

Study of final-state interactions of protons in neutrino-nucleus scattering with INCL and NuWro cascade models

Anna Ershova

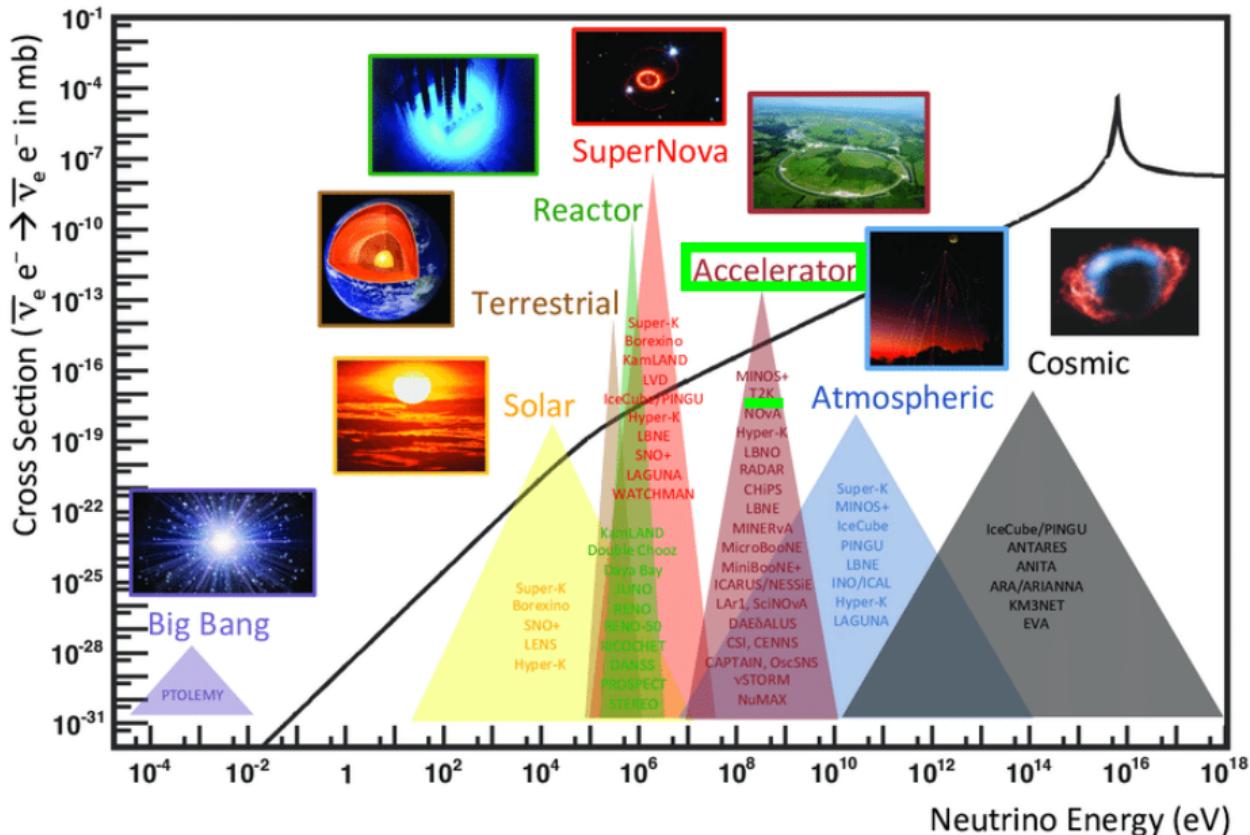
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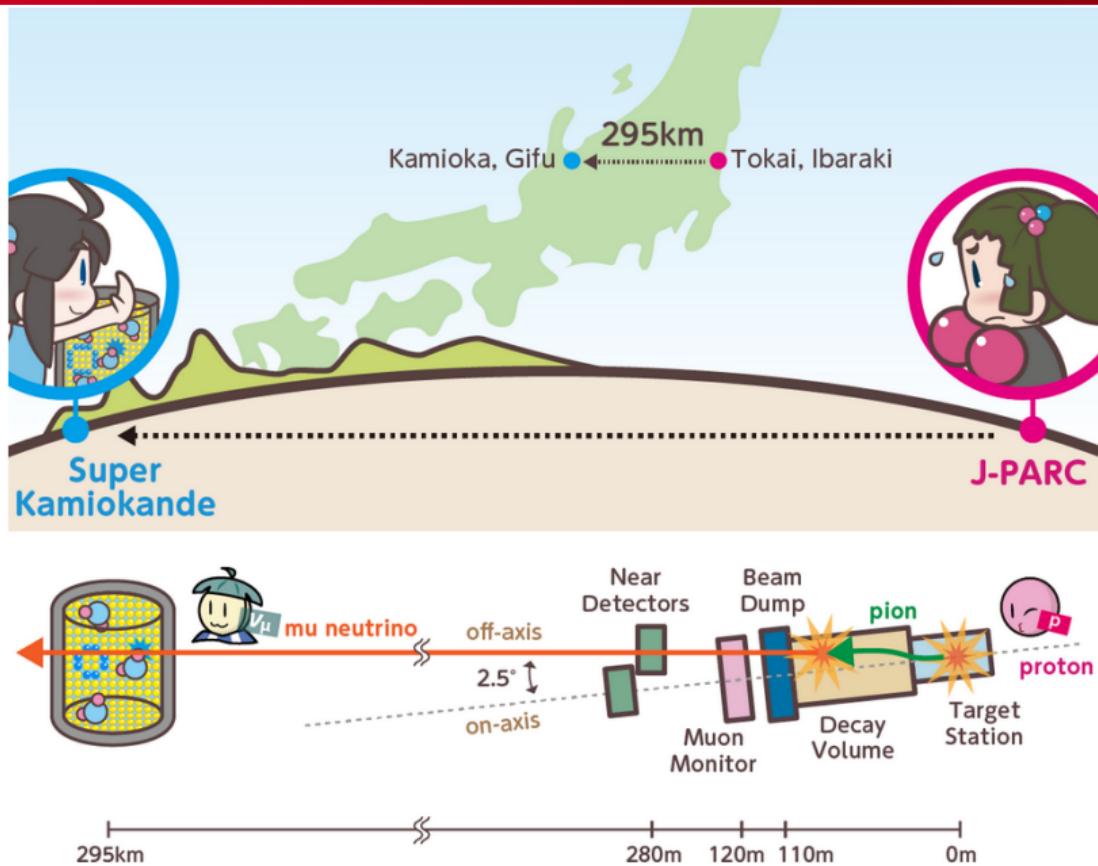
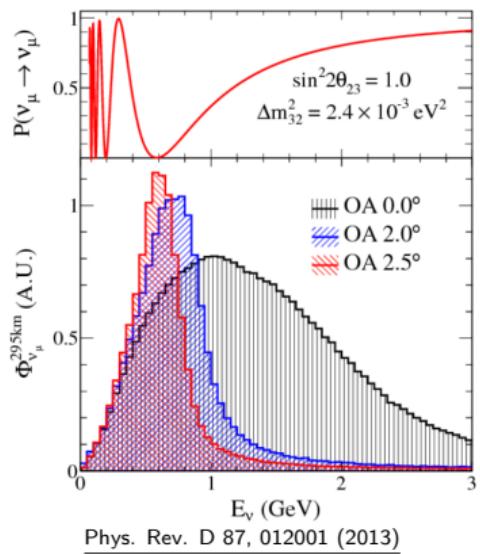
October 28, 2022



Neutrino studies



See also Vlada's presentation



$$\overbrace{R_{FD}^{\nu'}(E_\nu)}^{\nu \text{ rate}} = \underbrace{\Phi^\nu(E_\nu)}_{\nu \text{ flux}} \otimes \overbrace{P_{osc}^{\nu \rightarrow \nu'}}^{\nu \text{ oscillation probability}} \otimes \underbrace{\sigma^{\nu'}(E_\nu)}_{\nu \text{ cross-section}} \otimes \varepsilon^{\text{detector acceptance}}$$

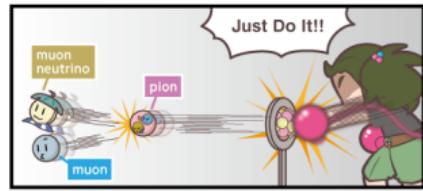
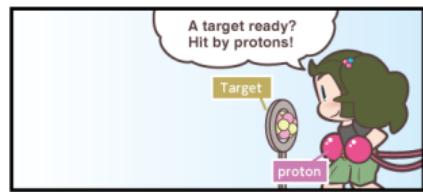
In order to get accurate neutrino rate, we need to **measure** neutrino energy **precisely**.



