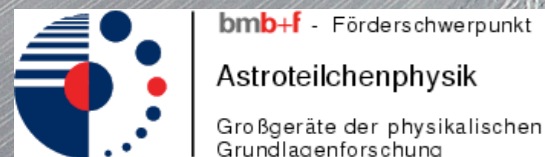
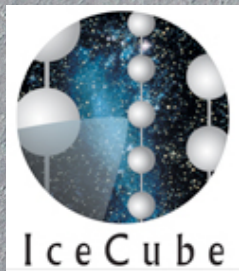


Status and Recent Results of the South Pole Acoustic Test Setup

Timo Karg
for the IceCube collaboration
Bergische Universität Wuppertal

ARENA 2010
29 June – 2 July 2010 in Nantes



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Aim of SPATS: Ice properties (10 – 100 kHz)

➔ Get realistic sensitivity estimate for an acoustic neutrino telescope in ice



- Speed of sound and its variation with depth
 - significant refraction would make vertex reconstruction difficult



- Attenuation length
 - determines sensor spacing / effective volume of neutrino detector
 - frequency dependence allows to determine attenuation mechanism



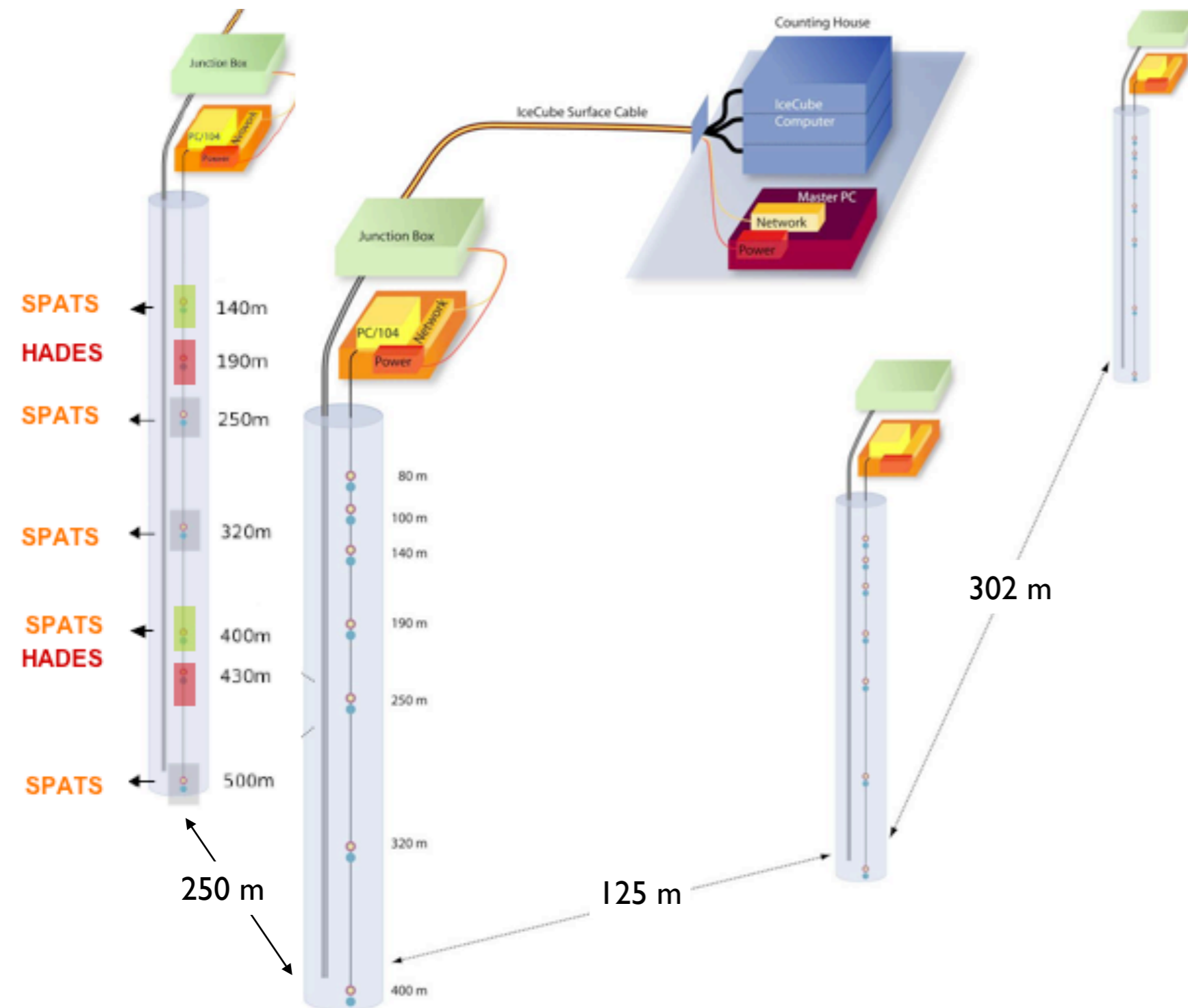
- Noise floor
 - determines energy threshold



- Transient noise sources
 - impulsive noise must be separated from neutrino signal

Hardware overview

South Pole Acoustic Test Setup (SPATS)



- 4 strings in IceCube drill holes
- instrumented depth: 80 m – 500 m
- per string:
 - 7 sensors
 - 7 transmitters
- String-PC
 - digitization
 - time stamping
 - monitoring (p,T)
- Master-PC
 - data storage
 - GPS clock
 - data transfer via satellite

Strings A, B, C installed in 2006/07
String D installed in 2007/08

SPATS stage design

Transmitter:

- ring shaped piezo ceramic coated in resin
- HV generator

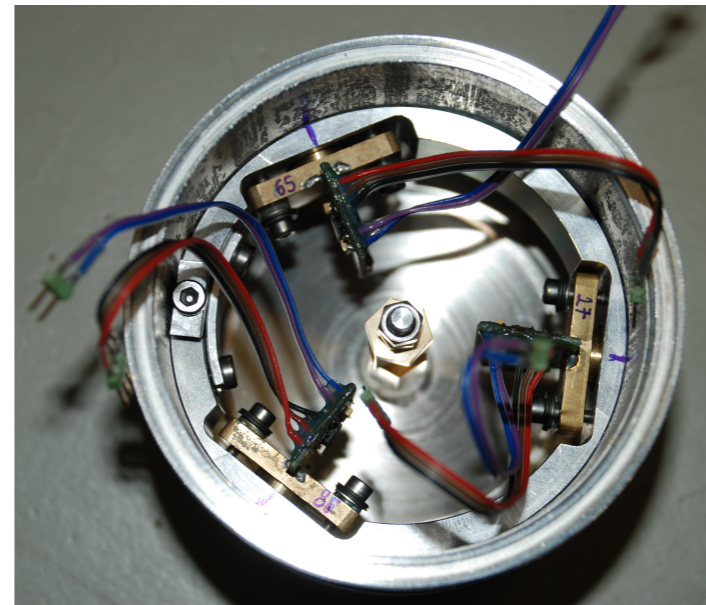
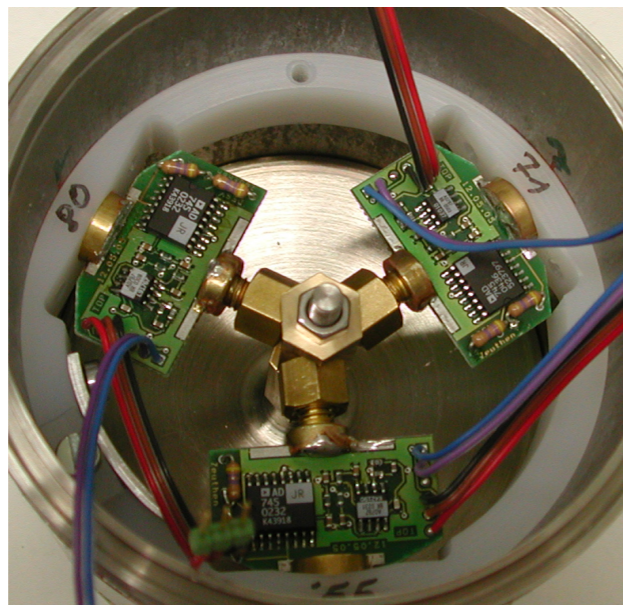
Sensor:

- 3 channels / sensor
- pre-amplifier
- analogue signal transmission
- steel pressure housing



