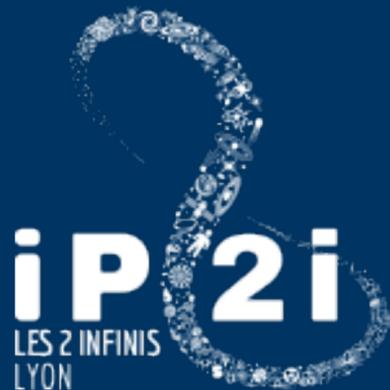


# Relic Challenges for Vector-Like Fermions as Connectors to a Dark Sector

GDR-InF Annual Workshop 2022 - Lyon

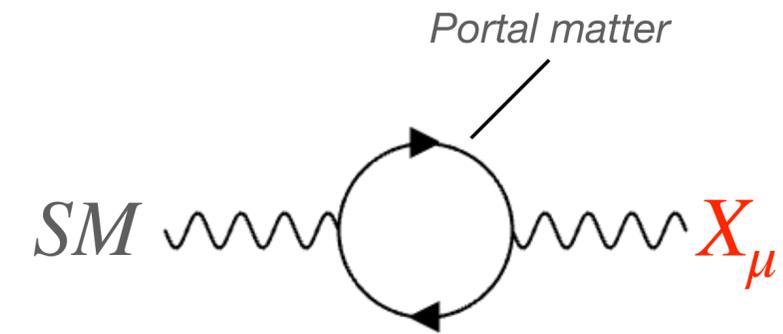
*AC, Gavin McGinnis, David E. Morrissey - arXiv:2209.14305*



Alexandre Carvunis - IP2I - 04/11/2022

# Motivations

## Dark Gauge Forces and their portals



- A dark gauge sector is well motivated by extensions of the SM (e.g. unification or DM)

$$\mathcal{G} = SU(3)_c \times SU(2)_L \times U(1)_Y \times \mathcal{G}'$$

- Can couple to the SM via connector fermions or *portal matter*
- $\mathcal{G}' = U(1)_x$ : Abelian dark forces can connect to the SM through kinetic mixing with  $U(1)_Y$  gauge boson –  $\Delta\mathcal{L} \propto B_{\mu\nu}X^{\mu\nu}$
- $\mathcal{G}' \neq U(1)_x$ : Non-abelian dark forces require dimension-8 connector operator –  $\Delta\mathcal{L} \propto (F_{\mu\nu})^2(X_{\mu\nu})^2$

# Motivation

## Challenges for portal matter

- Generically contains accidental symmetry (charge under  $\mathcal{G}'$ ) that makes portal matter stable. If they are produced in the early Universe, this leads to a relic density of exotic fermions.
- Charged (EM or QCD) relic particles are clearly problematic
- Neutral (or weakly interacting) relics receive cosmological bounds from Dark Matter density, direct and indirect DM detection experiments. *These bounds can rule out naive models.*
- **Using a minimal model, this work demonstrates these challenges and introduces two mechanisms to avoid them.**

# Minimal Model of Portal Matter to $U(1)_x$

$$\mathcal{G} = SU(3)_c \times SU(2)_L \times U(1)_Y \times U(1)_x$$

- Dark photon  $X_\mu$  with mass  $m_x = 15$  GeV, obtained from a dark Higgs or Stueckelberg mechanism. 15 GeV is a benchmark point

EWSB

Neutral fermions  $P^0$  and  $N$  mix

$$P \sim (1, 2, -1/2; q_x) \quad N \sim (1, 1, 0; q_x)$$

$$-\Delta\mathcal{L} = (\lambda\bar{P}\tilde{H}N + \text{h.c.}) + m_P\bar{P}P + m_N\bar{N}N$$

$$\tilde{H} = i\sigma_2 H^*$$

→

$$P^-, \psi_{1,2}$$

$$\begin{pmatrix} N \\ P^0 \end{pmatrix} = \begin{pmatrix} c_\alpha & s_\alpha \\ -s_\alpha & c_\alpha \end{pmatrix} \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix}$$

$$\tan(2\alpha) = \frac{2\lambda v}{m_P - m_N}$$

+ rotated interaction with Higgs + gauge bosons

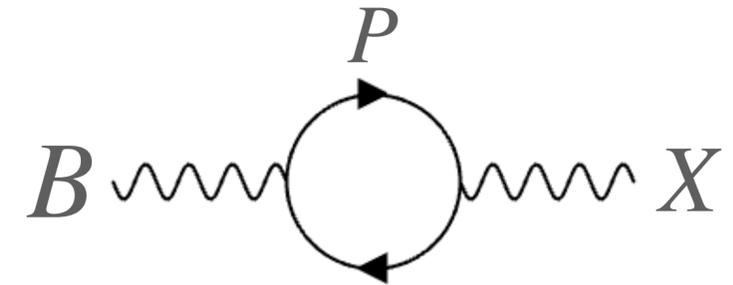
- $m_{1,2} = \frac{1}{2} \left[ m_N + m_P \mp \sqrt{(m_N - m_P)^2 + 4\lambda^2 v^2} \right] \rightarrow m_1 \leq m_P \leq m_2.$

- $\psi_1$  lightest fermion charged under  $U(1)_x$  and SM-neutral  $\rightarrow$  DM candidate

# The Model

## Portal Operators

- Kinetic mixing:  $-\mathcal{L} \supset \frac{\epsilon}{2c_W} B_{\mu\nu} X^{\mu\nu}$



- One loop contribution:

$$\Delta\epsilon \simeq -\frac{1}{3\pi} \sqrt{\alpha\alpha_x} \ln\left(\frac{\mu}{m_P}\right) \simeq - (3 \times 10^{-3}) \left(\frac{\alpha_x}{10\alpha}\right)^{1/2} \ln\left(\frac{\mu}{m_P}\right) \approx -10^{-3}$$