- **protons** hit the target and produce a **pion beam**
- **pions** decay and produce **neutrinos** and **muons**
- muons are stopped in the beam dump

Because of such construction of the neutrino beam, we **do not know** neutrino energy **precisely!**

How to make neutrinos

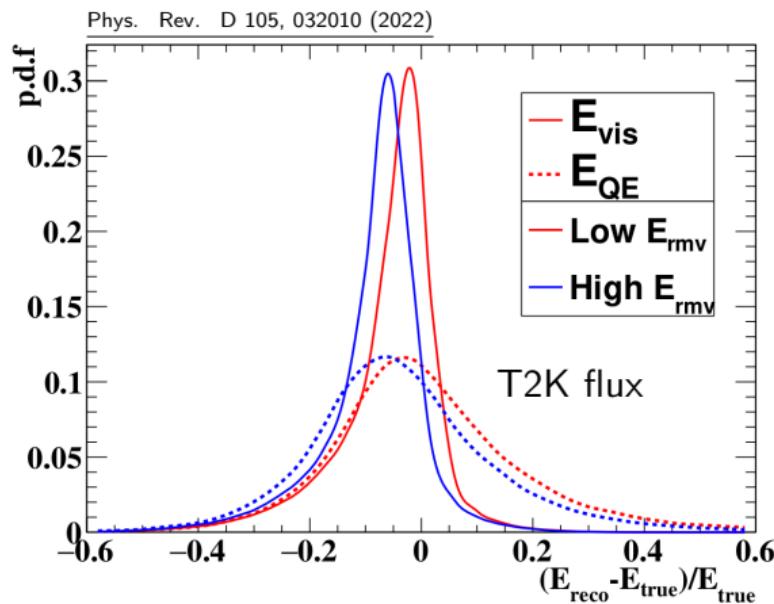


Energy reconstruction using only muon kinematics (works well for **quasi-elastic reaction**):

$$E_\nu^{QE} = \frac{m_p^2 - (m_n - E_B)^2 - m_\mu^2 + 2(m_n - E_B)E_\mu}{2((m_n - E_B) - E_\mu + p_\mu \cos\theta_\mu)}$$

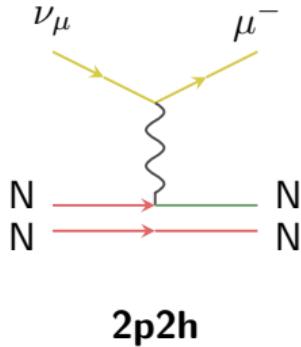
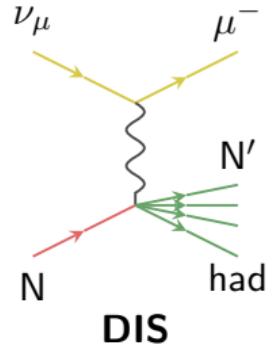
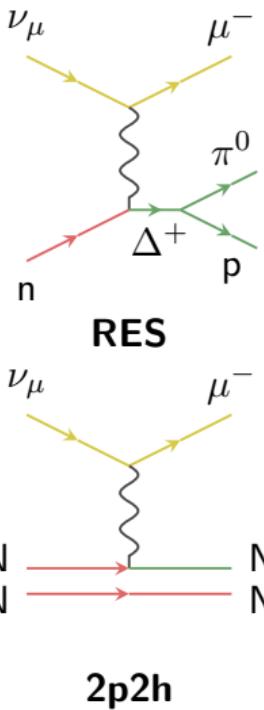
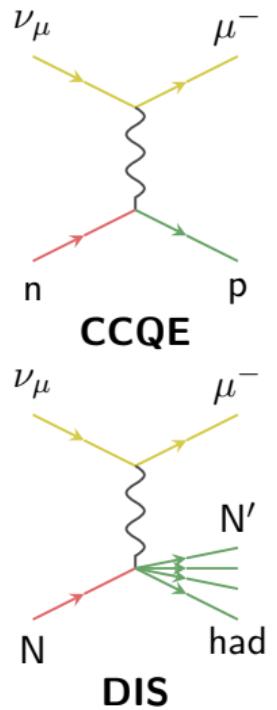
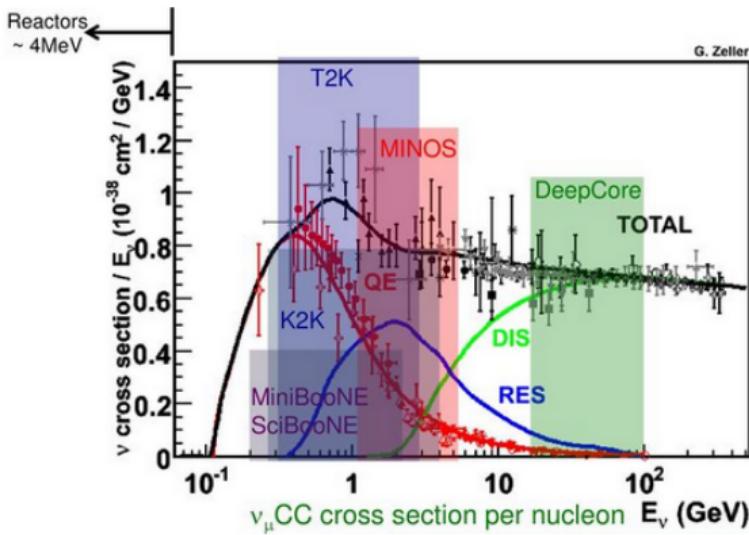
Energy reconstruction using **muon and kinetic energy of the nucleon**:

