# Cross-shell interactions at the N = 28 shell closure through ${}^{47}K(d,p\gamma)$ and ${}^{47}K(d,t\gamma)$ with MUGAST+AGATA+VAMOS

Charlie J. Paxman (University of Surrey) + e793s collaboration





Far from stability, magic numbers change due to relative shifting of orbital energies [1].

[1] T. Otsuka *et al.* Rev. Mod. Phys. **92**, 015002 (2020)

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48 <b>SC</b>	<sup>49</sup> Sc	50SC							
47 <b>Ca</b> 4.536 d		<sup>49</sup> Ca 8.718 m	<sup>50</sup> Ca	<sup>51</sup> Ca	<sup>52</sup> Ca	53 <b>Ca</b>	54 <b>Ca</b> 90 ms	<sup>55</sup> Ca <sub>22 ms</sub>	Z=20
46 <b>K</b> 96.3 s	47K 17.385	48 <b>K</b> 6.83 s	49 <b>K</b> 1.26 s	50 <b>K</b> 472 ms	<sup>51</sup> K 365 ms	<sup>52</sup> K 110 ms	<sup>53</sup> K 30 ms	<sup>54</sup> K	-
45 <b>Ar</b> 21.48 s		<sup>47</sup> Ar 123 s							
44 <b>CI</b> 562 ms		46 <b>CI</b> 232 ms							
43 <b>S</b> 265 ms		45 <b>S</b> 68 ms							
<sup>42</sup> P 48.5 ms		44 <b>P</b> 18.5 ms							
<sup>41</sup> Si <sup>20 ms</sup>		<sup>43</sup> Si 30 ms			-	3	Log of -1.5	Half-life 0	e[s] 1.5 3
40 10 ms		<sup>42</sup> AI 3 ms				🗌 Esti	mated		Unknown
<sup>39</sup> Mg		<sup>41</sup> Mg		Edi	ted ima	ge based	d on "Tl	he Colo	urful Nuclide Chart"
N	=2	28		peo edw	vard.sim	sics.anu ipson@	anu.edu	v~ecs10 1.au	ISI CHART

## Shell Evolution

Far from stability, magic numbers change due to relative shifting of orbital energies.

Predicted that N = 28 gap weakens, N = 32, 34 gaps emerge [2].

[2] T. Otsuka *et al*. Phys. Rev. Lett. **87**, 082502 (2001)

	<sup>56</sup> SC 26 ms	55 <b>SC</b> 96 ms	54 <b>SC</b> 526 ms	<sup>53</sup> Sc 24s	<sup>52</sup> Sc 82 s	<sup>51</sup> SC 12.4s	<sup>50</sup> Sc 102.5 s	<sup>49</sup> SC 57.18m	48 <mark>50</mark> 43.67 h
Z=20	55 <b>Ca</b> 22 ms	<sup>54</sup> Ca 90 ms	<sup>53</sup> Ca 461 ms	<sup>52</sup> Ca	51 <b>Ca</b>	50 <b>Ca</b> 13.45 s	<sup>49</sup> Ca 8.718 m	<sup>48</sup> Ca	47 <b>Ca</b> 4536 d
							48 <b>K</b> 6.83 s	47 <b>K</b> 17.38 s	46K 96.3 s
							<sup>47</sup> Ar 1.23 s	46 <b>Ar</b> 8.4 s	45 <b>Ar</b> 21.48 s
							46 <b>CI</b> 232 ms	45 <b>Cl</b> 413 ms	44 CI 562 ms
							45 <b>S</b> 68 ms	44 <b>S</b> 100 ms	<sup>43</sup> S 265 ms
							44 <b>P</b> 18.5 ms	43 <b>P</b> 35.8 ms	<sup>42</sup> P 48.5 ms
e [s] 1.5 3	Half-life 0	Log of -1.5	3	-			<sup>43</sup> Si <sup>30 ms</sup>	<sup>42</sup> Si 12.5 ms	<sup>41</sup> Si <sup>20 ms</sup>
Unknown		mated	🗌 Esti				<sup>42</sup> AI 3 ms		40AI 10 ms
ourful Nuclide Chart"	he Colo	d on "Tl	ge base	ted imag	Edi		<sup>41</sup> Mg	40 <b>Mg</b>	<sup>39</sup> Mg
03/chart	1/~ecs1( 1.au	ı.edu.au anu.edu	sics.anu vpson@	ple.phys vard.sim	peo edw				N

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### Shell Evolution

Far from stability, magic numbers change due to relative shifting of orbital energies.