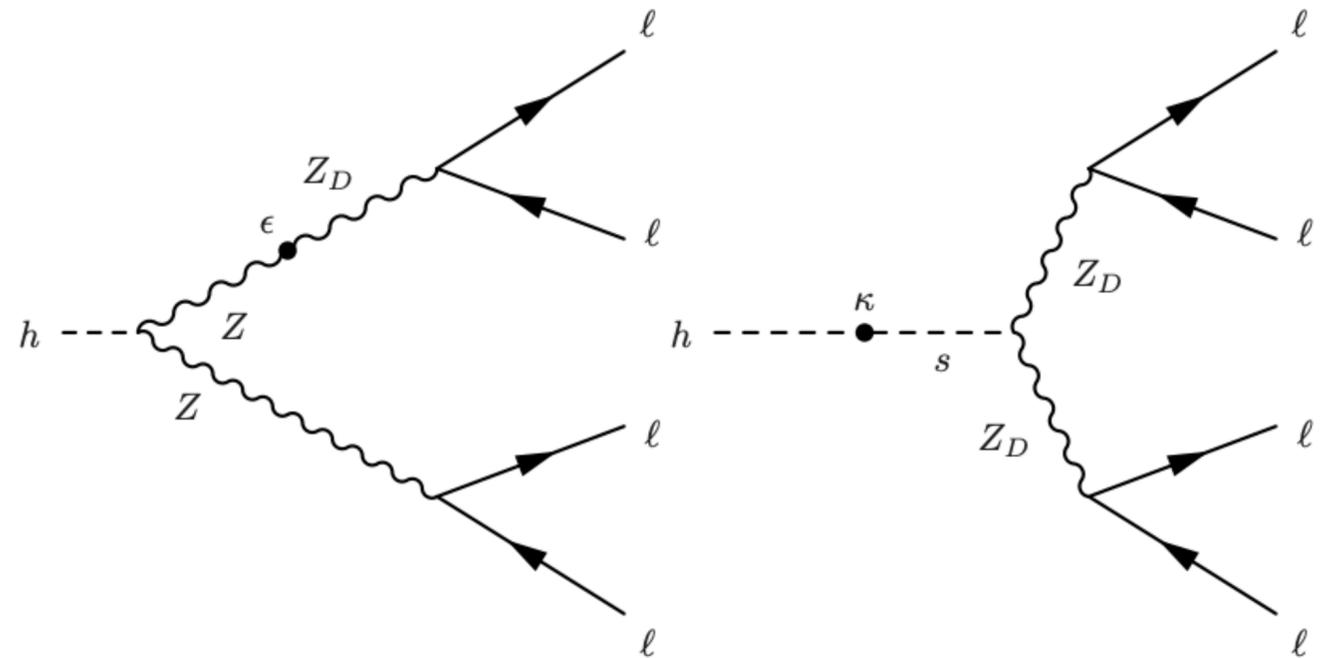
- Can be made arbitrarily small without fine tuning by introducing a mirror copy of the connector fermions with opposite  $U(1)_x$  charges
- For  $m_x = 15$  GeV, strongest direct bound comes from LHCb search for  $X^\mu \rightarrow \mu^+ \mu^-$  (1910.06926) finds  $|\epsilon| < 10^{-3}$

# Laboratory Bounds

## Higgs Decays

■ For  $m_1 < \frac{m_h}{2}$ ,  $h \rightarrow \psi_1 \bar{\psi}_1$  contributes to  $BR(h \rightarrow inv)$ . ATLAS limit (2202.07953) excludes the entire parameter region considered  $\lambda \geq 0.1$

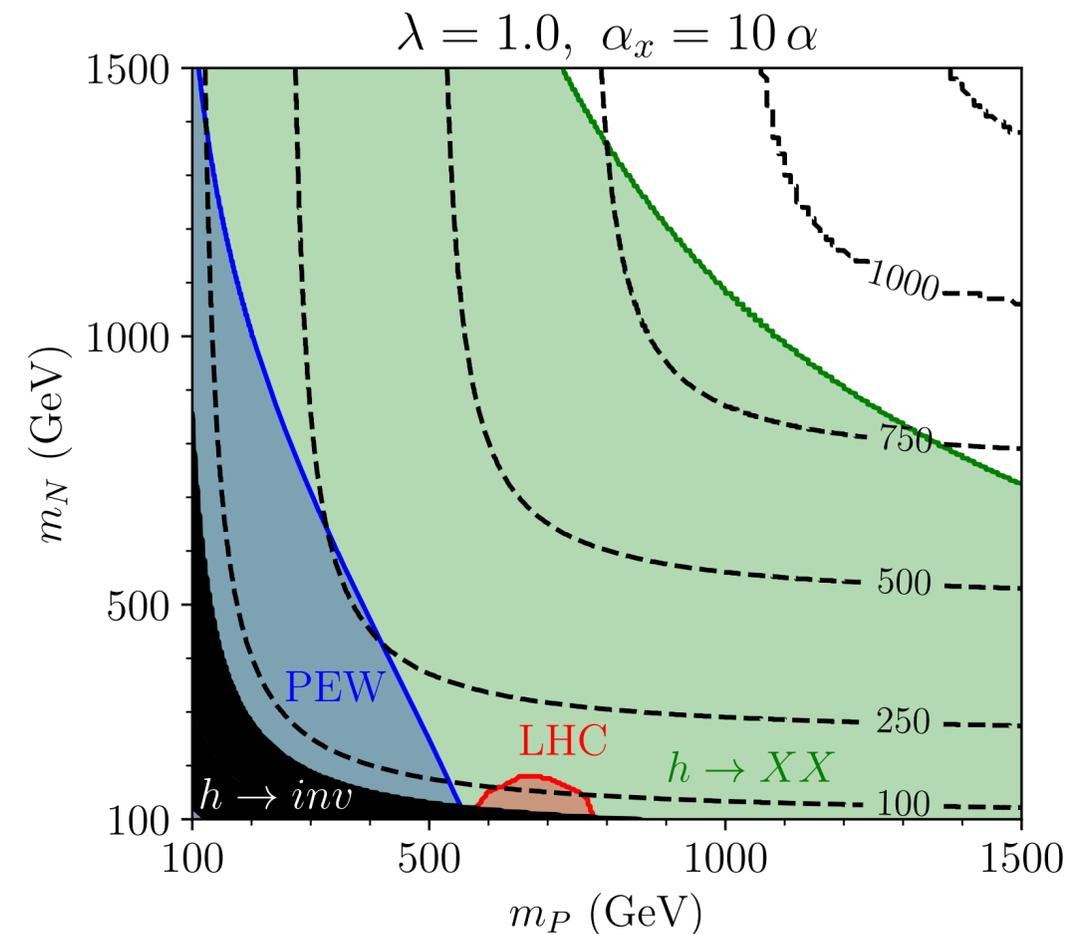
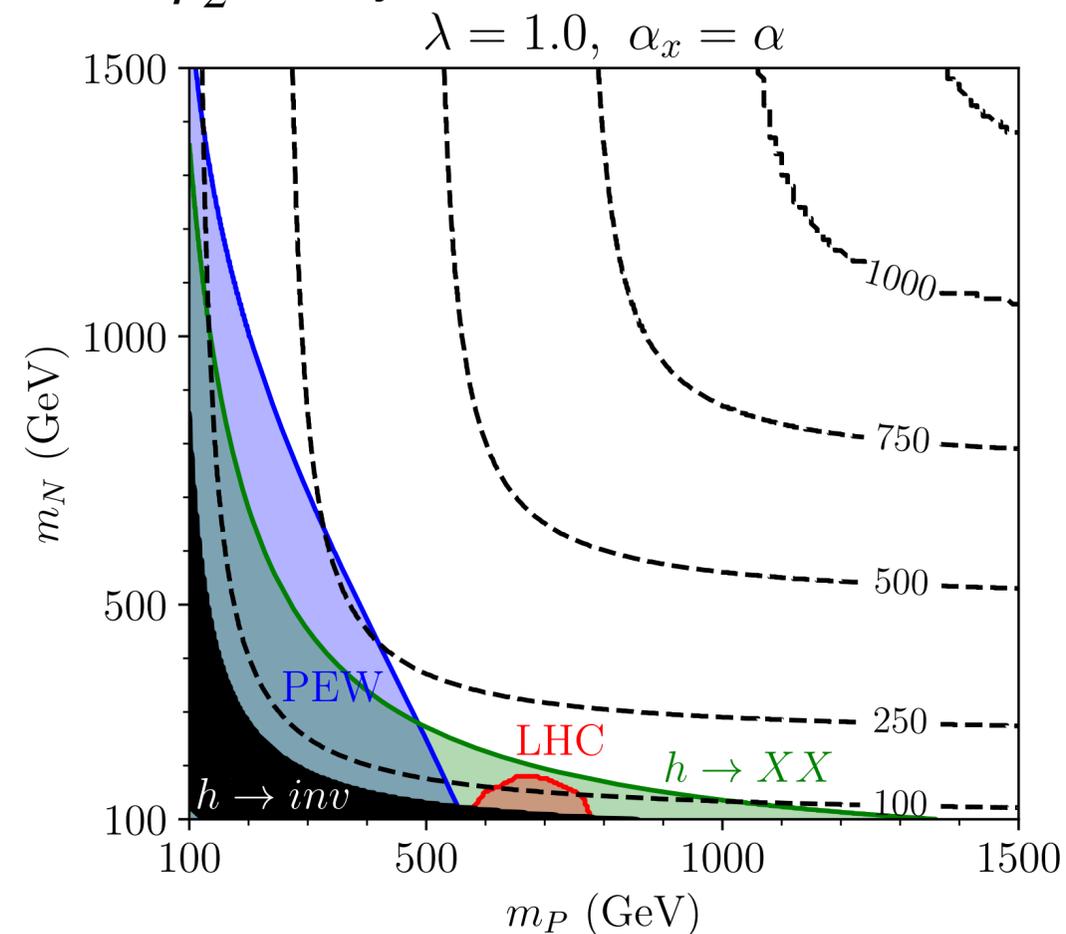
■ Contribution to  $h \rightarrow XX$  from heavy fermions the loop. ATLAS (2110.13673) and CMS (2111.01299) searches for  $h \rightarrow XX \rightarrow 4\ell$ . For  $m_x = 15$  GeV,  $BR(h \rightarrow XX) < 2.35 \times 10^{-5}$



