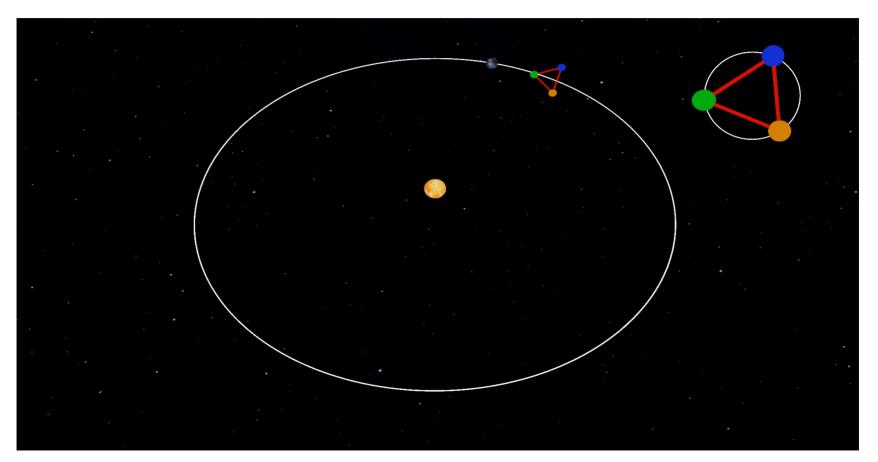
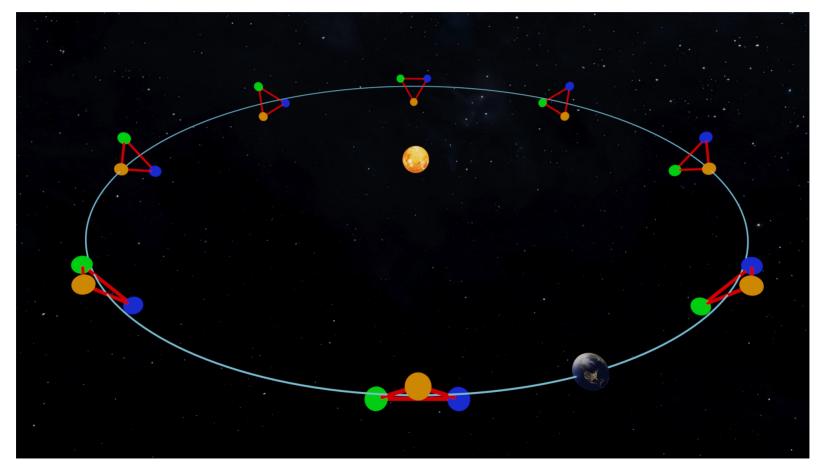
# Accelerating parameter estimation of Galactic binaries in the full LISA frequency band using Gaussian Process Regression



#### https://zenodo.org/record/6761175

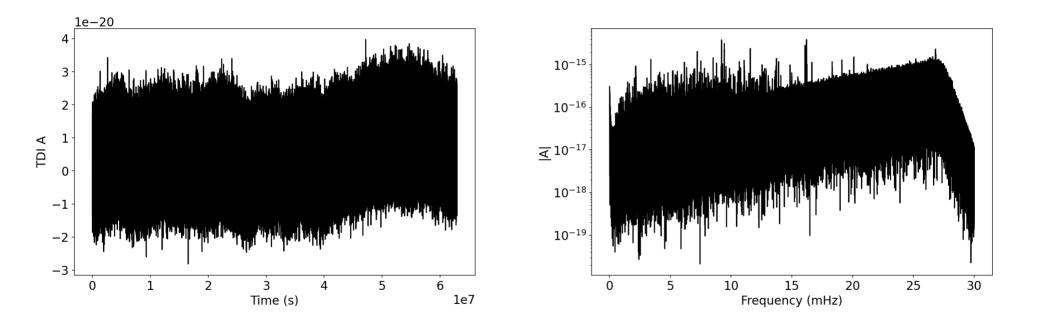
#### Stefan Strub, Luigi Ferraioli, Cedric Schmelzbach, Simon Stähler, Domenico Giardini

# Accelerating parameter estimation of Galactic binaries in the full LISA frequency band using Gaussian Process Regression

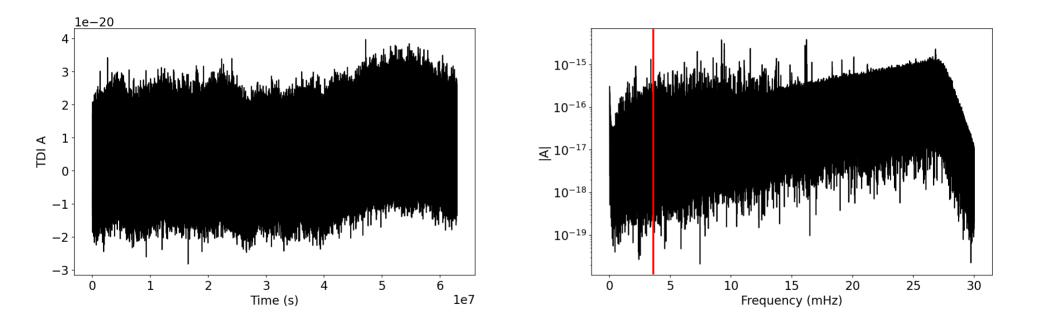


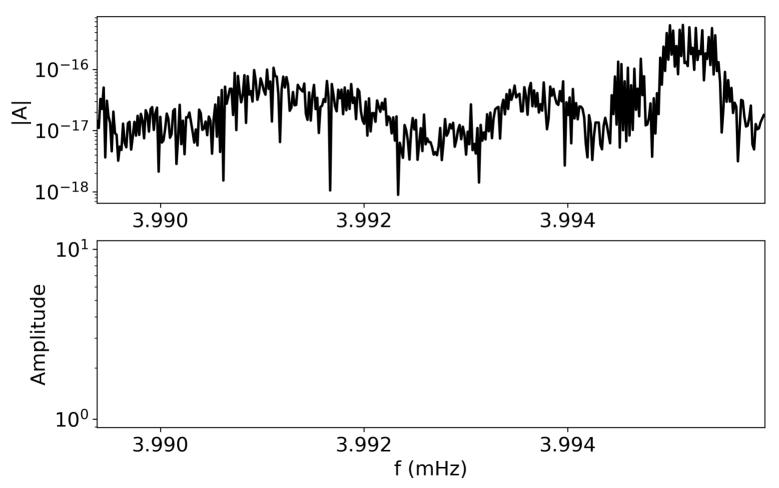
Stefan Strub, Luigi Ferraioli, Cedric Schmelzbach, Simon Stähler, Domenico Giardini