$$E_\nu^{vis} = E_\mu + T_N$$

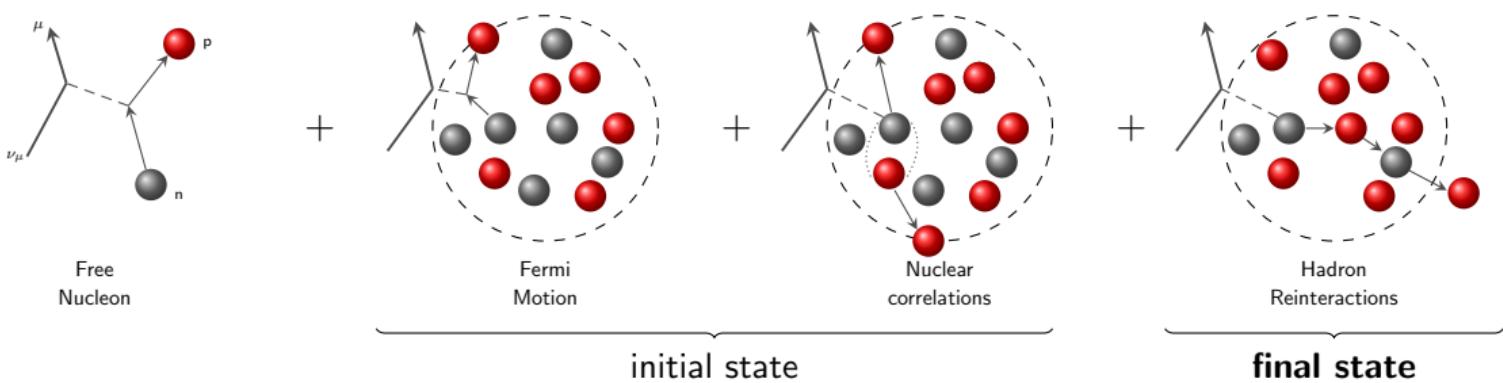


E_ν^{vis} , dashed line — QE formula
solid line — $\mu + N$ formula

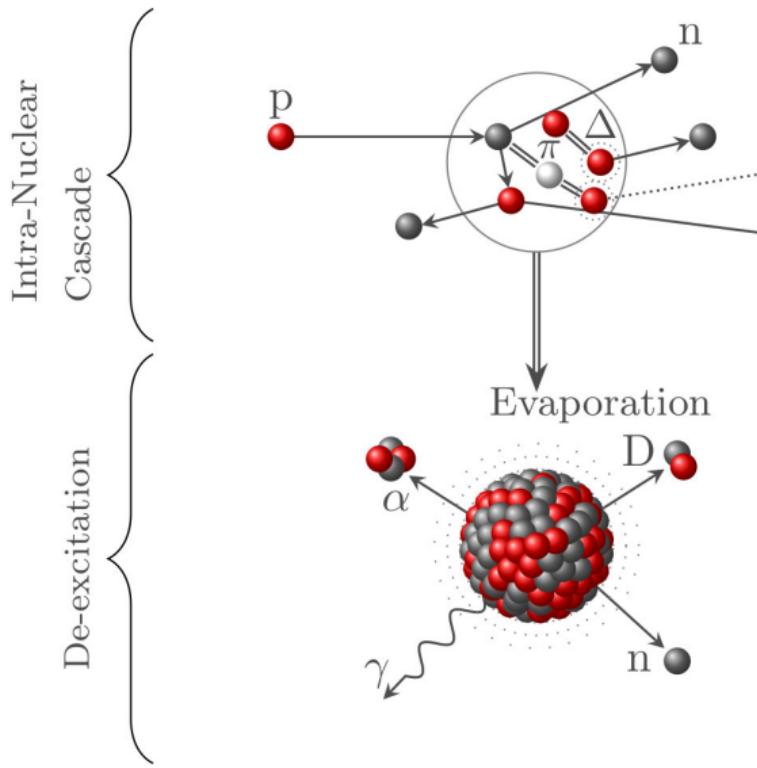
E_ν^{QE} works fine for the **CCQE** (Charged Current Quasi-Elastic) channel, where we have only μ^- and a proton in the final state. It is **less accurate** for other channels.



$\mu + N$ formula gives us more **opportunities**, but also it creates more **challenges** for modelling and we need to **understand better nuclear effects** also on neutrons and protons.



We will focus on **CCQE** ν reaction channel and the **Final State Interactions (FSI)** that are described by **cascade models**.



Projectiles: baryons (nucleons, Λ , Σ), mesons (pions and Kaons) or light nuclei ($A \leq 18$). **No neutrinos** yet! We use neutrino vertex from  **NuWro** (widely used ν -nucleus MC generator).

De-excitation: ABLA, SMM, GEMINI

Flexible tool: has been implemented in GEANT4 and GENIE

Phys. Rev. C 87, 014606 (2013)

Phys. Rev. C 90, 054602 (2014)

Phys. Rev. C 96, 054602 (2017)

Potential

Each nucleon in the nucleus has its **position and momentum** and moves **freely** in a square potential well. Nuclear model is essentially **classical**, with some additional ingredients to mimic quantum effects.

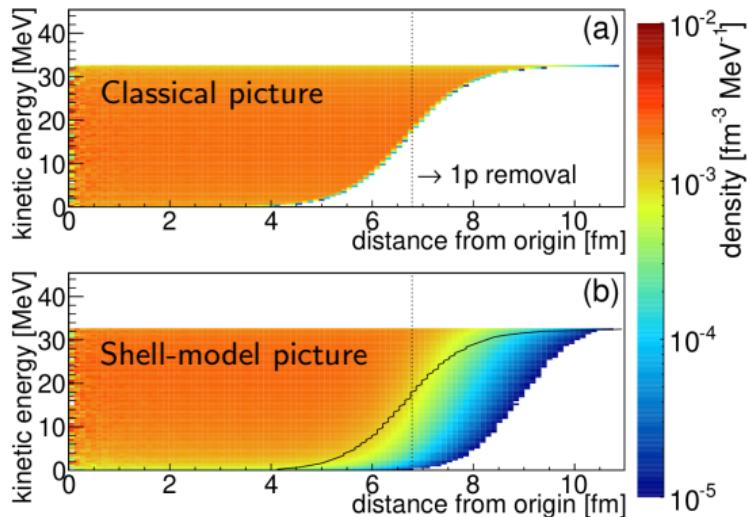
Pauli Blocking

- **strict:** blocked is $p < p_{Fermi}$
- **statistical:** count only nearby nucleons
- strict for the first event and statistical for the subsequent ones

Events inside cascade

- decay/collision
- reflection/transmission with probability to **leave the nucleus as a nuclear cluster**

Space-kinetic-energy density of protons in ^{208}Pb



Phys. Rev. C 91, 034602 (2021)

Potential

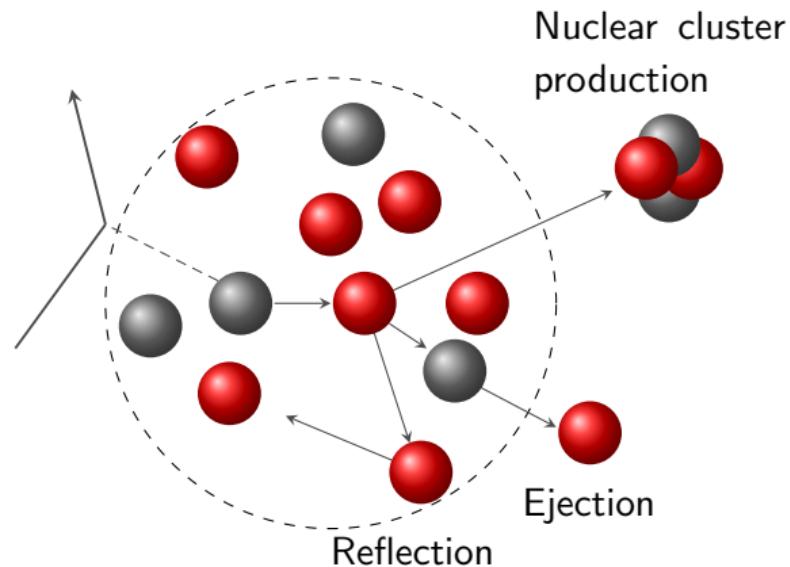
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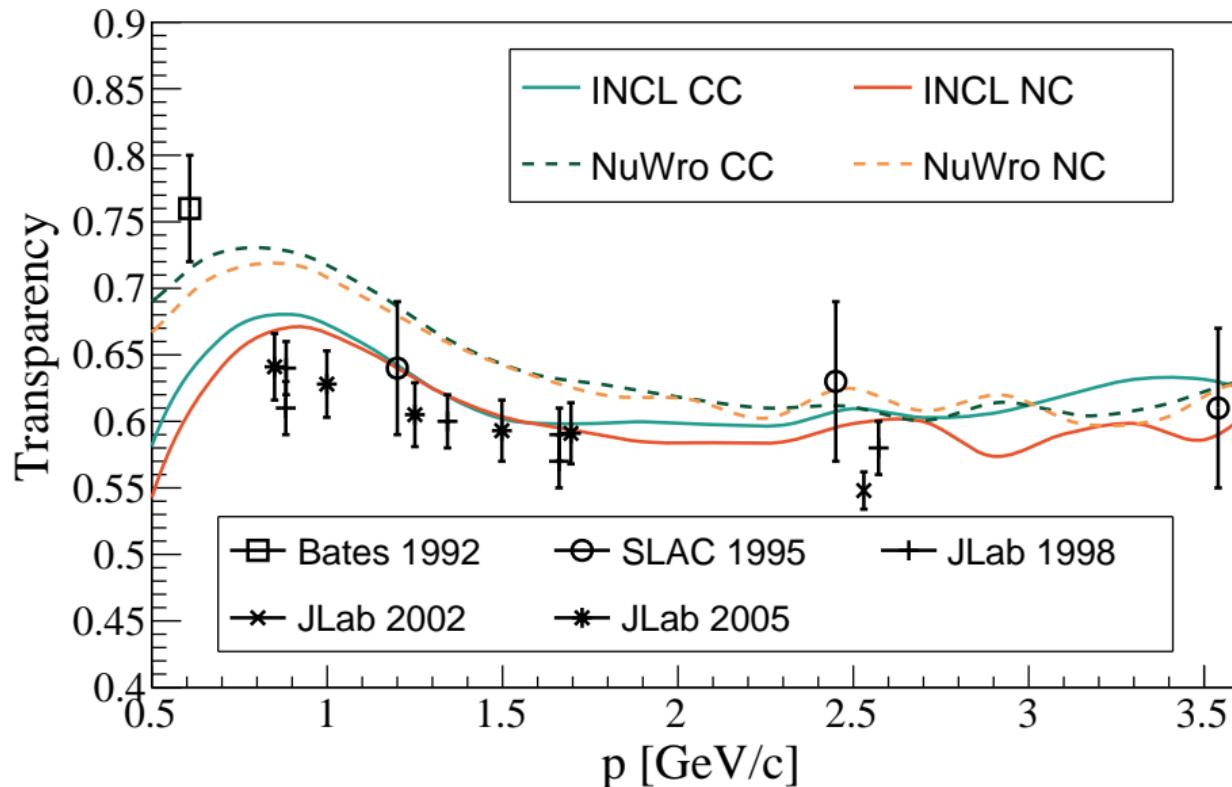
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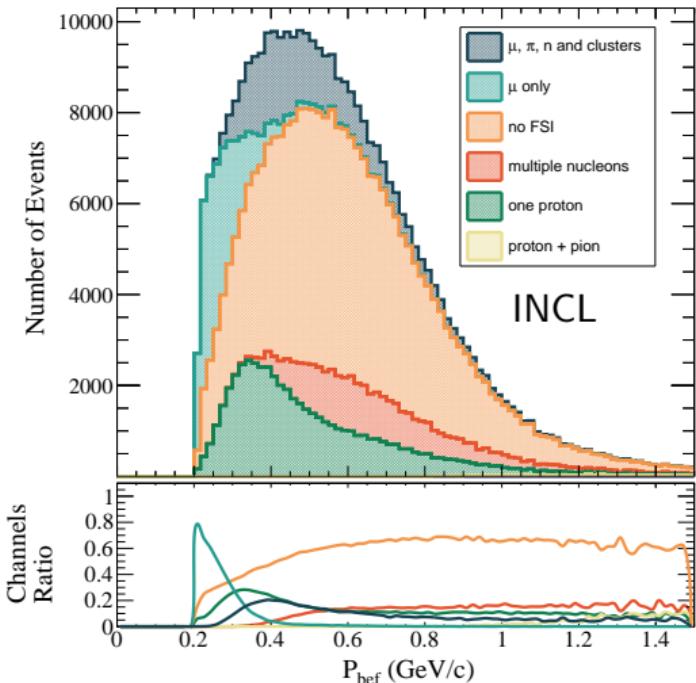
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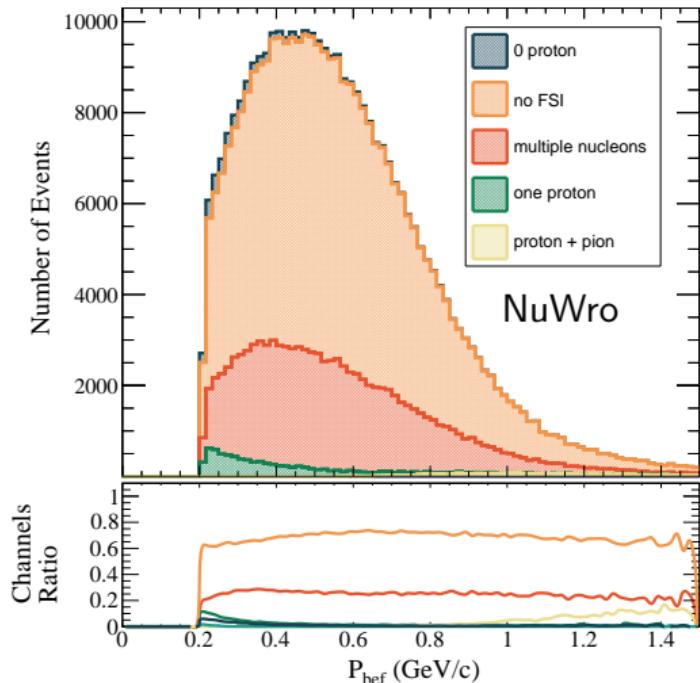




