



# **First Xenon Results With R2D2**

**Pierre Charpentier** 



## université de **BORDEAUX**

**IRN Neutrino** 





### Experimental Setup

### Latest Improvements And First Xenon Results



Pierre Charpentier



**IRN Neutrino** 

17/11/2022



# Introduction **Motivations: R2D2 - Rare Decays with Radial Detector**

search.

for  $\beta\beta0\nu$  decay searches by matching the following requirements:

Excellent energy resolution: 1% FWHM at <sup>136</sup>Xe Q<sub>ββ</sub> of 2.458 MeV

Low (zero) background: below 0.1 event / year

Large mass of isotope <sup>136</sup>Xe: ton scale experiment

<sup>1</sup>HP-TPC: High Pressure Time Projection Chamber <sup>2</sup> JINST 13 (2018) no.01, P01009 [arXiv:1710.04536] <sup>3</sup> SPC: Spherical Proportional Counter

**Pierre Charpentier** 

- R2D2 is an R&D project that explores a single anode HP-TPC<sup>1</sup> solution for BBOv
- Preliminary simulations<sup>2</sup> have shown that a SPC<sup>3</sup> could reach competitive sensitivity









# Principle And Advantages

## • Excellent energy resolution<sup>1</sup>.

## Low (zero) background.

- Excellent background rejection.
- Low material budget.

# Large mass of isotope <sup>136</sup>Xe. Easy isotope enrichment.

<sup>1</sup> A. Bolotnikov and B. Ramsey, Nucl.Instrum.Meth.A 396 1997 **Pierre Charpentier** 



## <sup>136</sup>Xe ( $\beta\beta$ emitter) enriched gas as medium



**IRN Neutrino** 





# Principle And Advantages

## Excellent energy resolution<sup>1</sup>.

## Low (zero) background.

- Excellent background rejection.
- Low material budget.

# Large mass of isotope <sup>136</sup>Xe. Easy isotope enrichment.

<sup>1</sup> A. Bolotnikov and B. Ramsey, Nucl.Instrum.Meth.A 396 1997 **Pierre Charpentier** 



## <sup>136</sup>Xe ( $\beta\beta$ emitter) enriched gas as medium



**IRN Neutrino** 





# Principle And Advantages

- Excellent energy resolution<sup>1</sup>.
- Low (zero) background.
  - Excellent background rejection.
  - Low material budget.
- Large mass of isotope <sup>136</sup>Xe. Easy isotope enrichment.

<sup>1</sup> A. Bolotnikov and B. Ramsey, Nucl.Instrum.Meth.A 396 1997 **Pierre Charpentier** 



## <sup>136</sup>Xe (ββ emitter) enriched gas as medium



**IRN Neutrino** 





# Principle And Advantages

- Excellent energy resolution<sup>1</sup>.
- Low (zero) background.
  - Excellent background rejection.
  - Low material budget.
- Large mass of isotope <sup>136</sup>Xe.
  - Easy isotope enrichment.

<sup>1</sup> A. Bolotnikov and B. Ramsey, Nucl.Instrum.Meth.A 396 1997 **Pierre Charpentier IRN Neutrino** 



## <sup>136</sup>Xe (ββ emitter) enriched gas as medium







# Introduction **Current Phases**

## Current prototype goal: Achieve 1% FWHM energy resolution at 2.458 MeV, $^{136}Xe Q_{\beta\beta}$ .

- First phase with Argon as detector medium and <sup>210</sup>Po as  $\alpha$  source.
  - Electronics and data acquisition
  - Sensor characterisation and improvement
  - Light readout<sup>1</sup>
- Second phase with Xenon as detector medium and <sup>210</sup>Po as  $\alpha$  source.
  - Gas purity development
  - Gas recirculation and recovery
  - Exploring CPC<sup>2</sup> solution
- <sup>1</sup> Nucl.Instrum.Meth.A 1028 (2022) 166382 [arXiv:2201.12621] <sup>2</sup> CPC: Cylindrical Proportional Counter

