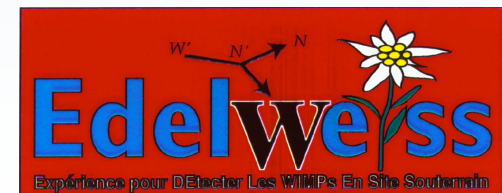


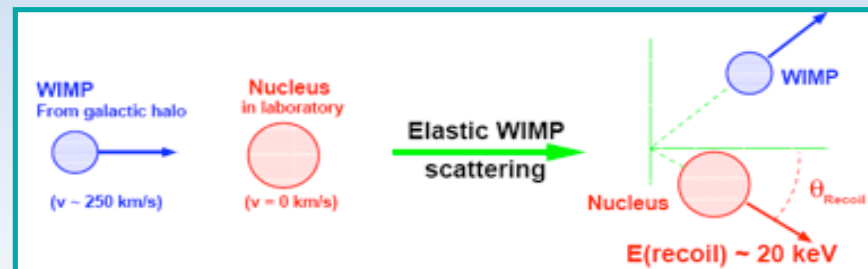
# First results of EDELWEISS II using Ge detectors with interleaved electrodes

A. S. Torrentó - CEA/IRFU



# WIMP detection

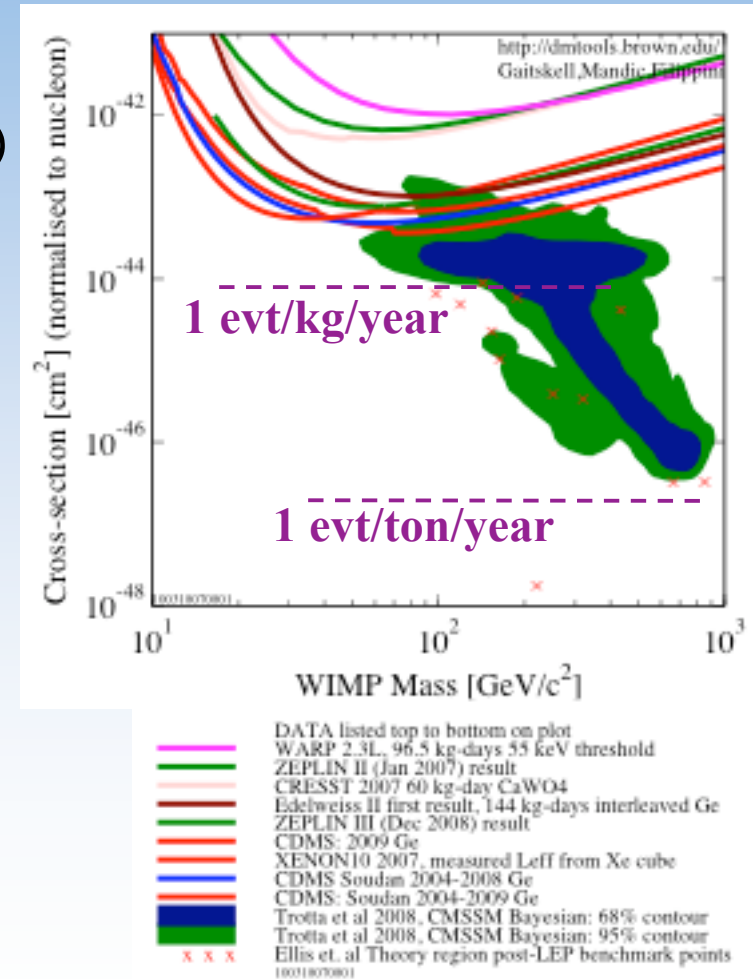
- Hypothesis: galactic halo filled with WIMPs
  - Direct: nuclear recoil from WIMP-nucleus scattering,  
 $E_r \sim 1-100 \text{ keV}$



- Indirect: WIMP annihilation products ( $\gamma$ -rays, X-rays,  $\nu$ , charged particles).