Proton momentum before FSI



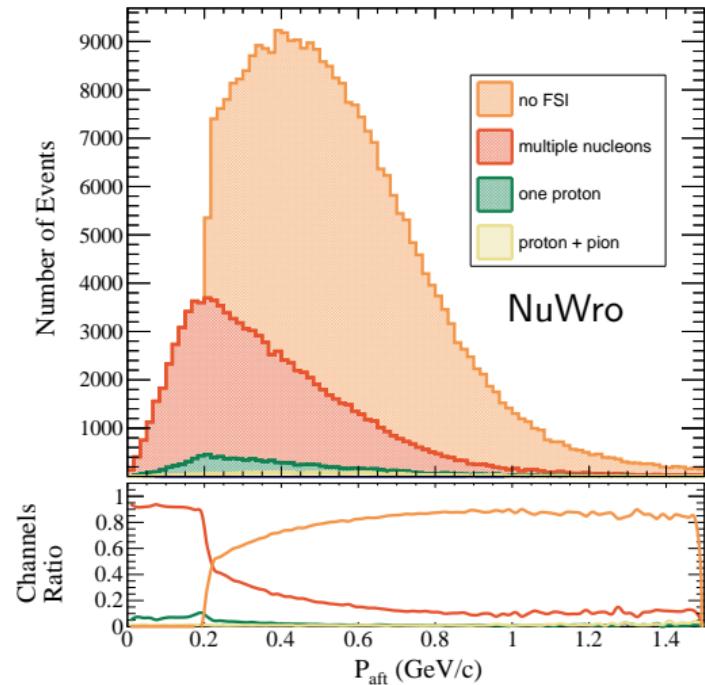
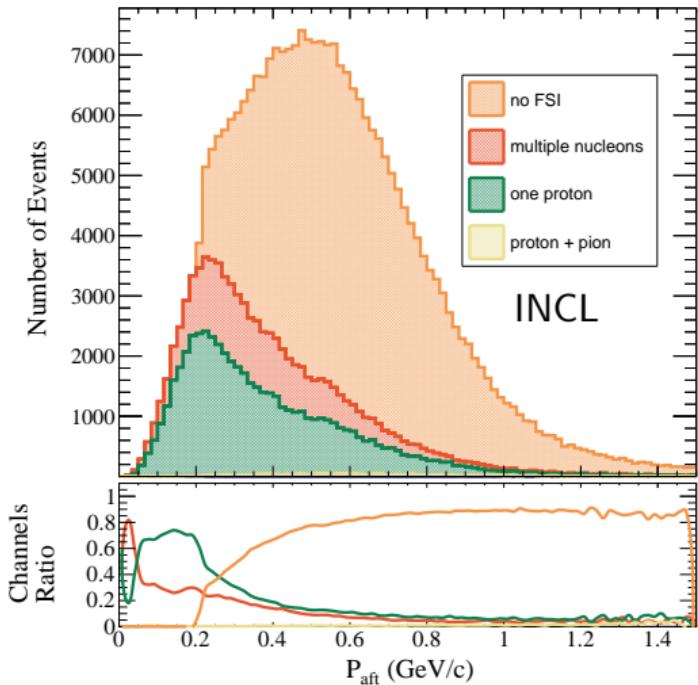
INCL



NuWro

INCL FSI simulation features a significant fraction of events **without a proton** in the final state, especially low momentum protons region.

Proton momentum after FSI



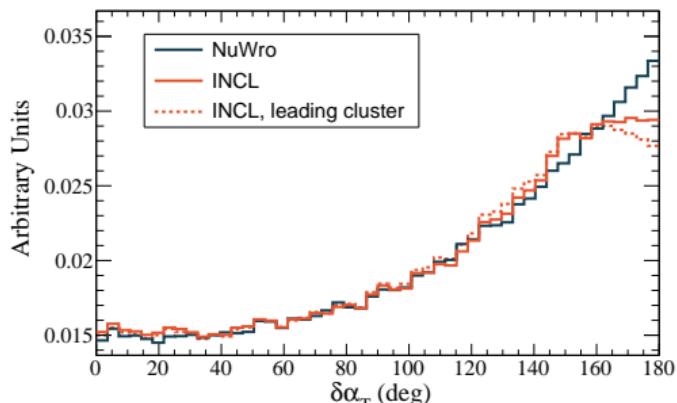
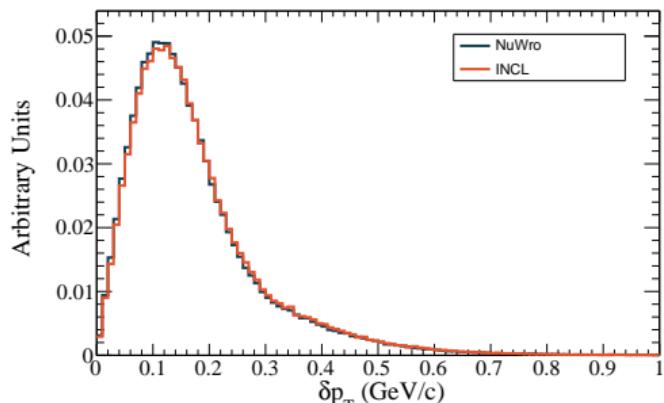
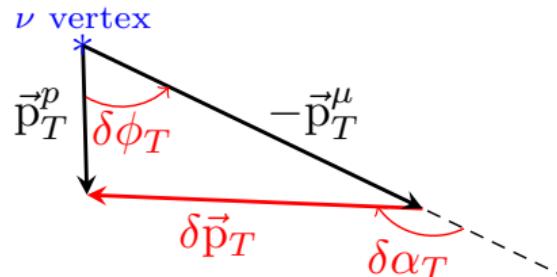
The FSI part of the distribution is **closer to zero** for NuWro → NuWro Pauli Blocking is **less strict** than INCL.

