

Estimate of the Galactic electron and positron fluxes at the Earth

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UNIVERSITÀ
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DI TORINO
ALMA UNIVERSITAS
TAURINENSIS



In collaboration with

Julien Lavalle

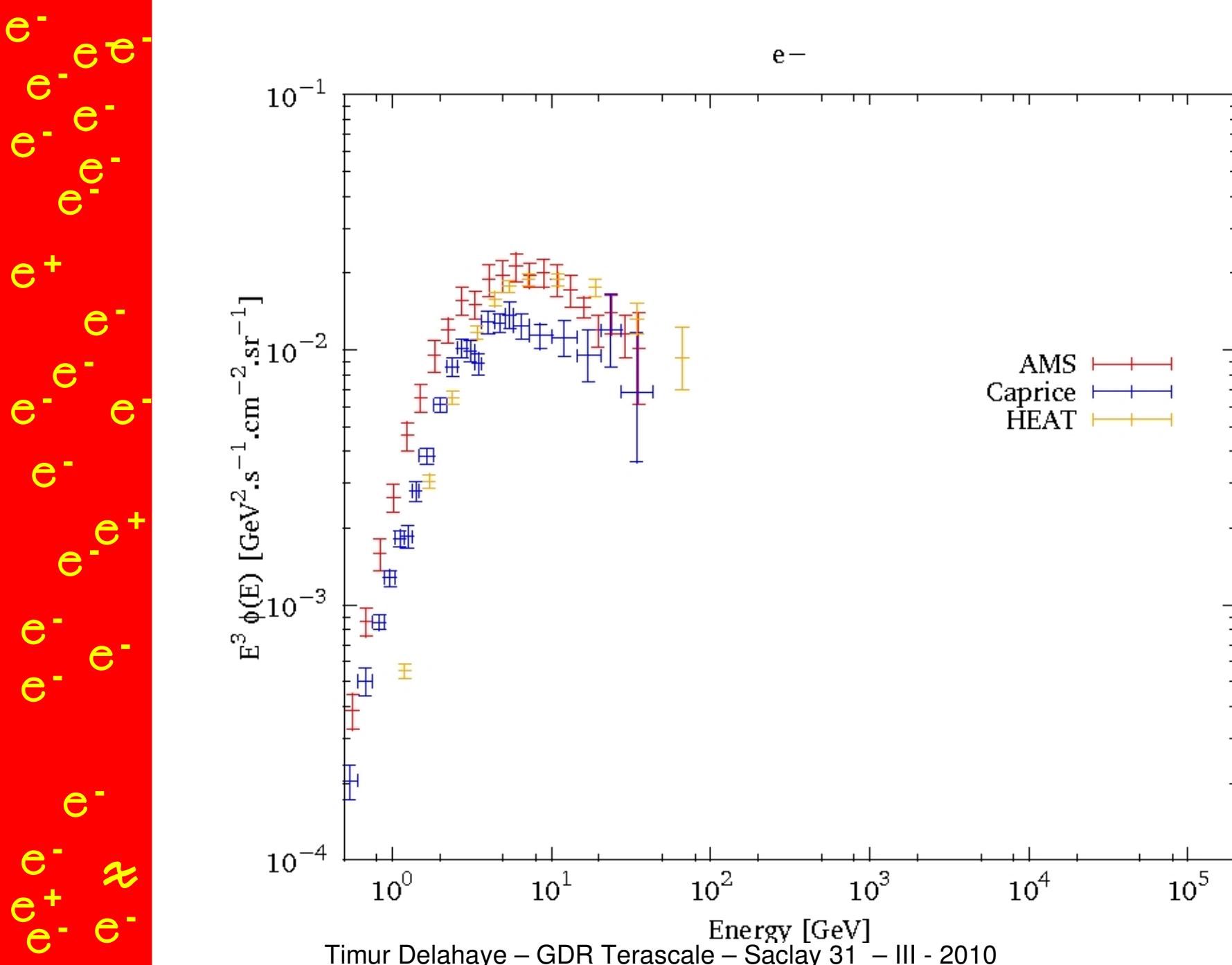
Roberto Lineros

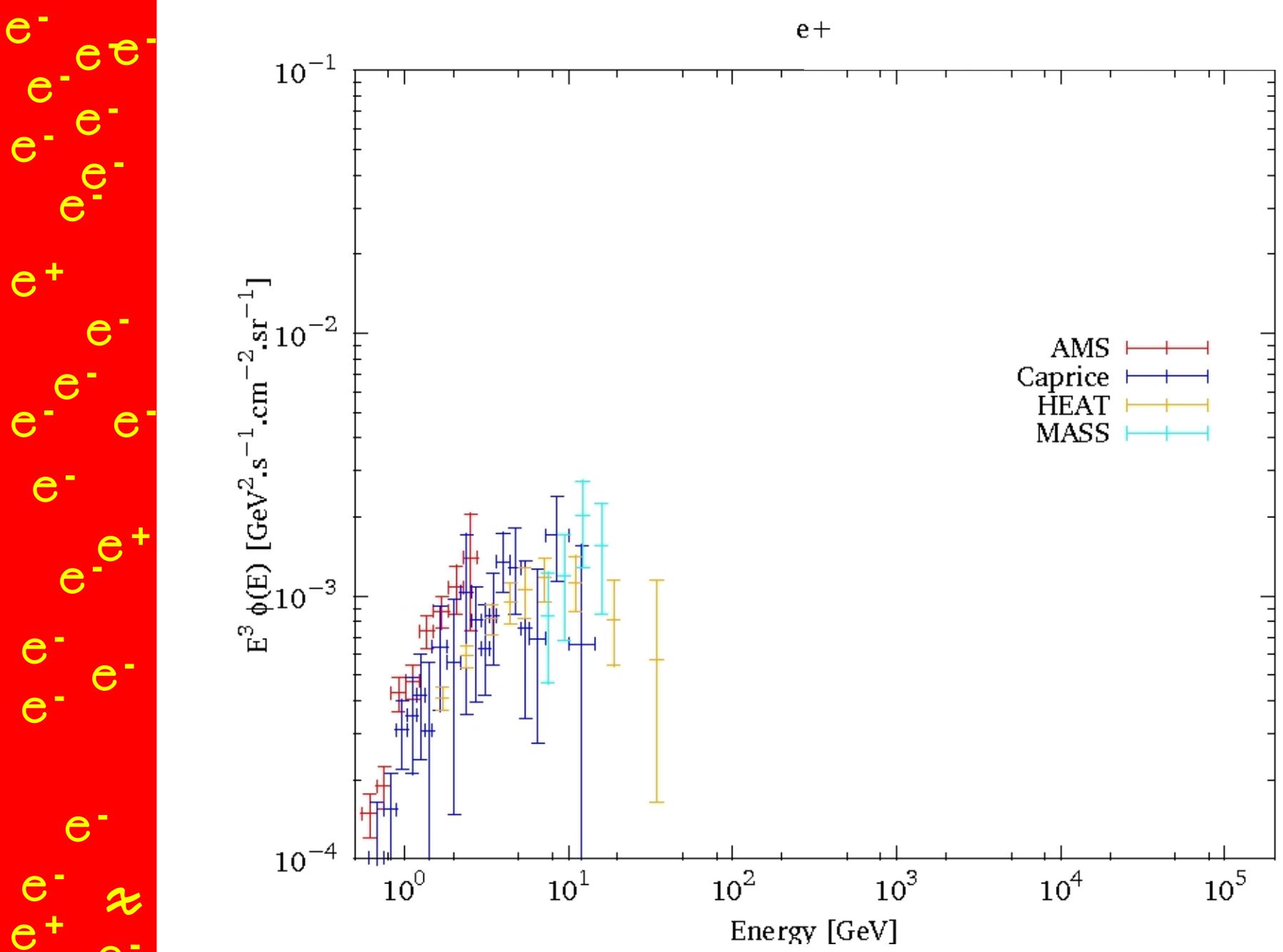
Fiorenza Donato

Nicolao Fornengo

Pierre Salati