# WIMP direct detection

- Measured quantity: interaction rate  $R$ 
  - limit in the parameter space  $(m_\chi, \sigma_{SI})$ 
    - $m_\chi \sim 10 \text{ GeV} - 10 \text{ TeV}$
    - $\sigma_{SI} \sim 10^{-6} - 10^{-12} \text{ pb}$  poorly constrained
  - $R$  very low,  $1-10^{-5} \text{ evt/kg\_detector/year}$
- Detector specifications to increase sensitivity to very small  $R$ :
  - Massive (kg - tons)
  - Low energy threshold ( $E_p \sim \text{few keV}$ )
  - Efficient rejection of all backgrounds ( $\gamma$ ,  $\beta$  radioactivity, CR-induced signals -  $\mu$ , neutrons)



# WIMP signatures

- Nuclear recoils (not electronic)
- Unique interactions (not multiple)
- Signals uniformly distributed in the detector volume (not surfaces)
- Exponential  $E_r$  spectrum (no peaks/other features)
- Directionality & annual modulation (relative movement Sun-WIMP galactic halo & Earth-Sun).
- $R \propto A^2$  of target nucleus (SI interactions)

# WIMP-signal detection & discrimination

- Different experimental approaches
  - Bolometric detectors:
    - Energy yield from WIMP-nucleus interaction measured as an increase in T of the absorber.
    - Discriminating measurement nuclear/electronic recoils
      - Ionisation signal: i.e. Ge, Si absorbers in EDELWEISS, CDMS
      - Scintillation yield: i.e.  $\text{CaWO}_4$  absorbers in CRESST.
  - Noble liquid detectors (Xe, Ar):
    - Scintillation yield, timing to both detect & discriminate nuclear recoils (XENON, ZEPLIN, WARP).

# EDELWEISS

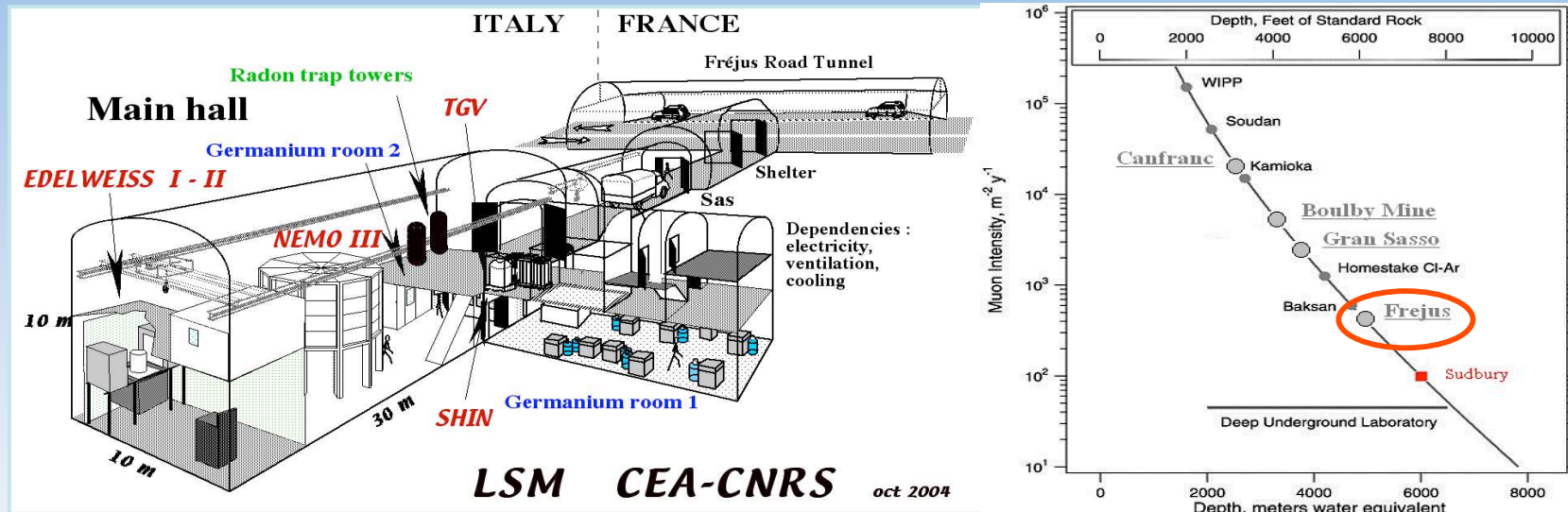
Direct search of WIMPs with Germanium bolometers

→ heat + ionisation signals, spin-independent coupling

- CEA Saclay (IRFU and IRAMIS)
- CSNSM Orsay (CNRS/IN2P3 + Univ. Paris Sud)
- IPNLyon (CNRS/IN2P3 + Univ. Lyon 1)
- Néel Grenoble (CNRS/INP)
- Karlsruhe Institute of Technology
- JINR Dubna
- Oxford University (joined in 2009)