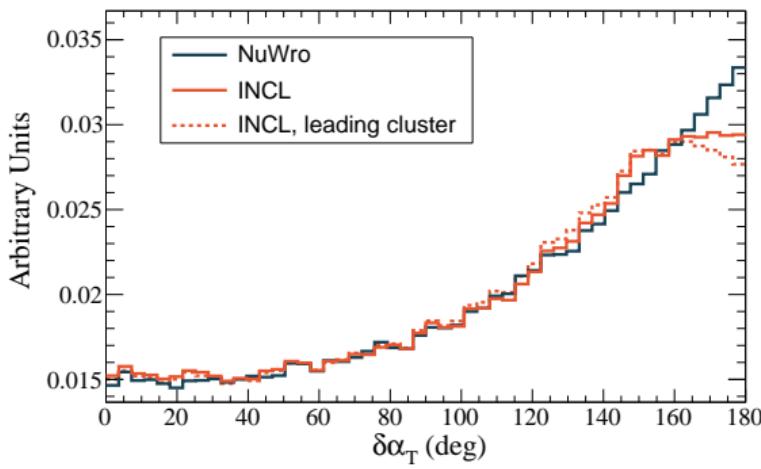
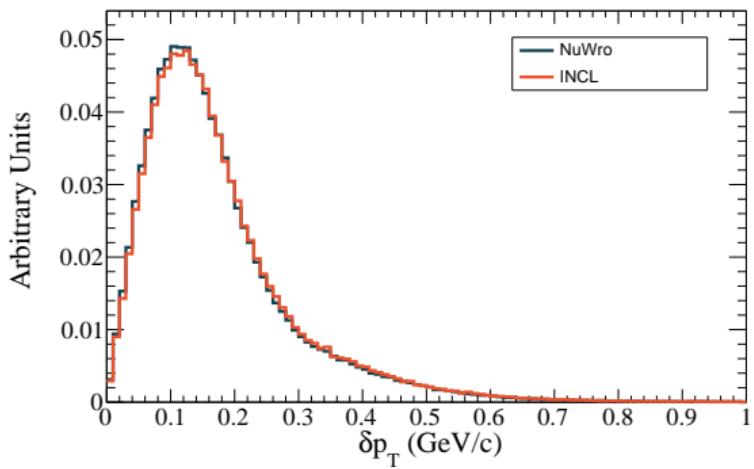
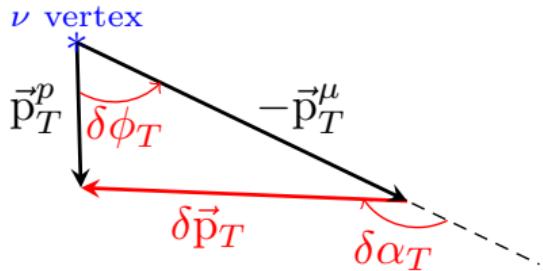
Variables of interest

We use **Single Transverse Variables (STV)** that allow to disentangle different effects for better FSI estimation. STV are **observable** and **measurable**.

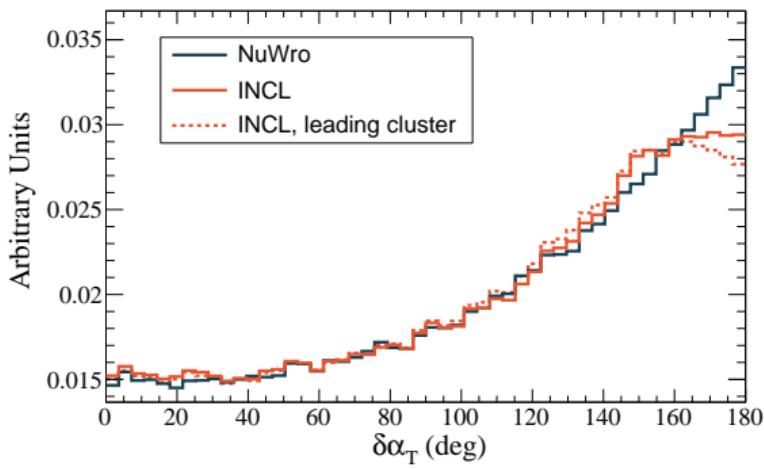
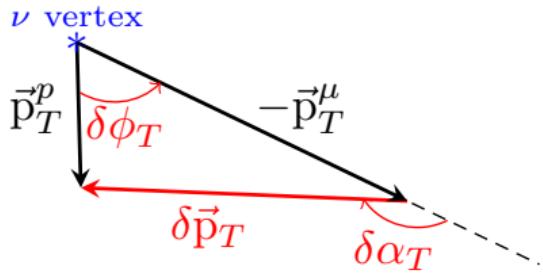
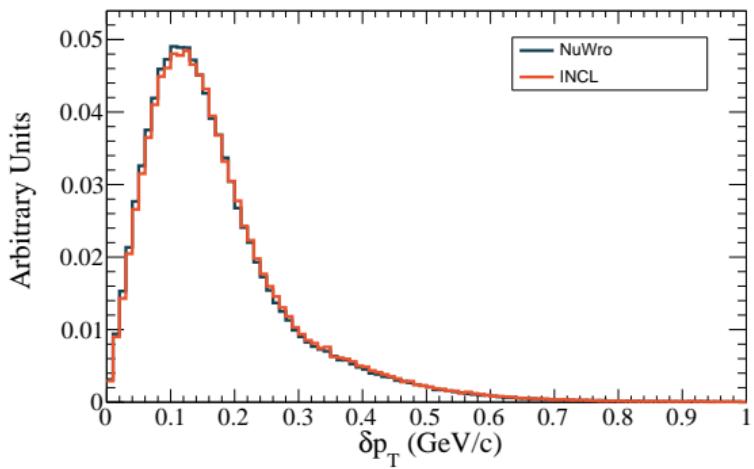
sensitive to FSI: $\delta\alpha_T = \arccos \frac{-\vec{k}'_T \cdot \delta\vec{p}'_T}{\vec{k}'_T \cdot \vec{p}'_T}$

sensitive to Fermi Motion: $\delta\vec{p}_T = \vec{p}_T^{\vec{p}} + \vec{p}_T^{\vec{\mu}} = \vec{p}_T^{\vec{n}}$

