



of Physics

Faculty

#### **Warsaw University** of Technology

Warsaw University of Technology; Faculty of Physics

e-mail: hanna.zbroszczyk@pw.edu.pl

#### GDR QCD Seminar, 6th October 2022





# Strong final state interactions

# seen by femtoscopy

#### Hanna Paulina Zbroszczyk

## Outline

#### Introduction

- QCD phase diagram
- HIC and femtoscopy method

#### Results

#### 1. Interactions

- a)Strong interactions between (anti)baryons
- b)Motivation to Y-N and Y-Y correlations
- c)Femtoscopy of strange baryons and their
  - interactions
- d)Possible **bound states**
- e)Coalescence production of deuterons
- f)Nonidentical particle correlation
- 2. QCD Phase diagram scan

Summary and conclusions



https://www.researchgate.net/figure/Color-Online-A-schematic-of-QCD-phase-diagram\_fig2\_357552969

# Introduction



## QCD Phase diagram of strongly interacting matter





#### Heavy-Ion collision and the HBT method







## Correlation femtoscopy





Size:  $\sim 10^{-15}$  m (fm) Time:  $\sim 10^{-23}$  s

Impossible to measure directly!

Femtoscopy (HIC) inspired by Hanbury Brown and Twiss interferometry method (Astronomy)

#### but!

- different scales,
- different measured quantities
- different determined quantities



## Traditional and non-traditional femtoscopy

Femtoscopy (originating from HBT): the method to probe geometric and dynamic properties of the source



Space-time properties  $(10^{-15}m, 10^{-23}s)$  can be determined due to two-particle correlations that arise due to: Quantum Statistics (Fermi-Dirac, Bose-Einstein); Final State Interactions (Coulomb, strong)

determined assumed  

$$C(k^*, r^*) = \int S(r^*) |\Psi(k^*, r^*)|^2 d^3 r^* = \frac{1}{2}$$

emission function  $S(r^*)$ 

 $\Psi(k^*, r^*)$  – two-particle wave function (includes e.g. FSI interactions)

 $Sgnl(k^*)$ - correlation function  $Bckg(k^*)$ 

measured  $Sgnl(k^*)$  $Bckg(k^*)$ 



## Traditional and non-traditional femtoscopy

If we assume we know the emission function, measured correlation function can be used to determine parameters of Final State Interactions



two-particle correlations that arise due to: Final State Interactions (Coulomb, strong)

assumed determined  

$$C(k^*, r^*) = \int S(r^*) |\Psi(k^*, r^*)|^2 d^3 r^* = \frac{1}{2} d^3 r$$

emission function

 $\Psi(k^*, r^*)$  – two-particle wave function (includes e.g. FSI interactions)

 $Sgnl(k^*)$  $\frac{\overline{Bckg(k^*)}}{Bckg(k^*)}$  - correlation function

- Space-time properties  $(10^{-15}m, 10^{-23}s)$  can be determined due to
- Quantum Statistics (Fermi-Dirac, Bose-Einstein);

measured  $Sgnl(k^*)$  $Bckg(k^*)$ 



## Traditional and non-traditional femtoscopy





#### What does femtoscopy measure?





Lisa MA, et al. 2005. Annu. Rev. Nucl. Part. Sci. 55:357-402



Lisa MA, et al. 2005. Annu. Rev. Nucl. Part. Sci. 55:357-402

# Homogeneity region





- <u>Strong</u>

#### Width of correlation function $\sim 1/R$

#### Bigger source and weaker correlation



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# Results



#### nature

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Published: 04 November 2015

#### Measurement of interaction between antiprotons

The STAR Collaboration

Nature527, 345–348 (2015)Cite this article9961Accesses47Citations368AltmetricMetrics

