

A painting depicting a large crowd of people in 19th-century attire standing on a balcony or terrace. They are looking out over a landscape with green hills and a body of water. In the sky, a bright comet with a long tail is visible, along with a rainbow. The scene is set at night or dusk, with a dark sky and a few stars.

Through the looking glass of the Standard Model with radioactive ion beams

Leendert Hayen

Colloque GANIL

September 28 2023

LPC Caen

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Introduction

CP violation tests

Exotic current searches

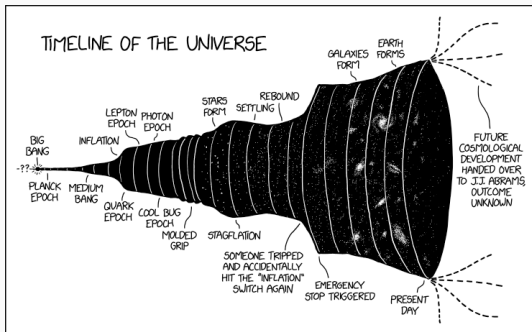
Superconducting sensors for dark matter & exotic currents

Summary & Outlook

Fundamental symmetries: *Precision measurements and symmetry tests of the Standard Model and Beyond*

Fundamental symmetries & RIBs

Fundamental symmetries: *Precision measurements and symmetry tests of the Standard Model and Beyond*

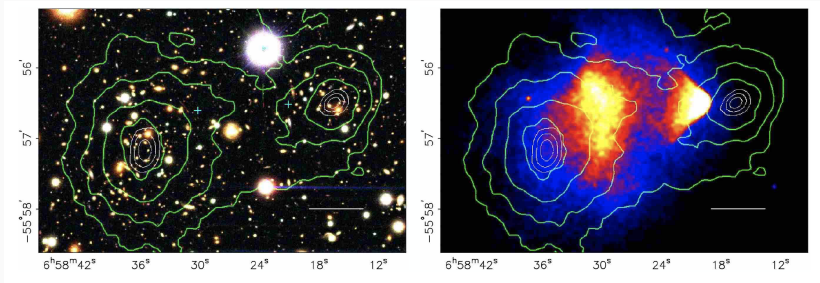


Extremely rich field with connections to

- Nature of neutrino's
- Dark matter
- Big bang nucleosynthesis
- Cosmology

Nature of dark matter

Visible matter \neq gravitational mass



Good indications we're missing 80% of matter

Nature of neutrinos and mass mechanism

Neutrinos are all around, but elusive!



We don't know

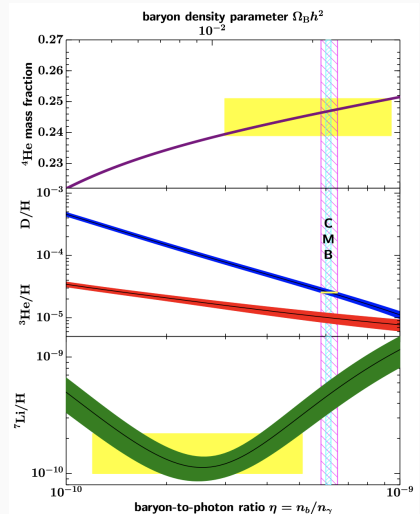
Image by Sandbox Studio

- their mass
- how to give them mass
- own antiparticle?
- CP violation?

Baryogenesis

Why are we even here?

$$\eta = \frac{n_B - n_{\bar{B}}}{n_\gamma} = \mathcal{O}(10^{-10})$$



Baryogenesis

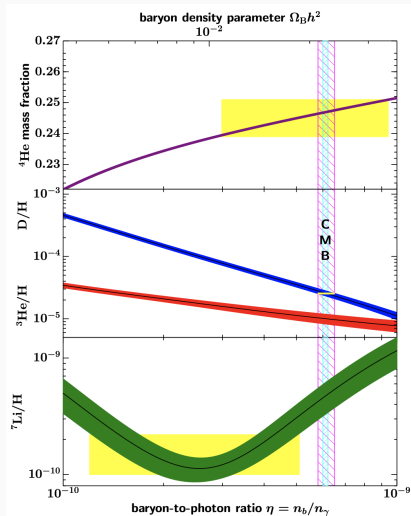
Why are we even here?