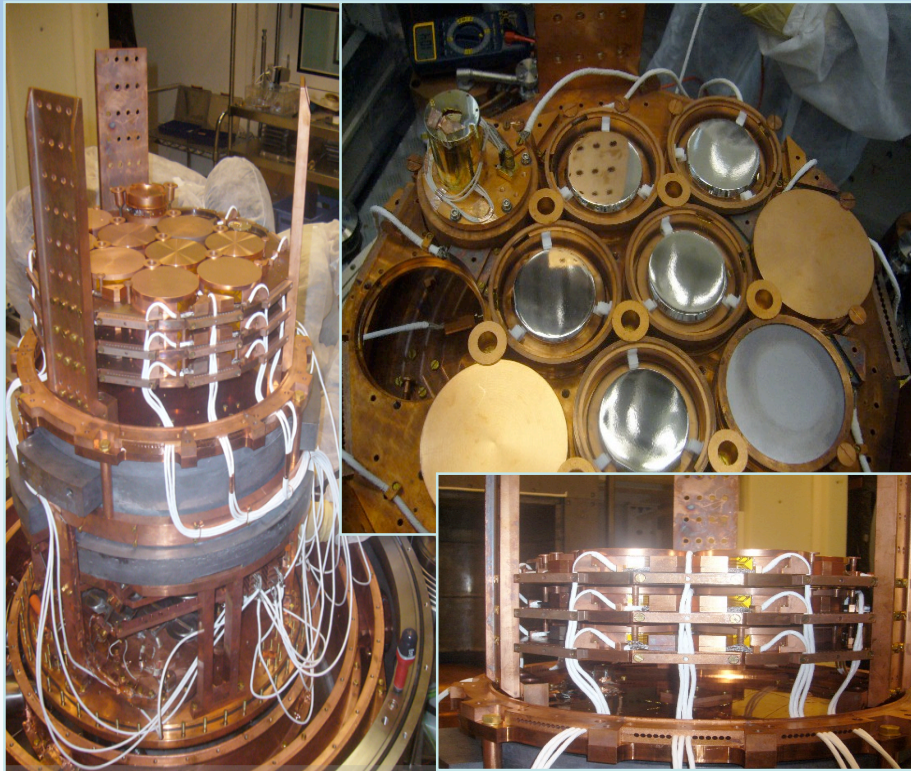


# EDELWEISS @ LSM



- Unique experimental site: Laboratoire Souterrain de Modane (LSM) in Fréjus Tunnel
- 4800 mwe depth: 4 muon/day/ $m^2$
- $10^{-6}$  neutrons ( $>1$  MeV)/ $cm^2/s$
- Deradonized air supply to reduce  $\beta$  radioactivity

# EDELWEISS II - Setup

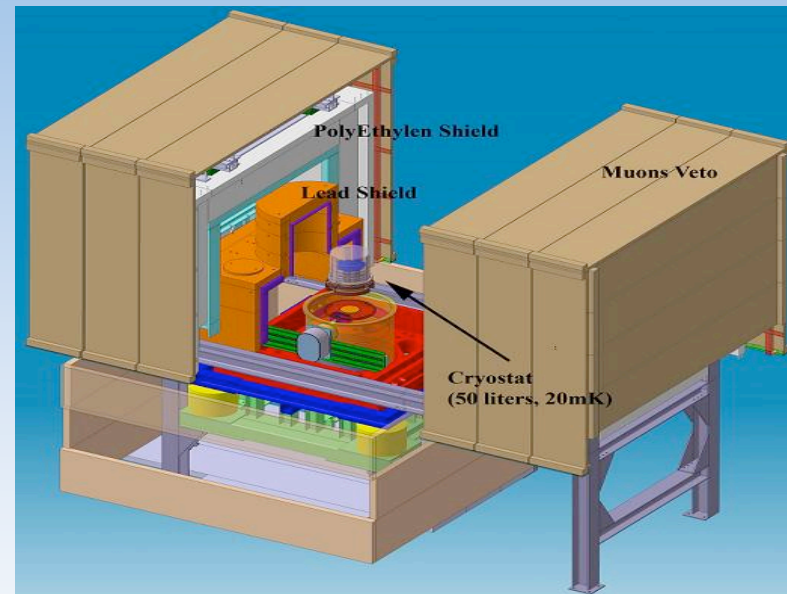


- Goal:  $10^{-8}$  pb,  $< 0.002$  evts/kg/d
- Currently 5 kg Ge, can host up to 40 kg
- Operation @ 18 mK with a dilution refrigerator controlled remotely.
- Strict control of material selection / Cleaning procedure / Environment
  - $\times 4$  reduction of  $\gamma$  background
  - (wrt EDW-I)

