

A 3D architectural rendering of the LSST (Large Synoptic Survey Telescope) observatory building, shown in white and grey, situated on a mountain ridge. The building has a complex, multi-level design with a large, dark, circular opening at the top. The background features a vast, hazy mountain range under a clear sky. The foreground shows a dirt road and a small, white, dome-shaped structure on the right.

# GRB detection with LSST

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# The Large Synoptic Survey Telescope

- 8.4m telescope 9.6 degree<sup>2</sup> field-of-view
- Survey the south hemisphere of the sky every 3-4 nights

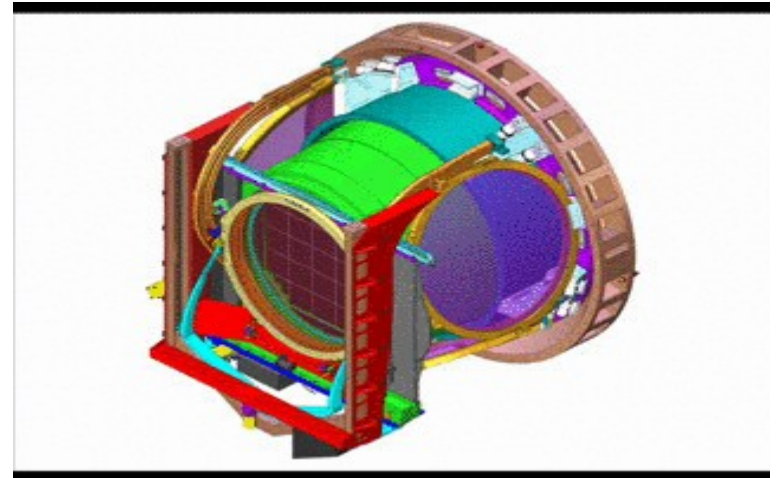
## The LSST Surveys :

### Wide-fast-deep (WFD)

- 18,000 sq deg
- A given object is observed in a given filter once every three days (up to redshift  $z \sim 0.4$ )

### Several Deep Drilling Fields (DDF)

- Each of 10 sq deg
- A given object is observed in a given filter several
- times during one night ( $0.4 < z < 1.2$ )



# Operations simulation (OpSim)

## Baseline cadence:

- Each field will be seen twice a night (2 visits) separated by 15-20 mins
- With subsequent revisits days later
- 825 visits per field over 10 years  
(u :54, g :70, r :178, i :180, z :160, y :166)

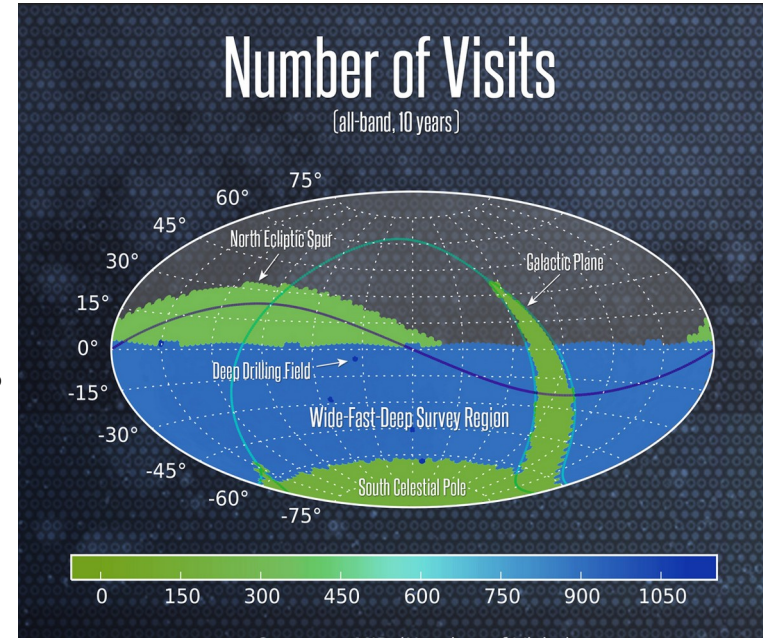
**OpSim** : simulates the field selection and image acquisition process of the LSST over the 10-year life of the planned survey

## The input of OpSim:

A metric defining the strategy parameters → area, revisit rate, observation time, ...