# Laboratory Bounds

 Precision EW: mixing of  $SU(2)_L$  singlet to doublet contributes to oblique parameters S,T,U [Peskin, Takeuchi - PRL 65 (1990) 964, PRD 46 (1992) 381]

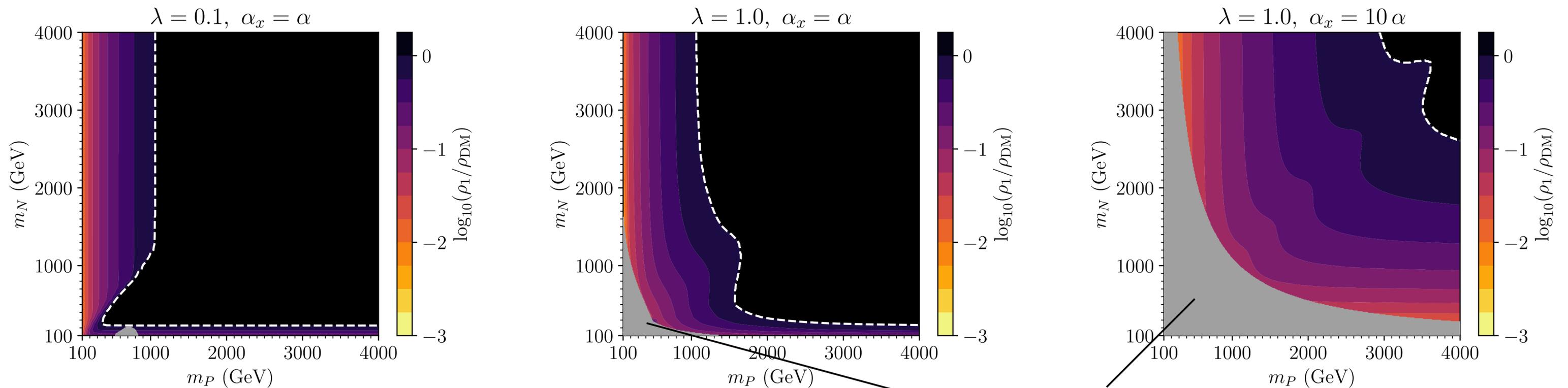
 The minimal doublet-singlet  $P - N$  model is analogous to Higgsino-Bino system. We use `Feynrules` interfaced with `MadGraph5` to calculate production cross-section. Remapping of ATLAS Higgsino-Bino search (2108.07586), including  $P^\pm$  and  $\psi_2$  decays to EW and h.



- From collider bounds only: Large viable parameter space for  $m_{N,P} \geq 100 - 700$  GeV. What about cosmological bounds ? <sup>7</sup>

# Bound from Dark Matter Relic Density

- Assuming that  $P^-$ ,  $\psi_2$ ,  $\psi_1$  thermally created in the early universe at  $T \geq m_1/20$  and then thermal freezeout.
- Annihilation via  $\psi_1\bar{\psi}_1 \rightarrow VV$  with  $V = X, Z, W$ , enhanced when  $m_P \approx m_1$  from coannihilation.
- Computation of relic density  $\rho_1$  using `Feynrules` and `MadDM`, yields upper bounds on  $m_N - m_P$ .