1 This article has been updated

#### Abstract

One of the primary goals of nuclear physics is to understand the force between nucleons, which is a necessary step for understanding the structure of nuclei and how nuclei interact with each other. Rutherford discovered the atomic nucleus in 1911, and the large body of knowledge about the nuclear force that has since been acquired was derived from studies made on nucleons or nuclei. Although antinuclei up to antihelium-4 have been discovered<sup>1</sup> and their masses measured, little is known directly about the nuclear force between antinucleons. Here, we study antiproton pair correlations among data collected by the STAR experiment<sup>2</sup> at the Relativistic Heavy Ion Collider (RHIC)<sup>3</sup>, where gold ions are collided with a centre-of-mass energy of 200 gigaelectronvolts per nucleon pair. Antiprotons are abundantly produced in such collisions, thus making it feasible to study details of the antiproton-antiproton interaction. By applying a technique similar to Hanbury Brown and Twiss intensity interferometry<sup>4</sup>, we show that the force between two antiprotons is attractive. In addition, we report two key parameters that characterize the corresponding strong interaction: the scattering length and the effective range of the interaction. Our measured parameters are consistent within errors with the corresponding values for proton-proton interactions. Our results provide direct information on the interaction between two antiprotons, one of the simplest systems of antinucleons, and so are fundamental to understanding the structure of more-complex antinuclei and their properties.

#### nature

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Article | Open Access | Published: 09 December 2020

#### Unveiling the strong interaction among hadrons at the LHC

**ALICE Collaboration** 

Nature588, 232–238 (2020)Cite this article9258Accesses6Citations231AltmetricMetrics

• A Publisher Correction to this article was published on 15 January 2021

• This article has been updated

#### Abstract

One of the key challenges for nuclear physics today is to understand from first principles the effective interaction between hadrons with different quark content. First successes have been achieved using techniques that solve the dynamics of quarks and gluons on discrete space-time lattices<sup>1,2</sup>. Experimentally, the dynamics of the strong interaction have been studied by scattering hadrons off each other. Such scattering experiments are difficult or impossible for unstable hadrons  $^{3,4,5,6}$  and so high-quality measurements exist only for hadrons containing up and down quarks<sup>7</sup>. Here we demonstrate that measuring correlations in the momentum space between hadron pairs<sup>8,9,10,11,12</sup> produced in ultrarelativistic proton-proton collisions at the CERN Large Hadron Collider (LHC) provides a precise method with which to obtain the missing information on the interaction dynamics between any pair of unstable hadrons. Specifically, we discuss the case of the interaction of baryons containing strange quarks (hyperons). We demonstrate how, using precision measurements of proton-omega baryon correlations, the effect of the strong interaction for this hadron-hadron pair can be studied with precision similar to, and compared with, predictions from lattice calculations<sup>13,14</sup>. The large number of hyperons identified in proton-proton collisions at the LHC, together with accurate modelling $^{15}$  of the small (approximately one femtometre) inter-particle distance and exact predictions for the correlation functions, enables a detailed determination of the short-range part of the nucleon-hyperon interaction.

## QCD Phase diagram of strongly interacting matter





- nucleon or / and nuclei.
- first time.
- more sophisticated anti-nuclei.



#### 1a) Strong interactions between anti-nucleons

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<b>1</b>	1.4 1.3 1.2 1.1 0.9
<u>1</u>	1.4 1.3 1.2 1.1 0.9 0.8

# 1a) Strong interactions between anti-nucleons



The scattering length f<sub>0</sub>: determines low-energy scattering.

The elastic cross section,  $\sigma_e$ , (at low energies) determined solely by the scattering length,  $\lim_{k \to 0} \sigma_e = 4\pi f_0^2$ 

 $d_0$  - the effective range of strong interaction between two particles. It corresponds to the range of the potential in an extremely simplified scenario - the square well potential.  $f_0$  and  $d_0$  - two important parameters of strong interaction between two particles. •

- Theoretical correlation function depends on: source size,  $k^*$ ,  $f_0$  and  $d_0$ . ۲

•  $f_0$  and  $d_0$  for the antiproton-antiproton interaction consistent with parameters

for the proton-proton interaction.