$$\eta = \frac{n_B - n_{\bar{B}}}{n_\gamma} = \mathcal{O}(10^{-10})$$

Naively,

Standard Model predicts **10-11 orders of magnitude smaller**

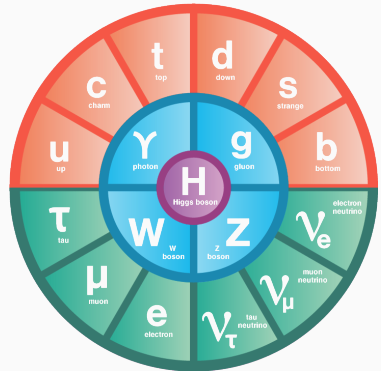
In more detail, it's simply impossible (Higgs is too heavy)



Meet the Standard Model

Three out of four
fundamental forces (no gravity):

Standard Model

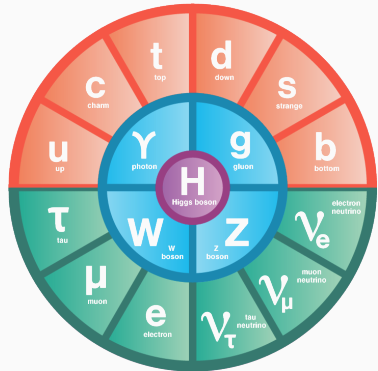


Meet the Standard Model

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Standard Model

18 free parameters



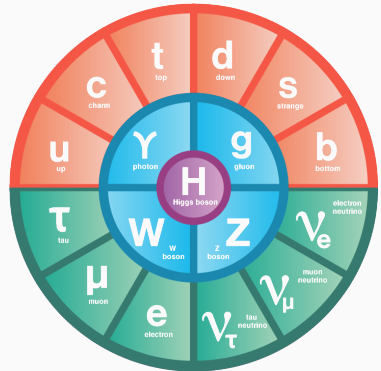
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Great (annoyingly so), consistent
with constraints at $\sim 10^{0-2}$ TeV



Meet the Standard Model

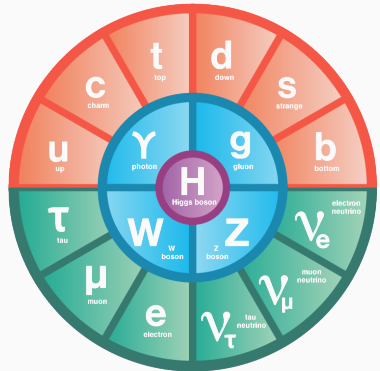
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Open questions: dark matter,
gravity, neutrino masses, ...



Introduction: Standard Model

What to do?

SM tests @ low energy: sensitive to **off-shell** exotic physics
(footprints rather than actual beast)

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Besides precision QED ($a_{e,\mu}, r_p, \dots$), weak interactions probe

- (C)P violation
- CKM unitarity
- Lorentz structure

Introduction: Standard Model

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All of these can be probed using **(nuclear) β decay** with RIBs!

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- (C)P violation **Today**
- CKM unitarity
- Lorentz structure **Today**

All of these can be probed using **(nuclear) β decay** with RIBs!

β decay introduction

Require Lorentz invariance, get Lee-Yang Hamiltonian (SM: $V - A$)

$$\mathcal{H}_\beta = \sum_{i=V,A,S,P,T} \bar{p} \mathcal{O}_i n \times \bar{e} \mathcal{O}_i (C_i + C_i' \gamma^5) \nu_e$$

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Differential decay rate

$$\Gamma \propto 1 + b \frac{m_e}{E_e} + a_{\beta\nu} \frac{\mathbf{p}_e \cdot \mathbf{p}_\nu}{E_e E_\nu} + \frac{\langle \mathbf{J} \rangle}{J} \cdot \left[A \frac{\mathbf{p}_e}{E_e} + B \frac{\mathbf{p}_\nu}{E_\nu} + D \frac{\mathbf{p}_e \times \mathbf{p}_\nu}{E_e E_\nu} \right]$$

