

Dissecting the interstellar medium of extremely distant galaxies with Gamma-ray bursts

Speaker

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Rencontres des Jeunes Physicien·ne·s
2 November 2022



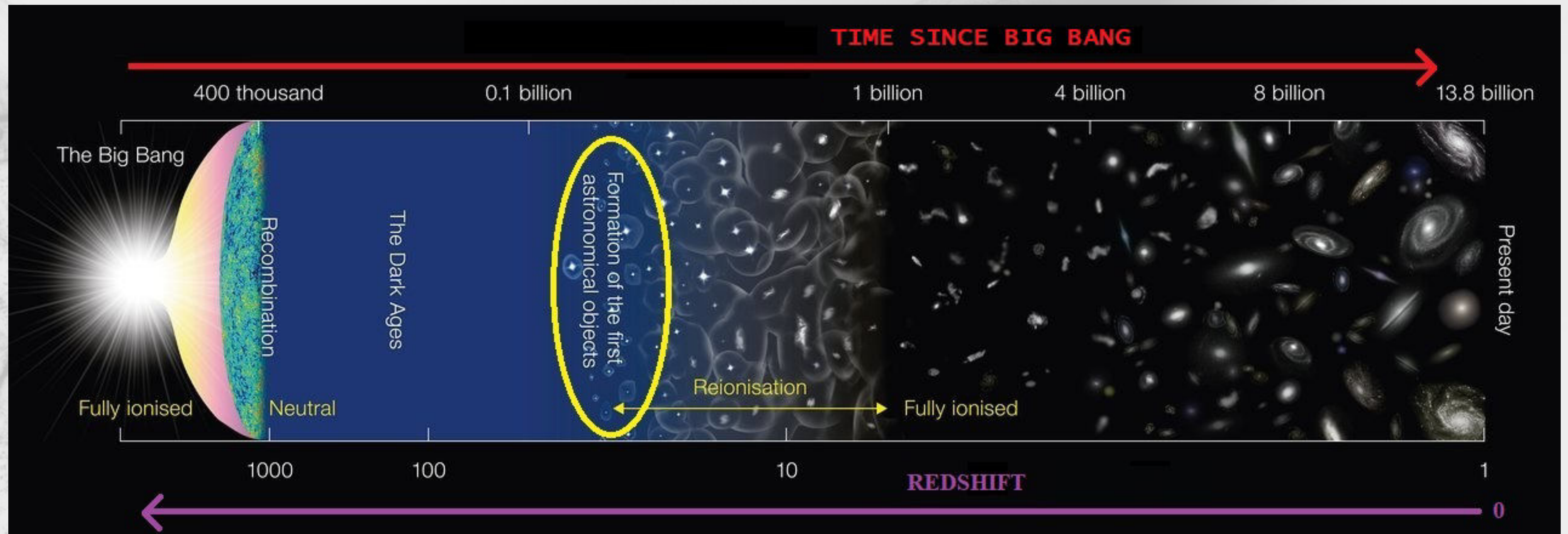
COLLÈGE
DE FRANCE
—1530—



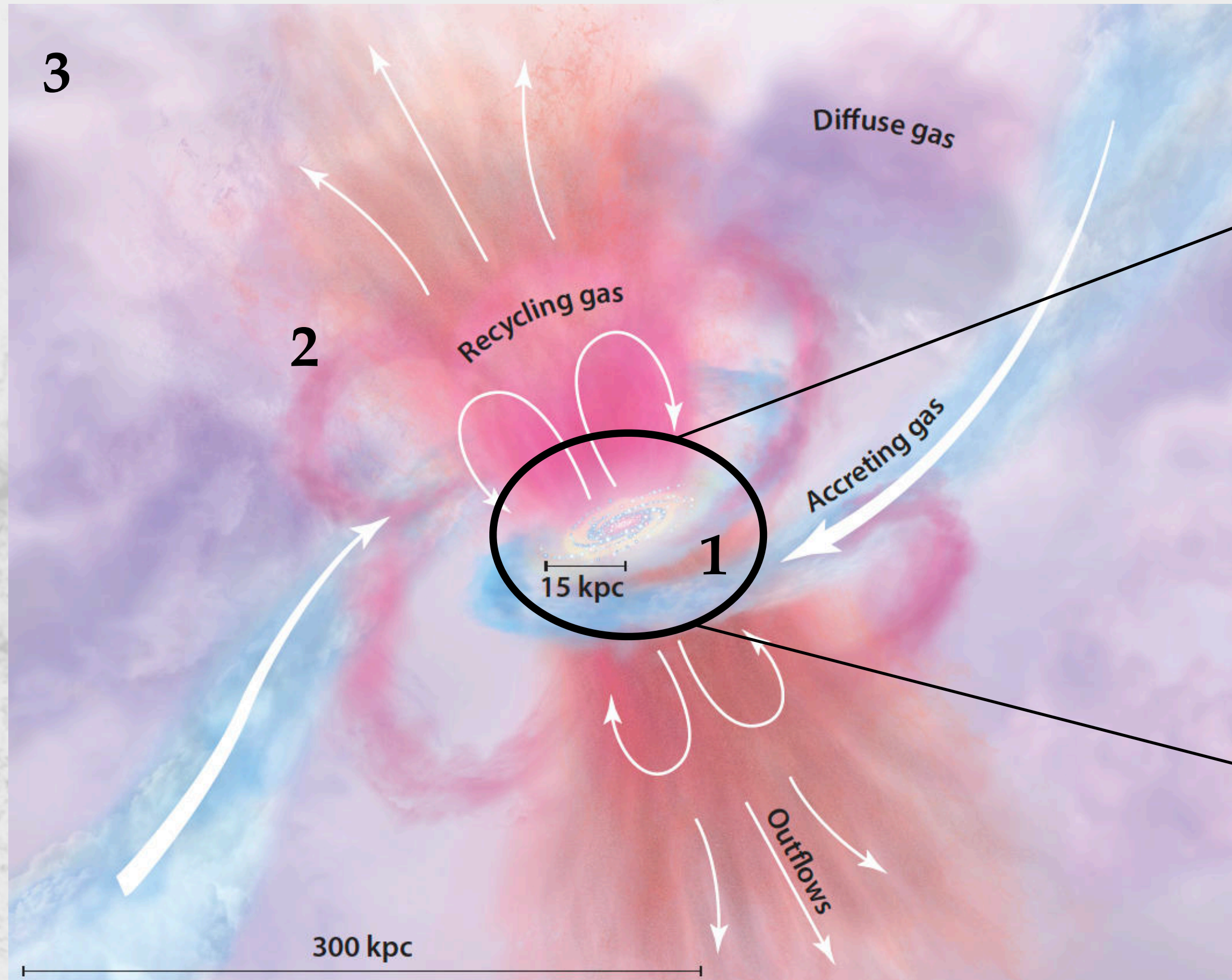
PSL
RESEARCH UNIVERSITY PARIS



- What are the first objects to be formed in the Universe?
- How do galaxies form and evolve?
- What is the interplay between star formation and the inter-stellar gas?



