

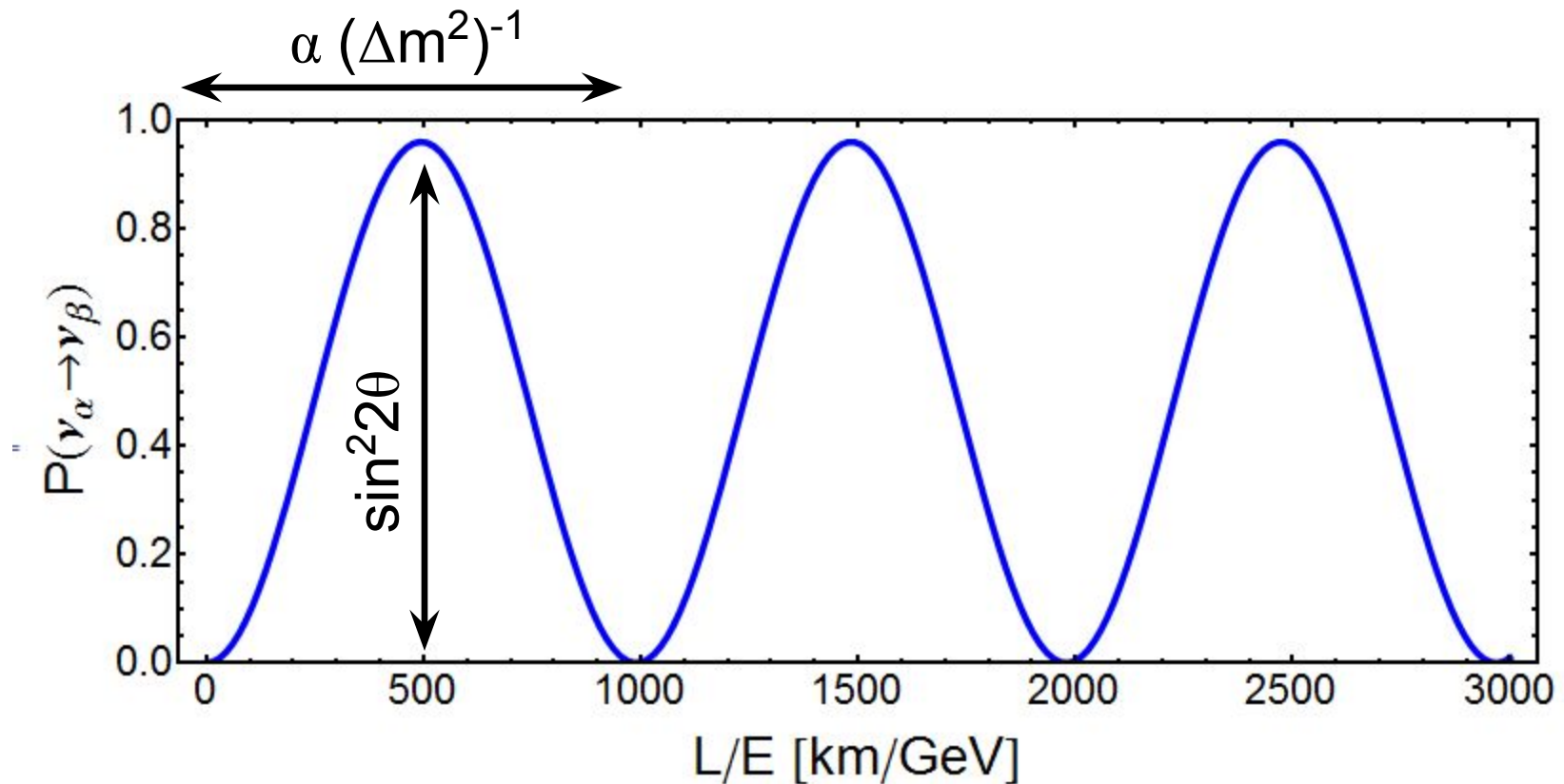
# Neutrino Oscillation Studies with KM3NeT/ORCA

João Coelho  
APC Laboratory  
16 November 2022



# Neutrino Oscillations

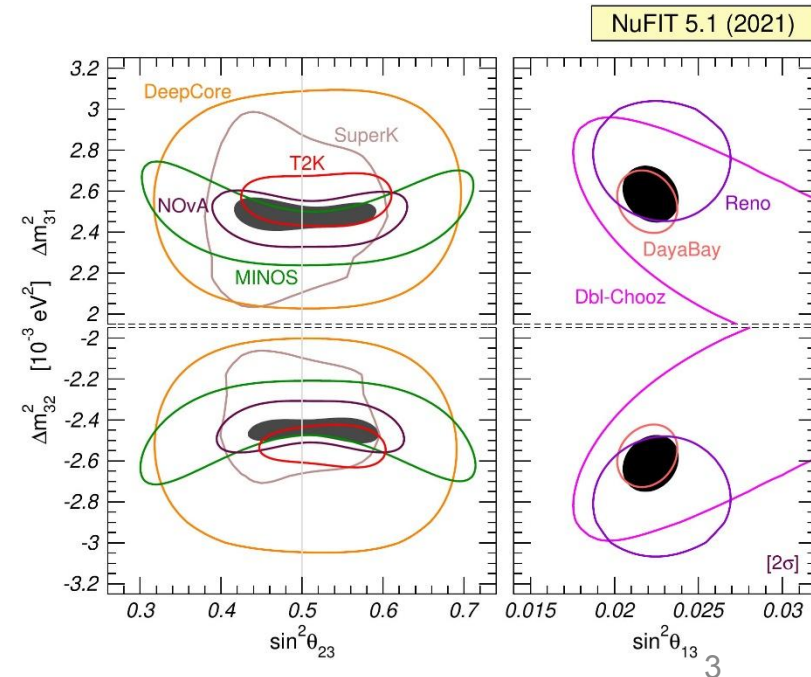
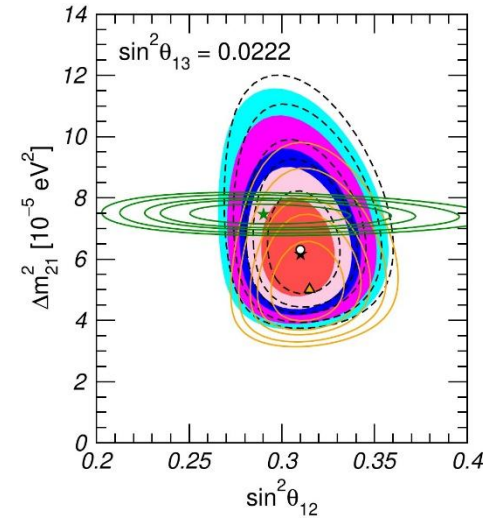
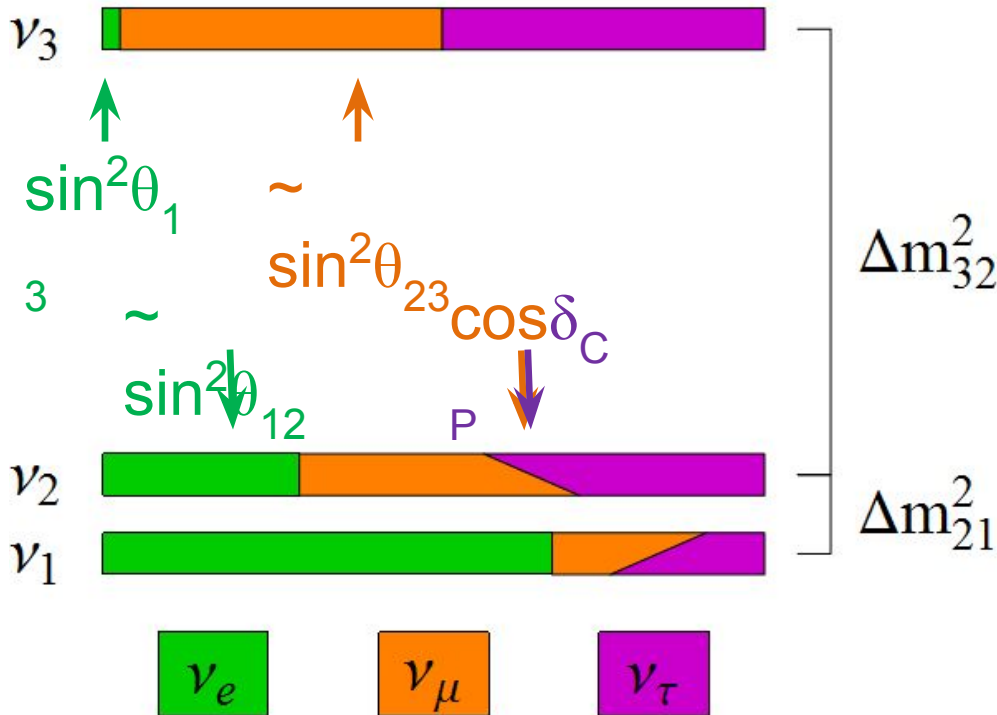
$$P(\nu_\alpha \rightarrow \nu_\beta) \approx \sin^2 2\theta \times \sin^2 \left( 1.27 \times \Delta m^2 [\text{eV}^2] \times L/E [\text{km/GeV}] \right)$$



# Neutrino Oscillations

JHEP 09 (2020) 178

- There are 3 neutrinos, so things are a bit more complicated
- Two independent differences in mass-squared ( $\Delta m_{21}^2$ ,  $\Delta m_{32}^2$ )
- 3 mixing angles ( $\theta_{12}$ ,  $\theta_{13}$ ,  $\theta_{23}$ ) and 1 CPV phase  $\delta_{CP}$



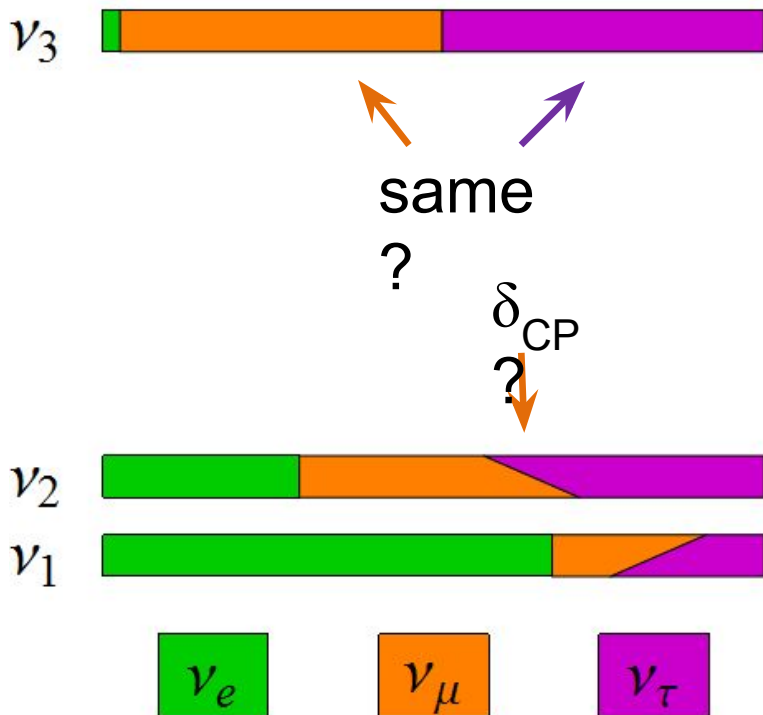
# Missing Pieces

symmetries

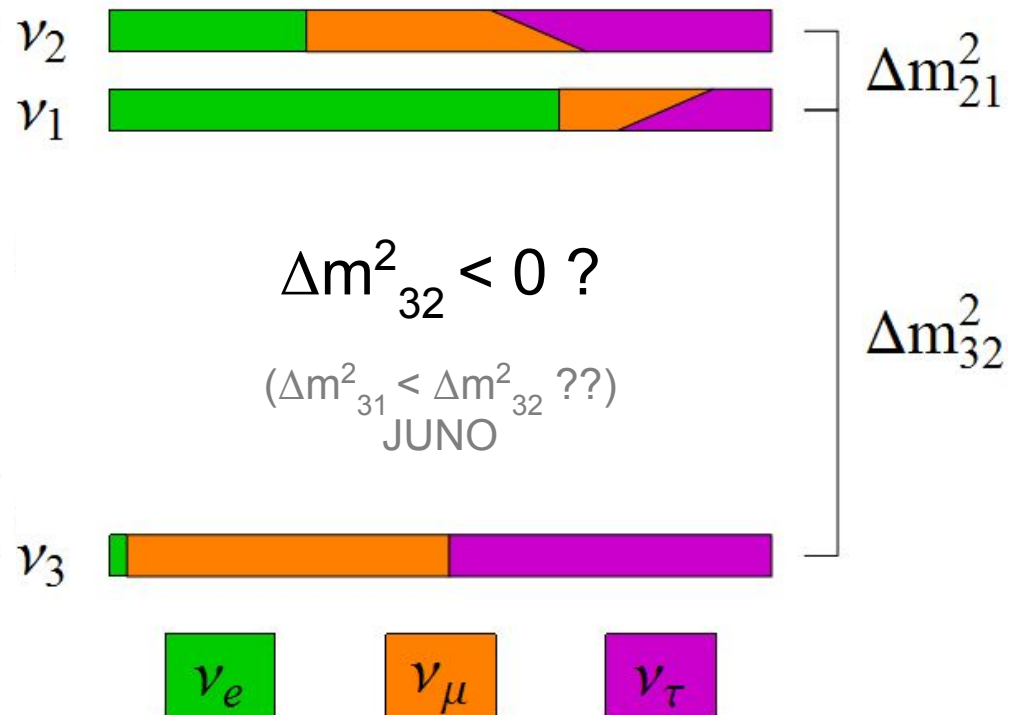
$$\sin^2 2\theta \times \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

- Is  $\theta_{23} = \pi/4$ ? Underlying symmetry?
- Do neutrinos violate CP? ( $\delta_{CP}$ )
- **What is the mass ordering? (Mass Hierarchy)**

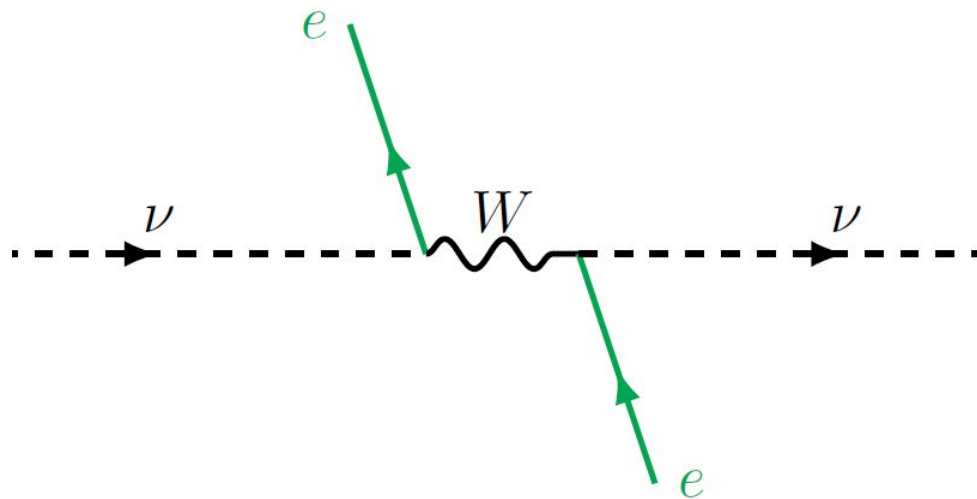
Normal Ordering



Inverted Ordering

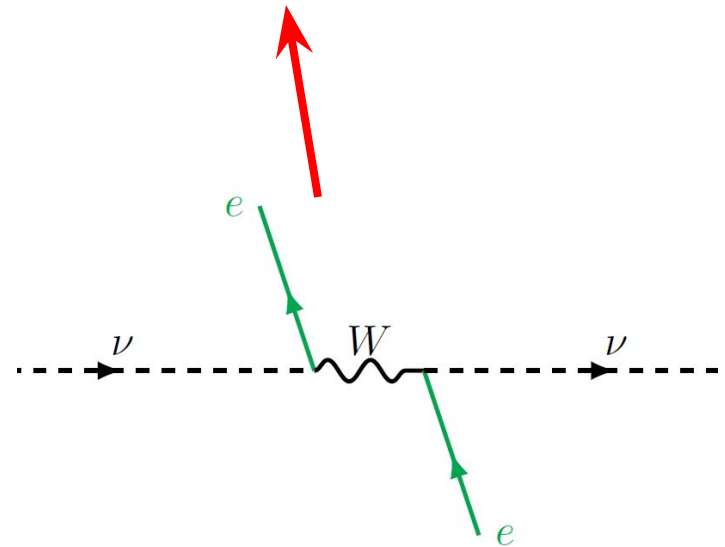
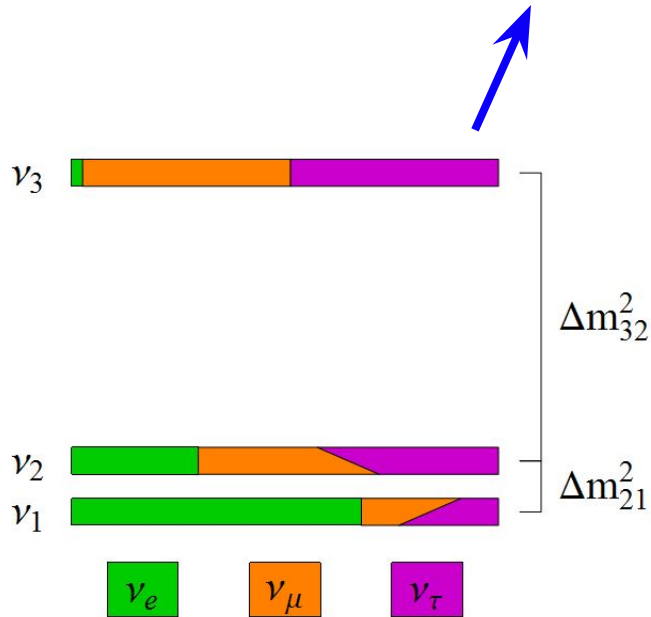


# One weird trick to measure the Mass Ordering...

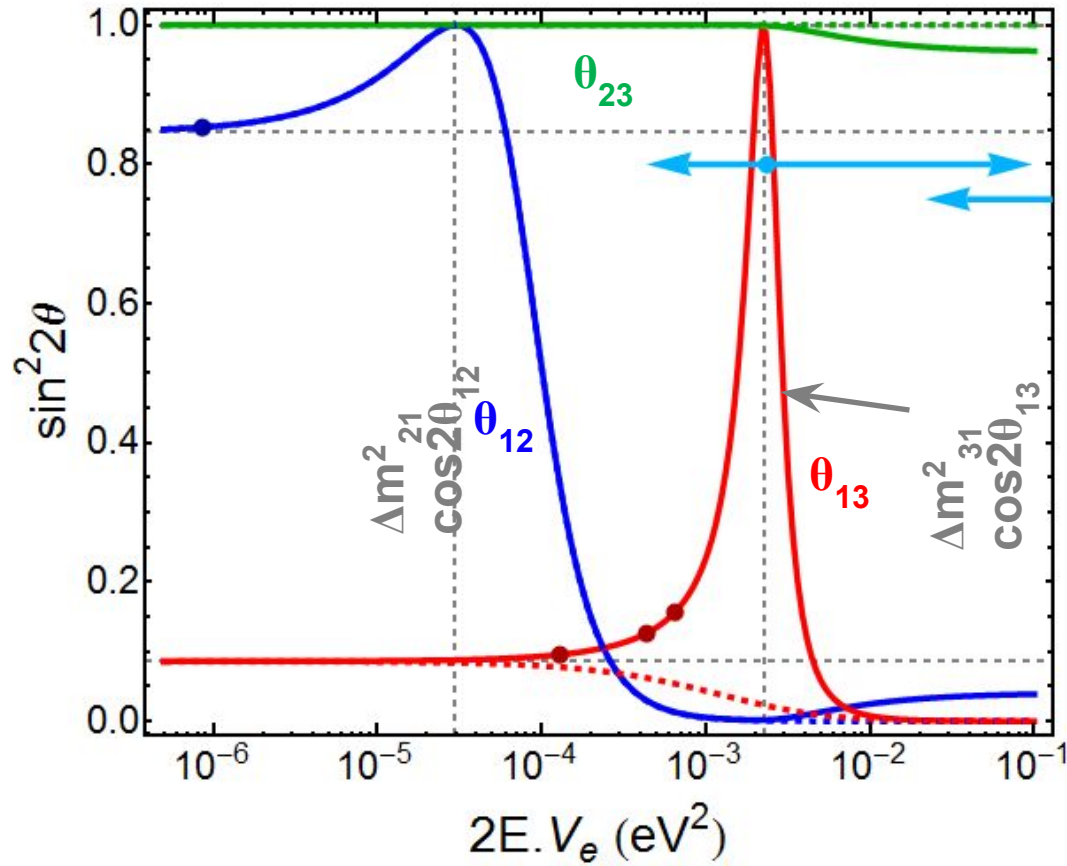


# Matter Effects

$$H_{eff} = U \overbrace{\begin{bmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} \end{bmatrix}}^{H_0} U^\dagger + \overbrace{V_e \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}}^V$$

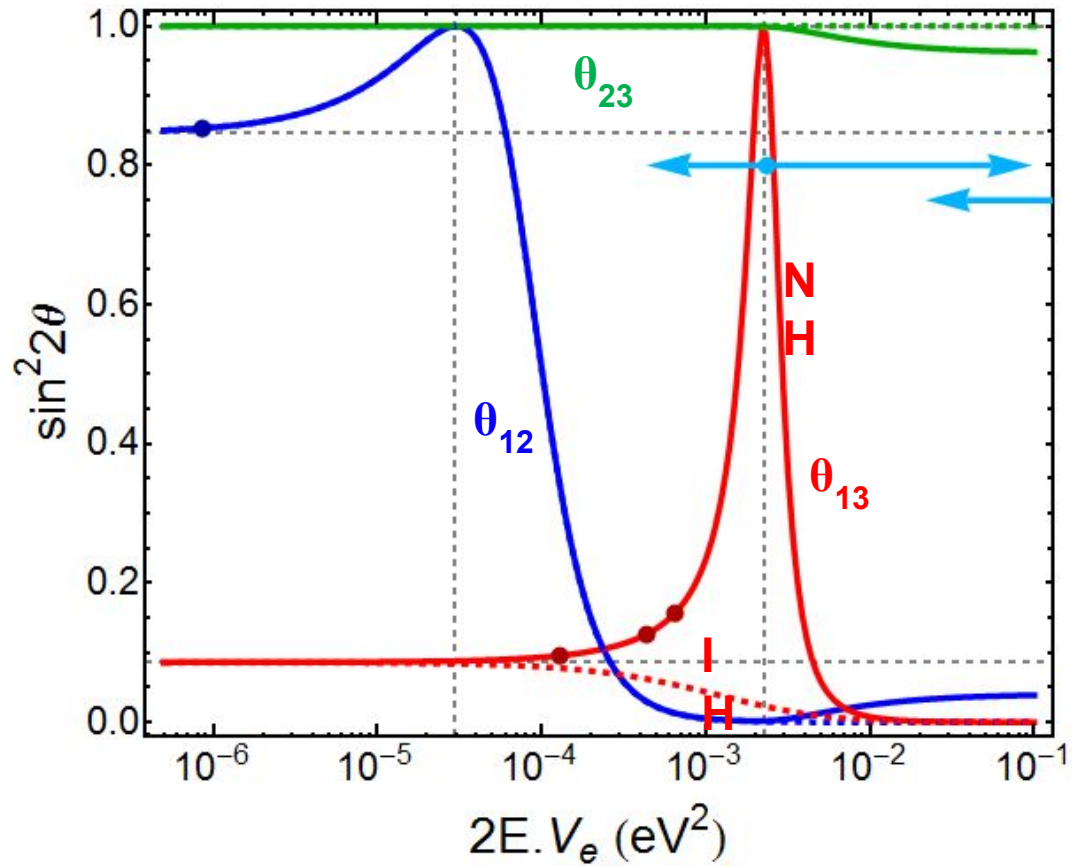


# Resonances



$$H_{eff} = U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m^2_{21}}{2E} & 0 \\ 0 & 0 & \frac{\Delta m^2_{31}}{2E} \end{bmatrix} U^\dagger + V_e \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}$$

