Feebly Interacting Particle searches with positrons



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IP2I – CNRS

04/11/2022

Based on 2209.09261 and 2211.xxxx with M. Raggi and E. Nardi



Outline

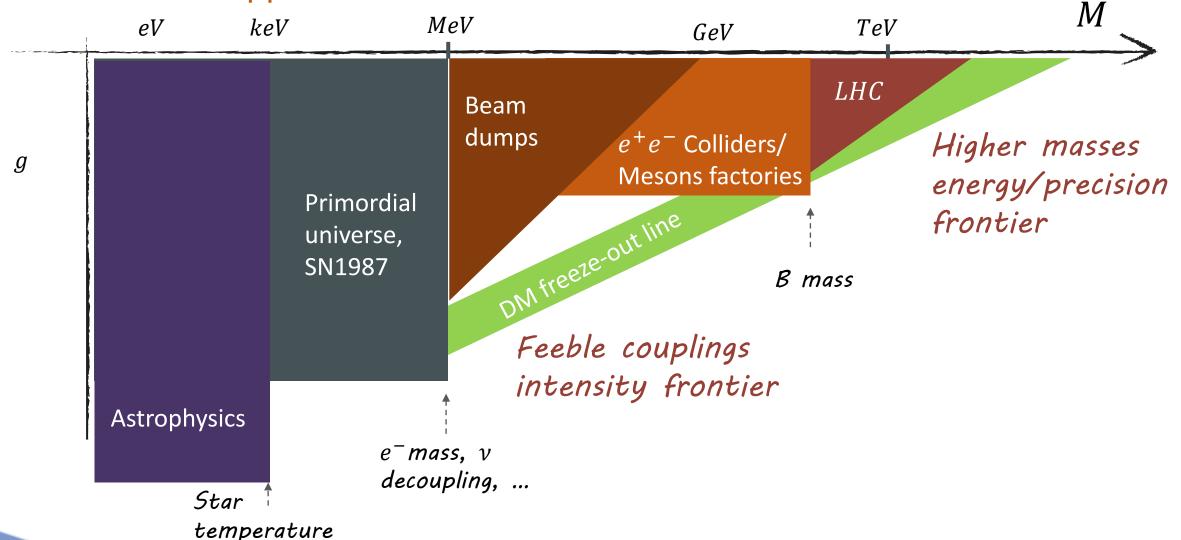
Introduction: using resonant production

Two experimental strategies for positron beams

Searches for X17 in PADME

Mapping the known particles: FIPs

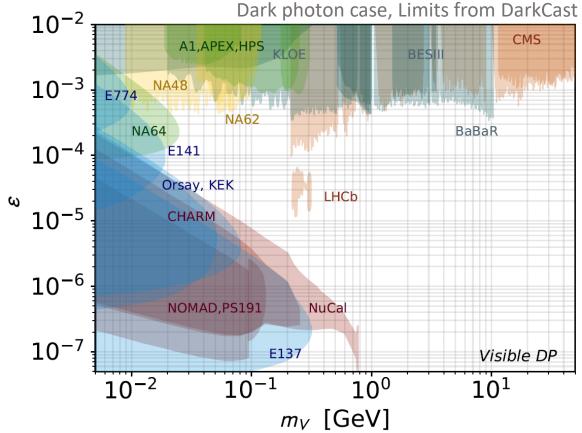
• Feebly Interacting Particles (FIPs) = "new neutral particle which interacts with the SM via suppressed new interactions"



Going forward in the MeV to GeV region?

 Closing the (in)-famous Mont's gap, between long-lived search and "visible" decay strategy

For the long-lived limit, we want to do better than experiments from the 80s-90s ...



Going forward in the MeV to GeV region?