• Descriptions of the interaction among antimatter (based on the simplest systems of anti-nucleons) determined.

• A quantitative verification of matter-antimatter symmetry in context of the forces responsible for the binding of (anti)nuclei.





## 1b) Inside a neutron star

#### Proceedings of the 12th Asia Pacific Physics Conference (APPC12) Downloaded from journals.jps.jp by 217.149.243.2 on 04/21/22

Proceedings of the 12th Asia Pacific Physics Conference JPS Conf. Proc. 1, 011003 (2014) ©2014 The Physical Society of Japan

#### **Nuclear Matter in Neutron Stars** -A Great Challenge in Nuclear Physics—

Hirokazu Tamura<sup>1</sup>

<sup>1</sup>Department of Physics, Tohoku University, Sendai 980-8578, Japan

*E-mail: tamura@lambda.phys.tohoku.ac.jp* 

(Received September 1, 2013)

Lack of our knowledge on the nuclear (hadronic) matter in the neutron star prevents us from understanding its internal structure. The elucidation of the neutron star matter is one of the most important and challenging subjects in nuclear physics today. The nuclear matter EOS (Equation Of State) which would describe pure neutron matter in the outer core of the neutron star is studied via various laboratory experiments particularly by using neutron-rich nuclei. Hyperons and/or kaons are expected to appear in the high density matter in the inner core, which is also experimentally investigated via hypernuclei and kaonic nuclei. Most of the EOS's including hyperons or kaons lead to a serious discrepancy against the large mass (~  $2M_{\odot}$ ) for the recently observed neutron stars, and resolution of this mystery requires new theoretical frameworks in nuclear physics which correctly describe high density matter, some of which may incorporate hadronic degrees of freedom with a deconfined quark matter phase.





## 1b) Neutron star puzzle

- Hyperons: expected in the core of neutron stars; conversion of N into Y energetically favorable.
- Appearance of Y: The relieve of Fermi pressure  $\rightarrow$  softer EoS  $\rightarrow$  mass reduction (incompatible with observation)



# 1b) Neutron star puzzle

- energetically favorable.
- (incompatible with observation).

• Hyperons: expected in the core of neutron stars; conversion of N into Y  $M_{\rm NS} \approx 1 \div 2 M_{\odot}$ R  $\approx 10-12 \text{ km}$ • Appearance of Y: The relieve of Fermi pressure  $\rightarrow$  softer EoS  $\rightarrow$  mass reduction  $\rho \approx 3 \div 5 \rho_0$ 2.5 "stiff" EoS The solution: a mechanism providing the additional pressure to make the EoS units] 2 stiffer. Gravitational mass M<sub>G</sub> [solar mass Possible mechanisms: 1.5 • Two-body YN & YY interactions • Chiral forces "soft" EoS • Hyperonic Three Body Forces • Quark Matter Core - Phase transition at densities lower than hyperon threshold - N+free Y 0.5 -N+Y0.5 1.5 Central baryon number density [fm<sup>-3</sup>]  $\rho_0 \approx 2.8 \times 10^{14} \text{ g/cm}^3$ 





# 1b) Neutron star puzzle

- energetically favorable.
- (incompatible with observation).

• Hyperons: expected in the core of neutron stars; conversion of N into Y  $M_{\rm NS} \approx 1 \div 2 M_{\odot}$ R  $\approx 10-12 \text{ km}$ • Appearance of Y: The relieve of Fermi pressure  $\rightarrow$  softer EoS  $\rightarrow$  mass reduction  $\rho \approx 3 \div 5 \rho_0$ 2.5 "stiff" EoS The solution requires a mechanism that could provide the additional pressure at units] high densities needed to make the EoS stiffer. Gravitational mass M<sub>G</sub> [solar mass Possible mechanisms: 1.5 • Two-body YN & YY interactions • Chiral forces "soft" EoS • Hyperonic Three Body Forces • Quark Matter Core - Phase transition at densities lower than hyperon threshold - N+free Y 0.5 -N+YA lot of experimental and theoretical effort to understand: E 12 1 - The KN interaction, governed by the presence of  $\Lambda(1405)$ 1.5 0.5 Central baryon number density [fm<sup>-3</sup>] - The nature of  $\Lambda(1405)$ , the consequences of KNN formation - K and  $\overline{K}$  investigated to understand kaon condensation  $\rho_0 \approx 2.8 \times 10^{14} \text{ g/cm}^3$ 