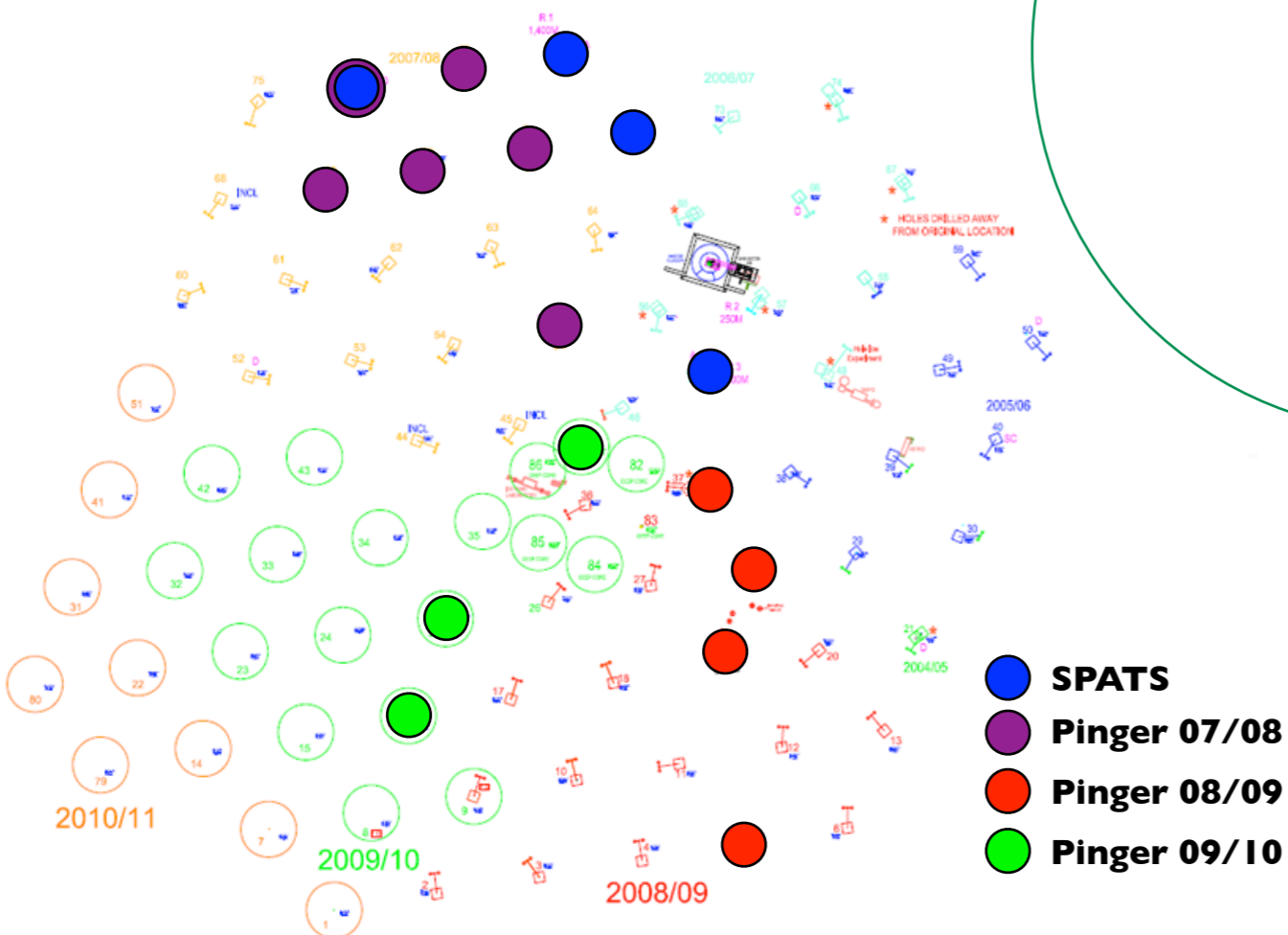
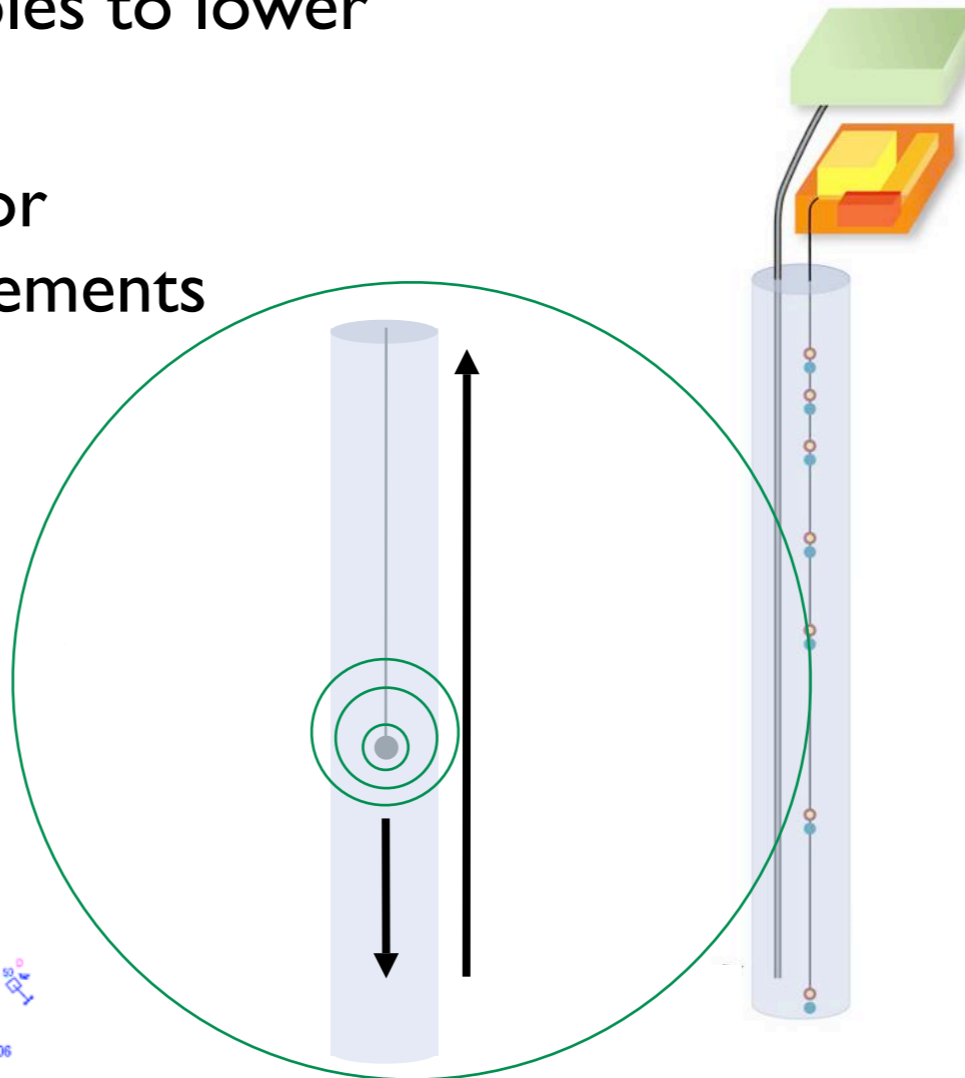
String-D:

- improved sensors: mechanical decoupling of sensor channels
- improved transmitters: higher power
- HADES: alternative sensor design with piezo ceramics outside the steel housing



Extended range: The retrievable pinger

- Use newly drilled IceCube holes to lower retrievable transmitter
 - increased distance range for attenuation length measurements
 - sound speed depth profile
 - relative sensor calibration



Speed of sound

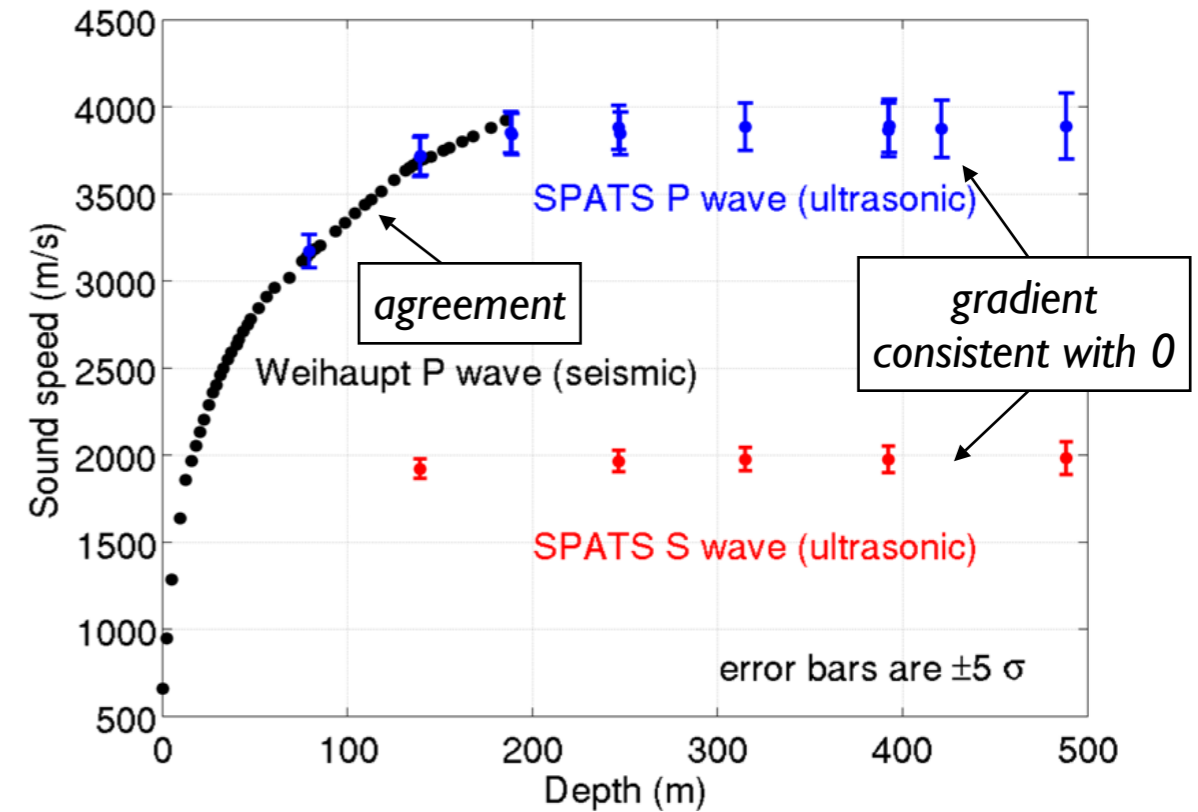
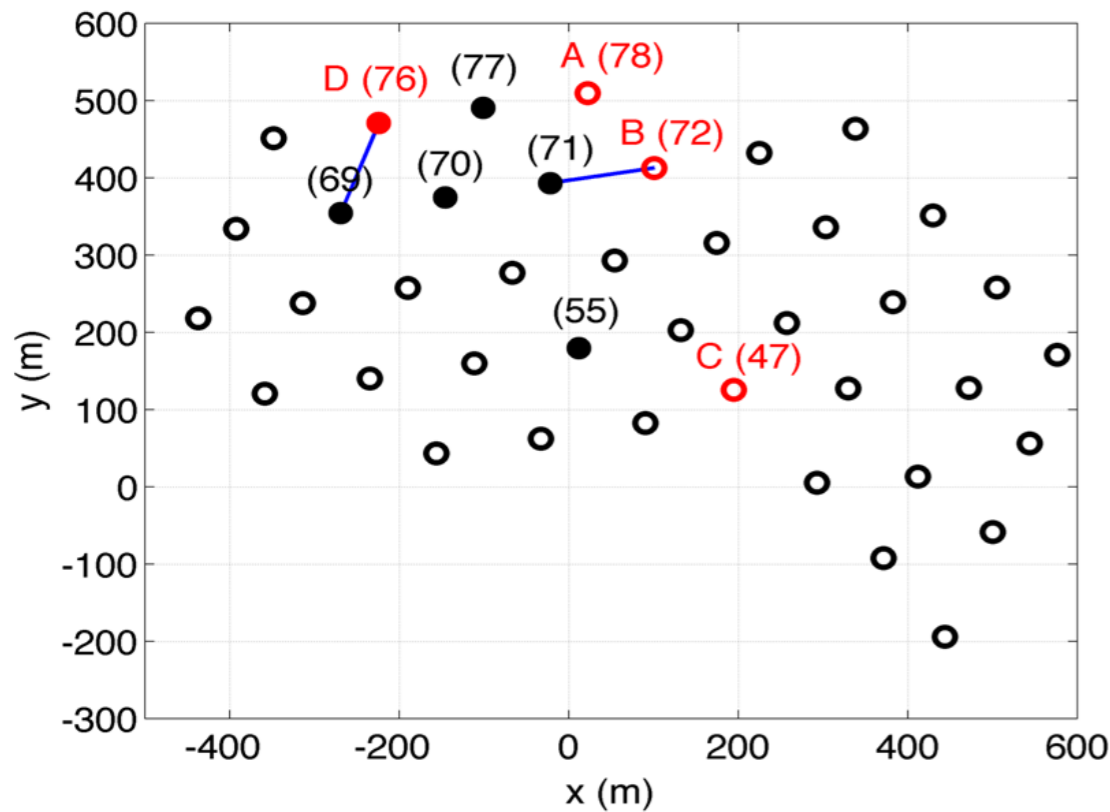
Astropart. Phys. **33** (2010) 277, arXiv:0909.2629 [astro-ph.IM]