Credits: ESO



Credits: Tumlinson et al. 2017

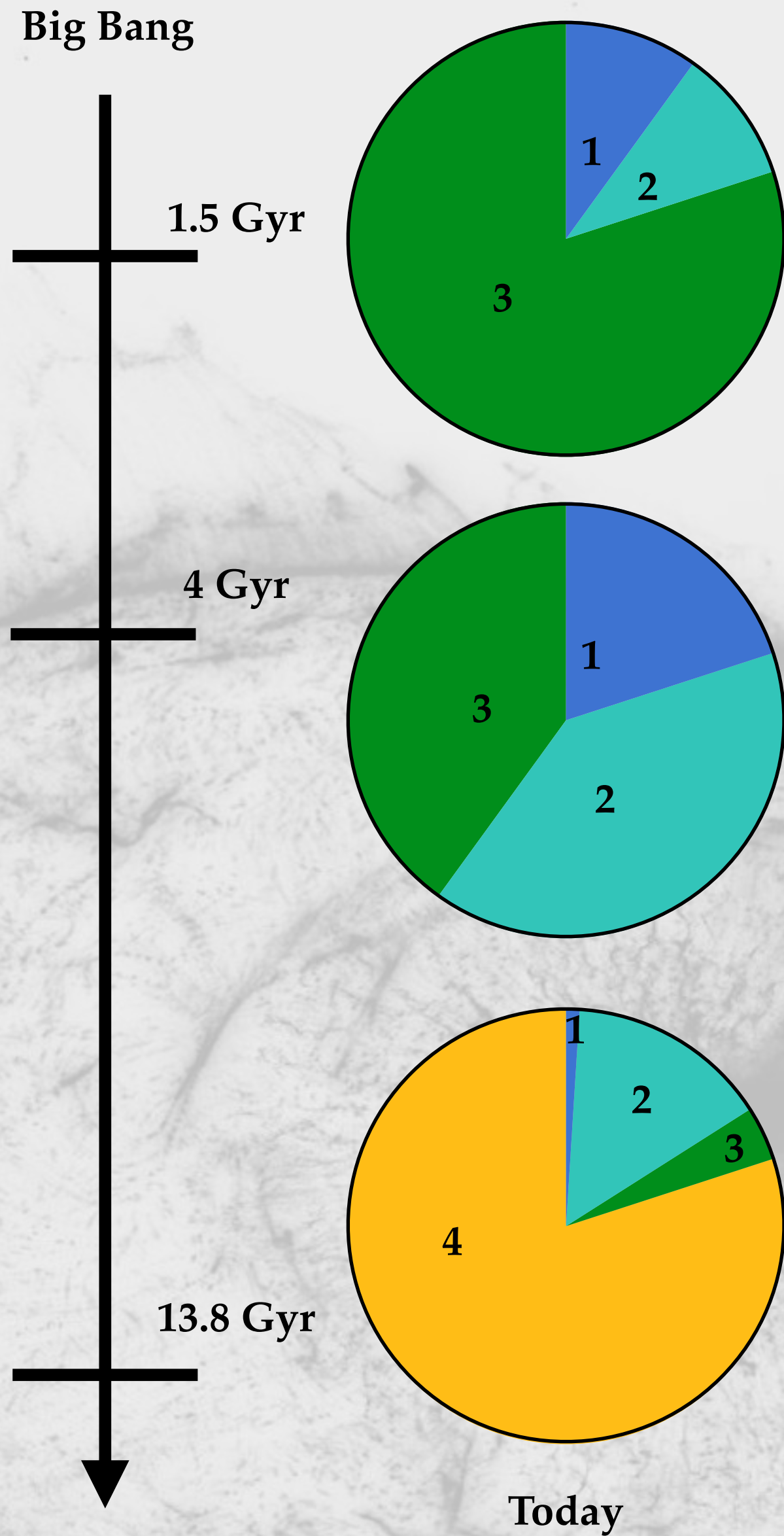


Credits: NASA

- Stars
- Neutral Gas
- Ionized Gas
- Dust

1. Inter-stellar medium (ISM)
2. Circum-galactic medium (CGM)
3. Inter-galactic medium (IGM)

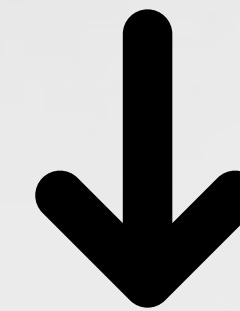
Stellar explosions contaminate the surrounding medium, composed by gas and dust, with heavy elements called "Metals"



- 1 Molecular Gas+Dust
- 2 Stars
- 3 Atomic Gas
- 4 Ionized gas

**In the early Universe
most of the gas in galaxies
is in the neutral phase**

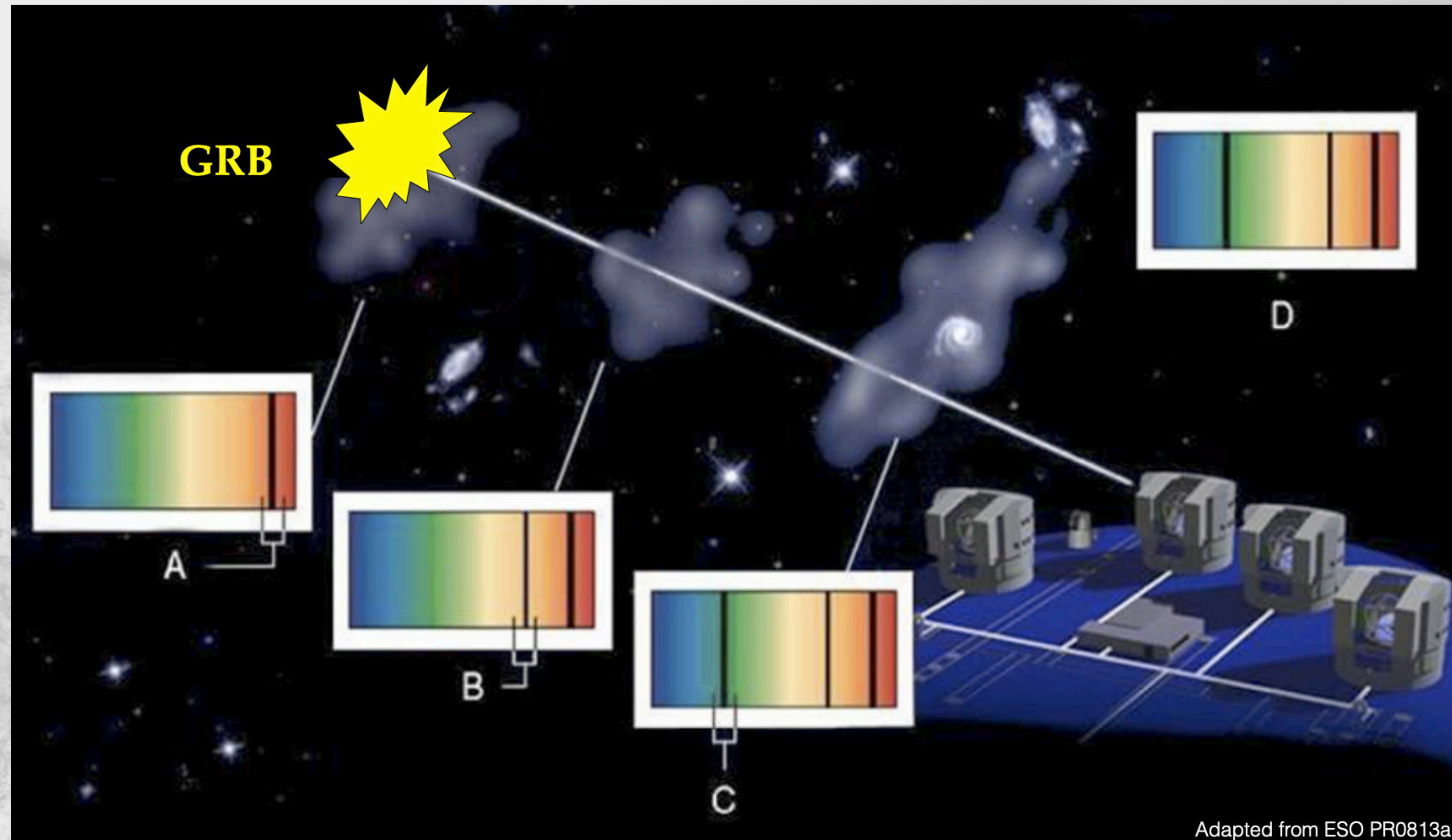
The neutral gas doesn't emit light
so it's hard to see...



Background Source

If a bright source shines through the gas,
it will absorb the light at different frequencies
depending on the atoms composing the gas

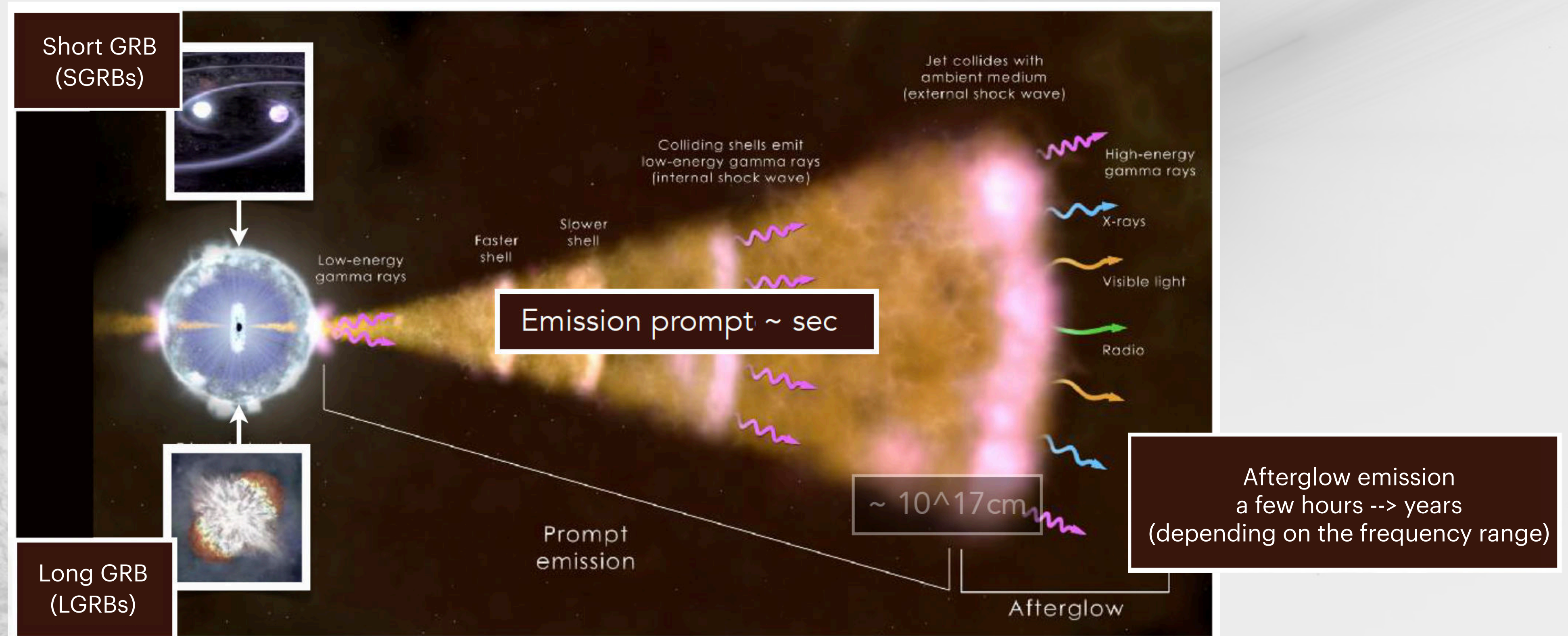
GRBs are unique powerful background sources to probe first galaxies



Adapted from ESO PR0813a

Credits: ESO

Ultra-Relativistic Jet produced by a new-born accreting black hole



SGRBs are linked with the merging of two compact objects.

