

Anomalous Magnetic Moment of the Muon in Quantum Electrodynamics

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Contents

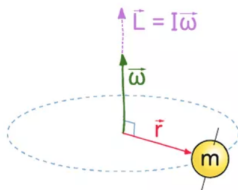
- ➊ Magnetic moment of charged lepton
- ➋ Leading order result
- ➌ D-dimensional regularization
- ➍ Renormalization
 - Renormalization
 - Renormalized functions
- ➎ NLO result
- ➏ Conclusion

Classical result

Let's assume that a charged lepton like a muon as a charged rigid body rotates around the z – axis with an angular velocity.

The magnetic moment can be obtain as

$$\vec{M} = \frac{e}{2m} \vec{L}. \quad (1)$$



Classical result

From here, Dirac made his prediction from the Dirac equation in the electromagnetic field.

$$(i\gamma^\mu \partial_\mu - m)\psi - e\gamma^\mu A_\mu \psi = 0. \quad (2)$$

Following by the non-relativistic Schrödinger-Pauli Hamiltonian equation that can explain the interaction of spin one-half particles:

$$H = \frac{1}{2m}(\vec{p} - e\vec{A})^2 - \frac{e}{m} \vec{S} \cdot \vec{B} + e\phi, \quad (3)$$

Taking out the magnetic moment component:

$$\vec{M} = \frac{e}{m} \vec{S} = g \frac{e}{2m} \vec{S} \quad (4)$$

where g is called the Landé g -factor.

Dirac equation's prediction

Compared to the classical result g must equal to 2. This holds true not only for muons but for any one-half spin elementary particles.

In higher-order correction, however, QED predicts that $g = 2 + \mathcal{O}(\alpha)$. The anomalous magnetic moment we are looking for refers to this small correction.

$$a = \frac{g - 2}{2}. \quad (5)$$

Anomalous magnetic moment from scattering amplitude

We will calculate the muon anomalous magnetic moment through the scattering matrix (**S-matrix**) of an imaginary process:

$$\mu^-(p_1, s_1) \longrightarrow \mu^-(p_2, s_2) + \gamma(p_3, \lambda_3), \quad (6)$$

The QED Lagrangian density of this system is given by

$$\begin{aligned} \mathcal{L} &= \mathcal{L}_{Dirac} + \mathcal{L}_{photon} + \mathcal{L}_{int} \\ &= \bar{\psi}(i\gamma^\mu \partial_\mu - m)\psi - \frac{1}{4}F^{\mu\nu}F_{\mu\nu} - \textcolor{red}{e}\bar{\psi}\gamma^\mu\psi\textcolor{red}{A}_\mu, \end{aligned} \quad (7)$$

Definition of the initial state and the final state is

$$|i\rangle = \sqrt{2E_1}a_{p_1}^{s_1\dagger}|0\rangle, \quad |f\rangle = \sqrt{4E_2E_3}a_{p_2}^{s_2\dagger}c_{p_3}^{\lambda_3\dagger}|0\rangle. \quad (8)$$

S-matrix

The scattering amplitude is defined

$$A = \langle f|S|i\rangle, \quad (9)$$

where the total scattering amplitude $A = A_0 + A_1 + A_2 + \dots$ is the sum of all leading terms in each order of perturbation theory, and the scattering matrix S can be expanded to Dyson's series as follows:

$$S = \sum_{n=0}^{\infty} \frac{i^n}{n!} \int_{-\infty}^{\infty} d^4x_1 \dots \int_{-\infty}^{\infty} d^4x_n T\{\mathcal{L}_{int}(x_1) \dots \mathcal{L}_{int}(x_n)\}. \quad (10)$$

Form factors

We expand the S-matrix and only keep the orders that have contribution to the scattering matrix, from which we can derive the muon's anomalous magnetic moment in QED at NLO.

$$\langle f|S|i\rangle = -ie\bar{u}(p_2)\left\{\gamma^\mu F_1(q^2) + \frac{i\sigma^{\mu\nu}}{2m}(p_1 - p_2)_\nu F_2(q^2)\right\}u(p_1)\epsilon_\mu^*, \quad (11)$$

where form factor F_1 and factor F_2 contain the influence of the EM field on the muon.

The g -factor now is defined

$$g = 2[F_1(0) + F_2(0)]. \quad (12)$$

The leading 2 is prediction of Dirac equation that we proved above.