The **Metrics Analysis Framework (MAF)** : a code package that enables the derivation of scientific results from the *OpSim* database

➔ measurement of scientific performance, number of detections of an event



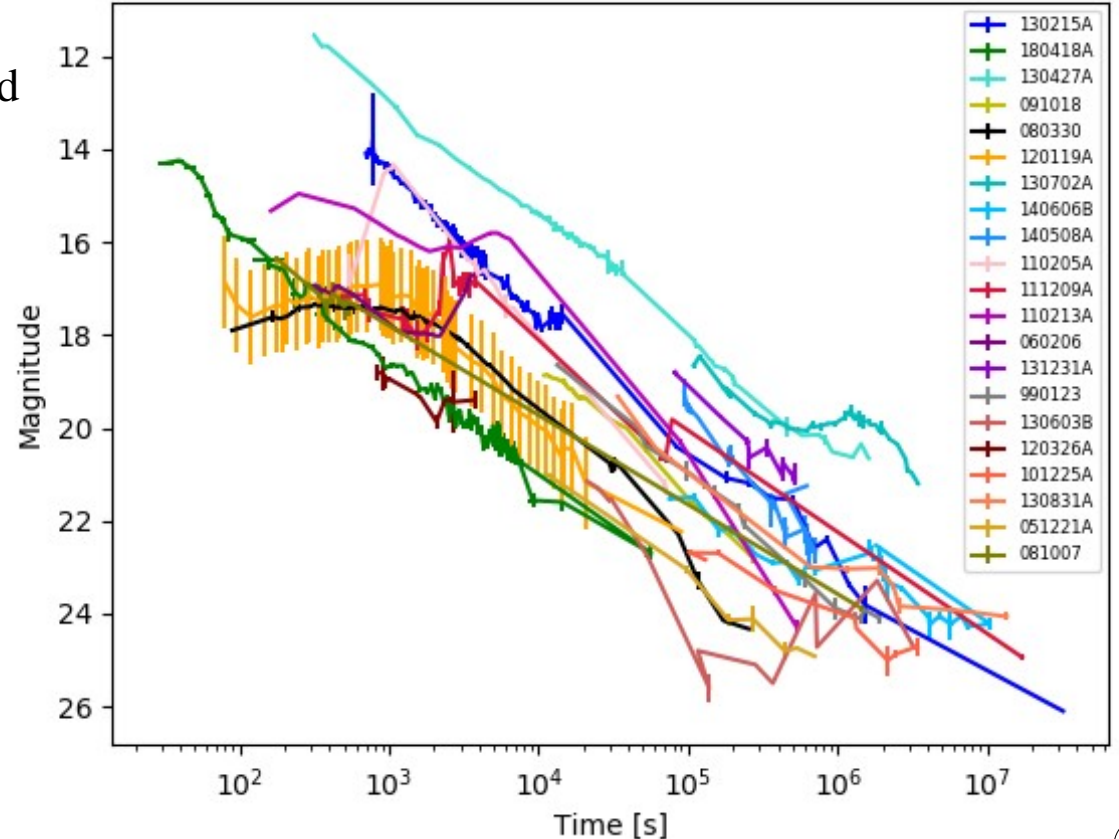


# Catalog of optical GRBs

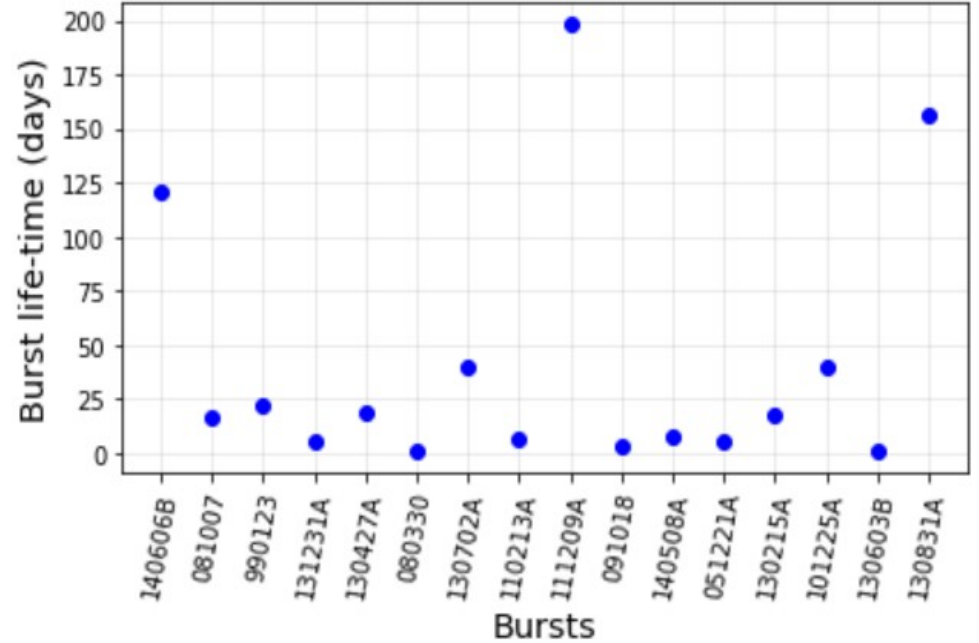