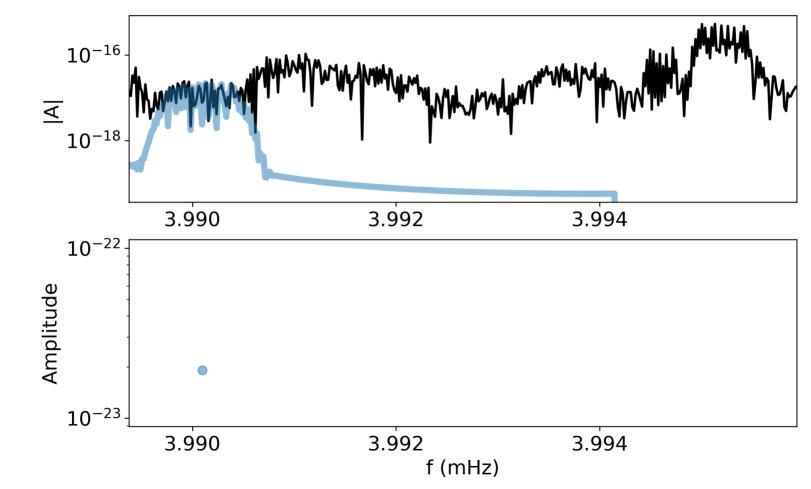
#### LDC1-4 (Radler)

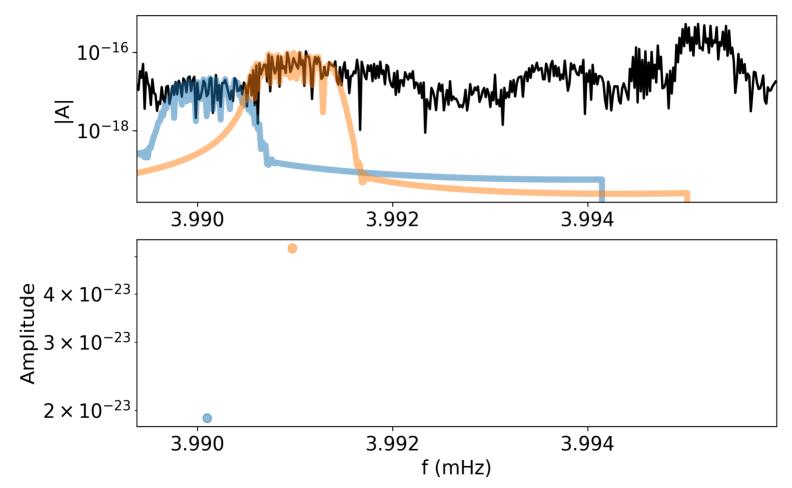


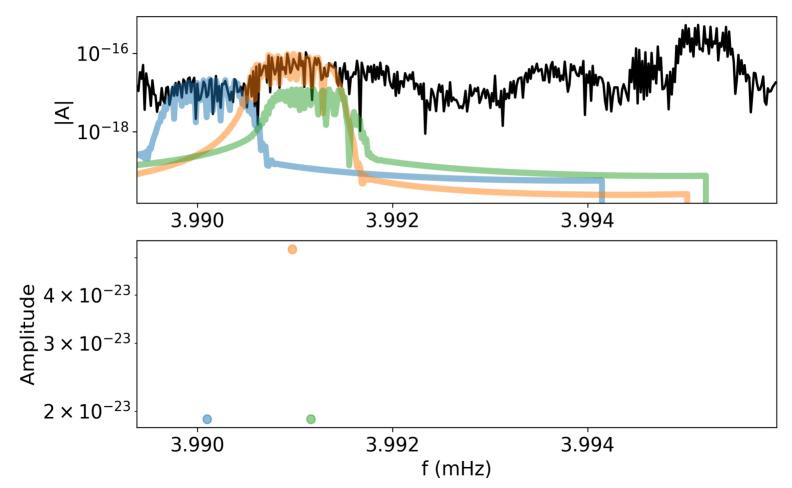
#### LDC1-4 (Radler)

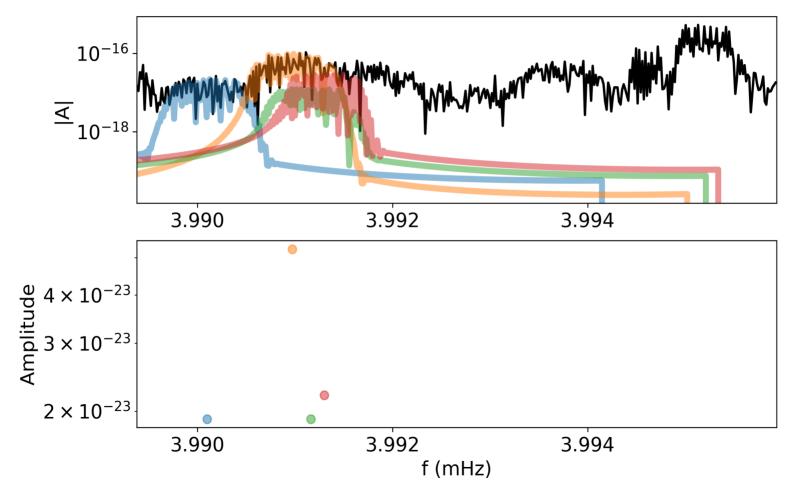


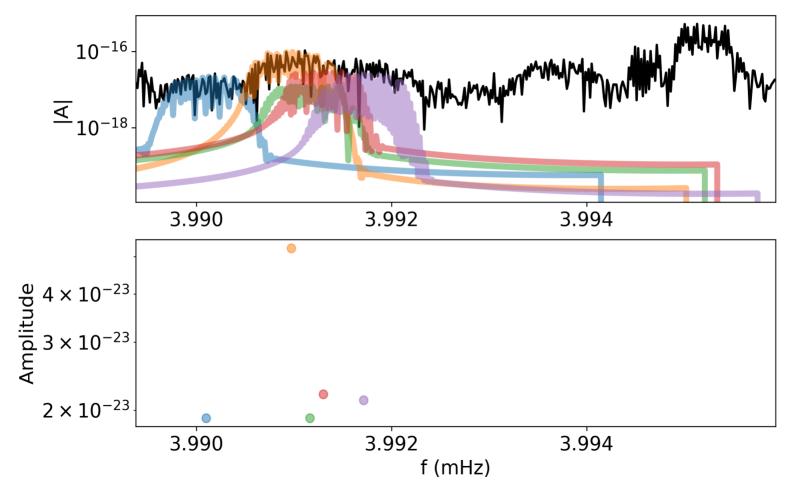


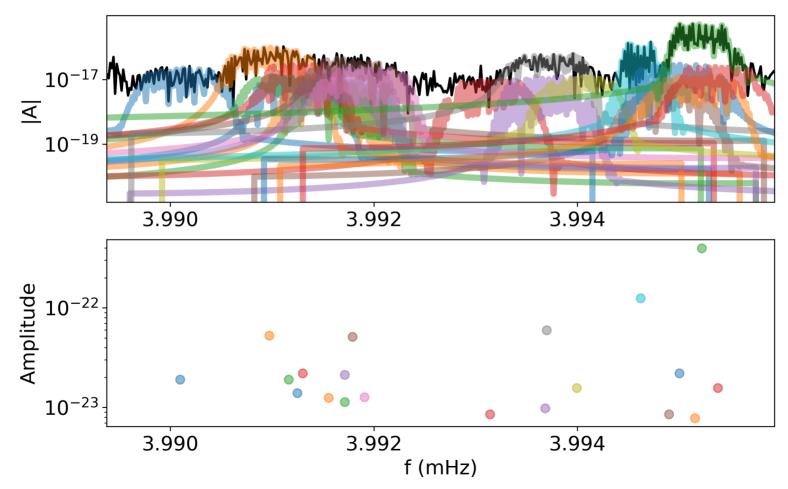












#### Fast best fit and slow posterior distribution: The best of both worlds

	Best fit	Bayesian posterior distribution
Possible method	Genetic Search Algorithm	Markov Chain Monte Carlo
Number of signal simulations	20'000 Fast	100'000 – 1'000'000 Slow
Uncertainty estimate	No	Yes