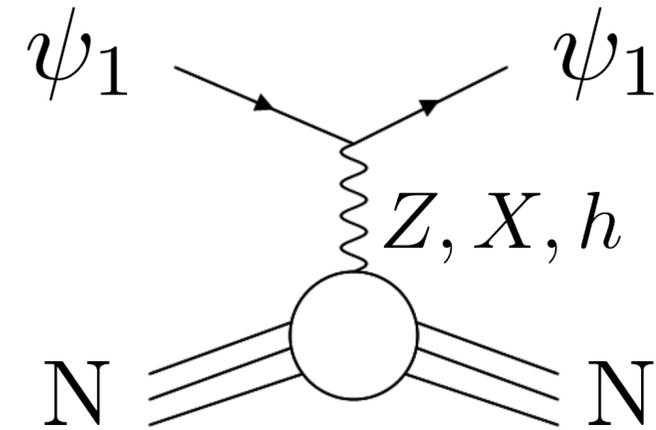
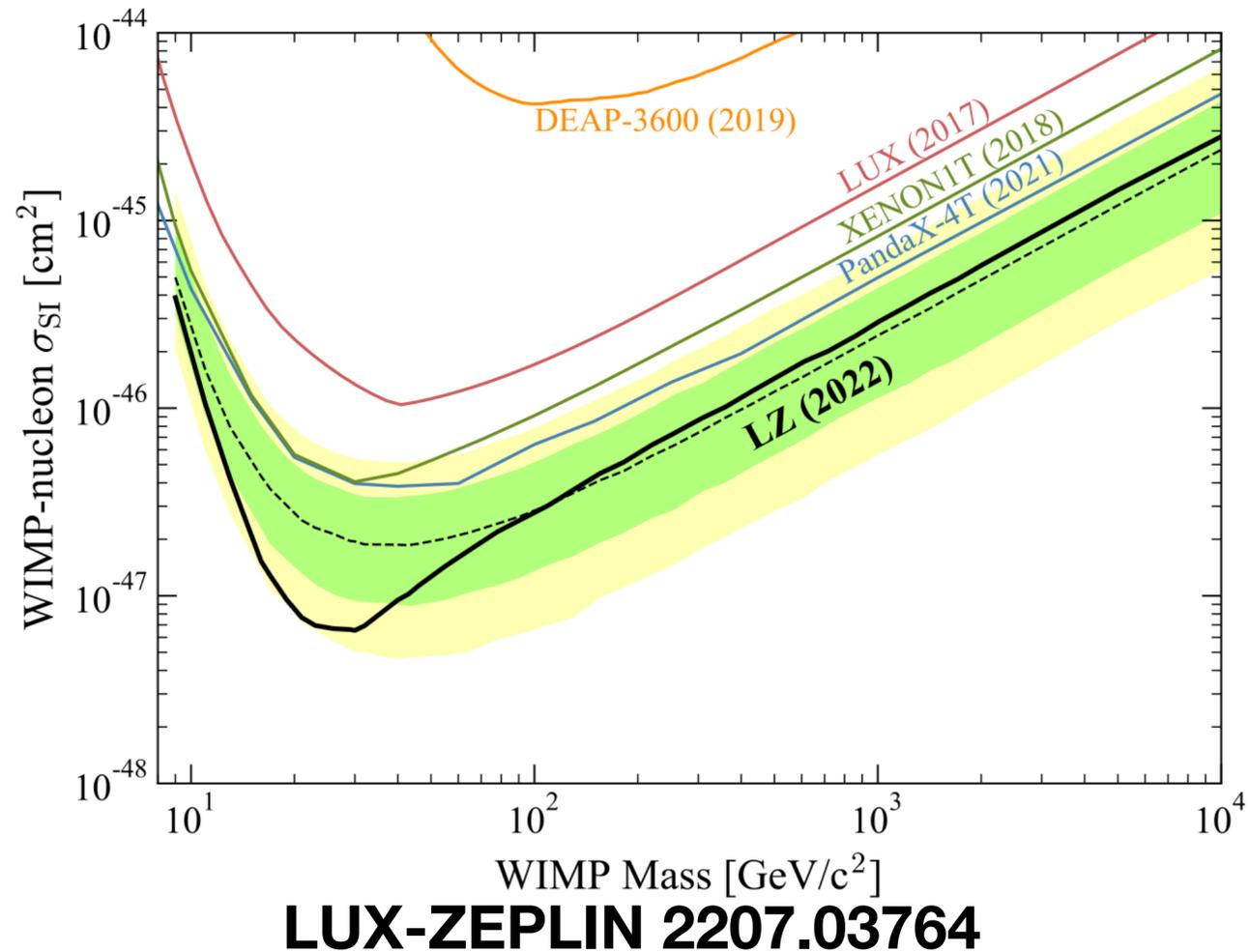


Collider bounds

# Bounds from Dark Matter Direct Detection

## Per-nucleon spin-independent scattering cross-section - $\sigma_{SI}$

- Three tree-level contributions
- Best bound for  $m_1 > 100$  GeV from LUX-ZEPLIN (2022)



$$\sigma_{SI} = \frac{\mu_n^2}{\pi} \left[ \frac{Zf_p + (A - Z)f_n}{A} \right]^2$$

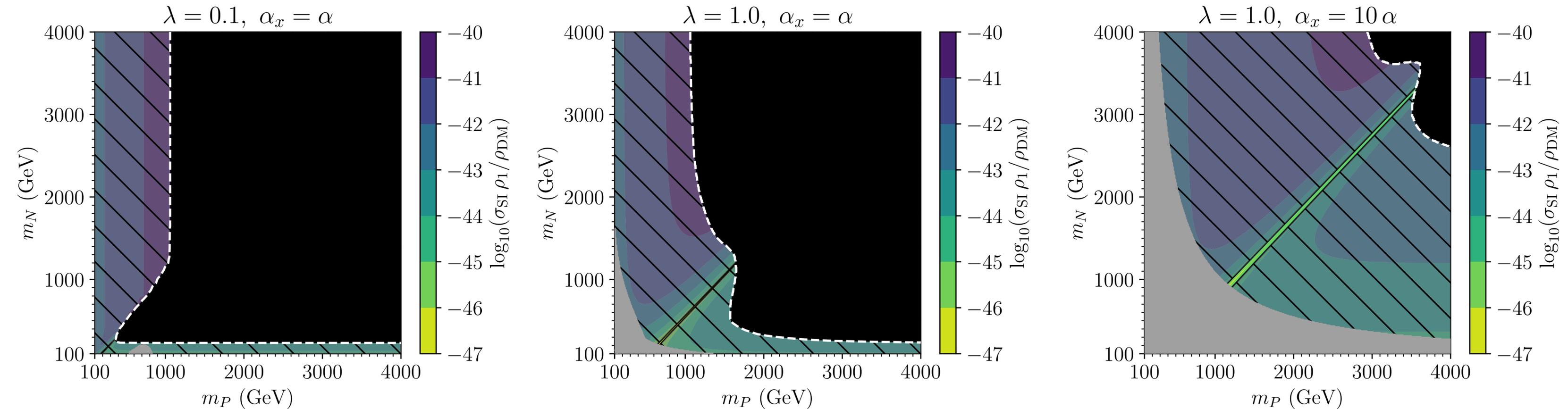
$$f_p = \overset{Z}{\frac{G_F}{\sqrt{2}} s_\alpha^2 (1 - 4s_W^2)} - \overset{X}{\frac{4\pi}{m_x^2} \epsilon \sqrt{\alpha\alpha_x}} - \overset{h}{\tilde{d}_p \left[ \frac{2}{9} + \frac{7}{9} \sum_q f_q^p \right]},$$

$$f_n = -\frac{G_F}{\sqrt{2}} s_\alpha^2 + 0 - \tilde{d}_n \left[ \frac{2}{9} + \frac{7}{9} \sum_q f_q^n \right].$$

# Bounds Direct Detection

## Rules out the minimal model

- $$\sigma_{\text{SI}} = \frac{\mu_n^2}{\pi} \left[ \frac{Zf_p + (A - Z)f_n}{A} \right]^2$$
, cancellation possible for  $\epsilon < 0$  around  $m_N \approx m_P$



- Minimal model almost entirely excluded by DD experiments. Can we avoid these bounds with minimal changes to the model?