&

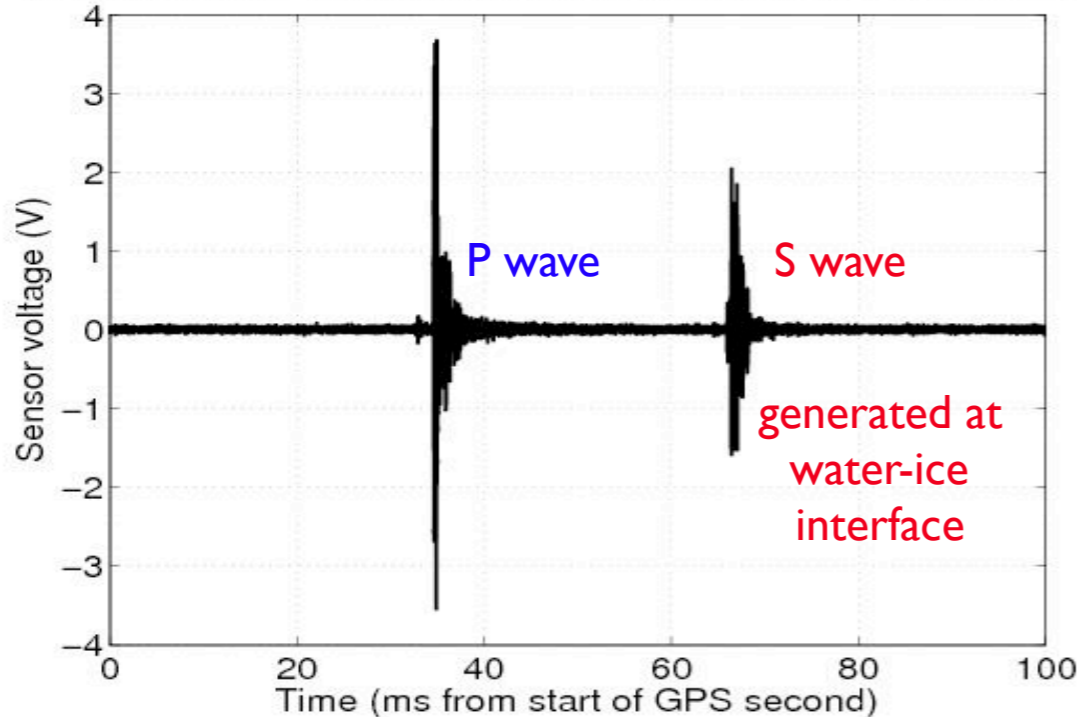
Attenuation length

submitted to Astropart. Phys., arXiv:1004.1694 [astro-ph.IM]

Sound speed depth profile



Hole 69 to DS7-0 (Run 30733, 488.4 m depth) average waveform



- 2 combinations, 125 m distance from pinger data season 2007-2008
- Better than 1% accuracy
- First measurement in situ for P and S waves

$$v_P(375m) = 3878 \pm 12 \text{ m/s}$$

$$v_S(375m) = 1975.8 \pm 8.0 \text{ m/s}$$

Attenuation length

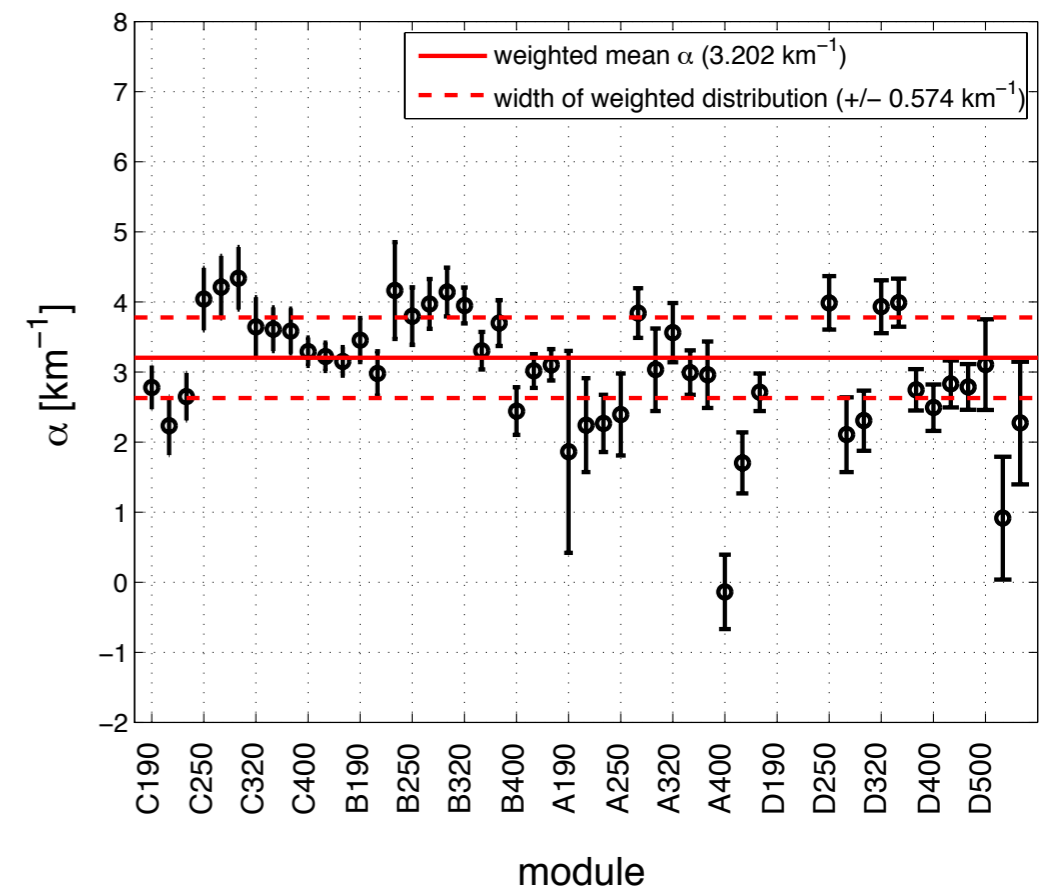
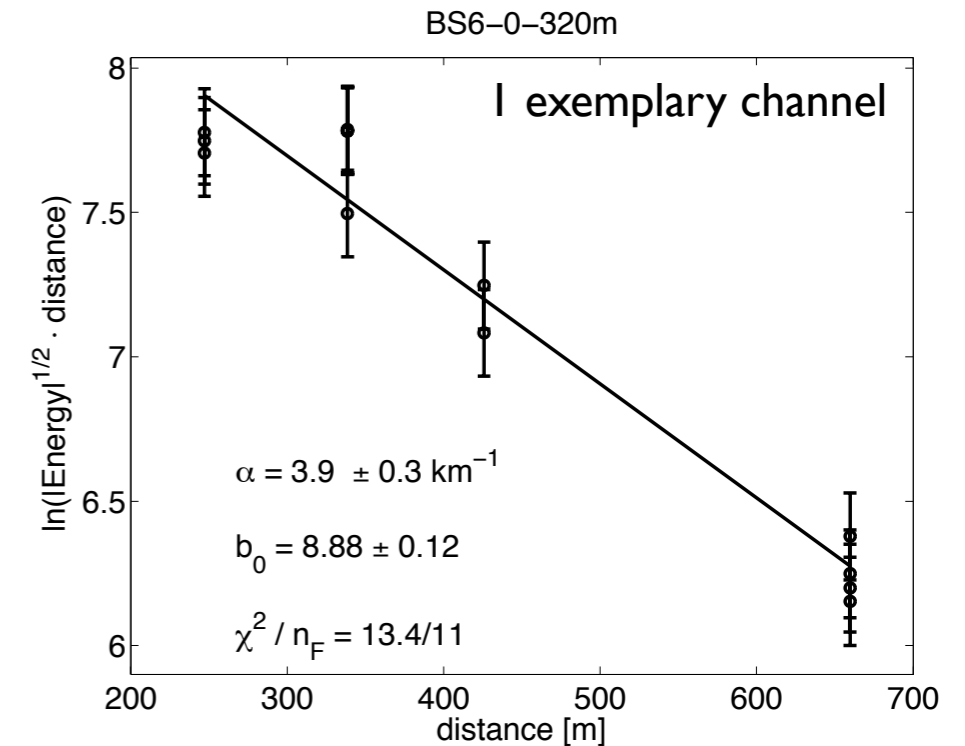
Expectation: several kilometers \leftrightarrow Measurement: 300 m

Analysis	α_{att} (km ⁻¹)
Pinger data (time domain)	3.20 \pm 0.57
Pinger data (frequency domain)	3.75 \pm 0.61
Inter-string data (same level)	3.16 \pm 1.05
Inter-string data (3-level ratios)	4.77 \pm 0.67
Transient events	3.64 \pm 0.29

- No significant evidence for depth dependence, but not excluded
- Unclear frequency dependence: absorption or scattering?
(analysis with new pinger data in progress)

