# Rare event searches with LiquidO technology

Diana Navas Nicolás On behalf of the LiquidO consortium



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### Some remarks to begin with ...

Neutrinos are rare events in this talk

 This presentation is not focused on other rare events (double beta, proton decay, dark matter ... )

 ... but its goal is to highlight the possible benefits of the LiquidO technology in the context of Deep Underground Physics, such as the extra handles for the control of cosmogenics

### Outline

- What is LiquidO?
- Physics potential
- Experimental validationFuture prospects and conclusions

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### A NEW APPROACH

Stochastic light confinement near its creation point by using opaque medium





**Transparent:** Today's technology Topology information washed-out Opaque: LiquidO technology Light clustering

### INNOVATIVE DETECTION TECHNIQUE

- Maximal light collection by a dense array of fibers connected to SiPMs
- ★ Low background, high efficiency
- ★ Fast time resolution (< 0.1 ns)
  - **Excellent spatial resolution** ( $\leq$  1 mm)



#### Top view: (x,y) projection



Bottom view: (x,y) projection

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#### NOVEL ENGINEERING SOLUTIONS



- ★ 1x Axis(Z) low cost & simplicity
- ★ (X,Y): topology  $\rightarrow$  mm resolution (robust)
- C: timing → few cm resolution



- ★ "1x" Axis(twisted-Z @  $\leq$ 10°) development
- ★ (X,Y): topology  $\rightarrow$  mm resolution (robust)
- ★ Z: topology  $\rightarrow$  ≤1cm resolution (robust)
- ★ (X,Y,Z): timing  $\rightarrow$  over-constrain & energy-flow

#### OPAQUE MEDIUM



### NoWaSH [prototype]

- ★ Opaque Liquid Scintillator
- ★ Linear Alkyl Benzene (~80 wt.%) + Paraffin Wax (~20 wt.%) + PPO (~0.3%)

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- ★ Opacity depends on paraffin concentration (changes crystalisation temperature)
- ★ Short scattering length and moderate absorption length
  - arXiv:1908.03334

LiquidO R&D extensive field: new  $\mu$ Crystal scintillators arXiv:1807.00628, emulsion...

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#### UNPRECEDENTED IMAGING CAPABILITIES

#### ENERGY DEPOSITION 1 MEV C+



- ★ Particle Identification (PID) is a major challenge in MeV neutrino detection.
- ★ Confinement of light into sphere around each ionization point
- A self-segmented detector!
   (no need to introduce dead material for segmentation)
- ★ Discrimination of individual  $e^+$ ,  $e^-$  and  $\gamma$  events @1MeV

### UNPRECEDENTED PID @MEV SCALE



#### $\star v_e/\overline{v}_e$ separation

- Background (cosmogenic)
   identification
- No segmentation (less background for MeV physics)



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- ★ Topology of deposited energy
- ★ "Energy flow": time pattern for the light to be collected
- Able to achieve e<sup>-</sup> separation from γ with efficiency > 85% and contamination ~10<sup>-3</sup> @2MeV

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#### POWERFUL BACKGROUND REJECTION



Signal : Background ~ 30 : 1Signal : Background  $\gtrsim 10 \times 30 : 1$ Background: few/dayBackground: few/year

#### PID + vertex reconstruction

#### POSSIBILITY OF DOPING

- ★ Relaxing scintillator transparency requirement
- ★ Unparalleled affinity for loading
- ★ Plenty of room to explore unconventional scintillators (e.g. ultra high light-yield)

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#### SOLAR NEUTRINOS

Indium loading will allow to perform precise pp solar neutrino physics and beyond

$$v_{e} + {}^{115}\text{In} \rightarrow {}^{115}\text{Sn} + e^{-} \text{ (LENS)} \text{ delay}$$

$$\tau = 4,76 \,\mu\text{s} \left[ \begin{array}{c} e_{1}^{-}/\gamma_{1} & \text{E} = 116 \,\text{keV} \, (e/\gamma) = 0.96 \\ + \\ \gamma_{2} & \text{E} = 497 \,\text{keV} \end{array} \right]$$