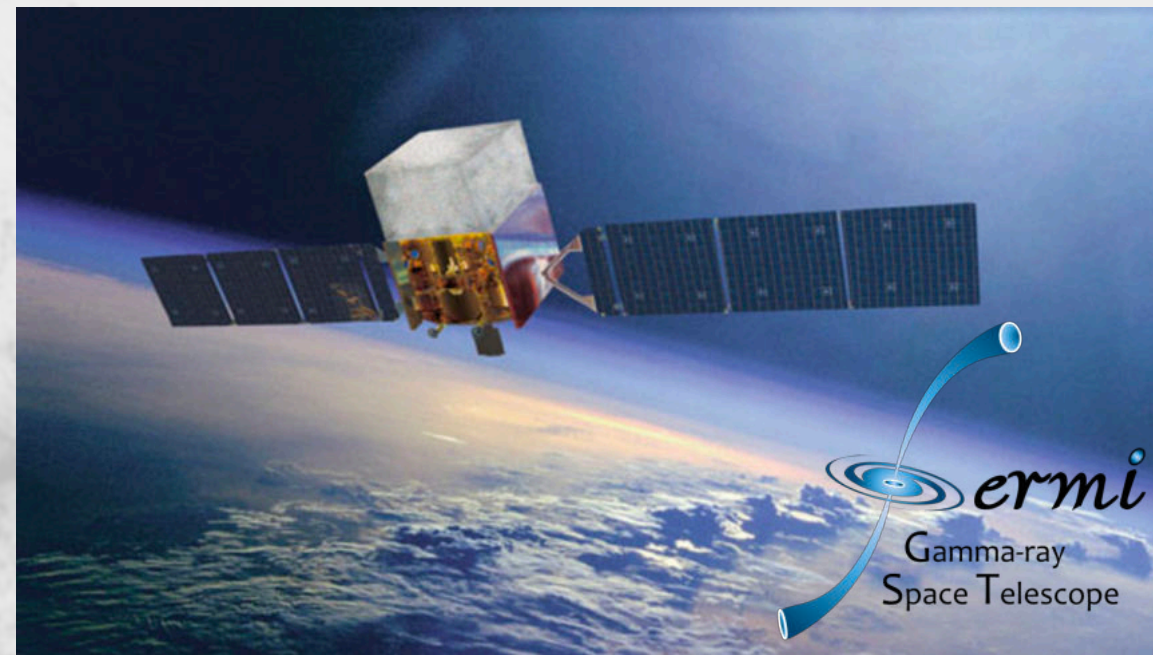
LGRBs are associated with the collapse of massive stars

THE GRB OBSERVATION



Satellite Alert

Fermi



Swift

Telescope Trigger

Minutes - Hours



Ground based Telescopes

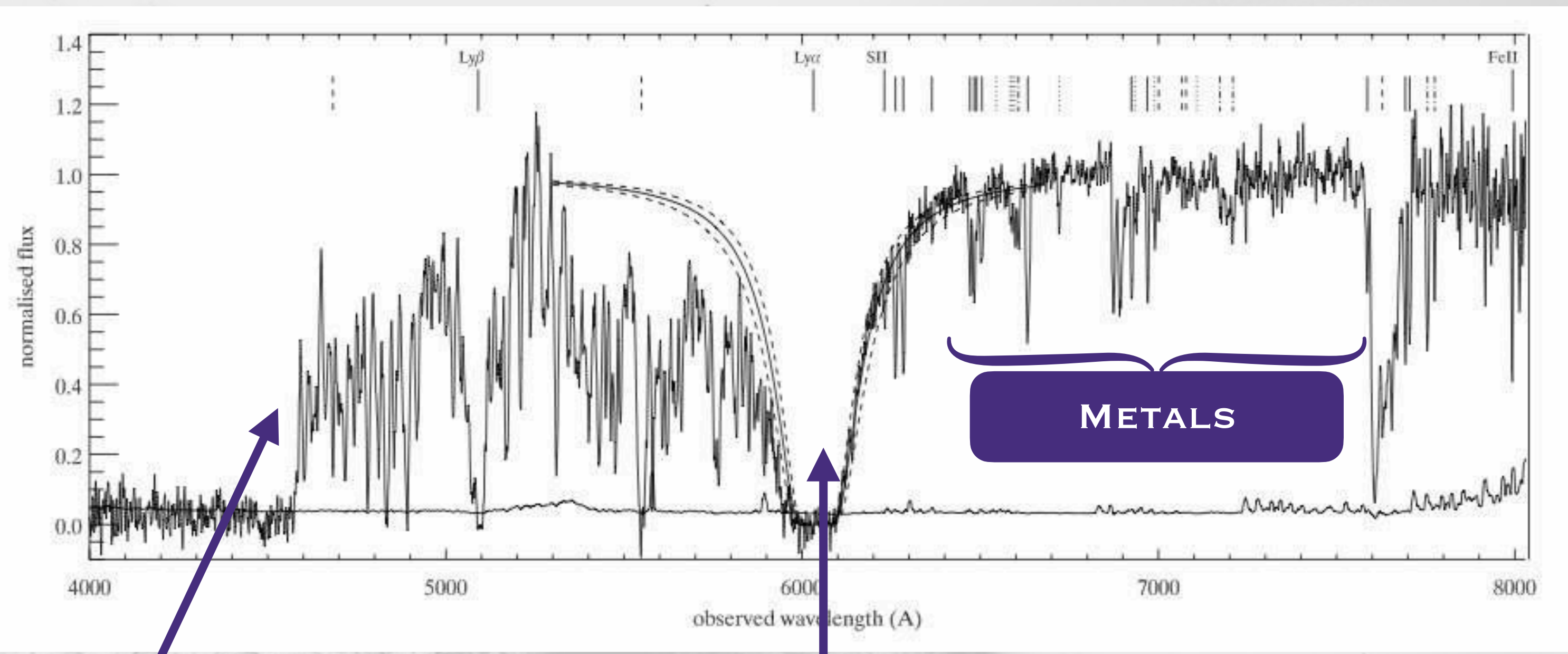


Spectroscopy

Photometry

LYMAN ALPHA FOREST
INTERGALACTIC MEDIUM (IGM)