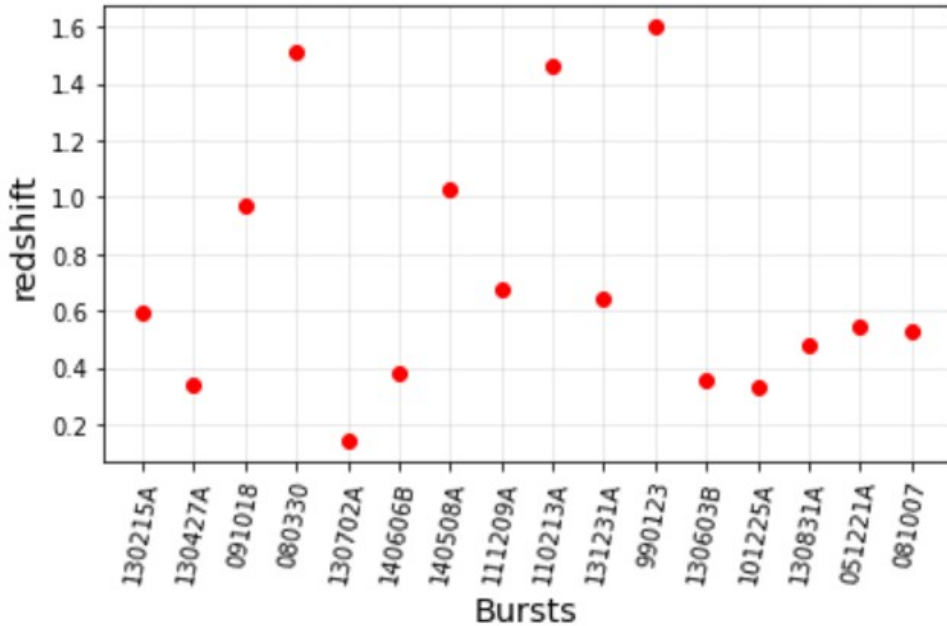
- A catalog of 549 LCs of GRBs detected by optical telescopes in r and R bands (provided by D.Turpin)

**Available information :** time, magnitudes and their errors

- Only 21 GRBs have sampled LCs in ‘r’
- 16 GRBs among 21 have a life-time longer than 1 day