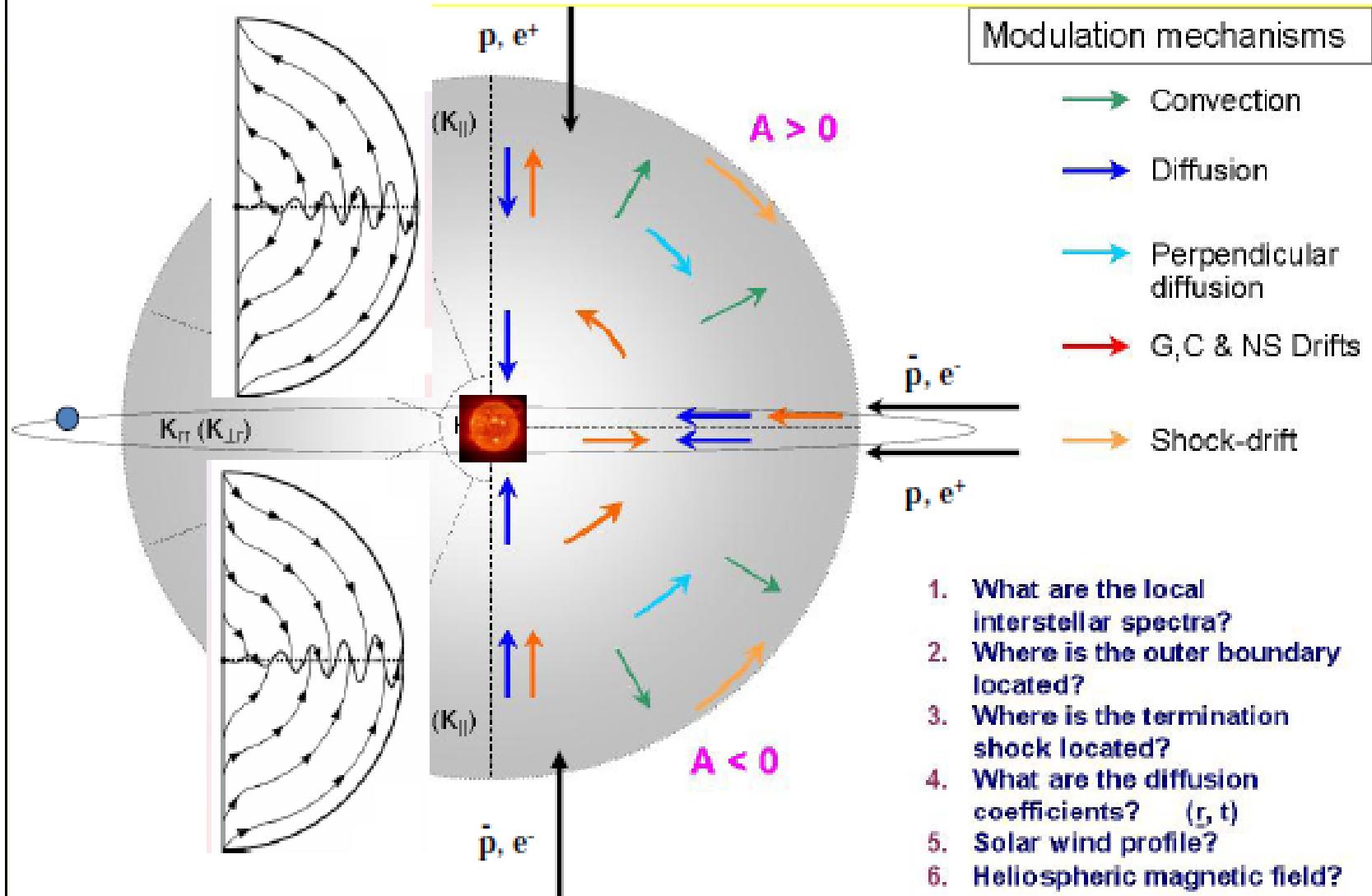
arXiv:1002.1910



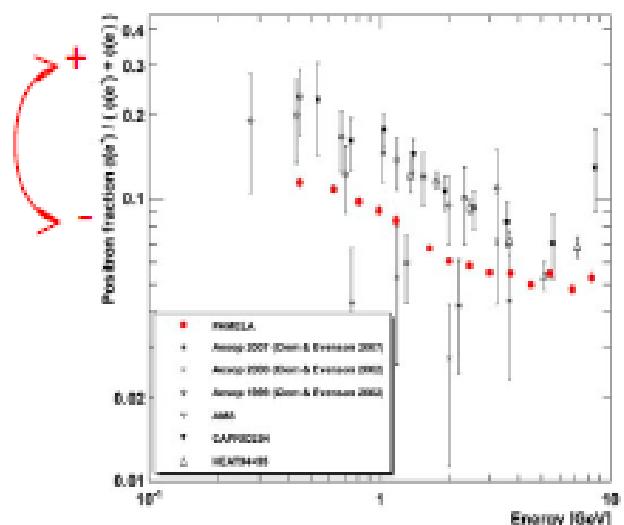
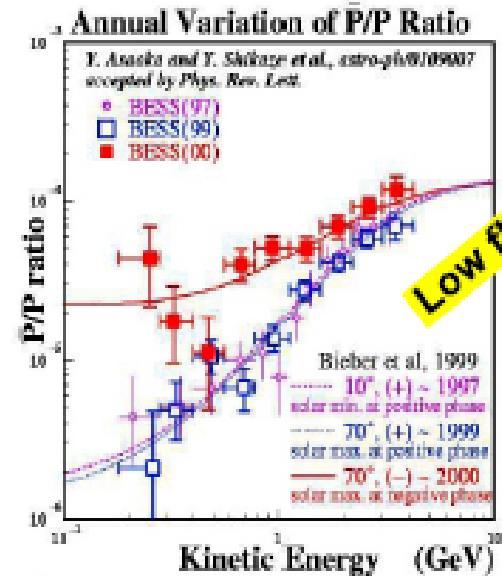
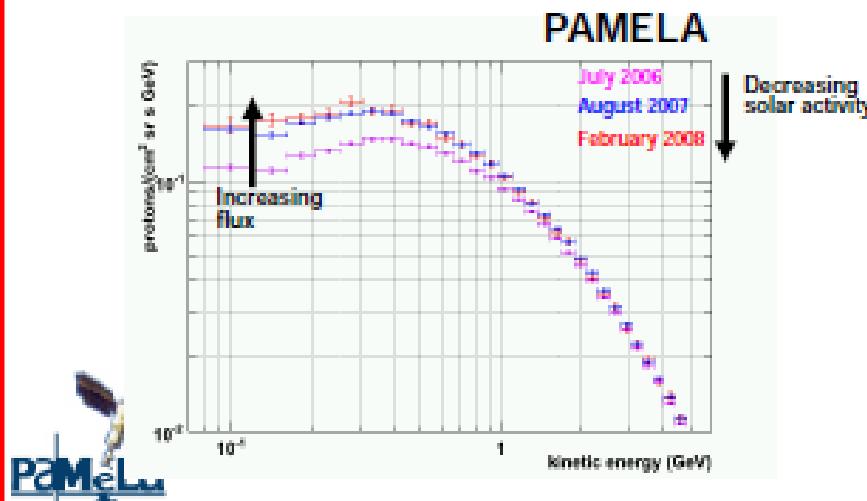
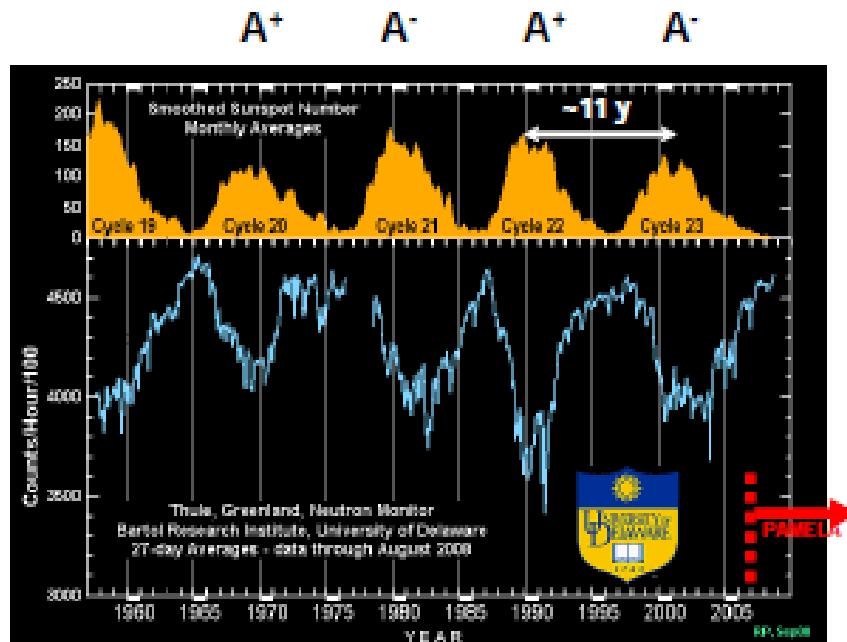