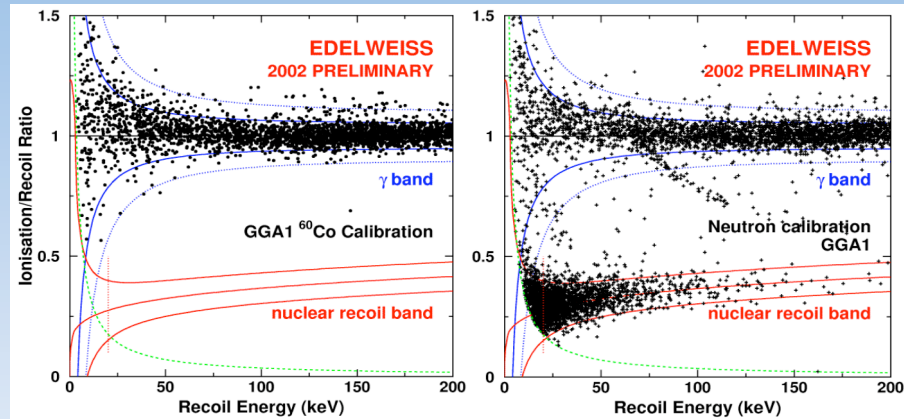
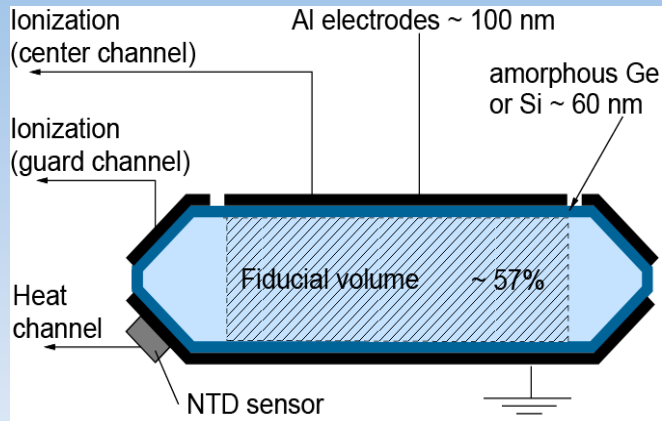


# EDELWEISS II - Shieldings

- Clean room & deradonised air ( $\beta$ )
- Active muon veto ( $\mu$ , n)
  - > 98% coverage
- 50-cm polyethylene shield (thermalise fast neutrons)
- 20-cm lead shield ( $\gamma$ )
- Moreover...
  - Radon detector
  - $^3\text{He}$ -neutron detector (thermal neutron monitoring)
  - Liquid scintillator neutron counter (studies of  $\mu$ -induced neutrons)
  - Remotely controlled  $^{60}\text{Co}$ ,  $^{133}\text{Ba}$  sources for detector regeneration &  $\gamma$  calibrations
  - AmBe source for neutron calibrations



# EDW I - Detectors

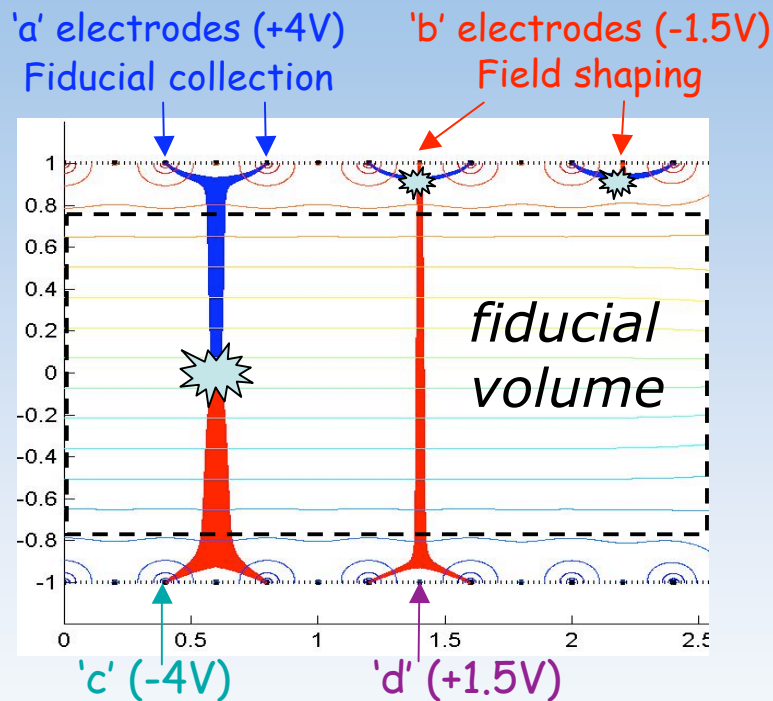


- NTD sensor: Heat measurement @ 20 mK
- Al electrodes on surfaces: Ionization measurement @ few V/cm
- Discriminating variable electronic/nuclear recoils :  $Q \sim \text{ionization/heat}$