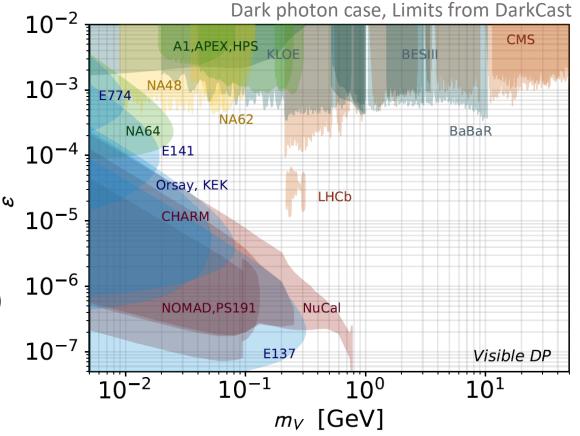
 Closing the (in)-famous Mont's gap, between long-lived search and "visible" decay strategy

→ For the long-lived limit, we want to do better than experiments from the 80s-90s ...

 Various experimental strategies pursued

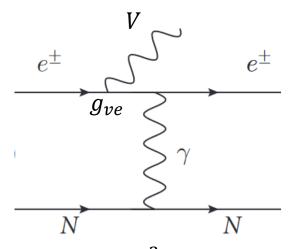
- → Higher boost factor (FPF, SHIP, etc...)
- → Smaller beam dump size (NA64) / Displaced vertices (LHCb)
- → More statistics for bump-search method (Belle-II) ... and many others!

This talk → use a different production mechanism

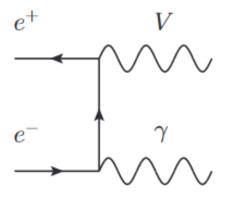


Going resonant ...

Bremsstrahlung process



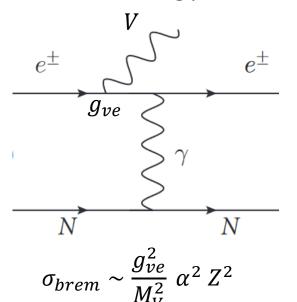
Associated annihilation process



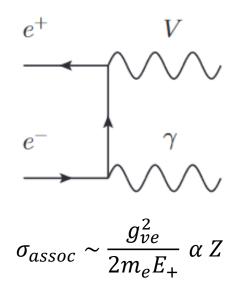
$$\sigma_{assoc} \sim \frac{g_{ve}^2}{2m_e E_+} \alpha Z$$

Going resonant ...

Bremsstrahlung process



Associated annihilation process





- → Cross-section x100 times larger
- →Scales as Z only

$$\sigma_{res} \sim \frac{g_{ve}^2}{2 m_e} \pi Z \delta(E_+ - E_{res})$$

 g_{ve}

- What are the trade-offs for resonant production ?
 - → First, we need to find positrons somewhere. Typically, this implies a certain loss in energy + beam intensity
 - → Then we need to hit the resonant energy (works mostly for 10-100 MeV range)

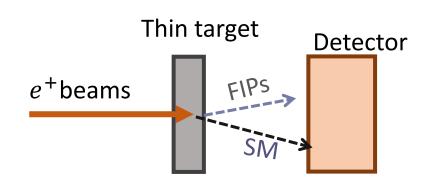
How to get to the exact energy?

• Study models with large invisible width $\Gamma_V^{inv} \rightarrow$ Typically extremely important for DM-motivated models !

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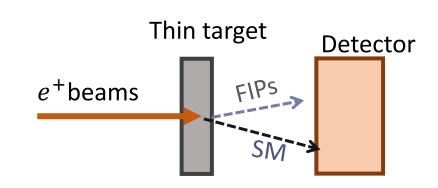
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- Vary the beam energy
- → "Scanning" procedure is required, varying the beam energy on non-negligible range See e.g. 1802.04756