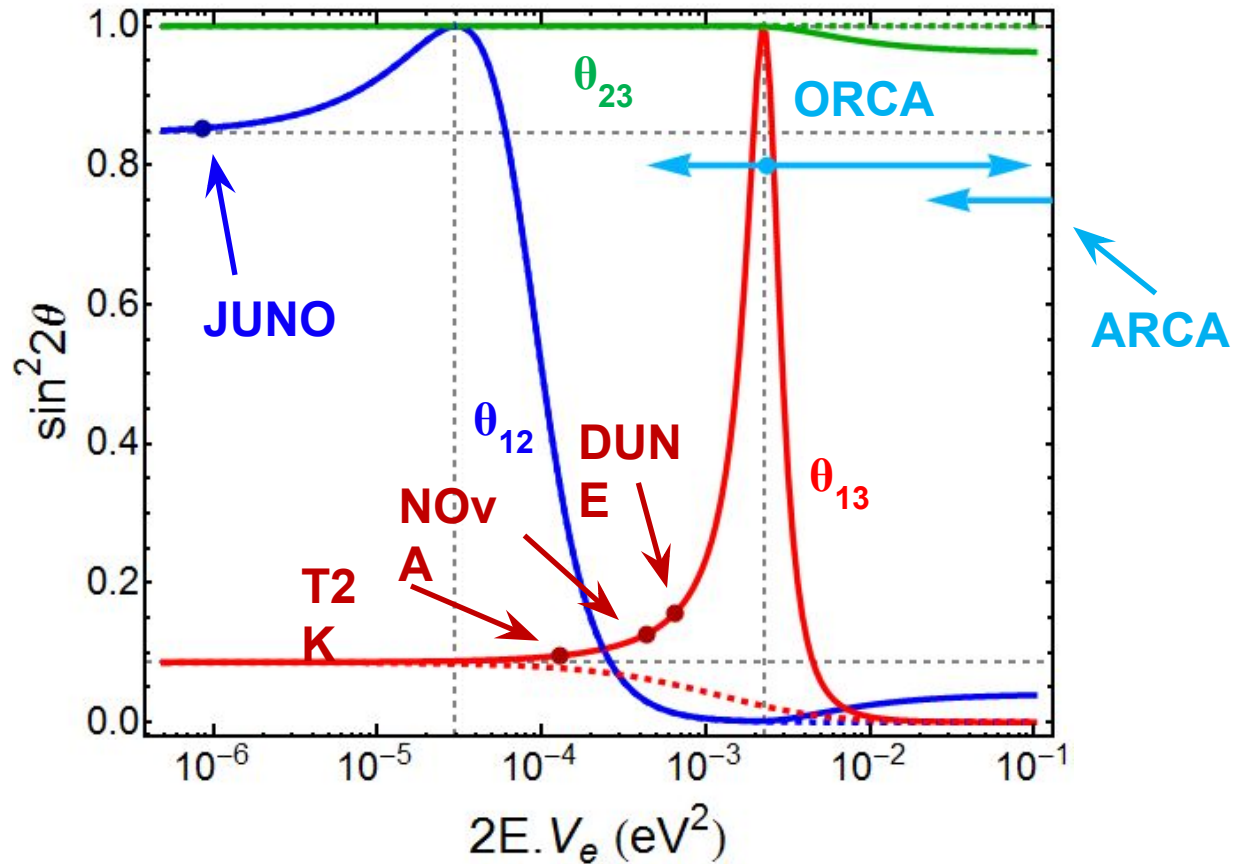
# Resonances



$$H_{eff} = U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} \end{bmatrix} U^\dagger + V_e \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}$$

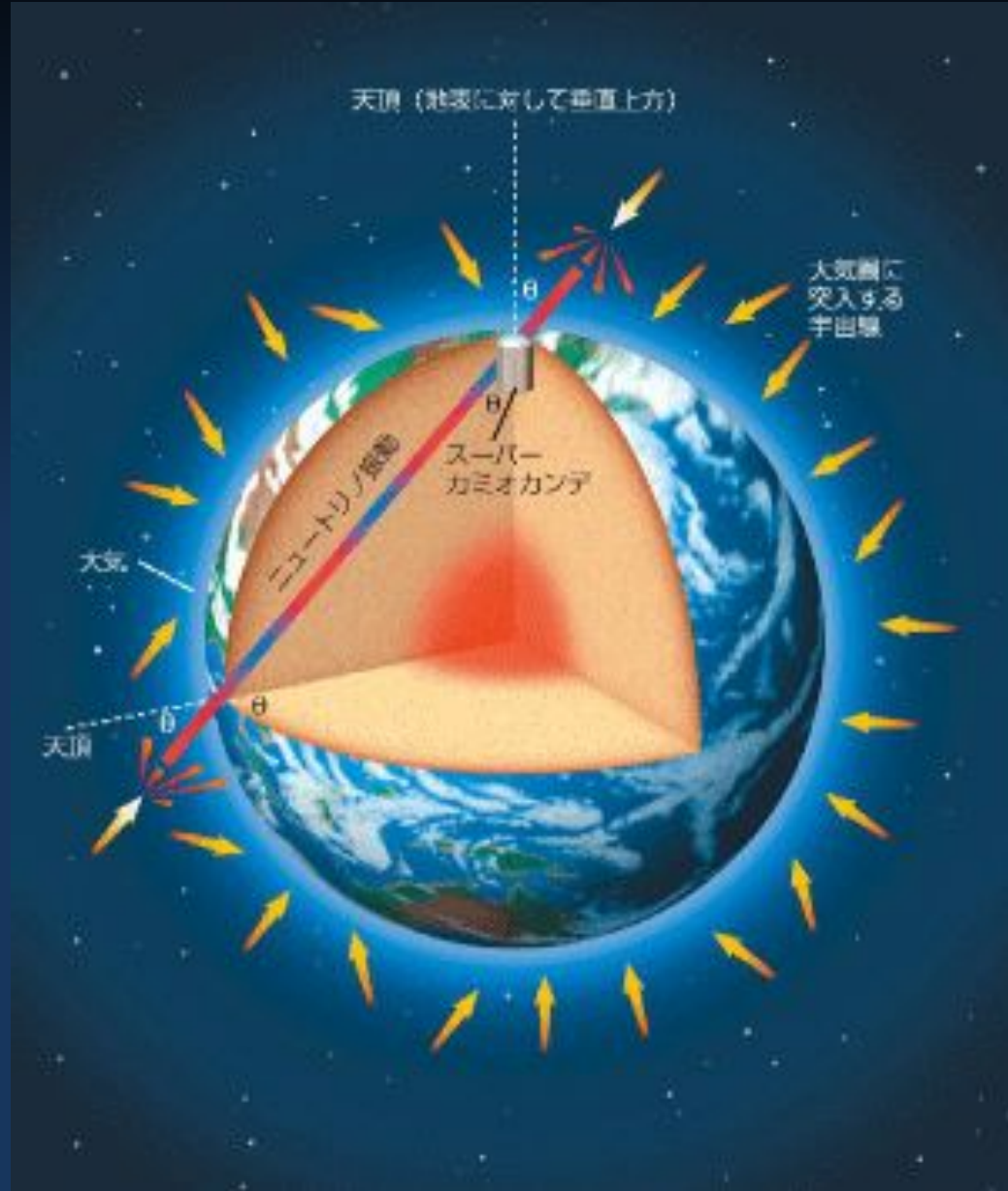


# Resonances

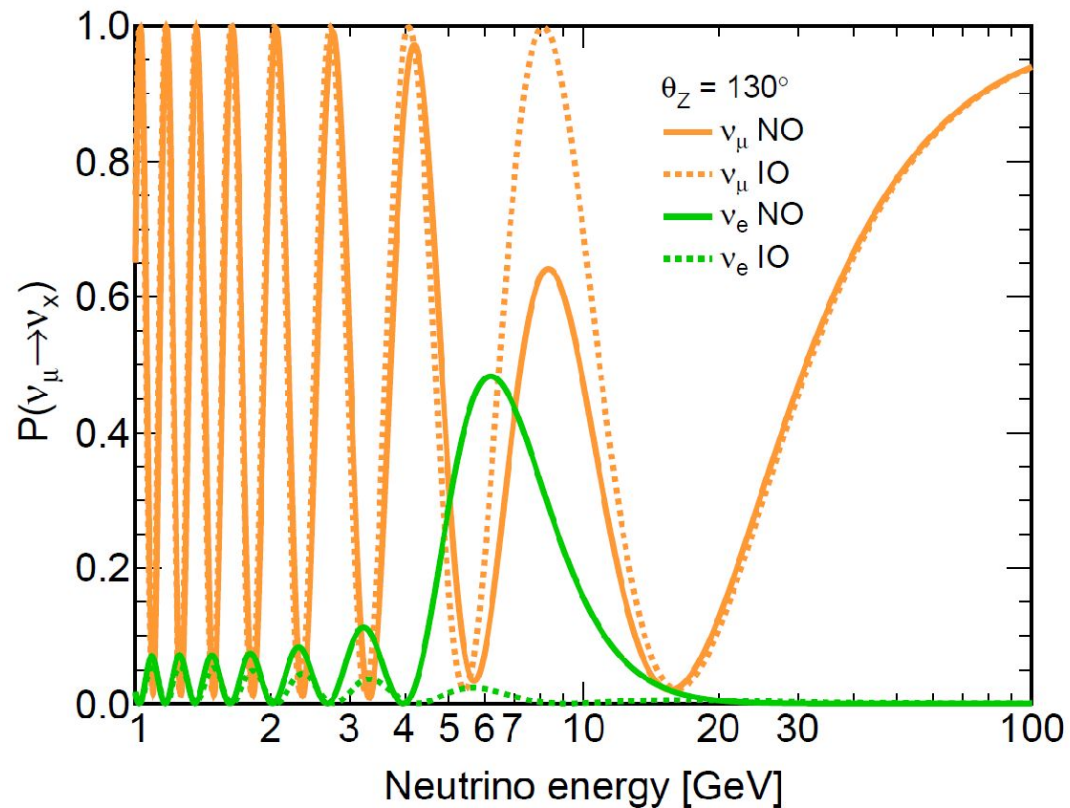
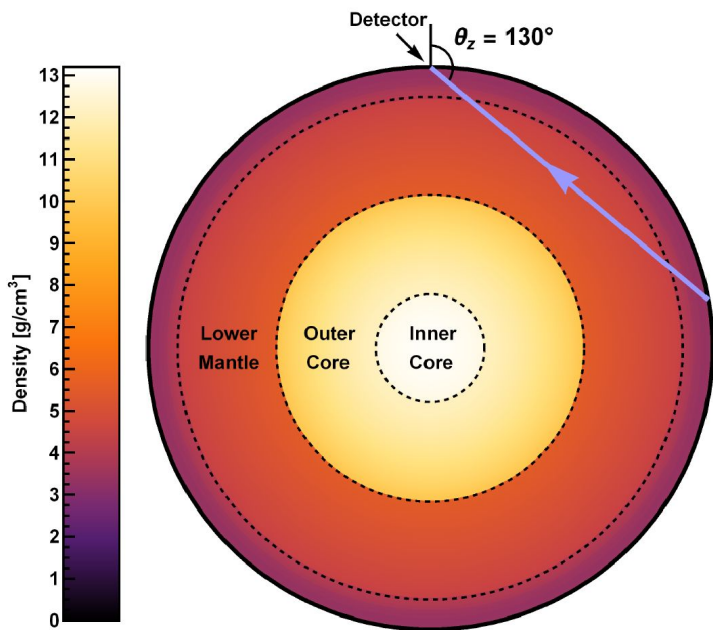


$$H_{eff} = U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} \end{bmatrix} U^\dagger + V_e \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}$$

# Atmospheric Neutrinos



# Sensitive to Mass Ordering at resonance energies



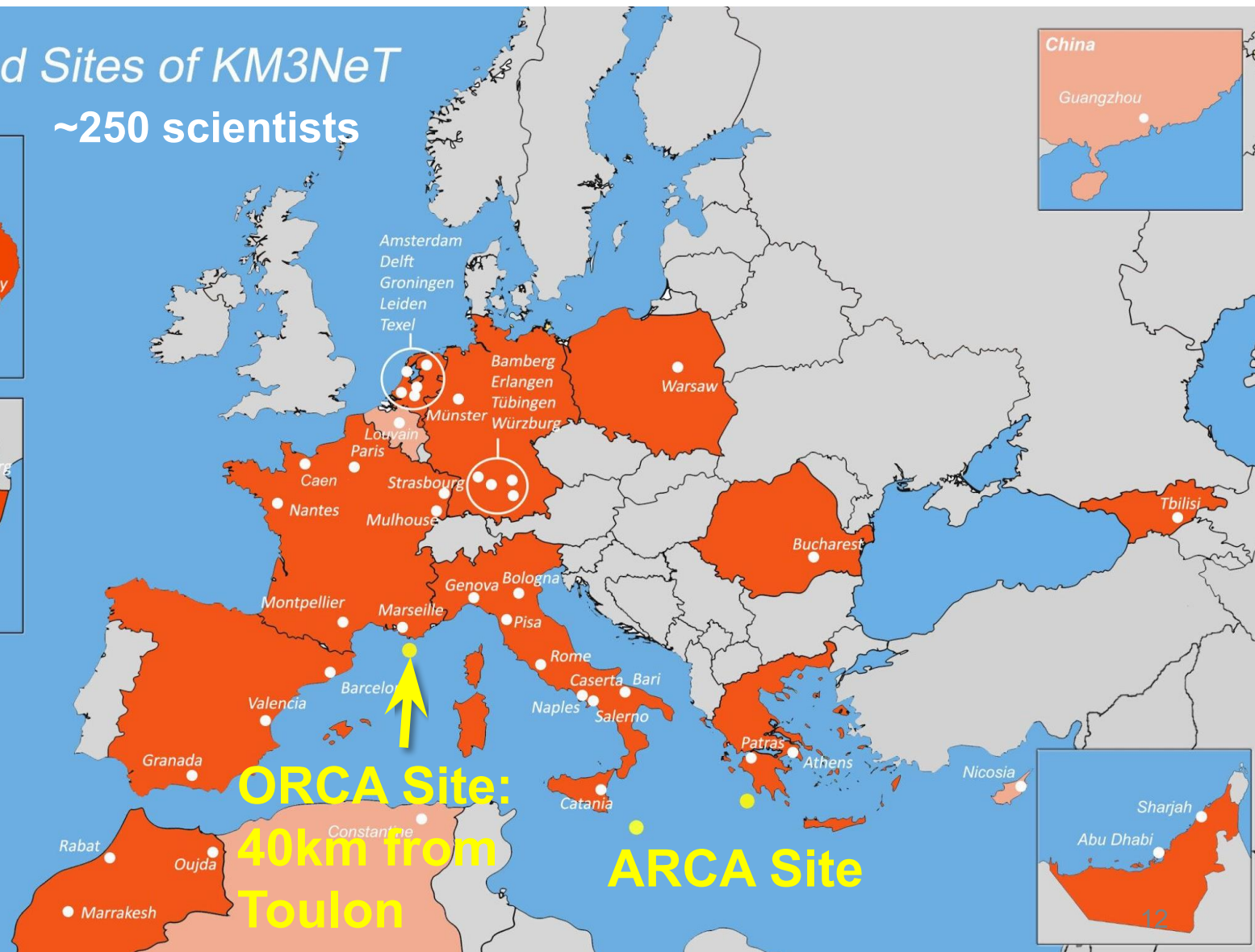


# KM3NeT Collaboration

Letter of Intent: J.Phys.G 43 (2016) 8, 084001

Cities and Sites of KM3NeT

~250 scientists



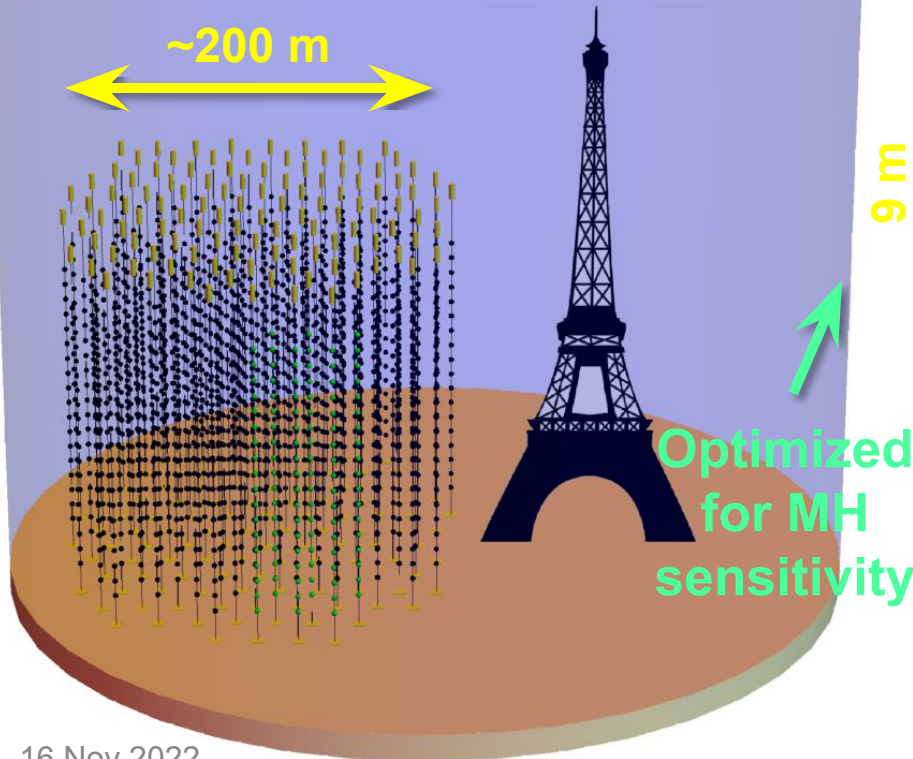
**ORCA Site:**  
40km from  
**Toulon**

**ARCA Site**

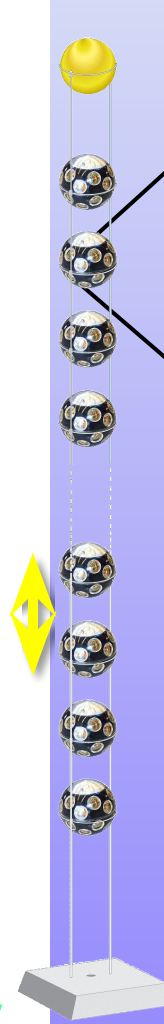
# The ORCA Detector

- **~7 Mt** instrumented
- **115** strings
- **18** DOMs / str
- **31** PMTs / DOM
- Total: **64k PMTs**

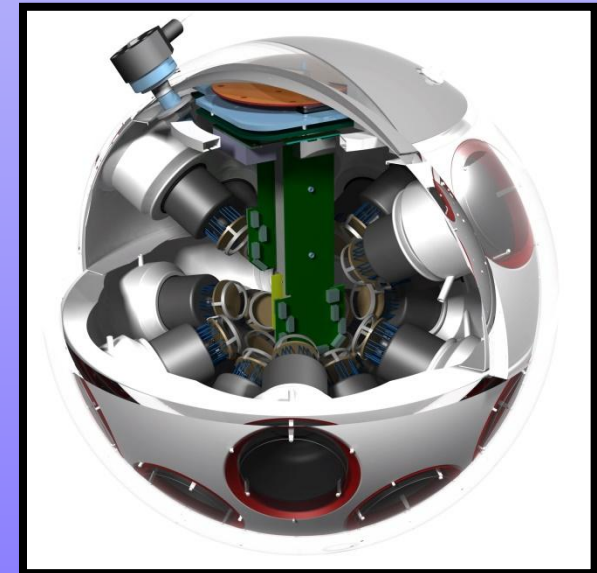
~200 m



9 m

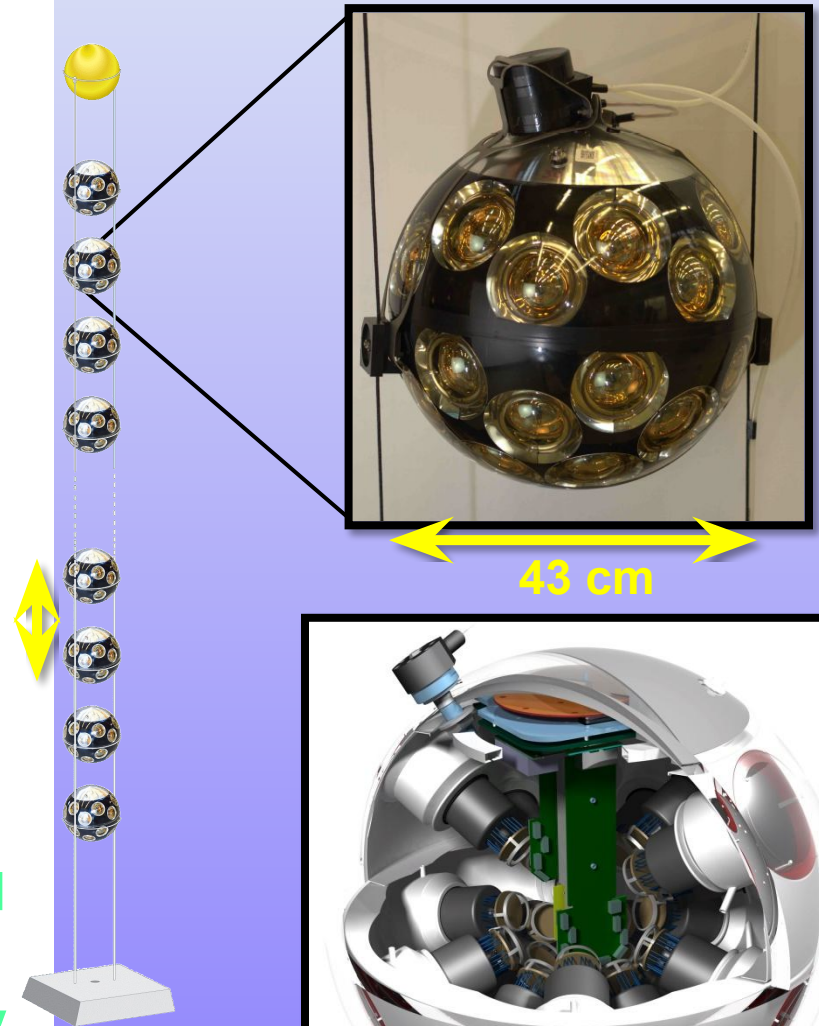


43 cm



# The ORCA Detector

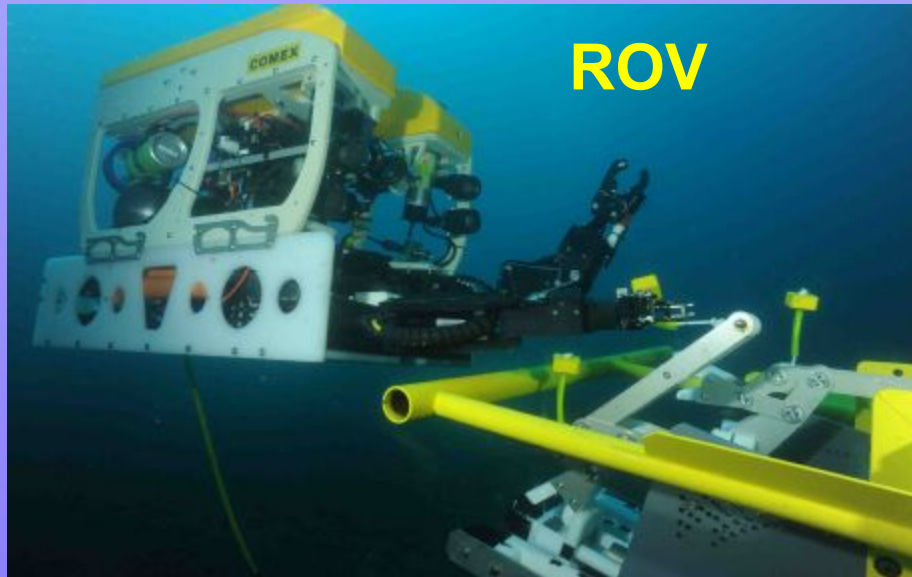
- **~7 Mt** instrumented
- **115 strings**
- **18 DOMs / str**
- **31 PMTs / DOM**
- Total: **64k PMTs**



# Deployment

- String fitted to Launcher Vehicle
- Delivered at a **depth of 2450m**
- ROV connects cable to junction box
- Boat triggers unfurling of the string
- Many nice videos on youtube channel:

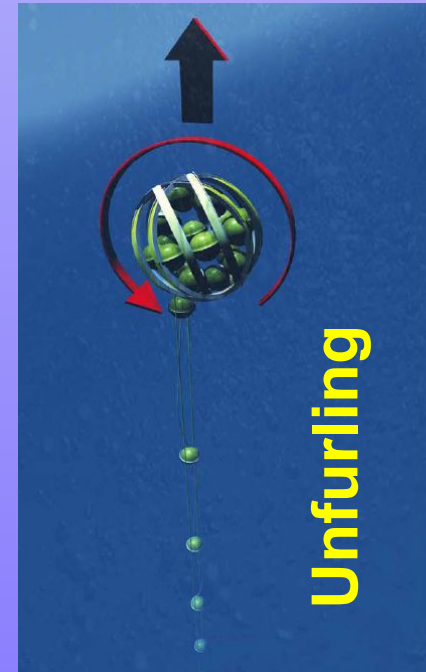
<https://www.youtube.com/user/KM3NeTneutrino>