$$\text{Ge} \rightarrow Q_{\text{electronic}} \sim 3 Q_{\text{nuclear}}$$

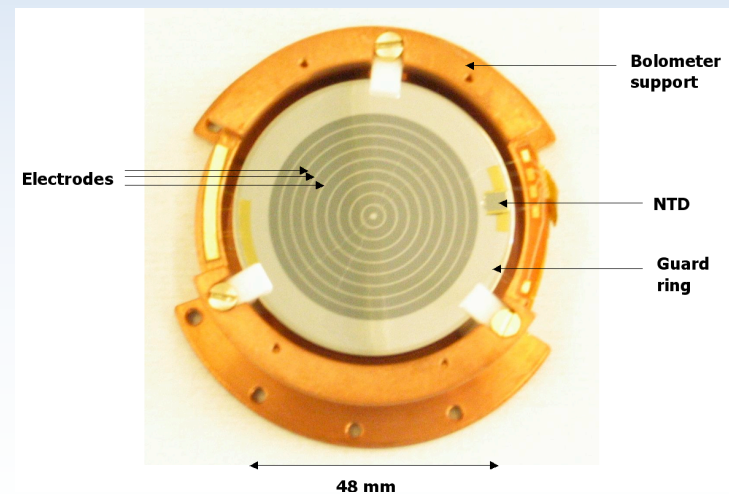
- Limitation: inefficient charge collection in surface interactions mimics nuclear recoils  $\rightarrow Q_{\text{surface}} \sim Q_{\text{nuclear}}$
- EDELWEISS I best limit:  $\sigma_{\text{SI}} \sim 10^{-6} \text{ pb}$

# EDW II - ID Detectors



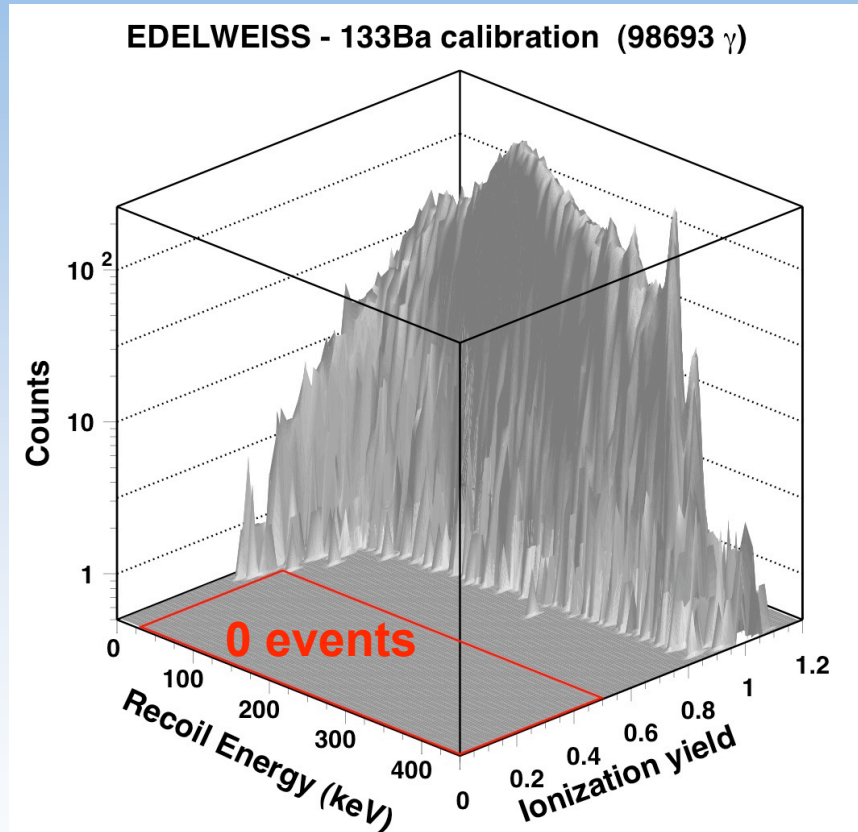
- First 200g detector built 2007
- 1x200g + 3x400g tested in 2008
- **10x400g running since beginning 2009**

- Keep the EDW-I NTD thermal detector
- Modify the E-field near the surfaces with interleaved electrodes
- Use 'b' and 'd' signals as vetos against surface events
- Redundancy provides high rejection efficiency