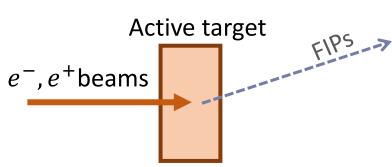


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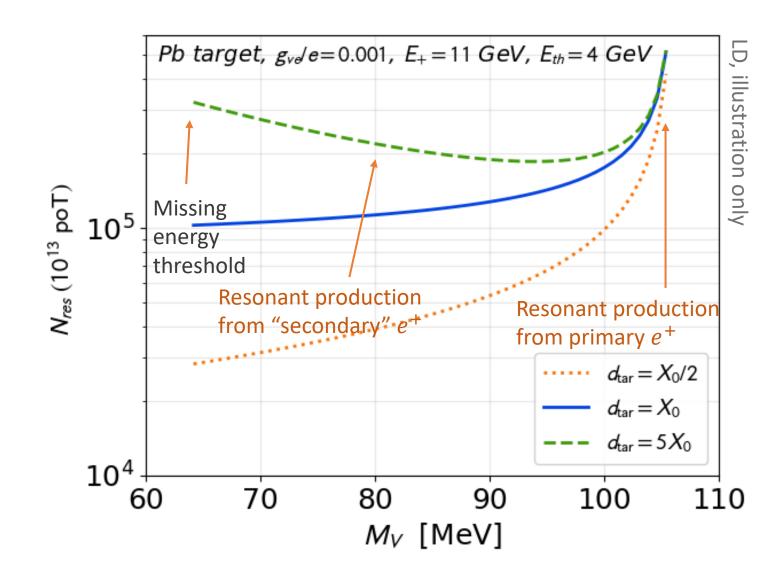


- Use energy loss and secondary e^+ production in the target to "scan" naturally various positron energies
 - → Requires a "not-too-thin" target to allow some evolution of the beam
 - → Works to a certain extent also in electron-based machines



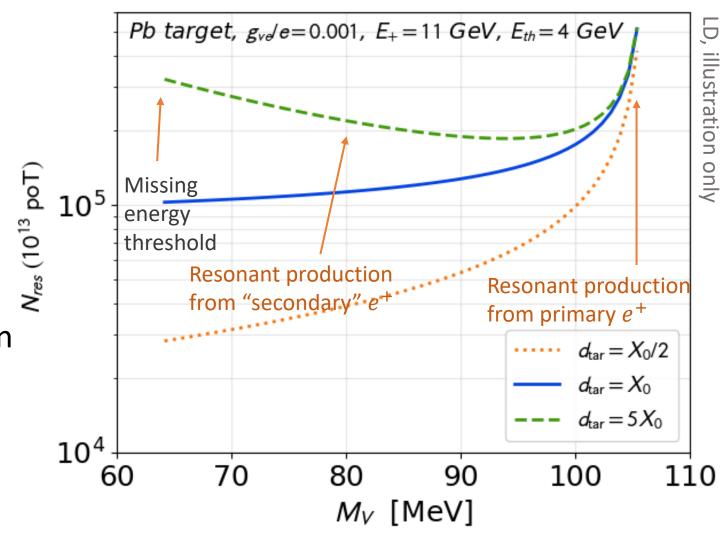
The thick target approach

- Use straggling and bremsstrahlung processes to degrade the beam energy
- Effective to probe a large range of masses without varying the beam energy too much



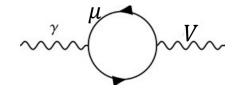
The thick target approach

- Use straggling and bremsstrahlung processes to degrade the beam energy
- Effective to probe a large range of masses without varying the beam energy too much
- But FIP production occurs directly in the shower
 - → Requires either a displaced signal or missing energy to escape background
 - →This works as soon as we have a coupling to neutrinos ...