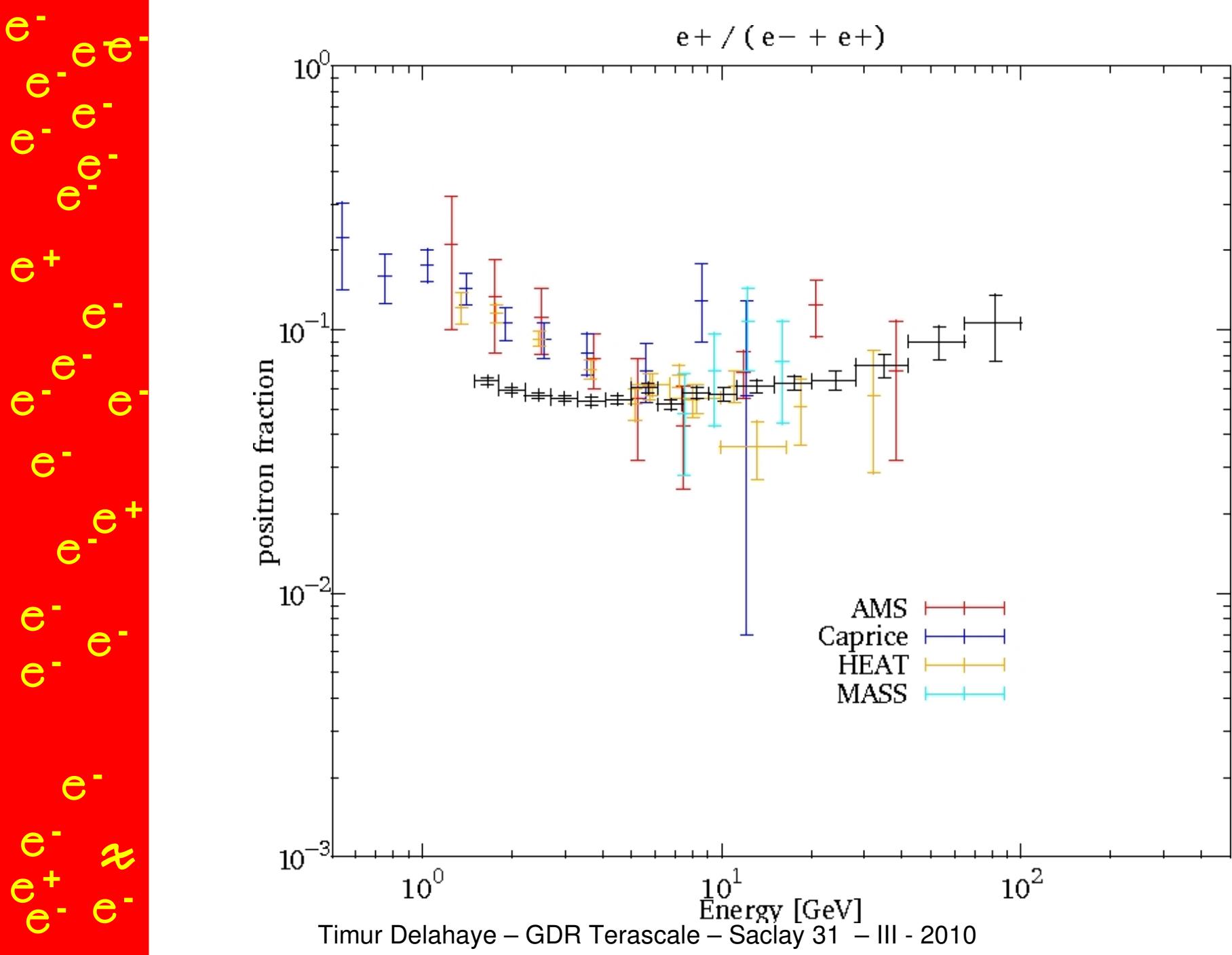
Charge dependent solar modulation

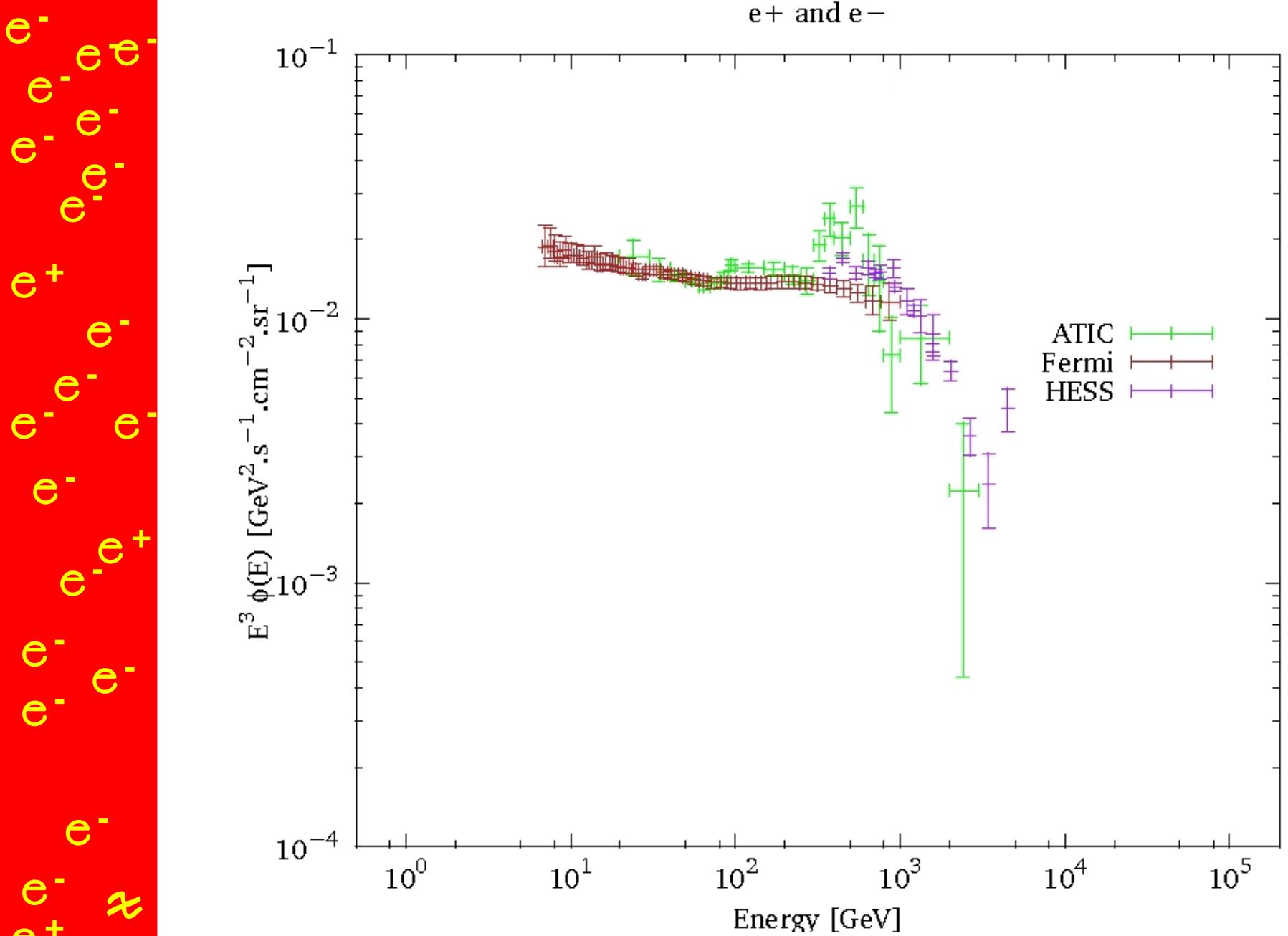
e⁻ e⁻
 α
e⁻ e⁻
e⁻ e⁻
e⁻ e⁻
e⁻ e⁻
e⁻ e⁻



Solar modulation







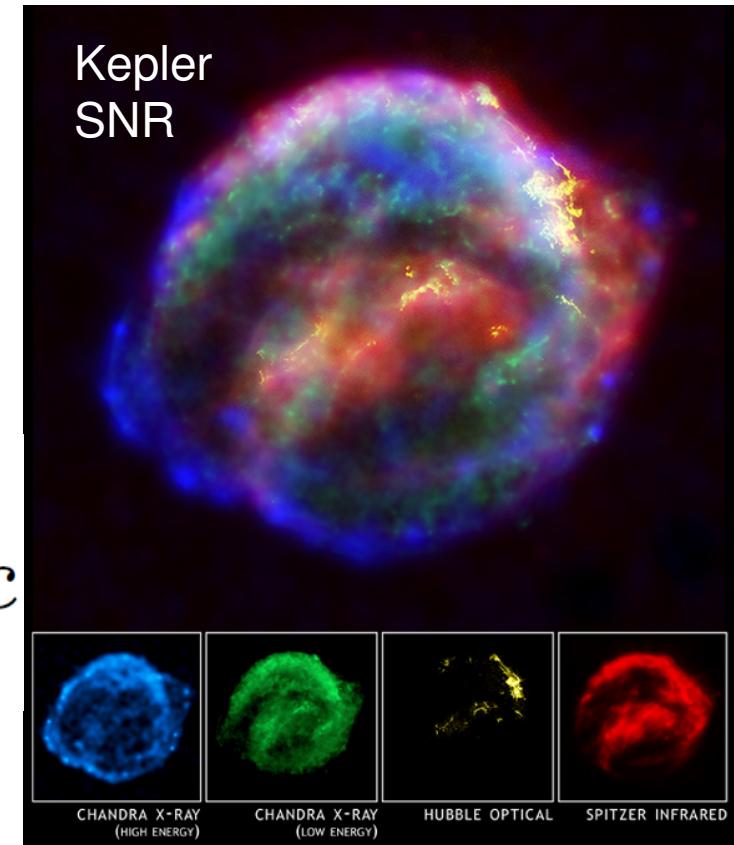
Model

- Diffusion equation

$$-K_0 E^\delta \Delta \Psi(r, z, E) + \partial_E [b_{loss}(E) \Psi(r, z, E)] = Q(r, z, E)$$