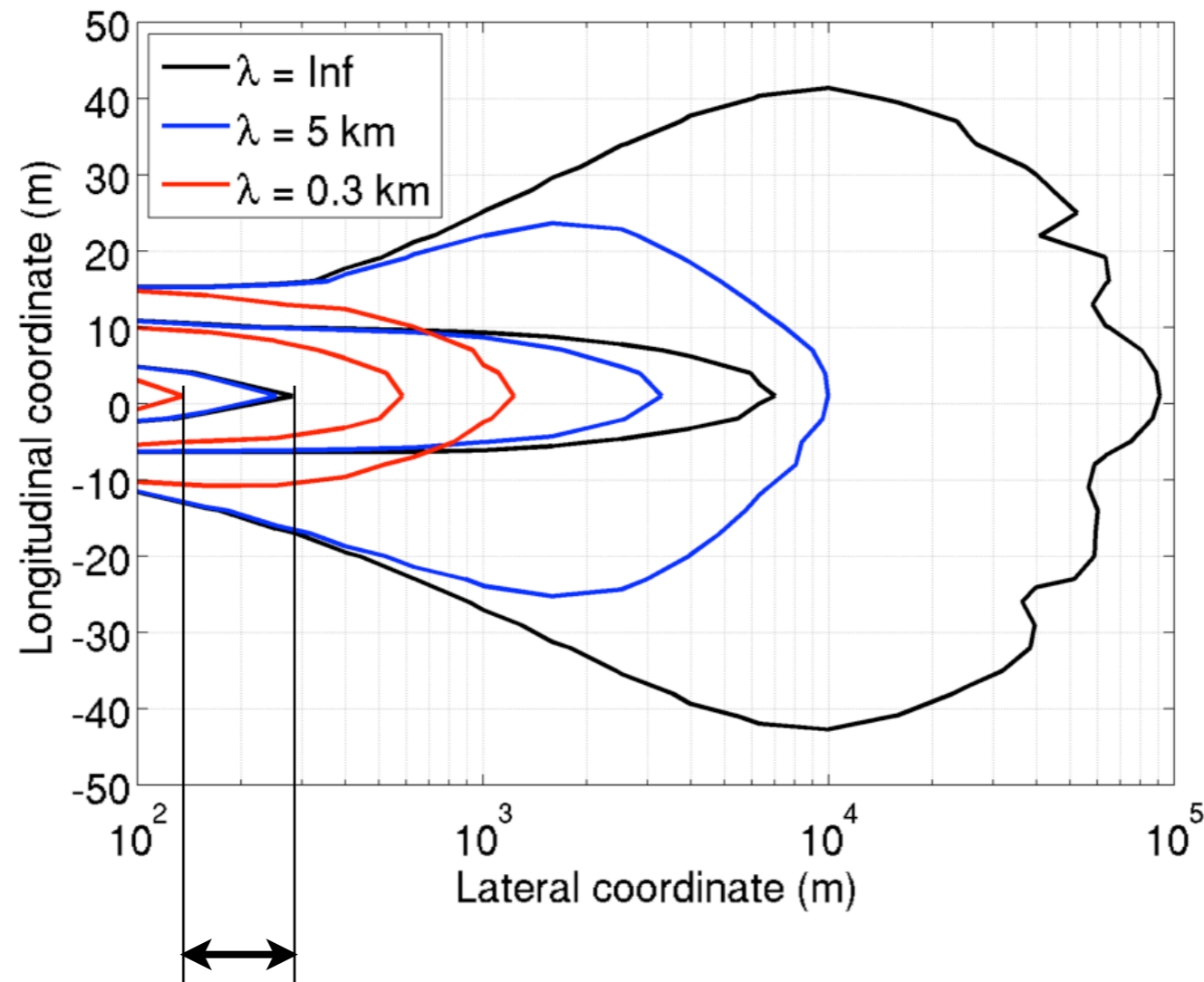
Pinger attenuation analysis

- Signal energy E calculated for each channel and over all pinger holes, noise subtracted from pinger-off runs
- Linear fit of $y = \ln(\text{distance} \times \sqrt{\text{energy}})$ yields attenuation coefficient α
- 48 independent measurements (sensor channels)
- Weighted mean value and width of distribution:
 $\alpha = 3.20 \pm 0.57 \text{ km}^{-1} \leftrightarrow$
 $\lambda = 312 \text{ m}^{+68 \text{ m}}_{-47 \text{ m}}$



Implications of short attenuation length

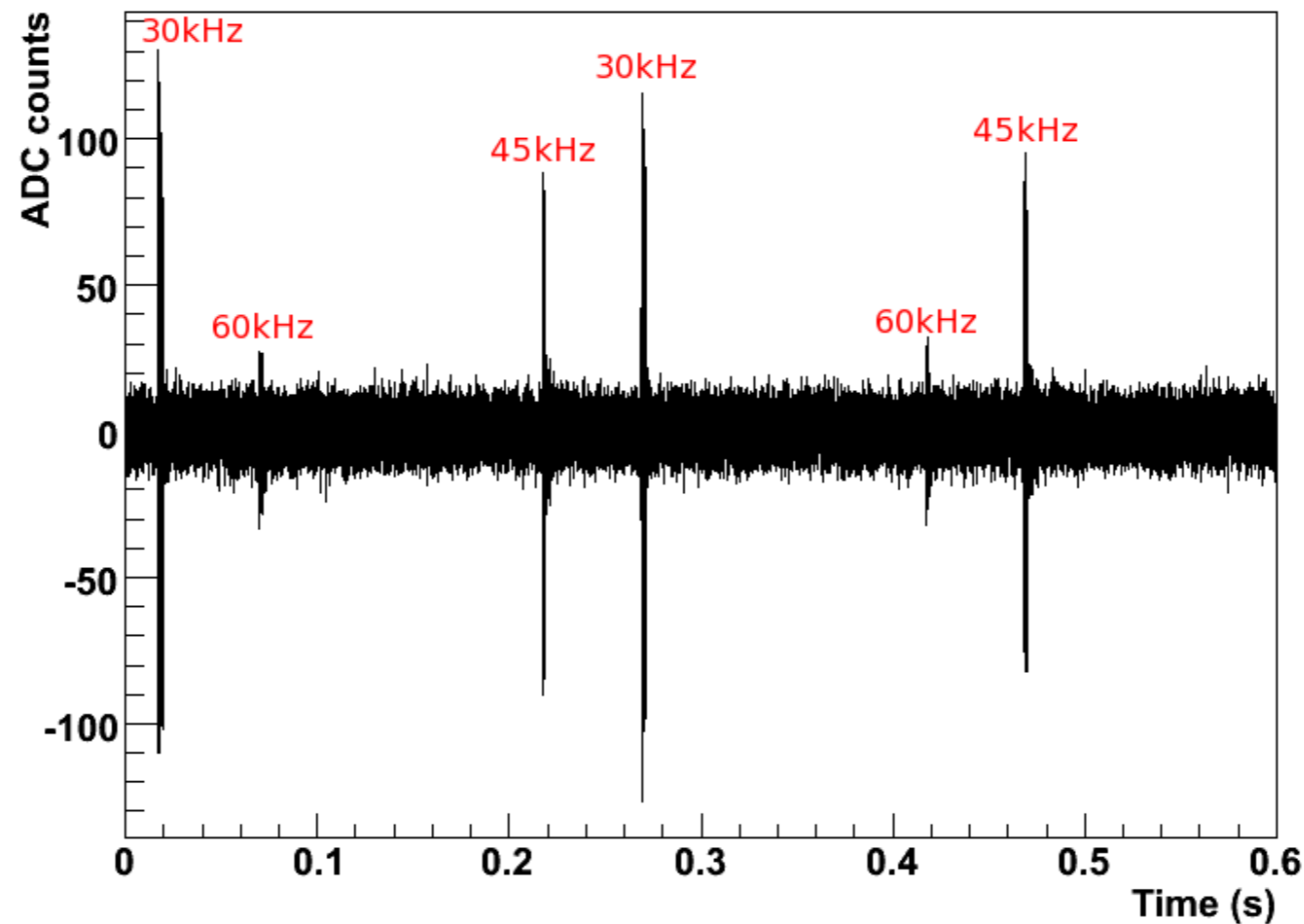
Pancake contours ($P_{\max} = 9$ mPa) for $E_{\nu} = 10^{18}$, 10^{19} , and 10^{20} eV



@ 10^{18} eV: Loose Factor ~ 2 in “visibility range”

- Short attenuation length only relevant at high ($> 10^{19}$ eV) energies
- At low energies (where expected flux is highest) $1/r$ decrease dominant