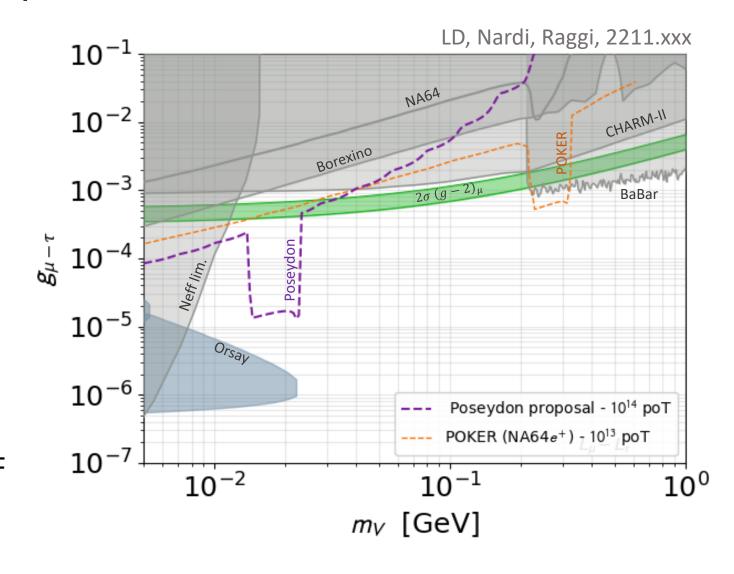


In practice: limits on $L_{\mu}-L_{ au}$ gauge boson

 Use radiatively generated kinetic mixing for the production stage

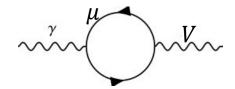


- Invisible decay into neutrinos
 - → Recent limits from NA64 promising!
- Projections for
 - \rightarrow Poseydon (based on the e^+ LNF beam at 0.5 GeV)
 - \rightarrow NA64- e^+ (with ~100 GeV beam) From 2206.03101

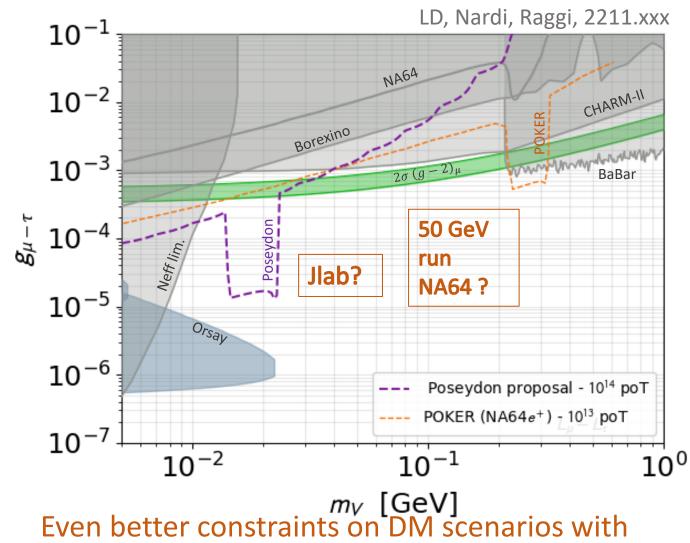


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"tree-level" e^{\pm} couplings...

Resonant production: thin target

- Main idea: use resonant production and search for visible FIP decay in a noisy environment
 - \rightarrow Mostly relevant with e^+e^- final states (hard to produce a FIP with mass above the di-muon threshold resonantly)
- Vary the beam energy, fit the background, and look for resonance
 - → Simple analysis strategy

$$\mathcal{N}_{X_{17}}^{\,\mathrm{per\ poT}}(E) = rac{\mathcal{N}_A Z
ho}{A} \ell_{\mathrm{tar}} rac{g_{ve}^2}{2m_e} f(E_{\mathrm{res}}, E)$$

With f the beam spread, typically modelised by a Gaussian distribution with spread δE

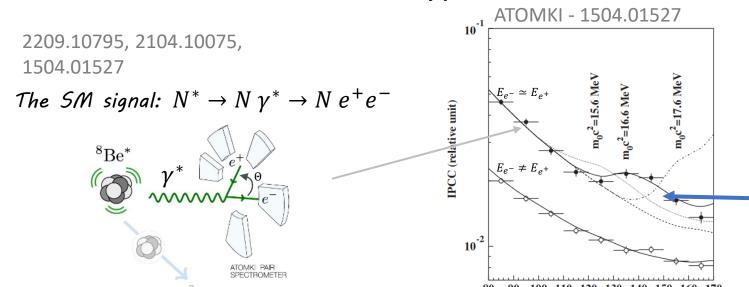
Thin target

Detector

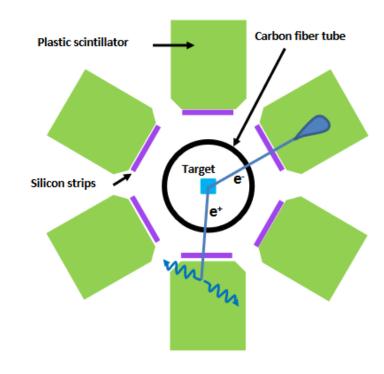
- Main background is from Bhabha scattering, but can be fitted directly from the data
 - → "Large angular acceptance" detector important to reduce the t-channel contribution

The X17 anomaly

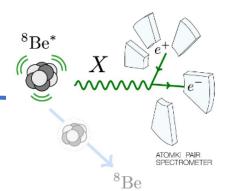
- The signal: a possible 17 MeV boson in the ATOMKI spectrometer?
 - ightharpoonup Production in excited nuclei 12 C, 8 Be and 4 He, followed by radiative decay $N^* \to N \ \gamma^* \to N \ e^+ e^-$
 - → Very large excess, cannot be statistical but from one experiment only.
- How can we test the NP hypothesis?