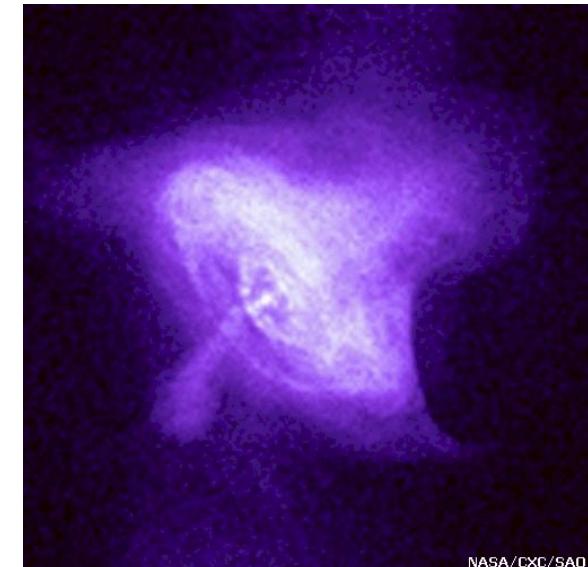
Sources

- Supernova remnants



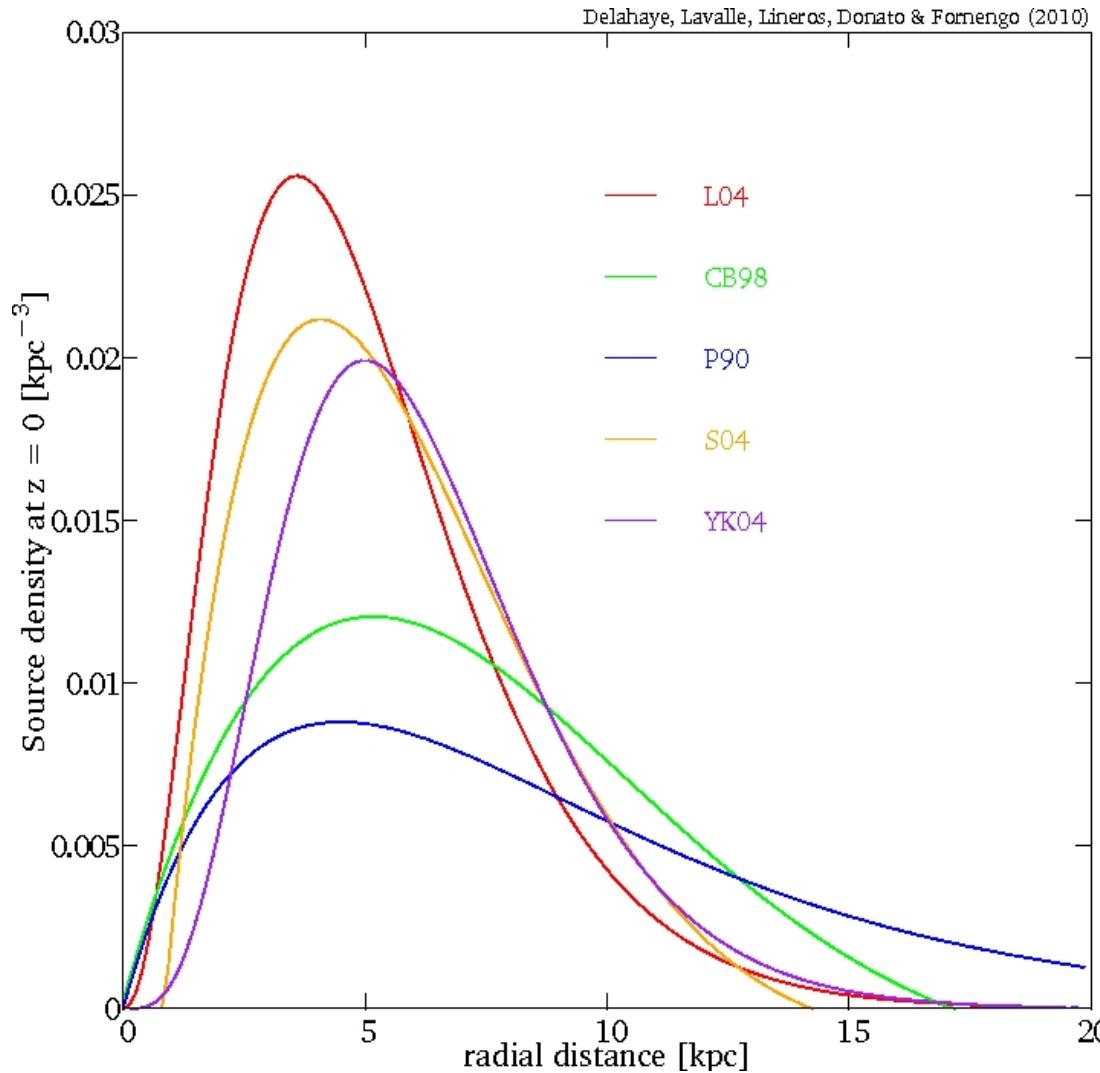
Sources

- Supernova remnants
- Pulsars

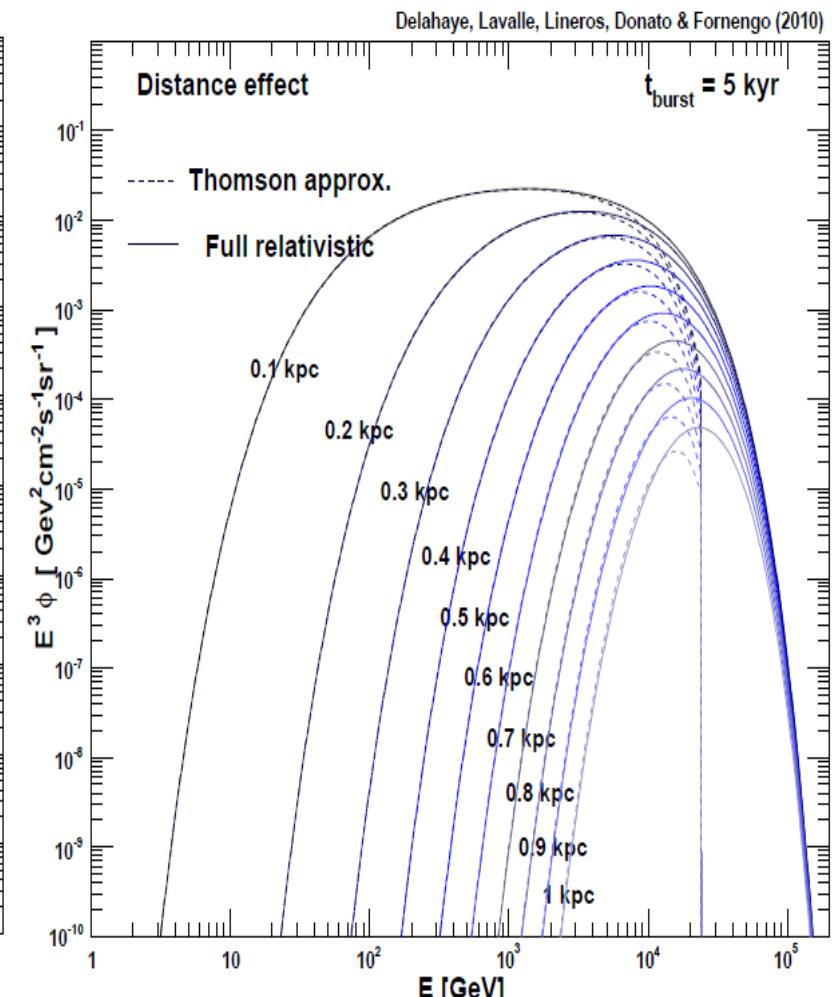
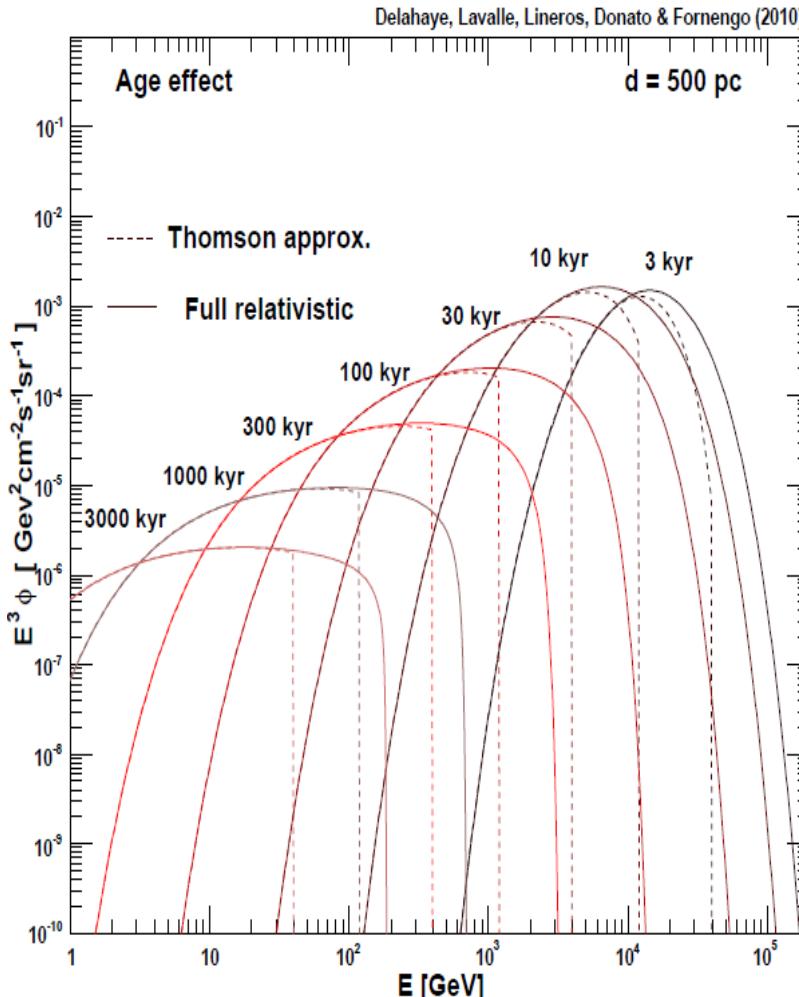