Further studies: The “multi-frequency” pinger 2009/10



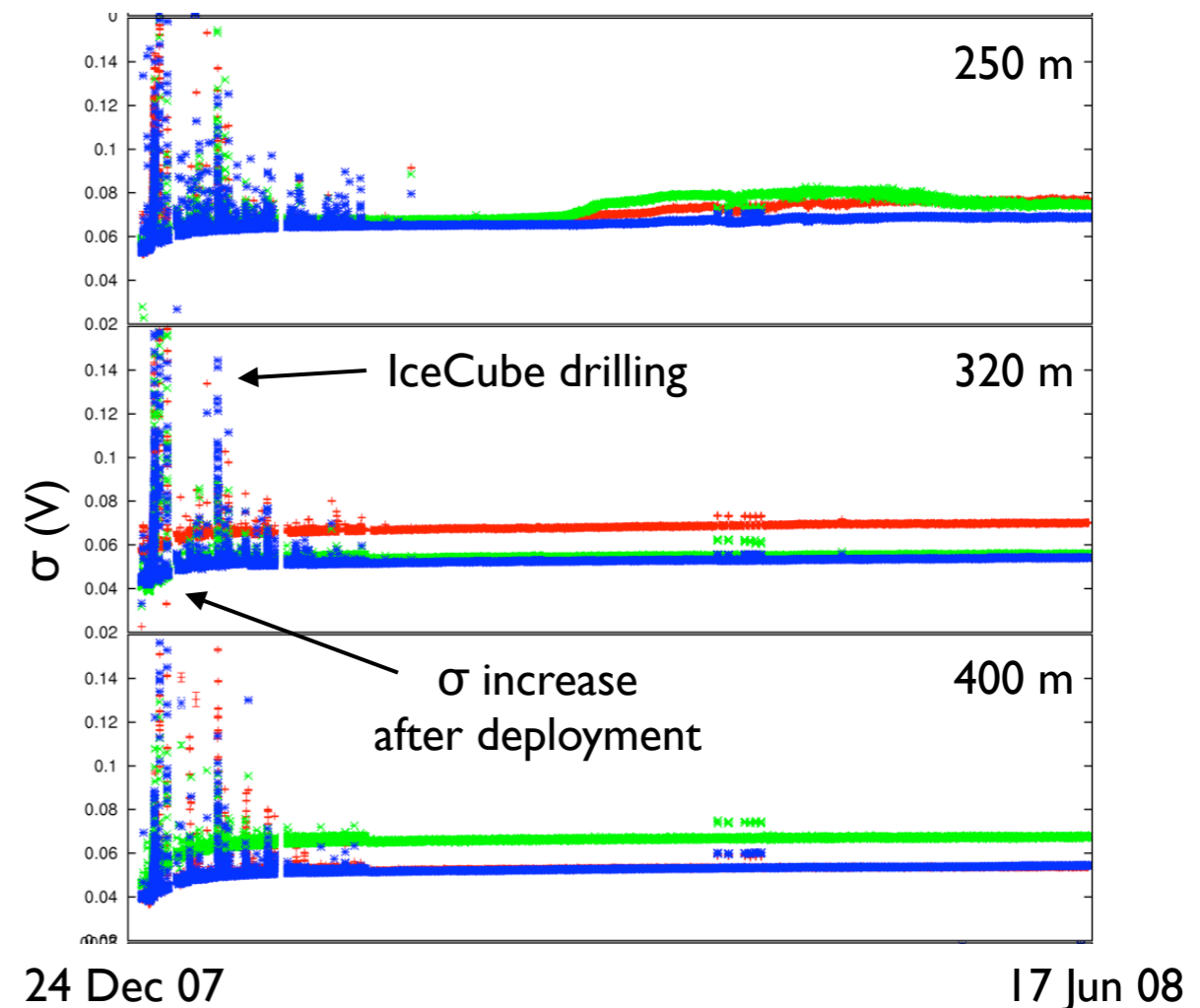
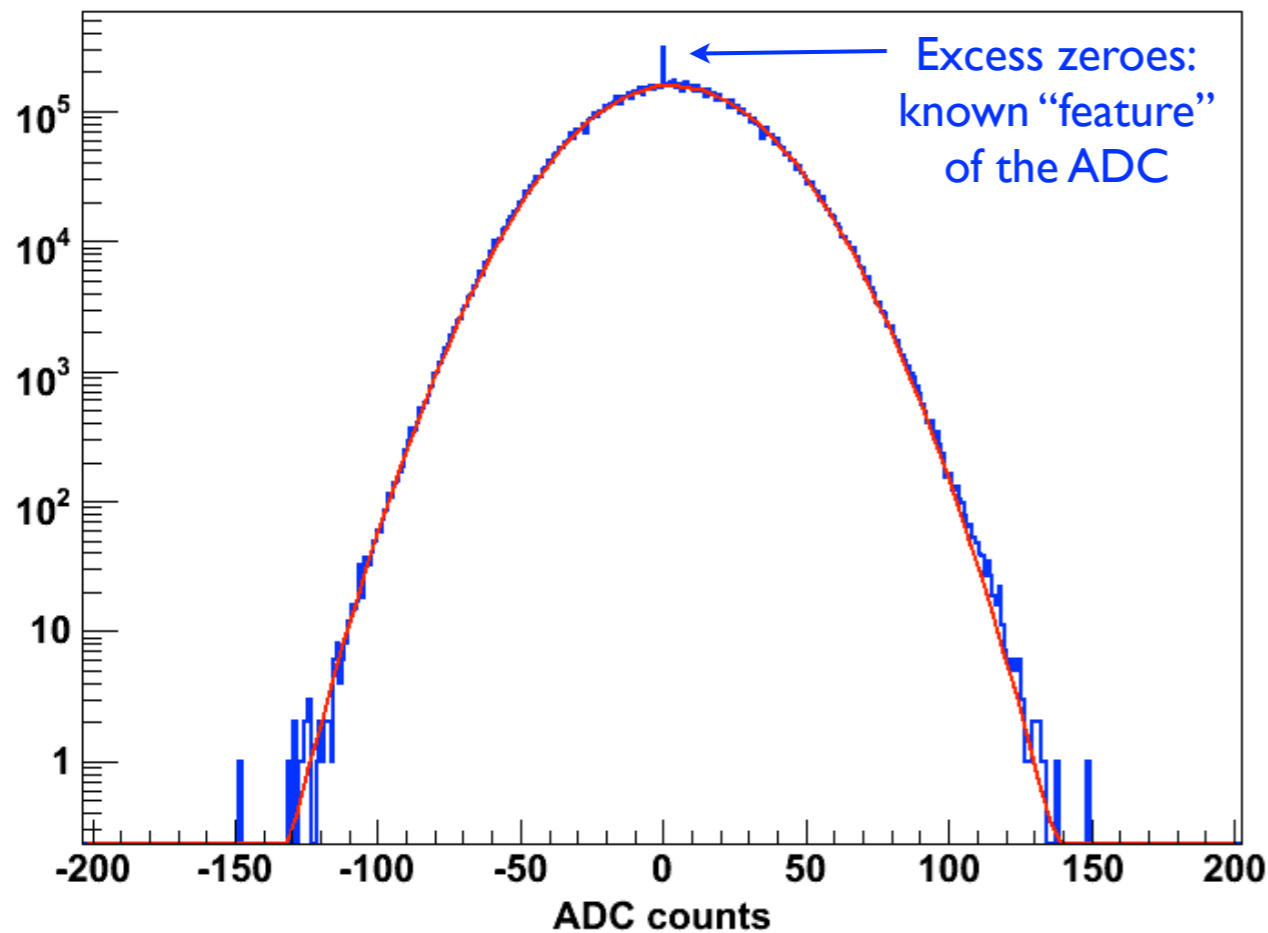
- Frequency dependence of attenuation length \Rightarrow attenuation mechanism
 - **Absorption**: frequency independent; **Scattering**: $\alpha_{\text{att}} \propto f^4$
- Deep stops (up to 1000 m) to measure sound speed and attenuation in deep ice and on inclined paths
 - Also interesting for Glaciology (ice crystal orientation)
- Data under study

Absolute noise level & Sensor calibration

see also
poster on sensor calibration
(L. Paul et al.)

Noise: properties and temporal evolution

- Gaussian and stable over long time
- Peaks correlated with IceCube drilling, inter-string data taking
- Hypothesis: freeze-in improves coupling to ice causing noise level to increase and then stabilize in the first couple months



Absolute noise level

- SPATS sensors have been calibrated in water at 0°C prior to deployment (relative to a reference hydrophone SensorTech SQ-03)
- In-situ calibration is challenging, but
 - ▶ Can study different effects separately in the lab
 - Temperature -50°C:
 - sensitivity increase by factor 1.5 in air
 - Increased static ambient pressure:
 - sensitivity stable within 30% in water at room temperature
 - Different coupling from medium to sensor (acoustic impedance)
 - under study (cf. calibration poster)



Wuppertal Water Tank
11 m³ water

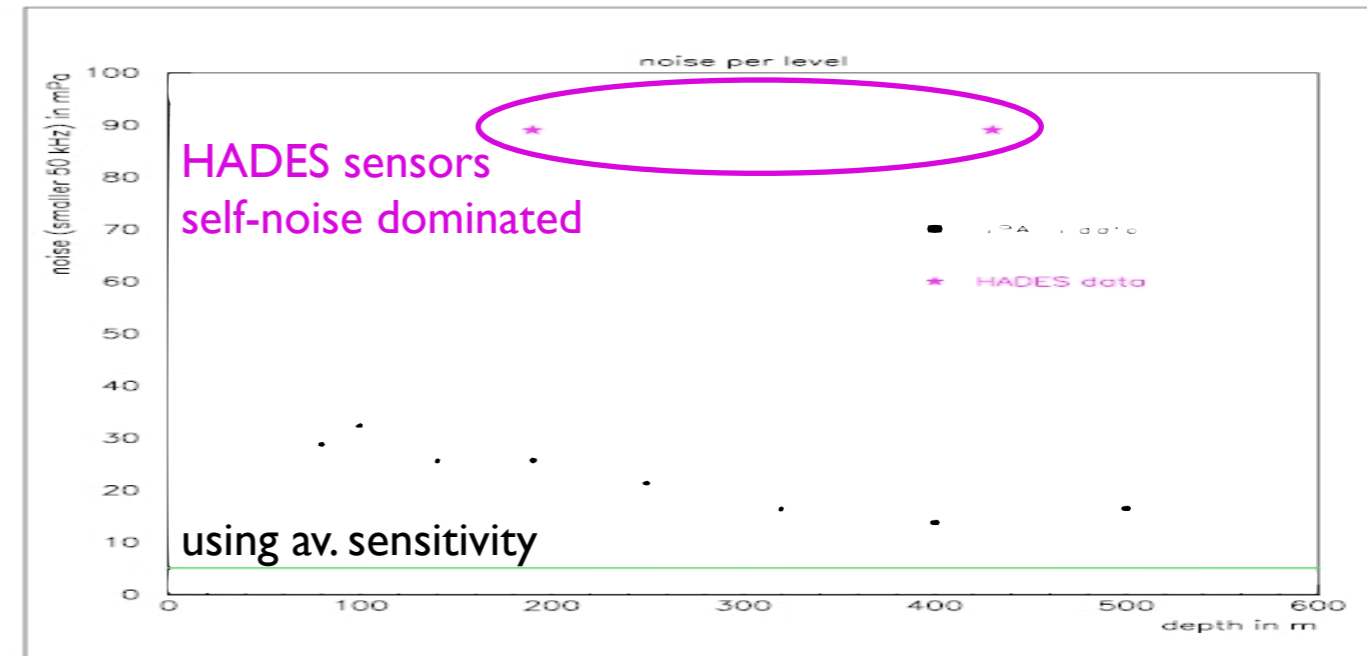
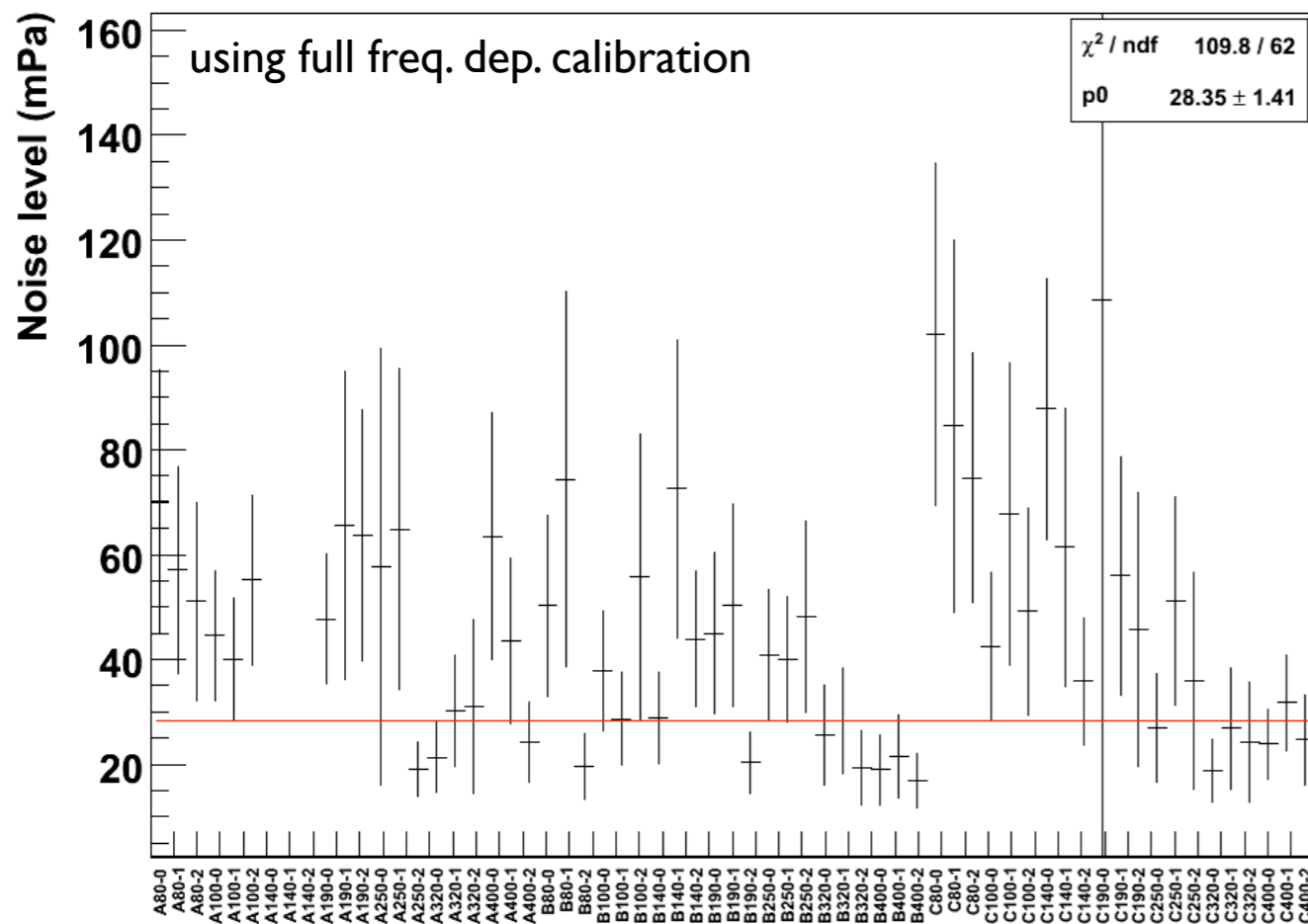


