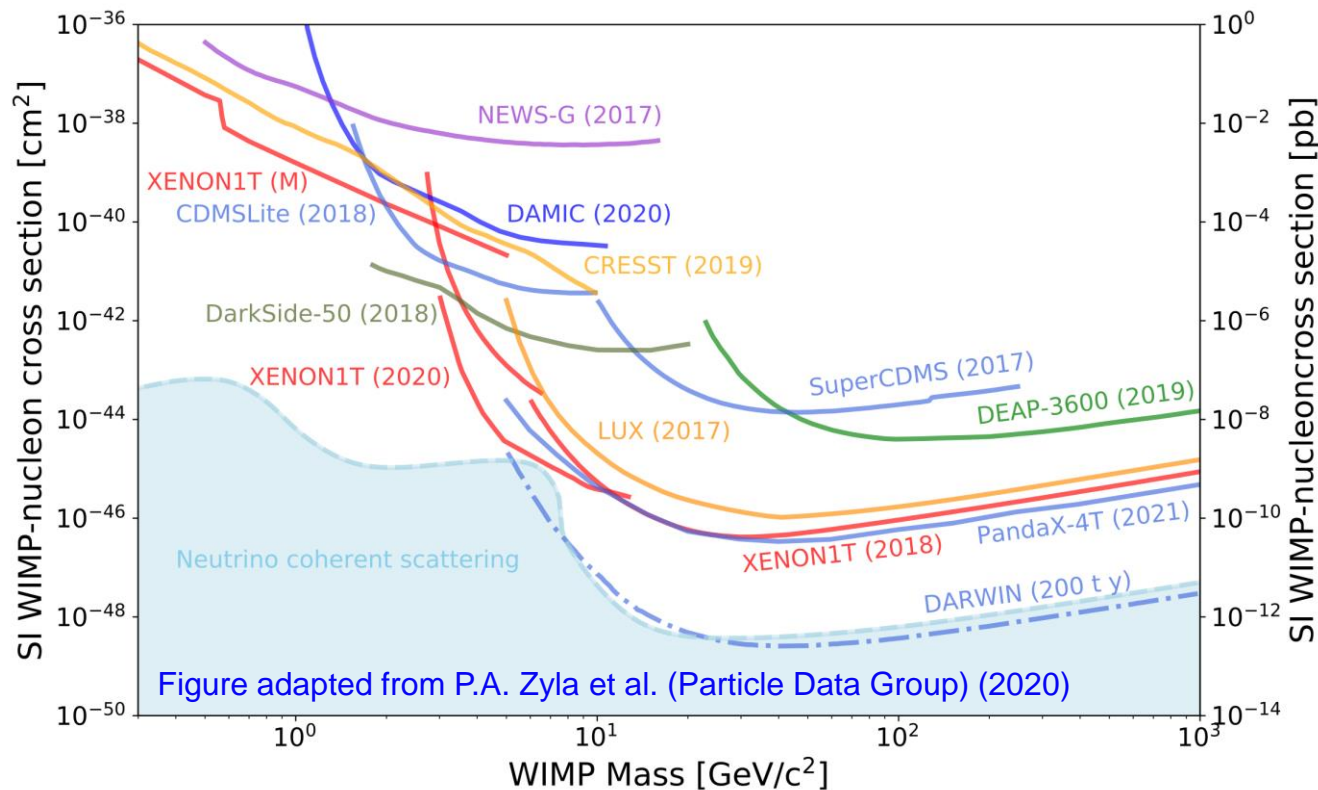
Elastic Scattering of
WIMPs off target nuclei
→ nuclear recoil



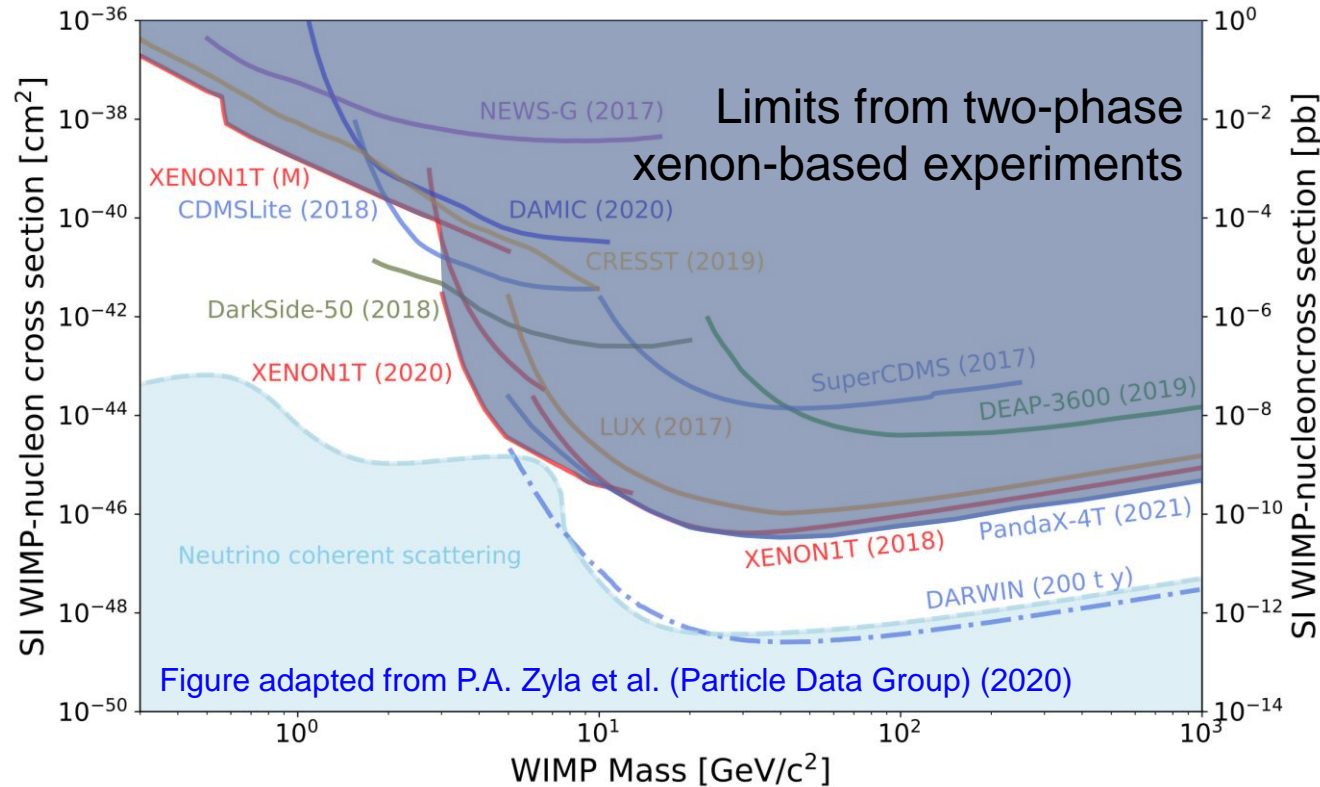
Recoil Spectra:



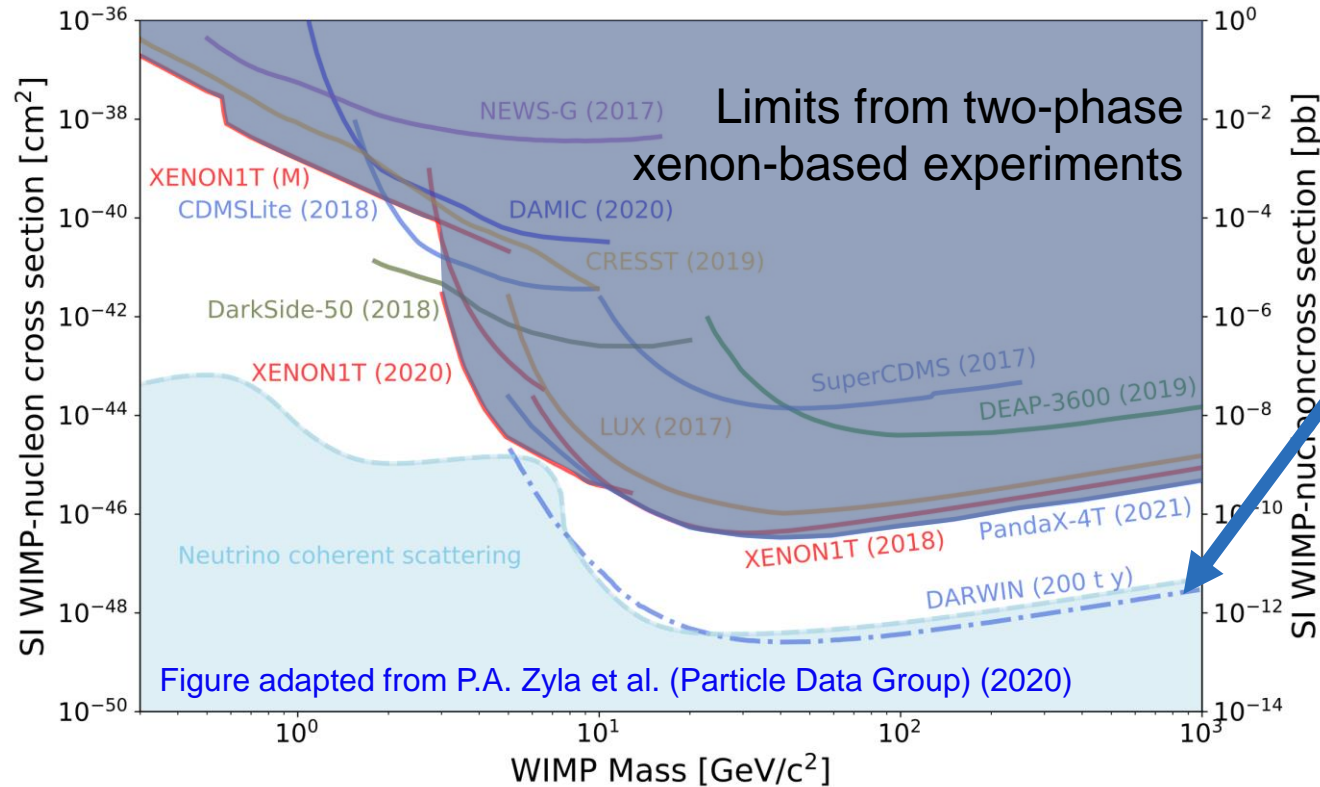
WIMP landscape (SI scattering)



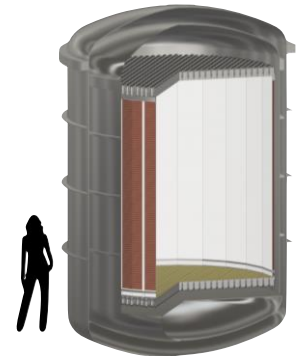
WIMP landscape (SI scattering)