$$Q_0 E^{-\gamma} e^{-E/E_c}$$

Source distribution



Discreteness



Source catalogues

- **Green** SNR catalogue

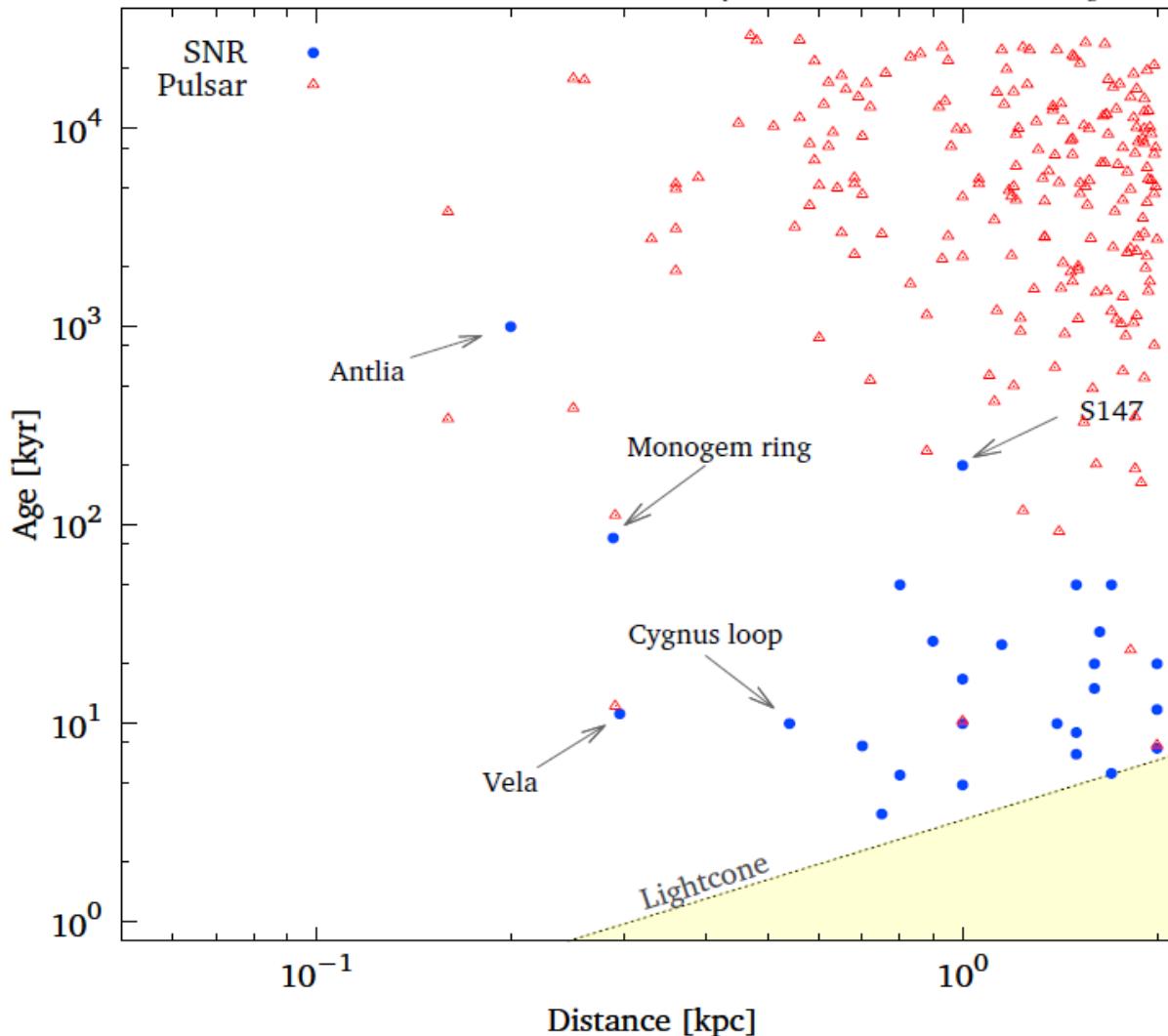
- <http://www.mrao.cam.ac.uk/surveys/snrs/>