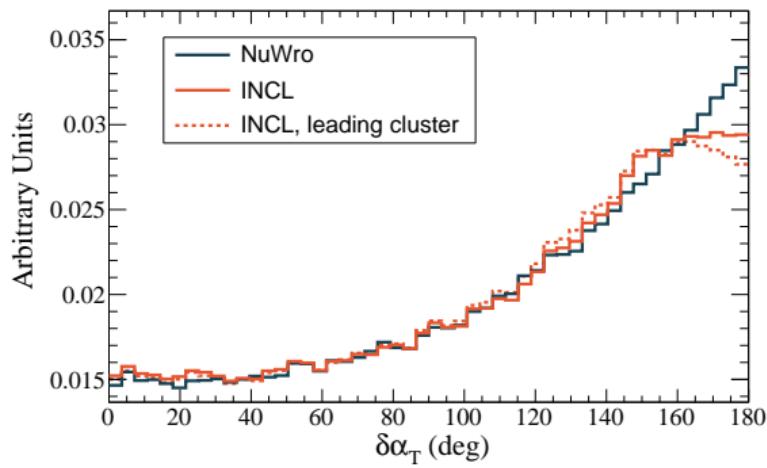
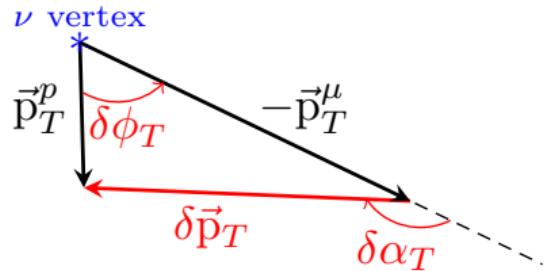
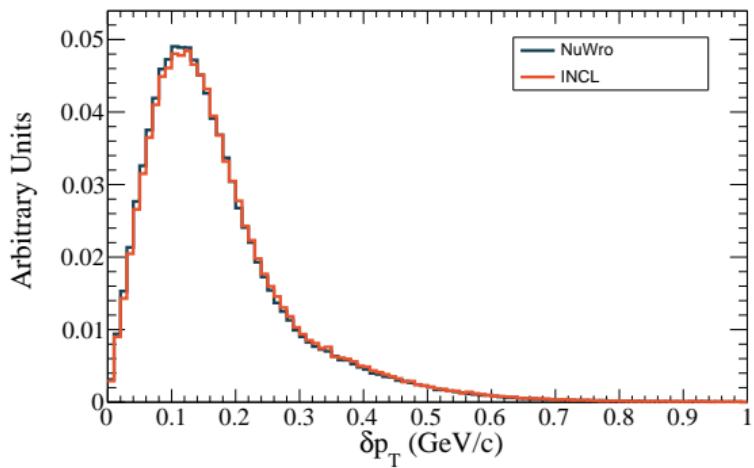




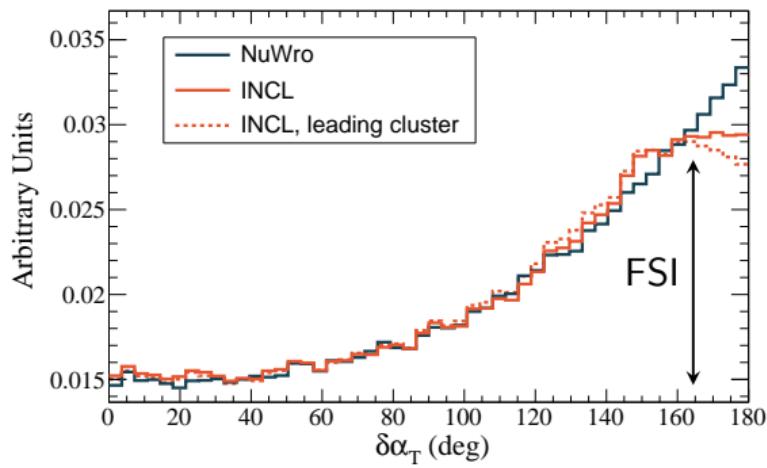
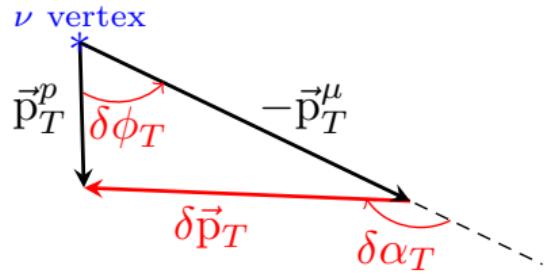
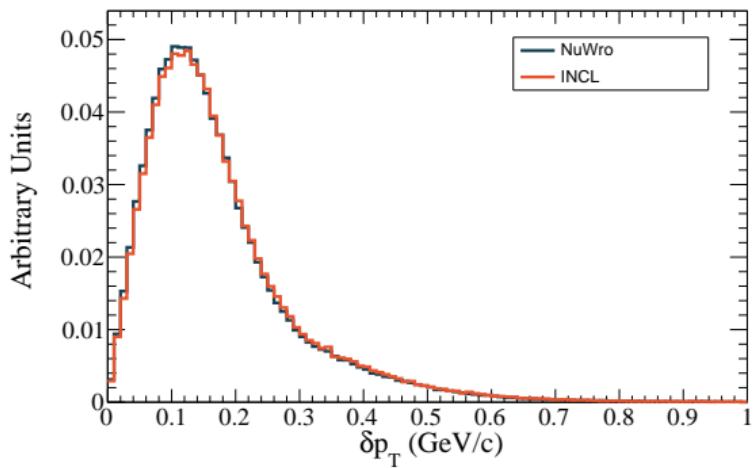
Fermi Motion

Fermi Motion FSI + SRC



Fermi Motion FSI + SRC



Current detector **threshold is too large**, so we **cannot see the difference** between INCL and NuWro.

Cuts (MeV):

$$p_\mu > 250$$

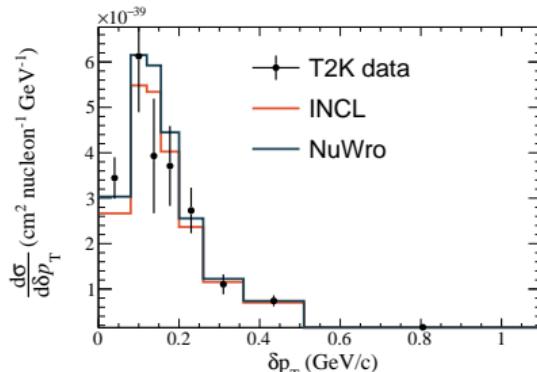
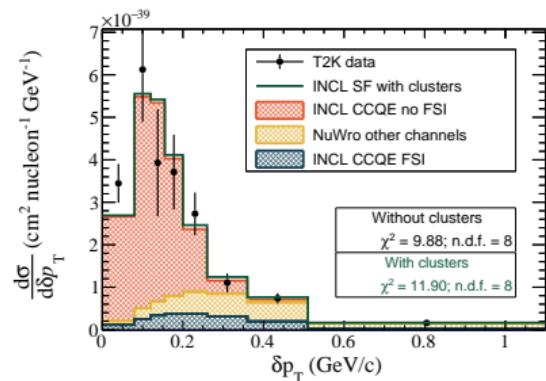
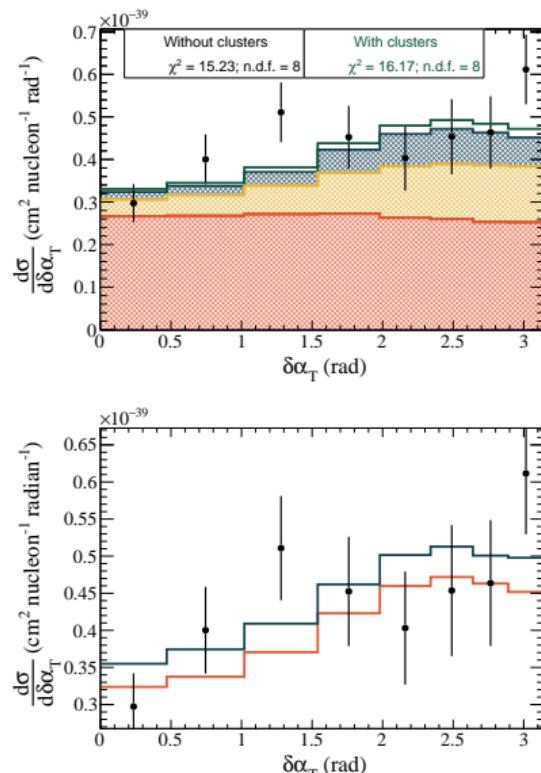
$$450 < p_p < 1000$$

$$\cos(\Theta_\mu) > -0.6$$

$$\cos(\Theta_p) > 0.4$$

Comparison to MINER ν A data
in backup

T2K data taken from
Phys. Rev. D, 98 032003 (2018)



Geant4 simulation of the **CH scintillator**.

How often do nuclear clusters travel enough to be reconstructed as a track?