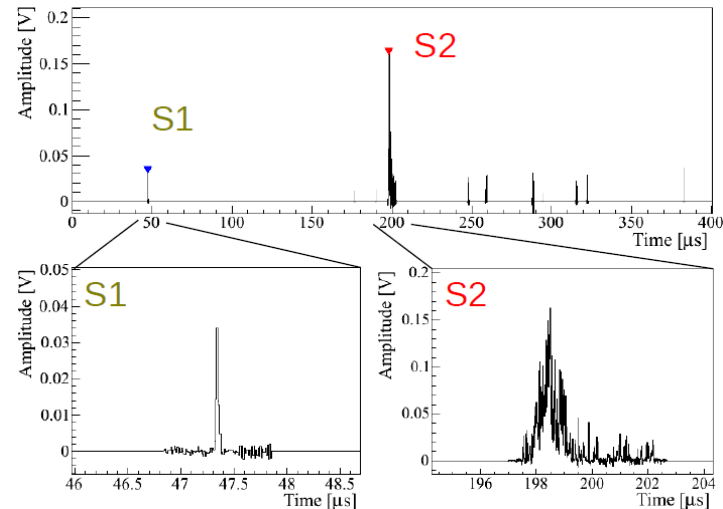
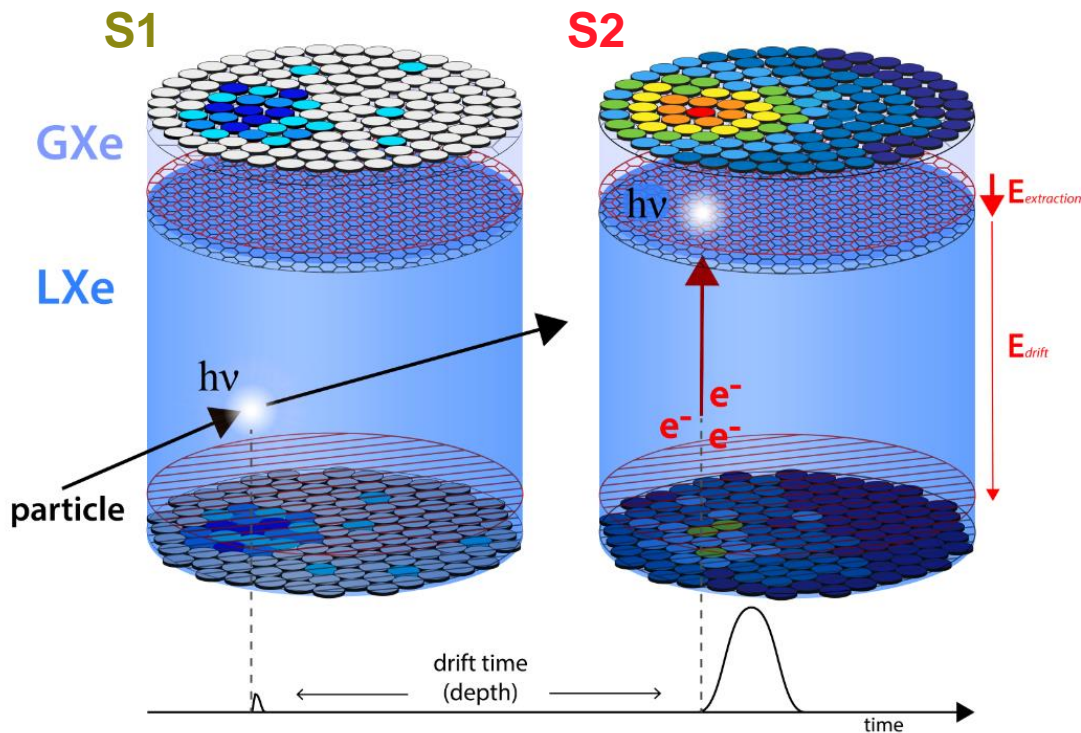
WIMP landscape (SI scattering)