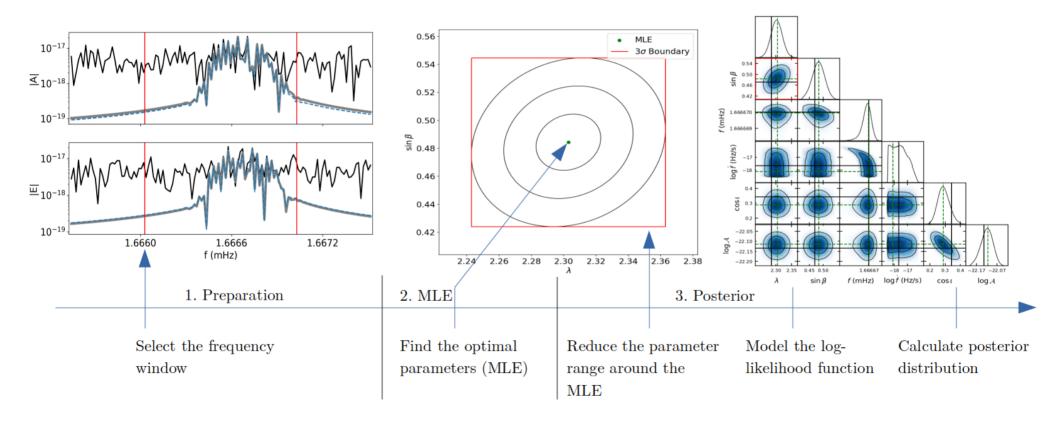
Log-likelihood:

$$\begin{split} \log p(d \mid \theta) &= -\frac{1}{2} \langle d - s\left(\theta\right) \left| d - s\left(\theta\right) \right\rangle \\ \theta &: \text{parameter set} \end{split}$$

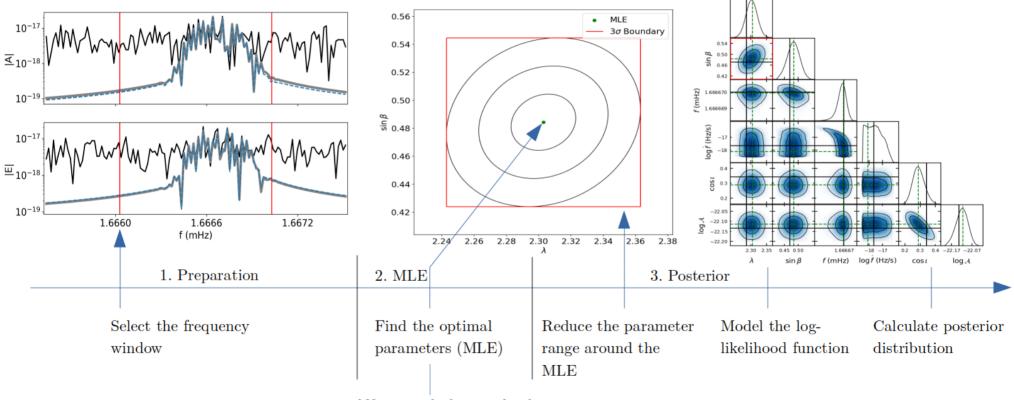
Scalar product:

$$\langle x(t) | y(t) \rangle = 4\mathcal{R}\left(\int_0^\infty \frac{\tilde{x}(f) \,\tilde{y}^*(f)}{S(f)} \, df\right)$$