ROV

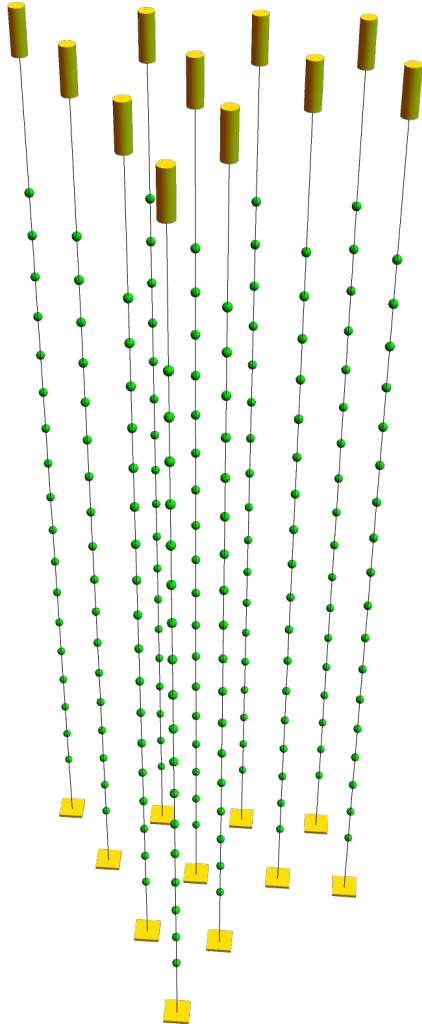
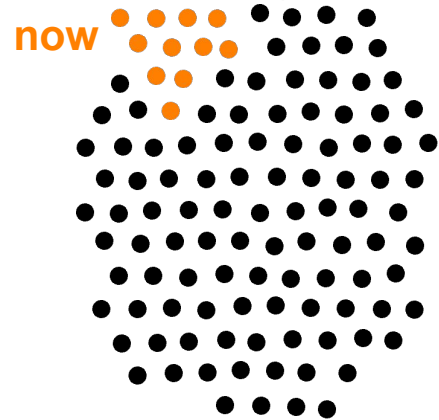


KM3NeT



Unfurling

# Production ongoing around the world

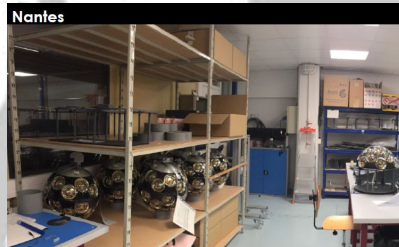


- **ORCA: 11 lines in the water since September 2022**
- **~20 more lines expected by end of 2023 to be deployed**

Amster  
dam



Nantes



Ath  
ens



Bolo  
gna



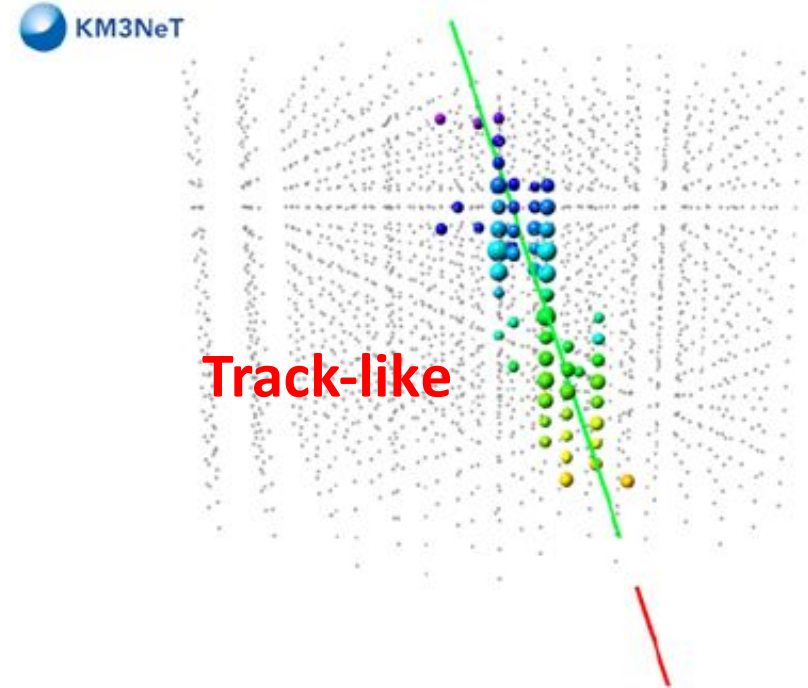
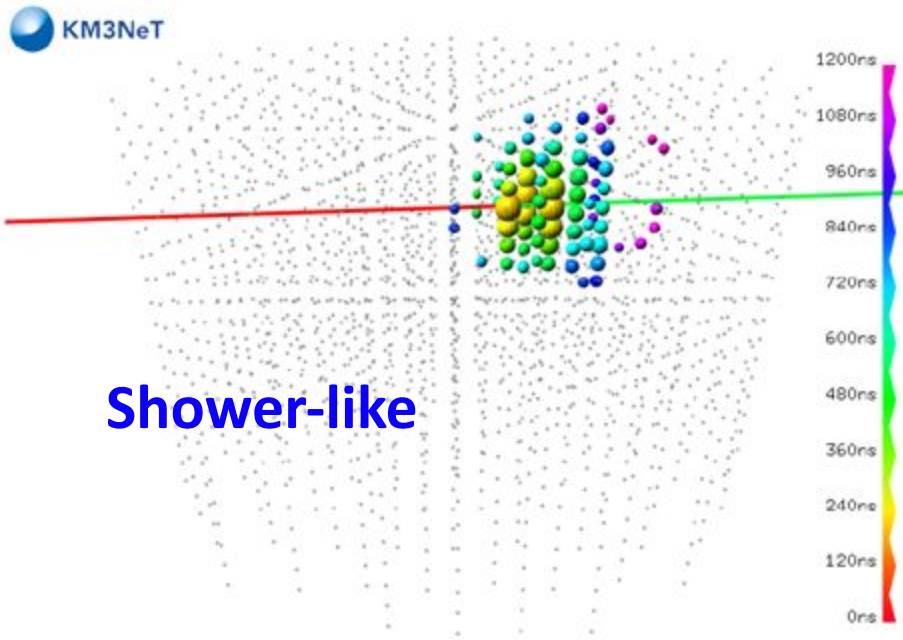
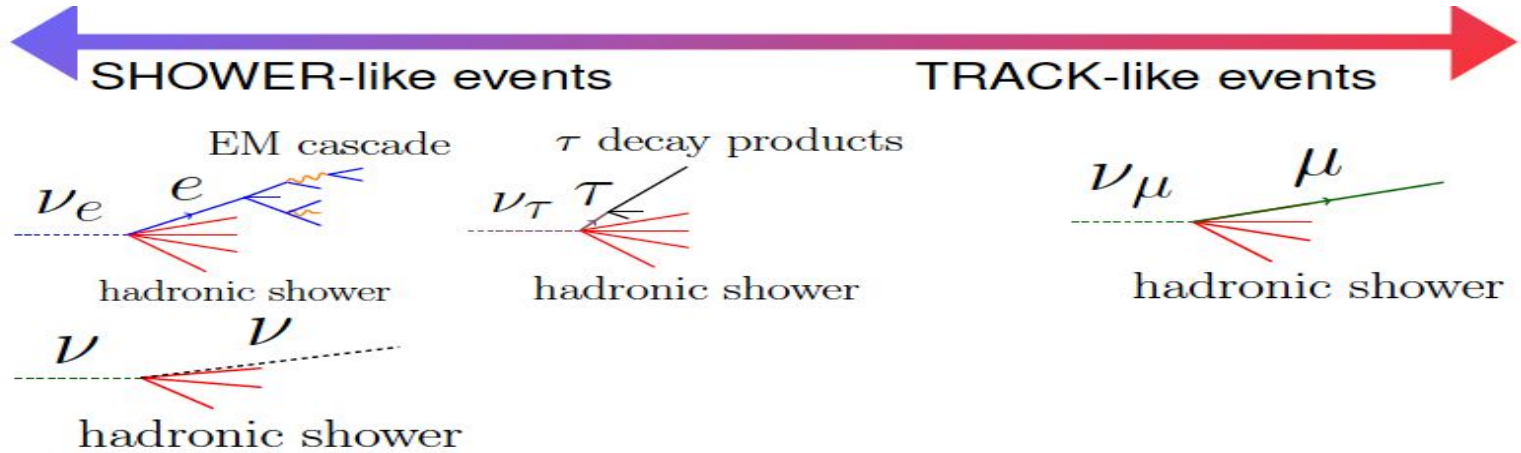
Nap  
les





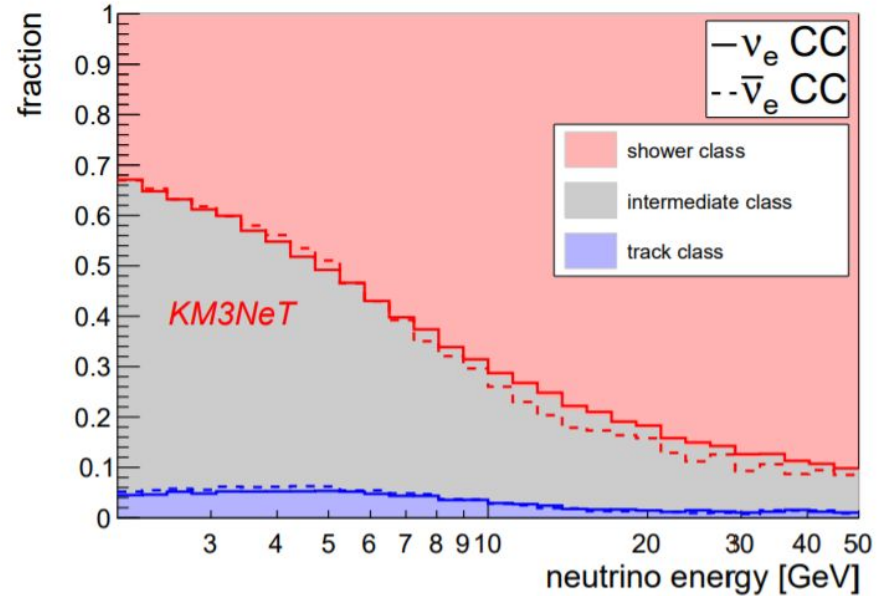
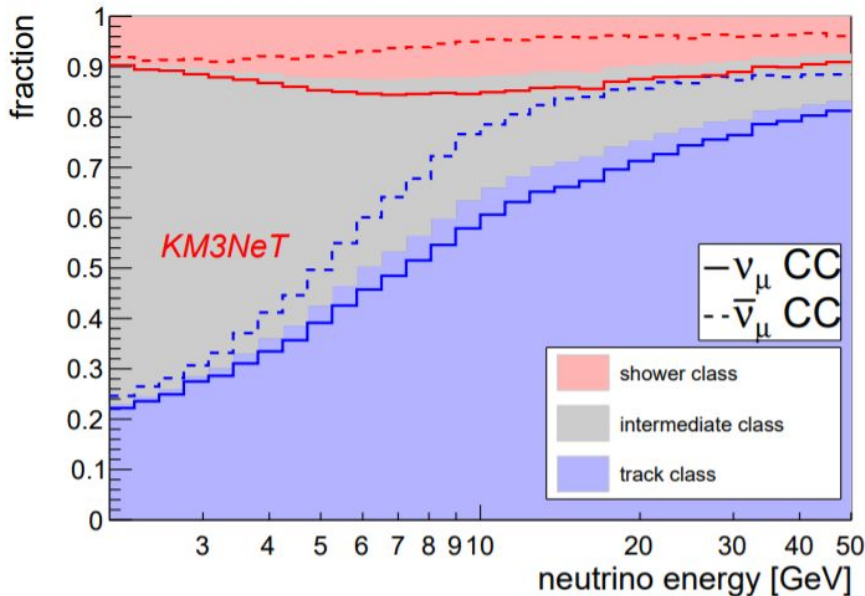
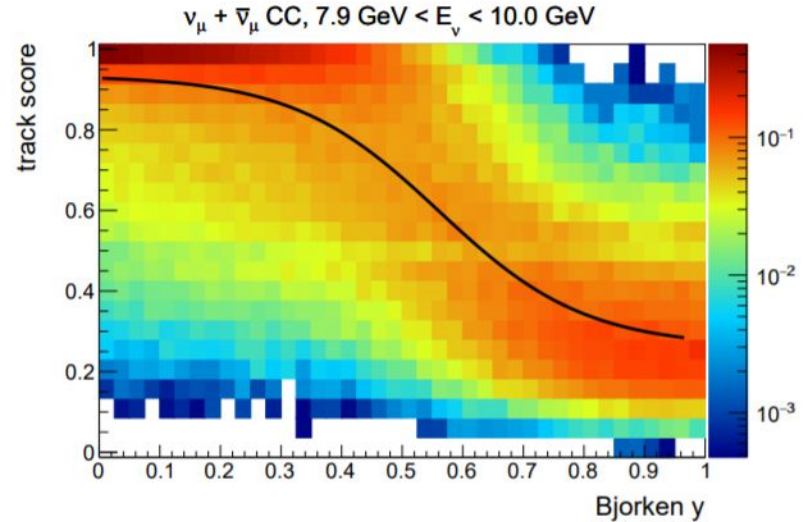
# Sensitivities

# Measuring Neutrinos



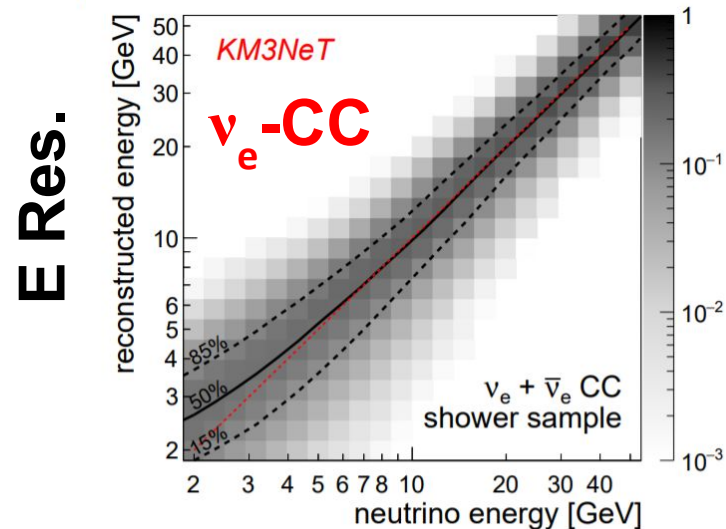
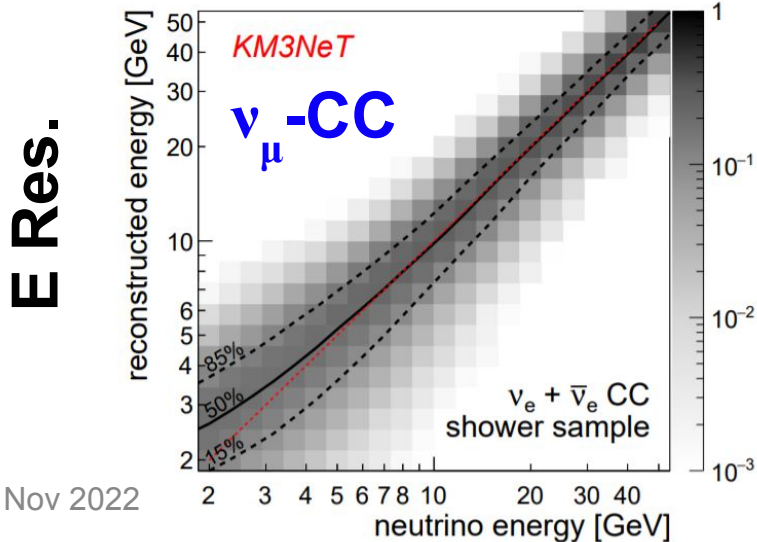
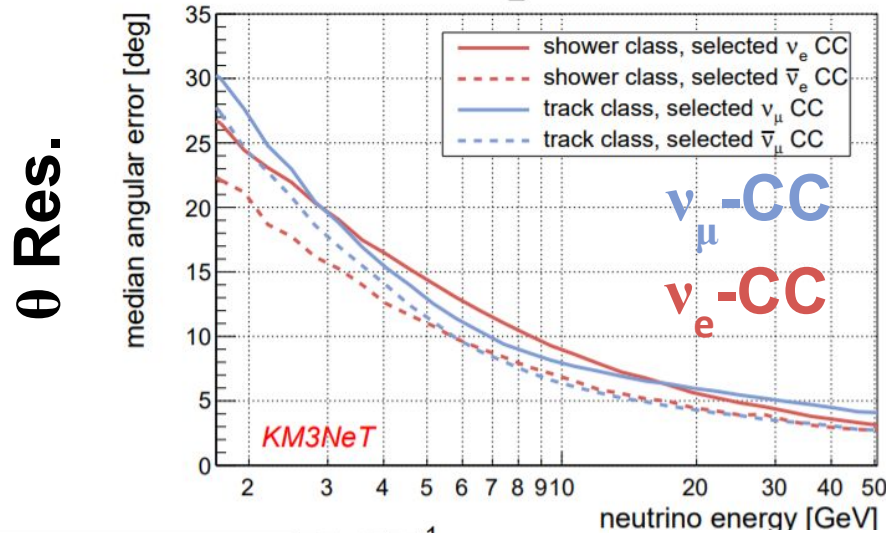
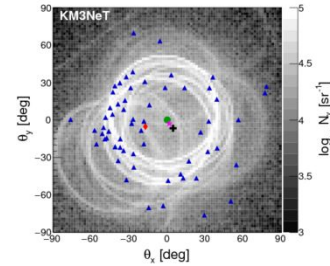
# Event Selection

- Split non-track events into shower class (track score < 0.3) and intermediate class ( $0.3 < \text{track score} < 0.7$ )
- Increases purity in shower class but keeps low purity intermediate sample for control
- $\nu_{\mu}$ -CC contamination mostly from highly inelastic interactions

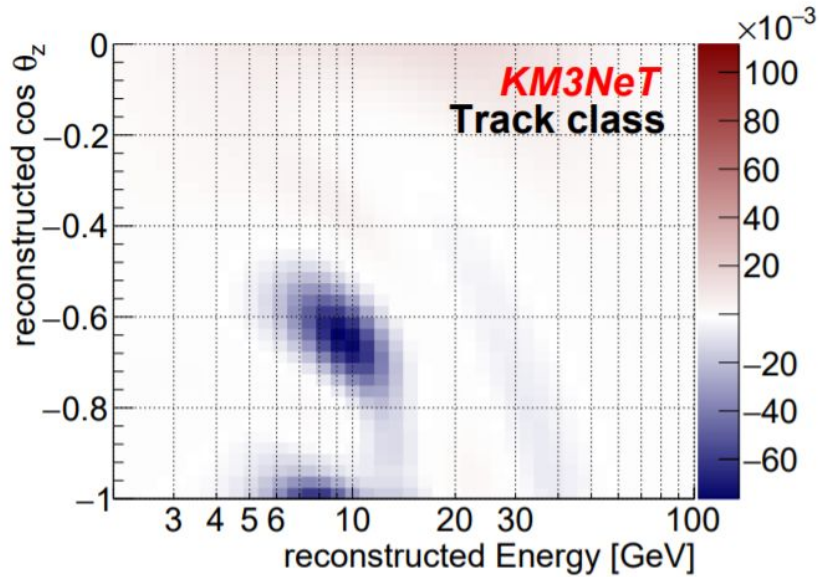


# Reco Performance

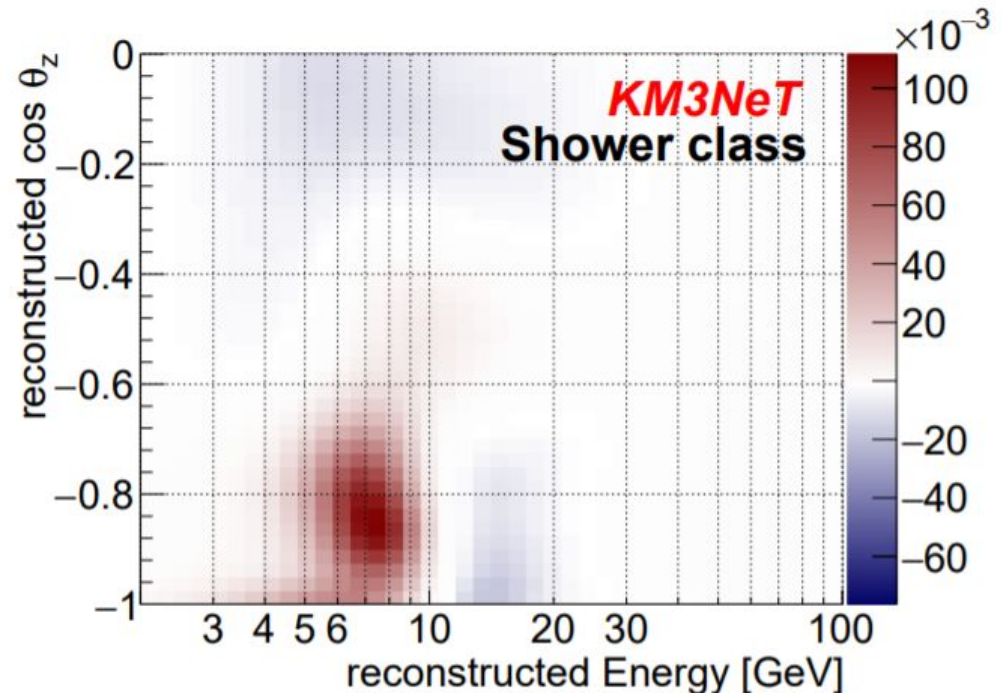
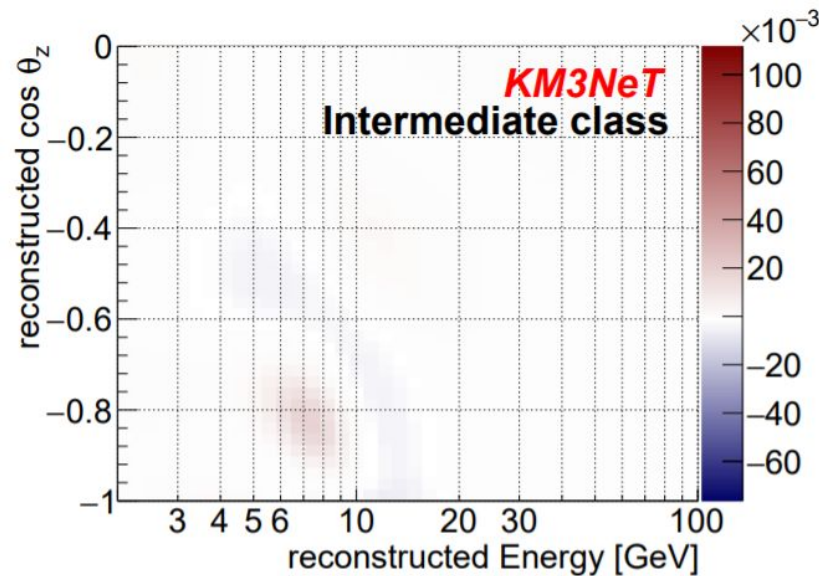
- Energy resolution:  $\sim 25\%$  (Close to limit [arXiv:1612.05621](https://arxiv.org/abs/1612.05621))
- Angular resolution: Better than 15 degrees at relevant energies