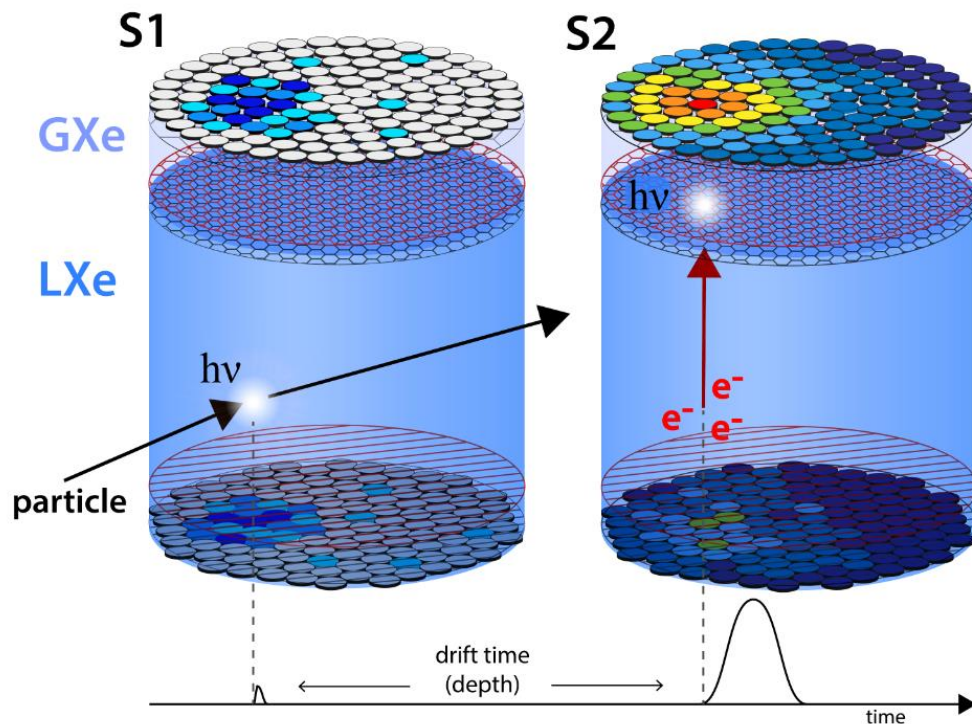
DARWIN -
the next generation



The dual-phase Xe TPC

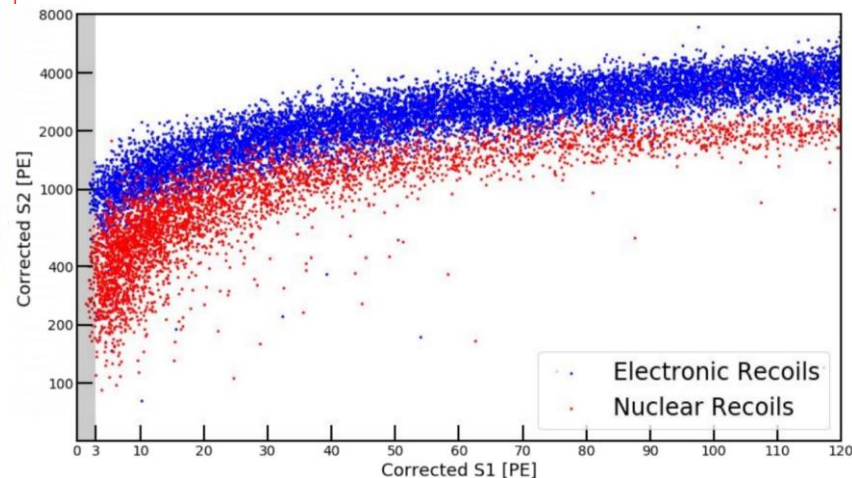


The dual-phase Xe TPC

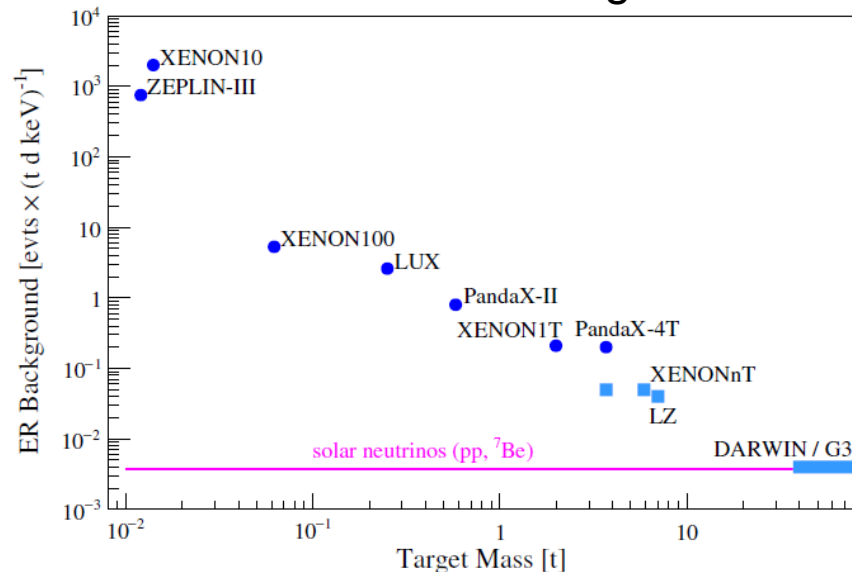


Combination of **S1** and **S2** allows for:

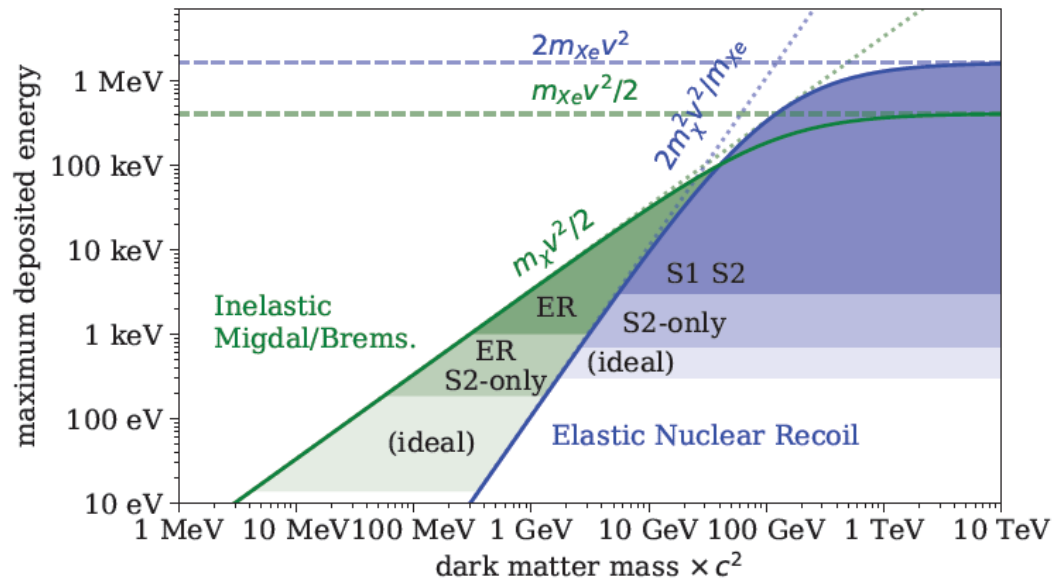
- 3D Position reconstruction
- Energy reconstruction
- **ER/NR** discrimination



Larger target mass
& lower background

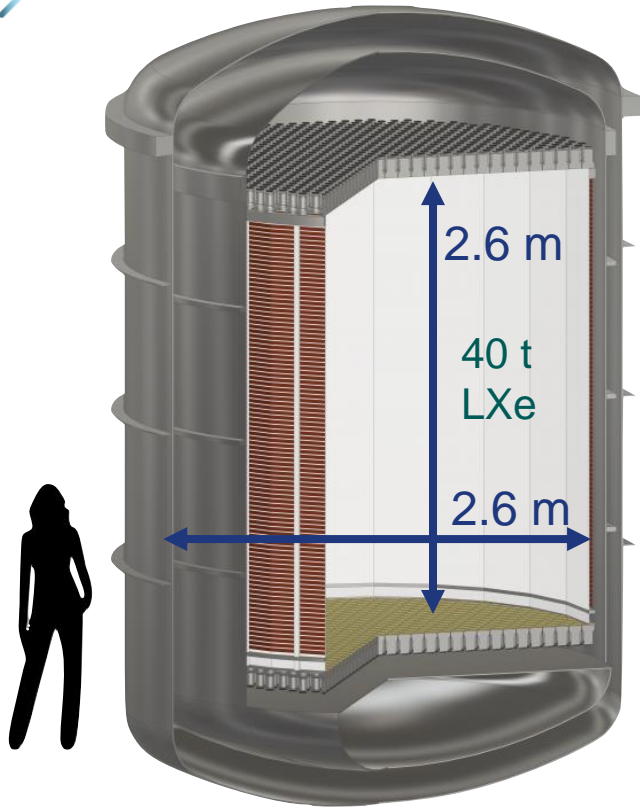


Interaction kinematics
→ lowering the threshold



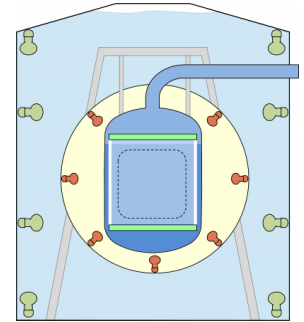
Community Whitepaper: [arXiv:2203.02309](https://arxiv.org/abs/2203.02309) (2022)