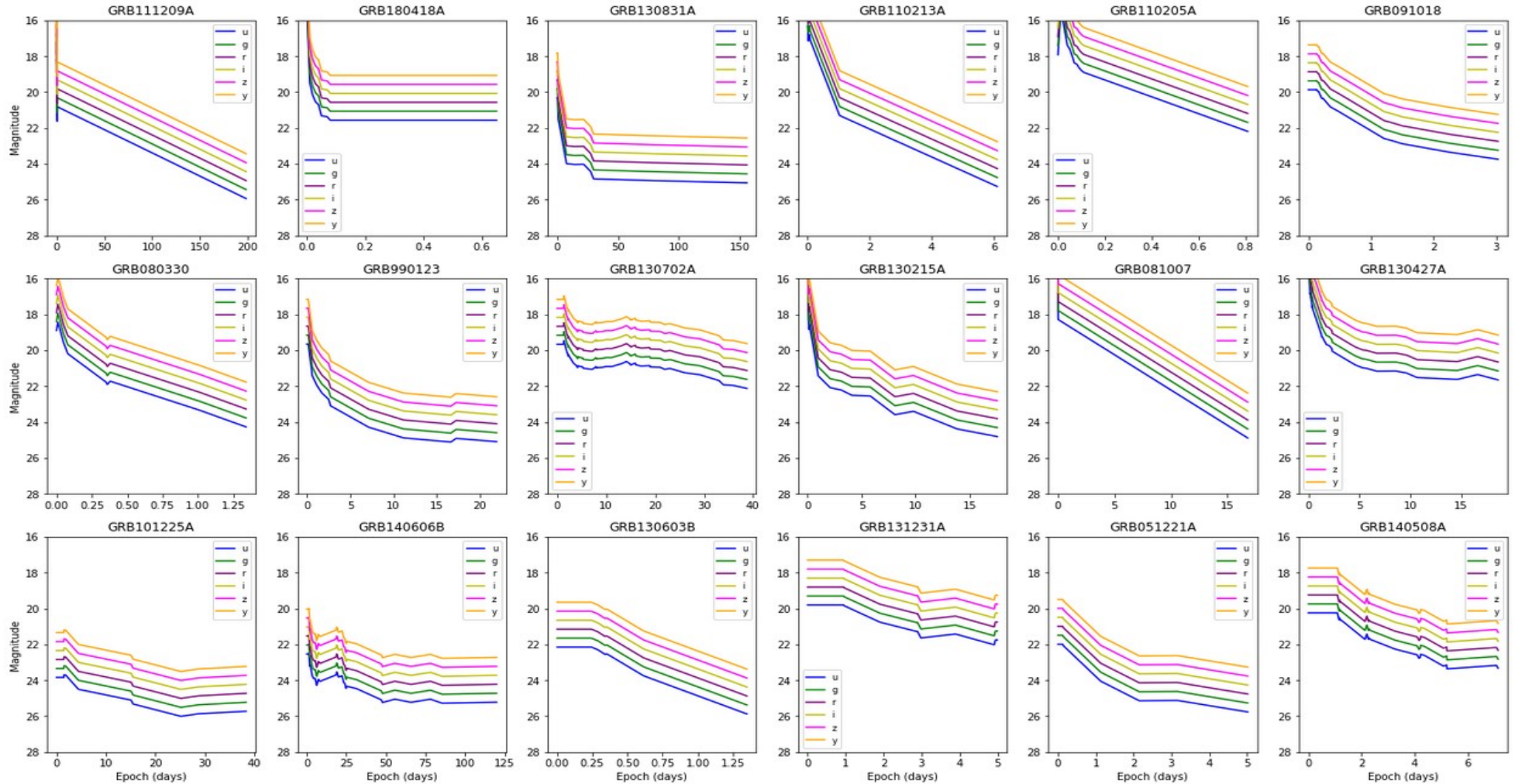


# Catalog of optical GRBs



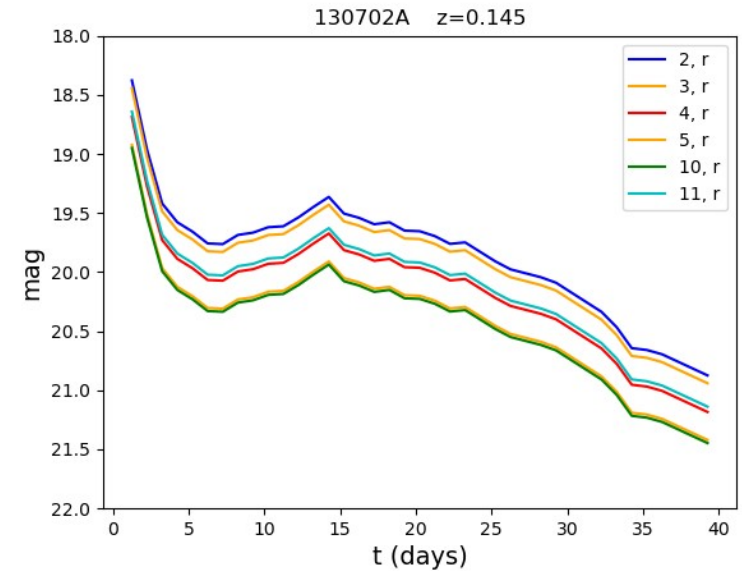
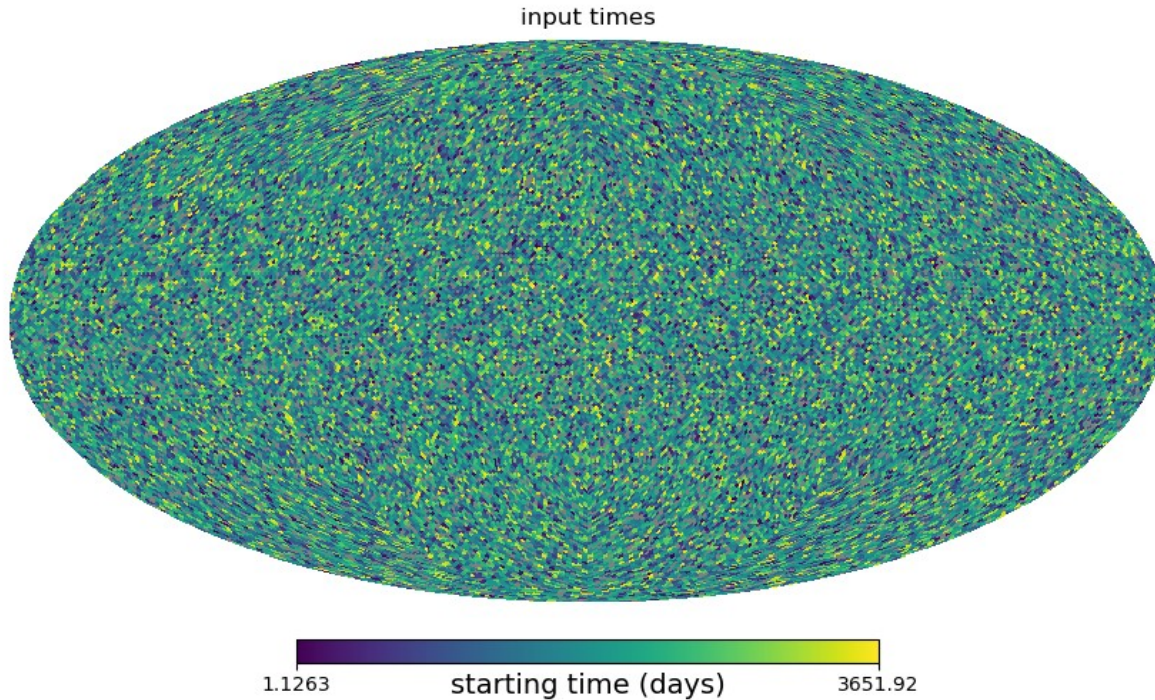
- Redshift between 0.145 and 1.6
- LC duration between 1.3 and 200 days
- Add fake LCs in ugizy with 0.5 difference between the consecutive filters
- Convert the redshift to distance (di) in Mpc using  $\Lambda$ CDM model ( $H_0 = 70$  km/sec/Mpc,  $\Omega_{m_0} = 0.3$ )