Predicted that N = 28 gap weakens, N = 32, 34 gaps emerge.

• Deformation at N = 28, Z<20 [3].

[3] H.L. Crawford *et al*. Phys. Rev. Lett **122**, 052501 (2019)



## Shell Evolution

Far from stability, magic numbers change due to relative shifting of orbital energies.

Predicted that N = 28 gap weakens, N = 32, 34 gaps emerge.

- Deformation at N = 28, Z<20.
- Doubly-magic <sup>52,54</sup>Ca from E(2<sub>1</sub><sup>+</sup>)
  [4].

[4] D. Steppenbeck *et al*. Nature **502**, 207 (2013)









Selective (d,p) reaction



Edited image based on "The Colourful Nuclide Chart people.physics.anu.edu.au/~ecs103/chart edward.simpson@anu.edu.au

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N=28



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ν

π

<sup>47</sup>K



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Methodology

SPIRAL1+  $^{47}$ K RIB @ 7.7MeV/u.  $5 \times 10^5$  pps,  $10^{-4}$  mass res.  $\rightarrow$  pure beam

TARGET 0.31(2) mg/cm<sup>2</sup> CD<sub>2</sub>



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VAMOS++ Zero degree; fast counting **Recoil timing** & reject C reactions



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MUGAST Light ejectile detection FWHM  $\approx$  300 keV in <sup>48</sup>K excitation.



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**Prompt γ-ray** emissions 16 ATC's @ 18 cm Pulse shape analysis, add-back & DC FWHM  $\approx$  7 keV @ 1.8 MeV;  $\beta$  = 0.16



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Light ejectile detection FWHM  $\approx$  300 keV in <sup>48</sup>K excitation.

### AGATA

**Prompt y-ray** emissions 16 ATC's @ 18 cm Pulse shape analysis, add-back & DC FWHM  $\approx$  7 keV @ 1.8 MeV;  $\beta$  = 0.16



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## MUGAST & VAMOS++



Unambiguous kinematic selection of reaction channel (d,d) elastic scattering provides internally consistent normalisation. (d,tγ) transfer allows for neutron hole investigation.

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## MUGAST & AGATA & VAMOS++



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## MUGAST & AGATA & VAMOS++



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## MUGAST & AGATA & VAMOS++



## MUGAST & AGATA & VAMOS++



p-γ Coinc.

Precise determination of state energies.

Construction of level scheme.

Clear isolation of specific states.









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### Angular Distributions

- Discriminate between p-wave (L=1) and fwave (L=3) transfer by differential cross section
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- Discriminate between p-wave (L=1) and fwave (L=3) transfer by differential cross section
- Comparison to data provides unambiguous L-transfer assignment
- Scaling factor between exp. and theory represents strength of population
  - Spectroscopic factor





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3.868 3.792

3.601

**3.254** 

2.908

2.407

Distinct regions of **p-wave** states and **f-wave** states.

**p-wave:** p<sub>1/2</sub>, p<sub>3/2</sub> **f-wave:** f<sub>5/2</sub>

Preliminary results suggest **mixed state** between the two regions.





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γ-ray transitions + ...



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#### Results Cross-shell interaction at N=28

### γ-ray transitions + L-transfers + ...





Results Cross-shell interaction at N=28

y-ray transitions + L-transfers + spectroscopic factors (vs. theory) = ...