⇒ **Find these form factors ?**

Leading order

The Feynman diagram corresponds to LO:

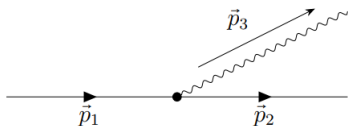


Figure 1: Tree-level Feynman diagram.

The scattering amplitude at LO can be obtained quickly as

$$A_1 = \bar{u}(p_2)(-ie\gamma^\mu)u(p_1)\epsilon_\mu^*. \quad (13)$$

At the LO, $F_1 = 1$ and $F_2 = 0$ followed by g -factor is equal to 2 \rightarrow satisfies the classical result.

The anomalous magnetic moment at LO:

$$a_\mu = \frac{g_\mu - 2}{2} = \frac{2(1 + 0) - 2}{2} = 0. \quad (14)$$

Next-to-leading order

The Feynman diagrams correspond to NLO:

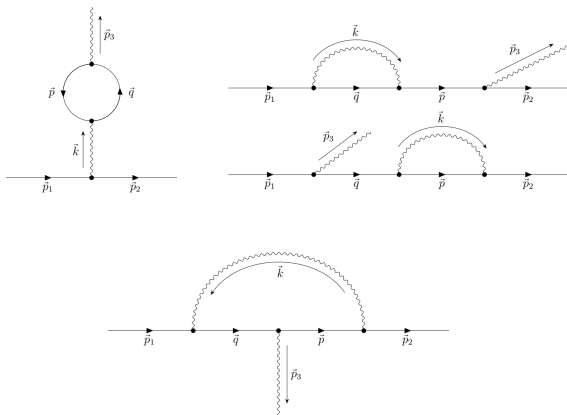


Figure 2: NLO connected Feynman diagrams.

Amputated Feynman diagram

Due to the on-shell condition, we only keep the amputated diagram.

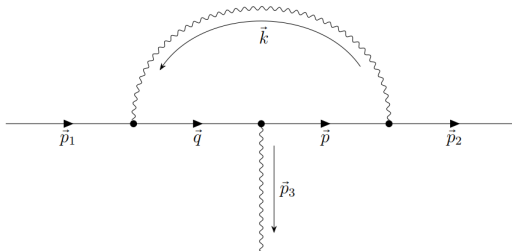


Figure 3: Amputated vertex correction.

$$i\mathcal{M}_v = -e^3 \bar{u}_{p_2} \int \frac{d^4 k}{(2\pi)^4} \frac{\gamma^\nu (\not{p}_2 + \not{k} + m) \gamma^\mu (\not{p}_1 + \not{k} + m) \gamma_\nu}{(k^2 - \lambda^2 + i\epsilon)[(p_1 + k)^2 - m^2 + i\epsilon][(p_2 + k)^2 - m^2 + i\epsilon]} u_{p_1} \epsilon_\mu^*. \quad (15)$$

UV and IR-divergence

This integral diverges at

- $x \rightarrow \infty$: UV-divergence.
 \Rightarrow We need to use the dimensional regularization and renormalization method to cancel UV-divergence.
- $x \rightarrow 0$: IR-divergence.
 \Rightarrow Cancelled thanks to the on-shell condition.

D-dimensional regularization

Calculation of loop integrals in $D = 4 - 2\epsilon$ dimension with ϵ is called dimensional regularization parameter.

$$\int \frac{d^4 q}{(2\pi)^4} = \mu^{4-D} \int \frac{d^D q}{(2\pi)^D}. \quad (16)$$

where the renormalization scale μ has dimension of energy.

Conventions

Conventions		
Space	4-Dimension	D-Dimension
Metric Tensor $g^{\mu\nu}$	$\mu, \nu = 0, 1, 2, 3$ $g_{\mu}^{\mu} = \delta_{\mu}^{\mu} = 4$	$\mu, \nu = 0, 1, \dots, D - 1$ $g_{\mu}^{\mu} = \delta_{\mu}^{\mu} = D$
Dirac Matrices	$\dim(\gamma^{\mu})=4$ $\gamma^{\mu}\gamma_{\mu} = 4 \cdot \mathbf{I}_4$ $\gamma^{\mu}\gamma_{\nu}\gamma_{\mu} = -2\gamma_{\nu}$ $\{\gamma^{\mu}, \gamma^{\nu}\} = 2g^{\mu\nu} \cdot \mathbf{I}_4$ $Tr(\mathbf{I}_4) = 4$	$\dim(\gamma^{\mu})=2^{D/2}$ $\gamma^{\mu}\gamma_{\mu} = D \cdot \mathbf{I}_D$ $\gamma^{\mu}\gamma_{\nu}\gamma_{\mu} = (2 - D)\gamma_{\nu}$ $\{\gamma^{\mu}, \gamma^{\nu}\} = 2g^{\mu\nu} \cdot \mathbf{I}_D$ $Tr(\mathbf{I}_D) = 4$ (per definition)
Properties	Preserve Lorentz invariant, Gauge invariant	