**Fig. 1.** A schematic illustration for nuclear (baryonic) matter Equation Of State (EOS) as a function of baryon density for symmetric nuclear matter, pure neutron matter, and charge-neutral baryonic matter with hyperons. The EOS is determined only for symmetric nuclear matter around  $\rho \sim (0.5-2)\rho_0$  and has large uncertainties particularly for pure neutron matter, as symbolically shown in the figure.

"To establish the EOS applicable to the neutron star has been one of the most important subjects in nuclear physics for a long time but has not been achieved yet."

*H.Tamura, JPS Conf. Proc.*, 011003 (2014)





# 1b) Equation Of States for different types of baryonic matter

Determining the Equation of State (EOS) of dense neutron-rich nuclear matter is a shared goal of both nuclear physics and astrophysics. Except possible phase transitions, the density dependence of nuclear symmetry  $Esym(\rho)$  is the most uncertain part of the EOS of neutron-rich nucleonic matter especially at supra-saturation densities.

"The Energy per nucleon  $E(\rho, \delta)$  in nuclear matter at density  $\rho$  and isospin asymmetry  $\delta = (\rho_n - \rho_p)/\rho$ is the most basic input for calculating the EOS of neutron star matter regardless of the models used."









- Experiment: More interest about Y–N and Y–Y interactions.
- **Theory**: Major steps forward have been made (Lattice QCD).
- Numerous theoretical predictions exist, but no clear evidence for any such bound states, despite many experimental searches.
- The existence of hypernuclei (confirmed by attractive Y–N interaction)  $\rightarrow$  indicates the possibility to bind Y to N.
- The measurement of the Y–N and Y–Y interactions leads to important implications for the possible formation of **Y–N** or **Y–Y bound states**.
- A precise knowledge of these interactions help to explore unknown structure of neutron stars.

#### 1b) Y-N and Y-Y interactions











# 1c) YN interactions at HADES



$$C_F(k) = \frac{C_{meas}(k)}{C_{LRC}(k)}$$
$$C_{LRC}(k^*) = 1 + ak + bk^2$$

#### 1c) Strange Baryon Correlations (Including $\Lambda$ Hyperons)





## 1d) Strange Baryon Correlations (including p- $\Omega$ )



	V1	$\mathbf{V}_{2}$	V3
Ebin [MeV]	_	6.3	26.9
a0 [MeV]	-1.12	5.79	1.29
reff [MeV]	-1.16	0.96	0.65

Scattering lenght positive, favor the hypothesis of  $p\Omega$  bound state



## 1d) Strange Baryon Correlations (Including $\Xi$ Hyperons)

First measurement of  $\Xi$ - $\Xi$  correlation at RHIC

Combination of quantum statistics, strong interaction, and Coulomb interaction.

Anti-correlation at Q < 0.25 GeV/c.

Lattice QCD/chiral EFT calculations indicate an attractive interaction potential, but not strong enough to form a bound state.









- High-precision measurement of the strong interaction (anti-correlation) between kaons and protons.
- A structure (ALICE in p+p collisions) observed around a relative momentum of 58 MeV/c in the measured correlation function of opposite charges in p+p collisions.