Results Cross-shell interaction at N=28

 $\gamma$ -ray transitions + L-transfers + spectroscopic factors (vs. theory) = state spin ( $J^{\pi}$ ) + ...


Results Cross-shell interaction at N=28

 $\gamma$ -ray transitions + L-transfers + spectroscopic factors (vs. theory) = state spin ( $J^{\pi}$ ) + structure





Results Cross-shell interaction at N=28







Shell model fails to predict 1- ground state.



Shell model fails to predict 1- ground state. Measured p-wave & f-wave states have smaller gap than theory

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11/16



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Results Cross-shell interaction at N=28



#### Collected simultaneous data for **adding** and **removing** a neutron.

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Results Cross-shell interaction at N=28



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#### More well-known nucleus than <sup>48</sup>K.







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Results Cross-shell interaction at N=28



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#### Obs. (d,t)

Shell model is limited by small phase space:

• No deep v(d<sub>5/2</sub>)

• No high v(p<sub>1/2</sub>)

No reduction in (d,t) spectroscopic factor, as observed in (d,p).

Small occupation of  $v(p_{3/2})$  ground state suggests **no blocking** of (d,p) transfer.









First experimental measurement of exotic π(s<sub>1/2</sub>)⊗ν(fp) interaction conducted by way of <sup>47</sup>K(d,pγ)<sup>48</sup>K.

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- Small (d,p) spectroscopic factors, exploring possible interpretations.

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- Range of **p-wave and f-wave states identified**, each with spin-parity assignments and spectroscopic factors.
- **Preliminary comparison** with shell models; qualitative observations suggest **overestimation of N=34 gap**.
- Small (d,p) spectroscopic factors, exploring possible interpretations.
- Complementary <sup>47</sup>K(d,tγ)<sup>46</sup>K results obtained; no evidence to suggest transfer is "blocked".

## THANK YOU

C.J. Paxman<sup>1</sup>, W.N. Catford<sup>1</sup>, A. Matta<sup>2</sup>, G. Lotay<sup>1</sup>, D.T. Doherty<sup>1</sup>, M. Assié<sup>3</sup>, E. Clément<sup>4</sup>, A. Lemasson<sup>4</sup>, D. Ramos<sup>4</sup>, F. Galtarossa<sup>3</sup>, L. Achouri<sup>2</sup>, D. Ackermann<sup>4</sup>, D. Beaumel<sup>3</sup>, L. Canete<sup>1</sup>, P. Delahaye<sup>4</sup>, J. Dudouet<sup>5</sup>, B. Fernández-Domínguez<sup>6</sup>, D. Fernández-Fernández<sup>6</sup>, F. Flavigny<sup>2</sup>, C. Fougéres<sup>4</sup>, G. de France<sup>4</sup>, S. Franchoo<sup>3</sup>, J. Gibelin<sup>2</sup>, V. Girard-Alcindor<sup>4</sup>, N. Goyal<sup>4</sup>, F. Hammache<sup>3</sup>, D.S. Harrouz<sup>3</sup>, B. Jacquot<sup>4</sup>, L. Lalanne<sup>3,4</sup>, C. Lenain<sup>2</sup>, J. Lois-Fuentes<sup>6</sup>, T. Lokotko<sup>2</sup>, F.M. Marqués<sup>2</sup>, I. Martel<sup>7</sup>, N.A. Orr<sup>2</sup>, L. Plagnol<sup>2</sup>, D. Regueira-Castro<sup>6</sup>, N. de Séréville<sup>3</sup>, J.-C. Thomas<sup>4</sup>, A. Utepov<sup>4</sup>.

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[5] IP2I Lyon[6] Univ. Santiago de Compostela[7] Univ. Huelva







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- First experimental measurement of exotic  $\pi(s_{1/2}) \otimes \nu(fp)$  interaction.
- Range of p-wave and f-wave states identified.
- Overestimation of N=34 gap by shell model.
- Small (d,p) spectroscopic factors.
- No evidence to suggest transfer is "blocked".