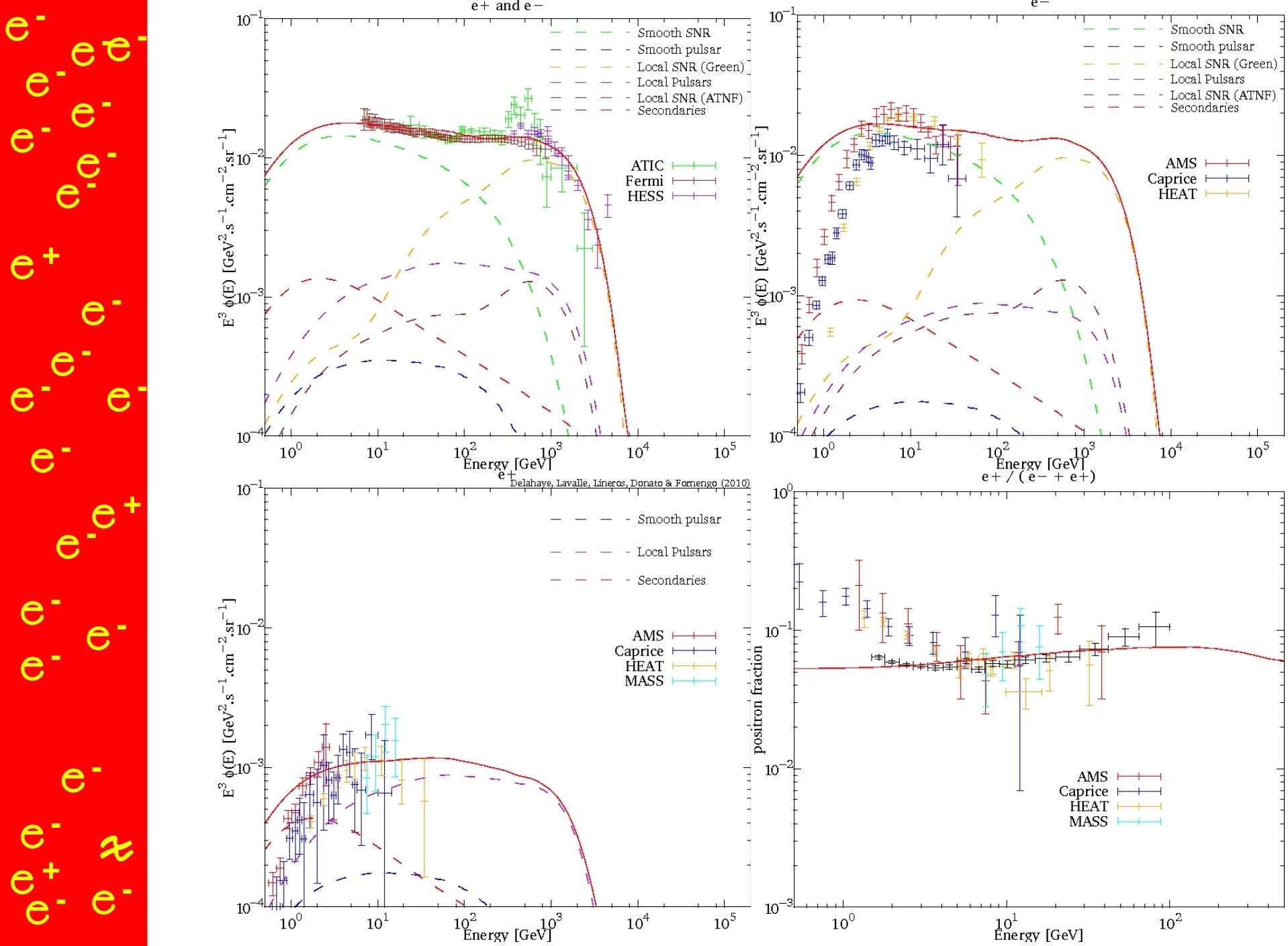
- **ATNF** pulsar catalogue

- <http://www.atnf.csiro.au/research/pulsar/psrcat/>

Local sources

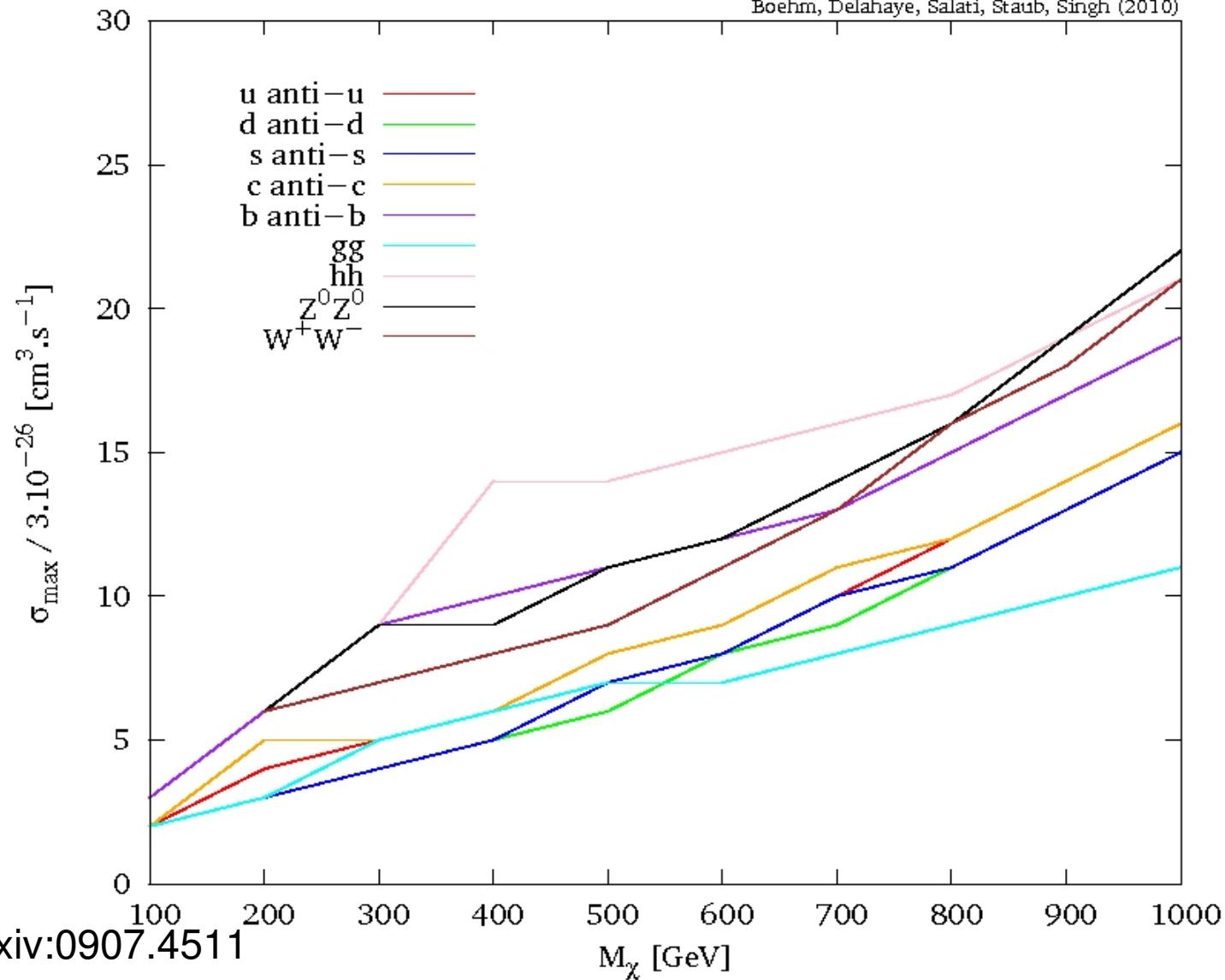
Delahaye, Lavalle, Lineros, Donato & Fornengo (2010)





Anti-protons

Boehm, Delahaye, Salati, Staub, Singh (2010)



arxiv:0907.4511

Conclusions

- No need for Dark Matter to explain current data
=>Does not mean that it is not there!
- Dark matter to quark coupling can be large

You want more ?