# 1d) Neutral kaons

# Parametrization - K<sup>0</sup><sub>s</sub>K<sup>0</sup><sub>s</sub>

 $\lambda$  - the correlation strength,  $R_{inv}$  - the size of the particle-emitting source. Lednicky & Lyuboshitz model includes strong FSI: [Sov.J.Nucl.Phys. 35, 770 (1982)]

$$CF(q_{inv}) = 1 + \lambda \left( e^{[-R_{inv}^2 q_{inv}^2]} + \frac{1}{2} \left[ \left| \frac{f(k^*)}{R_{inv}} \right|^2 + \frac{4\Re f(k^*)}{\sqrt{\pi}R_{inv}} F_1(q_{inv}R_{inv}) - \frac{2\Im f(k^*)}{\sqrt{\pi}R_{inv}} F_2(q_{inv}R_{inv}) \right] \right]$$
QS effect strong FSI through the f<sub>0</sub>(980) and a<sub>0</sub>(980) resonances

$$F_1(z)$$

$$F_{1}(z) = \int_{0}^{z} dx \frac{e^{x^{2}} - x^{2}}{z}, F_{2}(z) = \frac{1 - e^{x^{2}}}{z}$$

$$f(k^{*}) = \frac{1}{2} [f_{0}(k^{*}) + f_{1}(k^{*})], \quad f_{I}(k^{*}) = \frac{\gamma_{r}}{m_{r} - s - i\gamma_{r}k^{*} - i\gamma_{r}'k_{r}'}, \quad s = 4(m_{K}^{2} + k^{*2})$$

	$m_{f_0}\left[\frac{GeV}{c^2}\right]$	$\gamma_{f_0K\overline{K}}$	$\gamma_{f_0\pi\pi}$	$m_{a_0}\left[\frac{GeV}{c^2}\right]$	<i>Υ</i> <sub>α0</sub> K ¯ K	<b>γ</b> <sub>a0</sub> ππ
Antonelli [1]	0.973	2.763	0.5283	0.985	0.4038	0.3711
Achasov2001 [2]	0.996	1.305	0.2684	0.992	0.5555	0.4401
Achasov2003 [3]	0.996	1.305	0.2684	1.003	0.8365	0.4580
Martin [4]	0.978	0.792	0.1990	0.974	0.3330	0.2220

**Gaussian density distribution** (includes only QS effects):  $CF(q_{inv}) = 1 + \lambda e^{[-R_{inv}^2 q_{inv}^2]}$ 

[1] eConf C020620, THAT06 (2002), [2] Phys. Rev. D 63, 094007 (2001) [3] Phys. Rev. D 68, 014006 (2003), [4] Nucl. Phys. B 121, 514–530 (1977)



The strong final-state interaction has a significant effect on the neutral kaons correlation due to the near-threshold  $f_0(980)$  and  $a_0(980)$  resonances

#### 1d) Neutral kaons



# Parametrization - $K_s^0 K^{\pm}$

$$CF(k^*) = 1 + \frac{\lambda}{4} \left[ \left| \frac{f(k^*)}{R} \right|^2 + \frac{4\Re f(k^*)}{\sqrt{\pi}R} F_1(2k^*R) - \frac{2\Im f(k^*)}{\sqrt{\pi}R} F_2(2k^*R) \right]$$

$F_{1}(z) = \int_{0}^{z} dx \frac{e^{x^{2}} - x^{2}}{z}, F_{2}(z) = \frac{1 - e^{x^{2}}}{z}$ $f(k^{*}) = \frac{\gamma_{r}}{m_{r} - s - i\gamma_{r}k^{*} - i\gamma_{r}'k_{r}'}, \qquad s = 4(m_{K}^{2} + k^{*2})$						
	$m_{a_0}\left[\frac{GeV}{c^2}\right]$	Υ <sub>α0</sub> Κ κ	<b>γ</b> α <sub>0</sub> πη			
Antonelli [1]	0.985	0.4038	0.3711			
Achasov2001 [2]	0.992	0.5555	0.4401			
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Martin [4]	0.974	0.3330	0.2220			

[1] eConf C020620, THAT06 (2002), [2] Phys. Rev. D 63, 094007 (2001) [3] Phys. Rev. D 68, 014006 (2003), [4] Nucl. Phys. B 121, 514–530 (1977)