# Fix #1

## Mass splitting from a Majorana mass term

- Dark Higgs:  $\Phi \sim (1,1,0; -2q_x)$  allows new Yukawa coupling :

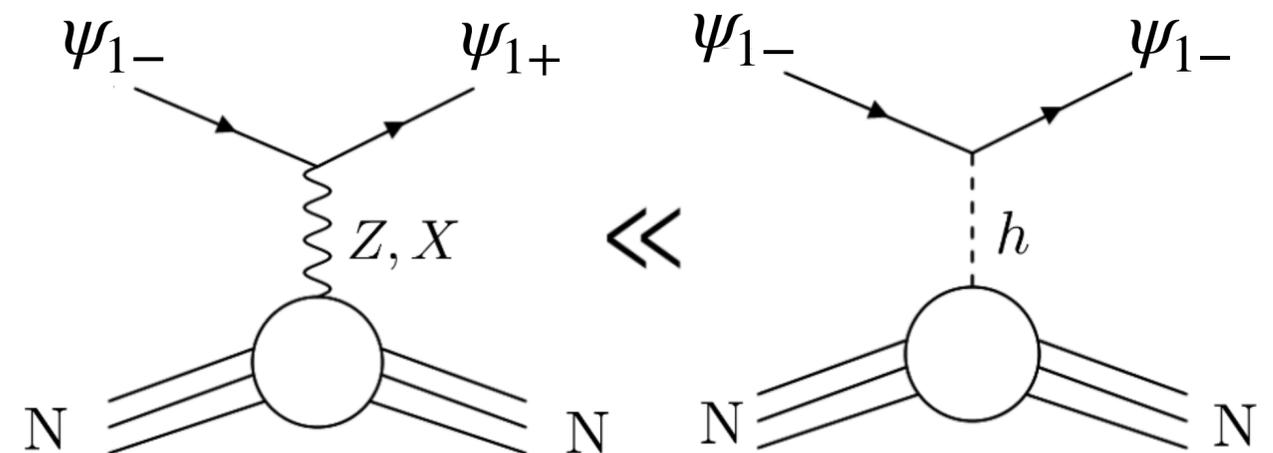
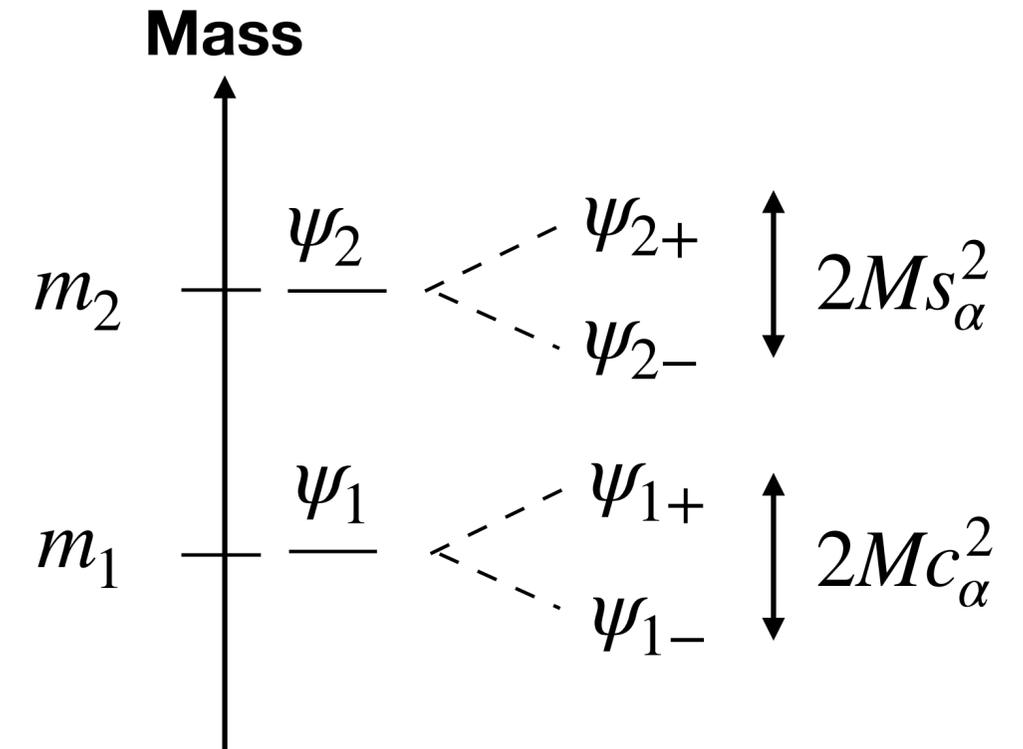
$$\Delta\mathcal{L} = y_N \Phi \bar{N}^C N + (h.c.)$$

- $N = (\chi_N \bar{\chi}_N^c)^T$ ,  $M = y_N \langle \Phi \rangle \rightarrow \Delta\mathcal{L} = M(\chi_N \chi_N + \bar{\chi}_N \bar{\chi}_N)$

- Splits Dirac Fermions  $\psi_{1,2}$  mass eigenstates into two pairs of Majorana fermions  $\psi_{1\pm}, \psi_{2\pm}$  with masses  $m_{1,2} \pm \Delta m_{1,2}$

$$-\mathcal{L} \supset -\frac{\lambda}{2\sqrt{2}} \sin(2\alpha + 2\gamma_-) h \bar{\psi}_{1-} \psi_{1-} - i \bar{\psi}_{1-} \gamma^\mu \psi_{1+} \left( \cos(\gamma_+ - \gamma_-) g_x X_\mu + [\cos(\gamma_+ - \gamma_-) - \cos(2\alpha + \gamma_+ + \gamma_-)] \frac{g_i}{4} Z_\mu \right)$$