Aachen Ice Tank
3 m³ bubble free ice



Uppsala Pressure Vessel
liquid filled,
cable feeds,
up to 1000 bar

Estimation of absolute noise level (10 – 50 kHz)

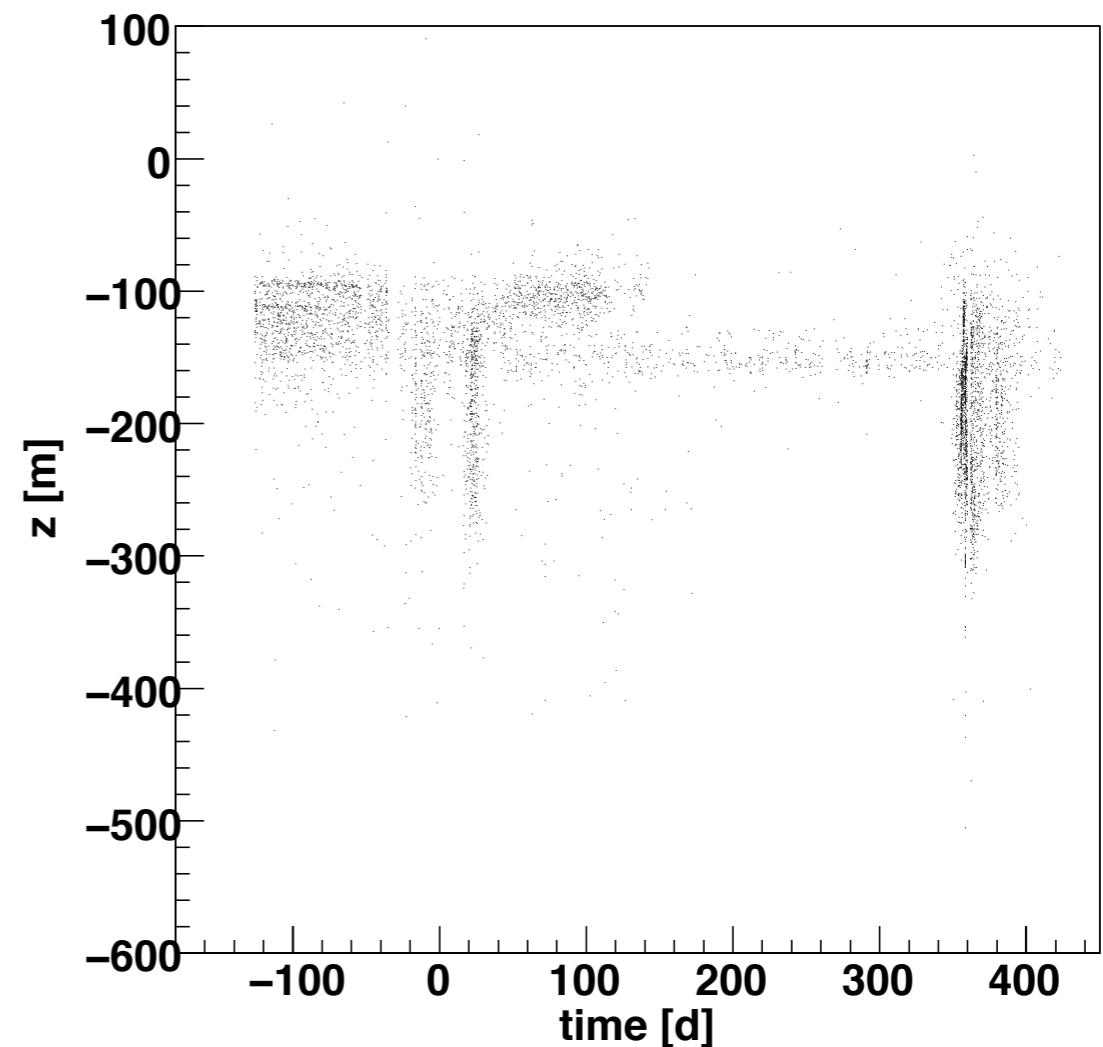
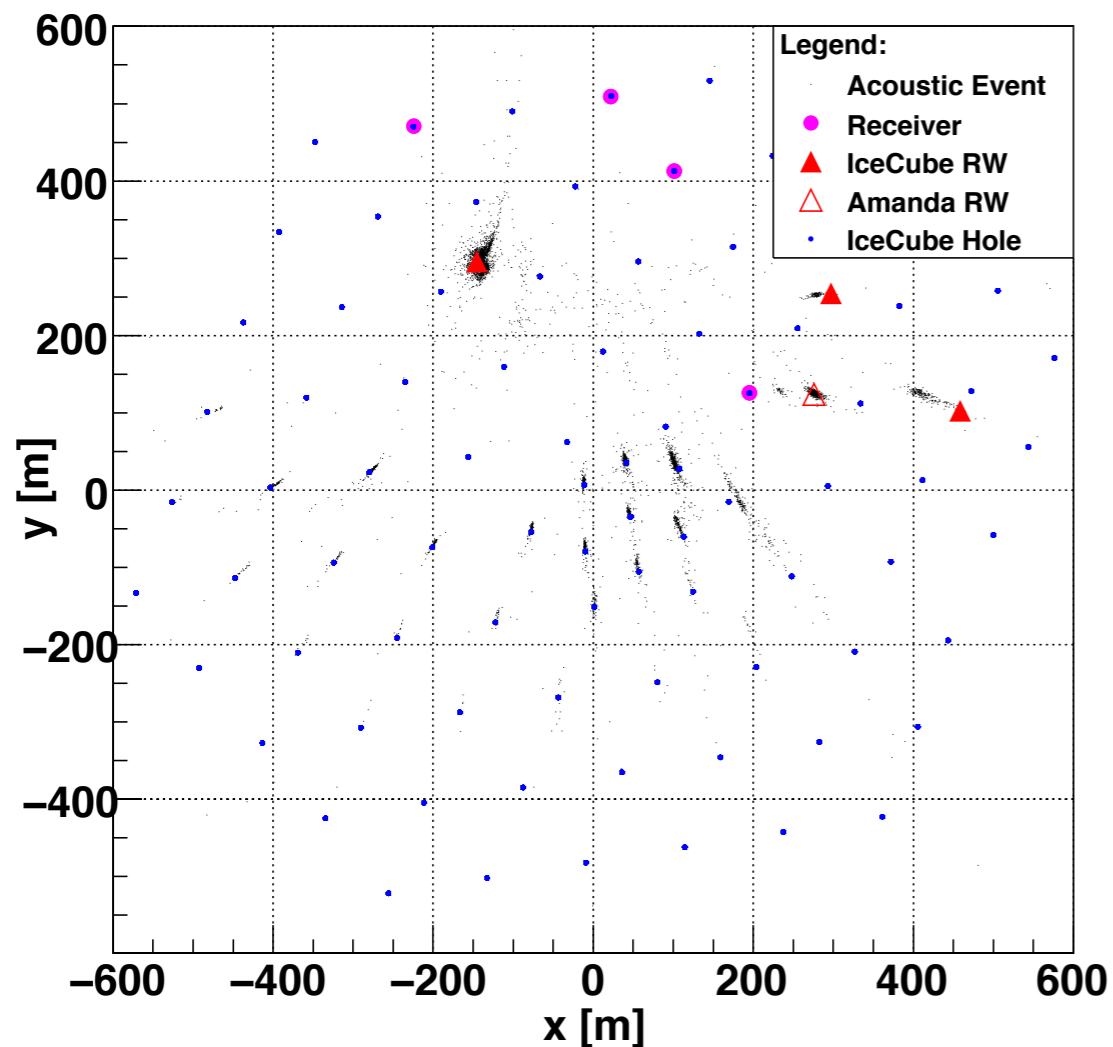


- Use sensitivity correction factor of 1.5 (from temperature)
- Sensitivity change due to freeze in under study
- Different approaches agree within a factor of 2
- In-situ measurement with different type of low noise glaciophone planned for 2010/11 season