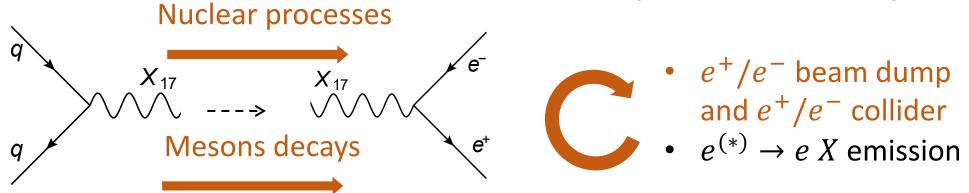
 Θ (deg.)



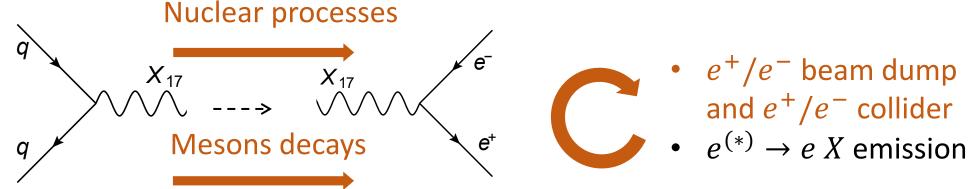
NP sigma: $N^* \rightarrow N V \rightarrow N e^+e^-$



PADME and the X17 boson, the perfect target



PADME and the X17 boson, the perfect target



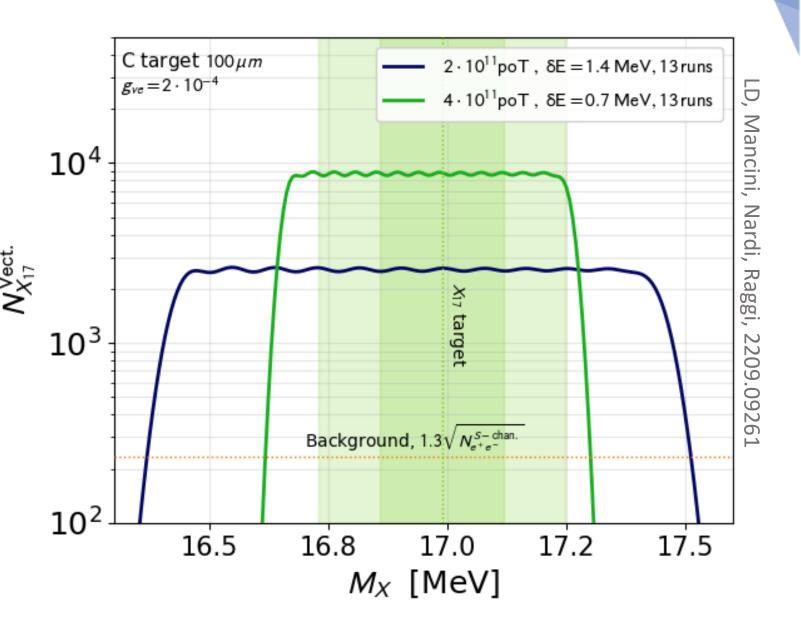
• We look for a light boson decaying to mostly to e^+e^- with mass:

$$M_X = \begin{cases} 16.70 \pm 0.35 \pm 0.50 \text{ MeV} \\ 17.01 \pm 0.16 \text{ MeV} \\ 16.94 \pm 0.12 \pm 0.21 \text{ MeV} \end{cases}$$
 *8Be 1504.01527 + cds.cern.ch/record/2312578

- The narrow mass range plus model-independent e^{\pm} couplings makes this anomaly a perfect target for a resonant search!
- The target energy range is [270 290] MeV \rightarrow perfectly adapted to e^+ beam in Frascati

