



Gluod_g*namics*



From initial gluons to hydrodynamics: Gluons inside hadrons and their thermalization

Oct. 24-25, Institut Pascal



Generalized parton distributions and energy-momentum tensor

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October 24, Institut Pascal, Orsay, France

Energy-momentum tensor (EMT)

EMT is a key **fundamental** object

It is the conserved current associated with invariance under **spacetime translations**

$$P^\mu = \int d^3x T^{0\mu}(x)$$

$$\phi(x+a) = e^{iP \cdot a} \phi(x) e^{-iP \cdot a} \Leftrightarrow i[P^\mu, \phi(x)] = \partial^\mu \phi(x)$$

It also plays the role of **source** for gravitation in the Einstein equations of **GR**

$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = \kappa T_{\mu\nu}$$

QCD EMT operator

$$T^{\mu\nu} = T_q^{\mu\nu} + T_g^{\mu\nu}$$

$$T_q^{\mu\nu} = \bar{\psi} \gamma^\mu \frac{i}{2} \overleftrightarrow{D}^\nu \psi$$

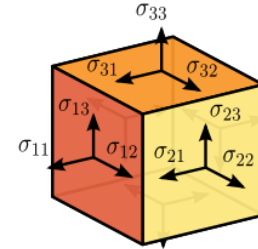
$$T_g^{\mu\nu} = -G^{\mu\lambda} G^\nu{}_\lambda + \frac{1}{4} g^{\mu\nu} G^2$$

Energy-momentum tensor (EMT)

Mass, spin and pressure are all encoded in the EMT

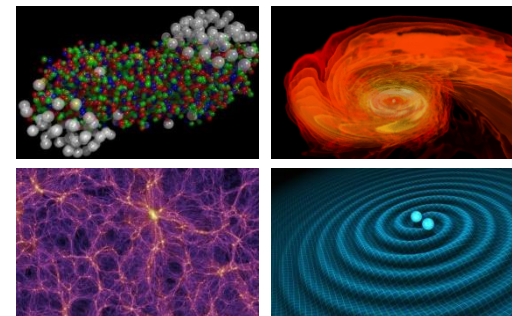
$$T^{\mu\nu} = \begin{bmatrix} \text{Energy density} & \text{Momentum density} & & \\ T^{00} & T^{01} & T^{02} & T^{03} \\ T^{10} & T^{11} & T^{12} & T^{13} \\ T^{20} & T^{21} & T^{22} & T^{23} \\ T^{30} & T^{31} & T^{32} & T^{33} \\ \text{Energy flux} & \text{Momentum flux} & & \end{bmatrix}$$

Shear stress
Normal stress (pressure)

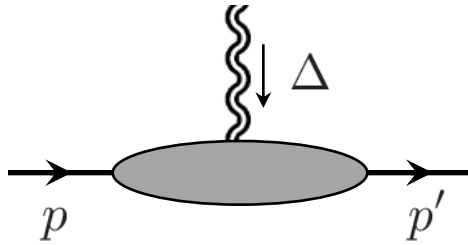


Central object for

- Nucleon mechanical properties
- Quark-gluon plasma
- Relativistic hydrodynamics
- Stellar structure and dynamics
- Cosmology
- Gravitational waves
- Modified theories of gravitation
- ...



Gravitational form factors (GFFs)



Poincaré symmetry constrains the form of the **EMT** matrix elements

Symmetrized variables $P = \frac{p' + p}{2}$, $\Delta = p' - p$, $t = \Delta^2$

$$p'^2 = p^2 = M^2 \Rightarrow \boxed{P \cdot \Delta = 0}$$

Spin-0 target $T^{\mu\nu} = \sum_a T_a^{\mu\nu}$

$$\langle p' | T_a^{\mu\nu}(0) | p \rangle = 2M \left[\frac{P^\mu P^\nu}{M} A_a(t) + \frac{\Delta^\mu \Delta^\nu - g^{\mu\nu} \Delta^2}{M} C_a(t) + M g^{\mu\nu} \bar{C}_a(t) \right]$$