#### 1d) Neutral kaons

Lednicky & Lyuboshitz model includes strong FSI: [Sov.J.Nucl.Phys. 35, 770 (1982)]

strong FSI through the  $a_0$  (980) resonance

$$\frac{e^{x^2} - x^2}{z}, F_2(z) = \frac{1 - e^{x^2}}{z}$$

$$\frac{r}{z}, S = 4(m_K^2 + 1)$$

## 1d) Neutral kaons



The  $a_0(980)$  FSI parametrization gives very good representation of the shape of the signal region in CF; the parametrization with the larger  $a_0(980)$  mass and decay coupling gives larger size of the source; Antonelli parametrization favors  $a_0(980)$  resonance as a tetraquark





The  $a_0(980)$  FSI parametrization gives very good representation of the shape of the signal region in CF; the parametrization with the larger  $a_0(980)$  mass and decay coupling gives larger size of the source; Antonelli parametrization favors  $a_0(980)$  resonance as a tetraquark

#### 1d) Neutral kaons



# 1e) Light nuclei formation at STAR



• First measurement of proton-deuteron and deuteron-deuteron correlation functions from STAR

• Proton-deuteron and deuteron-deuteron correlations qualitatively described by the Lednicky-Lyuboshits model; deuterondeuteron has larger emission source size than proton-deuteron

•Deuteron-deuteron correlations described better by the model including coalescence. Light nuclei are likely to be formed via coalescence.







## 1f) Kaon-Lambda

To measure scattering parameters of  $\Lambda - K$  pairs in all three charge combinations  $\Lambda - K^+$ ,  $O = \Lambda - K^-$  and  $\Lambda - K_S^0$ .

**Strong** force is **repulsive** in the  $\Lambda - K^+$  interaction and **attra**  $\Lambda - K_S^0$  interactions.

Different  $q - \bar{q}$  interactions between the pairs  $(s - \bar{s} \text{ in } \Lambda - K^+ \Lambda - K^-)$ 

Different net-strangeness for each system (S=0 for  $\Lambda - K^+$ , and

Source radii larger than expected from extrapolation from identical particle femtoscopic studies, due to separation in space-time of the single-particle  $\Lambda$  and K source distributions.





**ctive** in the 
$$\Lambda - K^-$$
 and

and 
$$\Lambda - K_S^0$$
 and  $u - \bar{u}$  in

$$|S=-2$$
 for  $\Lambda - K^-$ ).



## CBM (and HADES) challenges

#### Challenges in QCD matter physics –

diagnostic probes which are sensitive to the dense phase of the nuclear fireball. The goal of the CBM experiment at SIS100 ( $\sqrt{s_{NN}} = 2.7 - 4.9$  GeV) is to discover fundamental properties of QCD matter: the phase structure at large baryon-chemical potentials ( $\mu_B > 500$  MeV), effects of chiral symmetry, and the equation-of-state at high density as it is expected to occur in the core of neutron stars. In this article, we review the motivation for and the physics programme of CBM, including activities before the start of data taking in 2024, in the context of the worldwide efforts to explore high-density QCD matter.

arXiv:1607.01487v3 [nucl-ex] 29 Mar 2017



#### The CBM experimental setup together with the HADES detector

#### The scientific programme of the Compressed Baryonic Matter experiment at FAIR





## QCD Phase diagram of strongly interacting matter - intermediate energies







# Program Beam Energy Scan



RHIC Top Energy: 200 GeV p+p, p+Al, p+Au, d+Au, 3He+Au, Cu+Cu, Cu+Au, Ru+Ru, Zr+Zr, Au+Au, U+U1. QCD at high energy density/temperature 2. Properties of QGP, EoS

Beam Energy Scan: Au+Au 7.7-62 GeV 1. Search for turn-off of QGP signatures 2. Search for signals of the first-order phase transition 3. Search for QCD critical point 4. Search for signals of Chiral symmetry restoration

Fixed-Target Program: Au+Au = 3.0-7.7 GeV High baryon density regime with 420-720 MeV