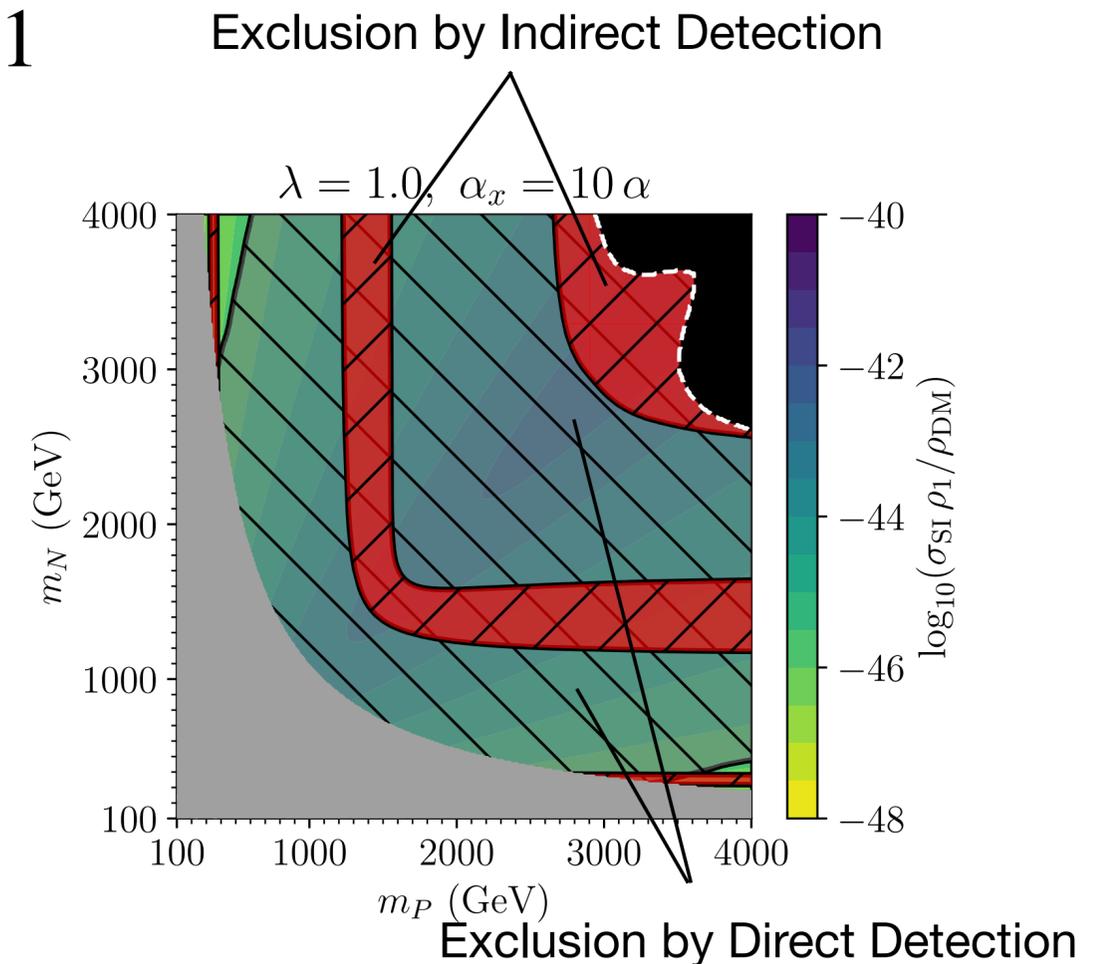
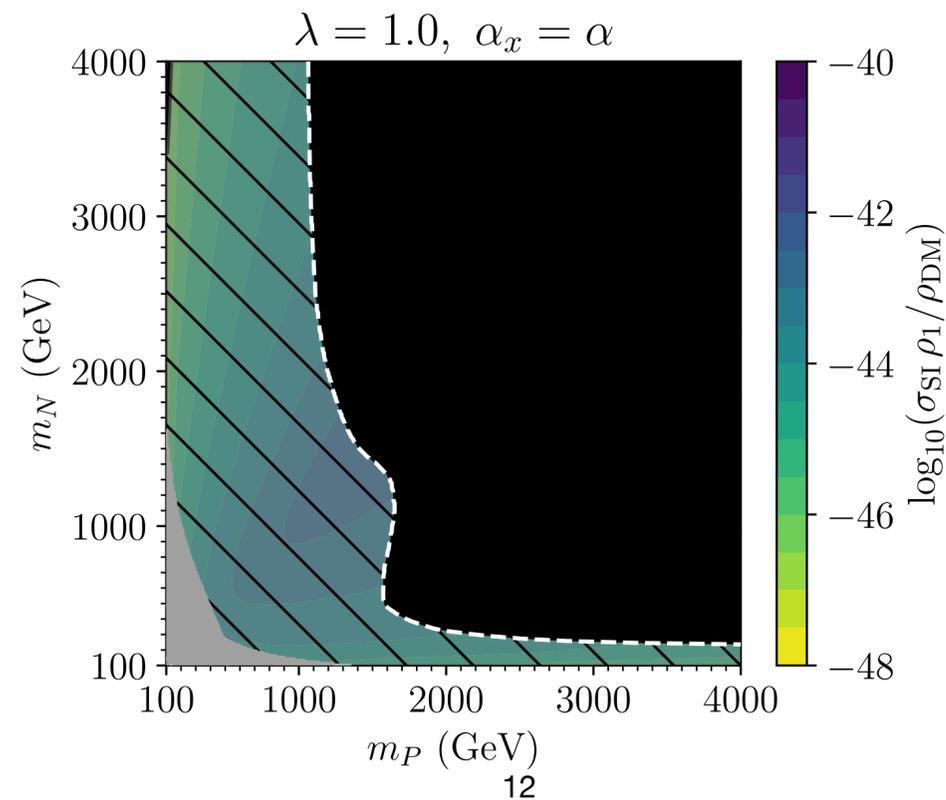
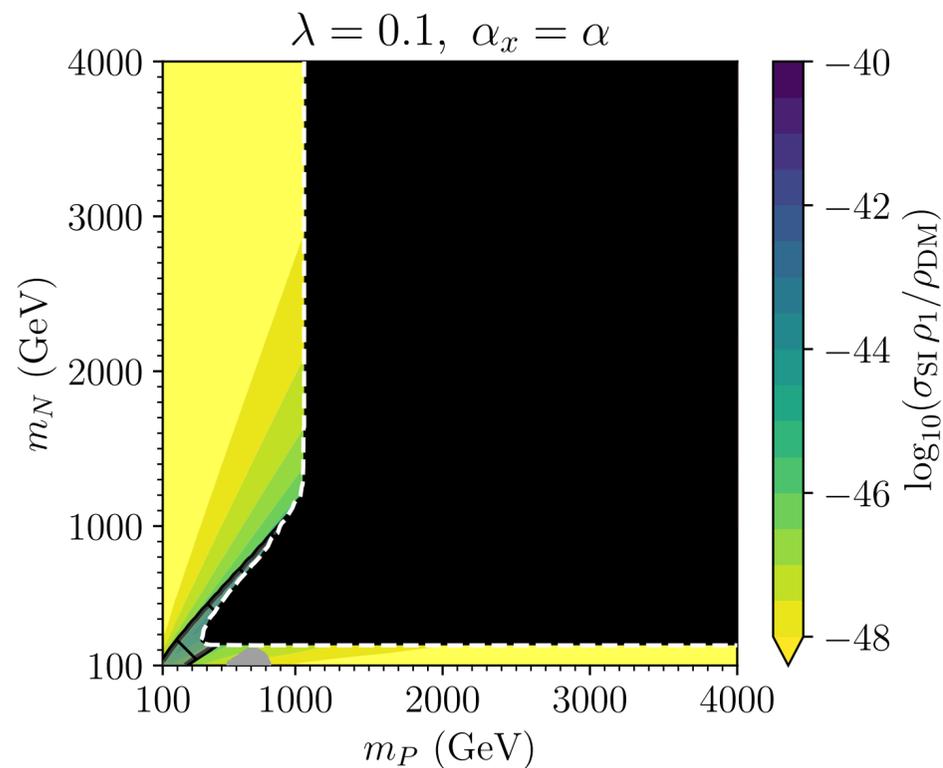
- Typical recoil energy in DD  $E_R \sim 100$  keV. Inelastic scattering with  $M \geq 200 - 500$  keV are kinematically suppressed