# NMO Sensitive Regions



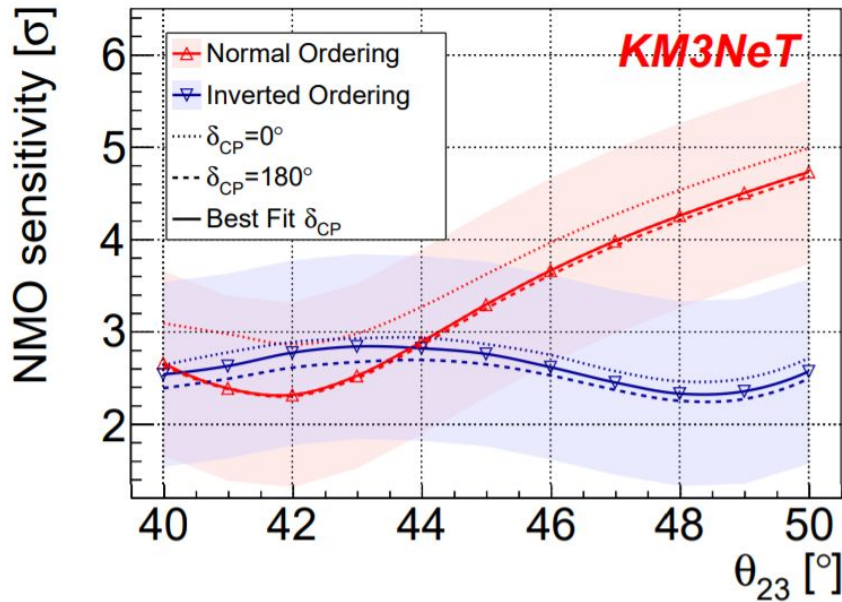
- Matter effects means NMO will impact very specific regions in L and E space
- Sensitivity dominated by shower class, but non-negligible contribution from track class
- Intermediate class has little sensitivity but helps constraint systematics



# Systematics

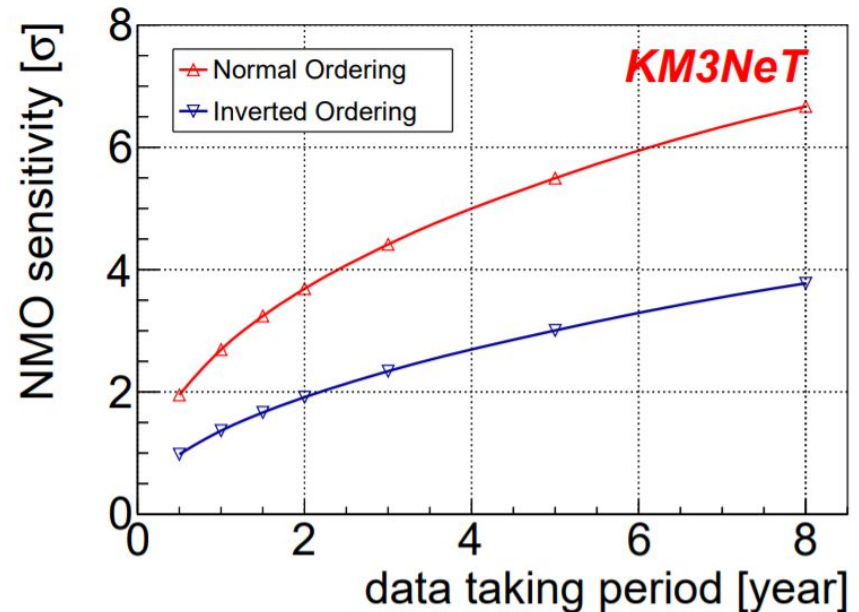
- A total of 12 parameters are included in the analysis to account for shape and normalization uncertainties due to a combination of flux, cross-section and detector systematics
- Many systematics can be left free with no priors and will be constrained by our own data in different regions of L and E
- Ratios between samples:
  - $\nu_e/\text{anti-}\nu_e$ : 7%
  - $\nu_\mu/\text{anti-}\nu_\mu$ : 5%
  - $\nu_e/\nu_\mu$ : 2%
- Normalization components:
  - NC component: Free
  - $\nu_\tau$  component : Free
  - Overall norm. of each class: Free
- Shape uncertainty:
  - Energy scale: 6%
  - Had. energy scale: 5%
  - Directional skew: 2%
  - Spectral index: Free
- Plus 4 oscillation parameters:
  - $\Delta m^2_{31}$ : Free
  - $\delta_{CP}$ : Free
  - $\sin^2\theta_{23}$ : Free
  - $\sin^2\theta_{13}$ : 3%

# Sensitivity Results



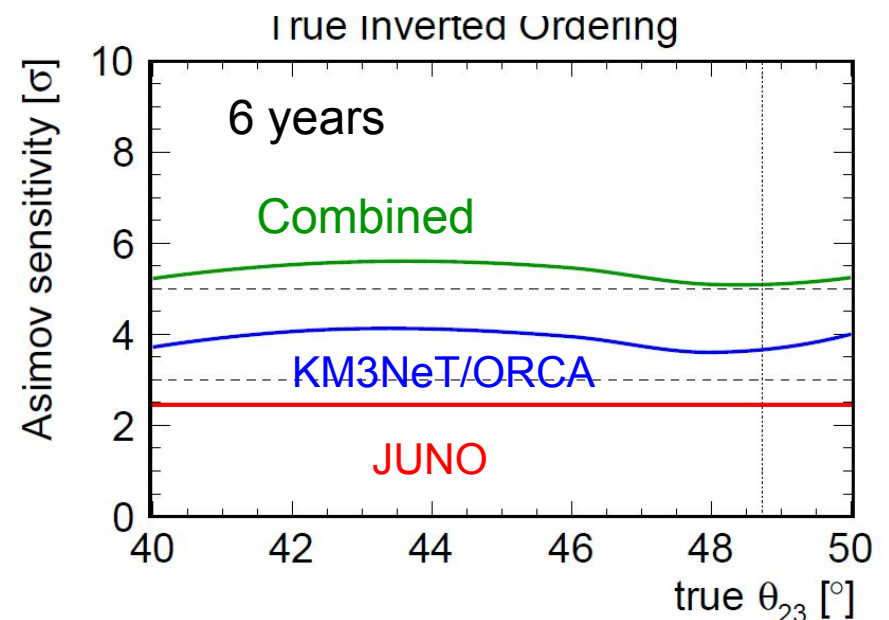
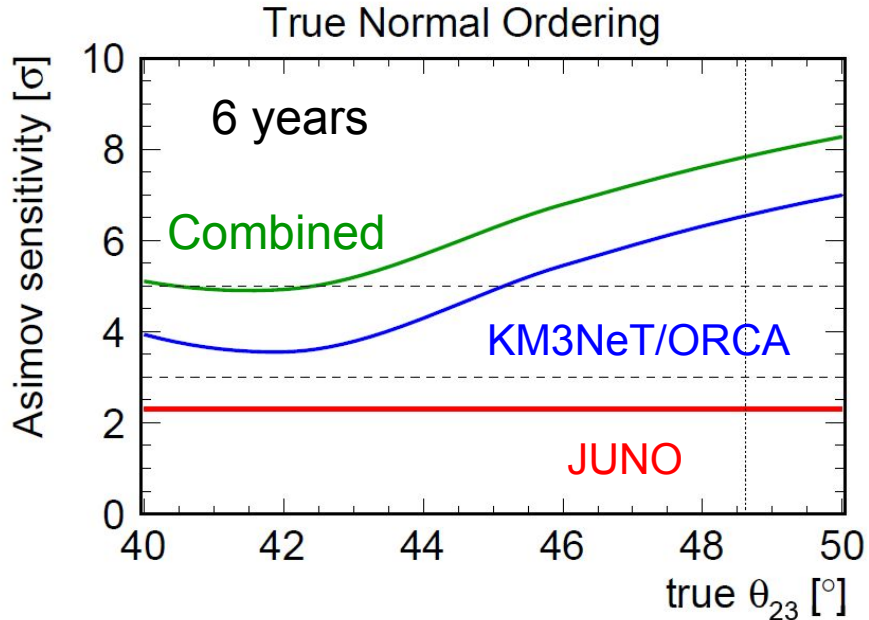
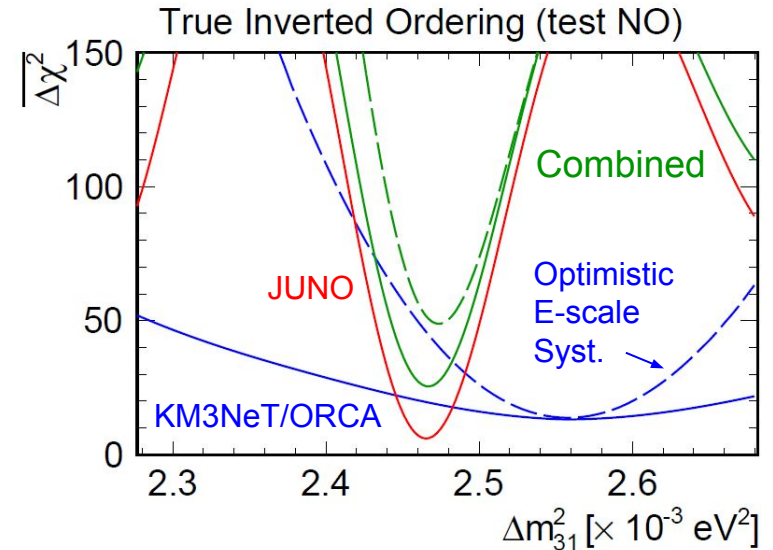
- **In 3 years** expect to achieve **4.4 $\sigma$**  (2.3 $\sigma$ ) if truth is NO (IO) at current best fit oscillation parameters
- **NO and upper octant** of  $\theta_{23}$  significantly enhances sensitivity (**5 $\sigma$  in ~4 years**)

- In the IO, sensitivity is less dependent on octant of  $\theta_{23}$
- **Worst case** scenarios should reach **3 $\sigma$  in ~5 years**
- The value of  $\delta_{cp}$  has small but non-negligible impact on sensitivity



# Synergy with JUNO

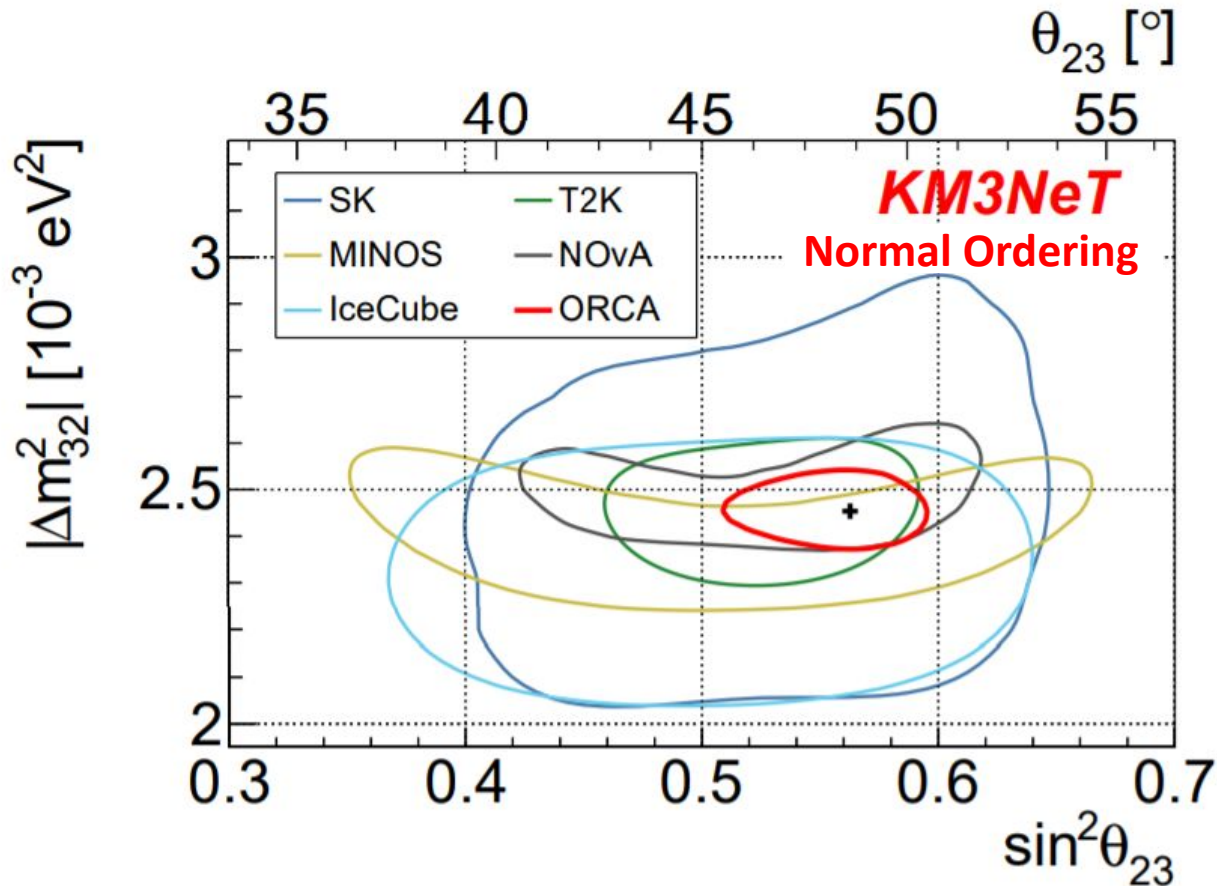
- Before 2030, JUNO and KM3NeT may be able to **definitively measure the NMO at  $5\sigma$**  by exploiting a synergy between measurements of  $\Delta m_{31}^2$
- Combined sensitivity is better than the simple sum of their  $\Delta\chi^2$





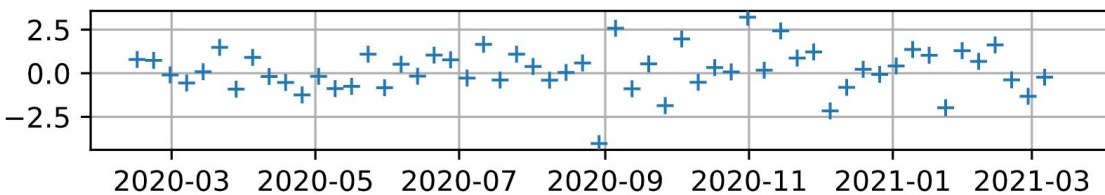
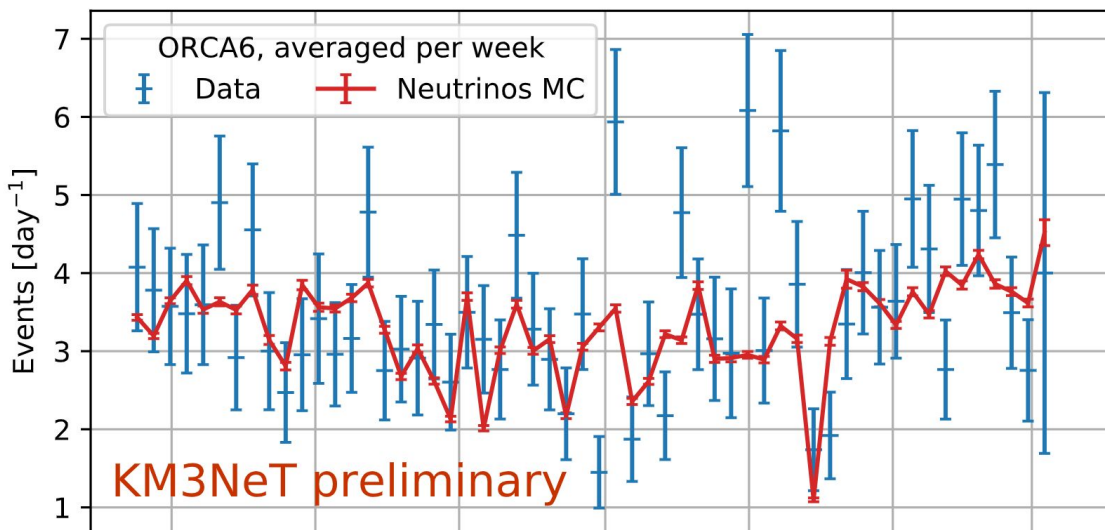
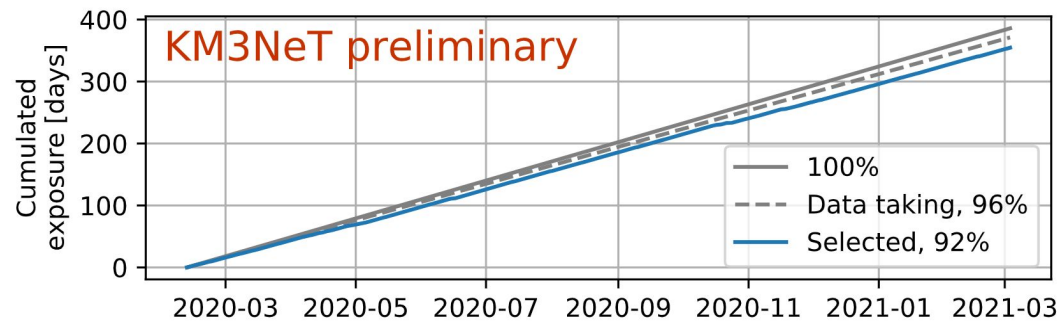
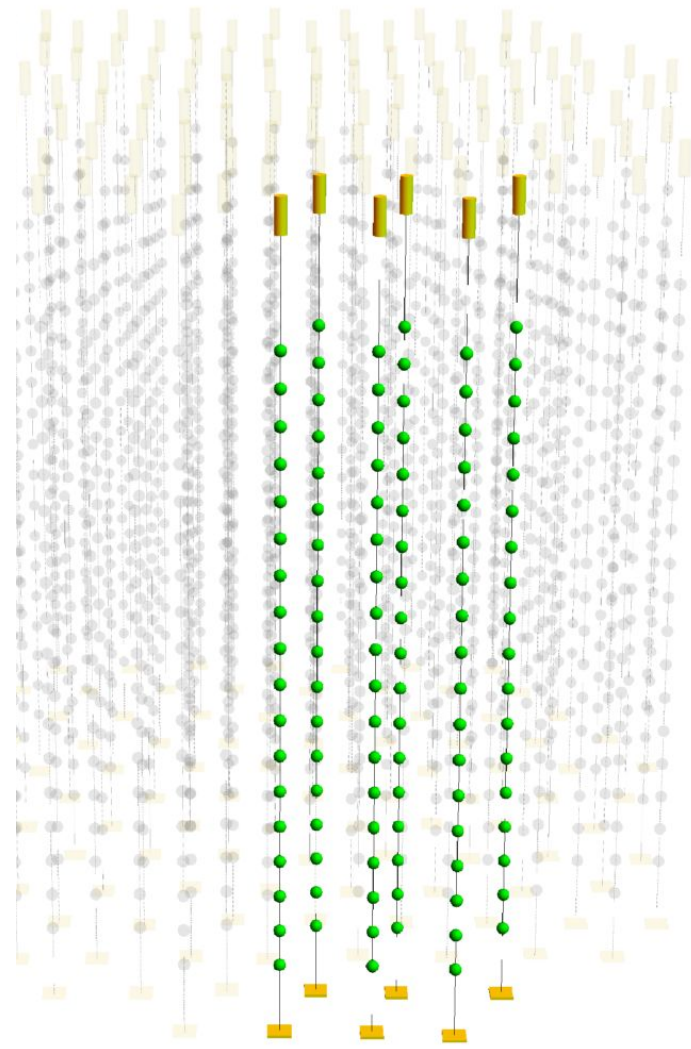
# Other Parameters

- High statistics and excellent resolution → Measure  $\Delta m_{32}^2$  and  $\sin^2\theta_{23}$
- **Competitive sensitivity with LBL experiments**
- Achieve **4.5%** prec. in  $\Delta m_{32}^2$  and **~10%** in  $\sin^2\theta_{23}$



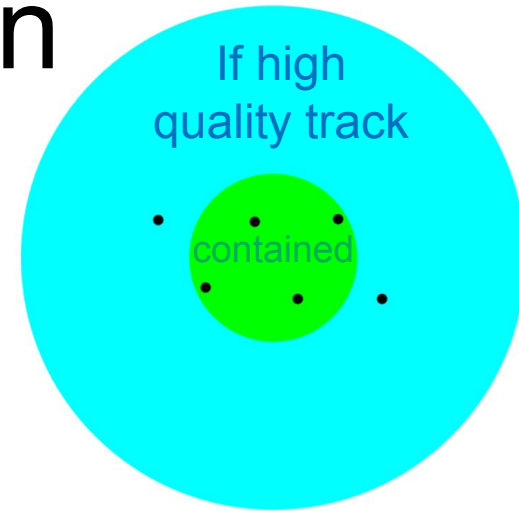
# First Results

# Over 90% up time for 1 year



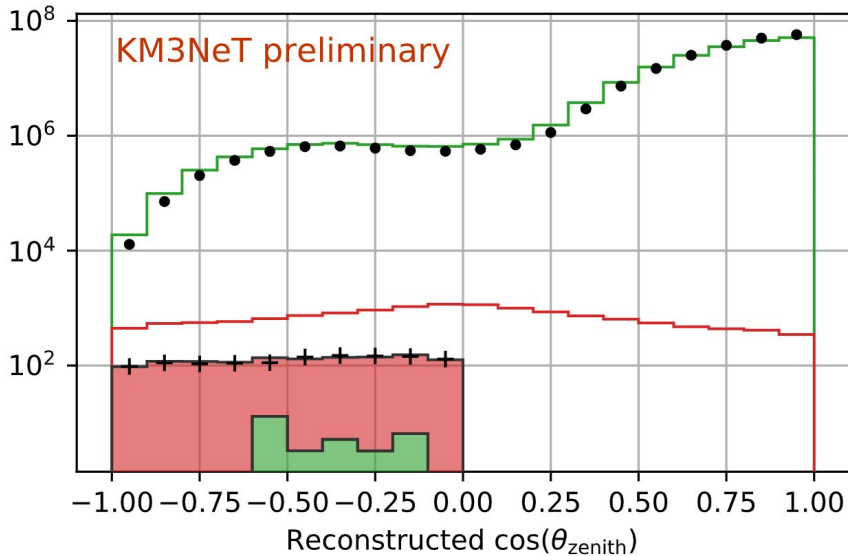
# Neutrino Selection

- Simplified selection for track-like events only
- Minimum number of hits matching Cherenkov angle
- Vertex contained in fiducial volume
- High quality track reconstruction, increasing with radius
- Reject poorly understood high energy tail (dE/dx based)



ORCA6, 354.6 days.

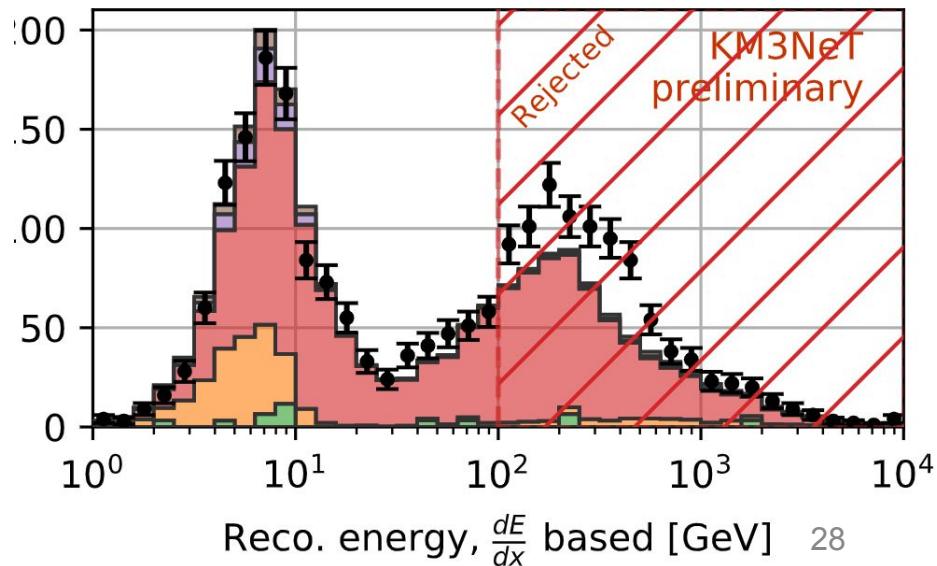
♣	Data	Atm. mu	Atm. nu
+	Data, $\nu$ sel.	Atm. mu, $\nu$ sel.	Atm. nu, $\nu$ sel.