### **Pierre Charpentier**



**IRN Neutrino** 





# Introduction **Current Phases**

- First phase with Argon as detector medium and <sup>210</sup>Po as  $\alpha$  source.
  - Electronics and data acquisition
  - Sensor characterisation and improvement
  - Light readout<sup>1</sup>
- Second phase with Xenon as detector mediur
  - Gas purity development
  - Gas recirculation and recovery
  - Exploring CPC<sup>2</sup> solution
- <sup>1</sup> Nucl.Instrum.Meth.A 1028 (2022) 166382 [arXiv:2201.12621] <sup>2</sup> CPC: Cylindrical Proportional Counter

**Pierre Charpentier** 

## Current prototype goal: Achieve 1% FWHM energy resolution at 2.458 MeV, <sup>136</sup>Xe $Q_{\beta\beta}$ .





### **IRN Neutrino**





# Introduction **Current Phases**

- First phase with Argon as detector medium and <sup>210</sup>Po as  $\alpha$  source.
  - Electronics and data acquisition
  - Sensor characterisation and improvement
  - Light readout<sup>1</sup>
- Second phase with Xenon as detector medium and <sup>210</sup>Po as  $\alpha$  source.
  - Gas purity development
  - Gas recirculation and recovery
  - Exploring CPC<sup>2</sup> solution

<sup>1</sup> Nucl.Instrum.Meth.A 1028 (2022) 166382 [arXiv:2201.12621] <sup>2</sup> CPC: Cylindrical Proportional Counter

**Pierre Charpentier** 

Current prototype goal: Achieve 1% FWHM energy resolution at 2.458 MeV,  $^{136}Xe Q_{\beta\beta}$ .



**IRN Neutrino** 





# **Experimental setup Prototype setup evolution at LP2I Bordeaux**



## **SPC-1** (2018) 40 cm Ø Up to 1 bar<sup>1</sup>



<sup>1</sup> No Pressure certification <sup>2</sup> Pressure certification

**Pierre Charpentier** 



## **CPC-1** (2022) 1m x 37 cm Ø Up to 1 bar<sup>1</sup>

Made at



### 17/11/2022

### **IRN Neutrino**



## Ourification:

- High purity is a strong requirement.
- Circulation inside cold getters.
- Recirculation:
  - Recirculation system.
  - Controlled flow.
- Recovery: First design by
  - Creation of a cryopumping system.
  - Pressure controlled valve.

### **Pierre Charpentier**





## Output Purification:

- High purity is a strong requirement.
- Circulation inside cold getters.
- Recirculation:
  - Recirculation system.
  - Controlled flow.
- Recovery: First design by
  - Creation of a cryopumping system.
  - Pressure controlled valve.

### **Pierre Charpentier**





### 17/11/2022

### **IRN Neutrino**



## Output Purification:

- High purity is a strong requirement.
- Circulation inside cold getters.
- Recirculation:
  - Recirculation system.
  - Controlled flow.
- Recovery: First design by
  - Creation of a cryopumping system.
  - Pressure controlled valve.

### **Pierre Charpentier**





### 17/11/2022

### **IRN Neutrino**



## Output Purification:

- High purity is a strong requirement.
- Circulation inside cold getters.
- Recirculation:
  - Recirculation system.
  - Controlled flow.
- Recovery: First design by



- Creation of a cryopumping system.
- Pressure controlled valve.

### **Pierre Charpentier**





### **IRN Neutrino**





## Output Purification:

- High purity is a strong requirement.
- Circulation inside cold getters.
- Recirculation:
  - Recirculation system.
  - Controlled flow.
- Recovery: First design by



- Creation of a cryopumping system.
- Pressure controlled valve.

### **Pierre Charpentier**



## Hot getter coming in 2023 ...

### **IRN Neutrino**



## **Experimental setup**

# CPC

Since May 2022 a new prototype is under study. A CPC exploiting the existing electronic chain, pumping and gas management system.