Scanning strategy

- Several runs depending on the beam spread δE
 - \rightarrow Smaller spread implies lower background as the signal a "bump" with spread δE
 - → Currently only LNF's accelerator complex can provide a positron beam and vary its energy
- Quite a lot of theory involved: radiative return effects with use of NLO, continuous vs discrete description for narrow width



Projections for PADME – X17

 Complete simulation based on the current PADME setup

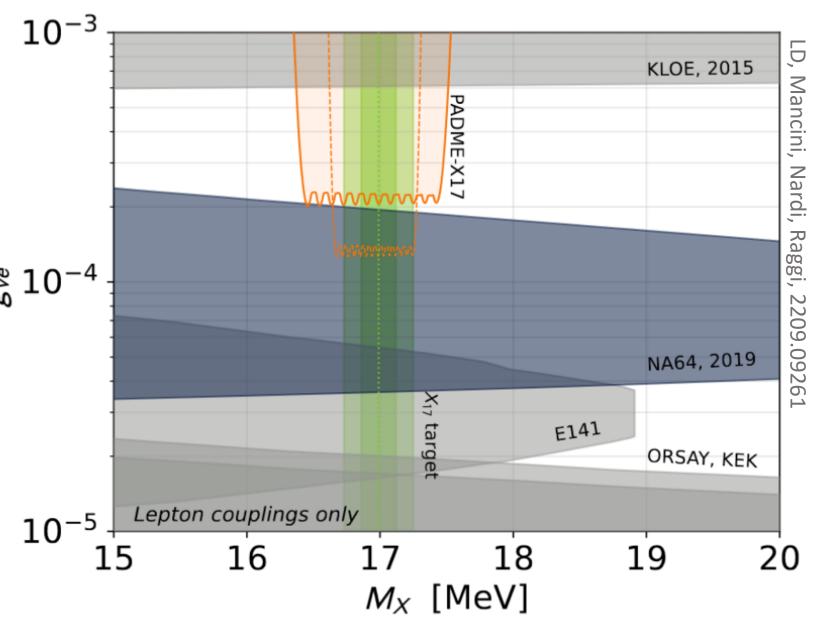
 \rightarrow Conservative: 2 · 10¹¹ PoT, a 0.5% beam spread

 \rightarrow Aggressive: $4 \cdot 10^{11}$ total PoT, a 0.25% beam spread

• Rely on e^+e^- couplings only

→relevant for dark photon

 Serve as a first "test run" for these kind of analysis + test the vector X17 NP hypothesis



Conclusion

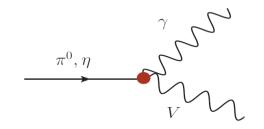
Conclusion

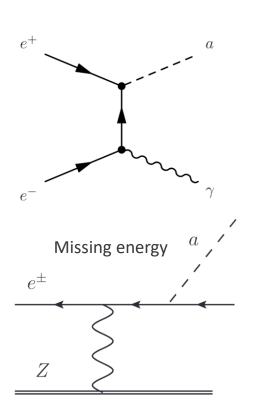
- Positron-based facilities allow to leverage resonant production to significantly increase signal rates
 - → With the cost of having to scan over a large energy ranges
- In the future, this may be leveraged either from
 - → Thick target + missing energy (works both for DM-models and for neutrino decays)
 - → Thin target with visible decays + scan in beam energy
- In both cases, close to an order of magnitude improvement compared to existing limits
- First experimental example of the later strategy will be completed shortly by the PADME collaboration, with the X17 anomaly as a target