Table 1: Conventions in D-dimension compared to 4-dimension.

Counterterms

The Lagrangian that we used at the tree level is not sufficient to describe the divergences that appear in the Feynman diagrams, so we need a tiny impact to modify the Lagrangian with quantities called "bare quantities" (denoted by subscript 0).

$$\mathcal{L}_0 = -\frac{1}{4}(\partial^\mu A_0^\nu - \partial^\nu A_0^\mu)^2 + \bar{\psi}_0(i \not{\partial} - m_0)\psi_0 - e_0 \bar{\psi}_0 \gamma^\mu \psi_0 A_{0\mu}. \quad (17)$$

The bare quantities are conventionally associated with the finite renormalized quantities by renormalization parameters Z_i as

$$\psi_0 = \sqrt{Z_2}\psi \equiv (1 + \delta_2)\psi, \quad A_0^\mu = \sqrt{Z_3}A^\mu \equiv (1 + \delta_3)A^\mu. \quad (18)$$

$$m_0 = Z_m m \equiv m + \delta_m, \quad e_0 = Z_e e \equiv e + \delta_e. \quad (19)$$

Counterterm Lagrangian

Applying these expansions into the bare Lagrangian, we derive the Lagrangian for renormalized perturbation theory:

$$\begin{aligned}
 \mathcal{L}_0 &= -\frac{1}{4}F^{\mu\nu}F_{\mu\nu} + \bar{\psi}(i\not{\partial} - m)\psi - e\bar{\psi}\gamma^\mu\psi A_\mu \\
 &\quad -\frac{1}{4}\delta_3 F^{\mu\nu}F_{\mu\nu} + \delta_2 \bar{\psi}(i\not{\partial} - m)\psi - \delta_m \bar{\psi}\psi - (\delta_e + \delta_2 + \frac{1}{2}\delta_3)e\bar{\psi}\gamma^\mu\psi A_\mu \\
 &= \mathcal{L}_R + \delta\mathcal{L}.
 \end{aligned}
 \tag{20}$$

- By rescaling the Lagrangian density, we can split the Lagrangian density into two terms, which are the renormalized Lagrangian (\mathcal{L}_R) and the counterterm Lagrangian ($\delta\mathcal{L}$).
- The renormalization parameters are fixed by the on-shell renormalization condition ($p^2 = m^2$).

Feynman rules

Feynman rules for counterterm vertex:

$$\mu \sim \text{wavy line} \text{---} \bigotimes \text{---} \text{wavy line} \nu = -i\delta_3 k^2 g^{\mu\nu}.$$

Figure 4: Photon counterterm vertex.

$$\text{---} \rightarrow \text{---} \bigotimes \text{---} \rightarrow \text{---} = i\delta_2 (\not{p} - m) - i\delta_m.$$

Figure 5: Fermion counterterm vertex.

$$\text{---} \swarrow \bigotimes \searrow \text{---} \text{---} \uparrow \text{wavy line} = -ie\gamma^\mu (\delta_2 + \delta_e + \tfrac{1}{2}\delta_3).$$

Figure 6: Photon-fermion counterterm vertex.