- Inox Tube: 1m50 x 20cm Ø
- Copper cathode: 1m x 17.5 cm Ø
- Tungsten anode: 20 µm Ø
- <sup>210</sup>Po source

**Electric Field:** 

- SPC:  $\propto \frac{1}{n^2}$
- CPC:  $\propto \frac{1}{r}$  (far from the edges)





### **IRN Neutrino**



## Latest Improvement

# Updated Argon Result: SPC – Proportional

Former published<sup>1</sup> measurement with SPC-1 in ArP2<sup>2</sup> spanned from 200 mbar to 1.1 bar using <sup>210</sup>Po  $\alpha$  of 5.3 MeV.

New measurement with SPC-2:

- Measurement up to 3 bar in proportional mode.
- Resolution between 1% and 1.3%.
- Anode radius: 1 mm.
- Limits:
  - HV: with small radius anode the electric field is too weak to collect all electrons at the cathode.
  - Gas purity: at higher pressure even small electronegative impurities induce important signal reduction.

<sup>1</sup> JINST 16 (2021) 03, P03012 [arXiv:2007.02570]

<sup>2</sup> ArP2: Argon (98%) and CH<sub>4</sub> (2%) mix

**Pierre Charpentier** 





**IRN Neutrino** 





# Latest Improvement Updated Argon Result: SPC – Ionisation

- SPC-2: Resolution measure for <sup>210</sup>Po  $\alpha$  of 5.3 MeV.
- is to use a larger sensor tip.
- The downside is a lower gain on the signal and thus a ionisation working mode.
- Anode with 3 mm radius at 1 bar:
  - HV: 700 V instead of 1900 V in proportional
  - Spread: 2.5 ADU  $\rightarrow$  DAQ limitation

### The electric field at the cathode is too weak to collect all the charges. A workaround **Best Result** SPC ionisation Entries Entries 45 Mean Std Dev 35 ⊢ Gain = 125⊢ HV = 700V20 FWHM resolution = 8.2% 15 10 5 80 70 90 110 100

## Integral: ~70 ADU $\rightarrow$ resolution ~8% FWHM independently of the gas pressure

**IRN Neutrino** 





# Latest Improvement **CPC** Argon result

for the same HV compared to the SPC.

First measurement in proportional mode (1 bar, 900 V):

1.2% resolution FWHM.

Test in ionisation mode (1 bar, 200 V):

• 4.9% resolution FWHM.

This result is better with respect to the SPC since such resolution is dominated by baseline fluctuation which is much smaller by decoupling HV and signal.

## First tests with ArP2 show a much lower noise due to the separation of the signal from the HV. The 1/r electric field dependence results in a higher field at the cathode



**IRN Neutrino** 



# **First Xenon Results** SPC Xenon: Main Difficulties

Switching from Ar to Xe implied a lot of challenges to overcome. Aside from the previously discussed technical consideration: Gain (ne) Argon

- Xe electrons drift time is one order of magnitude larger than Ar.
- Electronegative impurities become more
   critical. Purity is paramount.
- A stronger electric field is needed across the whole medium.
  - Higher  $HV \rightarrow$  higher noise.
  - Larger anode  $\rightarrow$  lonisation mode only.



**IRN Neutrino** 





## 250 mbar

- HV scan: 800 V up to 1400 V
- Optimal: 1300 V
  - Integral: 118 ADU
  - Sigma: 1.9 ADU
  - Resolution: 3.8%

Higher pressure measurement would require a gas purity level that is not achieved yet with the current setup. Even after several days of recirculation through the getters, electron attachment is still present.



Resolution: 7%





First measurement: 500 mbar; 900 V; 24 hours of recirculation. Still dominated by attachment issue...

- Image: But 2.3% of resolution in proportional mode.
- Down to 1.8% if a rise time cut is applied to reject  $\alpha$  particles with partial deposit.

Second measurement: 1 bar; 1200 V; 48 hours of recirculation.

Attachment reduced, thus integral has increased and it is less direction dependent.

