BAYESIAN INFERENCE FOR LISA DATA ANALYSIS

Quentin Baghi, CEA Saclay

Tuesday, November 15th, 2022

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Rencontre du groupe de travail "méthodes d'analyse des données" du GdR oG





- 1. Challenges of LISA data analysis
- 2. Overview of used Bayesian concepts
- 3. Towards the future







- > The analysis of LISA data will be drastically **different from current ground-based detection**:
 - ♦ Numerous superimposed sources ≠ isolated events
 - Different time scales, larger waveform cycles observed
 - ◆ Signal-dominated measurement ≠ noise-dominated



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Research problem



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 - Stochastic noise
 - Instrumental transients (glitches)
 - Non-stationarities
 - Spectral lines
 - ✤ Data gaps

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Disturbances





- What kind of data will LISA measure?
 - + Fractional frequency deviations (relative doppler shits) from 27 interferometers
 - + Times series sampled at 4 Hz, observed over 4+ years with 89% duty cycle
 - Dominated by laser frequency noise
 - + After pre-processing, obtain 3 time-delay interferometry (TDI) data streams (X, Y, Z)



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- What is the strategy to analyse the data?
 - ✦ Bayesian framework: probe the parameters + number of model components posterior

$$p(\boldsymbol{\theta}, k | \boldsymbol{d}) = \frac{p(\boldsymbol{d} | \boldsymbol{\theta}, k) p(\boldsymbol{\theta}, k)}{p(\boldsymbol{d})}$$

Ced

CONS

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<u>Cea</u>

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Data vector. For example d=(X, Y, Z)
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$$p\left(\boldsymbol{d} \mid \boldsymbol{\theta}, k\right) = \frac{1}{\sqrt{(2\pi)^{N} |\boldsymbol{\Sigma}(\boldsymbol{\theta})|}} \exp\left\{\left(\boldsymbol{d} - \boldsymbol{h}(\boldsymbol{\theta}, k)\right)^{\dagger} \boldsymbol{\Sigma}(\boldsymbol{\theta})^{-1} \left(\boldsymbol{d} - \boldsymbol{h}(\boldsymbol{\theta}, k)\right)\right\}$$

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GW signals:
$$\boldsymbol{h}(\boldsymbol{\theta}, k) = \sum_{j=1}^{k} \boldsymbol{h}_{j}(\boldsymbol{\theta}_{j})$$

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$$h(\theta, k) = \sum_{j=1}^{k} h_j(\theta_j)$$
 Stochastic processes: $\Sigma(\theta) = \sum_{i=1}^{p} \Sigma_i(\theta_i)$

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> 1 Sample for $p(\boldsymbol{\theta}_{\mathrm{GB}}|\boldsymbol{y}, \boldsymbol{\theta}_{\mathrm{others}})$ 2 Sample for $p(\boldsymbol{\theta}_{\mathrm{MBHB}}|\boldsymbol{y}, \boldsymbol{\theta}_{\mathrm{others}})$



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- > The number of overlapping sources (especially Galactic binaries) is not know in advance
- Need to estimate the optimal number of sources



- Algorithm: reverse-jump Markov-chain Monte Carlo (RJMCMC)
- Allow for parallel computing by splitting the frequency-domain data into segments



- In the case of missing data points or gaps
- Example: interrupted science data due to antenna repointing
- Consequence: both the signal and the covariance become expensive to compute
- One strategy is data augmentation [Baghi et al, 2019]

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[Baghi et al, in prep.]



$$\tilde{d} = T_h \tilde{h} + T_n \tilde{n}$$

[Baghi et al, in prep.]



TDI data
$$\longrightarrow \tilde{d} = T_h \tilde{h} + T_n \tilde{n}$$

[Baghi et al, in prep.]





[Baghi et al, in prep.]





[Baghi et al, in prep.]





[Baghi et al, in prep.]





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 $\boldsymbol{\Sigma}(f) = \mathbf{R}_h(f) S_h(f, \theta_h) + \mathbf{R}_n(f) S_n(f, \theta_n)$

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- We use **Bayesian concepts which require acceleration**:
 - ✤ Data representation and segmentation
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 - Reduced order modelling / heterodyned likelihood [See Sylvain Marsat's talk!]
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- Framework for research: **the LISA Data Challenges**
 - Collaborative playground https://lisa-ldc.lal.in2p3.fr/
 - Progressively increases the number of source types in "enchiladas" + instrumental realism
 - Writing of the LISA Data Analysis Living Reviews: sources, analysis methods, acceleration algorithms, specificities of LISA, challenges



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Thank you for your attention !



BACKUP SLIDES



What makes gravitational noise in the milihertz band?



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In our galaxy: pairs of orbiting white dwarfs



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One billion light-years away: collision of black holes



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In the entire universe: a cosmic gravitational wave background? Possibly farther away: merging supermassive black holes