GRB 050730 $z=3.968$



LYMAN LIMIT

HI (LYMAN ALPHA)
NEUTRAL HYDROGEN

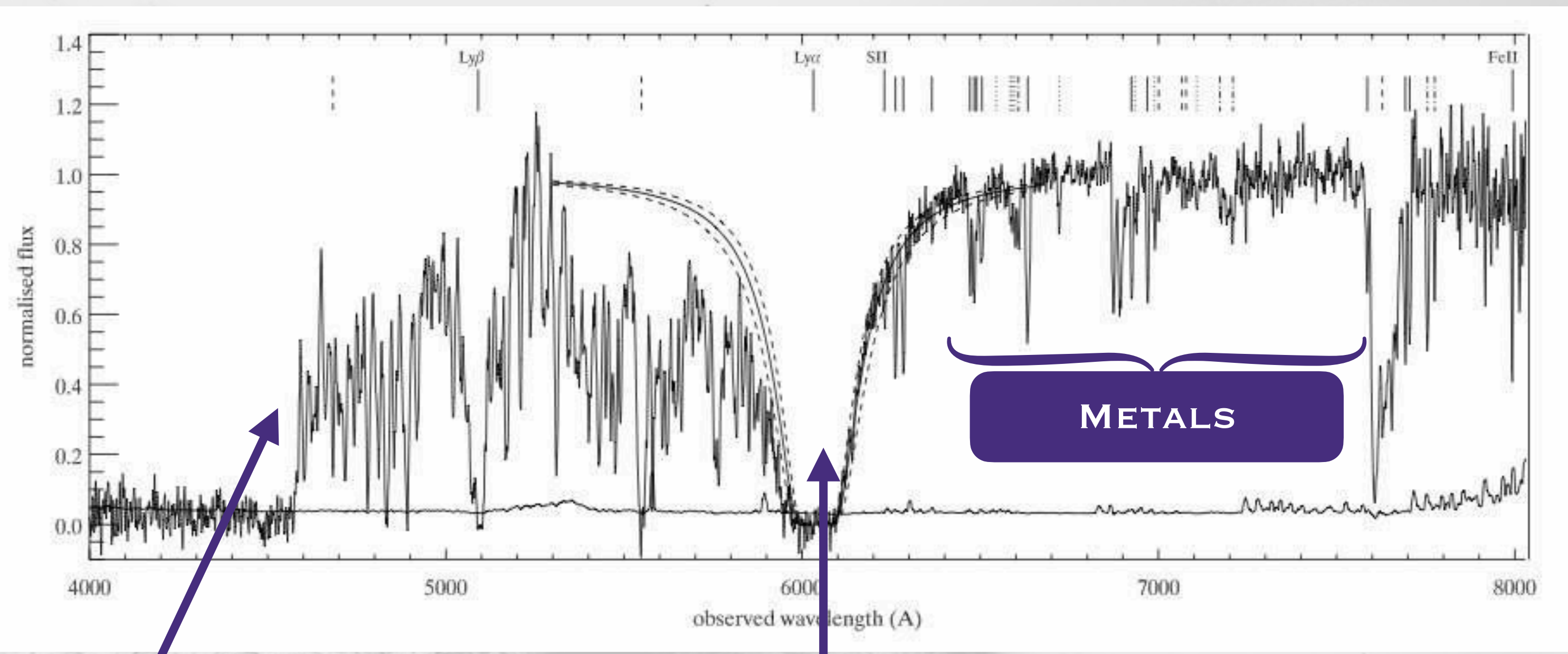
R. L. C. Starling+2005

From the analysis of the absorption lines we can measure:

- *Redshift of the absorbers*
- *Column densities of the ions of different chemical elements (neutral hydrogen N_{HI} and Metals N_{X})*

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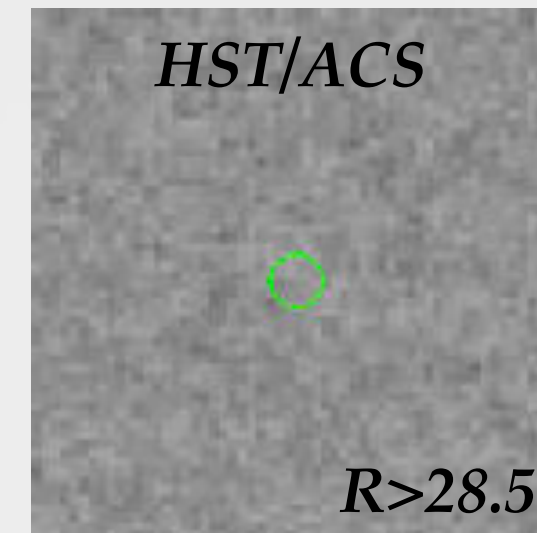
METALS

R. L. C. Starling+2005

From the analysis of the absorption lines we can measure:

- ➔ *Redshift of the absorbers*
- ➔ *Column densities of the ions of different chemical elements (neutral hydrogen N_{HI} and Metals N_X)*

Tool to select and study in detail faint star-forming galaxies



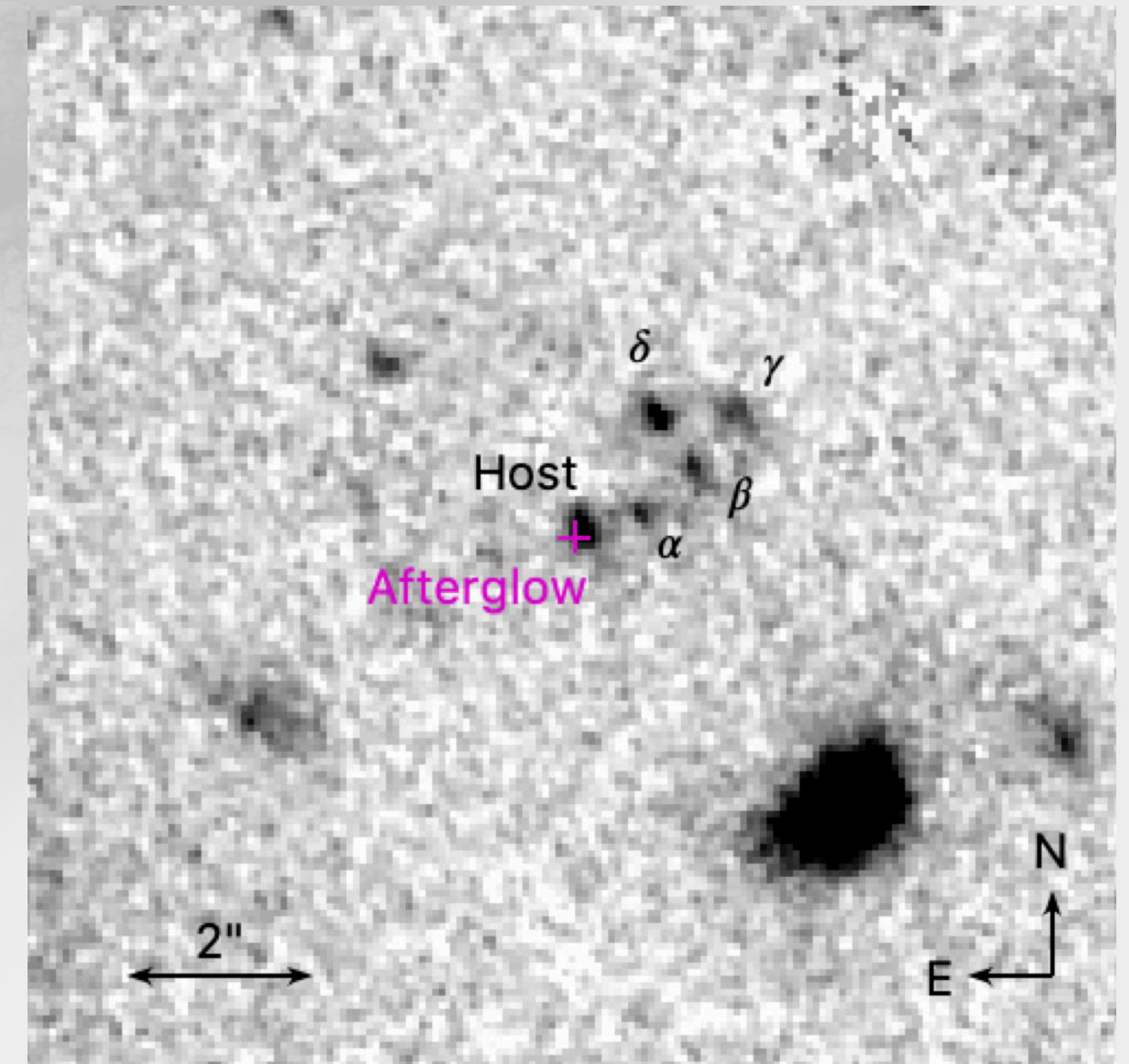
$R > 28.5$ Chen+2005

OBSERVATIONS

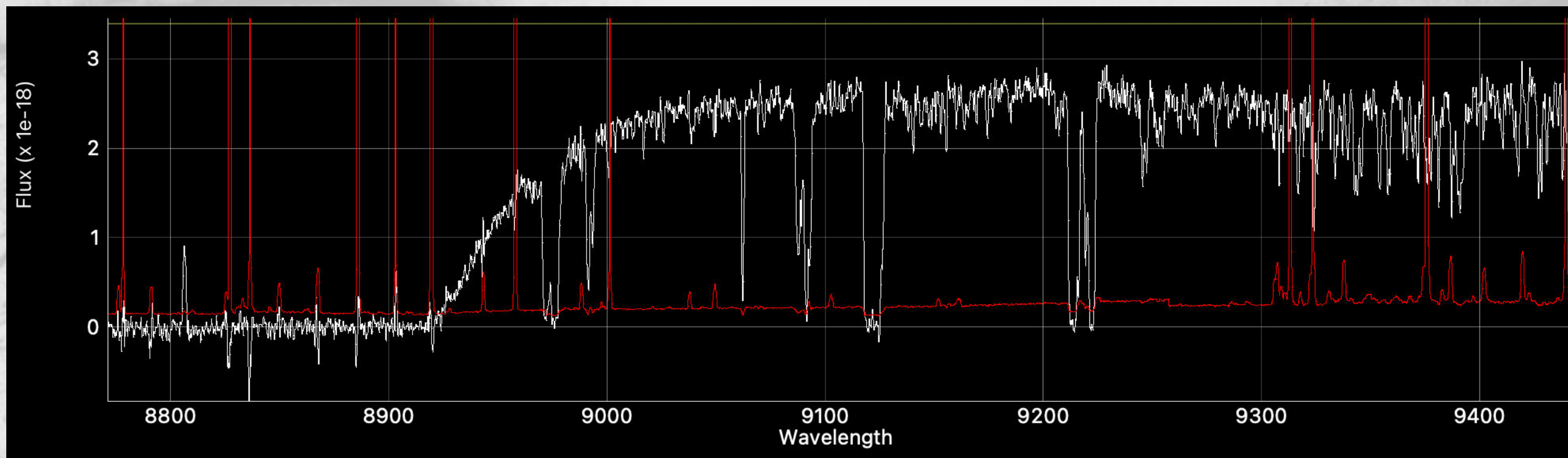
**European Southern Observatory (ESO)
Very Large Telescope (VLT)/X-shooter Spectrum**
After ~2.53h (obs frame) from the GRB trigger