Non-conserved

$$0 = \langle p' | \partial_\mu T^{\mu\nu}(x) | p \rangle = i \Delta_\mu \langle p' | T^{\mu\nu}(x) | p \rangle \Rightarrow \sum_a \bar{C}_a(t) = 0$$

Translation symmetry $\phi(x + a) = e^{iP \cdot a} \phi(x) e^{-iP \cdot a}$

Gravitational form factors (GFFs)

Spin-1/2 target

$$\langle p', s' | T_a^{\mu\nu}(0) | p, s \rangle = \bar{u}(p', s') \Gamma_a^{\mu\nu}(P, \Delta) u(p, s)$$

$$\Gamma_a^{\mu\nu}(P, \Delta) = \frac{P^\mu P^\nu}{M} A_a(t) + \frac{\Delta^\mu \Delta^\nu - g^{\mu\nu} \Delta^2}{M} C_a(t) + M g^{\mu\nu} \bar{C}_a(t) \\ + \frac{P^{\{\mu} i \sigma^{\nu\} \lambda} \Delta_\lambda}{2M} J_a(t) - \frac{P^{[\mu} i \sigma^{\nu] \lambda} \Delta_\lambda}{2M} S_a(t)$$

$$x^{\{\mu} y^{\nu\}} = x^\mu y^\nu + x^\nu y^\mu$$

$$x^{[\mu} y^{\nu]} = x^\mu y^\nu - x^\nu y^\mu$$

NB: Because of the Dirac equation, alternative but equivalent parametrizations may look quite different !

Gordon identity $\bar{u}(p', s') \gamma^\mu u(p, s) = \bar{u}(p', s') \left[\frac{P^\mu}{M} + \frac{i \sigma^{\mu\nu} \Delta_\nu}{2M} \right] u(p, s)$

Gravitational form factors (GFFs)

Spin- j target

$$\langle p', s' | T_a^{\mu\nu}(0) | p, s \rangle = \bar{\eta}(p', s') \Gamma_a^{\mu\nu}(P, \Delta) \eta(p, s)$$

$$\begin{aligned} \Gamma_a^{\mu\nu}(P, \Delta) = & \frac{P^\mu P^\nu}{M} A_a(t) + \frac{\Delta^\mu \Delta^\nu - g^{\mu\nu} \Delta^2}{M} C_a(t) + M g^{\mu\nu} \bar{C}_a(t) \\ & + \frac{i P^{\{\mu} \Sigma^{\nu\} \lambda} \Delta_\lambda}{2M} J_a(t) - \frac{i P^{[\mu} \Sigma^{\nu] \lambda} \Delta_\lambda}{2M} S_a(t) \\ & + \text{higher spin multipoles} \end{aligned}$$

Integer $j > 0$ $9j + 2$ GFFs

Half-integer $j > 1/2$ $9j - \frac{1}{2}$ GFFs

Same parametrization holds for hadrons, nuclei, ...

Poincaré sum rules

Four-momentum conservation

$$p^\mu = \sum_a \frac{\langle p, s | P_a^\mu | p, s \rangle}{\langle p, s | p, s \rangle} \Rightarrow \begin{cases} \sum_a A_a(0) = 1 \\ \sum_a \bar{C}_a(0) = 0 \end{cases}$$

LF momentum fraction $\langle x \rangle_a = \frac{\langle P_a^+ \rangle}{p^+} = A_a(0)$

Mechanical equilibrium !

$$\langle \int d^3r T_a^{\mu\nu} \rangle_{\text{rest}} = M \begin{pmatrix} A_a(0) + \bar{C}_a(0) & 0 & 0 & 0 \\ 0 & -\bar{C}_a(0) & 0 & 0 \\ 0 & 0 & -\bar{C}_a(0) & 0 \\ 0 & 0 & 0 & -\bar{C}_a(0) \end{pmatrix} \Leftrightarrow \begin{pmatrix} \varepsilon_a & 0 & 0 & 0 \\ 0 & p_a & 0 & 0 \\ 0 & 0 & p_a & 0 \\ 0 & 0 & 0 & p_a \end{pmatrix} V$$