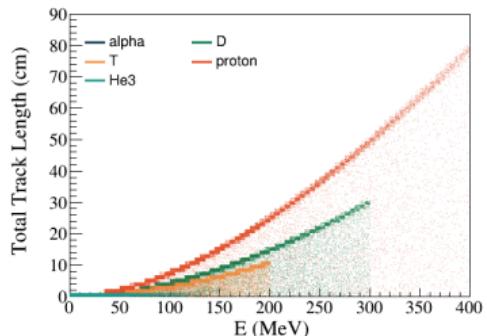
	α	^3He	T	D	proton
Travels > 1 cm, %	0.3	1.3	60	72	87
Travels > 3 cm, %	0	0	34	51	74

Can we **identify** nuclear clusters?

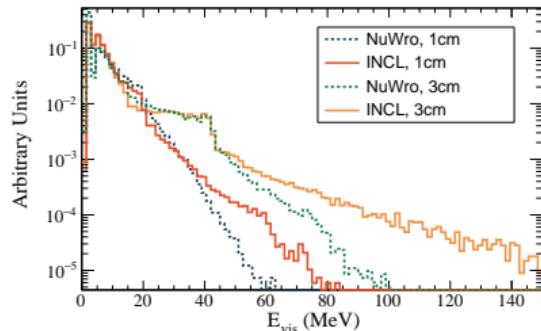
- **Deuteron** can be misidentified ($\sim 20\%$ of events) as **protons**
- **Tritium** can be misidentified ($\sim 10\%$) equally as **proton** or **deuteron**
- Main source of misidentification: **inelastic events**

Nuclear clusters contribute to the **vertex activity** that needs to be accounted for to avoid the ν **energy reconstruction bias**.

Total track length vs. kinetic energy



Vertex activity per event

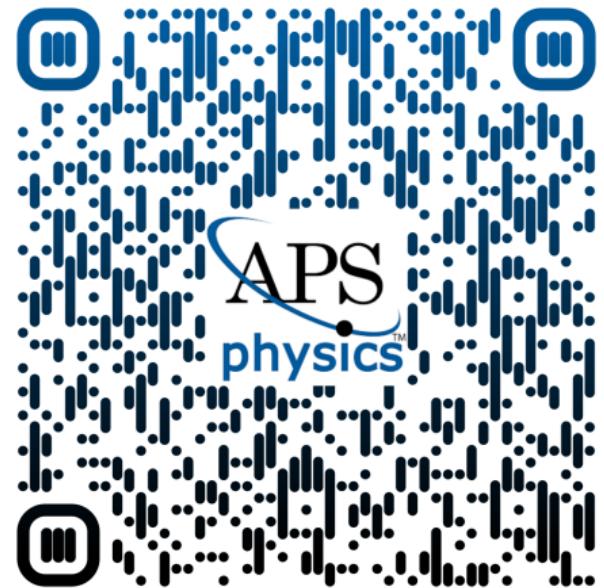


- We have compared the simulation of the final-state interactions between the **NuWro** and **INCL** cascade models in CCQE events
- Differences in the FSI models:
 - INCL FSI simulation features a significant fraction of events **without a proton** in the final state, especially low momentum protons region
 - INCL tends to **re-absorb** other particles produced during the cascade
 - An essential novelty of this study is the **simulation of nuclear cluster production** by INCL in FSI of neutrino interactions

A correct FSI simulation is **crucial** to achieving an **accurate ν energy reconstruction**.

Future prospects:

- We want to repeat the same study for the **antineutrinos**: the leading particle will be **neutron** and its modelling is crucial for the upgrade
- We want to continue the study of the detector response **of clusters**
 - **CCQE implementation** in INCL
 - **Neutron secondary interactions**: using INCL implemented in **Geant4**

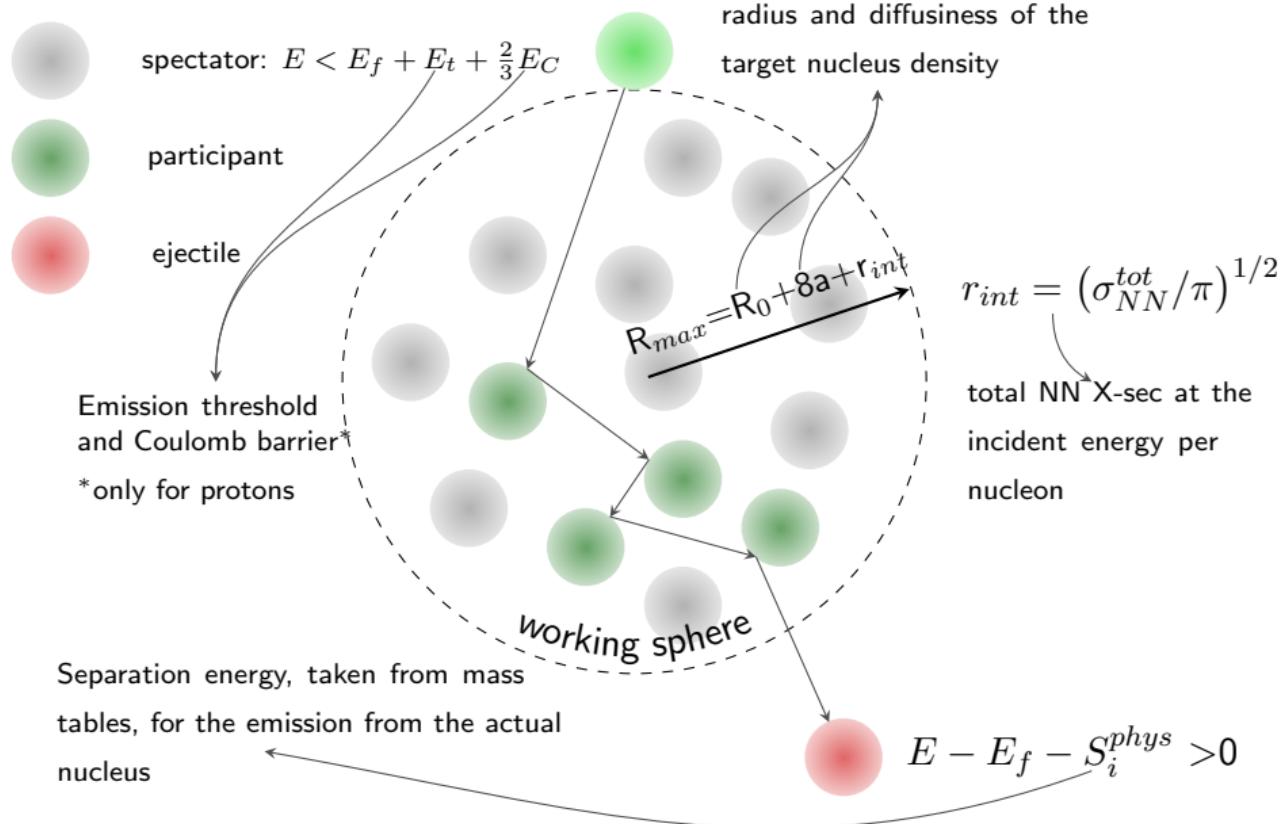


The paper has been published:

► Phys.Rev.D, 106 032009 (2022)

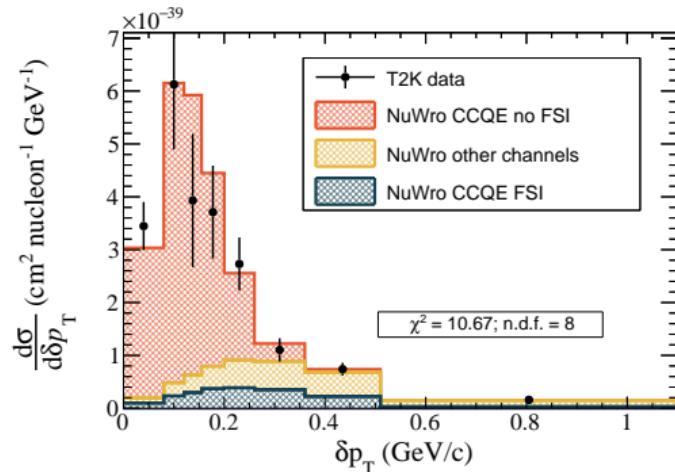
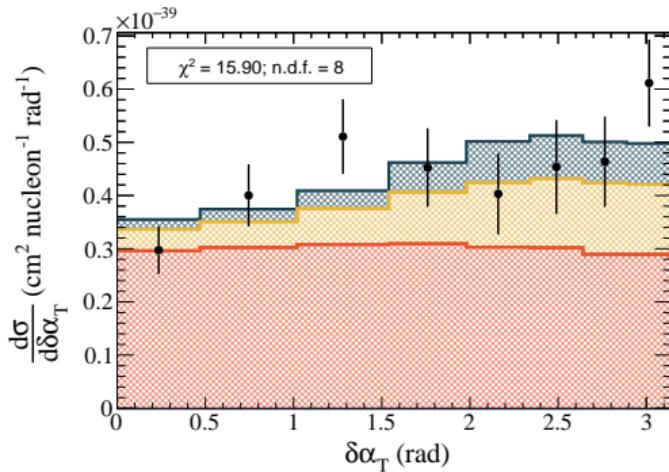
BACK UP