Backup

Accelerator facilities (currently) available

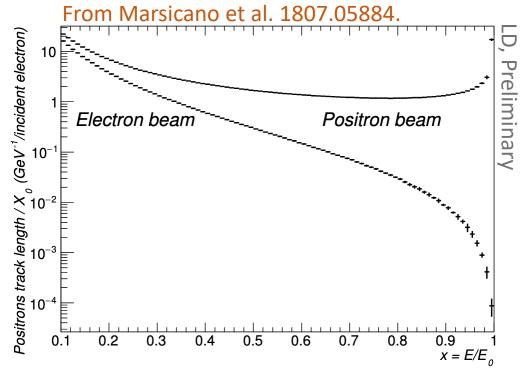
- Intensity beam dumps: typically, p machines (beam neutrinos exp, SHiP).
 - Large backgrounds + protophobia of X17 + far away detectors → Challenging for X17
- e^+e^- colliders (BaBaR, Belle-II ...)
 - Good production rates, large luminosity, but also background control and the small p_T for the e^+e^- pair → Still interesting avenue for X17 (displaced vertices?)
- e^+e^- beam dumps: typically, e^+ or e^- machine (NA64, PADME, MAGIX, etc...)
 - Large production rates, can search for displaced vertices or reconstruct the e^+e^- pair \rightarrow particularly suitable





• Rare meson/lepton decays -> Promising, but with model-dependence

Secondary positron production

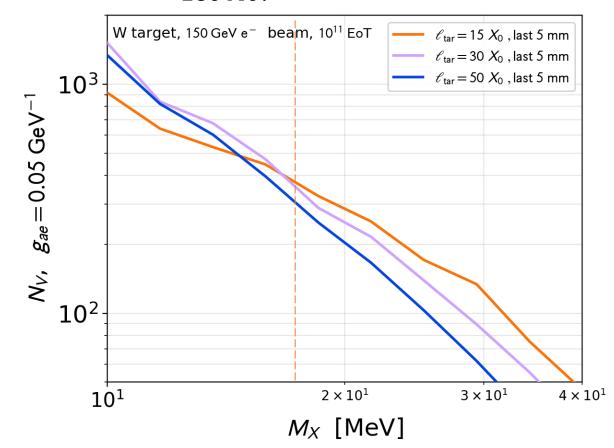


→ X17 resonant production occurs at any point in the target, including at the end

→ Background from the residual shower likely to swamp the signal

 A secondary positron population build up the shower "convert energy to statistics"

$$N_{e^{+}}^{X17} \sim \frac{E_{ini}/2}{280 \, MeV} N_{e^{-}}^{ini}$$



Energy matters for decay lengths!

Bremsstrahlung extracts most of the energy of the beam

$$\gamma_{X17}\ell_{X17} \sim 3 \text{ cm} \left(\frac{E_{X17}}{100 \text{ GeV}}\right) \left(\frac{17 \text{ MeV}}{m_{X17}}\right) \left(\frac{3 \cdot 10^{-4}}{g_{Xe}}\right)^2$$
 (Vector) $\gamma_{X17}\ell_{X17} \sim 3 \text{ cm} \left(\frac{E_{X17}}{100 \text{ GeV}}\right) \left(\frac{17 \text{ MeV}}{m_{X17}}\right) \left(\frac{0.5 \text{ GeV}^{-1}}{g_{Xe}}\right)^2$ (ALP) experiments

By contrast, X17 from resonant production have relatively low energy

$$E_{X17}^{\mathrm{res}} = \frac{m_{X17}^2}{2 \, m_e} \simeq 280 \, \mathrm{MeV}$$
 \longrightarrow $\gamma_{X17}^{\mathrm{res}} \simeq 15$ \longrightarrow Displaced signatures viable only for the lowest allowed couplings

• However, resonant production implies that the decay production satisfy precisely both $E_{X17}^{\rm res} \simeq 280~{\rm MeV}$ and $m_{ee} \simeq m_{X17}$

Fixing notations: explicit Lagrangian for X17

• An axion-like particle (ALP) a, interacting via $\bar{f}\gamma^{\mu}\gamma^{5}f$

$$\mathcal{L} \subset \frac{1}{2} (\partial_{\mu} a)(\partial^{\mu} a) - \frac{1}{2} m_{a}^{2} a^{2} + \sum_{f = \ell, q} \frac{g_{af}}{2} (\partial_{\mu} a) \, \bar{f} \, \gamma^{\mu} \gamma^{5} f \qquad \qquad \underbrace{\begin{array}{c} g_{af} \text{ corresponds to} \\ \frac{Q_{af}}{f_{a}} \text{ in Daniele} \\ \text{Alves's talk} \end{array}}_{\text{Alves's talk}}$$