# Fix #1

## Mass splitting from a Majorana mass term

- For  $M \ll m_1$  and  $M \geq 10 \text{ MeV}$  relic density remains identical.
- For  $\alpha_x = 10\alpha$ , annihilation at late time  $\psi_1 - \bar{\psi}_1 \rightarrow XX$  probed notably by distortion of the power spectra of the CMB measured by Planck. Excluded region exhibits a band structure corresponding to enhancement through the formation of bound states, dependent on  $m_x/m_{DM}$
- Bounds from direct searches remain unchanged for  $\Delta m/m_x \ll 1$



# Fix #2

## Decay Through Lepton Mixing

- Avoid overabundance of relic portal fermions by allowing them to decay quickly to SM
- Dark Higgs field  $\phi \sim (1,1,0; q_x)$

$$-\mathcal{L} \supset \lambda_a \phi \bar{P}_R L_{La} + (h.c.), \quad a = e, \mu, \tau$$

- $\langle \phi \rangle = \eta$  induces mixing with leptons,  $\psi_1 \rightarrow \nu_{La} \phi$  and  $\psi_1 \rightarrow \nu_{La} X$

$$\tau \simeq (6.61 \times 10^{-8} \text{ s}) \left( \frac{10^{-9}}{\lambda_a s_\alpha} \right)^2 \frac{1 \text{ TeV}}{m_1}$$

- As long as the couplings are not exceedingly small,  $\lambda_a s_\alpha \gtrsim 10^{-12}$  these decays occur before primordial nucleosynthesis and neutrino decoupling, and will generally be safe from cosmological bounds
- Simultaneously contributes to  $BR(\tau \rightarrow \mu\gamma)$ ,  $BR(\mu \rightarrow e\gamma)$  and  $\Delta a_{e,\mu}$
- Challenge: Can  $\psi_1$  decay fast enough while avoiding bounds from LFV and  $\Delta a_{e,\mu}$  ?

# Fix #2

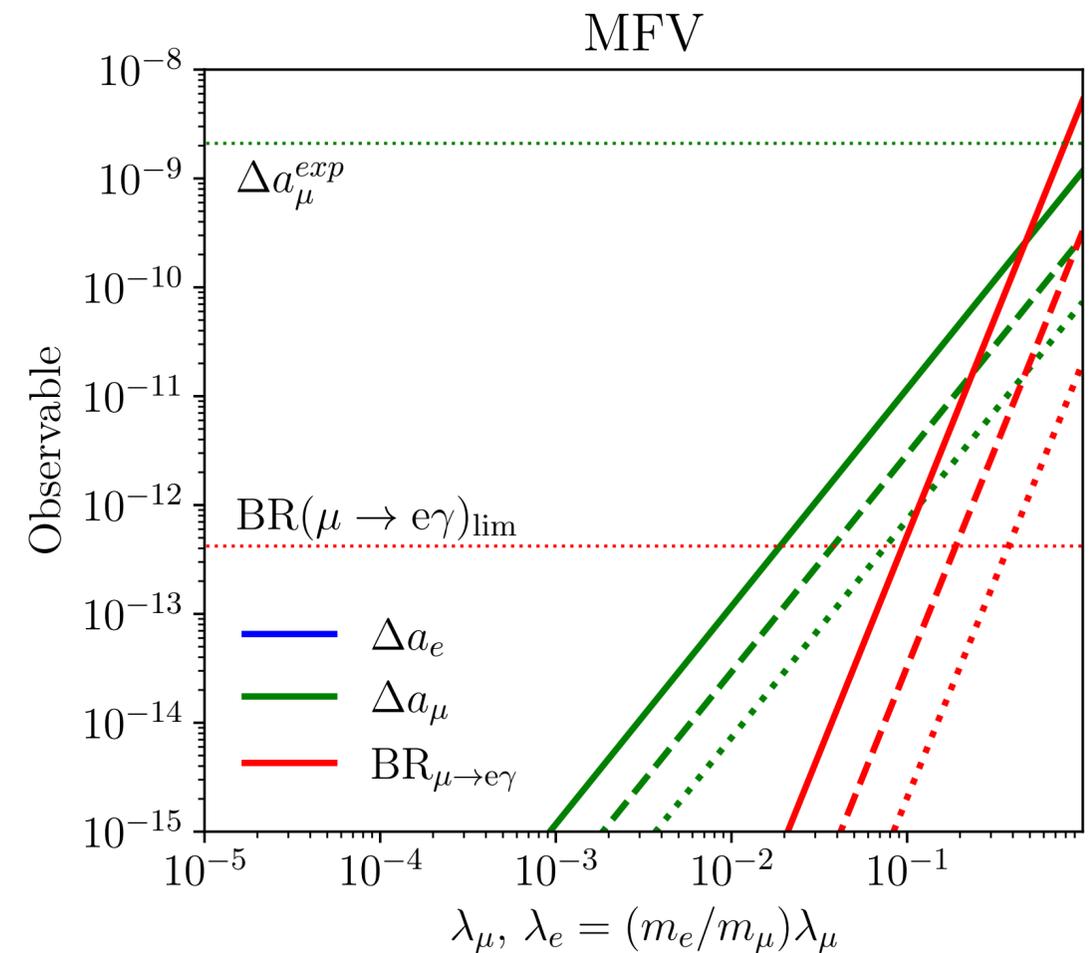
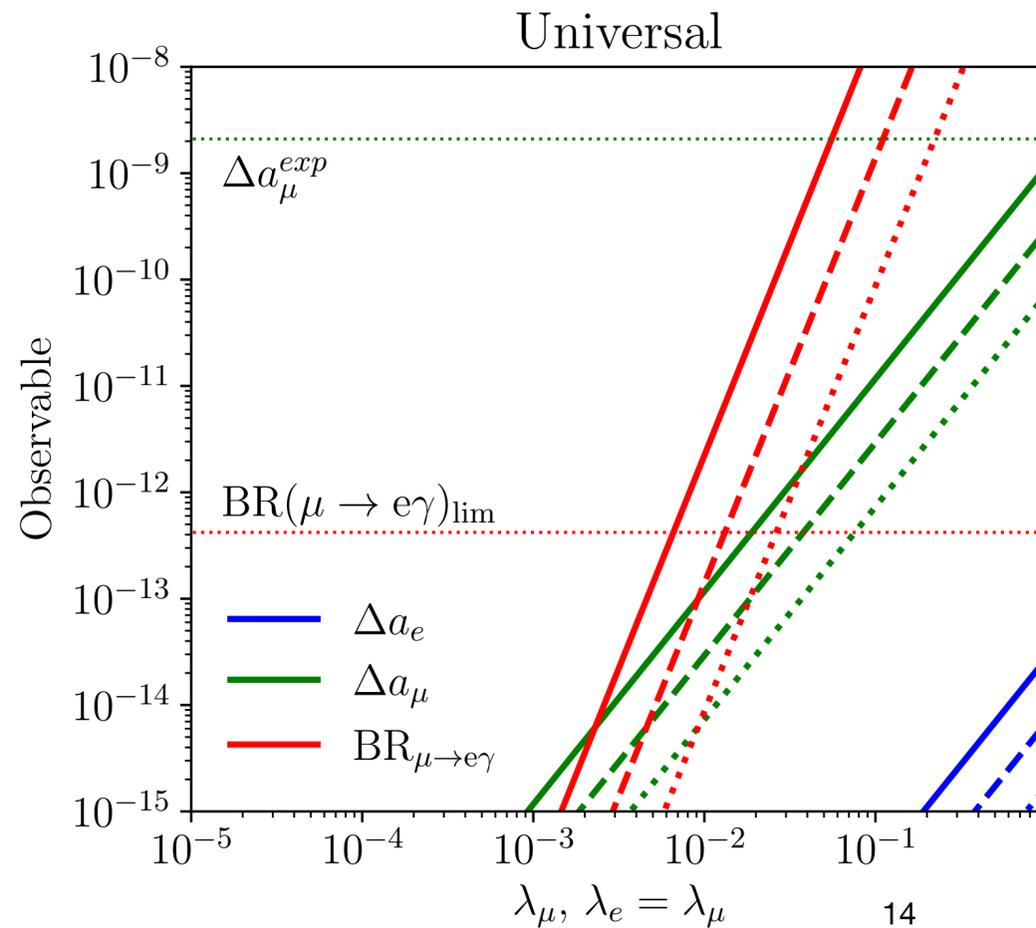
## Decay Through Lepton Mixing