# Catalog of optical GRBs



# Available data

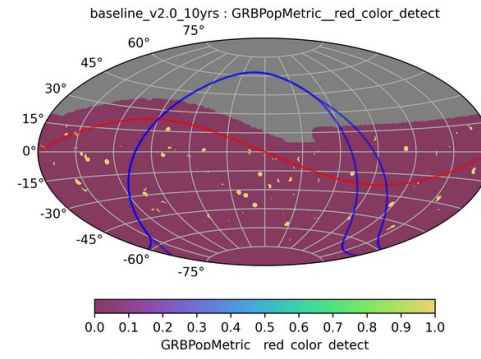
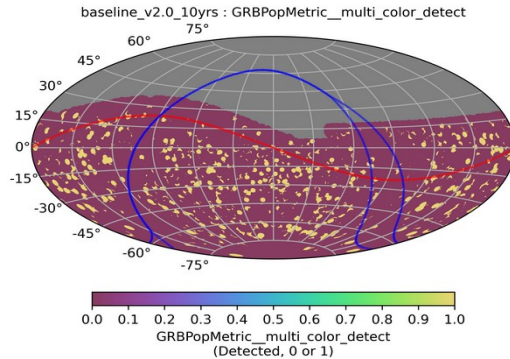
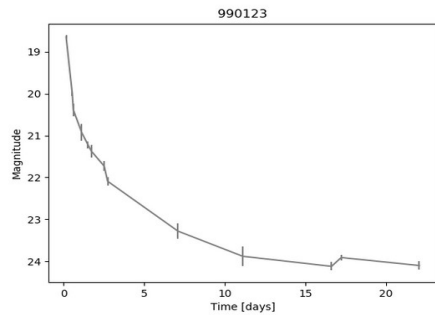
- For each GRB : Distribute  $10^5$  LCs uniformly in volume (increasing as a function of distance<sup>3</sup>, following <https://arxiv.org/pdf/2106.06820.pdf>)
- Distance range [di-100, di+100]