Partial pressure

Angular momentum conservation

$$\langle J^i \rangle = \sum_a \frac{\langle p, s | \int d^3r \epsilon^{ijk} r^j \frac{1}{2} T_a^{\{0k\}} | p, s \rangle}{\langle p, s | p, s \rangle} \Rightarrow \sum_a J_a(0) = j$$

$$\partial_\mu (x^\alpha T^{\mu\beta} - x^\beta T^{\mu\alpha} + S^{\mu\alpha\beta}) = 0 \Rightarrow \begin{cases} S_q(t) = j G_A^q(t) \\ S_g(t) = 0 \end{cases}$$

Orbital AM **Intrinsic AM**

Let us recap

Spin-1/2 target

$$\langle p', s' | T_a^{\mu\nu}(0) | p, s \rangle = \bar{u}(p', s') \Gamma_a^{\mu\nu}(P, \Delta) u(p, s)$$

$$\Gamma_a^{\mu\nu}(P, \Delta) = \frac{P^\mu P^\nu}{M} A_a(t) + \frac{\Delta^\mu \Delta^\nu - g^{\mu\nu} \Delta^2}{M} C_a(t) + M g^{\mu\nu} \bar{C}_a(t) \\ + \frac{P^{\{\mu} i \sigma^{\nu\}\lambda} \Delta_\lambda}{2M} J_a(t) - \frac{P^{[\mu} i \sigma^{\nu]\lambda} \Delta_\lambda}{2M} S_a(t)$$

$A_a(0) \leftrightarrow$ **Momentum**

$\bar{C}_a(0) \leftrightarrow$ **Pressure**

$J_a(0) \leftrightarrow$ **Total angular momentum**

$S_q(0) \leftrightarrow$ **Intrinsic angular momentum**

$C_a(0) \leftrightarrow$ **?**

Also, what do we learn from the t -dependence ?

Spatial distributions

Phase-space approach

$$\langle \psi | O(x) | \psi \rangle = \int \frac{d^3 P}{(2\pi)^3} d^3 R \rho_\psi(\vec{R}, \vec{P}) \langle O \rangle_{\vec{R}, \vec{P}}(x)$$

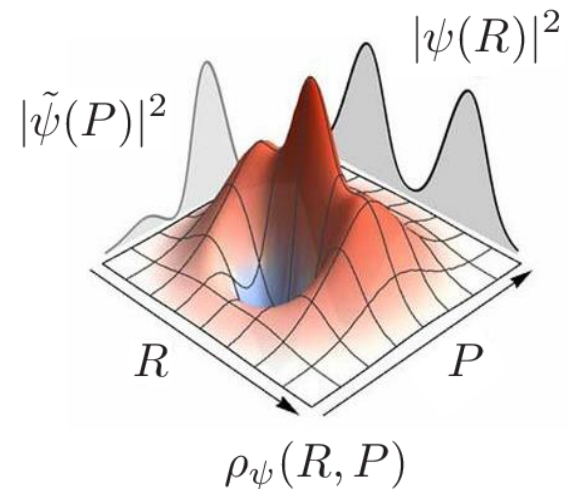
**Wigner
distribution**

$$\begin{aligned} \rho_\psi(\vec{R}, \vec{P}) &= \int d^3 z e^{-i\vec{P}\cdot\vec{z}} \psi^*\left(\vec{R} - \frac{\vec{z}}{2}\right) \psi\left(\vec{R} + \frac{\vec{z}}{2}\right) \\ &= \int \frac{d^3 q}{(2\pi)^3} e^{-i\vec{q}\cdot\vec{R}} \tilde{\psi}^*\left(\vec{P} + \frac{\vec{q}}{2}\right) \tilde{\psi}\left(\vec{P} - \frac{\vec{q}}{2}\right) \end{aligned}$$

$$\psi(\vec{r}) = \int \frac{d^3 p}{(2\pi)^3} e^{i\vec{p}\cdot\vec{r}} \tilde{\psi}(\vec{p})$$