Yes !

- $\Delta a_{\ell_a} = + \frac{\lambda_a^2}{96\pi^2} \left( \frac{m_a}{m_P} \right)^2$

- $\text{BR}(\mu \rightarrow e\gamma) = \frac{12\pi^3}{m_\mu^4} \frac{\alpha}{G_F^2} \left( \frac{\lambda_e}{\lambda_\mu} \right)^2 \times (\Delta a_\mu)^2$  ,  $\text{BR}(\tau \rightarrow \mu\gamma) = \frac{12\pi^3}{m_\mu^4} \frac{\alpha}{G_F^2} \left( \frac{\lambda_\tau}{\lambda_\mu} \right)^2 \times (\Delta a_\mu)^2 \times \text{BR}(\tau \rightarrow \mu\nu\bar{\nu})$

$m_P = 100 \text{ GeV}$  ———  
 $200 \text{ GeV}$  - - - -  
 $400 \text{ GeV}$  ·····



# Fix #2

## Collider signature

- $\psi_1$  is now unstable and can decay to visible particles in colliders.  
Signatures are:  $X \rightarrow f\bar{f}$  and  $\phi \rightarrow XX$
- $\lambda_a s_\alpha \leq 10^{-10}$  yields similar signature as the original setup (long-lived in the detector), may be visible in far detectors such as FASER, MATHUSLA,...
- For larger couplings,  $\psi_1$  decays promptly on typical collider timescale. In the limit  $m_x \ll m_1$ , dark vector decay product will be boosted  $\rightarrow$  lepton jets.  
Similar searches exist for  $h \rightarrow XX \rightarrow 4\ell$  at ATLAS and CMS. To our knowledge, there is no directly remappable existing analysis to constrain our model.

# Conclusion

- We study a minimal model of connector fermions to a dark gauged  $U(1)_x$
- A naive model is ruled out as it includes a DM candidate ruled out by DD searches.
- Fix #1 : Through a small Majorana mass term for  $N$ , DM candidate scatters inelastically in DD experiments for vector bosons exchanges. Model is viable for  $\lambda \simeq 0.1$
- Fix #2 : Couple  $P$  to LH SM leptons, so that  $\psi_1$  decays in the early Universe. Viable for  $10^{-12} \lesssim \lambda_a s_\alpha \lesssim 10^{-3}$ , could be searched at collider

Merci !

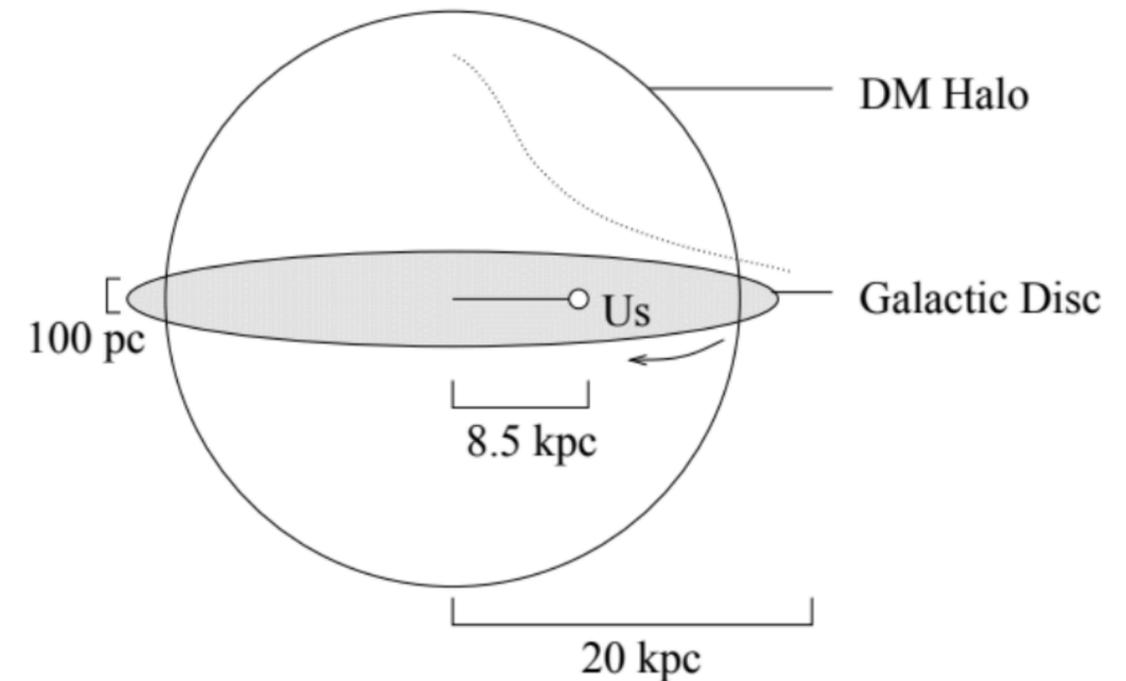
# Backup

# Direct Detection with Inelastic Scattering

- Typical velocity of DM particles relative to the Earth:  $v \sim 10^{-3}$

- $$E_R = \frac{2\mu_N^2 v^2 \cos^2(\theta)}{m_N} \sim 100 \text{ keV}$$

- $$\mu_N = \frac{m_N m_\chi}{m_N + m_\chi}$$



# Higgs Portal Operator

## Higgs Portal Operators

- $\mathcal{L} \supset \frac{\alpha_x}{6\pi} \frac{\lambda^2}{m_1 m_2} H^\dagger H X_{\mu\nu} X^{\mu\nu}$  in the limit  $m_h \ll m_{1,2}$