The DARWIN baseline design



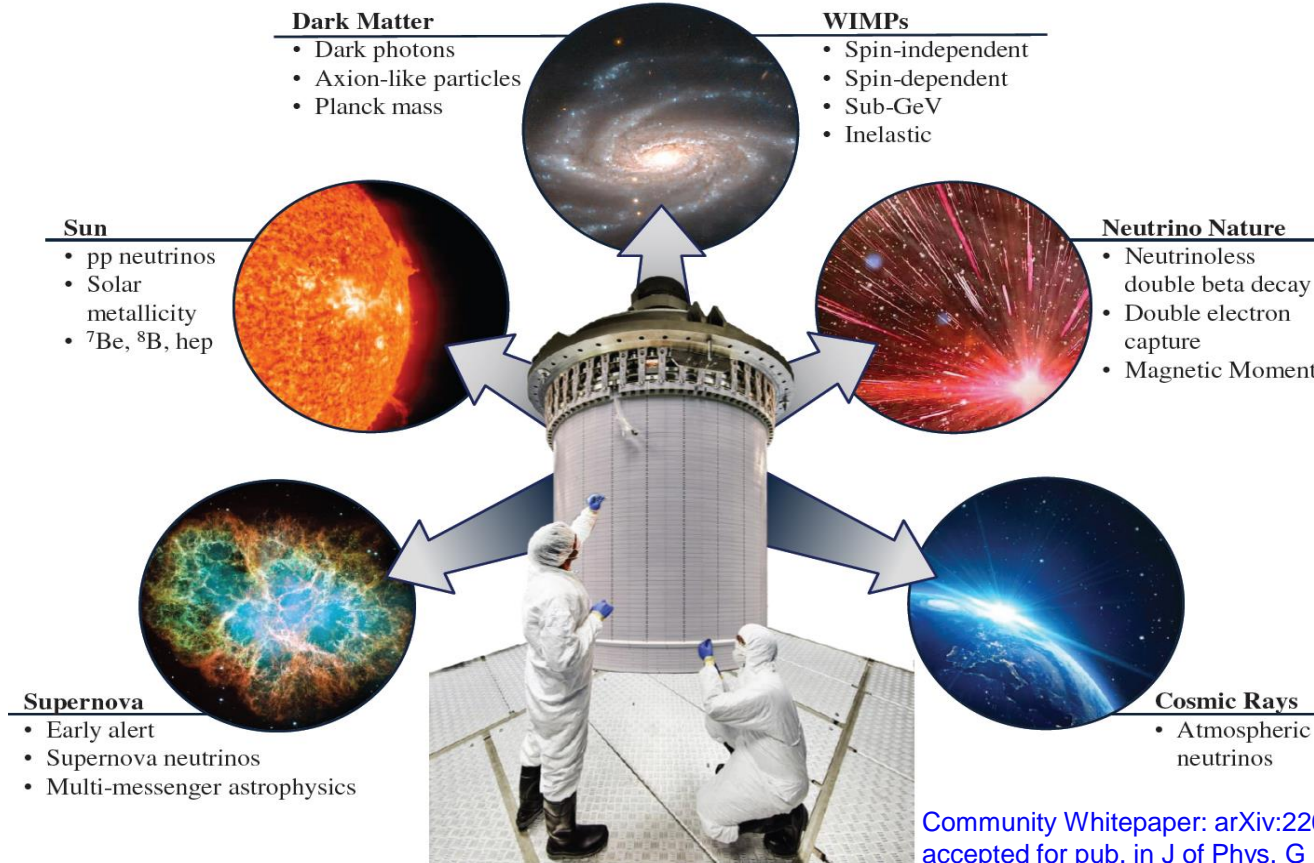
DARk matter WImp search with liquid xenon

- Two-phase LXe/GXe TPC; aspect ratio 1 ($\varnothing=h$)
- 50 t total LXe (40 t active target)
- Top and bottom photosensors (~ 1800 3" XENON PMTs)
- PTFE reflectors and Cu field-shaping rings
- In-situ purification plus krypton and radon distillation (background mitigation)
- Veto detectors: water Cerenkov for muons with Gd doping for neutrons



DARWIN Collaboration, JCAP 1611 (2016) 017

Science channels for DARWIN



Science channels for DARWIN



Melih Kara

Dark Matter

- Dark photons
- Axion-like particles
- Planck mass

WIMPs

- Spin-independent
- Spin-dependent
- Sub-GeV
- Inelastic

Sun

- pp neutrinos
- Solar metallicity
- ${}^7\text{Be}$, ${}^8\text{B}$, hep

Neutrino Nature

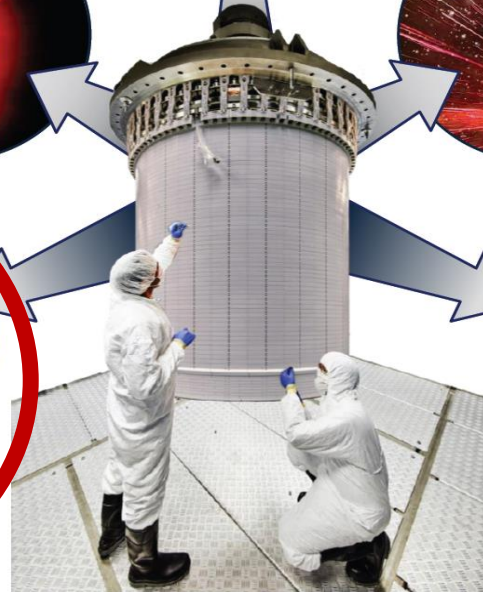
- Neutrinoless double beta decay
- Double electron capture
- Magnetic Moment

Supernova

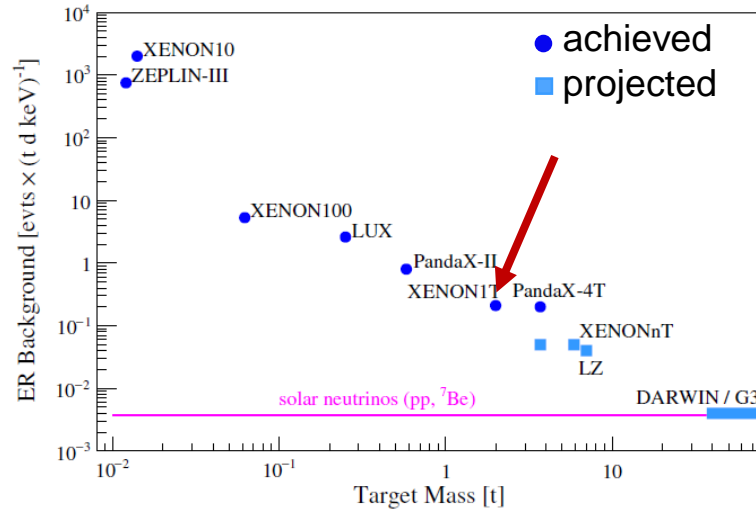
- Early alert
- Supernova neutrinos
- Multi-messenger astrophysics