+ higher order



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Baryon asymmetry

Already discussed baryon asymmetry:

$$\eta = \frac{n_B - n_{\bar{B}}}{n_\gamma} = \mathcal{O}(10^{-10})$$

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$$\eta = \frac{n_B - n_{\bar{B}}}{n_\gamma} = \mathcal{O}(10^{-10})$$

Baryon asymmetry can arise **dynamically**
on three conditions (Sakharov 1967)

- B-number violation
- CP violation
- Departure from equilibrium

First two present in SM but **way too little**,
Higgs too heavy for last



CP violation searches with MORA

$$\Gamma \propto 1 + b \frac{m_e}{E_e} + a_{\beta\nu} \frac{\mathbf{p}_e \cdot \mathbf{p}_\nu}{E_e E_\nu} + \frac{\langle \mathbf{J} \rangle}{J} \cdot \left[A \frac{\mathbf{p}_e}{E_e} + B \frac{\mathbf{p}_\nu}{E_\nu} + D \frac{\mathbf{p}_e \times \mathbf{p}_\nu}{E_e E_\nu} \right]$$

D term is T -odd, non-zero when $\text{Im}(C_i) \neq 0$

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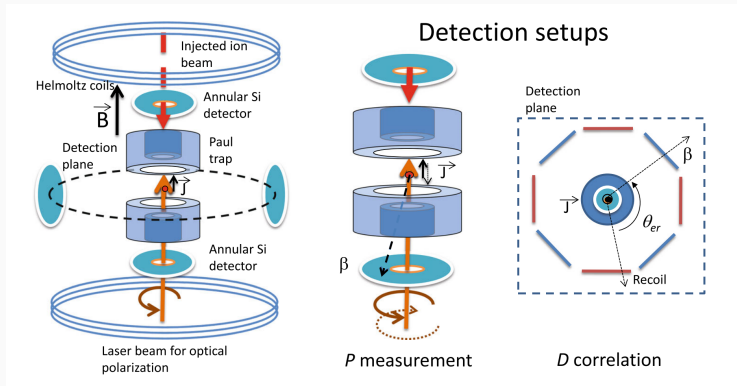


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Fierz interference: Spectrum shape

Allowing exotic interactions (C_S, C_T) modifies β spectrum

$$P(E_e) = \text{Standard Model} \times \left(1 + b_F \frac{m_e}{E_e} \right)$$

Fierz interference: Spectrum shape

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Fierz interference

$$b_F \propto \text{Re} \left[|M_F|^2 (C_S C_V^* + C_S' C_V'^*) + |M_{GT}|^2 (C_T C_A^* + C_T' C_A'^*) \right]$$

Fierz interference: Spectrum shape

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Promising to directly measure spectra, but also tricky

- Detector linearity, energy losses, pile-up, . . .
- Theory spectrum calculation

See also talks by M. Kanafani, M. Lejoubioux

Naviliat-Cuncic, Gonzalez-Alonso PRC 94, 035503

LH *et al.*, RMP 90 015008

Mirror experimental status

Community is investi(gati)ng in different ideas



with new spectroscopy techniques & traps

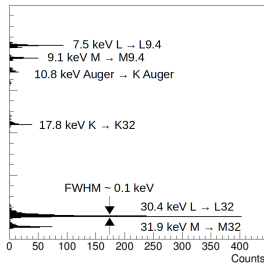
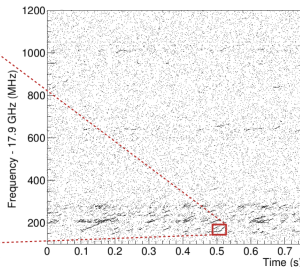
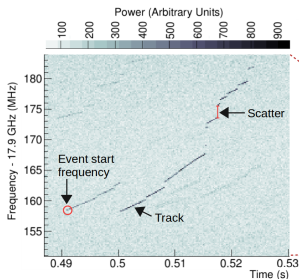
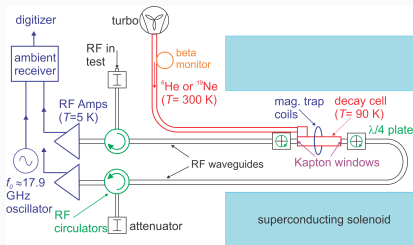
with additional great progress in the $A = 8$ system