Quasi-probabilistic interpretation

$$\begin{aligned} \int d^3 R \rho_\psi(\vec{R}, \vec{P}) &= |\tilde{\psi}(\vec{P})|^2 \\ \int \frac{d^3 P}{(2\pi)^3} \rho_\psi(\vec{R}, \vec{P}) &= |\psi(\vec{R})|^2 \end{aligned}$$



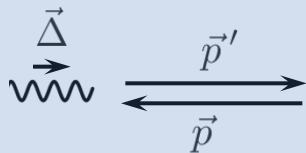
Spatial distributions

Internal distribution (for a state localized in phase-space)

$$\begin{aligned}\langle O \rangle_{\vec{R}, \vec{P}}(\vec{x}) &= \int \frac{d^3\Delta}{(2\pi)^3} e^{-i\vec{\Delta} \cdot (\vec{x} - \vec{R})} \frac{\langle P + \frac{\Delta}{2} | O(0) | P - \frac{\Delta}{2} \rangle}{2P^0} \\ &= \langle O \rangle_{\vec{0}, \vec{P}}(\vec{r}), \quad \vec{r} = \vec{x} - \vec{R}\end{aligned}$$

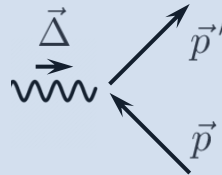
Elastic frames $\Delta^0 = \frac{\vec{P} \cdot \vec{\Delta}}{P^0} \stackrel{!}{=} 0$ (no energy transfer \rightarrow same initial and final boost factor)