Hubble Space Telescope (HST)/WFC3 Image
After ~250 days (obs frame) from the GRB trigger



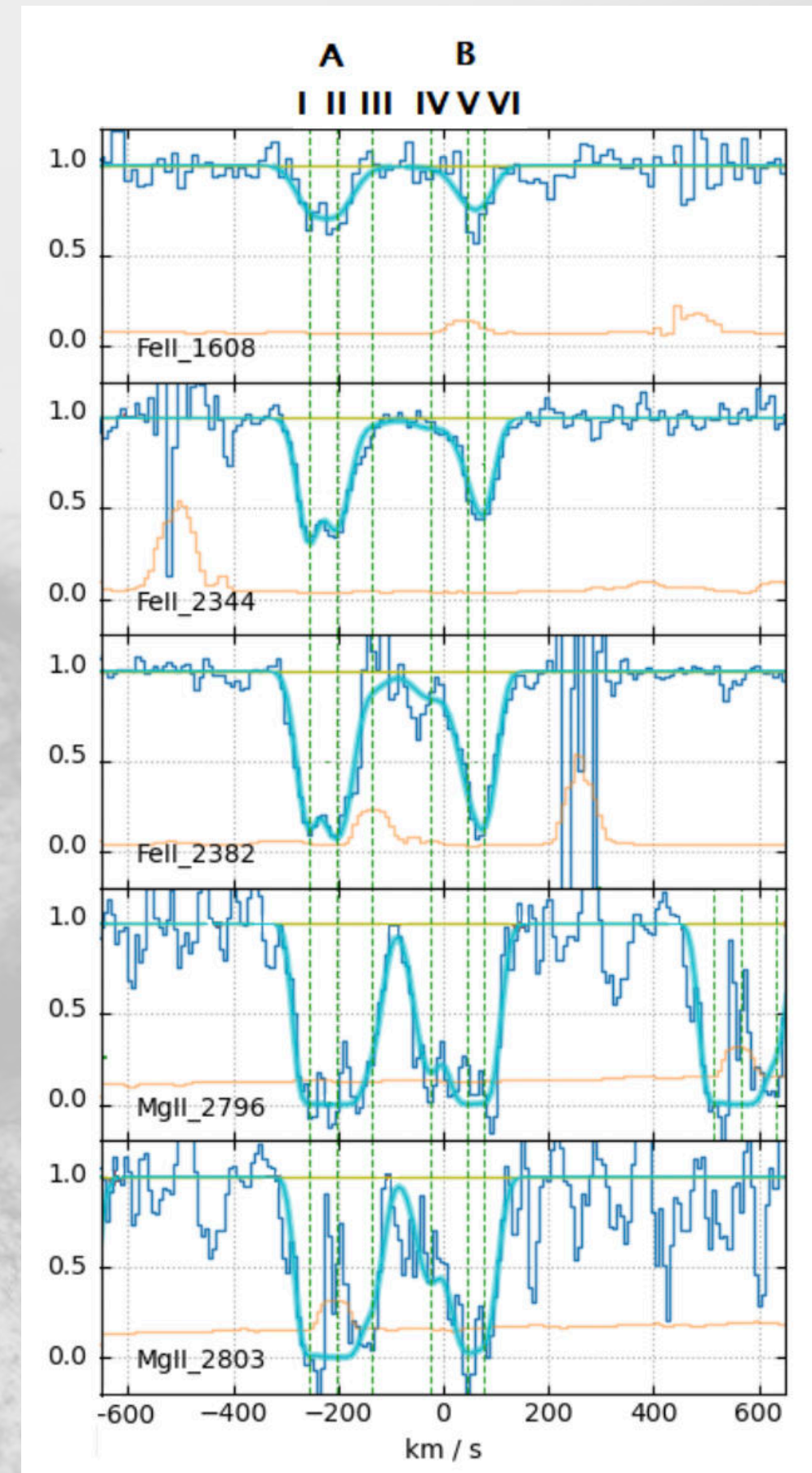
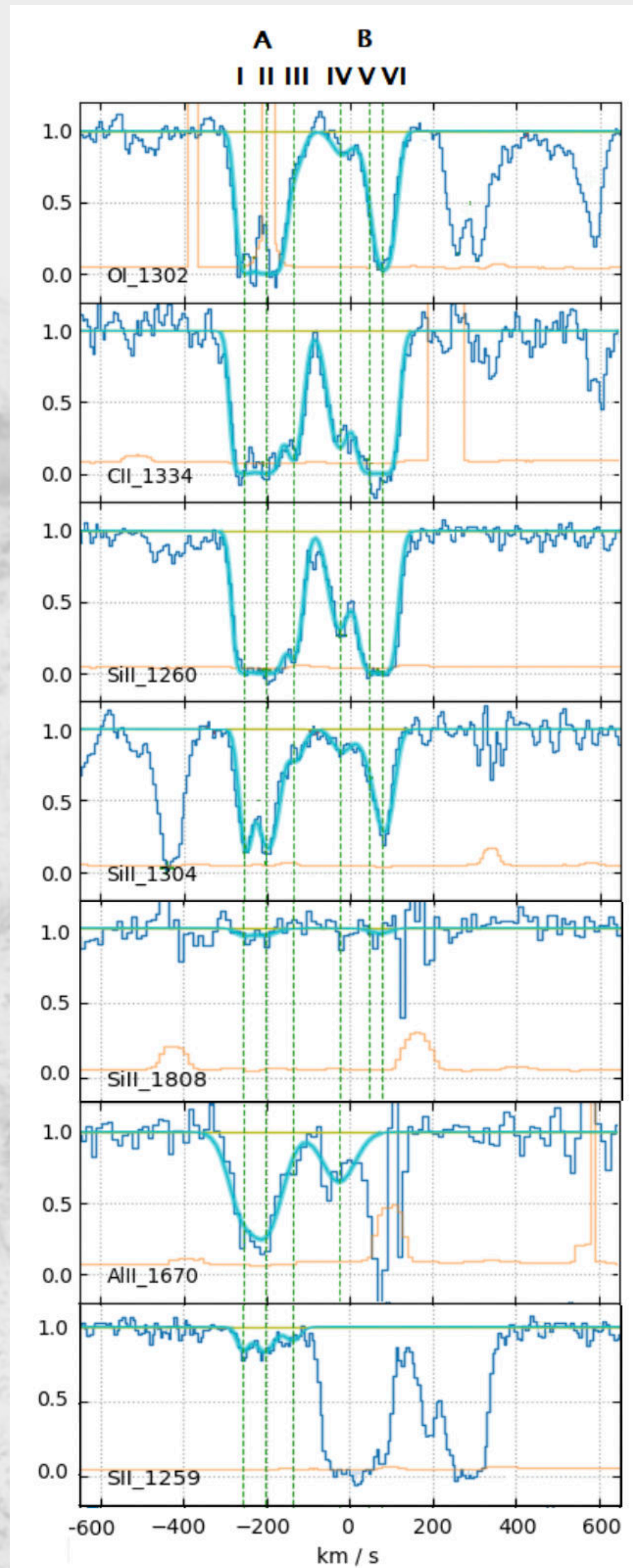
A. Saccardi, S.D. Vergani, A. De Cia et al. submitted



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GRB 210905A
Redshift $z=6.312$
 $t=0.875$ Gyr