#### The pipeline

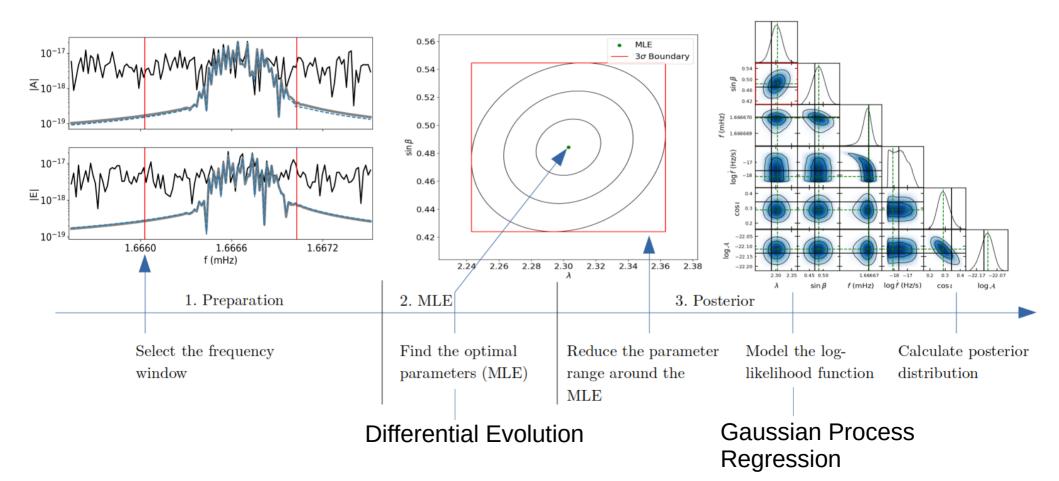


### The pipeline

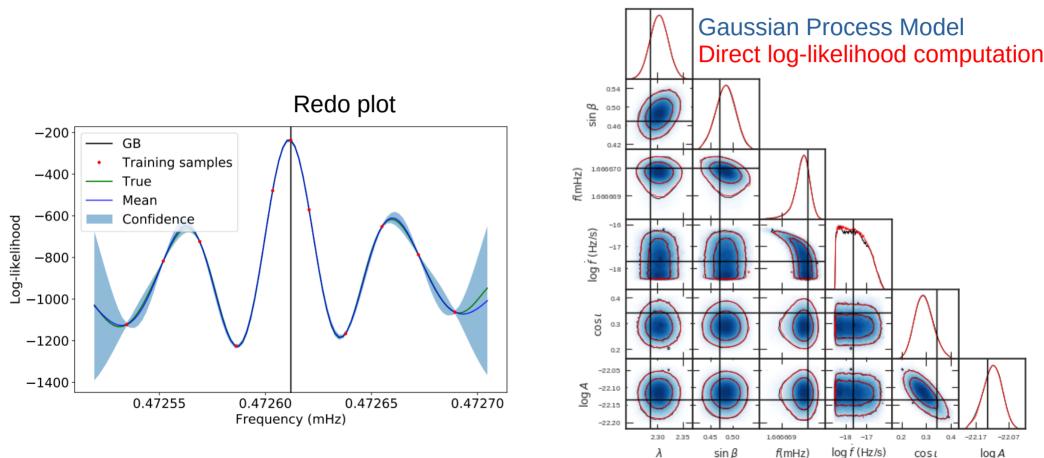


**Differential Evolution** 

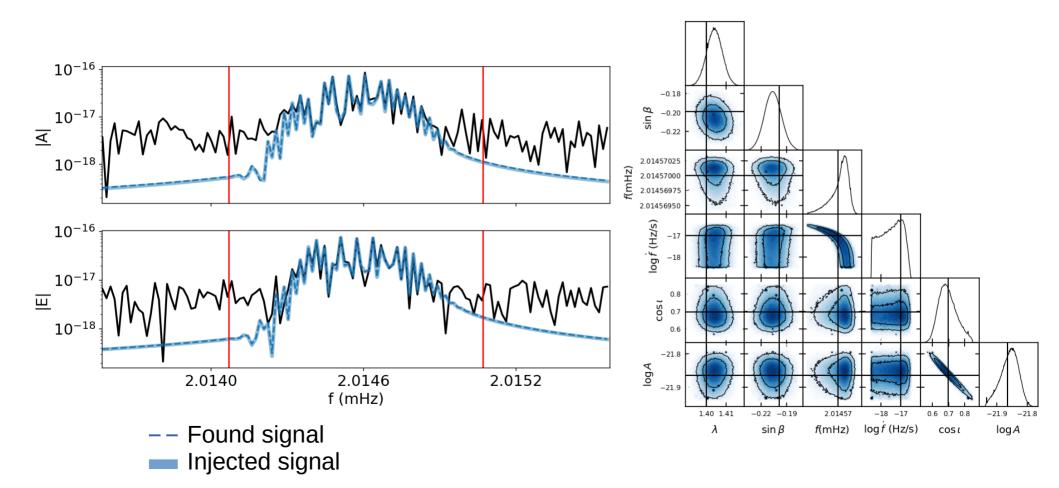
### The pipeline



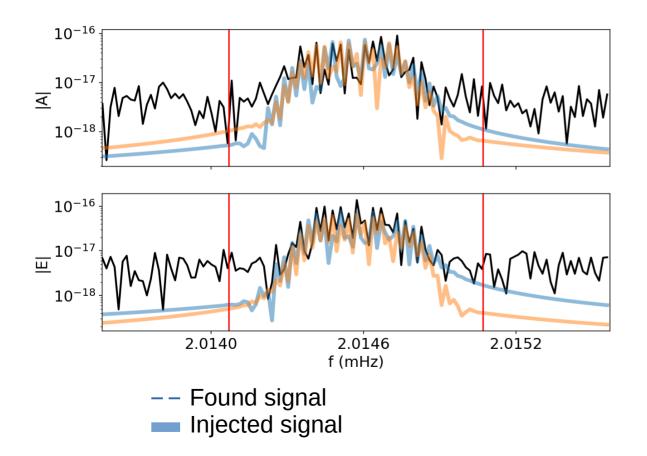
# Step 3: Gaussian Process Regression to model the likelihood for rapid MCMC sampling



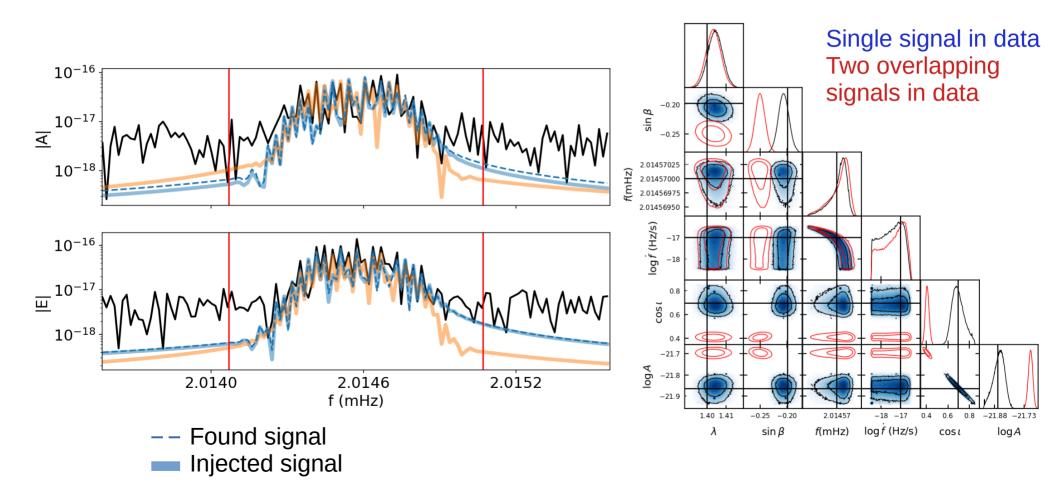
#### From single to overlapping signals



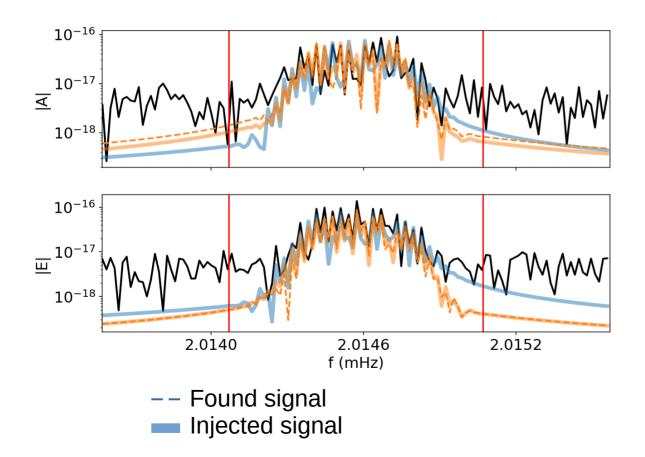
#### Two overlapping signals which only differ at sky location