Cosmic Rays

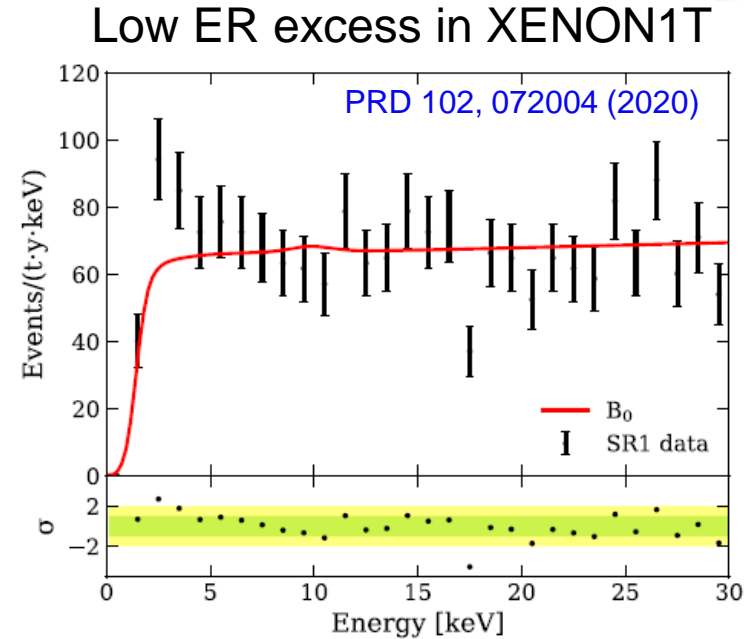
- Atmospheric neutrinos

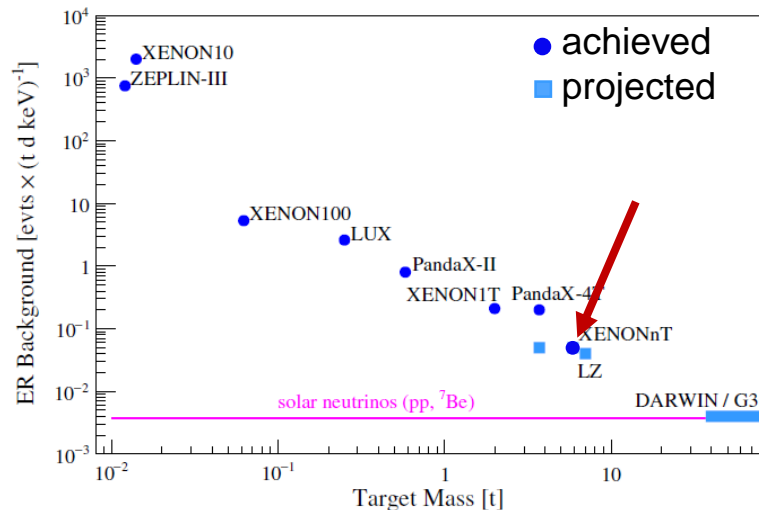


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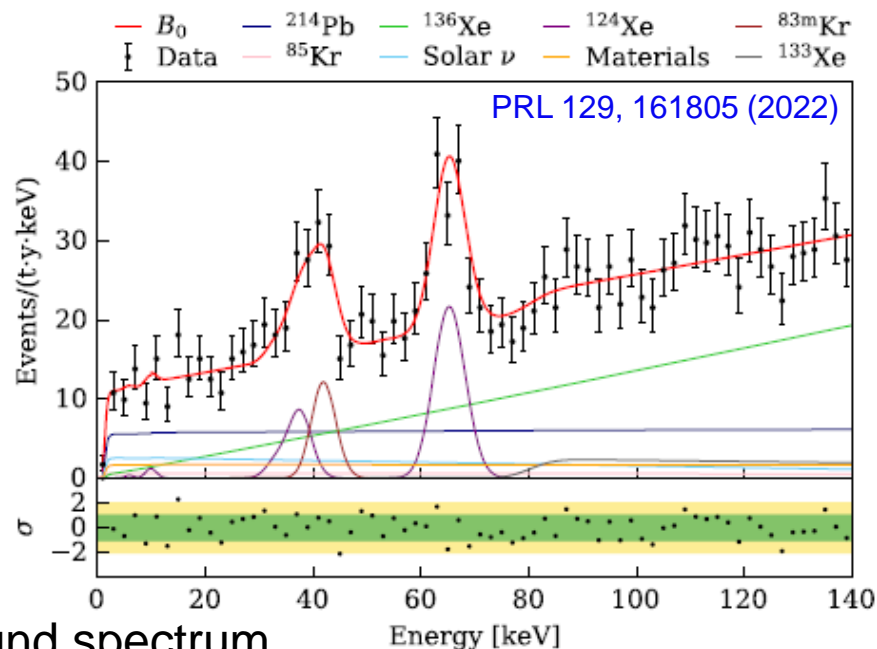


- solar axions?
- bosonic DM?
- neutrino magnetic moment?
- tritium?