16 Nov 2022

ORCA6, 354.6 days

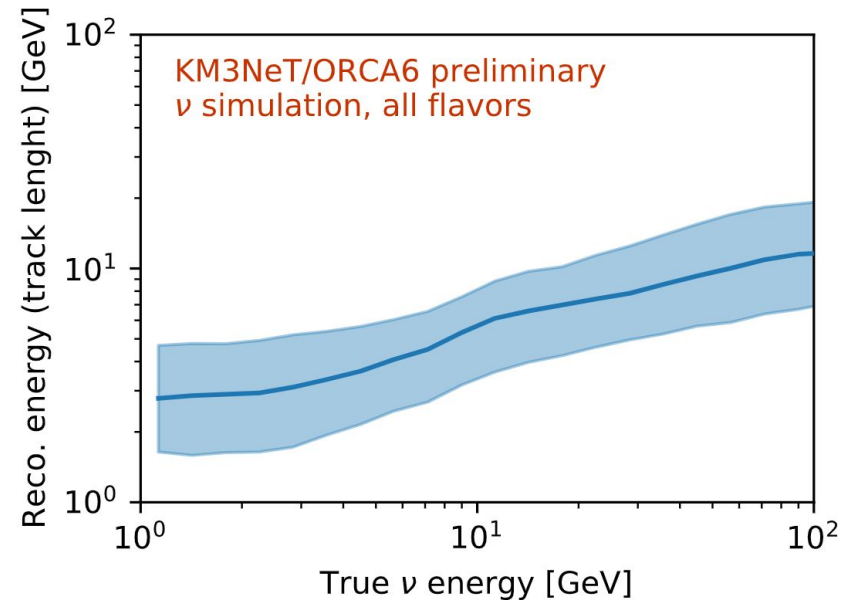
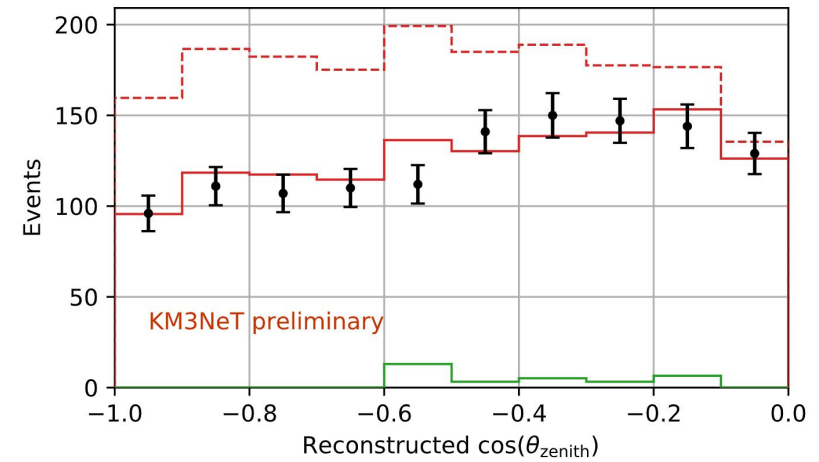
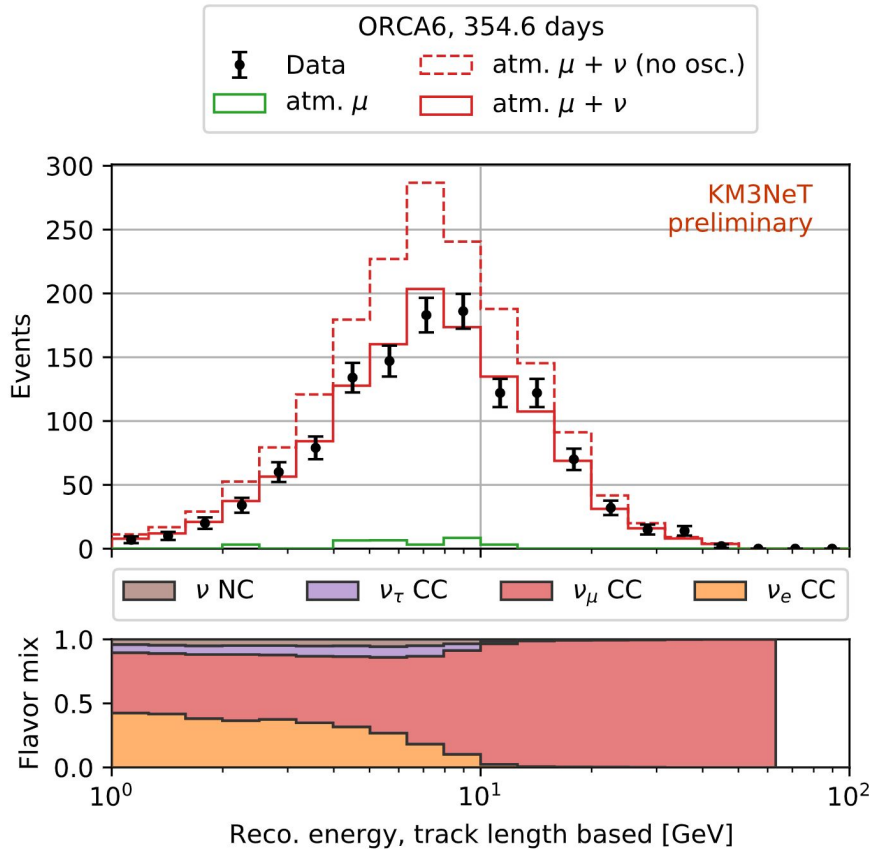
♣	Data	$\nu_e$ CC	$\nu_\tau$ CC
■	Atm. $\mu$	$\nu_\mu$ CC	$\nu$ NC



28

# ORCA6 Data

- Energy estimate based on track length
- Poor energy resolution, but robust
- Good agreement with oscillated MC

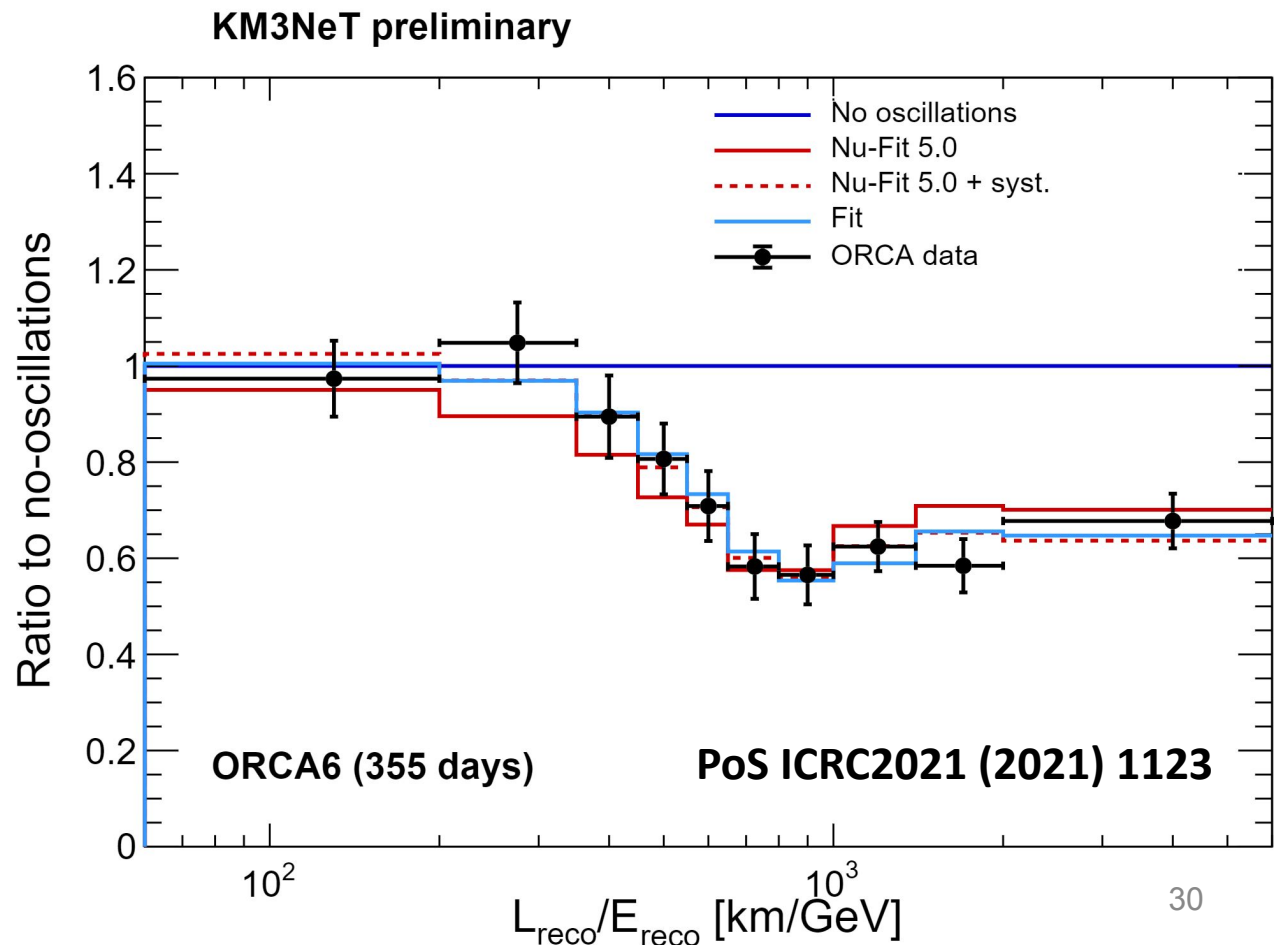


# Clear Evidence of Oscillations

- First oscillation dip already visible even with suboptimal resolution
- **Oscillation hypothesis confirmed at  $5.9\sigma$**
- Data prefers oscillation minimum at higher L/E than expected from global fits
- Our systematics do cover this effect however, so results are consistent ( $1.9\sigma$ )

## Systematics

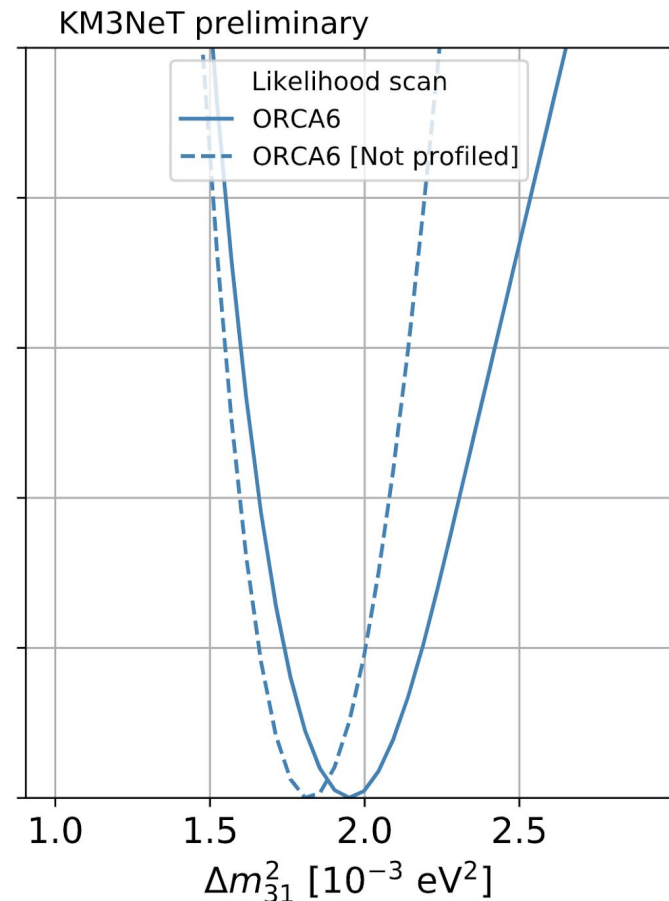
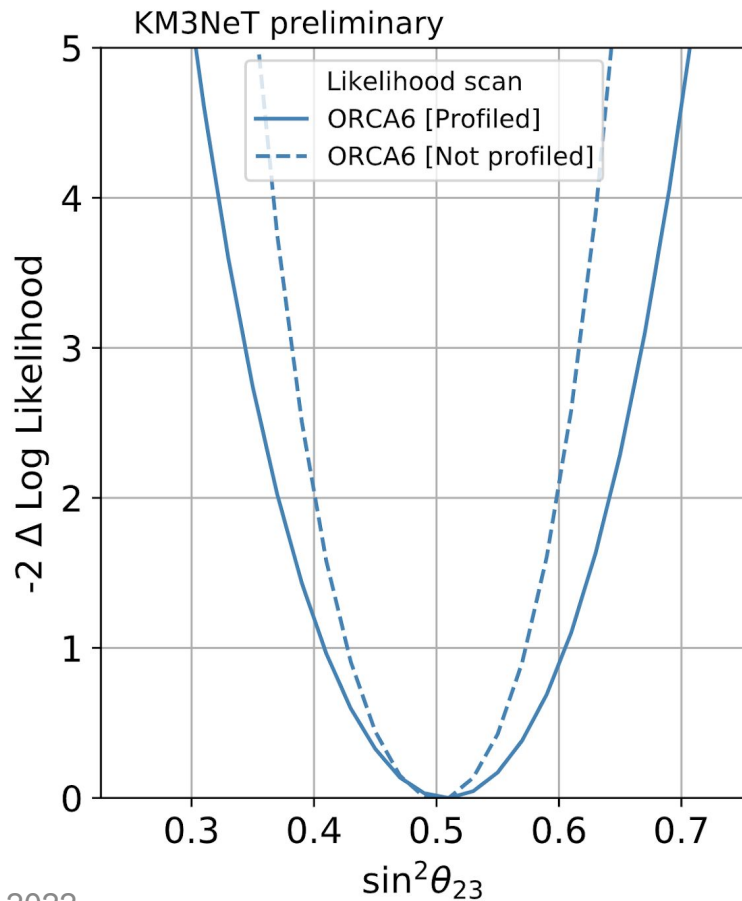
Parameter	Treatment
	Free/Fixed/Prior
$\theta_{12}$ [deg]	Fixed
$\theta_{13}$ [deg]	Fixed
$\theta_{23}$ [deg]	Free
$\Delta m_{31}^2$ [ $10^{-3}$ GeV $^2$ ]	Free
$\Delta m_{21}^2$ [ $10^{-5}$ GeV $^2$ ]	Fixed
$\delta_{CP}$ [deg]	Fixed
Normalisation	Free
Spectral index	Prior: 10%
$n_{\nu_{up}}/n_{\nu_{horiz}}$	Prior: 7%
$n_{\nu_{\mu}}/n_{\nu_{\bar{\mu}}}$	Prior: 10%
$n_{\nu_e}/n_{\nu_{\bar{e}}}$	Prior: 10%
$n_{\nu_{\mu}}/n_{\nu_e}$	Prior: 3%
$n^{NC}$	Prior: 10%
$n_{\tau}^{CC}$	Prior: 20%
Energy scale	Prior: 10%



# First Oscillation Measurement

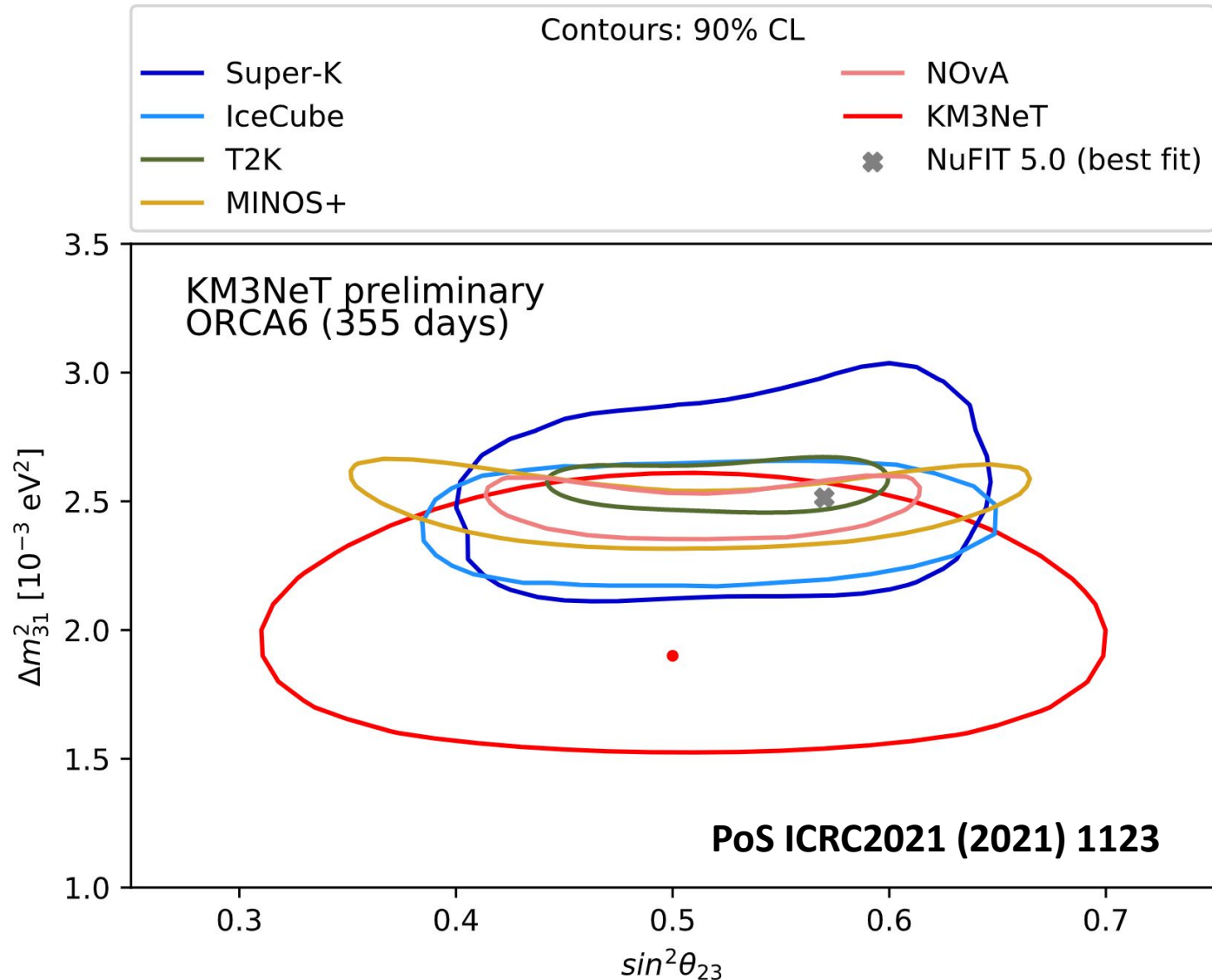
$$\Delta m_{31}^2 = 1.95_{-0.21}^{+0.24} \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} = 0.50 \pm 0.10$$



# KM3NeT is in the game!

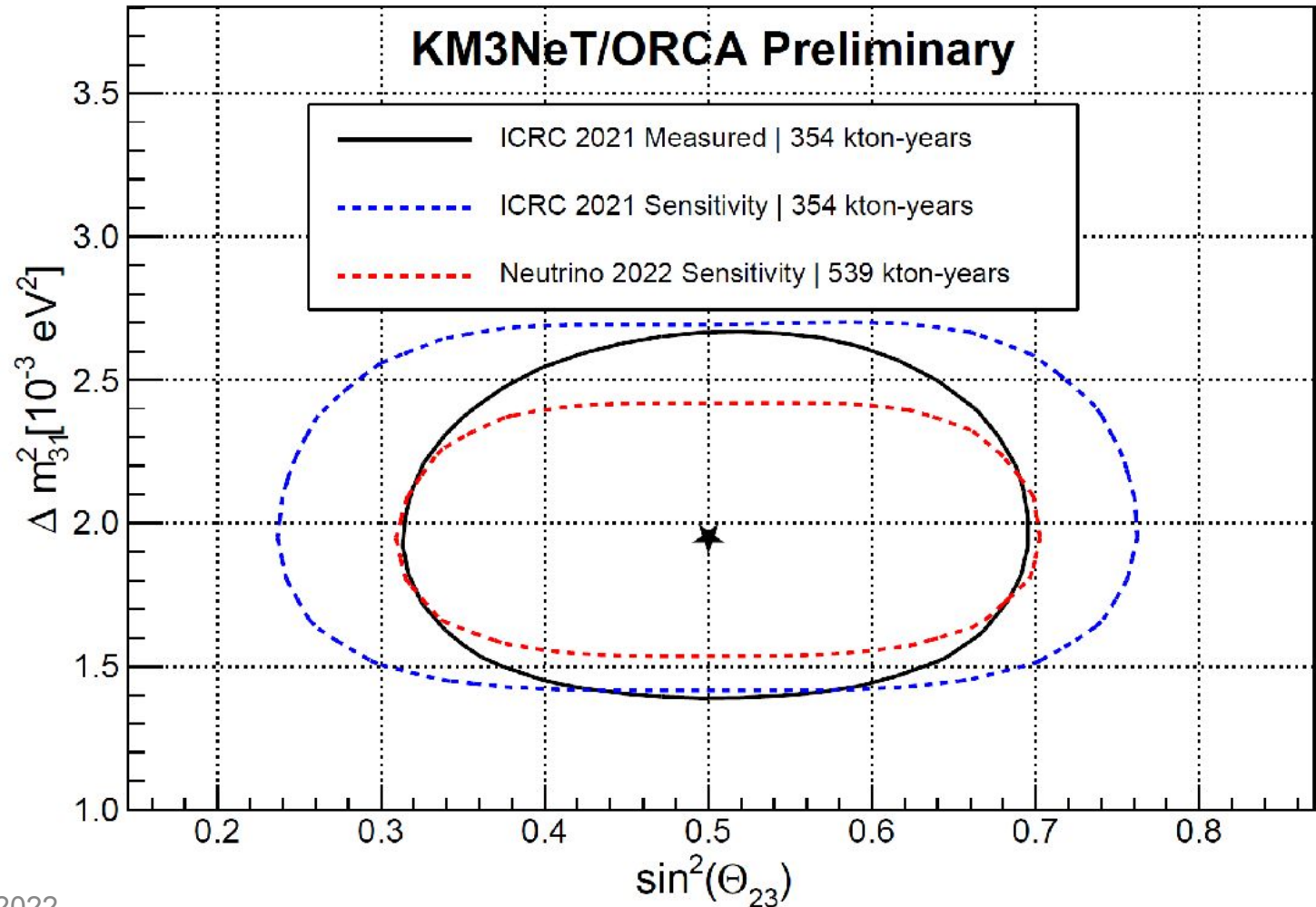
- First oscillation results already interesting with only 5% of the detector in 1 year





# Updated Results Soon!

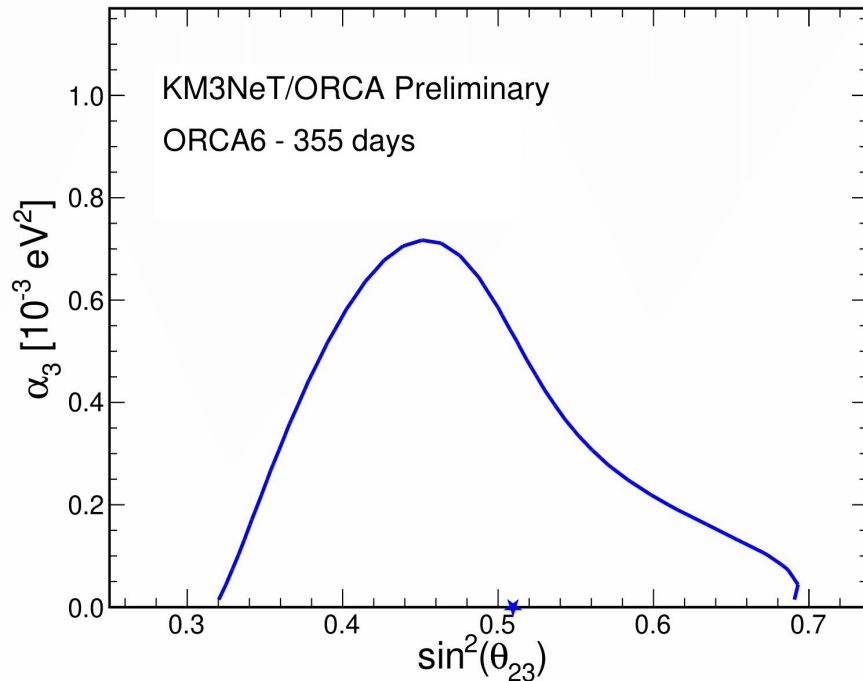
- 50% more data
- 4x more statistical power from analysis improvements