Renormalized functions

Let's go to the renormalized vertex function (denoted by a hat) definition:

$$\begin{aligned}
 \mu \text{ wavy line} \text{---} \text{circle} \text{---} \nu \text{ wavy line} &= \hat{\Gamma}^{\mu\nu}(k) \\
 &= \Gamma^{\mu\nu}(k) + \delta\Gamma^{\mu\nu}(k) = -ik^2 g^{\mu\nu} + \underbrace{i\Pi^{\mu\nu}(k) - ik^2 g^{\mu\nu} \delta_3}_{i\hat{\Pi}^{\mu\nu}(k)}.
 \end{aligned}$$

$$\begin{aligned}
 \text{---} \rightarrow \text{circle} \rightarrow \text{---} &= \hat{\Gamma}^{ff}(p) \\
 &= \Gamma^{ff}(p) + \delta\Gamma(p) = i(\not{p} - m) \underbrace{-i\Sigma(p) + i\delta_2(\not{p} - m) - i\delta_m}_{-i\hat{\Sigma}(p)}.
 \end{aligned}$$

$$\begin{aligned}
 \mu \text{ wavy line} \text{---} \text{circle} \begin{array}{l} \nearrow \text{---} \\ \searrow \text{---} \end{array} &= \hat{\Gamma}^{Aff}(p, p') \\
 &= \Gamma^{Aff}(p, p') + \delta\Gamma^{Aff}(p, p') = \underbrace{-ie\gamma^\mu - ie\Lambda^\mu(p, p') - ie\gamma^\mu(\delta_e + \delta_2 + \frac{1}{2}\delta_3)}_{-ie\hat{\Lambda}^\mu(p, p')}.
 \end{aligned}$$

Renormalization conditions

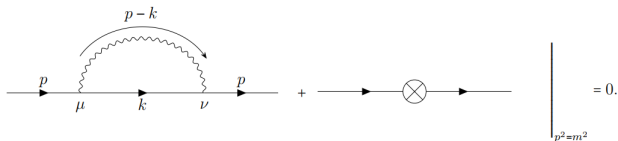
By imposing renormalization conditions, the renormalization constants are fixed to cancel out the UV-divergence.

- Condition 1 (Dirac equation):

$$\hat{\Sigma}(p)u(p) = 0$$

$$\Rightarrow \frac{\delta m}{m} = -\Sigma_V(m^2) - \Sigma_S(m^2). \quad (21)$$

- Condition 2:



$$\left[\lim_{p^2 \rightarrow m^2} \frac{\not{p} + m}{p^2 - m^2} \hat{\Sigma}(p) \right] u(p) \stackrel{!}{=} 0$$

$$\Rightarrow \delta_2 = \Sigma_V(m^2) + 2m^2(\Sigma_V'(m^2) + \Sigma_S'(m^2)). \quad (22)$$

Renormalization conditions

Condition 3:

$$\left. \mu \xrightarrow{k} \text{loop} \xrightarrow{k} \nu + \mu \text{---} \bigotimes \text{---} \nu \right|_{p^2=m^2} = 0.$$

$$\left[\lim_{k^2 \rightarrow 0} \frac{-ig_{\nu\rho}}{k^2} \hat{\Pi}^{\mu\nu}(k) \right] \epsilon_\mu(k) \stackrel{!}{=} 0$$

$$\Rightarrow \delta_3 = \Pi'_T(0). \quad (23)$$

Condition 4:

$$\left. \mu \xrightarrow{p_3} \text{vertex} \xrightarrow{p} \text{loop} \xrightarrow{p_2, q} + \mu \text{---} \bigotimes \text{---} \text{vertex} \xrightarrow{p_1, p_2} \right|_{p^2=m^2} = 0.$$

$$\bar{u}(p) \hat{\Lambda}^\mu(p, p) u(p) = 0$$

$$\Rightarrow \delta_e = -\frac{1}{2} \delta_3. \quad (24)$$

Summary

Renormalization procedure

Define the new parameters → Separate the bare parameters into renormalized parameters and counterterm parameters → Choose renormalization conditions to fix the counterterms.

Vertex correction

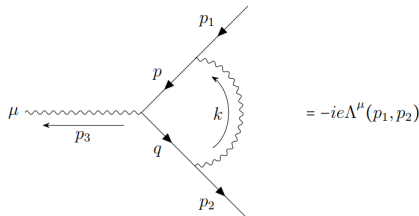


Figure 7: Vertex correction.