Transient sources & Event reconstruction

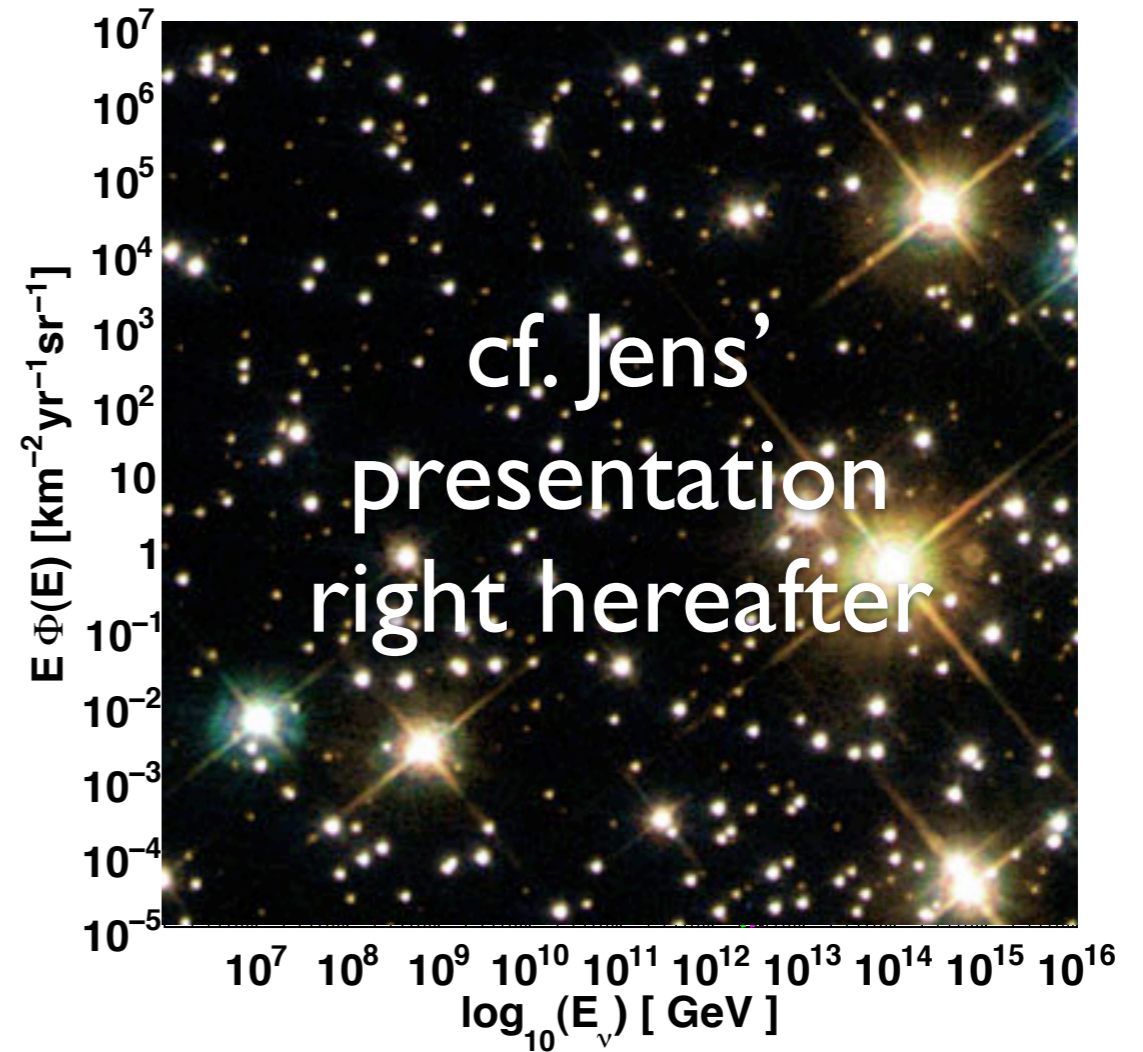
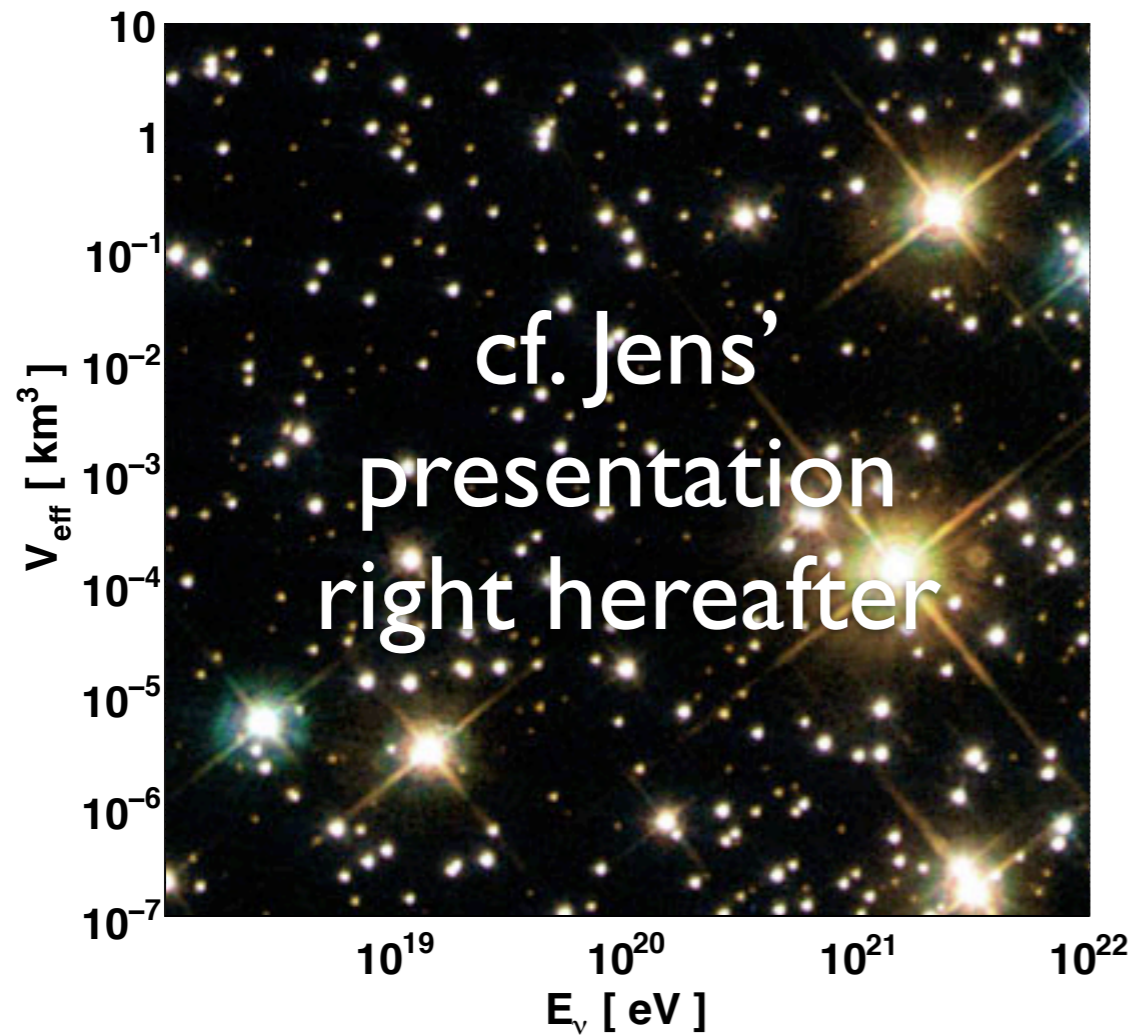
see also
talk on SPATS sensitivity
(J. Berdermann)

Spatial and temporal distribution of vertices



- Transient data recorded for 45 minutes of every hour
 - nearly 2 years of data available
- Only shallow (top 200 m) transients observed outside IceCube drilling
- Refreezing of close IceCube holes and Rod-wells are main (only) sources

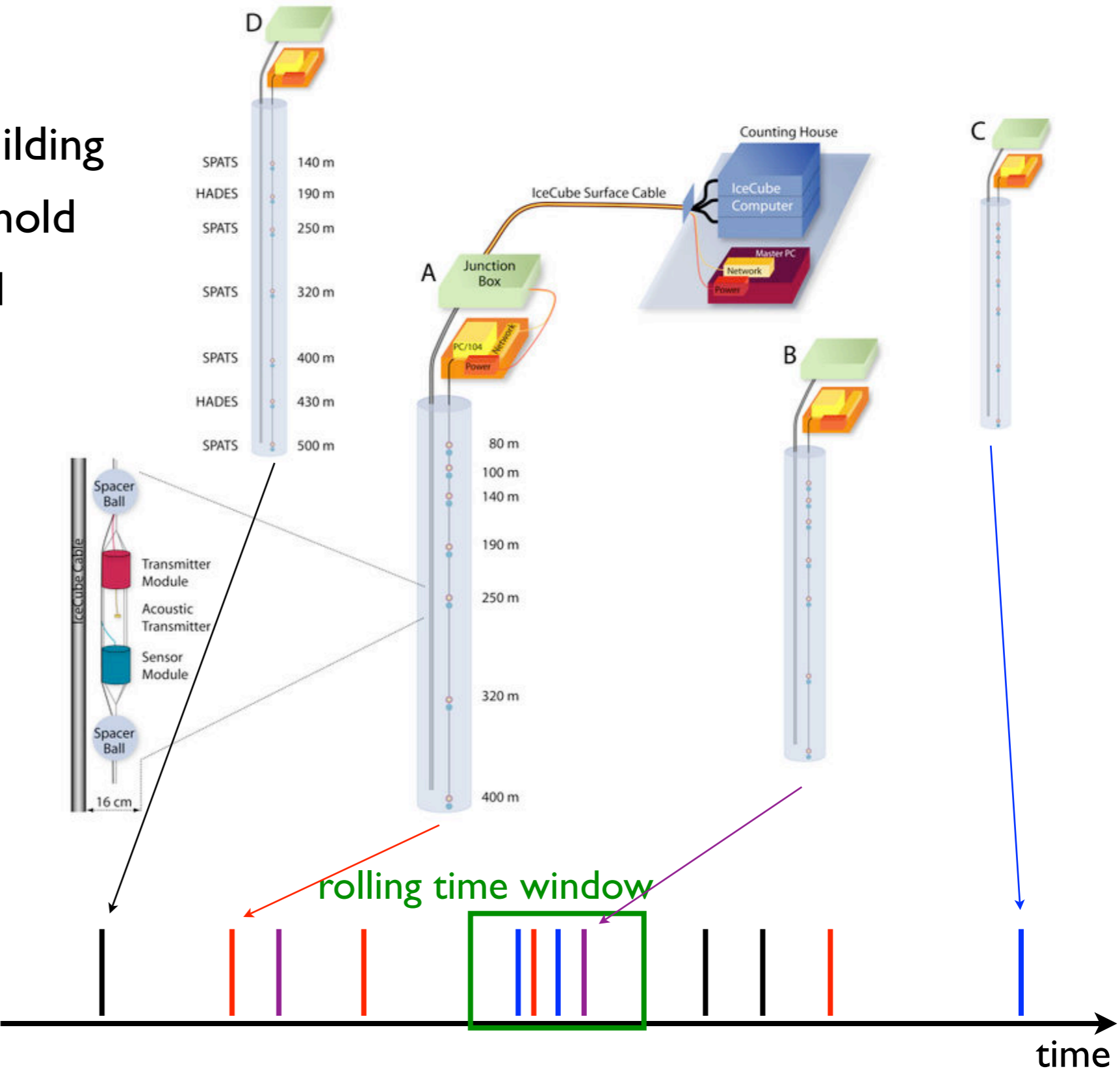
SPATS sensitivity



- Simulate ν interactions outside IceCube volume and below 200 m (no refraction)
- Trigger threshold 50 mPa (estimated)
- Very promising result for a test setup