ANALYSIS

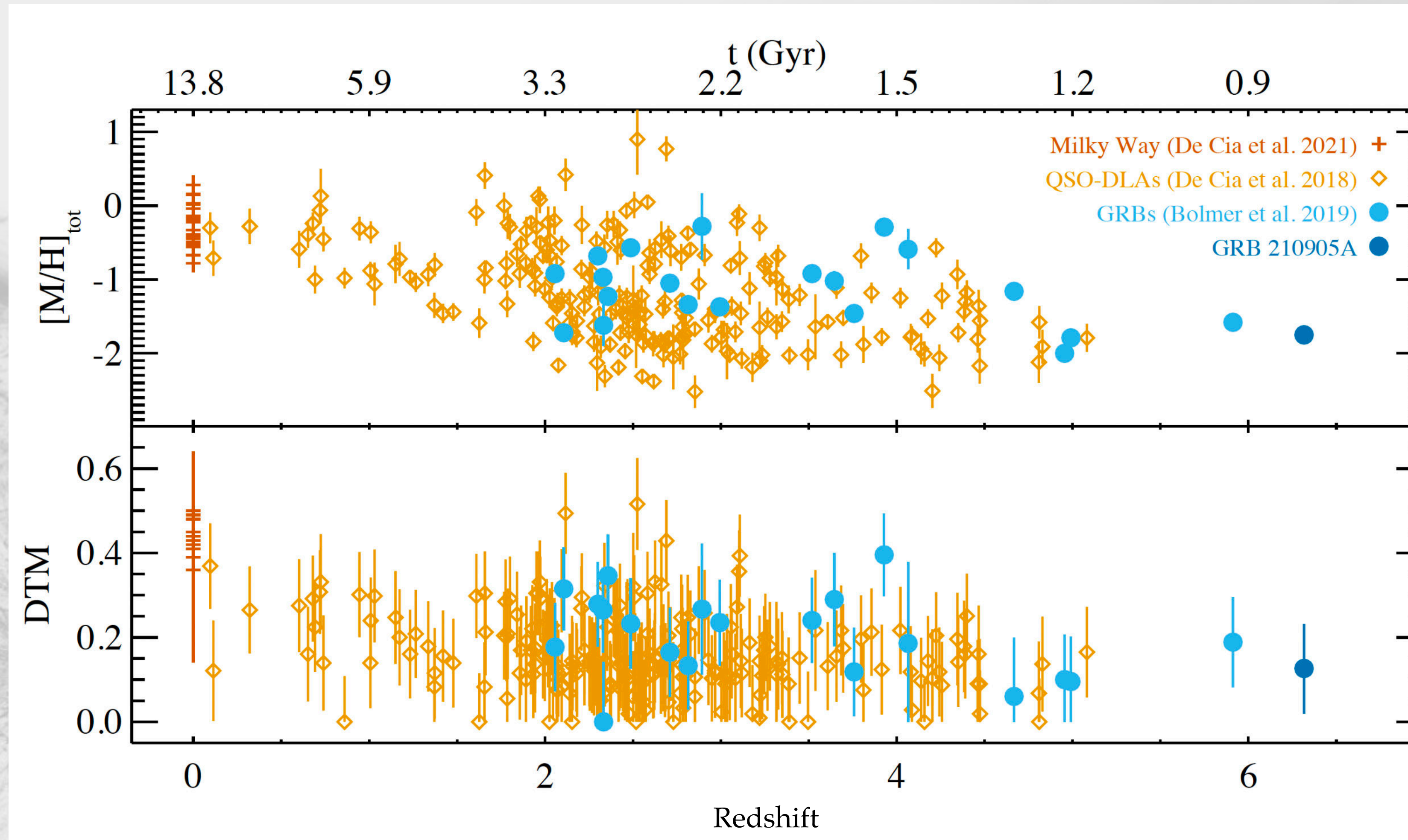


A. Saccardi, S.D. Vergani, A. De Cia et al. submitted

From the absorption properties:

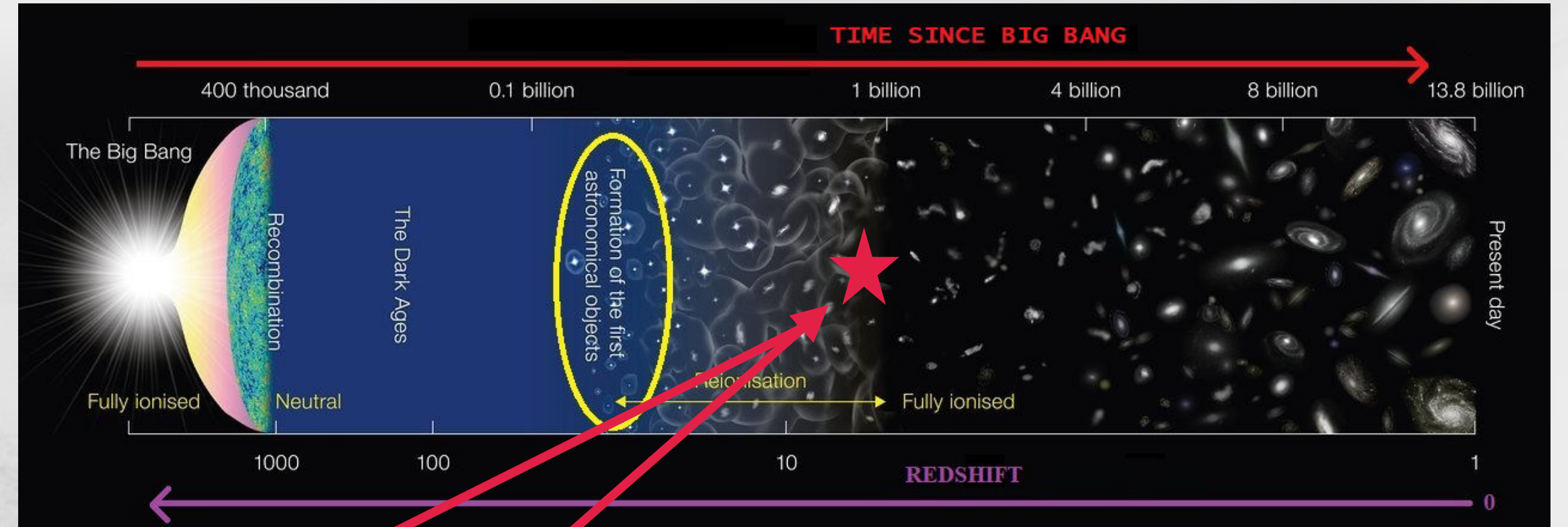
- ➔ *Metallicity and dust depletion*
- ➔ *The distance of the gas clouds*
- ➔ *Kinematic of the gas*
- ➔ *Chemical composition*

