# STUDY OF THE TENSOR FORCE CONTRIBUTION IN THE OXYGEN ISOTOPES USING QFS REACTIONS 

Barrière Antoine, Mozumdar Nikhil, Sorlin Olivier<br>and the R3B Collaboration



## Study of the evolution of proton $p_{3 / 2}-p_{1 / 2} \mathrm{SO}$ splitting between ${ }^{16} \mathrm{O}$ and ${ }^{22} \mathrm{O}$

In the shell model framework, the nuclear interaction can be divided in several parts:

- Central
- Spin-orbit
- Tensor

A review[1] and global fit based mostly on stable nuclei data reproduces the SO splitting with the following function :

$$
\Delta_{S O}=\frac{24.5}{n}(l+1 / 2) A^{-0.597}
$$

The factor $\sim \mathrm{A}^{-2 / 3}$ is the fingerprint of the role of the SO interaction in this trend (surface term).


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Estimation of the proton gap $0 p_{1 / 2}-0 p_{3 / 2}$ :
in ${ }^{16} \mathrm{O}-->7.02 \mathrm{MeV}$
in ${ }^{22} \mathrm{O}$--> 5.81 MeV
Deviations from this trend may be due to the tensor force contribution


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## Role of tensor force in the $\mathbf{O}$ isotopic chain

Tensor interaction [3]:


The tensor force should then reduce the spin-orbit splitting $0 p_{1 / 2}-O p_{3 / 2}$ (proton) in the $O$ isotopic chain, when the $0 d_{5 / 2}$ (neutron) is filled.

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## Experimental setup

$R^{3} B$ setup at GSI
Cocktail beam from FRS (2 beam settings)
Beam energy: ~550MeV/A

QFS reaction e.g. ( $p, p n$ ) and ( $p, 2 p$ ) in LiH
Complete measurement in inverse kinematics:


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Complete measurement in inverse kinematics:


LOS (t)


MusLi ( $\Delta \mathrm{E}$ ) + MWPC ( $\mathrm{x}, \mathrm{y}$ )


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Complete measurement in inverse kinematics:

- Incoming nucleus;
- Light particles and gammas emitted from the target;


CALIFA $(\theta, \phi, \Delta E)$ Array of Csl crystals


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Complete measurement in inverse kinematics:

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Complete measurement in inverse kinematics:

- Incoming nucleus;
- Light particles and gammas emitted from the target;
- Outgoing fragment;
- Neutrons.


NeuLAND ( $x, y, z, t$ ) with 26 plans


## Identification of the incoming nuclei

2 settings were used :


$\Delta E$ (MusLi) --> Z
(Bethe-Bloch)

$$
A / Q=\frac{B \rho_{0} \cdot\left(1-\operatorname{Pos} S 2 / D_{S 2-C C}\right)}{3.10716 \cdot \beta \cdot \gamma}
$$

Identification of the N isotopes from the ${ }^{22} \mathrm{O}(p, 2 p)^{21} \mathrm{~N}$ reaction


## Population of bound states in the ${ }^{22} \mathrm{O}(p, 2 p)^{21} \mathrm{~N}$ reaction




CALIFA (Barrel + Endcap): 1504 crystals, including 480 with 2 electronic gains (Endcap)

## Population of bound states in the ${ }^{22} \mathrm{O}(p, 2 p)^{21} \mathrm{~N}$ reaction




$$
\begin{array}{ll}
\begin{array}{l}
\text { Resolution: } \\
\sigma=40 \mathrm{keV} \text { at } \mathrm{E}_{\nu}=1.170 \mathrm{MeV}
\end{array} & \varepsilon_{\gamma}(1.170 \mathrm{MeV})=26 \% \\
=>\text { FWHM } / \mathrm{E}_{\gamma}=8 \%
\end{array} \quad \begin{aligned}
& \text { Results }: \frac{N\left(3 / 2_{1}^{-}\right)}{N\left(1 / 2_{G S}^{-}\right)}=\frac{N\left(3 / 2_{1}^{-}\right)}{N_{\text {incl }}-N\left(3 / 2_{1}^{-}\right)}=19.0(31) \%
\end{aligned}
$$

=> Among the bound states, the $1 / 2^{-}{ }_{\text {Gs }}$ is the most populated by the $(p, 2 p)$ reaction.

## NeuLAND and the invariant mass method

Alignment of the time difference of the two PMTs for each bar, using cosmic rays (muons) and a tracking algorithm.


Portion of NeuLAND crossed by a cosmic ray particle.


Use of the invariant mass method:
$M_{i n v}=\sqrt{\left(\sum_{i=0}^{N} E_{i}\right)^{2}-\left(\sum_{i=0}^{N} p_{i}\right)^{2}}$
with the energy and momentum of the fragment and neutron(s)

$$
E_{r e l}=M_{i n v}-\sum_{i=0}^{N} m_{i}
$$



Population of neutron unbound states in the ${ }^{22} \mathrm{O}(p, 2 p)^{21} \mathrm{~N}$ reaction

$S_{2 n}=6.75 \mathrm{MeV}$


Fit of the $E_{\text {rel }}$ distribution requires to take into account:

- Time resolution of NeuLAND --> already done;
- Effective resolution of the beta(fragment)
--> straggling + detector resolution --> still ongoing.


## Conclusions and perspectives



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First estimation of the proton gap in ${ }^{22} \mathrm{O}$ :

$$
\begin{aligned}
& <0 p_{3 / 2}>_{s . p .} \geq \frac{E^{*}\left(3 / 2_{1}^{-}\right) N\left(3 / 2_{1}^{-}\right)+N_{1 n} E_{1 n}^{*}}{N\left(3 / 2_{1}^{-}\right)+N_{1 n}} \\
& <0 p_{3 / 2}>_{\text {s.p. }}-<0 p_{1 / 2}>_{s . p .} \geq 5.46 \mathrm{MeV}
\end{aligned}
$$



## Conclusions and perspectives



## Perspectives:

- Study of the ${ }^{21} \mathrm{O}$ momentum via ${ }^{22} \mathrm{O}(\mathrm{p}, \mathrm{pn})$
( 6 neutrons in the neutron orbital $0 d_{5 / 2}$ ?);
- 1n spectroscopy, w. gamma-neutron coincidences;
- Analysis of unbound states decaying by $2 n$ emission.



## Conclusions and perspectives



## Backup Slides

## Tracking around the target



FOOTs mapping : 4 on each side +4 in-beam

## NeuLAND calibration

PMTs time for one bar -> position within the bar + ToF (target-bar)

1) Alignment of the time difference of the two PMTs for each bar, using cosmic rays and a tracking algorithm


Time difference left/right or bottom-top PMTs vs Bar Id, for on-spill events


Portion of NeuLAND crossed by a cosmic ray particle.

For each bar, we can define:
Time diff : T1-T2
-> Effective speed of light
Time synch : T1 + T2
-> Global offset
(1) $=>$ diff correction
(2) $=>$ synch correction
(3) $=>$ both of them

## NeuLAND calibration

PMTs time for one bar -> position within the bar + ToF (target-bar)

1) Alignment of the time difference of the two PMTs for each bar, using cosmic rays and a tracking algorithm


Time difference left/right or bottom-top PMTs vs Bar Id, for on-spill events
2) Fine tuning with the gamma peak


## Califa and the QFS reactions




## Modifications of the gamma simulations



Crystal Id vs energy (simulation): 1274 keV peak


Use of simulations in gamma analysis:
-> To get the energy of the transitions
-> To obtain the associated cross section