$$|\vec{P}| = 0$$

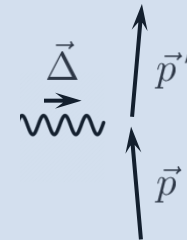


Breit (or brick-wall or rest) frame

$$|\vec{P}| \neq 0$$



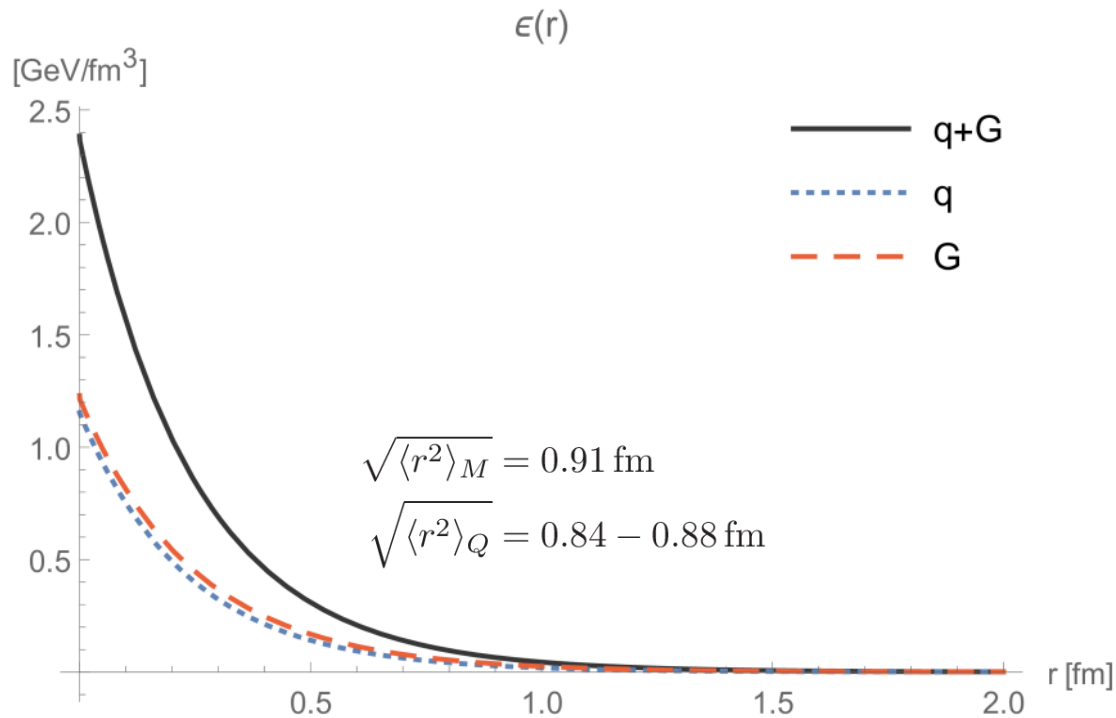
$$|\vec{P}| \gg M$$



Infinite-momentum frame (IMF)

Energy distribution (3D Breit frame)

$$\langle T^{00} \rangle_{\vec{0}, \vec{0}}(\vec{r}) = \int \frac{d^3 \Delta}{(2\pi)^3} e^{-i\vec{\Delta} \cdot \vec{r}} \frac{\langle P + \frac{\Delta}{2} | T^{00}(0) | P - \frac{\Delta}{2} \rangle}{2P^0} \Big|_{\vec{P}=\vec{0}}$$



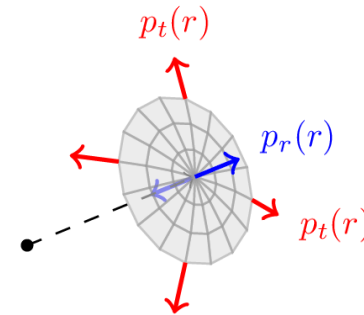
Multipole model for the gravitational form factors

$$F(t) = \frac{F(0)}{(1 + t/\Lambda^2)^n}$$

Pressure distributions (3D Breit frame)

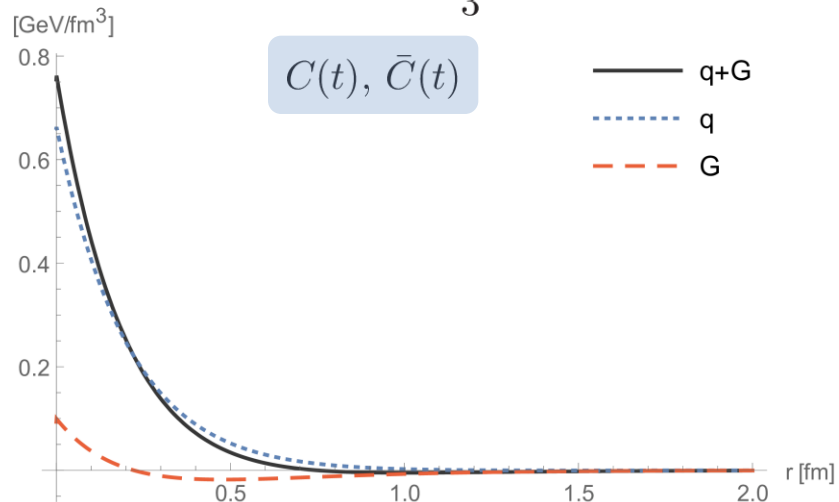
$$\langle T^{ij} \rangle_{\vec{0}, \vec{0}}(\vec{r}) = \int \frac{d^3 \Delta}{(2\pi)^3} e^{-i\vec{\Delta} \cdot \vec{r}} \frac{\langle P + \frac{\Delta}{2} | T^{ij}(0) | P - \frac{\Delta}{2} \rangle}{2P^0} \Big|_{\vec{P}=\vec{0}}$$

$$= \delta^{ij} p(r) + \left(\frac{r^i r^j}{r^2} - \frac{1}{3} \delta^{ij} \right) s(r)$$