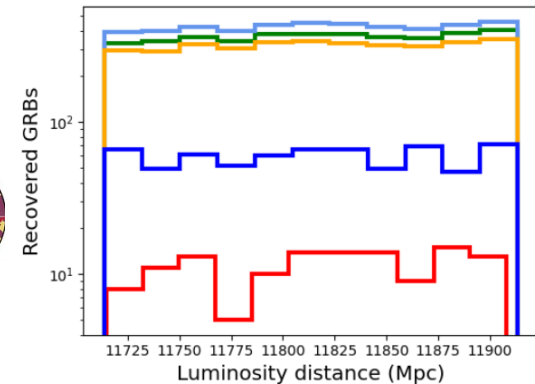
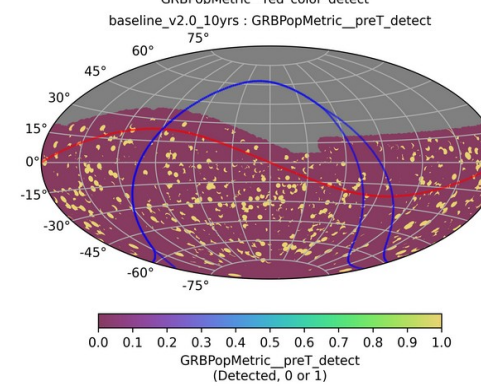
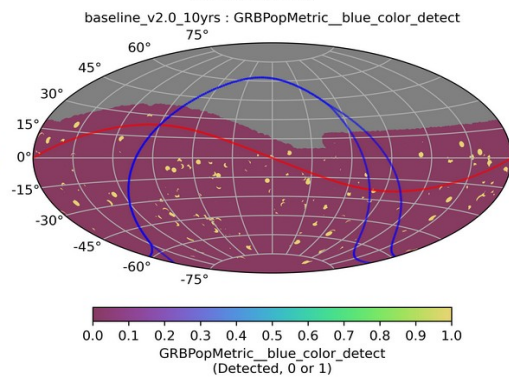
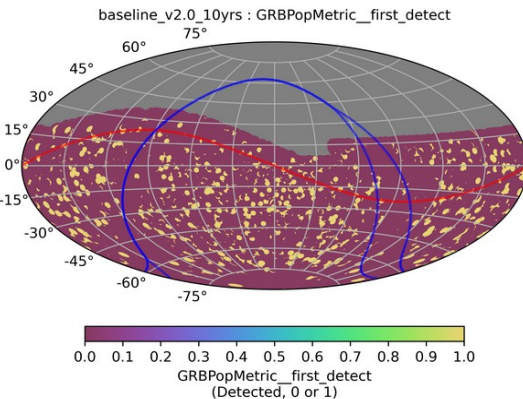
# Define the detectability criteria

- *first\_detect*: at least two detections ( $> 5\sigma$ ) during the first 5 days
- *preT\_detect*: at least two detections during the first 3 days
- *multi\_color\_detect*: Color-based simple detection criteria: detect at least twice, with at least two filters
- *red\_color\_detect*: Number of detected points in izey bands (4 detections)
- *blue\_color\_detect*: Detected at least in ugr colors (4 detections)