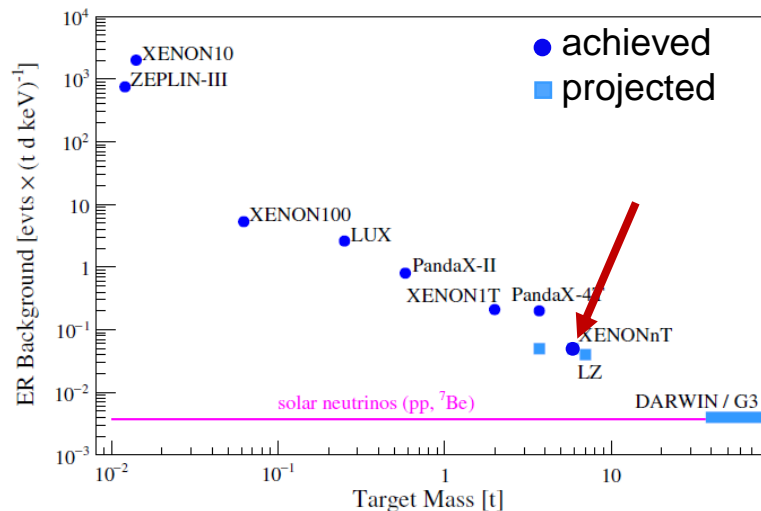




Low ER in XENONnT

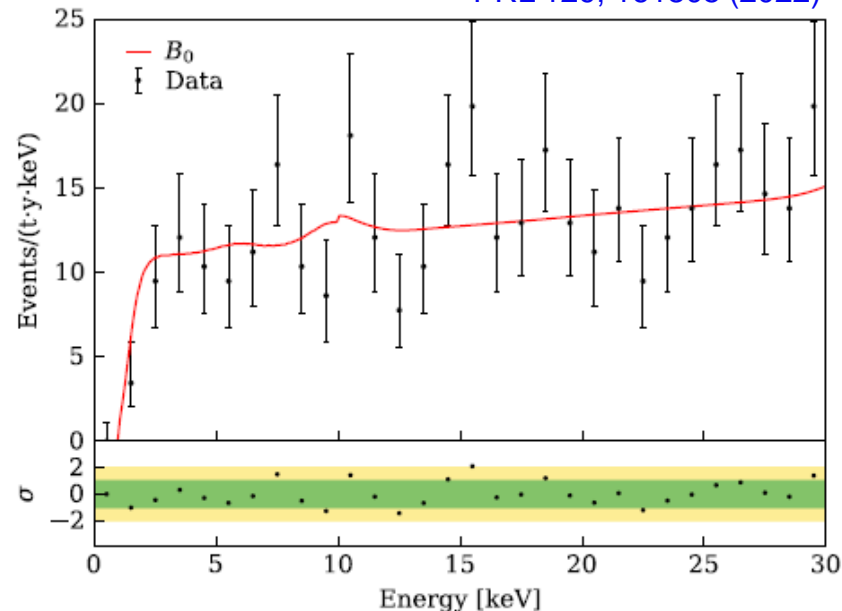


For the first time, shape of low-E background spectrum dominated by **second order weak decays** ($2\nu\beta\beta$ of ^{136}Xe , $2\nu\text{ECEC}$ of ^{124}Xe)

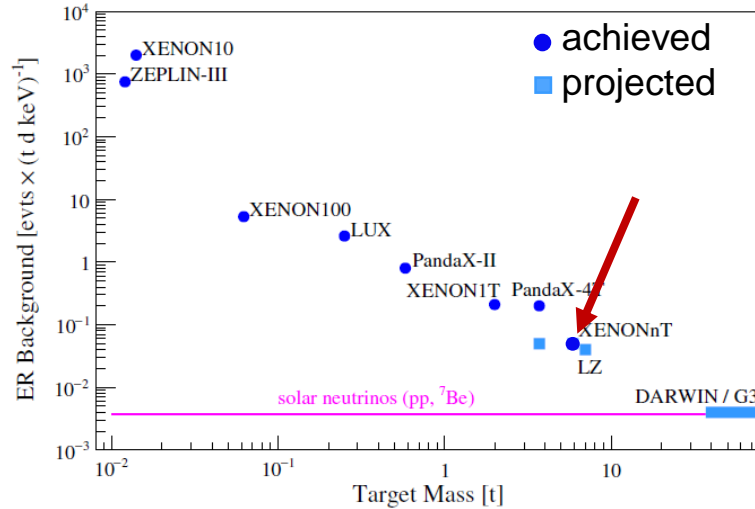


Low ER in XENONnT

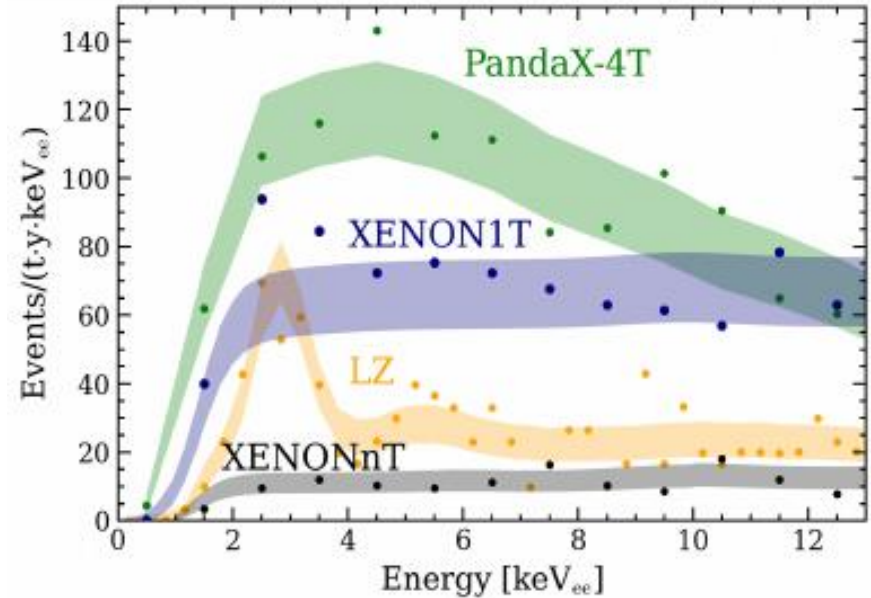
PRL 129, 161805 (2022)