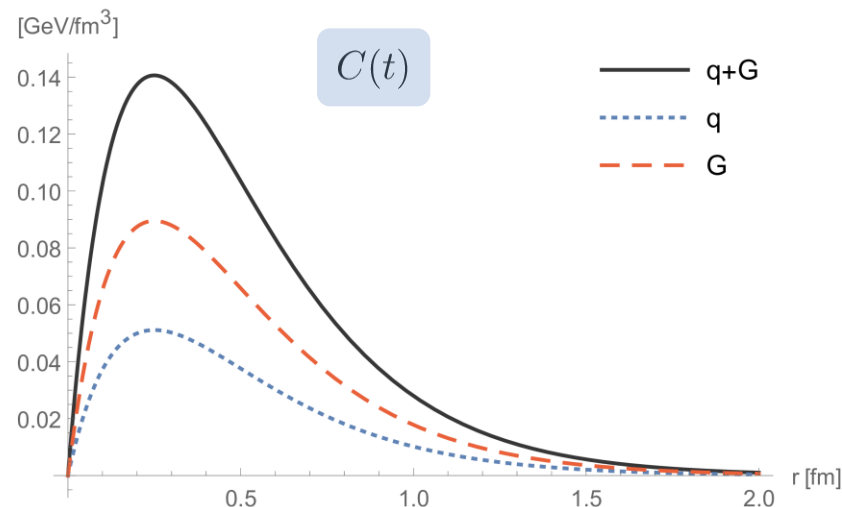
Isotropic pressure

$$p(r) = \frac{p_r(r) + 2p_t(r)}{3}$$



Pressure anisotropy

$$s(r) = p_r(r) - p_t(r)$$



How to access GFFs ?

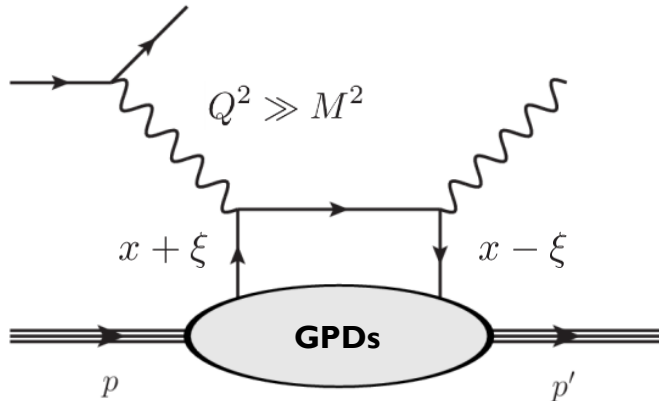
Measuring *directly* GFFs is currently not realistic ...

... but a spin-2 exchange can be mimicked by two spin-1 exchanges !

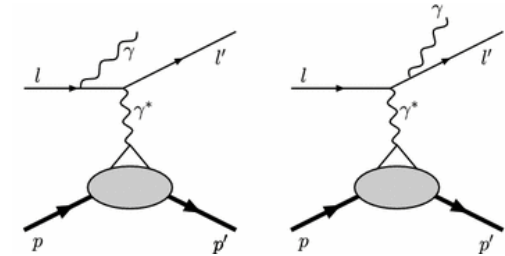
Target must remain intact  exclusive reaction

Generalized PDFs

Deeply virtual Compton scattering (DVCS)



interferes with



Bethe-Heitler

Correlator (in $A^+ = 0$ gauge)

$$\frac{1}{2} \int \frac{dz^-}{2\pi} e^{ixP^+z^-} \langle p', s' | \bar{\psi}(-\frac{z}{2}) \gamma^+ \psi(\frac{z}{2}) | p, s \rangle \Big|_{z^+ = z_\perp = 0}$$

$$= \frac{1}{2P^+} \bar{u}(p', s') \left[\gamma^+ H_q(x, \xi, t) + \frac{i\sigma^{+\lambda} \Delta_\lambda}{2M} E_q(x, \xi, t) \right] u(p, s)$$