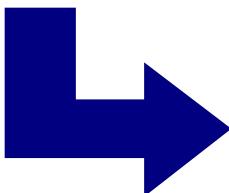
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Open questions

- Secondary production in SNRs

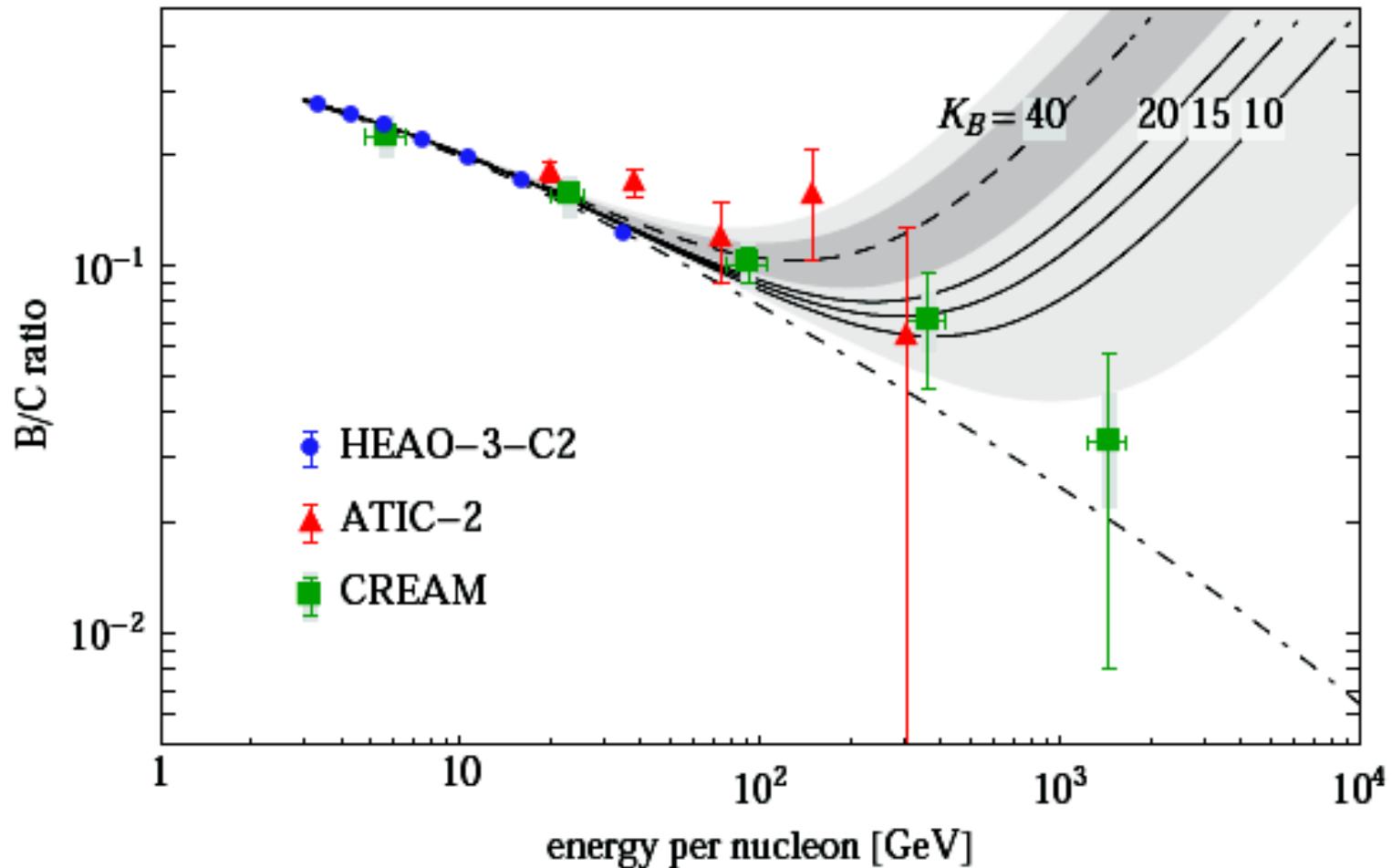
Blasi & Serpico (PRL103, 2009)

Mertsch, Ahlers & Sarkar (PRD80 & PRL103, 2009)



anti-protons & B/C

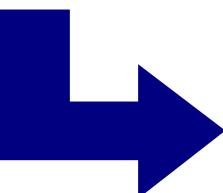
Open questions



Ahlers, Mertsch & Sarkar PRD 80, 2009

Open questions

- Fermi features
 - Expected (local astrophysical sources)
 - Dark Matter



positron absolute spectra