# BSM Searches

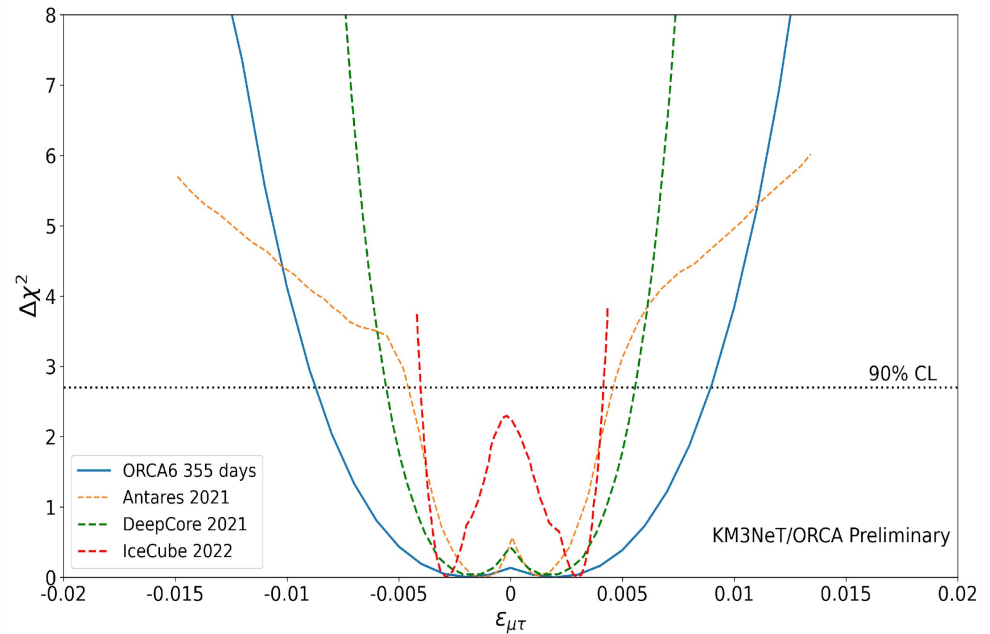
- First data already constraining some new physics scenarios
- Neutrino lifetime  $\tau_3/m_3 \equiv 1/\alpha_3 > 2.4$  ps/eV (comparable to LBL limits)
- NSI parameter  $|\varepsilon_{\mu\tau}| < 0.009$  (comparable to world best limits)

## Neutrino Decay



<https://zenodo.org/record/6758959>

## Non-Standard Interactions

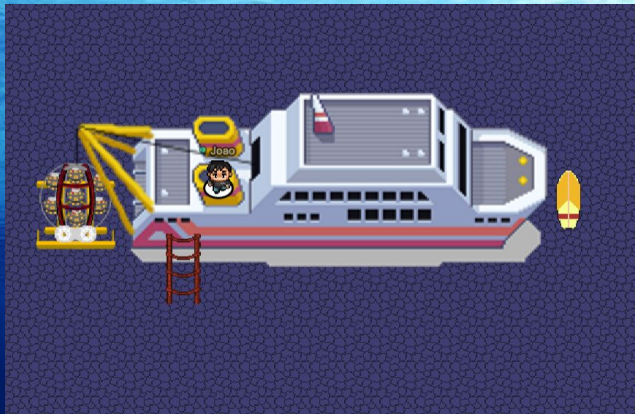


<https://zenodo.org/record/6785232>

# Summary

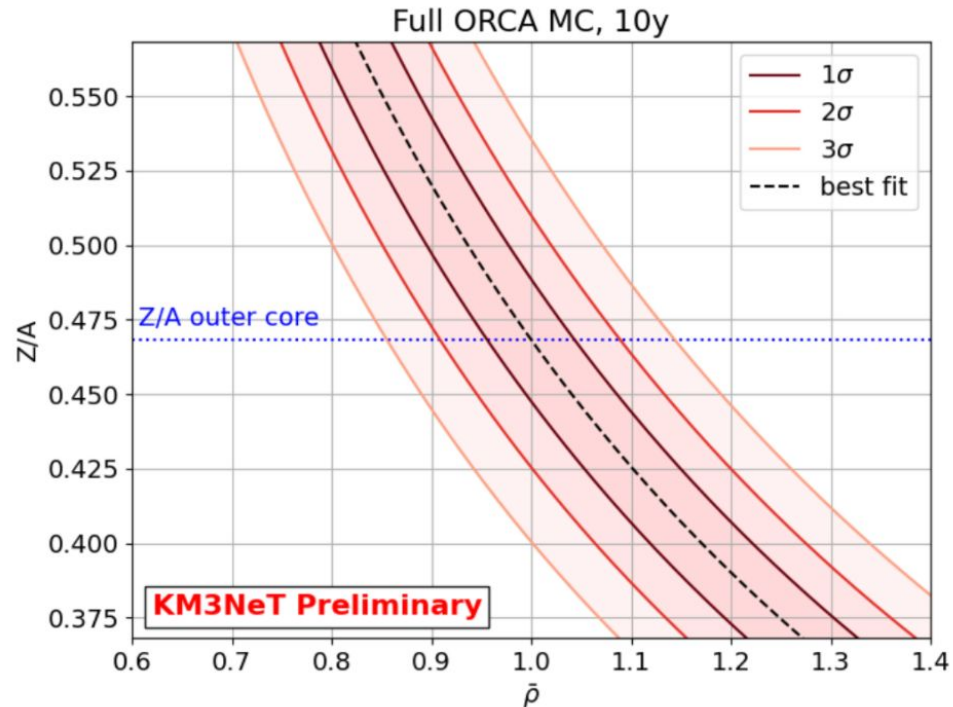
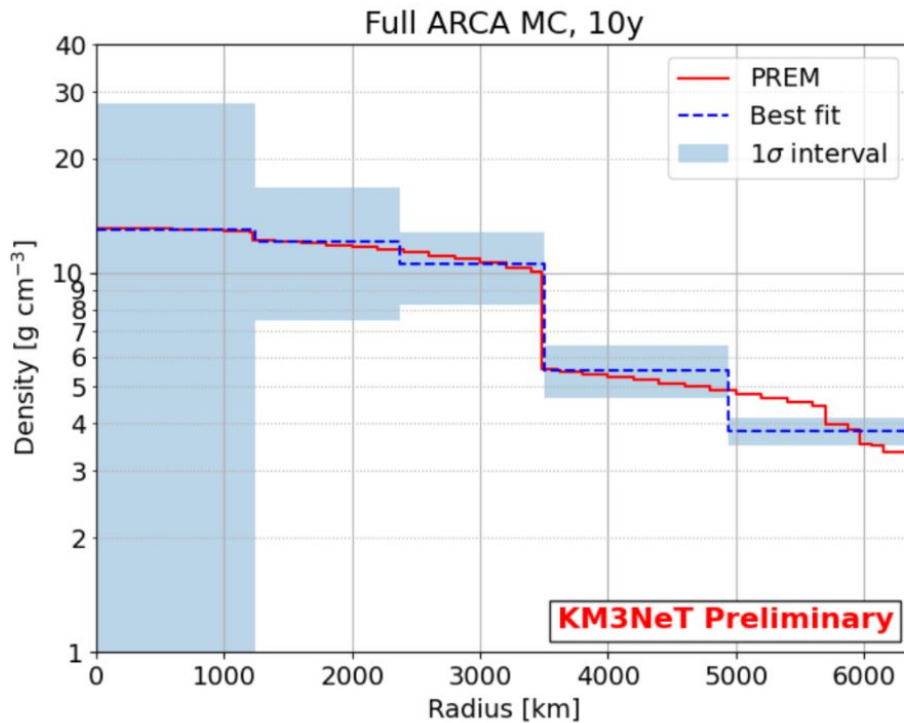
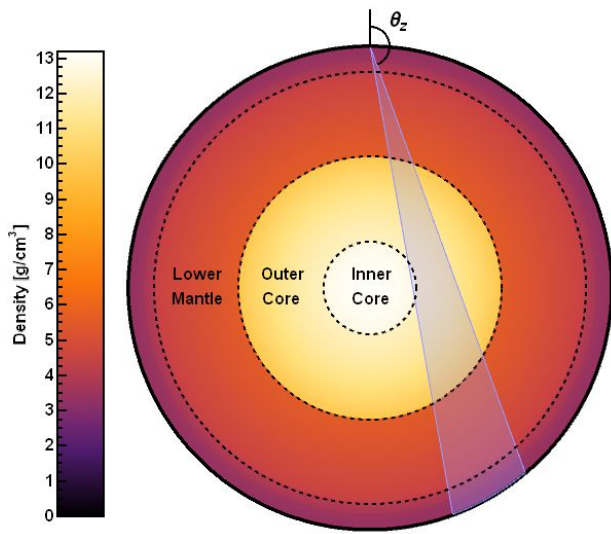
- **ORCA is here and taking data!**
- Main goal is to determine the **Neutrino Mass Ordering**
- Will also improve measurements of  $\Delta m_{32}^2$  and  $\theta_{23}$
- First results:  
$$\Delta m_{31}^2 = 1.95_{-0.21}^{+0.24} \times 10^{-3} \text{ eV}^2$$
$$\sin^2 \theta_{23} = 0.50 \pm 0.10$$
- Lots of potential for other searches: sterile neutrinos, NSI, earth tomography, tau neutrinos, etc.
- Stay tuned for more results soon

# Thank you!



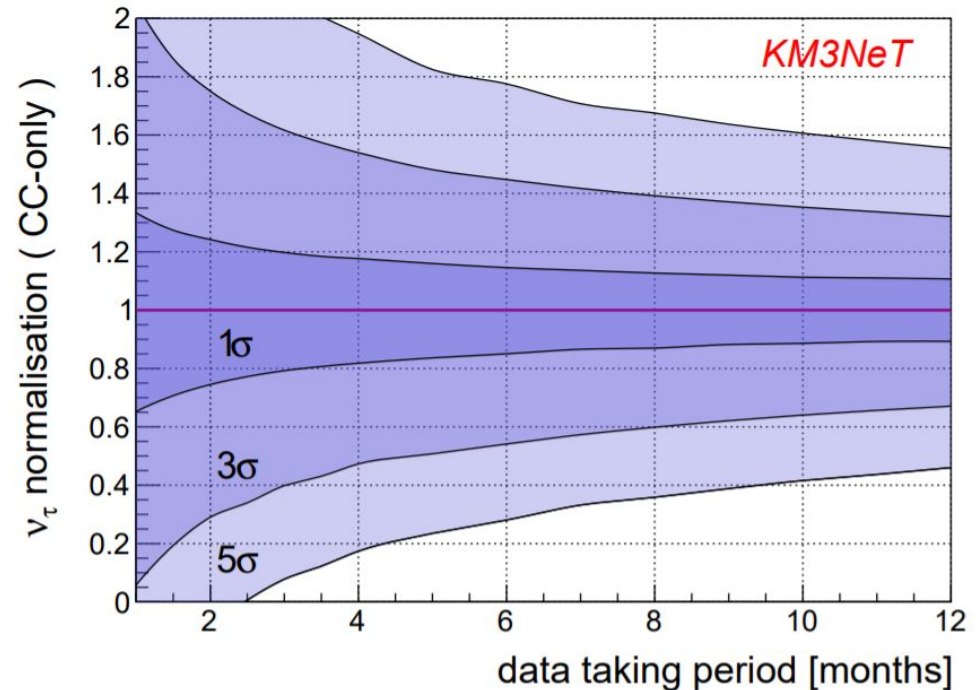
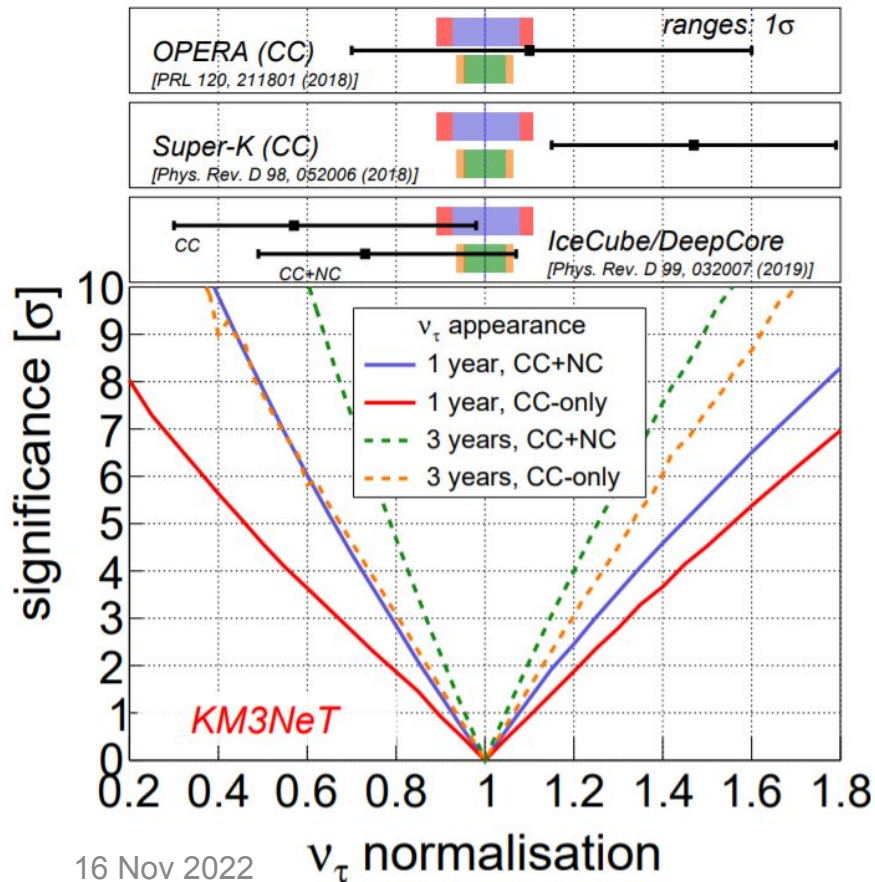
# Earth Tomography

- Use both absorption and oscillations to measure the density and composition of the deep Earth
- The sensitivity to the chemical composition is particularly interesting as there is no other known method to probe it directly



# Tau Appearance

- Atmospheric neutrinos are also an excellent probe of  $\nu_\tau$  appearance
- KM3NeT will be able to constrain the nt component to 7% level in 3 years
- Measurement can be used to probe the unitarity of the PMNS matrix
- Tau appearance can be confirmed with  $5\sigma$  confidence in 2.5 years

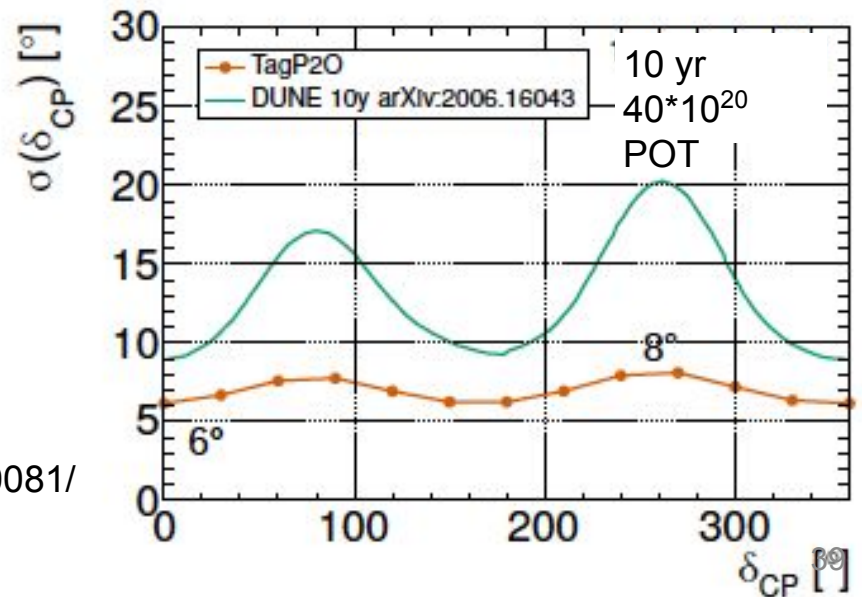
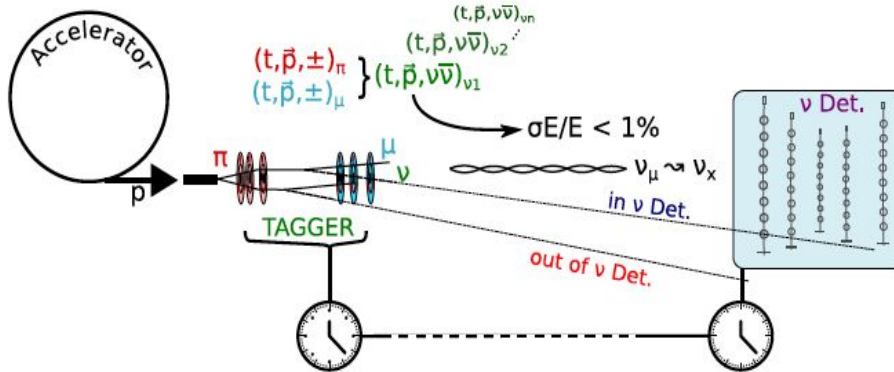
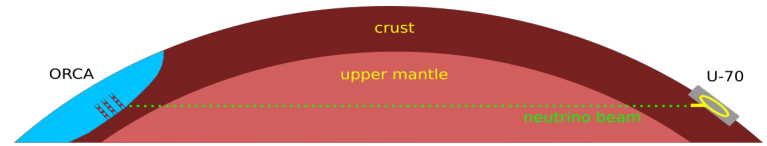




# New idea: Tagged Protvino to ORCA

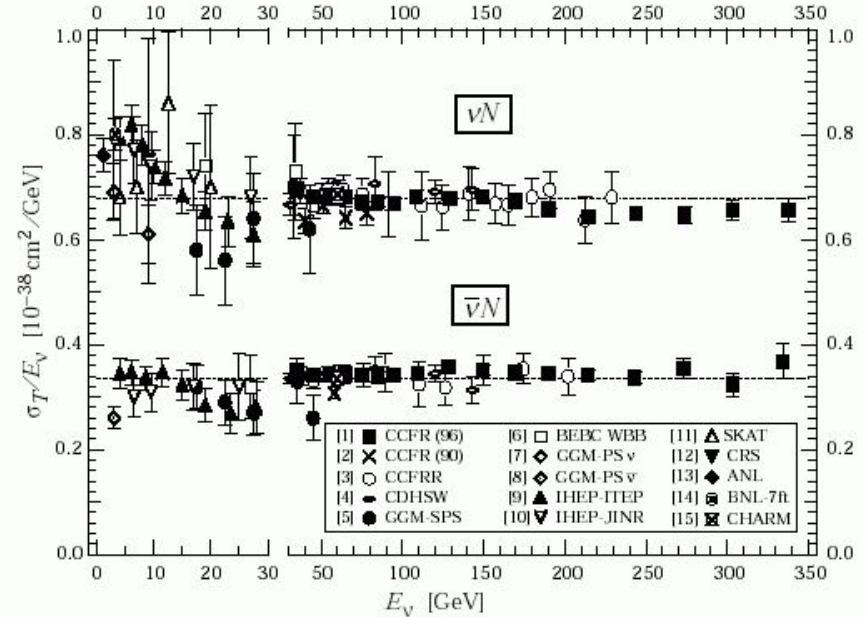
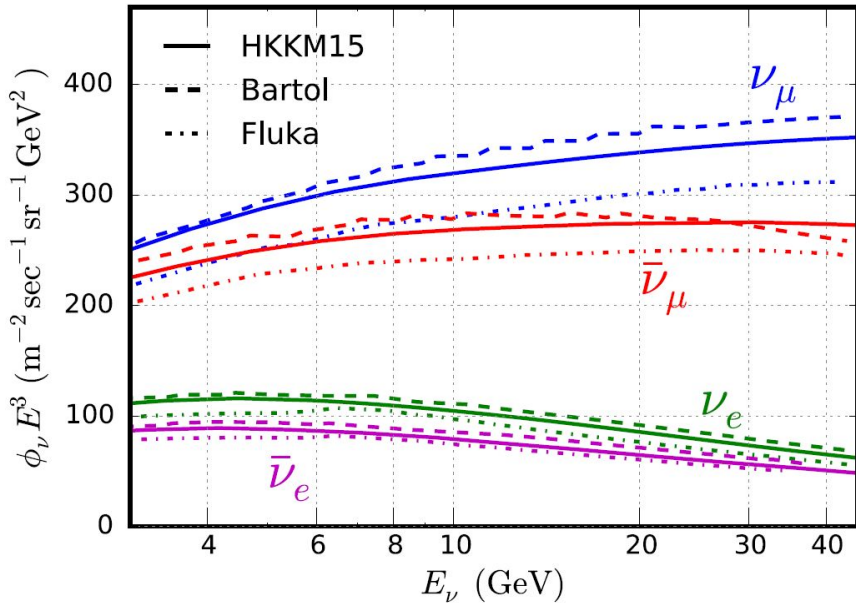
A. V. Akindinov et al.,  
"Letter of Interest for a Neutrino Beam from Protvino to KM3NeT/ORCA"  
<https://arxiv.org/abs/1902.06083>

- Neutrino Beam from Protvino to ORCA
- Baseline 2590 km
- First oscillation maximum 5.1 GeV
- Sensitivity to mass hierarchy and CPV
- Huge detector -> relax beam power
- **New idea -  $\nu$  tagging at source:**



Mathieu Perrin-Terrin@NuTel2021  
<https://agenda.infn.it/event/24250/contributions/130081/>

# Atmospheric Neutrinos



- Factor of  $\sim 2$  between  $\nu_e$  and  $\nu_\mu$
- Factor of  $\sim 2$  between  $\nu_\mu$  and  $\bar{\nu}_\mu$
- $\nu_\mu + \text{anti-}\nu_\mu = (\nu_\mu + \text{anti-}\nu_\mu + \nu_e + \text{anti-}\nu_e) \rightarrow (\nu_\mu + \text{anti-}\nu_\mu)$



# Resonance Formulas

$$\sin^2 2\theta_{13}^m \equiv \sin^2 2\theta_{13} \left( \frac{\Delta m_{31}^2}{\Delta^m m^2} \right)^2$$

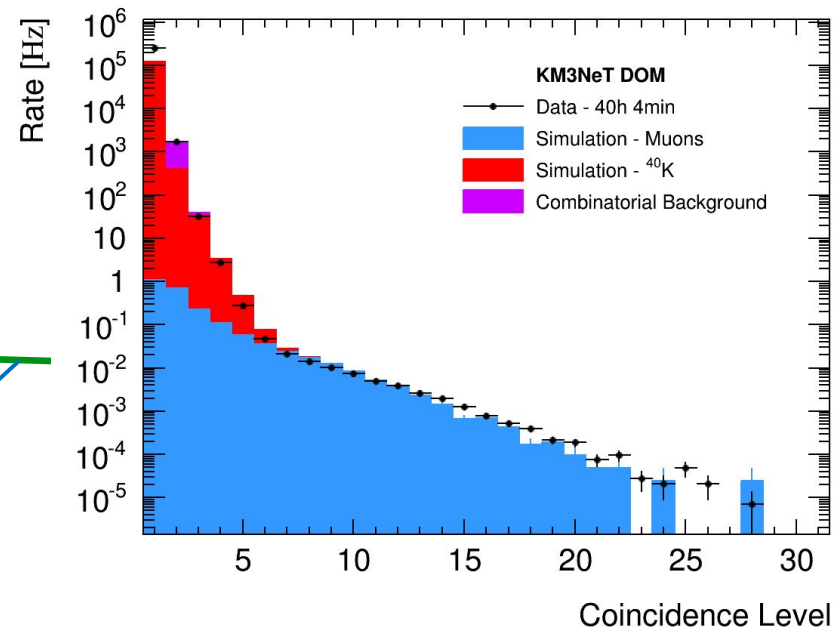
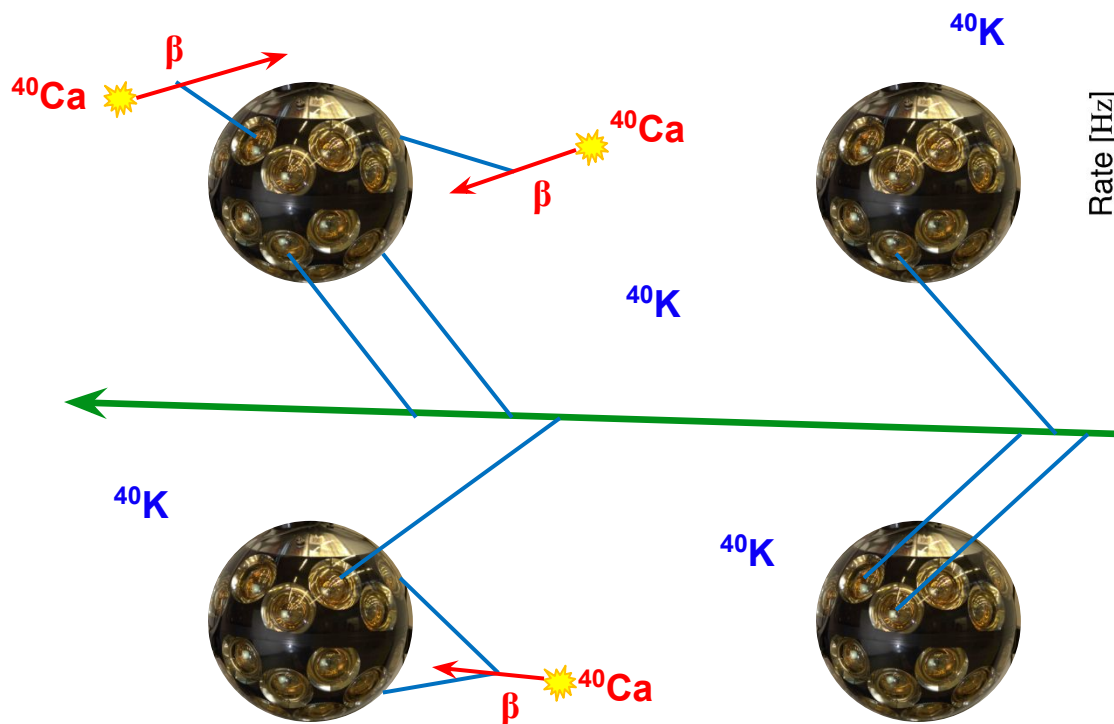
Depends on  
**sign of  $\Delta m_{31}^2$  (MH)**

$$\Delta^m m^2 \equiv \sqrt{(\Delta m_{31}^2 \cos 2\theta_{13} - 2 E_\nu A)^2 + (\Delta m_{31}^2 \sin 2\theta_{13})^2},$$

$$E_{\text{res}} \equiv \frac{\Delta m_{31}^2 \cos 2\theta_{13}}{2\sqrt{2} G_F N_e} \simeq 7 \text{ GeV} \left( \frac{4.5 \text{ g/cm}^3}{\rho} \right) \left( \frac{\Delta m_{31}^2}{2.4 \times 10^{-3} \text{ eV}^2} \right) \cos 2\theta_{13}.$$

# Trigger

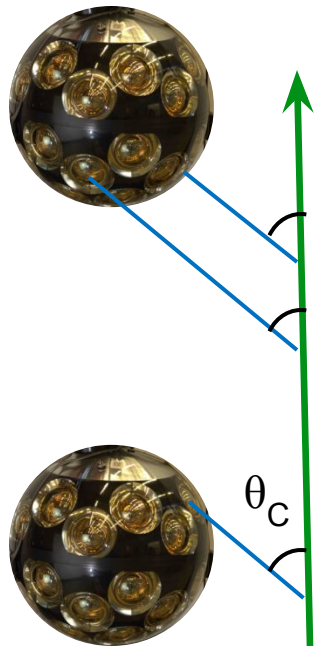
- Optical background mostly from  $^{40}\text{K}$  decays in the water
- Measured: **8 kHz** uncorr., **340 Hz** level-two coinc. / PMT [Eur. Phys. J. C 74, 3056 (2014)]
- Look for coincidences in time and PMT direction to reduce trigger rate.
- Causality further restricts space and time correlations for extra power.
- Final trigger rate  **$\sim 59$  Hz**, with **70%** of events containing a cosmic ray **muon**.