$$\begin{aligned}
 & \Lambda^\mu(p_1, p_2) \\
 &= -ie^2 \mu^{4-D} \int \frac{d^D k}{(2\pi)^D} \frac{\gamma^\nu (\not{p}_2 + \not{k} + m) \gamma^\mu (\not{p}_1 + \not{k} + m) \gamma_\nu}{(k^2 - \lambda^2 + i\epsilon)[(p_1 + k)^2 - m^2 + i\epsilon][(p_2 + k)^2 - m^2 + i\epsilon]} \\
 &= \frac{e^2}{(4\pi)^2} \frac{(2\pi\mu)^{4-D}}{i\pi^2} \int d^D k \frac{\gamma^\nu (\not{p}_2 + \not{k} + m) \gamma^\mu (\not{p}_1 + \not{k} + m) \gamma_\nu}{(k^2 - \lambda^2 + i\epsilon)[(p_1 + k)^2 - m^2 + i\epsilon][(p_2 + k)^2 - m^2 + i\epsilon]}.
 \end{aligned}
 \tag{25}$$

Let's consider this sophisticated numerator with gamma matrices properties in D-dimension:

$$\gamma^\nu \gamma^\mu \gamma_\nu = (2 - D) \gamma^\mu, \quad (26)$$

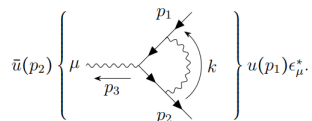
$$\gamma^\nu \gamma^\mu \gamma^\rho \gamma_\nu = 4g^{\mu\rho} + (D - 4) \gamma^\mu \gamma^\rho, \quad (27)$$

$$\gamma^\nu \gamma^\rho \gamma^\mu \gamma^\sigma \gamma_\nu = -2\gamma^\sigma \gamma^\mu \gamma^\rho - (D - 4) \gamma^\rho \gamma^\mu \gamma^\sigma. \quad (28)$$

Expanding the numerator, we got

$$\begin{aligned} \mathcal{N}^\mu &= \gamma^\nu \not{p}_2 \gamma^\mu \not{p}_1 \gamma_\nu + \gamma^\nu \not{p}_2 \gamma^\mu \not{k} \gamma_\nu + m \gamma^\nu \not{p}_2 \gamma^\mu \gamma_\nu \\ &\quad + \gamma^\nu \not{k} \gamma^\mu \not{p}_1 \gamma_\nu + \gamma^\nu \not{k} \gamma^\mu \not{k} \gamma_\nu + m \gamma^\nu \not{k} \gamma^\mu \gamma_\nu \\ &\quad + m \gamma^\nu \gamma^\mu \not{p}_1 \gamma_\nu + m \gamma^\nu \gamma^\mu \not{k} \gamma_\nu + m^2 \gamma^\nu \gamma^\mu \gamma_\nu \\ &= -2(\not{p}_1 \gamma^\mu \not{p}_2 + \not{k} \gamma^\mu \not{p}_2 + \not{p}_1 \gamma^\mu \not{k}) - (D - 4)(\not{p}_2 \gamma^\mu \not{p}_1 + \not{p}_2 \gamma^\mu \not{k} + \not{k} \gamma^\mu \not{p}_1) \\ &\quad + 4m(p_1 + p_2 + 2k)^\mu + (D - 4)m(\not{p}_2 \gamma^\mu + \not{k} \gamma^\mu + \gamma^\mu \not{p}_1 + \gamma^\mu \not{k}) \\ &\quad + (2 - D) \not{k} \gamma^\mu \not{k} + (2 - D)m^2 \gamma^\mu. \end{aligned} \quad (29)$$

Consequently, sandwich the vertex function:



$$\begin{aligned}
 & \bar{u}(p_2) (-ie\Lambda^\mu(p_1, p_2)) u(p_1) \epsilon_\mu^* \\
 = & \bar{u}(p_2) (-ie\gamma^\mu) u(p_1) \epsilon_\mu^* \frac{\alpha}{4\pi} \left\{ 4B_0(m^2, 0, m) - 3B_0(p_3^2, m, m) + (4m^2 - 2p_3^2)C_0 - 2 \right\} \\
 & + \bar{u}(p_2) (-ie) u(p_1) \epsilon_\mu^* (p_1 + p_2)^\mu \frac{2m}{4m^2 - k^2} \frac{\alpha}{4\pi} \left\{ 2B_0(p_3^2, m, m) - 2B_0(m^2, 0, m) + 2 \right\}.
 \end{aligned} \tag{30}$$

\Rightarrow Of course, this is not UV-convergence \rightarrow we need to add the counterterm diagram.