• A light vector V^{μ} , potentially with both vector and axial couplings

$$\mathcal{L} \supset -\frac{1}{4} V_{\mu\nu} V^{\mu\nu} + \frac{1}{2} M_V^2 \, V_\mu V^\mu + \sum_{f=\ell,q} V_\mu \, \bar{f} \, (g_{Vf} + \gamma^5 g_{Af}) f \, \longrightarrow \, \begin{array}{c} g_{Vf} \text{ corresponds to} \\ e \varepsilon_f \text{ in Jonathan} \\ \text{Feng's and Tim Tait's} \\ \text{talks} \end{array}$$

Most of the e^+/e^- -driven production rates shown in the rest of the talk satisfy approximately:

 $m_{\ell}g_{a\ell}\longleftrightarrow g_{V\ell}$

 e^+/e^- -driven production rates are pretty agnostic concerning the X17 nature/couplings

The X17: couplings

- Need a large couplings to quarks, but the actual couplings target depends on the X17 nature
 - → As a reference for the vector case

$$|g_{Vu}+2\ g_{Vd}| \sim [0.6\cdot 10^{-3}, 3\cdot 10^{-3}]$$
 See e.g. 1608.03591, assumed BR_{ee} at 1. $|2g_{Vu}+g_{Vd}| \lesssim 0.4\cdot 10^{-3}$ Huge couplings! Protophobia needed to escape NA48

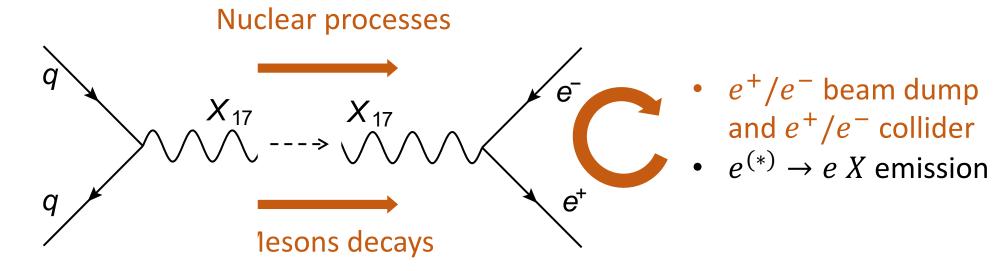
- It has to decay (mostly) visibly into e^+e^-
 - → For ATOMKI result, coupling with electron constrained only by a lower limit to ensure decay length smaller than ~cm (we will discuss it in detail in this talk)
 - → Can have an invisible BR to e.g. a new dark sector particles but leads to even larger coupling to quarks
 - \rightarrow Strong constraints on neutrinos interactions from $v_e e^-$ scattering experiments

X17: widths and productions

• Combined, the above requirements imply that the X17 must have a tiny width, mostly driven by the e^+e^- decay

Vector case
$$\Gamma_X \sim \frac{g_{Ve}^2}{12\pi} M_V \sim 0.5 \text{ eV} \times \left(\frac{g_{Ve}}{0.001}\right)^2 \qquad \qquad \text{More challenging}$$
 to produce it on resonance

Altogether we have the following situation



Rare decays searches

- Rare decays probes are both extremely effective in probing X17, often at the price of a large model dependence
- Mesons decay probes (example from mostly last year)

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 \begin{array}{c} & \text{hep-ex/0610072} \\ \hline \\ 0 & \pi^0 \rightarrow \gamma V_{17}, \text{ for vector states: NA48 bounds implies proto-phobic} \\ \hline \\ 0 & J/\Psi \text{ decays, charm couplings only} \\ \hline \\ 0 & B^* \rightarrow B \ V_{17}, D^* \rightarrow D \ V_{17} \text{ for vector states} \\ \hline \\ 0 & \pi^0 \rightarrow a_{17} \rightarrow e^+e^-, K \rightarrow \pi(\pi)a_{17}, K \rightarrow \mu\nu \ a_{17} \\ \hline \\ 0 & \pi^0 \rightarrow a_{17} \ a_{17} \ a_{17} \text{ and other multi-leptons final states} \\ \hline \\ 0 & \pi^0 \rightarrow a_{17} \ a_{
```

- If flavour-violation, many more available channels both in lepton decays and in "standard" flavoured meson decay.
- Also radiative emission from μ decay