# EXTRA SLIDES



J. Papuga *et al.* PRC 90, 034321 (2014)

Laser spec. at COLLAPS

<sup>47</sup>K structure:  $\pi(s_{1/2})^1 \pi(d_{3/2})^4$


Small spectroscopic factors are not believed to be quenching:

- Papers below find no quenching in transfer reactions.
- 48K not far from stability, so optical potentials are well known

PRL 110, 122503 (2013) PHYSICAL REVIEW LETTERS 22 MARCH 2013

## Limited Asymmetry Dependence of Correlations from Single Nucleon Transfer

F. Flavigny,<sup>1,2</sup> A. Gillibert,<sup>1</sup> L. Nalpas,<sup>1</sup> A. Obertelli,<sup>1</sup> N. Keeley,<sup>3</sup> C. Barbieri,<sup>4</sup> D. Beaumel,<sup>5</sup> S. Boissinot,<sup>1</sup> G. Burgunder,<sup>6</sup> A. Cipollone,<sup>4,7,8</sup> A. Corsi,<sup>1</sup> J. Gibelin,<sup>9</sup> S. Giron,<sup>5</sup> J. Guillot,<sup>5</sup> F. Hammache,<sup>5</sup> V. Lapoux,<sup>1</sup> A. Matta,<sup>5</sup> E. C. Pollacco,<sup>1</sup> R. Raabe,<sup>6,2</sup> M. Rejmund,<sup>6</sup> N. de Séreville,<sup>5</sup> A. Shrivastava,<sup>6</sup> A. Signoracci,<sup>1</sup> and Y. Utsuno<sup>10</sup>

PHYSICAL REVIEW C 92, 041302(R) (2015)

## New findings on structure and production of <sup>10</sup>He from <sup>11</sup>Li with the $(d, {}^{3}\text{He})$ reaction

A. Matta,<sup>1,2</sup> D. Beaumel,<sup>1</sup> H. Otsu,<sup>3</sup> V. Lapoux,<sup>4</sup> N. K. Timofeyuk,<sup>2</sup> N. Aoi,<sup>3</sup> M. Assié,<sup>1</sup> H. Baba,<sup>3</sup> S. Boissinot,<sup>4</sup> R. J. Chen,<sup>3</sup> F. Delaunay,<sup>5</sup> N. de Sereville,<sup>1</sup> S. Franchoo,<sup>1</sup> P. Gangnant,<sup>6</sup> J. Gibelin,<sup>5</sup> F. Hammache,<sup>1</sup> Ch. Houarner,<sup>6</sup> N. Imai,<sup>7</sup> N. Kobayashi,<sup>8</sup> T. Kubo,<sup>3</sup> Y. Kondo,<sup>8</sup> Y. Kawada,<sup>8</sup> L. H. Khiem,<sup>9</sup> M. Kurata-Nishimura,<sup>3</sup> E. A. Kuzmin,<sup>13</sup> J. Lee,<sup>3</sup> J. F. Libin,<sup>6</sup> T. Motobayashi,<sup>3</sup> T. Nakamura,<sup>8</sup> L. Nalpas,<sup>4</sup> E. Yu. Nikolskii,<sup>3,13</sup> A. Obertelli,<sup>4</sup> E. C. Pollacco,<sup>4</sup> E. Rindel,<sup>1</sup> Ph. Rosier,<sup>1</sup> F. Saillant,<sup>6</sup> T. Sako,<sup>8</sup> H. Sakurai,<sup>3</sup> A. M. Sánchez-Benítez,<sup>10,11</sup> J-A. Scarpaci,<sup>1</sup> I. Stefan,<sup>1</sup> D. Suzuki,<sup>1</sup> K. Takahashi,<sup>8</sup> M. Takechi,<sup>3</sup> S. Takeuchi,<sup>3</sup> H. Wang,<sup>3</sup> R. Wolski,<sup>12</sup> and K. Yoneda<sup>3</sup>