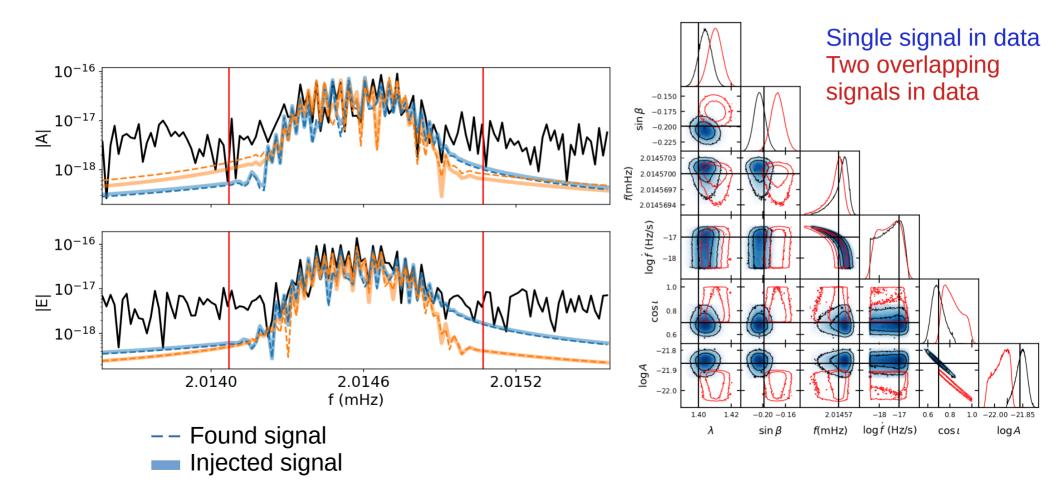
#### Finding only 1 signal has large bias



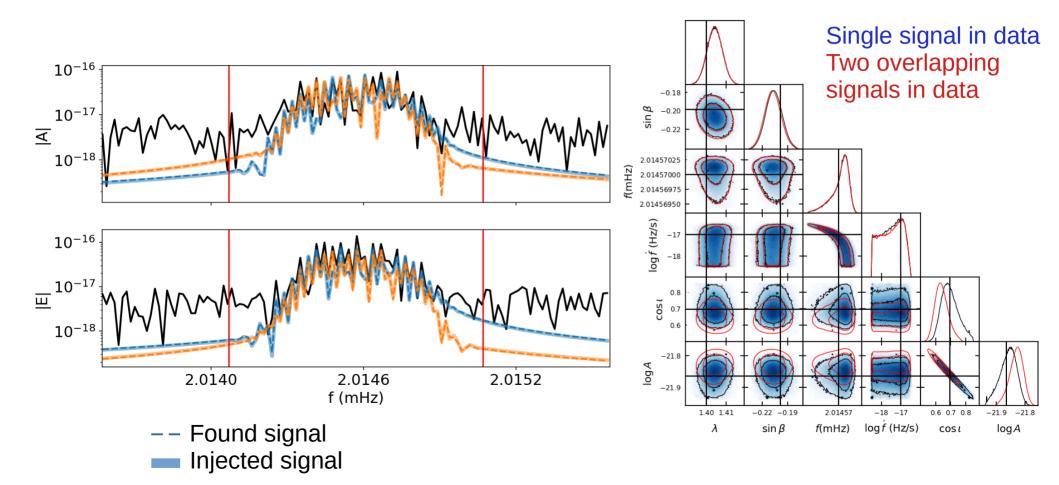
#### First find the orange signal and subtract

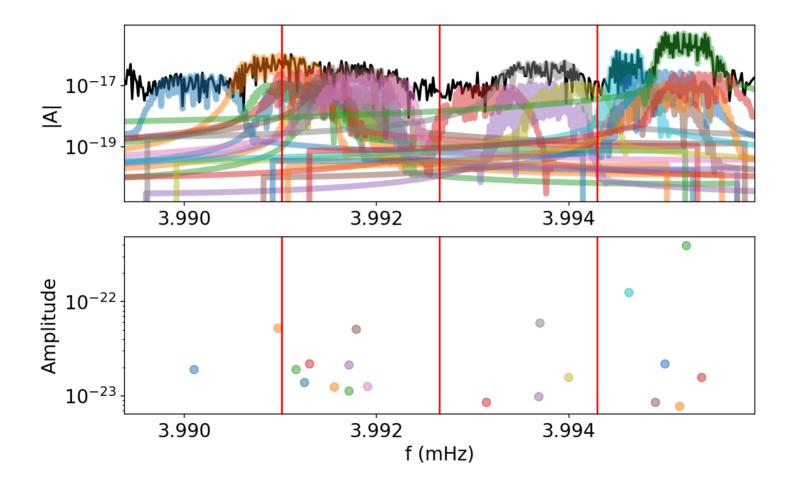


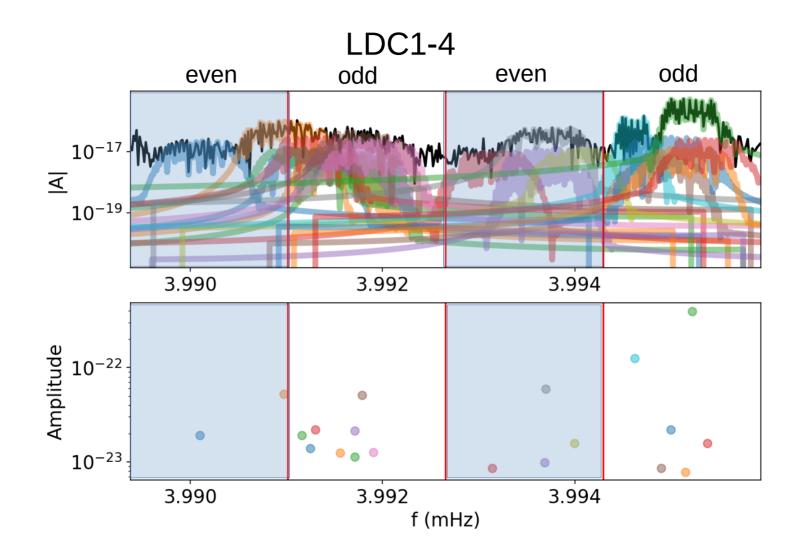
#### The bias due to the orange signal is reduced but still there



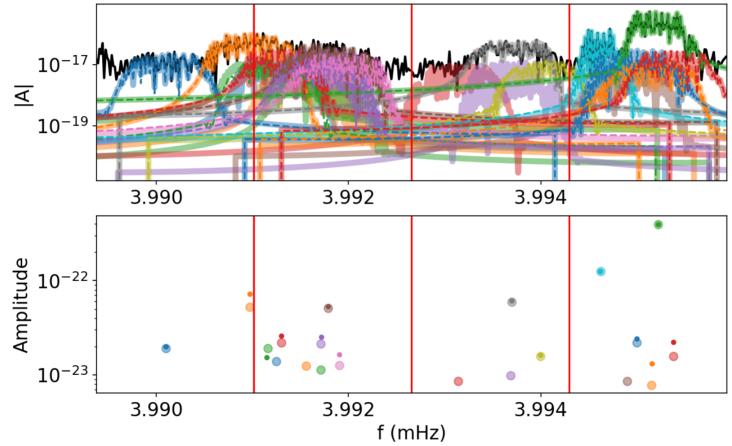
# Solution: Perform a global fit among both signals with initial guess from former biased solution