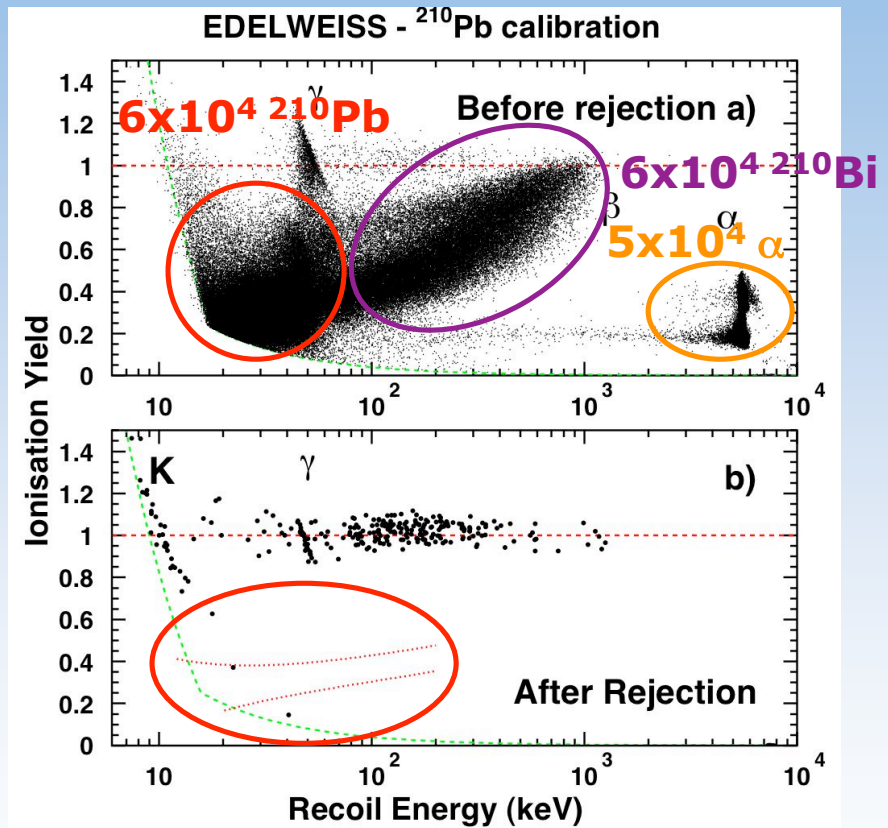


# ID Detectors - overall backg rejection

[Phys Lett B 681 \(2009\) 305-309 \[arXiv:0905.0753\]](#)