Momentum transfer

$$\xi = -\frac{\Delta^+}{2P^+} = \frac{p^+ - p'^+}{p^+ + p'^+}, \quad t = \Delta^2 = (p' - p)^2$$

Generalized PDFs

Link with other non-perturbative functions

$$H_q(x, 0, 0) = f_q(x)$$

PDF

$$\int dx H_q(x, \xi, t) = F_1^q(t)$$

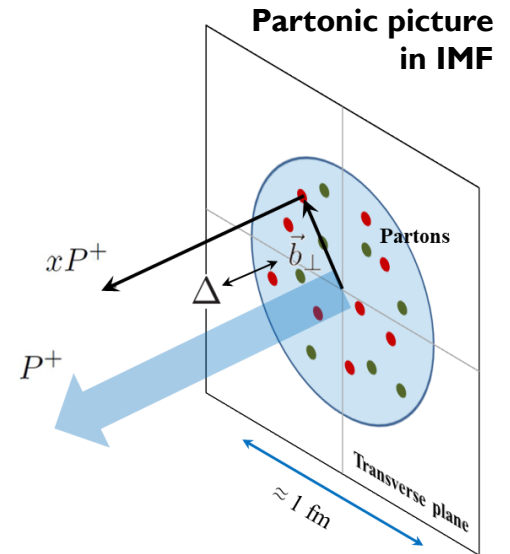
$$\int dx E_q(x, \xi, t) = F_2^q(t)$$

Electromagnetic
form factors

2+1D imaging (in $A^+ = 0$ gauge)

$$\rho(x, \vec{b}_\perp) = P^+ \int \frac{d^2 \Delta_\perp}{(2\pi)^2} e^{-i\vec{\Delta}_\perp \cdot \vec{b}_\perp} \frac{\langle p' | j^+(xP^+) | p \rangle}{2P^+} \Big|_{\Delta^+ = 0}$$

$$j^\mu(k^+) = \int \frac{dz^-}{2\pi} e^{ik^+ z^-} \bar{\psi}\left(-\frac{z}{2}\right) \gamma^\mu \psi\left(\frac{z}{2}\right) \Big|_{z^+ = z_\perp = 0}$$



Generalized PDFs

Link with gravitational form factors

$$T_q^{++}(0) = (\bar{\psi} \gamma^+ \frac{i}{2} \overleftrightarrow{D}^+ \psi)(0)$$

$$= 2(P^+)^2 \int dx x \frac{1}{2} \int \frac{dz^-}{2\pi} e^{ixP^+z^-} \bar{\psi}(-\frac{z}{2}) \gamma^+ \psi(\frac{z}{2}) \Big|_{z^+=z_\perp=0}$$

Operator entering GPD correlator !



$$\int dx x H_q(x, \xi, t) = A_q(t) + 4\xi^2 C_q(t)$$
$$\int dx x \frac{1}{2} [H_q + E_q](x, \xi, t) = J_q(t)$$

A similar reasoning applies to the gluonic sector

Conclusions

- The **energy-momentum tensor** contains key information about the system internal structure
- **Poincaré symmetry** constrains the structure of the **EMT** matrix elements in terms of gravitational form factors
- **Spatial distributions** of energy, pressure, ... can be expressed in terms **Fourier transforms** of the gravitational form factors
- **Generalized parton distributions** are off-forward extensions of the usual **PDFs** and can be measured in e.g. **DVCS**
- **x-moments** of **GPDs** give access to electromagnetic form factors and to some of the gravitational form factors

Some references

- **Diehl, Phys. Rep. 388 (2003) 41**
- **Leader, Lorcé, Phys. Rep. 541 (2014) 3, 163**
- **Polyakov, Schweitzer, IJMPA33 (2018) 26, 1830025**
- **Lorcé, Moutarde, Trawinski, Eur. Phys. J. C79 (2019) 89**

... and references therein !