Delay  $\begin{bmatrix} \overline{0} \\ 0 \\ -50 \\ -50 \\ 1 \\ 10 \\ 10 \end{bmatrix} \begin{bmatrix} 0 \\ 0 \\ 0 \\ -50 \\ 1 \\ 10 \\ 0 \end{bmatrix} \begin{bmatrix} 0 \\ 0 \\ 0 \\ -50 \\ 0 \\ -50 \\ 1 \\ -50 \\ -50 \\ 0 \\ -50 \\ -50 \\ -50 \\ -50 \\ -50 \\ 0 \\ -50 \\ -50 \\ 0 \\ -50 \\ -50 \\ -50 \\ 0 \\ -50$ 

THREE-FOLD COINCIDENCE BG-less ( $\lesssim 10^{-15}$ g/g) Low threshold = 114 keV (95.5% of pp v<sub>e</sub>)

 $E_{vis}=E_{\nu}-\text{threshold}$ 



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### FROM MEV TO MULTIMEV



- ★ Powerful PID
- ★ Energy Flow
- ★ Tracking (mm)
- ★ Directionality
- ★ dE/dx (range)

### FROM MEV TO MULTIMEV



### FROM MULTIMEV TO GEV



#### NUCLEON DECAY



- ★ Largest achievable free protons densitiy (scintillator)
- ★ Very high-efficiency
- ★ Full topological information and sign-ID for some channels through final Michel-e<sup>-</sup> (magnetize det)

#### NEUTRINOLESS DOUBLE BETA DECAY

- ★ Key advantages: background control and ability to load well beyond current limits
- ★ Looks like a very promising path for reaching deep into the normal ordering region!



#### SUPERNOVA NEUTRINOS, <sup>40</sup>K GEONEUTRINOS, REACTOR ANTINEUTRINOS...

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#### MICRO-LIQUIDO: FIRST EXPERIMENTAL PROOF OF PRINCIPLE

Article | Open Access | Published: 21 December 2021

#### Neutrino physics with an opaque detector

#### LiquidO Consortium

<u>Communications Physics</u> **4**, Article number: 273 (2021) | <u>Cite this article</u> **2530** Accesses | **3** Citations | **23** Altmetric | <u>Metrics</u>



More light collected by the fibres near the light source with the opaque sample!

#### MINI-II (UPGRADE): LIQUIDO'S PROTOTYPE DATA TAKING SINCE 2021



- $\star$  ~10 L with 64 readout fibres including a 3" PMT
- ★ e- from monoenergetic beam (<sup>90</sup>Sr) [0.4-1.8 MeV]
- ★ Stochastic light confinement observation
- ★ T cycle [5,40]°C powerful T control system (Chiller)
- ★ Very fast electronics: fast low-power custom preamplifier with sub-ns rise time
- ★ 64-channel WaveCatcher system for waveform digitization (ps time resolution)

#### WATER: SINGLE e CHERENKOV ONLY



Low amount of light: Cherenkov only & transparent (LiquidO's lowest acceptance)

Validation detector's integral timing readout

### LAB: CHERENKOV $\oplus$ Scintillation



~8.7x more light due to LAB's scintillation [no PPO]

Cherenkov excites the scintillator — loss ≥50% (optimisation)

Normalised Response (A.U.)

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1.8 MeV e

- 1.6 MeV e

— 1.4 MeV e — 1.2 MeV e

- 1.0 MeV e

- 0.8 MeV e

0.6 MeV e

LAB (no PPO)

250

300

Time (ns)

transparent

200

150



50

100

# Experimental validation

### TRANSPARENT LS: CHERENKOV VS SCINTILLATION

★ Transparent liquid scintillator (LS)

Remarkable separation using only timing
 Cherenkov light production threshold

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NW at 40°C "Transparent" — effectively like LAB or Water Cherenkov reduced by paraffine? — under investigation

#### NW@5°C: CHERENKOV ⊕ SCINTILLATION



~2x more light due to LiquidO's aggressive scattering

**Faster collection and better light containment** 

Formation topology → stochastic light confinement → LiquidO

#### OPACITY DEMONSTRATION



#### LIGHT YIELD EXPLORATION



~80% light collected within 5 cm's

Effective detected light yield >120PE/MeV [@ SiPM]