ID 400g  $\rightarrow 10^{-5}$   $\gamma$  rejection

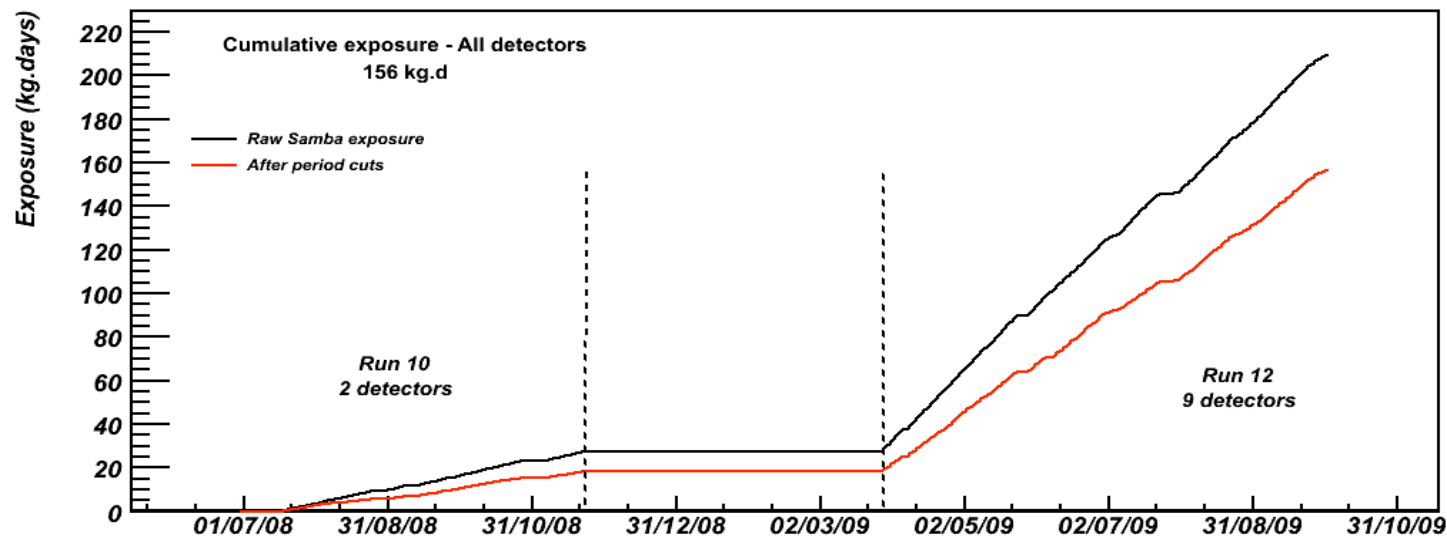


ID 200g  $\rightarrow 10^{-5}$   $\beta$  rejection

Meet specifications needed for a  $10^{-8}$  pb dark matter search

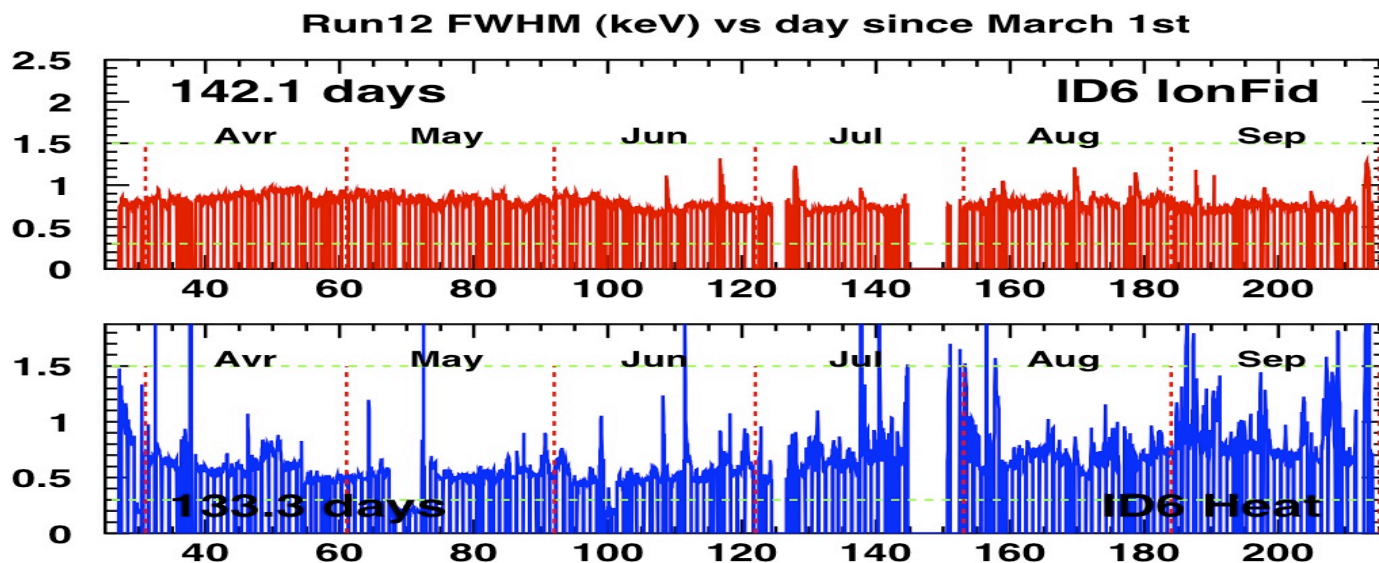
# WIMP search data set

- **Jul-Nov 2008:** validation runs ID detectors ( $2 \times 200\text{g}$ ), 20 kg.d
- **Apr-Sept 2009:** physics run ( $10 \times 400\text{g}$ ) ID detectors, 144 kg.d  
→ 166 kg.d of data presented
- **Oct 2009 - Spring 2010:** physics run continues