New technology: CRES

Cyclotron Radiation Emission Spectroscopy

$$f = \frac{|q|}{2\pi} \frac{B}{m_e + E_{kin}}$$

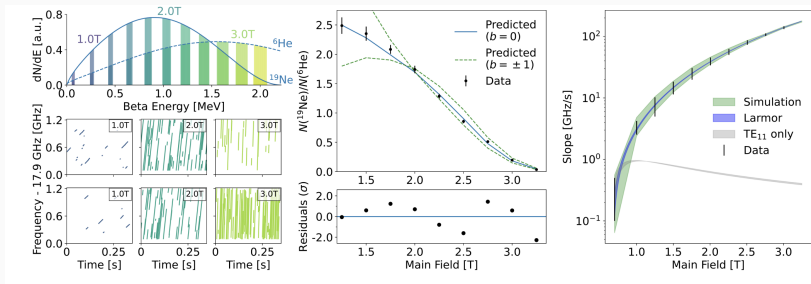
${}^6\text{He}$ and ${}^{19}\text{Ne}$



Physical Review Letters 131 (2023), 082502

New technology: CRES

Use ratio method: ${}^6\text{He}$ and ${}^{19}\text{Ne}$ have opposite b_F sign



Ratio means many systematic effects cancel to first order

Physical Review Letters 131 (2023), 082502

Investigating coupling to ion trap

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Spectroscopy experiments currently focused on β (e^-/e^+), but extremely demanding

- Detector linearity, energy losses, pile-up, . . .
- Theory spectrum calculation

Naviliat-Cuncic, Gonzalez-Alonso PRC 94, 035503; LH *et al.*, RMP 90 015008

β recoil spectroscopy

Spectroscopy experiments currently focused on β (e^-/e^+), but extremely demanding

- Detector linearity, energy losses, pile-up, . . .
- Theory spectrum calculation

Naviliat-Cuncic, Gonzalez-Alonso PRC 94, 035503; LH *et al.*, RMP 90 015008

Instead, **recoil** spectroscopy has interesting features

- Compressed energy range (<keV instead of \sim MeV)
- Electron capture gives single recoil peak
- Sensitive to β - ν correlation for β^\pm decay

Meet superconducting tunnel junctions

- Two electrodes separated by a thin insulating tunnel barrier
- Superconducting energy gap Δ is of order $\sim \text{meV}$
 → High Energy Resolution ($\sim 1 \text{ eV}$)
- Timing resolution on the order of $10 \mu\text{s}$, making it among the fastest high-resolution quantum sensors available
 → "High" Rate (10^4 s^{-1} per pixel)

← Ideal for RIB experiments at ISAC

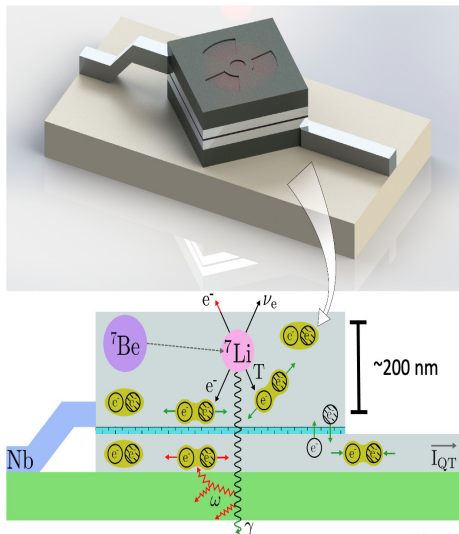
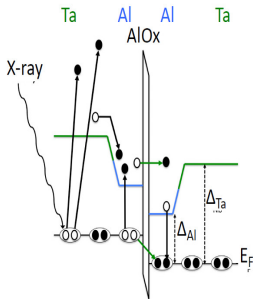
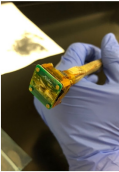
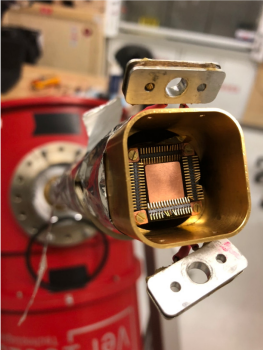
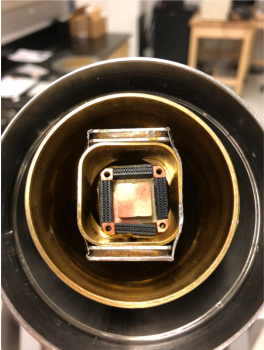
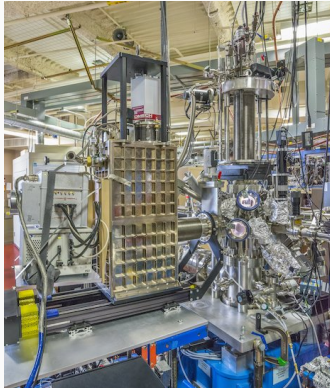