≥250PE/MeV — optimisation (ongoing engineering)

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TOPOLOGY PID (NO TIMING)

#### PID e/ $\gamma$ should be $\geq$ 100:1 rejection @ $\geq$ 90%



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### Future prospects

#### PROJECTS SEQUENCE AND TIMELINE

#### MINI- $\gamma$ project





**vAM-Otech project** Applied & Innovation R&D





UK Research and Innovation

~ 5 tons 2022-2027 + LPET-Otech project (medical approach) 2022-2024 ANR funded

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# Future prospects SUPER CHOOZ PROJECT

CNIS



**Chooz-B:** Reactor Cores

#### Chooz-A: Cavern Reactor Core

Ultra Near Detectors-

•LiquidO technology •Mass: ≤5ton

•Overburden: ≤3m

•Baseline: ≤30m

#### Chooz-A: Super Far Detector

LiquidO technology
Mass: ~ I Okton

•Overburden: ≤100m

•Baseline: ~I km

the Meuse river

 $\leq$  10 ktons

≥2030

### LiquidO consortium



LiquidO consortium consists of more than 70 scientists in 22 academic institutions (universities and/or laboratories) in over 10 countries

LiquidO@IN2P3: CPPM, IJCLab, IPHC, LNCA, LP2IB, Subatech (6 laboratories)

More info https://liquido.ijclab.in2p3.fr

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### Conclusions

- LiquidO is a novel detector concept  $\rightarrow$  opaque LS
  - Light confiment close to its creationpoint by a short scattering length
  - Light collection by a dense grid of wavelength shifting fibers + SiPMs
- LiquidO enables highly efficient PID
  - Event-by-event topological discrimination power
  - Powerful background discrimination
- Possibility of loading dopants at high concentrations
- Extensive and exciting physics potential (solar, reactor, SN, geoneutrinos...)
- Experimental validation of light confinement (MINI-II prototype)
- Future projects under development!

# Thank you!

# Back-up



- 1-cm-pitch lattice running along the z-axis
- Probability of misidentifying a γ as an e<sup>-</sup> vs. the efficiency of selecting e<sup>-</sup>
- mean scattering length  $\lambda s$  of either 1 mm or 5 mm
- photon detection efficiency ε of 3% (fibre trapping efficiency (~10%) and SiPM QE (~50%)
- The grey curve shows the probability of misidentifying a 2 MeV  $\gamma$  as an e<sup>-</sup> is estimated to be at the 10<sup>-2</sup> level with an efficiency of 87% for  $\lambda$ s=5 mm.

#### **BEAM PHYSICS**



**GeV-neutrino interactions** 

MINI-II SET UP



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- WATER: Cherenkov (transparent)
- LAB: Scintillator + Cherenkov (transparent)
- NoWASH-0: LAB + PPO
- NoWASH-20: addition of paraffin wax into the LAB scintillator with PPO shifter
  - Scattering lengths in the mm range without significant absorption
  - Light yield losses are small
  - Radiopurity estimates are promising
  - Open the possibility of using techniques for high loadings which were rejected from attenuation length limitations



### Electronics





- Analog memory digitiser using SAMLONG chip (CNRS/IJCLab + CEA/IRFU)
- Sampling frequency from 400 MHz to 3.2 GHz
- ADC dynamic range of 2.5 V coded on 12 bits (1 ADC tick = 0.61035 mV)
- Exists in 2, 8 or 16 ch desktop format, and 64 ch mini-crate
- Fast readout: potential resolution  $\leq$ 100ps



#### Light collection efficiency



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#### Reactor neutrino background



### scaling? much already demonstrated by NOvA...

#### common technology but not methodology

- •scintillator: ✓ (yield improvement)
- •fibres: ✔
- light collection system: √ (improvement?)
  photo-detector: √ (APD→SiPM OK?)
  different optimisation: R&D

### GeV OK!! But ~I MeV physics @ IOkton? (R&D)<sub>Anatael Cabrera (CNRS-IN2P3)</sub> — IJCLab / Université Paris-Saclay (Orsay)