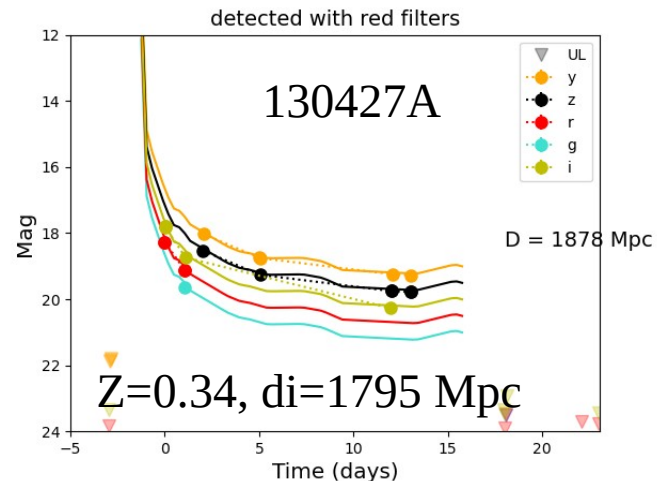
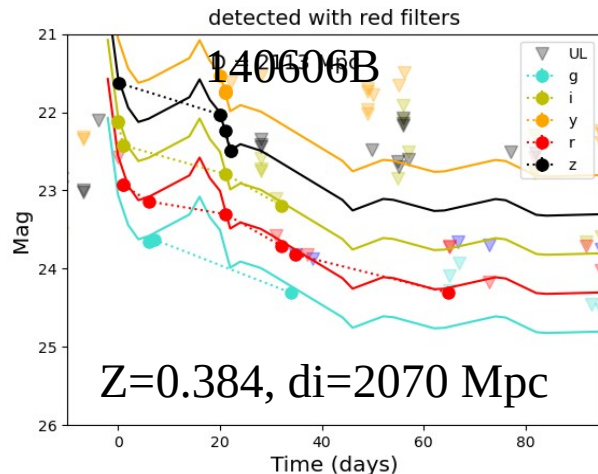
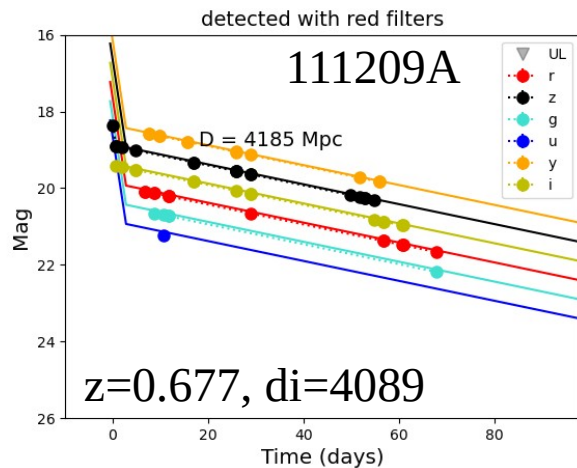
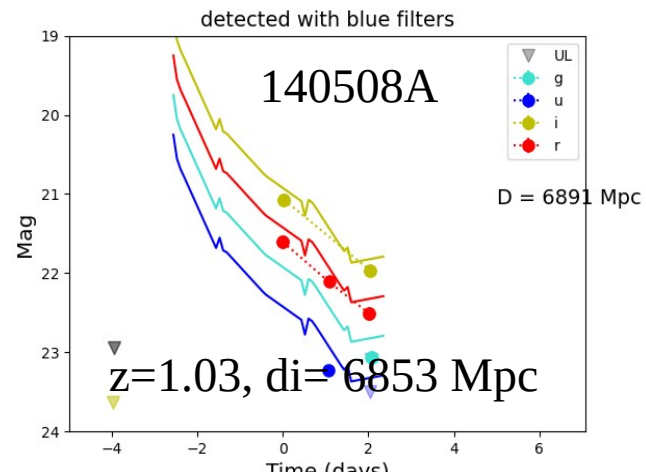
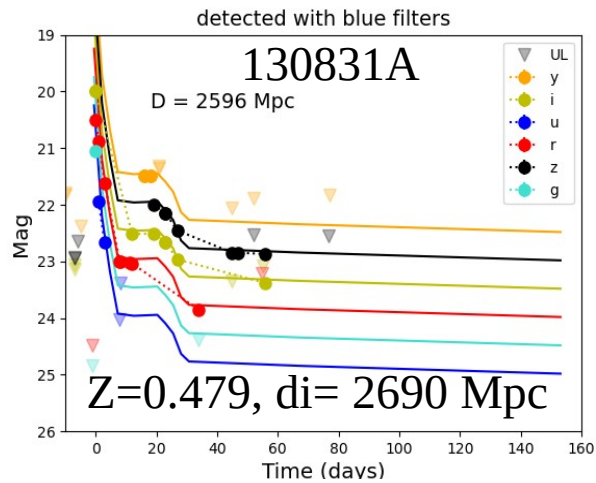
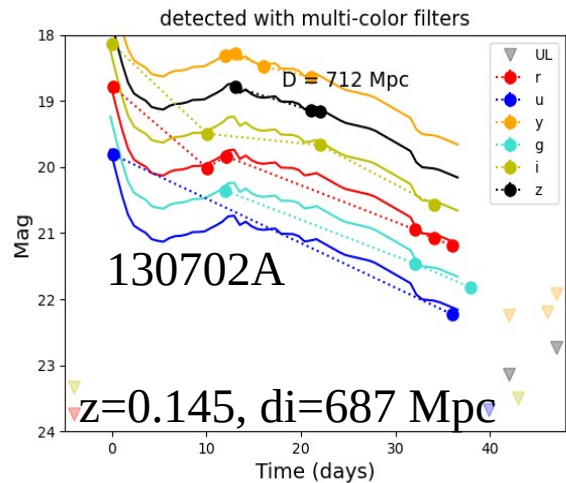
For other GRBs see [here](#)