Image courtesy S. Fretwell (Mines)

Superconducting tunnel junctions



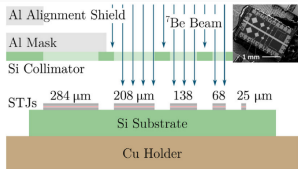


^7Be electron capture

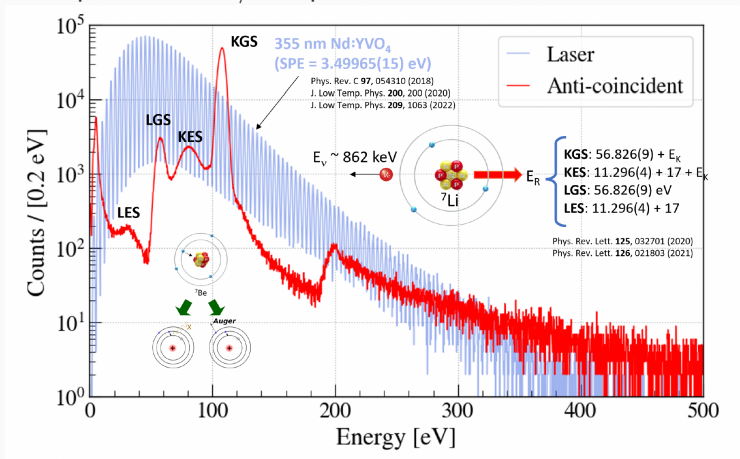
- Responsible for ^7Li creation in stars
- Essential contribution to solar neutrino spectrum

Measurement campaign

1. Implantation at ISAC (TRIUMF)
2. Ship to LLNL
3. Cool down and measure



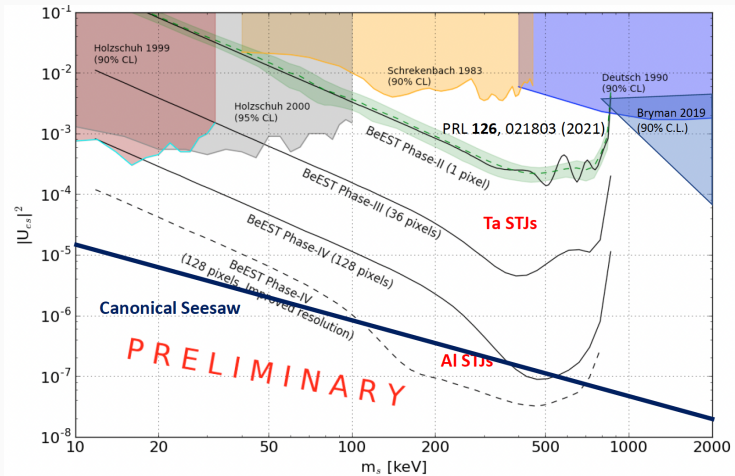
Most precise ${}^7\text{Be}$ L/K capture measurement



Constraints on MeV-scale sterile neutrino's

PRL 126 (2021) 021803; PRL 125 (2020), 032701

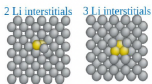
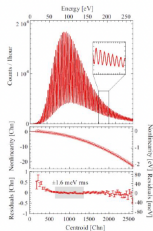
BeEST sterile neutrino limits



Phase I already into new territory

BeEST up to phase IV

Phases of the BeEST Experiment



Phase-III

Scaling to 36- and 112-Pixel Arrays of Ta-Based STJs

Phase-IV

Operation of 128-Pixel Arrays of Al-Based STJs in Dilution Refrigerator



Phase-II

First Limits and Precision Device Characterization



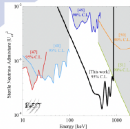
Phase-I

Proof of Concept

2018

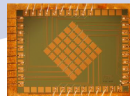


2020



2022

2024



2026



Can we do the same thing at radioactive ion beam facilities?