No excess above background observed
 → XENON1T excess not from new physics

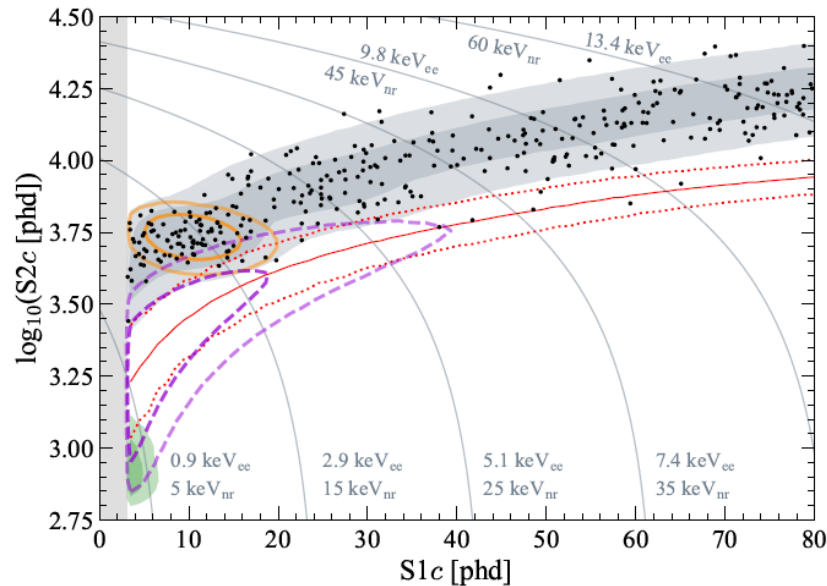


Low ER in XENONnT

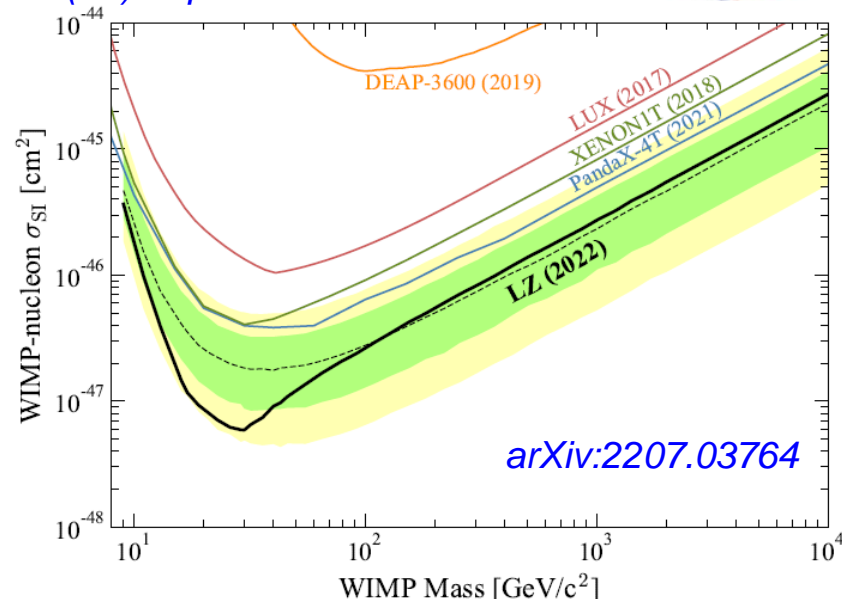


world-record **low ER background level**: (15.8 ± 1.3) evts/(txyr \times keV)

First Dark Matter Search Results from the LUX-ZEPLIN (LZ) Experiment



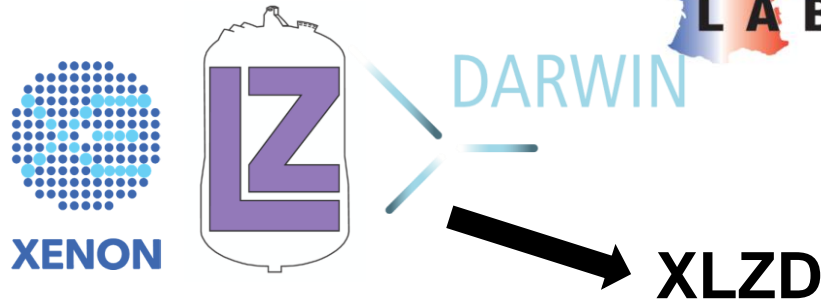
1 σ /2 σ contours of
 best fit ER background model
³⁷Ar background component
 30 GeV/c² WIMP model
⁸B solar neutrinos
 NR band (10%/90% quantiles)



arXiv:2207.03764

Non-Blind analysis
of 0.90 t \times y data
 → fiducial volume of 5.5t
 → very large downward
 fluctuation of background

Recent Developments (III)



- Memorandum of Understanding signed July 6, 2021 by 106 research group leaders from 16 countries
- Intl consortium w/ complementary areas of expertise
- 1st joint meeting April 26-27, 2021 online
- 2nd meeting June 27-29, 2022 at KIT

Same technology – Different detectors apart from different TPC design choices... different cryostats, LXe/neutron vetoes, LXe purification, Kr and Rn removal, calibration methods, DAQ paradigms, etc etc



<https://xlzd.org/>

A Next-Generation Liquid Xenon Observatory for Dark Matter and Neutrino Physics

J. Aalbers,^{1,2} K. Abe,^{3,4} V. Aerne,⁵ F. Agostini,⁶ S. Ahmed Maouloud,⁷ D.S. Akerib,^{1,2} D.Yu. Akimov,⁸ A.K. Al Musalhi,¹⁰ F. Alder,¹¹ S.K. Alsum,¹² L. Althueser,¹³ C.S. Amarasinghe,¹⁴ F.D. Amaro,¹⁵ A. T. I. Anderson,^{1,2} R. Andrieu,⁷ N. Angelides,¹⁶ E. Angelino,¹⁷ I. Angevaere,¹⁸ V.C. Antochi,¹⁹ D. Ant...

>600 authors from
XLZD + theory

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⁴⁰Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA
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⁴²Vatican Observatory. Castel Gandolfo. V-00120. Vatican City State

>100 institutions

The nature of dark matter and properties of neutrinos are among the most pressing issues in contemporary particle physics. The dual-phase xenon time-projection chamber is the leading technology to cover the available parameter space for Weakly Interacting Massive Particles (WIMPs), while featuring extensive sensitivity to many alternative dark matter candidates. These detectors can also study neutrinos through neutrinoless double-beta decay and through a variety of astrophysical sources. A next-generation xenon-based detector will therefore be a true multi-purpose observatory to significantly advance particle physics, nuclear physics, astrophysics, solar physics, and cosmology. This review article presents the science cases for such a detector.

Community Whitepaper:
arXiv:2203.02309
accepted for pub. in J Phys. G

[1] S. Ritz et al., *Building for Discovery: Strategic Plan for U.S. Particle Physics in the Global Context*, https://science.energy.gov/-/media/hep/hepap/pdf/May-2014/FINAL_P5_Report_Interactive_060214.pdf (2014).
C. Zeitnitz, *Particle Physics Strategy in Germany*, <https://indico.desy.de/indico/event/20166/contribution/6/material/slides/0.pdf> (2018).

[1258] N. Tyurin, *PARTICLE PHYSICS IN RUSSIA* (2012).
[1259] T. Nakada et al., *The European Strategy for Particle Physics Update 2013* (2013).
[1260] C. Bai et al., *Neutrinoless Double Beta Decay: A Study of Strategic Development by Chinese Academy of Sciences*, in *Chinese* (2020).
[1261] P. J. W. Faulkner et al. (GridPP), *J. Phys. G* **32**, N1 (2006).
[1262] D. Britton et al., *Phil. Trans. Roy. Soc. Lond. A* **367**, 2447 (2009).

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D. Supernova Neutrinos (NR)

1. Galactic Supernova Neutrinos
2. Pre-Supernova Neutrinos
3. Supernova Early Warning System
4. Diffuse Supernova Neutrinos

→ Melih Kara

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