Standard INCL cascade



	Channel	NuWro	INCL
	no protons	1.37%	19.47%
	protons	98.63%	80.53%
no proton	absorption	4.45%	39.49%
	neutron+ π production	3.40%	0.60%
	π production	0.21%	0%
	neutron knock-out	91.4%	29.58%
	nuclear cluster knock-out	0%	30.33%
proton	1 proton, no FSI	70.38%	68.49%
	1 proton only with FSI	2.45%	19.21%
	1p+nucleons/nucl. clusters	26.21%	11.68%
	1p+ π production	0.96%	0.62%

NuWro comparison to T2K data



Comparison to MINER ν A data

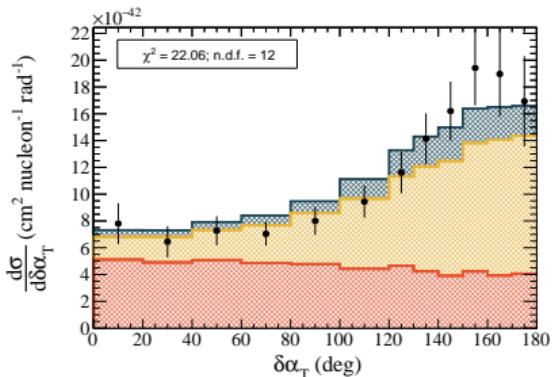
Cuts (MeV):

$$1500 < p_\mu < 10000$$

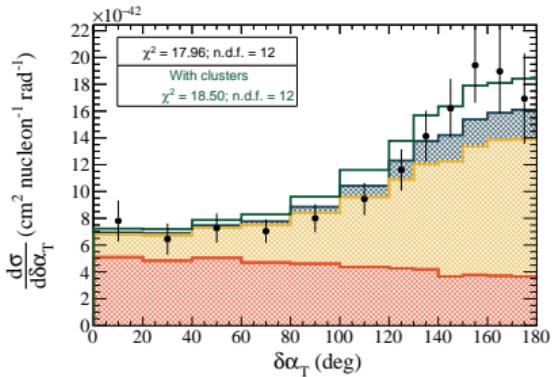
$$450 < p_p < 1200$$

$$\Theta_\mu < 20$$

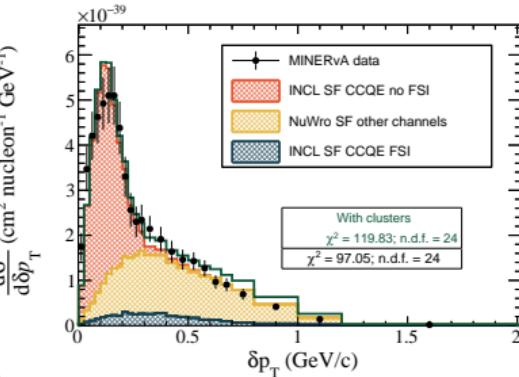
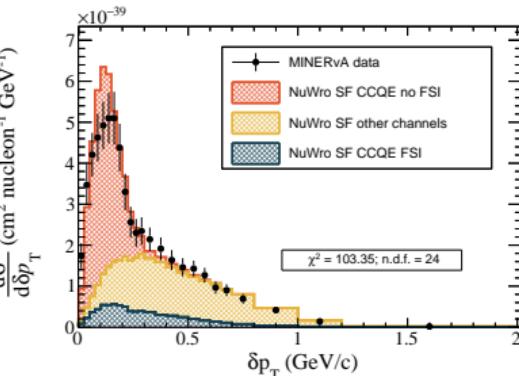
$$\Theta_p < 70$$

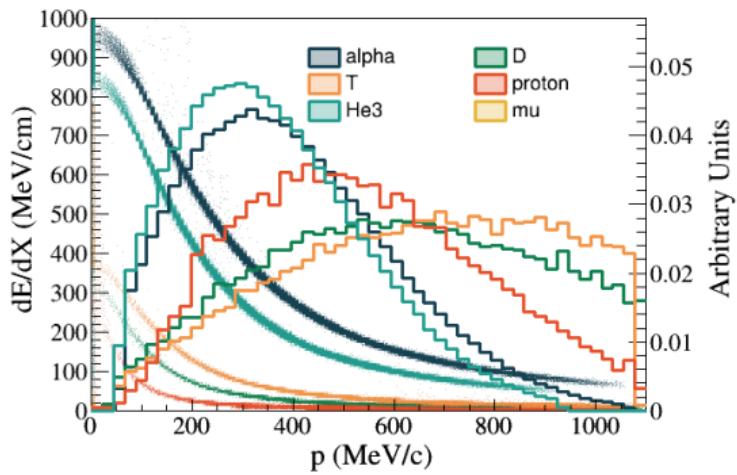


NuWro

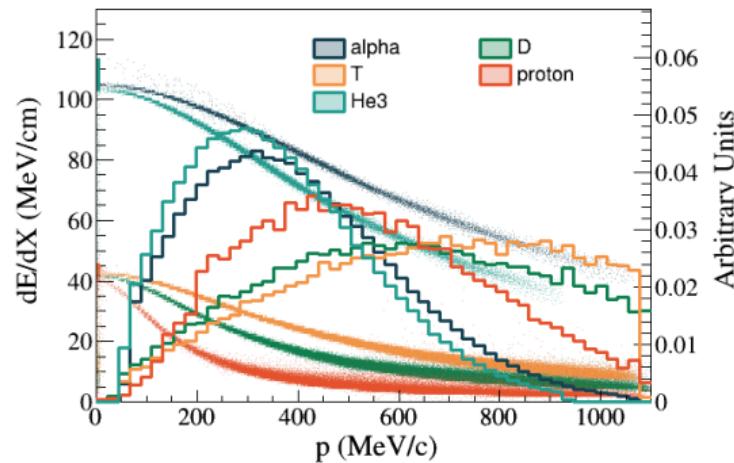


INCL

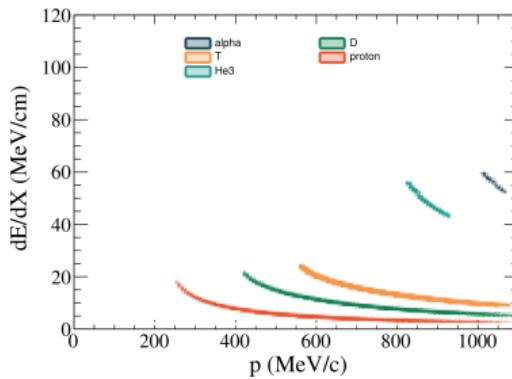
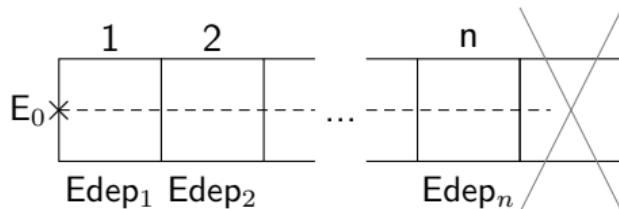




Energy loss by ionization



Visible energy loss by ionization (with Birks correction)



- **initial kinetic energy** E_0 is reconstructed as a sum of energy deposits along the whole track
- **momentum after passing 1 cm** is reconstructed using **5** mass hypotheses
- **for each momentum hypothesis**, the $\frac{dE}{dX}_{rec}$ is calculated using the $\frac{dE}{dX}$ dependence on momentum plot
- $\chi^2 = \sum \frac{(\frac{dE}{dX}_{sim} - \frac{dE}{dX}_{rec})^2}{\sigma^2}$ is calculated for each hypothesis
- we choose hypothesis with the **lowest** χ^2