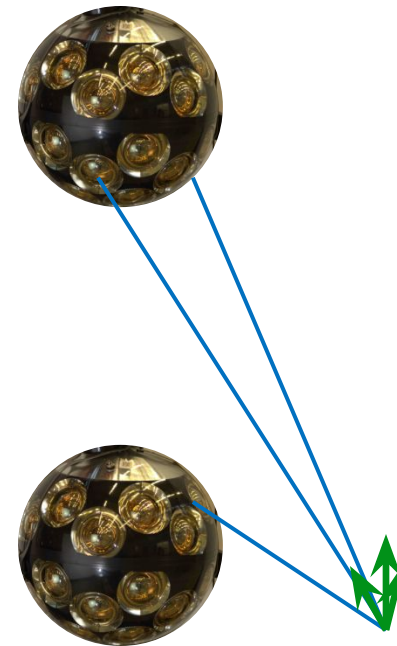
# Reconstruction

- 1) Start with a track or shower hypothesis
- 2) Use **causality** to perform a robust **hit selection**
- 3) Find **vertex** and **direction** that best match hit pattern
- 4) Estimate track range for computing **track energy** (0.24 GeV / m)
- 5) Estimate **Shower energy** and direction from hit distribution after initial fit to the vertex position and time

## Track Hypothesis

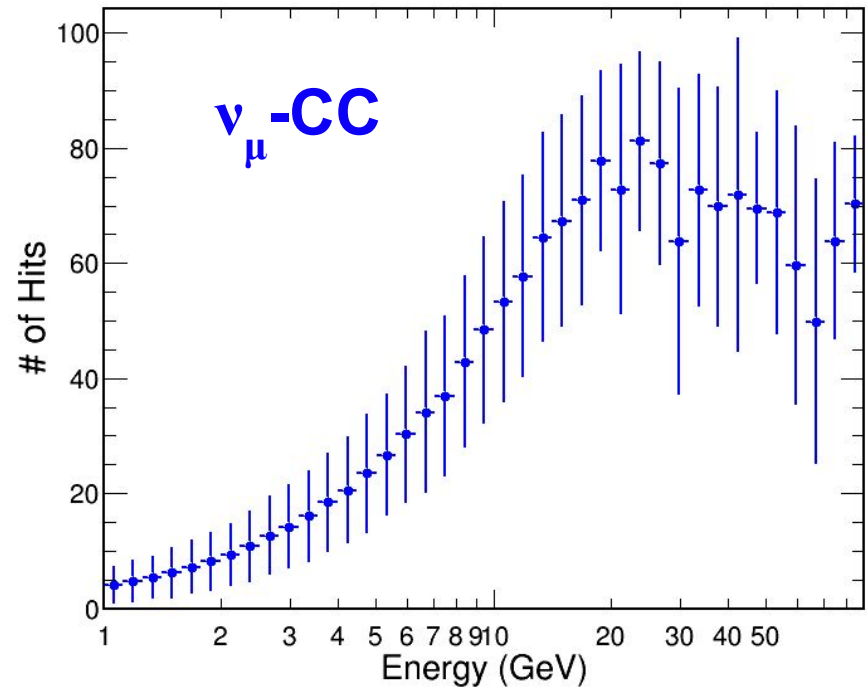
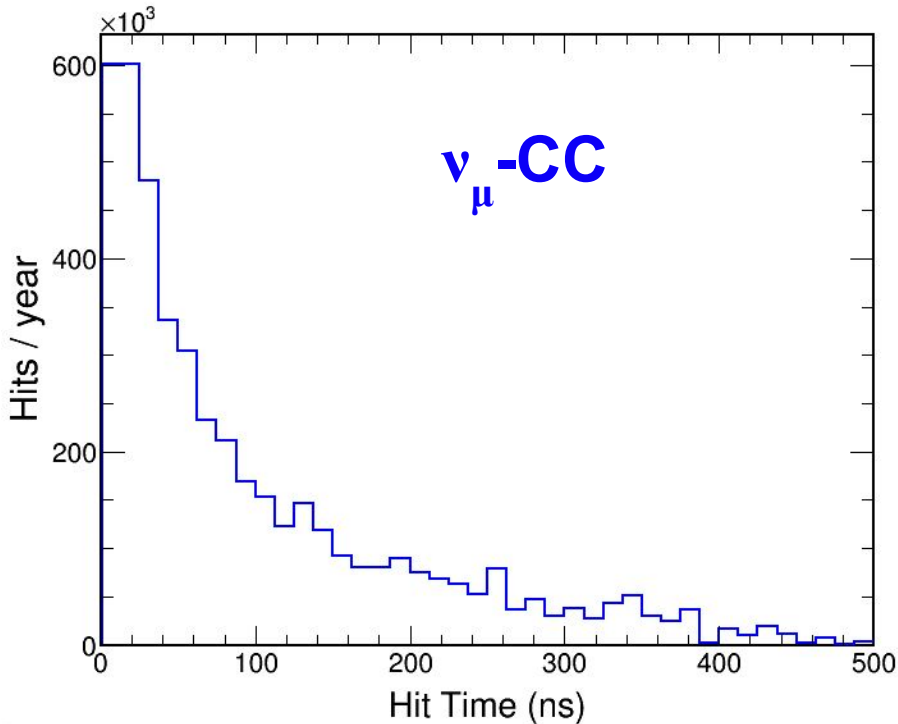
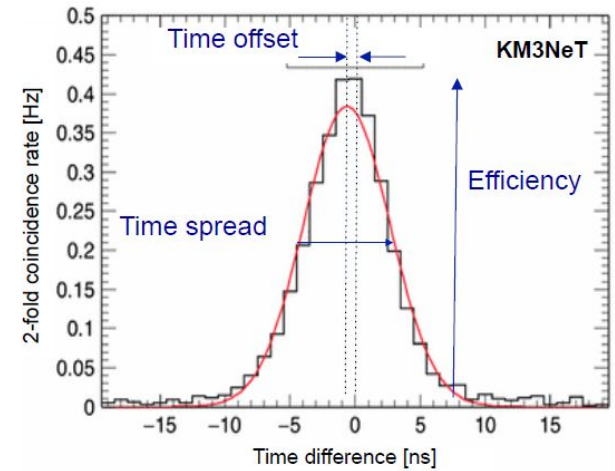


## Shower Hypothesis



# Optical Noise

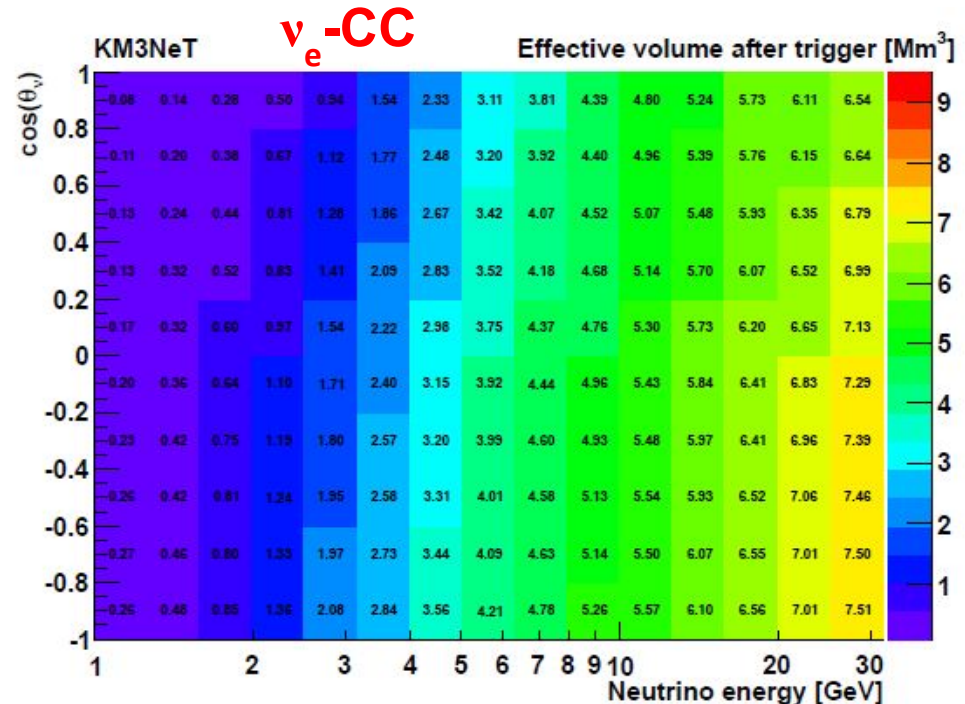
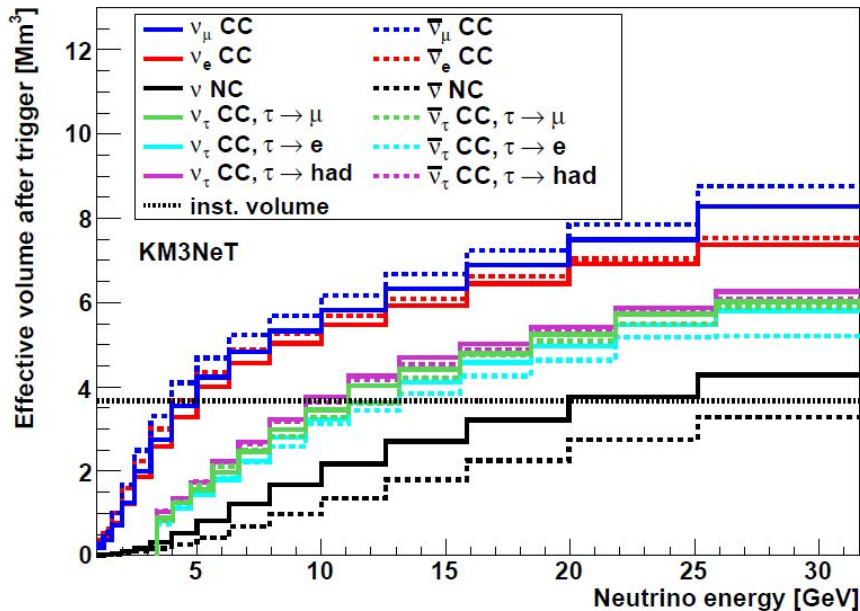
- Optical background in full detector **~500 MHz**.
- Neutrino events  $\sim 40$  hits in a **~500 ns** window
- Expect **250 noise hits** ( $\sim 15\%$  purity)
- Trigger approach **~5 ns time residuals**
- Calibrated using 2-fold coincidences
- Can achieve **~3 noise hits** per trigger ( $>90\%$  purity)



# Trigger Performance

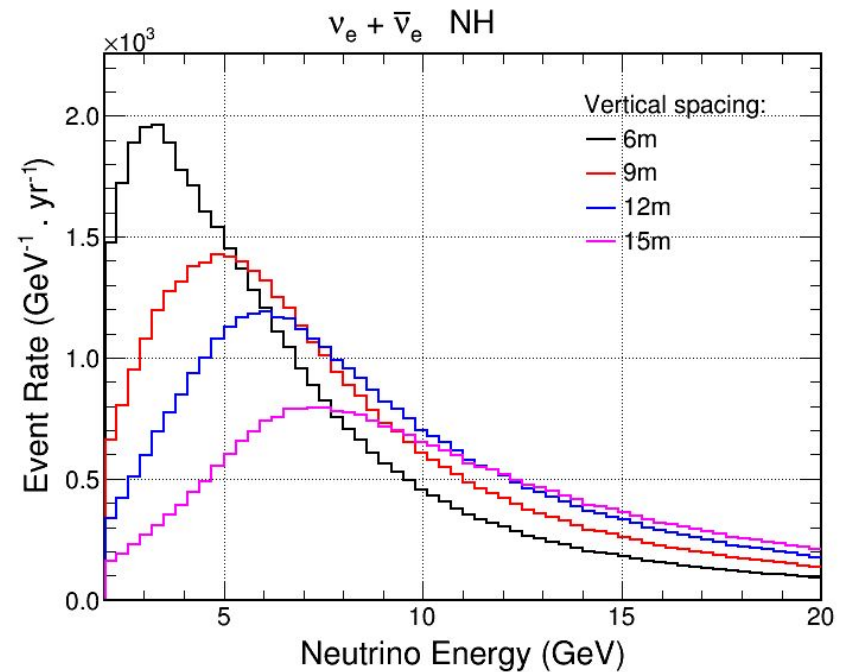
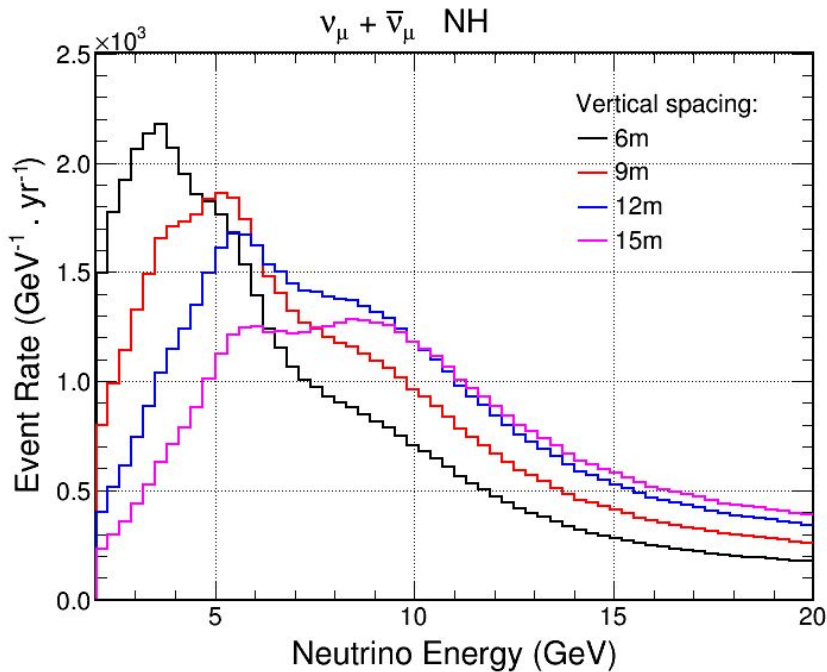
- Input a **conservative** noise rate of **10 kHz** uncorr. (**500Hz** level-two coinc.)
- Achieve a total triggered rate of **59 Hz**
- About **70%** of events contain a **muon** (41 Hz)
- High efficiency for  $\nu_\mu$  and  $\nu_e$  above 4 GeV
- Slightly more efficient for up-going neutrinos (Larger PMT coverage)

**Neutrino Rate:  
~ 1  $\nu$  / 10 min**



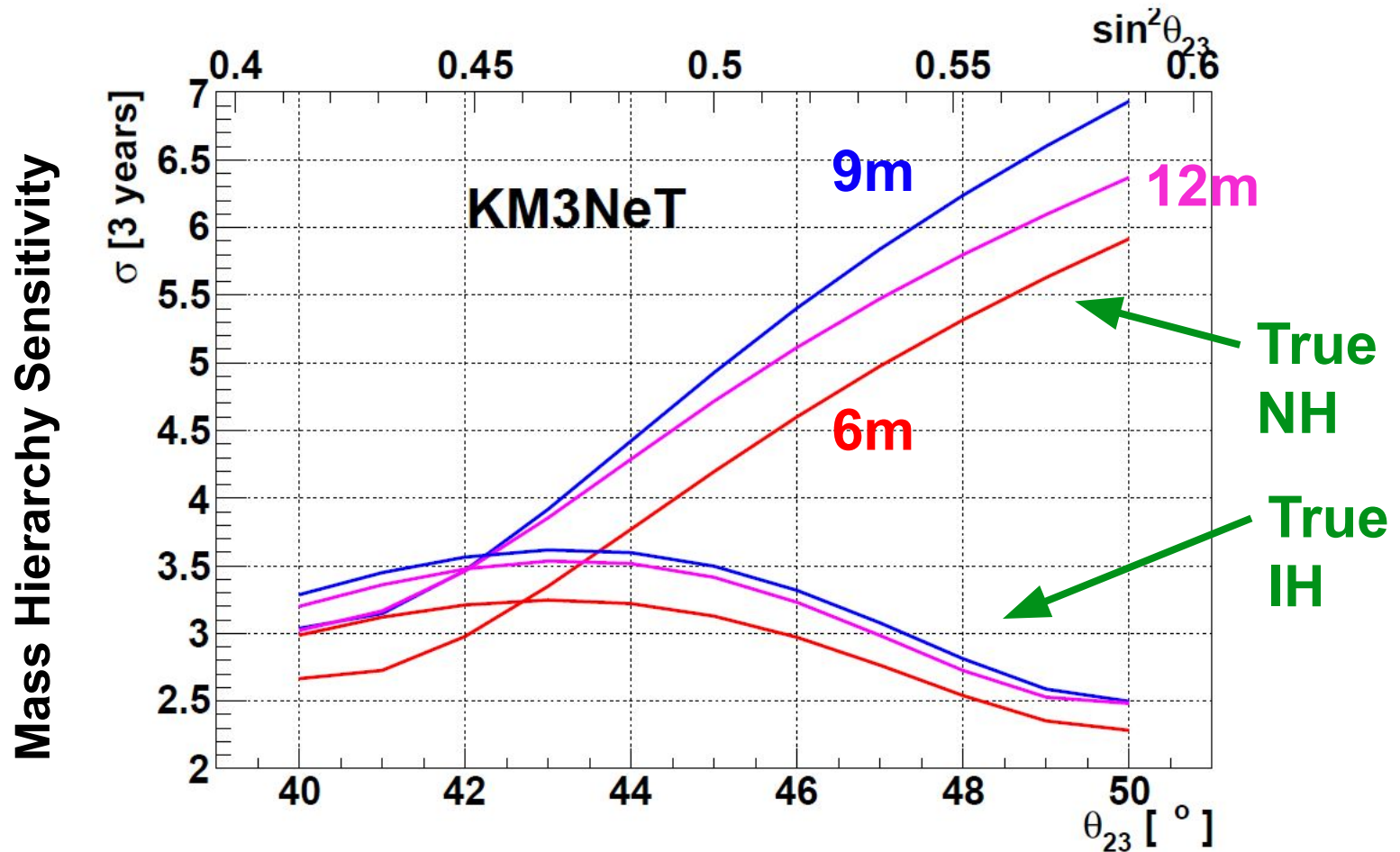
# Optimizing DOM Spacing

- Simulated small (6m) vertical spacing detector
- Mask off 1/3, 1/2 or 2/3 of DOMs to emulate larger spacing
- **Smaller spacing** enables measurement at **lower energies**
- **Larger spacing improves statistics** due to larger volume
- **Tune spacing** to obtain maximum sensitivity to **Mass Hierarchy**



# Simple $\chi^2$ Analysis

- 9m spacing achieves best sensitivity



# Beyond the Standard Model



# Extended Models

## Non-Standard Interactions (NSI)

Arbitrary  
Perturbation  
n

$$H_{eff} = U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} \end{bmatrix} U^\dagger + V_e \begin{bmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{bmatrix}$$

## Sterile Neutrinos (3+N Flavours)

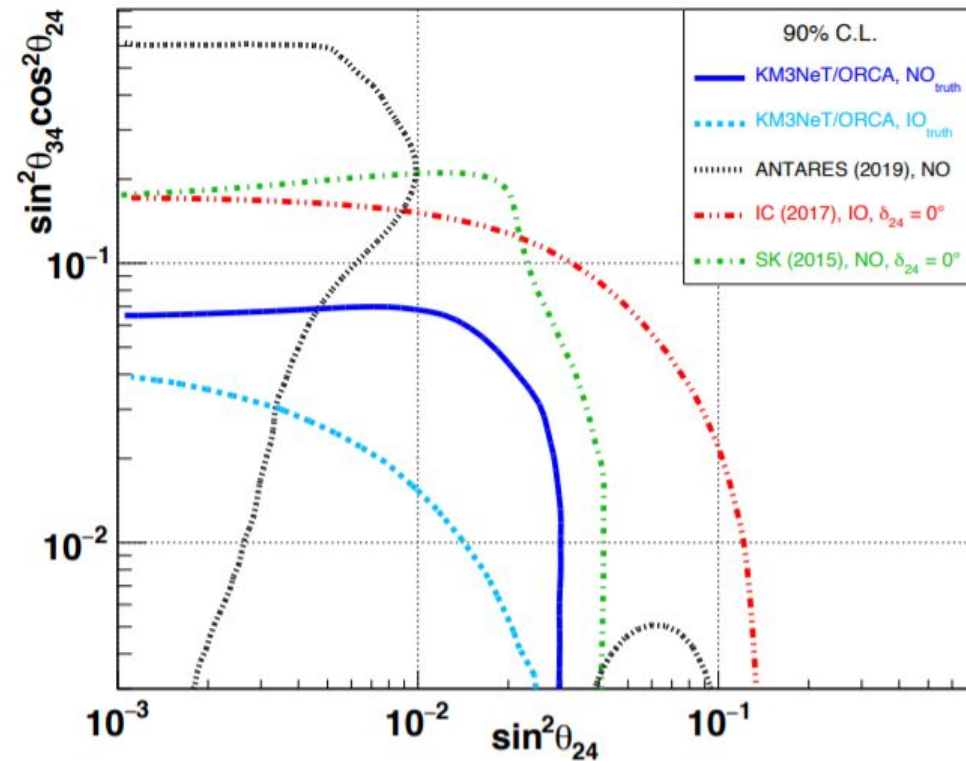
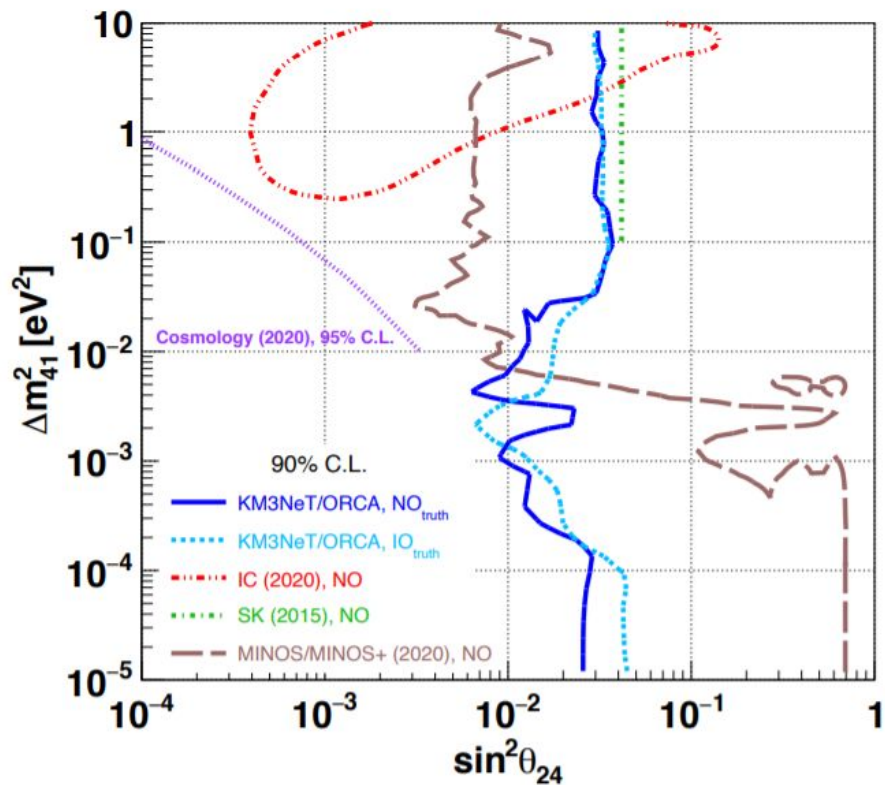
$$H_{eff} = U_S \begin{bmatrix} 0 & 0 & 0 & 0 & \cdots \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 & 0 & \cdots \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} & 0 & \cdots \\ 0 & 0 & 0 & \frac{\Delta m_{41}^2}{2E} & \cdots \\ \vdots & \vdots & \vdots & \vdots & \ddots \end{bmatrix} U_S^\dagger + \begin{bmatrix} V_e & 0 & 0 & 0 & \cdots \\ 0 & 0 & 0 & 0 & \cdots \\ 0 & 0 & 0 & 0 & \cdots \\ 0 & 0 & 0 & V_n/2 & \cdots \\ \vdots & \vdots & \vdots & \vdots & \ddots \end{bmatrix}$$