Next-to-leading order result

Applying $p_3 = 0$ due to on-shell condition to renormalized vertex function:

$$\begin{aligned}
 & \bar{u}(p_2) \hat{\Gamma}^{Aff}(p_3, -p_1, p_2) \Big|_{p_3=0} u(p_1) \epsilon_\mu^* \\
 &= -ie \bar{u}(p_2) \hat{\Lambda}^\mu(p_1, p_2) \Big|_{p_3 \rightarrow 0} u(p_1) \epsilon_\mu^* \\
 &= -ie \bar{u}(p_2) \Lambda^\mu(p_1, p_2) \Big|_{p_3 \rightarrow 0} u(p_1) \epsilon_\mu^* - ie \bar{u}(p_2) \gamma^\mu \delta_2 u(p_1) \epsilon_\mu^* \\
 &= -ie \bar{u}(p_2) \frac{i\sigma^{\mu\nu}}{2m} (p_1 - p_2)_\nu u(p_1) \epsilon_\mu^* \frac{\alpha}{4\pi} \left\{ \underbrace{4B_0(m^2, 0, m) - 4B_0(0, m, m) - 6}_{=2} \right\} \\
 &= -ie \bar{u}(p_2) \left\{ \frac{i\sigma^{\mu\nu}}{2m} (p_1 - p_2)_\nu \frac{\alpha}{2\pi} \right\} u(p_1) \epsilon_\mu^*.
 \end{aligned} \tag{31}$$

Anomalous magnetic moment at NLO

Total scattering amplitude at NLO:

$$\begin{aligned}
 A^{\text{NLO}} &= A_1 + A_3 \\
 &= -ie\bar{u}(p_2) \left\{ \gamma^\mu 1 + \frac{i\sigma^{\mu\nu}}{2m} (p_1 - p_2)_\nu \frac{\alpha}{2\pi} \right\} u(p_1) \epsilon_\mu^*.
 \end{aligned} \tag{32}$$

At NLO, $F_1 = 1$ and $F_2 = \alpha/2\pi$, then the Landé g -factor that matches the QED prediction as

$$g_\mu = 2 + \frac{\alpha}{\pi} + \mathcal{O}(\alpha^2). \tag{33}$$

Anomalous magnetic moment of muon at NLO

$$a_\mu = \frac{g_\mu - 2}{2} = \frac{2F_2(0)}{2} = \frac{\alpha}{2\pi} \approx 0.0011614. \tag{34}$$

Conclusions

Here are some results of the anomalous magnetic moment of the muon:

$$\begin{aligned}
 a_{\mu}^{\text{LO}} &= 0, & (\text{LO}) \\
 a_{\mu}^{\text{NLO}} &= \frac{\alpha}{2\pi} \approx 116140973 \times 10^{-11}, & (\text{NLO}) \\
 a_{\mu}^{\text{BNL}} &= 116592089(54)(33) \times 10^{-11}, & (\text{experimental value}) \\
 a_{\mu}^{\text{FNAL}} &= 116592061(41) \times 10^{-11}. & (\text{experimental value})
 \end{aligned} \tag{35}$$

Where the fine-structure constant $\alpha \approx 1/137.035999084$ is available on Particle Data Group.

a_{μ}^{BNL} is the result from E821 experiment at Brookhaven National Lab (BNL).

a_{μ}^{FNAL} is the result from the Fermilab National Accelerator Laboratory (FNAL).

Higher-order correction

The anomalous magnetic moment in QED has been computed up to five loops

$$a_{\mu}^{\text{QED}1} = 0.5 \frac{\alpha}{\pi} + 0.765857420 \left(\frac{\alpha}{\pi} \right)^2 + 24.05050985(23) \left(\frac{\alpha}{\pi} \right)^3 \\ + 130.8785(60) \left(\frac{\alpha}{\pi} \right)^4 + 751.0(9) \left(\frac{\alpha}{\pi} \right)^5 + \dots \quad (36)$$

The anomalous magnetic moment in higher-order correction:

$$a^{NNLO} = 116554190 \times 10^{-11}, \\ a^{N^3LO} = 116584332 \times 10^{-11}, \\ a^{N^4LO} = 116584713 \times 10^{-11}, \\ a^{N^5LO} = 116584718 \times 10^{-11}. \quad (37)$$

¹hep-ph/0507249,arXiv:0706.3496v2,hep-ph/0512330

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Thank you for your listening!