RESULTS

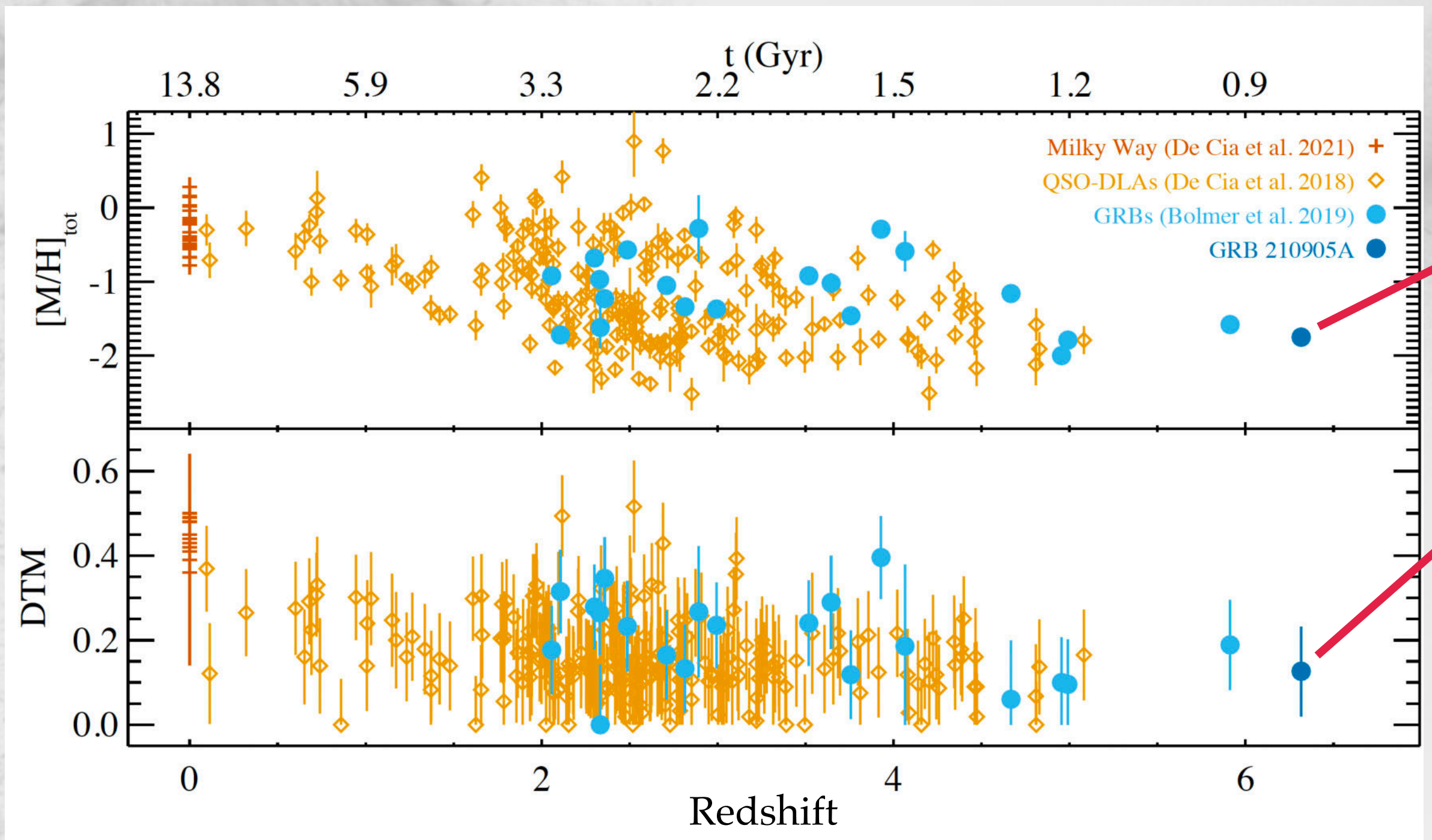


A. Saccardi, S.D. Vergani, A. De Cia et al. submitted

RESULTS



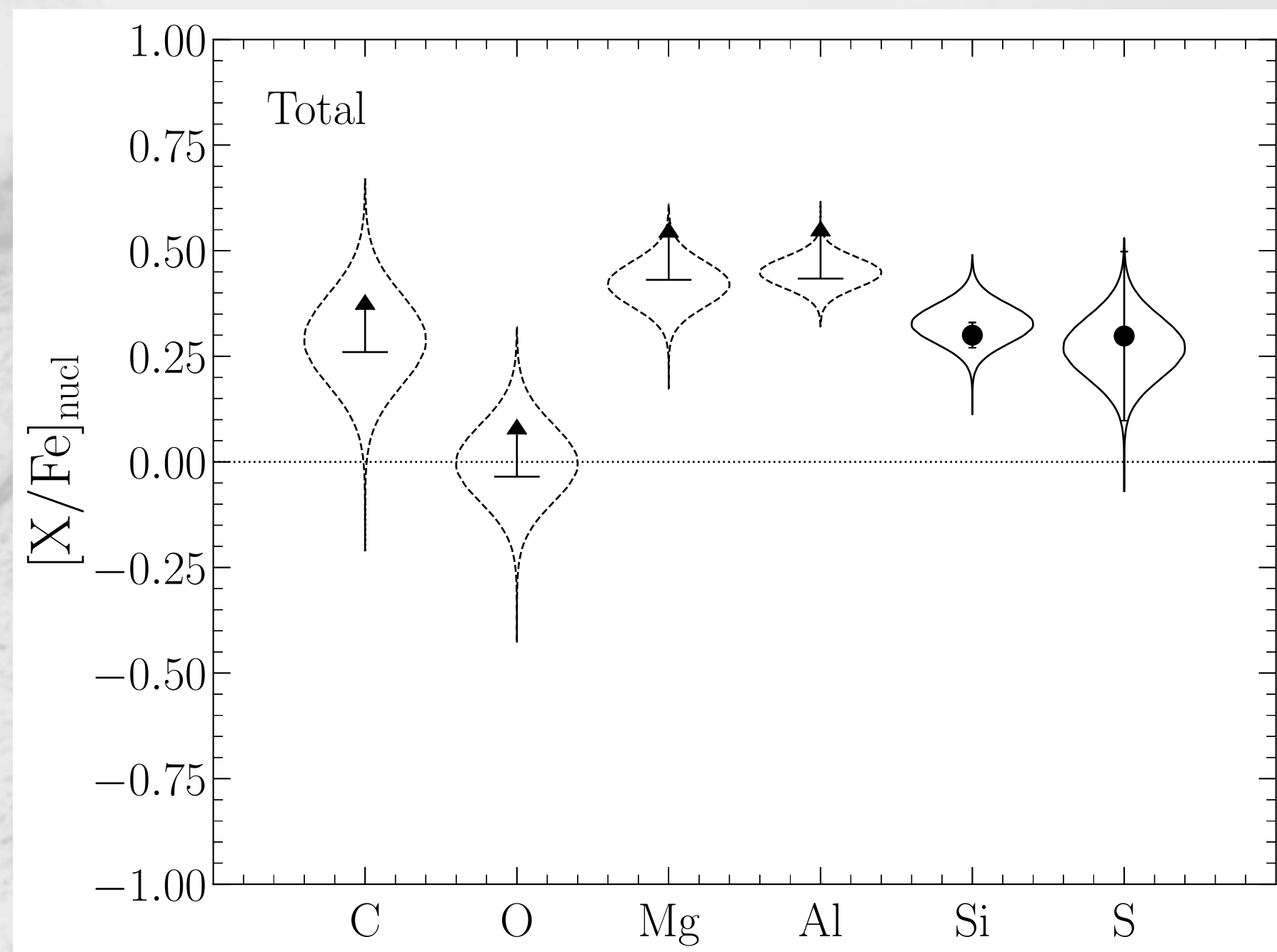
Credits: ESO



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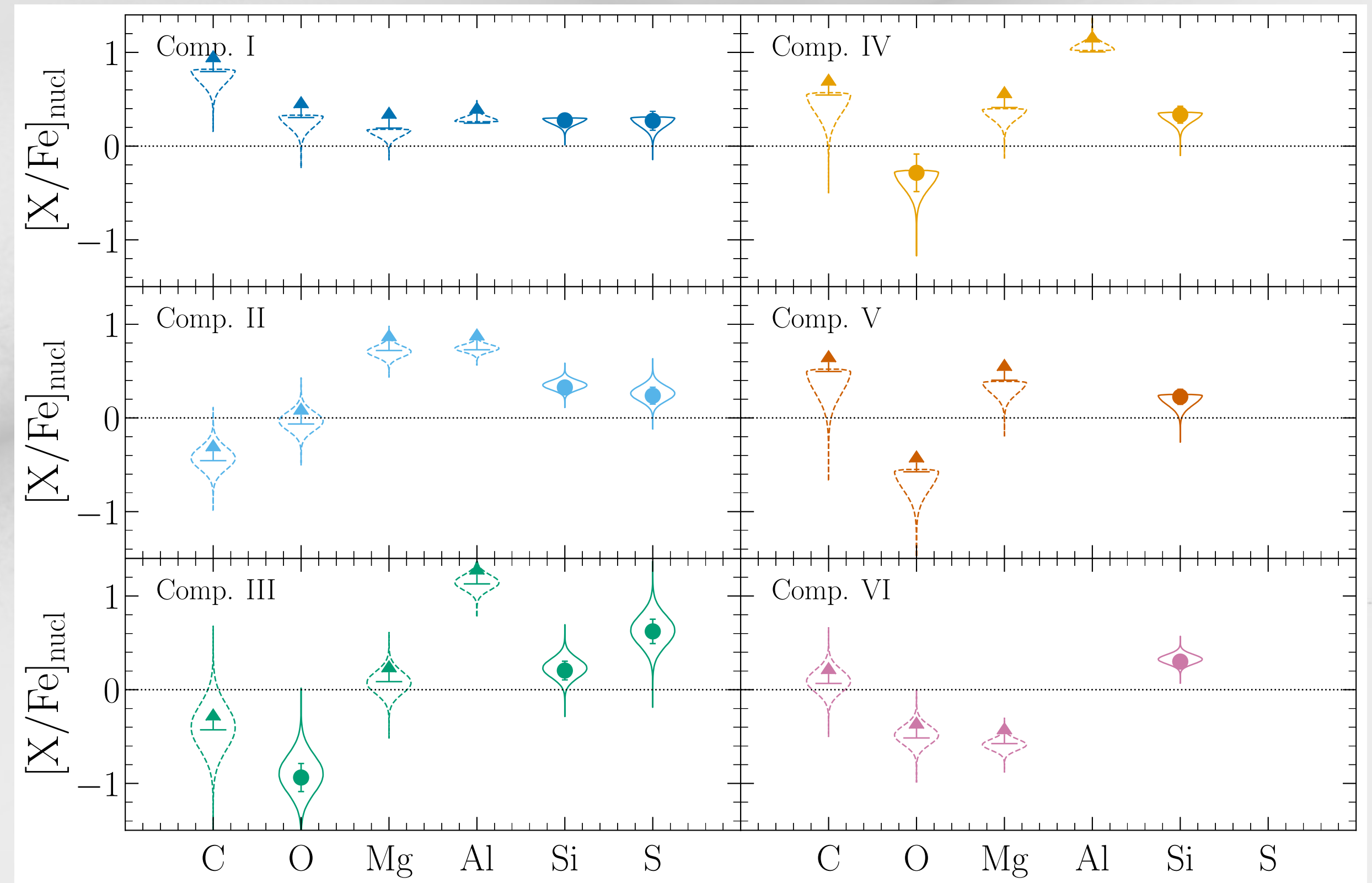
RESULTS