- preT\_detect
- first\_detect
- multi\_color\_detect
- red\_color\_detect
- blue\_color\_detect

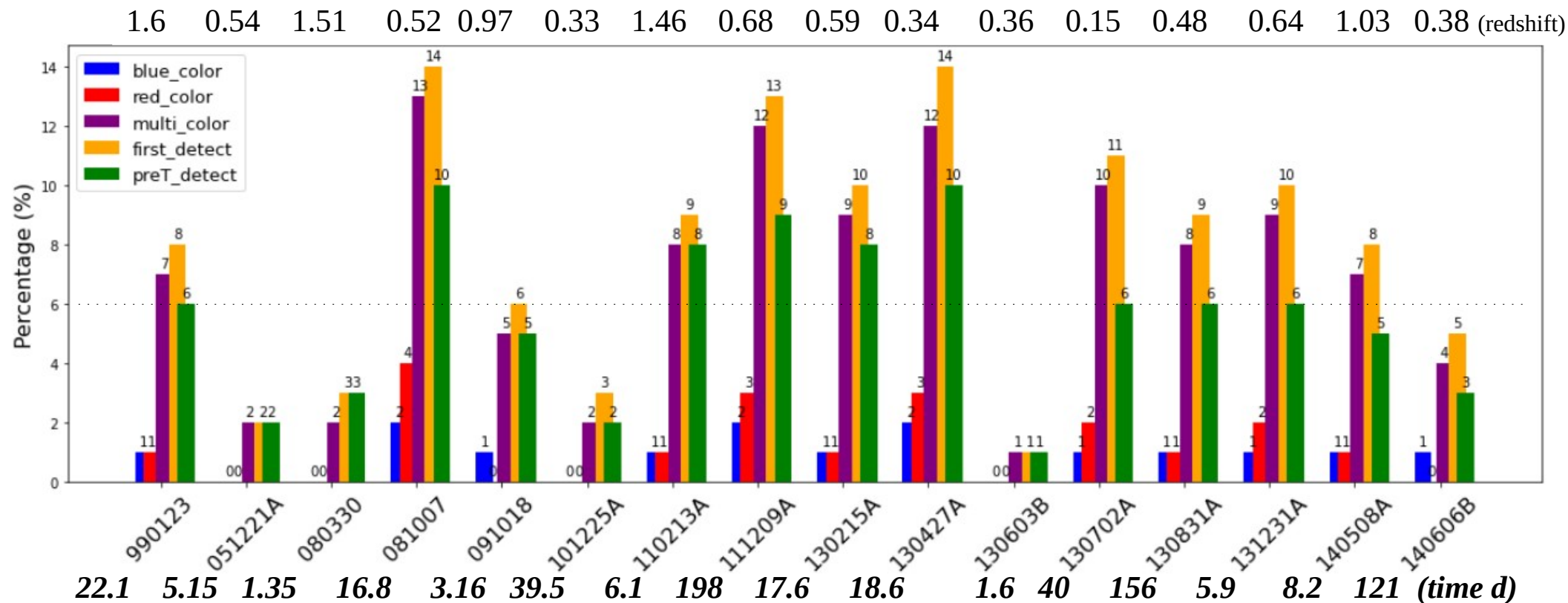




# Output GRB Lcs with LSST



# Detectability with LSST



- Short optical GRB LCs (<6 days) have a detection probability of less than 6%
- The maximum detectability for each GRB is found with the criterium first\_detect (detecting at least two events during the first five days)

# *To-do list*

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- Assume a constant color variation (g-r) or (r-i) according to literature to convert R into r data ([transformation](#))
- Define suitable criteria for GRB analysis
- Compare between the detectability with the baseline cadence and the rolling, the ddf, the presto-color and the long-gap cadences
- Understand the effect of each of the survey strategies on the detectability of a given GRB LC
- Construct the output LCs of LSST observations and compare them with the real LCs