DAQ software upgrade

- Online coincidence building
 - lower trigger threshold
- Currently being tested at South Pole





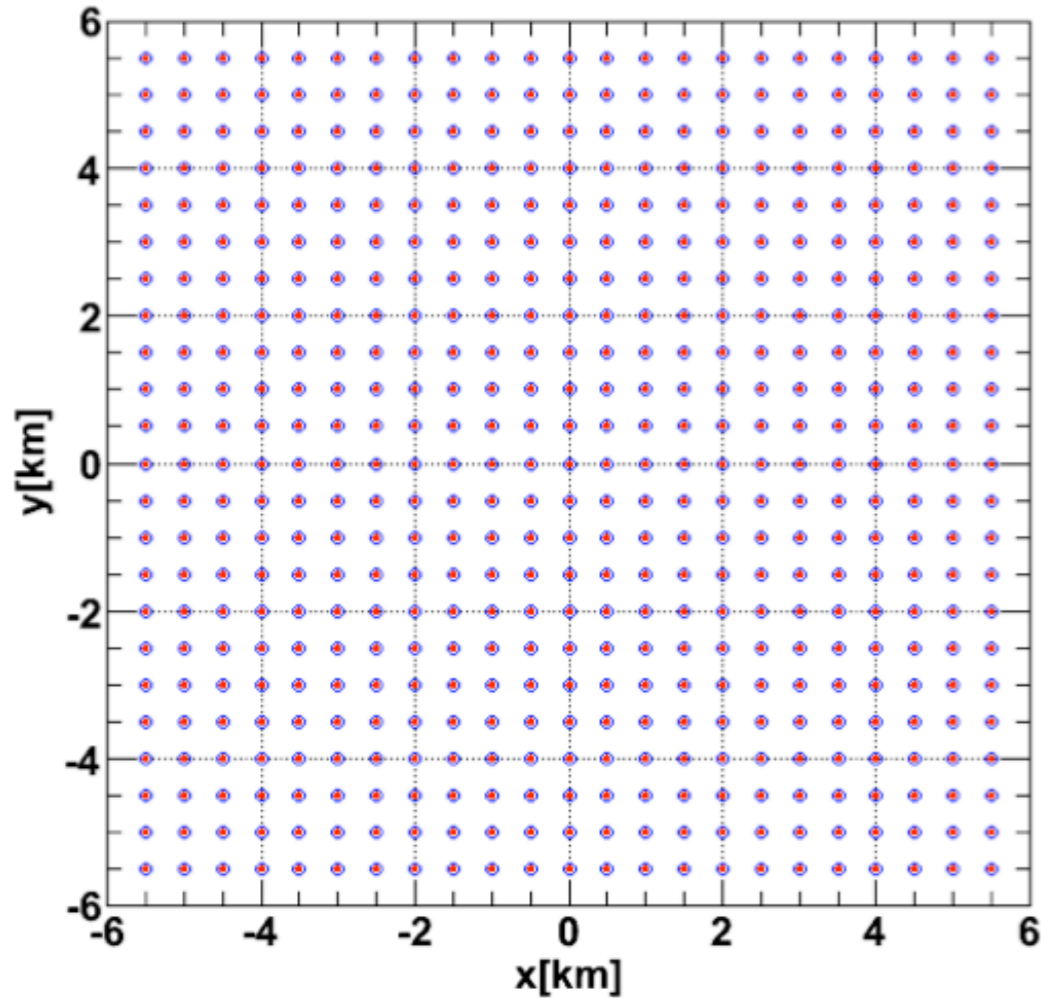
Implications for acoustic neutrino detection at South Pole

Implications for acoustic neutrino detection

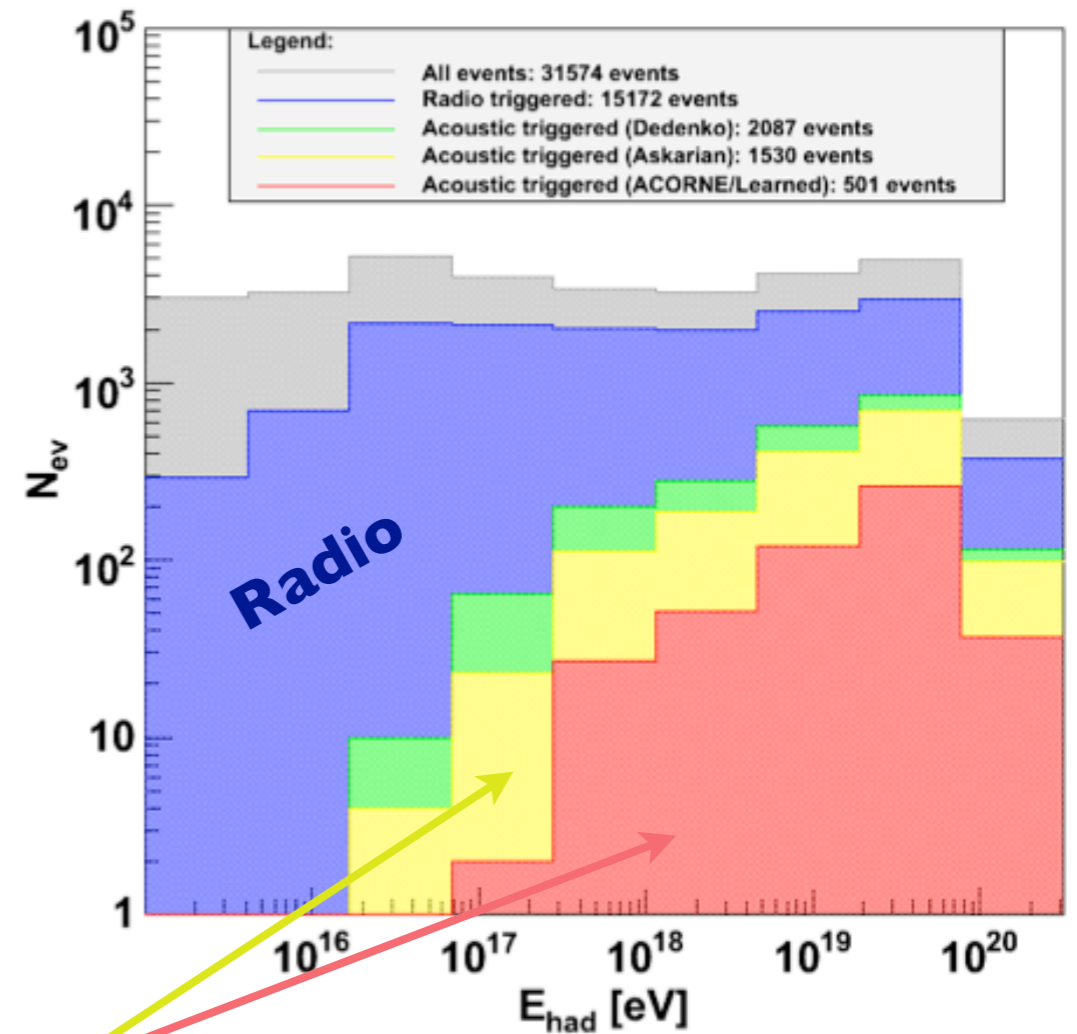
- If acoustic noise level is low:
 - low noise sensors can compensate for short attenuation
- Acoustics can add valuable additional information in hybrid neutrino detector (e.g. in combination with radio)
 - reduce systematic uncertainties
 - increase confidence in signal in absence of calibration beam (e.g. atmospheric muons / neutrinos in optical neutrino telescopes)
- Possible scenario:
 - at “low” energies: use radio for triggering: signal in single acoustic sensor significant
 - at “high” energies: separate(?) acoustic trigger: use hybrid information for direction/energy reconstruction

Possible radio/acoustic hybrid scenario

25 km², string spacing 500 m (441 strings)
8 acoustic sensors between 200 m and 300 m



Energy (StringNr. ≥ 1 , Sensors ≥ 1)



Radio/acoustic coincidences

Detection Mode	Trigger	Learned	Askaryan
1Sensor	Radio	3%	10%
4Sensors (1 String)	Acoustic	1%	5%
4Sensors (3 Strings)	Acoustic	0.1%	3%

Acoustic GZK events / year

	Learned	Askaryan
ν_μ	0.4	1.7
ν_e	0.2	0.8
Total	0.6	2.5

Aim of SPATS: Mission nearly accomplished

- Speed of sound and Refraction
 - speed of sound constant below 200 m
no refraction
- Attenuation length
 - $\lambda \approx 300$ m (20% uncertainty), factor 30 smaller than expected
possible explanation: larger influence of scattering
 - frequency and depth dependence under investigation
- Noise floor
 - Gaussian and stable
 - Comparable to deep sea (with reasonable assumptions)
better results soon to come
- Transient noise
 - Small rate and all deep events from identified sources

Open questions and Plans

- Absolute noise level
 - Deployment of low noise sensor pre-calibrated in ice planned for 2010/11
 - Study mechanism of surprisingly short attenuation length
 - Interest from Glaciology community
 - Data available from “multi-frequency” pinger
- ➔ Have robust sensitivity estimate for acoustic technique at South Pole within next 12 months