The overall host galaxy



A. Saccardi, S.D. Vergani, A. De Cia et al. submitted

Component-by-component

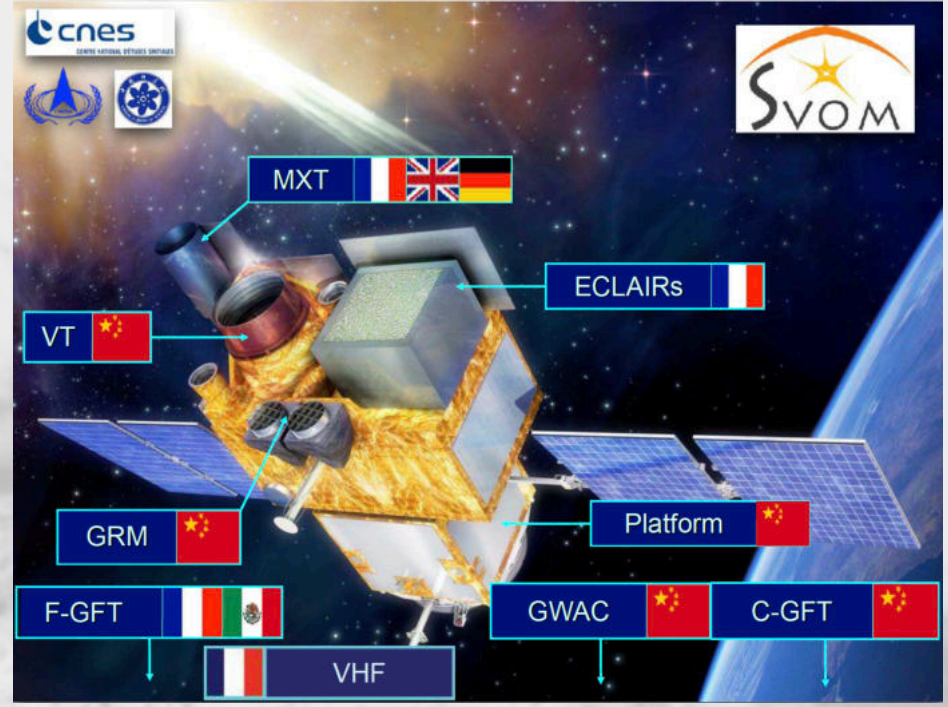


A. Saccardi, S.D. Vergani, A. De Cia et al. submitted

HIGH REDSHIFT GRBs

FUTURE OBSERVING FACILITIES

SVOM

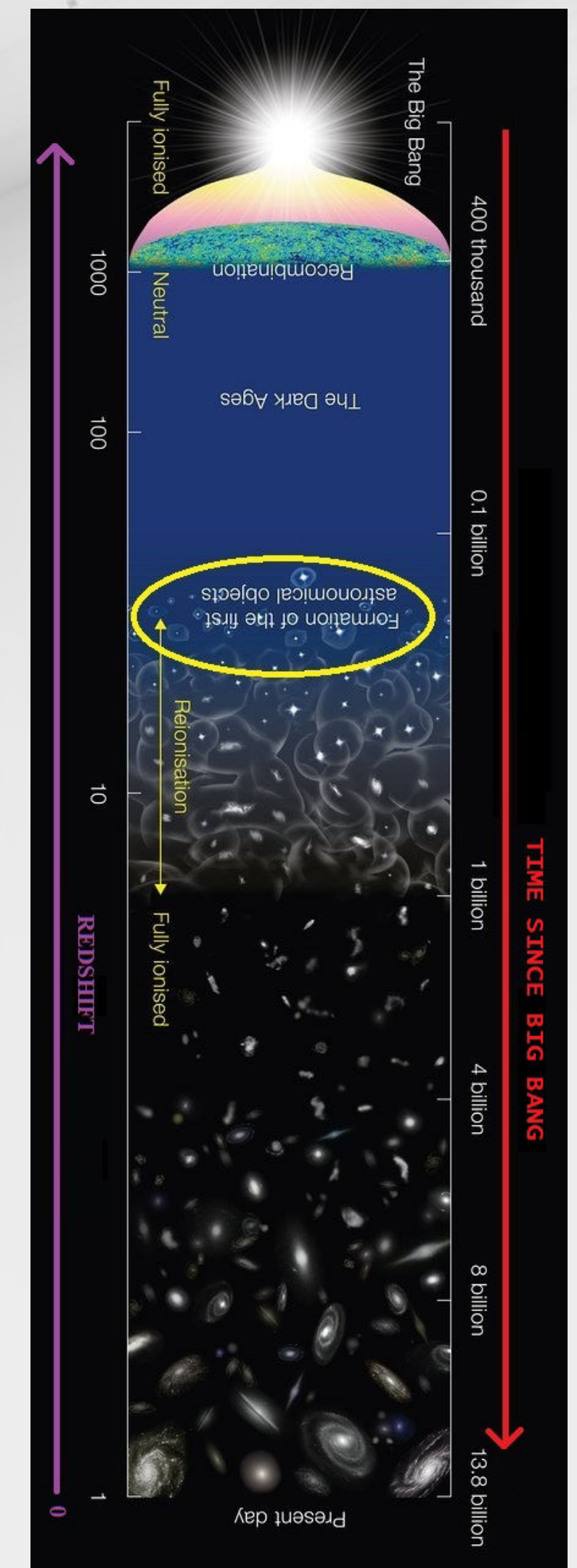
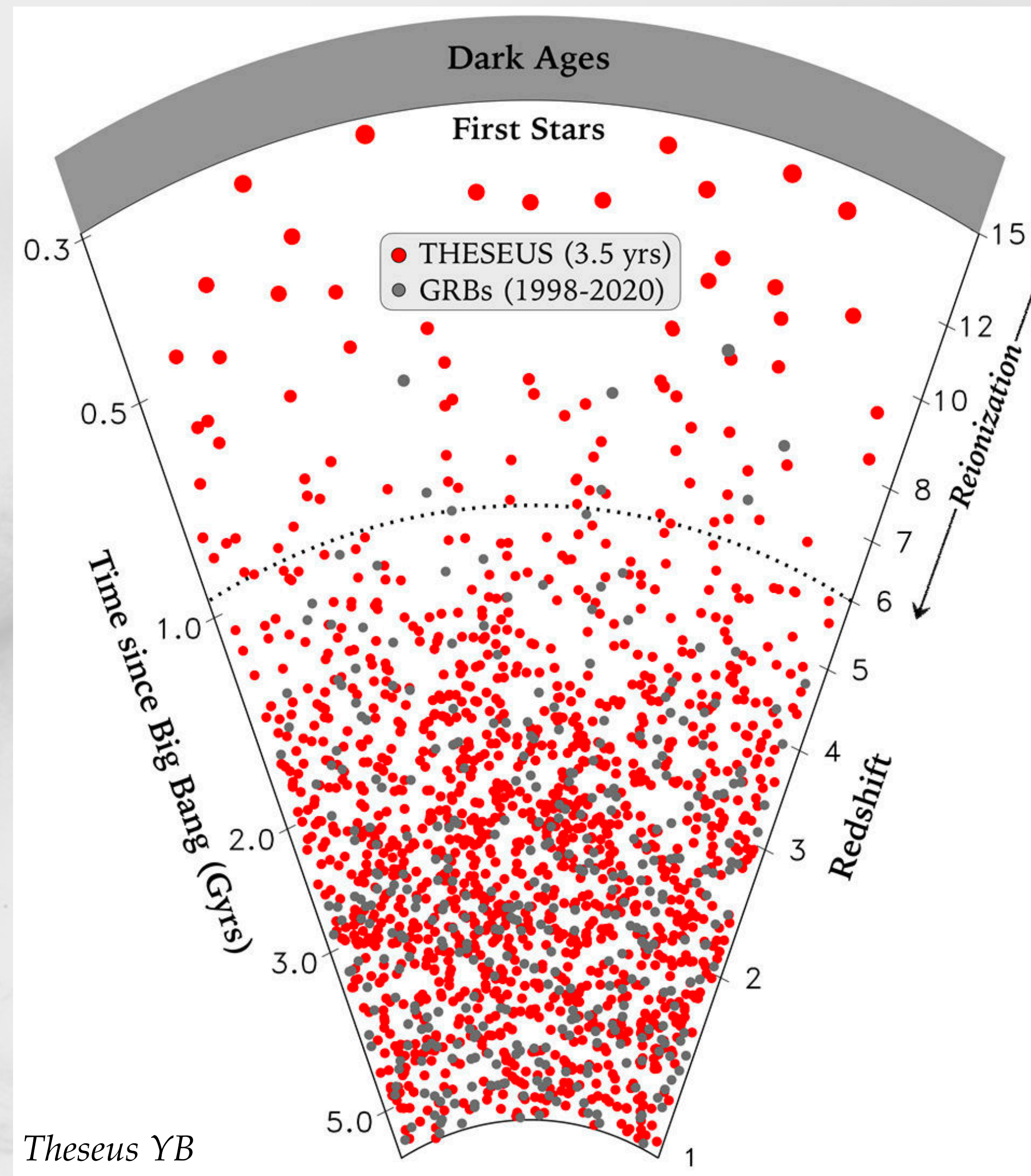


<https://www.svom.eu/>

THESEUS



<http://www.isdc.unige.ch/theseus>



Credits: ESO

- The formation and evolution of galaxies are key for the current extragalactic astrophysics**
 - Neutral gas dominates galaxies in the early Universe**
 - Bright background sources are needed to study the neutral gas**
 - GRBs are very powerful tools for this kind of studies**
- Thanks to GRB 210905A we were able to obtain unique and detailed information of the neutral gas and its chemical composition for a galaxy when the Universe was $t \sim 0.9$ Gyr**
 - The future is bright especially in France thanks to SVOM and we hope THESEUS**

THANKS FOR YOUR ATTENTION

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Rencontres des Jeunes Physicien·ne·s
2 November 2022



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