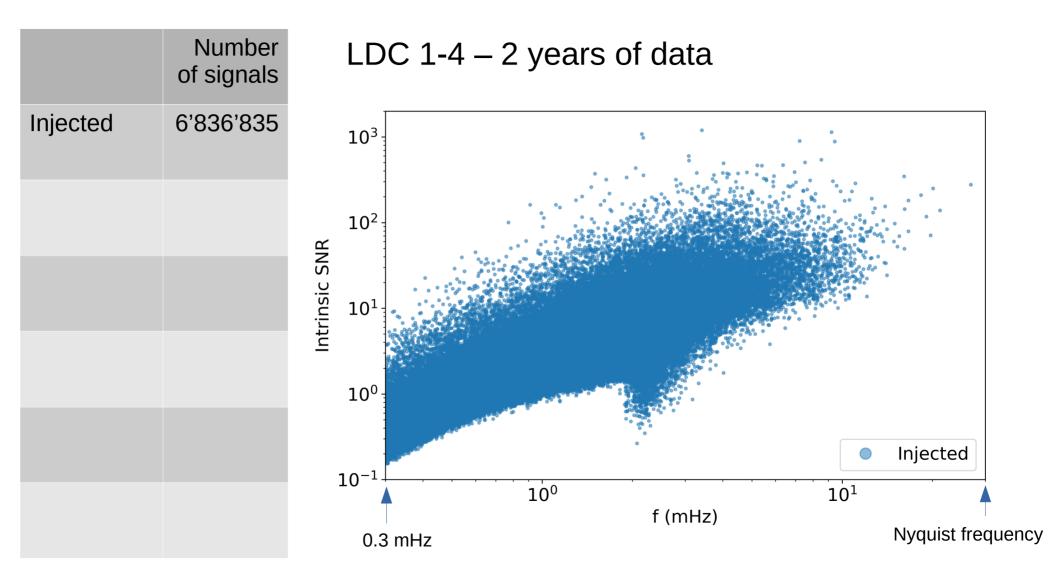


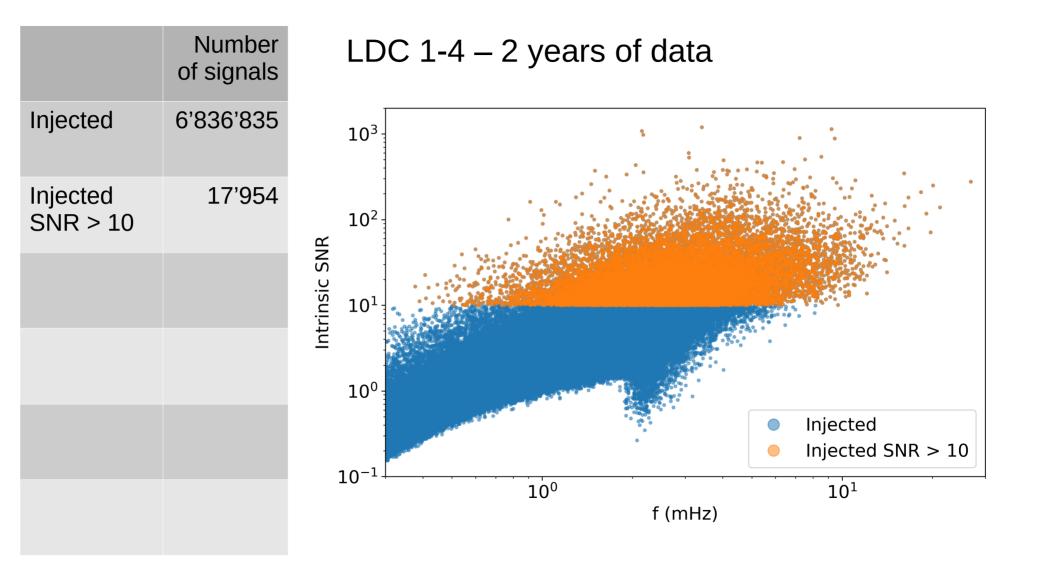


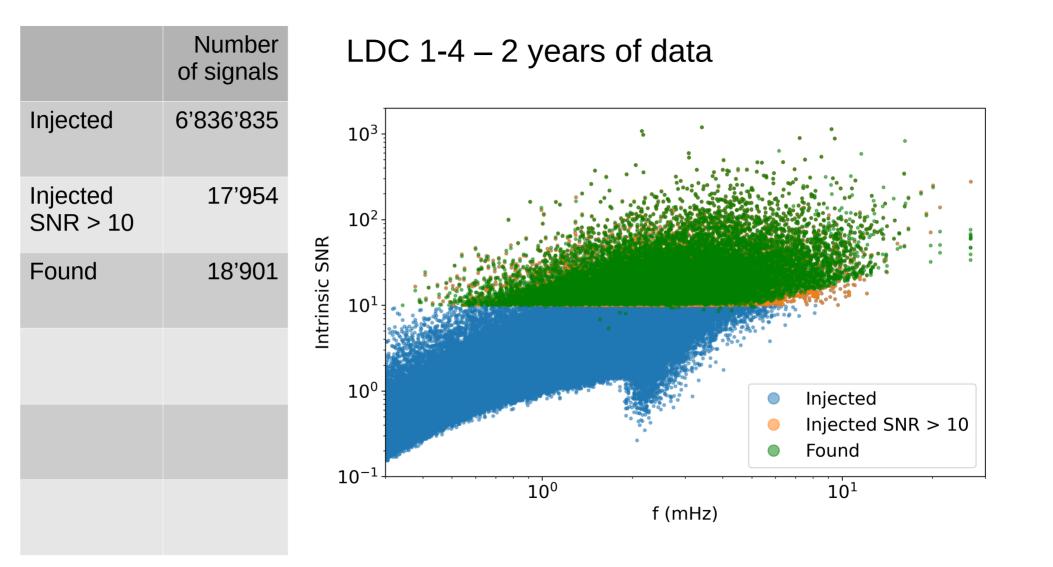


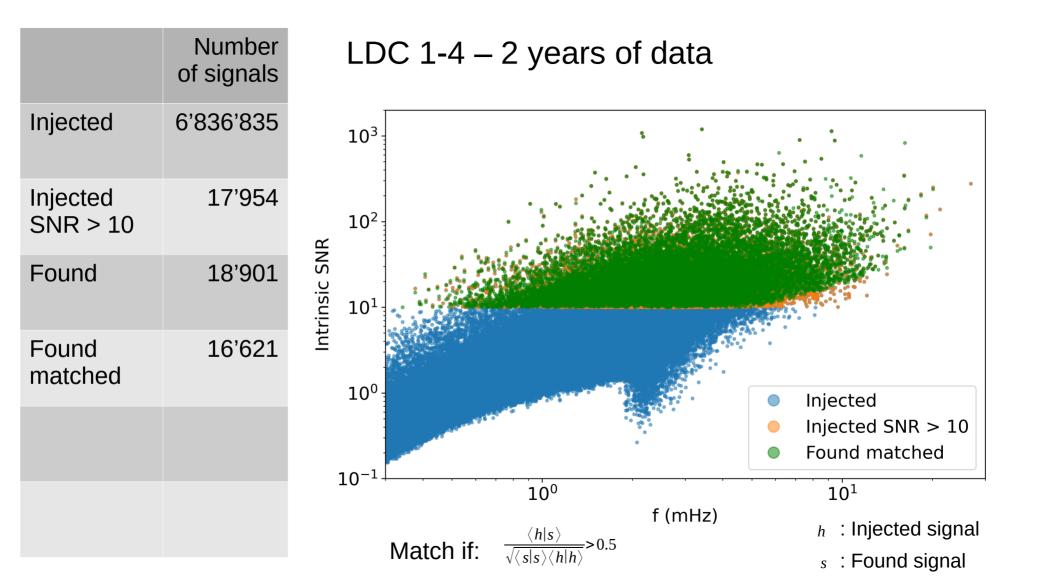
LDC1-4 14 out of 20 signals found

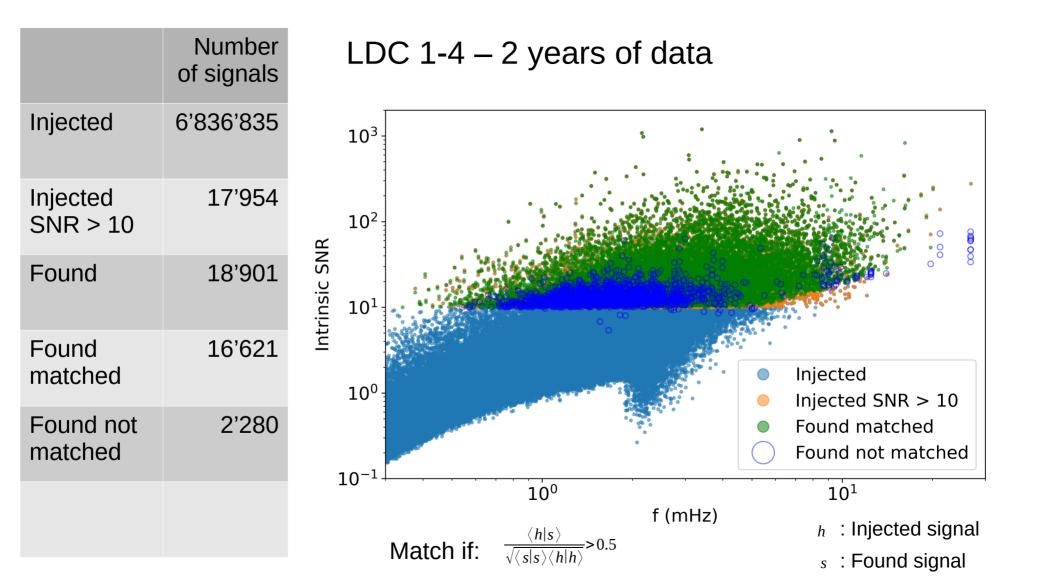


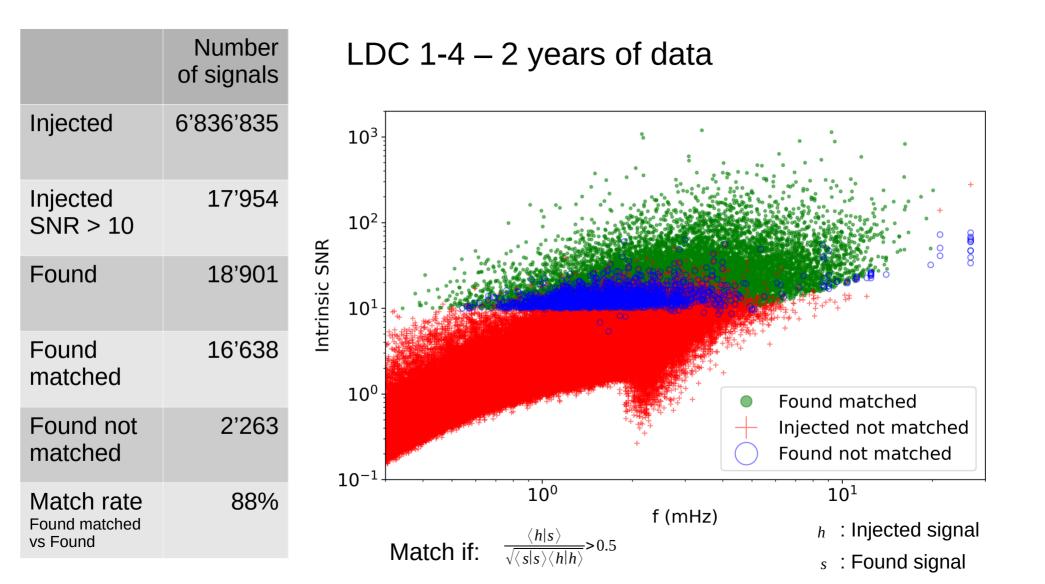


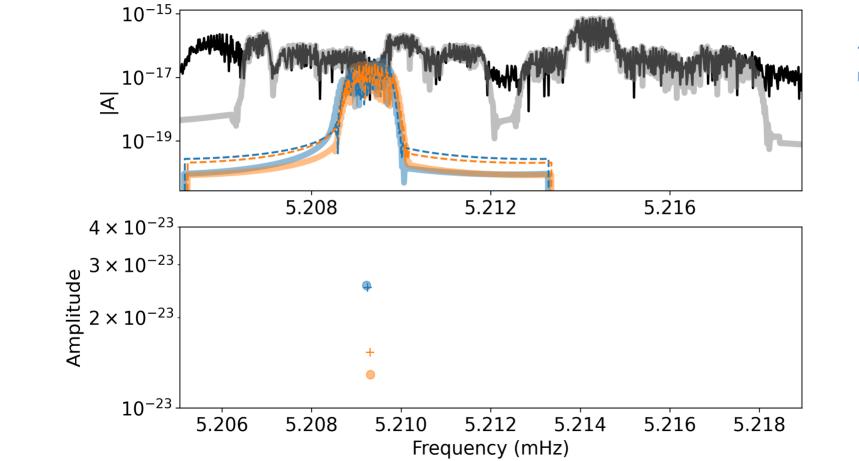










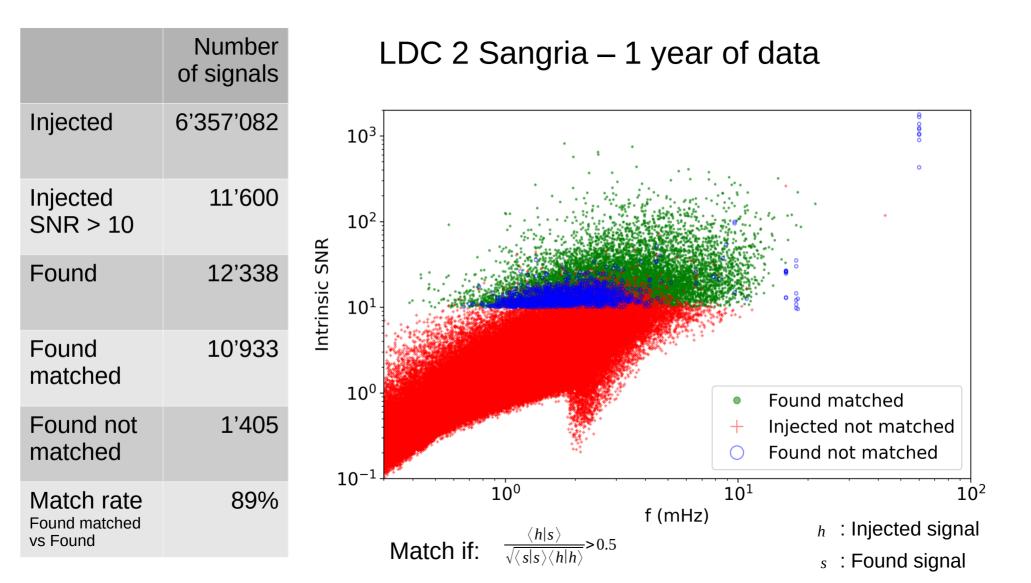


Found signal
Injected signal

#### LDC 1-4 – 2 years of data CPU time

Windows	Number of Windows	Max number of signals	CPU Time [h]	CPU Time Parallelized [min]