# Identical pion femtoscopy

- $\rightarrow$  R<sub>side</sub> spatial source evolution in the transverse direction
- $\rightarrow$  R<sub>out</sub> related to spatial and time components
- $\rightarrow R_{out}/R_{side}$  signature of phase transition
- $\rightarrow$  R<sub>out</sub><sup>2</sup>- R<sub>side</sub><sup>2</sup> =  $\Delta \tau^2 \beta_t^2$ ;  $\Delta \tau$  emission time
- $\rightarrow$  R<sub>long</sub> temperature of kinetic freeze-out and source lifetime
  - $C(\vec{q}) = (1 \lambda)$  $\times \exp\left(-q_o^2 R\right)$

HBT source determined for wide range of collision energy;

Non-monotonic behavior seen in three directions

$$k_{\rm Coul}(q_{\rm inv})\lambda R_o^2 - q_s^2 R_s^2 - q_l^2 R_l^2 - 2q_o q_s R_{os}^2 - 2q_o q_l R_{ol}^2$$













$$R_{\mu,n}^{2}\cos(n\Phi) \qquad (\mu = o, s, l, ol) \qquad \epsilon_{PP} = \frac{\sqrt{(\sigma_{y}^{2} - \sigma_{x}^{2})^{2} + \sigma_{y}^{2}}}{\sigma_{x}^{2} + \sigma_{y}^{2}}$$
$$\epsilon_{F} = \frac{\sigma_{y}^{\prime 2} - \sigma_{x}^{\prime 2}}{\sigma_{y}^{\prime 2} + \sigma_{x}^{\prime 2}} \approx 2\frac{R_{s,0}^{\prime 2}}{R_{s,0}^{2}}$$

$$R_{\mu,n}^{2} \sin(n\Phi) \qquad (\mu = os) \\ \sigma_{x}^{2} = \{x^{2}\} - \{x\}^{2} \text{ and } \sigma_{y}^{2} = \{y^{2}\} - \{$$

$$R_{out}^2 - R_{side}^2 = \beta_t^2 \Delta \tau^2$$

Visible peak in 
$$\frac{R_{out}}{R_{side}}(\sqrt{s_{NN}})$$
 near the  $\sqrt{s_{NN}} \simeq 20$ 

QCD calculations predict a peak near to the QGP transition threshold - signature of first-order phase transition?

Theoretical attention from hydro and transport models needed

) GeV



#### How to measure a phase transition?



vHLEE+UrQMD model verify sensitivity of HBT measurements to the first-order phase transition



Phys.Rev. C96 (2017) no.2, 024911

vHLLE (3+1)-D viscous hydrodynamics: Iu. Karpenko, P. Huovinen, H. Petersen, M. Bleicher; Phys.Rev. C 91, 064901 (2015), arXiv:1502.01978, 1509.3751

HadronGas + Bag Model  $\rightarrow 1^{st}$ order PT; P.F. Kolb, et al, PR C 62, 054909 (2000)

Chiral EoS  $\rightarrow$  crossover PT (XPT); J. Steinheimer, et al, J. Phys. G 38, 035001 (2011)









# Conclusions & Summary

- •Descriptions of the interaction among antimatter (based on the simplest systems of anti-nucleons) determined.
- •A quantitative verification of matter-antimatter symmetry in context of the forces responsible for the binding of (anti)nuclei.
- •Scattering length is positive and favor  $\Lambda \Lambda$  bound state hypothesis
- •Scattering length is positive and favor  $p \Omega$  bound state hypothesis
- •Searched for  $\Xi \Xi$  bounds states have started
- •Antonelli parametrization of  $K_S^0 K^{\pm}$  strong interactions favors  $a_0(980)$  resonance as a tetraquark
- •d-d CF described better by the model including coalescence
- •Light nuclei are likely to be formed via coalescence
- •Two-particle correlations are useful tool to explore unknown QCD phase diagram area



# Back-up slides

# Nonidentical particle correlations