• 2.9% of resolution. Down to 1.8% if a rise time cut is applied

**IRN Neutrino** 





First measurement: 500 mbar; 900 V; 24 hours of recirculation.

- Still dominated by attachment issue...
- In But 2.3% of resolution in proportional mode.
- Obvious Down to 1.8% if a rise time cut is applied to reject  $\alpha$  particles with partial deposit.

Second measurement: 1 bar; 1200 V; 48 hours of recirculation.

- Attachment reduced, thus integral has increased and it is less direction dependent.
- 2.9% of resolution. Down to 1.8% if a rise time cut is applied



17/11/2022

### **IRN Neutrino**



First measurement: 500 mbar; 900 V; 24 hours of recirculation.

- Still dominated by attachment issue...
- In But 2.3% of resolution in proportional mode.
- Down to 1.8% if a rise time cut is applied to reject  $\alpha$  particles with partial deposit.

Second measurement: 1 bar; 1200 V; 48 hours of recirculation.

- Attachment reduced, thus integral has increased and it is less direction dependent.
- 2.9% of resolution. Down to 1.8% if a rise time cut is applied



17/11/2022

**IRN Neutrino** 



First measurement: 500 mbar; 900 V; 24 hours of recirculation.

- Still dominated by attachment issue...
- In But 2.3% of resolution in proportional mode.
- Output Down to 1.8% if a rise time cut is applied to reject  $\alpha$  particles with partial deposit.

Second measurement: 1 bar; 1200 V; 48 hours of recirculation.

- Attachment reduced, thus integral has increased and it is less direction dependent.
- 2.9% of resolution. Down to 1.8% if a rise time cut is applied.



**IRN Neutrino** 



# **First Xenon Results CPC Xenon: Cosmic Background**

- Output Content SPC, the geometry and orientation of our CPC prototype makes it  $\widehat{\underline{g}}$ more sensible to cosmic muons background.
- The energy deposit of a muon in Xe at 1 § bar is significantly enough degrade the energy resolution of the  $\alpha$  particles.
- This explain the right hand tail of the CPC reconstructed integral distribution.
- Nevertheless the final experiment shall take place in underground facilities, avoiding such inconveniences.

Pierre Charpentier



### **IRN Neutrino**



# **First Xenon Results CPC Xenon: Cosmic Background**

- Output Content SPC, the geometry and orientation of our CPC prototype makes it  $\widehat{\mathbf{P}}$ more sensible to cosmic muons background.
- The energy deposit of a muon in Xe at  $1\bar{a}$ bar is significantly enough degrade the energy resolution of the  $\alpha$  particles.
- This explain the right hand tail of the CPC reconstructed integral distribution.
- Nevertheless the final experiment shall take place in underground facilities, avoiding such inconveniences.



### **IRN Neutrino**



# **First Xenon Results CPC Xenon: Cosmic Background**

- Output Content SPC, the geometry and orientation of our CPC prototype makes it more sensible to cosmic muons background.
- The energy deposit of a muon in Xe at 1 bar is significantly enough degrade the energy resolution of the  $\alpha$  particles.
- This explain the right hand tail of the CPC <sup>100</sup> reconstructed integral distribution.
- Nevertheless the final experiment shall of take place in underground facilities, avoiding such inconveniences.

**Pierre Charpentier** 



**IRN Neutrino** 





## 

- while the CPC has reached 1.2% up to 1 bar.
- 1.4 % is obtained at 1 bar in CPC.
- resolution goal. Hot getter coming in 2023.
- Test of a small CPC up to 40 bar in Xenon in 2023.

The SPC and CPC geometries were tested and compared in both Ar and Xe.

R2D2 new SPC setup allows to obtain resolution under 1.4% in Ar up to 3 bar

The first measurement in Xenon were performed with success and resolution of

Some efforts remain, especially concerning gas purity which is crucial for the 1%

**IRN Neutrino** 











## More infos: https://r2d2.in2p3.fr

Pierre Charpentier



# The End



### 17/11/2022

### **IRN Neutrino**