Even	3485	3	1.6K	27
Odd	3485	10	3.1K	53
Even	3485	10	2.5K	42
Total	10455		7.2K	122

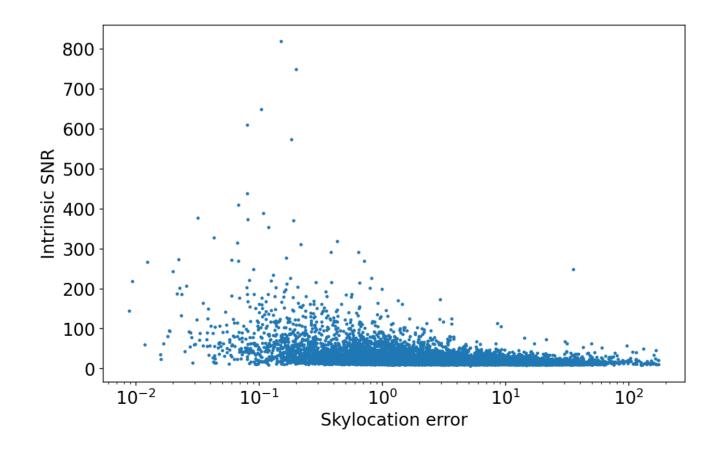
From Maudes' introduction slides: Time estimate was ~200K CPU-hours

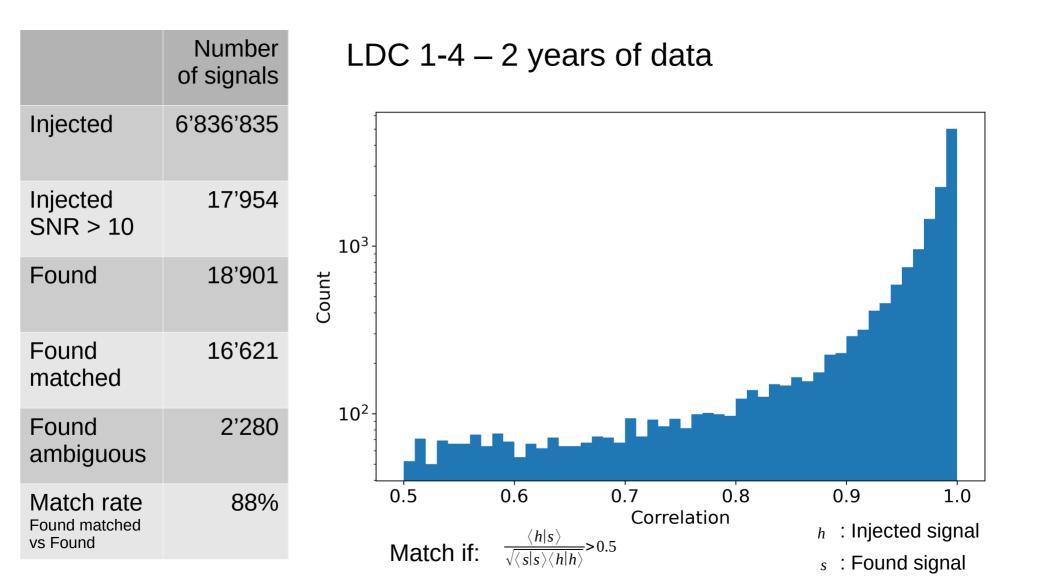


#### Conclusion

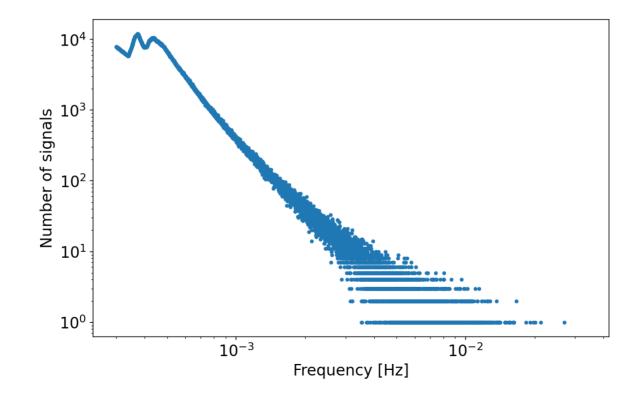
- Fast Bayesian parameter estimation of Galactic Binaries (Strub et al. 2022, 10.1103/PhysRevD.106.062003)
- Genetic Algorithms are efficient in finding the best fit solution
- Gaussian Process Regression accurately models the log likelihood function (so far only for unimodal posteriors, faint signals with SNR < 7 could be multimodal)</li>
- Robust pipeline to extract GBs from the full Galaxy
- Low computational cost: Only 7.2K CPU-hours for 2 years of data
- Low latency detection due to parallelization: ~2h for full galaxy with 3'500 CPU threads

#### Errors

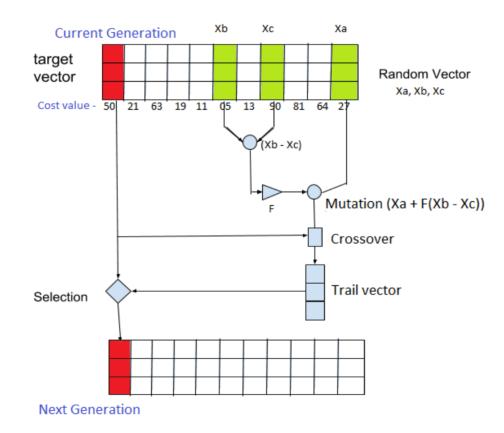




#### LDC 1-4 Number of signals per frequency window



#### **Differential Evolution**



Source: Medium, Abhishek Patel