NC  
Contribution

$$U_S = U_{N-1,N} \cdots U_{34} U_{24}^{(c)} U_{14}^{(c)} U_{23} U_{13}^{(c)} U_{12}$$

# Sterile Neutrinos

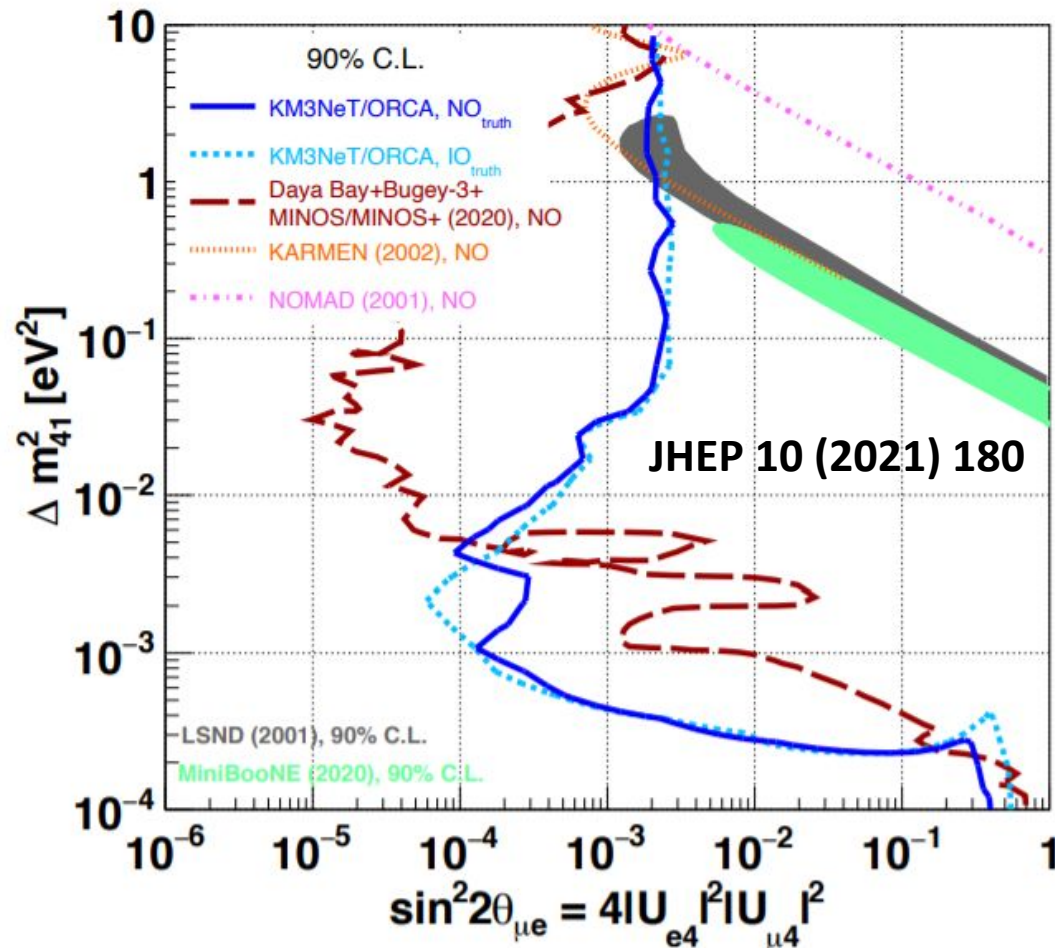
- Explores very low  $\Delta m_{41}^2$  values due to longer baselines
- World leading sensitivity to  $U_{\tau 4}$  coupling



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# Sterile Neutrinos

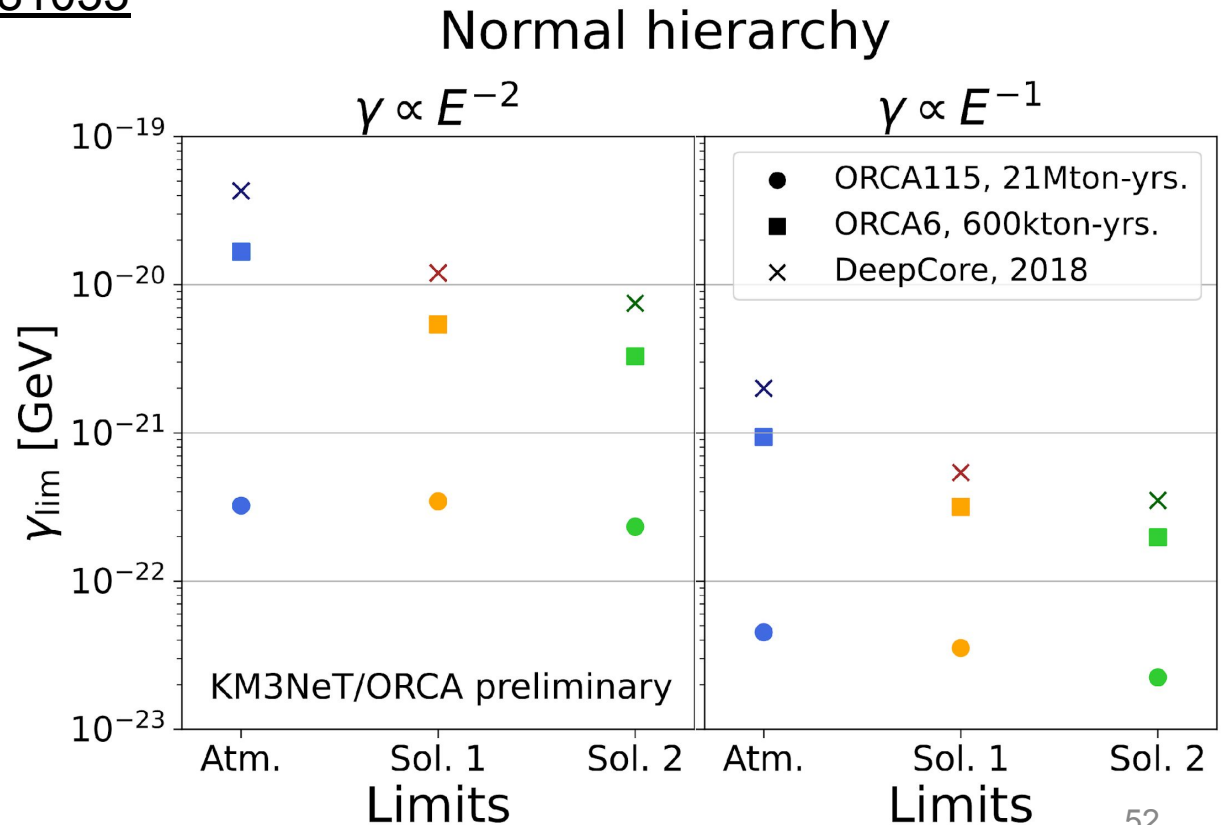
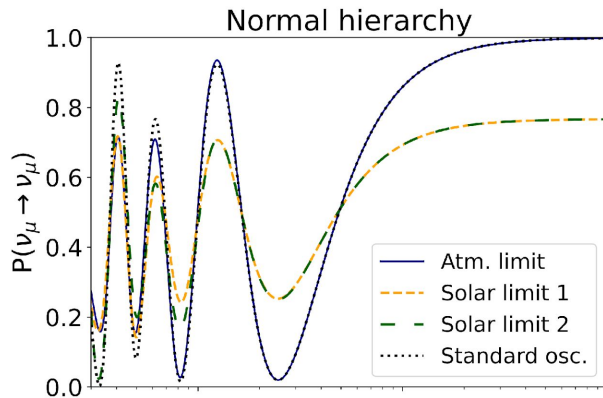
- Explores very low  $\Delta m_{41}^2$  values due to longer baselines
- World leading sensitivity to  $U_{\tau 4}$  coupling
- Probing LSND/MiniBooNE anomaly in single experiment



# Decoherence

- Measures possible loss of coherence of neutrinos due to interactions with the environment around it
- Proposed as a possible signal of Quantum Gravity
- World leading sensitivity to some decoherence modes

<https://zenodo.org/record/6781033>



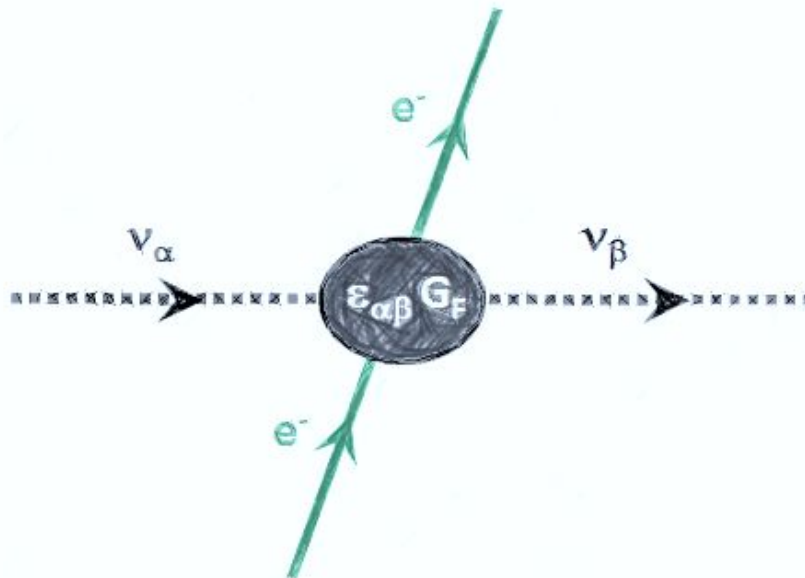
# Non-Standard Interactions

# NSI Motivation

## Non-Standard Interactions (NSI)

Arbitrary  
Perturbation  
n

$$H_{eff} = U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} \end{bmatrix} U^\dagger + V_e \begin{bmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{bmatrix}$$



Dimension-  
6

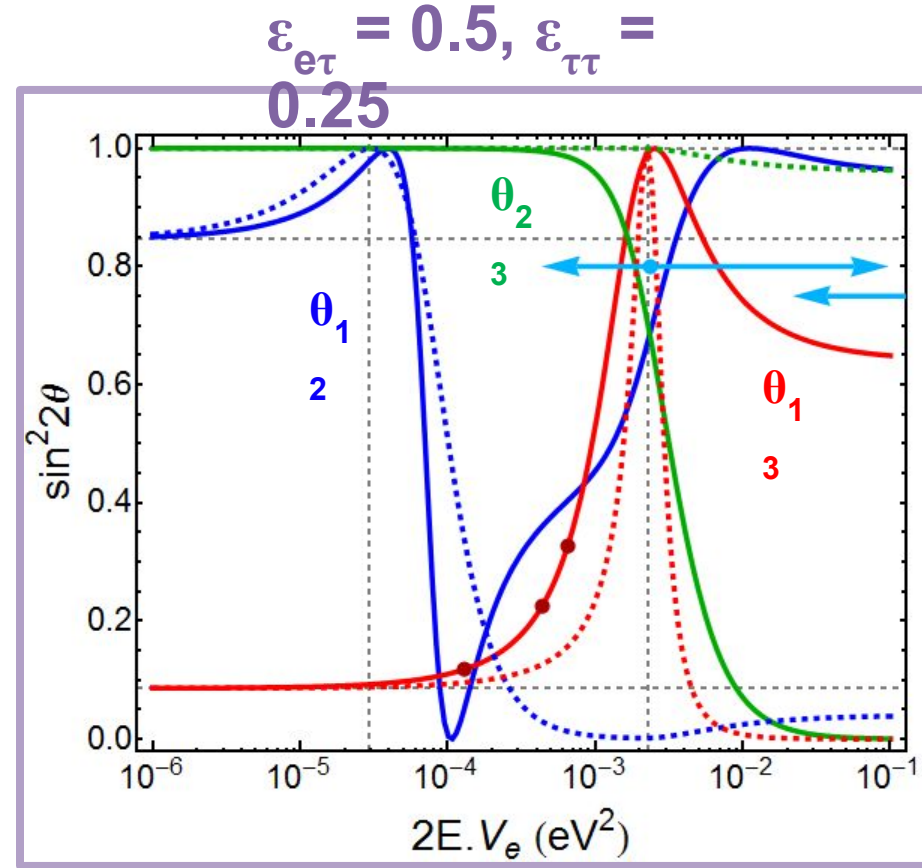
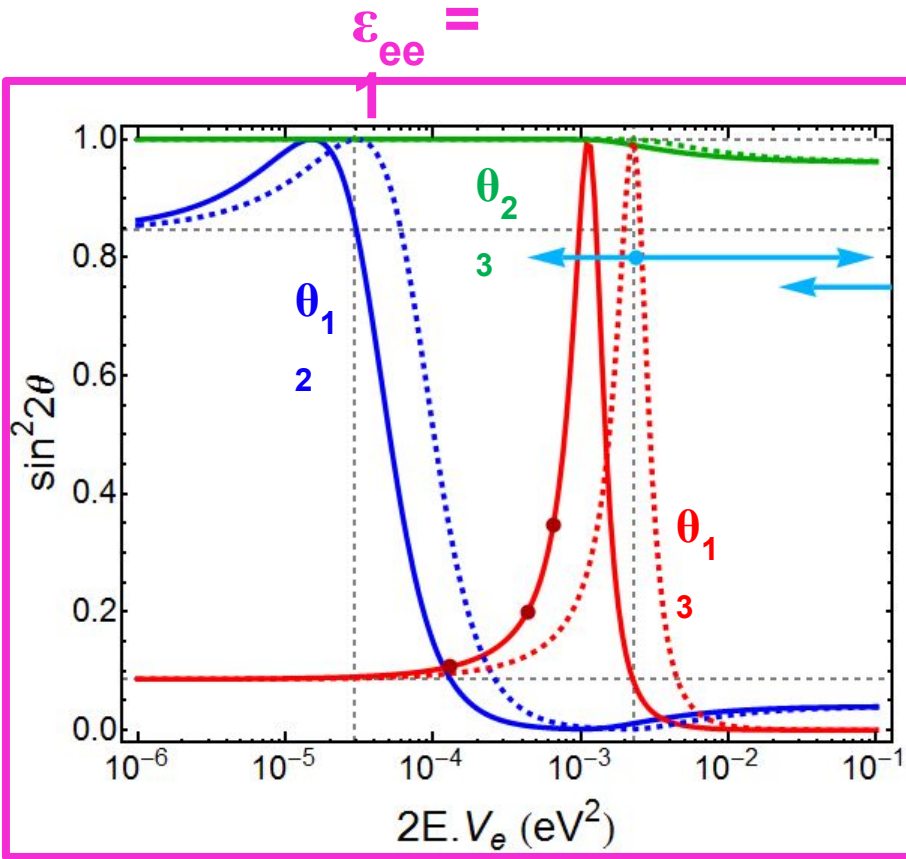
Naturalness

$$\epsilon \propto \frac{m_W^2}{m_X^2}$$

TeV  
Scale  
 $\sim 10^{-2}$

$$\mathcal{L}_{\text{NSI}} = -2\sqrt{2}G_F \epsilon_{\alpha\beta}^{ff'C} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{f} \gamma_\mu P_C f')$$

# Resonances w/ NSI



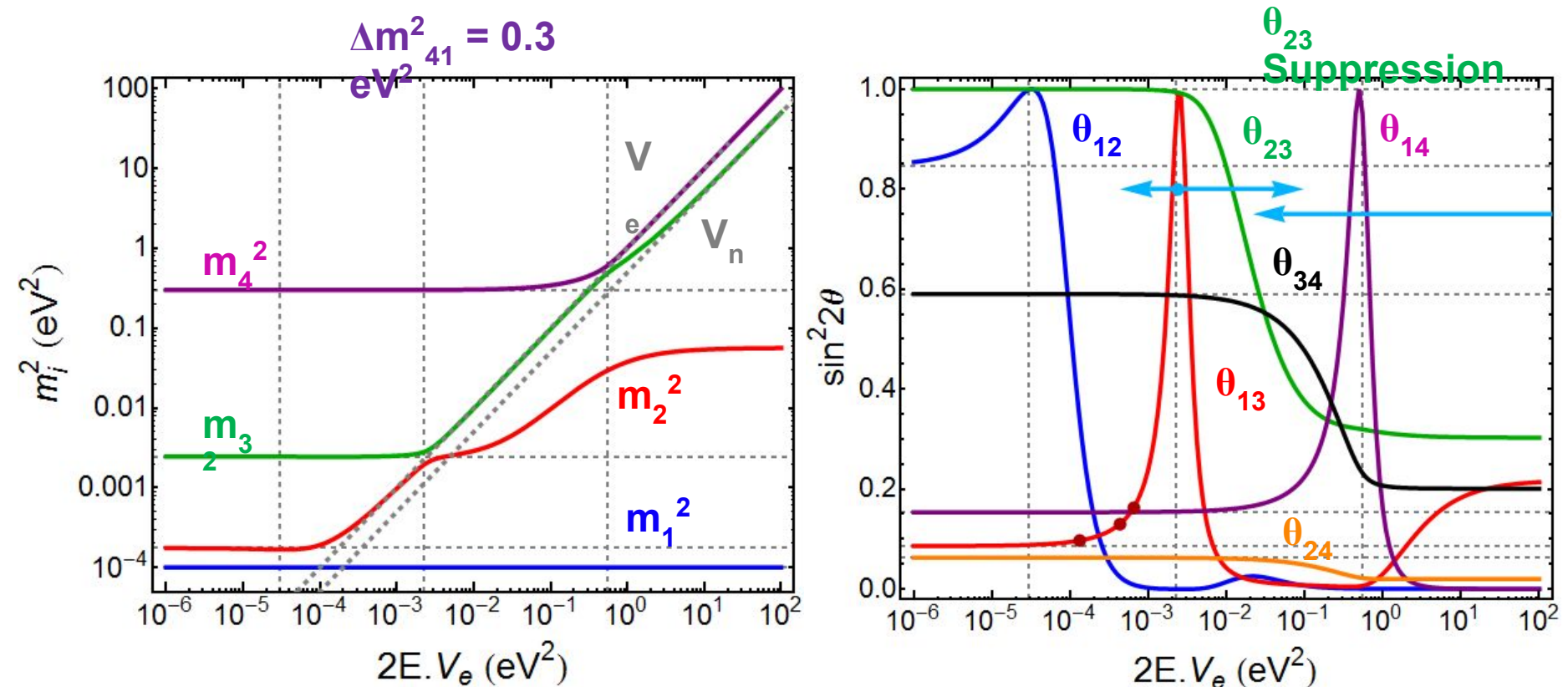
$$H_{eff} = U \begin{bmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} \end{bmatrix} U^\dagger + V_e \begin{bmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{bmatrix}$$

# Sterile Neutrinos



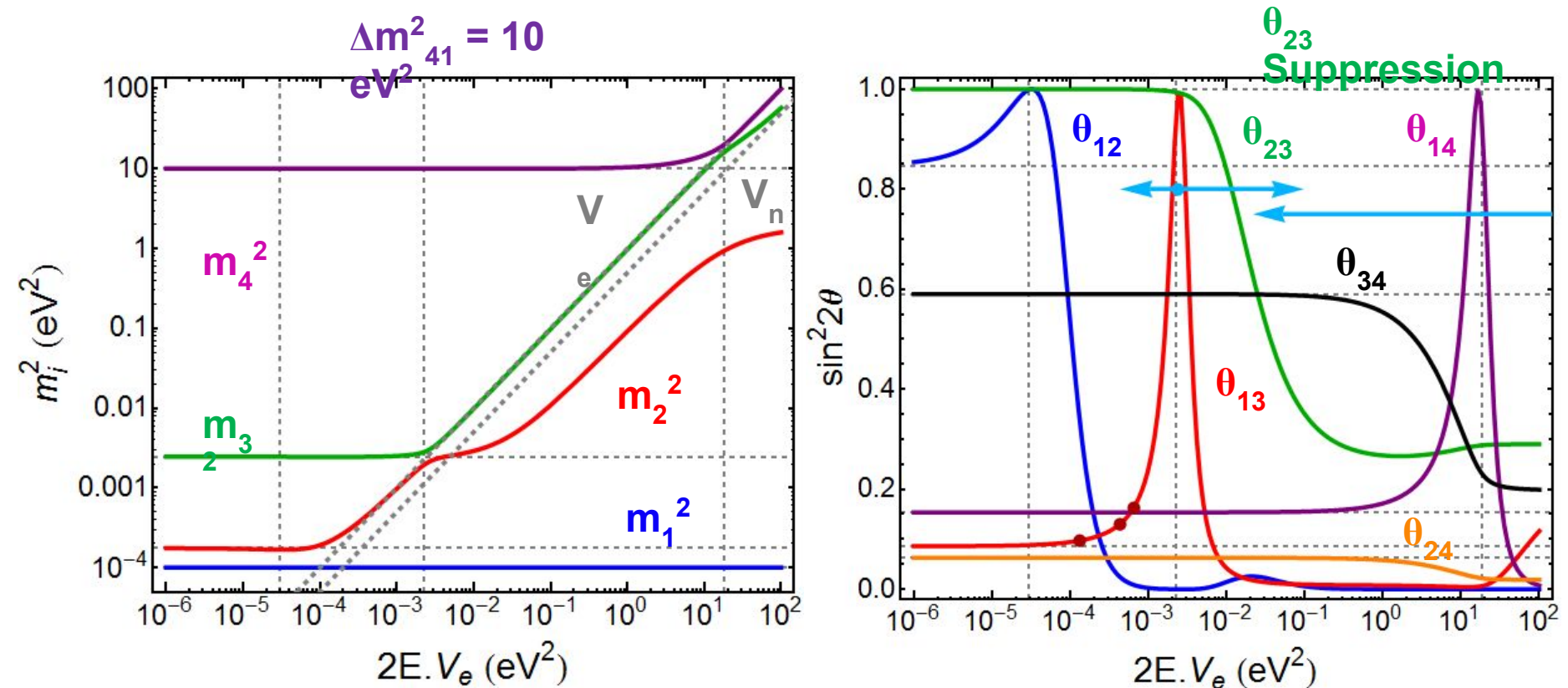
# Resonances w/ Steriles

- New resonant peak due to  $\Delta m_{41}^2$
- Some intermediate behaviour between  $\theta_{13}$  and  $\theta_{14}$  resonances
- $\theta_{23}$  suppression seems to be fairly independent of  $\Delta m_{41}^2$



# Resonances w/ Steriles

- New resonant peak due to  $\Delta m_{41}^2$
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# Resonances w/ Steriles

- New second order resonance also depends on CP phases
- Very rich structure with interplays between  $U_{\mu 4}$  and  $U_{\tau 4}$
- New paper out:  
<https://arxiv.org/abs/2107.00344>