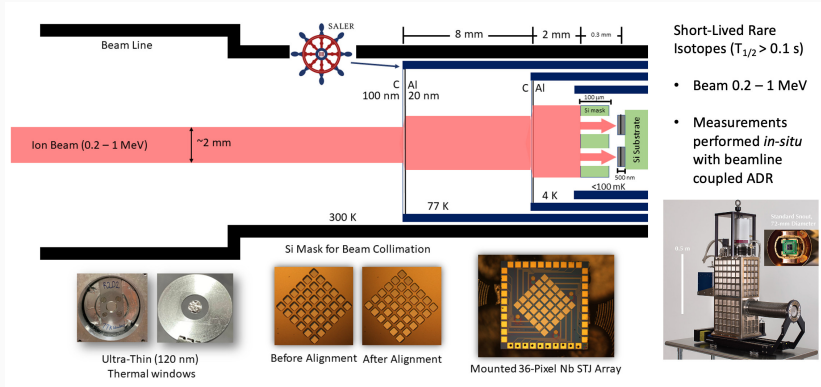
Introducing



Superconducting Array for Low Energy Radiation

Superconducting tunnel junctions

Concept to couple to beam line



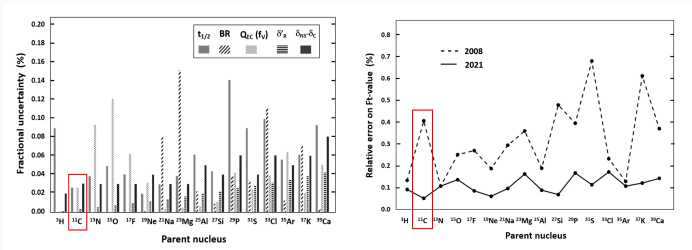
Portability allows easy installation (ISAC, SPIRAL2, FRIB, ISOLDE, ...)

Commissioning in Fall 2023 at FRIB

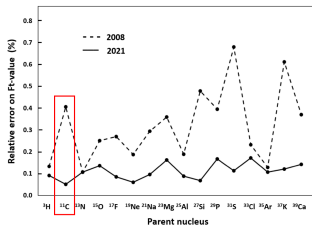
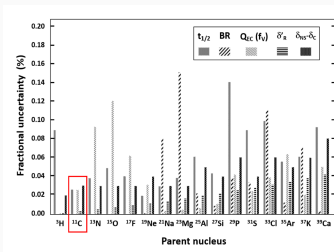


Detector array will arrive late October!

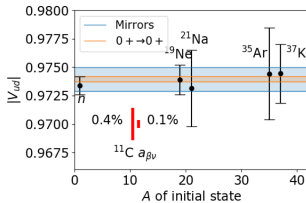
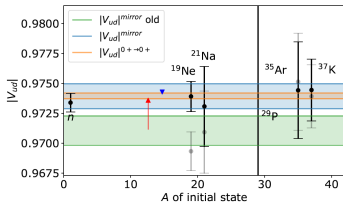
^{11}C first physics target (long $t_{1/2}$, unreachable with traps!)



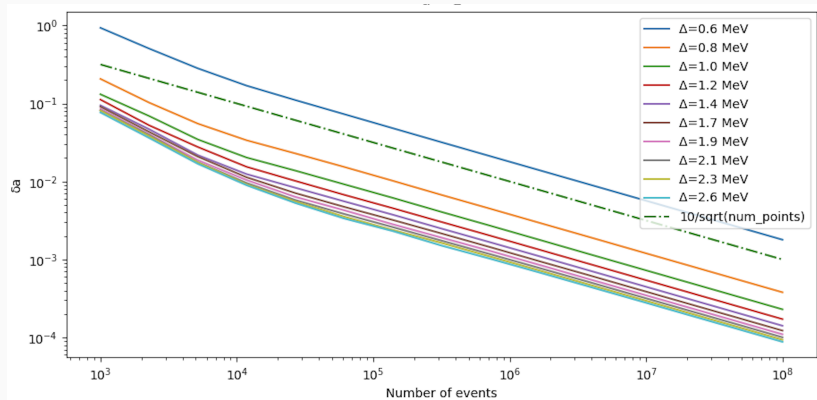
^{11}C first physics target (long $t_{1/2}$, unreachable with traps!)



Excellent V_{ud} sensitivity



Statistical sensitivity (Preliminary!)



1E8 events (~ 1 day) statistical sensitivity $\delta V_{ud} \approx 3.7 \times 10^{-4}$

(Full superallowed data set is 3.2×10^{-4})

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MORA collaborators

MORA collaborators:



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R.P.De Groot
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X.Fléchard
E.Liénard
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M.Kowalska
G. Neyens



M. González-Alonso



A.Falkowski
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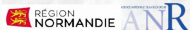
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V.Virtanen



BeEST & SALER



NC STATE UNIVERSITY



NC STATE UNIVERSITY



U.S. DEPARTMENT OF
ENERGY

Office of
Science

Summary & Outlook

Fundamental symmetries lives at the interface, connections to many different fields

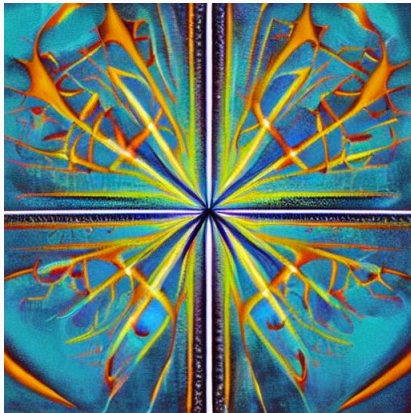
Nuclear β decay searches provide **crucial input** through variety of experiments, haven't even talked about

- Neutrino mass measurements
- Dark matter searches
- Big bang nucleosynthesis
- ...

Quantum sensing in β spectroscopy incoming, highly promising!

Thank you

Thank you!



β decay symmetries according to Stable Diffusion

Introduction: Weak interaction & CKM matrix

Cabibbo-Kobayashi-Maskawa matrix relates weak and mass eigenstates

$$\begin{pmatrix} d \\ s \\ b \end{pmatrix}_w = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}_m$$

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Unitarity requires

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1$$

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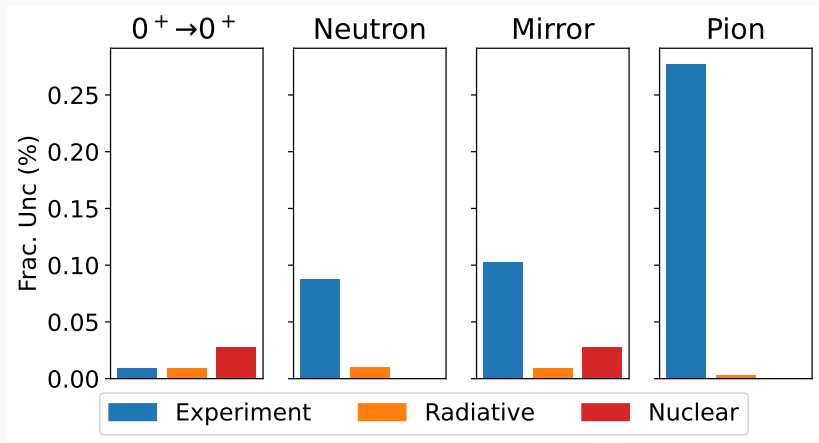
$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1$$

(nuclear) β decay, meson decay (π , K), $|V_{ub}|^2 \sim 10^{-5}$

Violations are sensitive to **TeV scale** new physics!

CKM unitarity: V_{ud} precision

Four (\sim)competitive channels of extracting V_{ud}



Status of $0^+ \rightarrow 0^+$ **great nuclear structure triumph**

2018-2020 reanalysis nuclear structure current bottleneck

CKM unitarity: Current status

Signs of non-unitarity at few σ level...

Disagreement between $K/2$ and $K/3$ $|V_{us}|$ 'Cabibbo angle anomaly'