# WIMP search DAQ efficiency

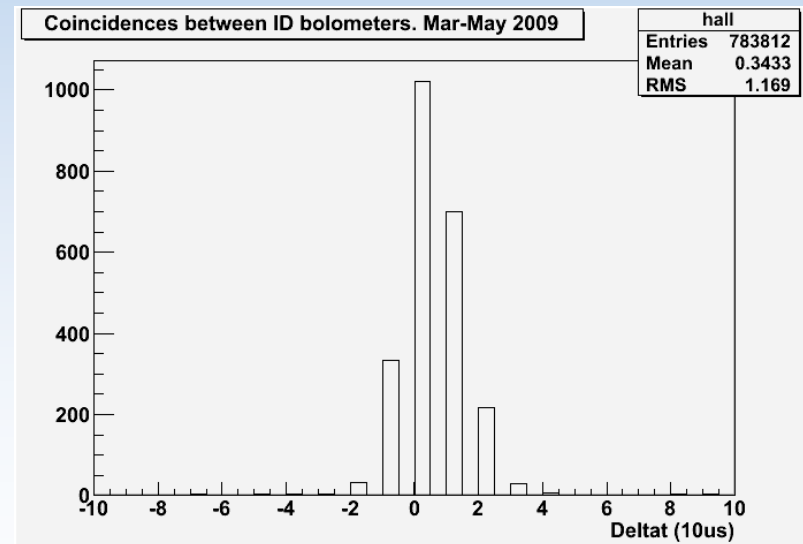
- DAQ running 80% of time (maintenance operations)
  - 6% of running time:  $\gamma$ , n calibrations
  - Working detectors: heat + both fiducial electrodes + 3 veto & guard / 4  
(due to redundancy, background rejection is not affected)
    - 9/10 detectors
    - 10th detector (1 veto + 1 guard not working) OK a posteriori but not included in present analysis
- **proves reliability of IDs in real conditions of WIMP search**
- Average resolutions:  $\sigma \sim 400$  eV ionization, 500 eV heat



# WIMP search analysis

- 2 independent processing chains
- Heat/Ionisation pulse fit with optimal filtering using instantaneous noise spectra
- Selection cuts:
  - Period selection based on baseline noises:  $\varepsilon = 80\%$
  - Pulse reconstruction quality ( $\chi^2$ ):  $\varepsilon = 97\%$
  - Fiducial cuts based on ionization signals (160g)
  - Nuclear recoil:  $\varepsilon = 90\%$
  - Gamma rejection:  $\varepsilon = 99.99\%$
  - Bolo-bolo & bolo-veto coincidence:  $\varepsilon > 99\%$
- WIMP search threshold fixed a priori

$$E_{\text{recoil}} > 20 \text{ keV}$$



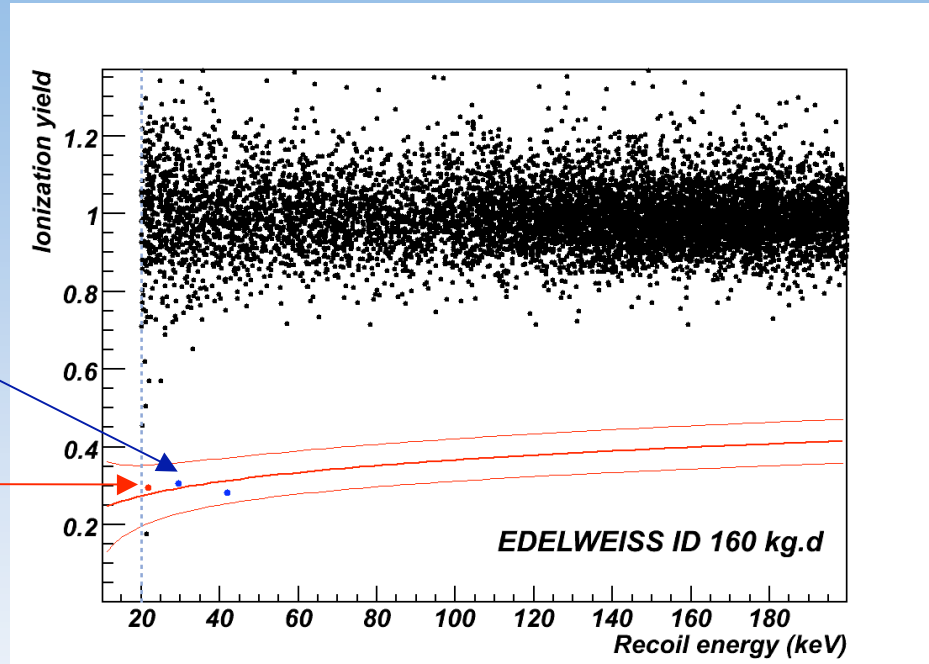
# WIMP search results

arXiv:0912.0805

160 kg.d  $\times$  90% NR band = 144 kg.d

Coincidences bolo-bolo+veto  
with  $E_r=27, 43$  keV  
 $\Rightarrow$  muon-induced neutrons  
in fiducial volume

«WIMP candidate»  $E_r = 21$  keV



## Background estimation : work in progress

First estimation *from previous calibrations/simulations*:

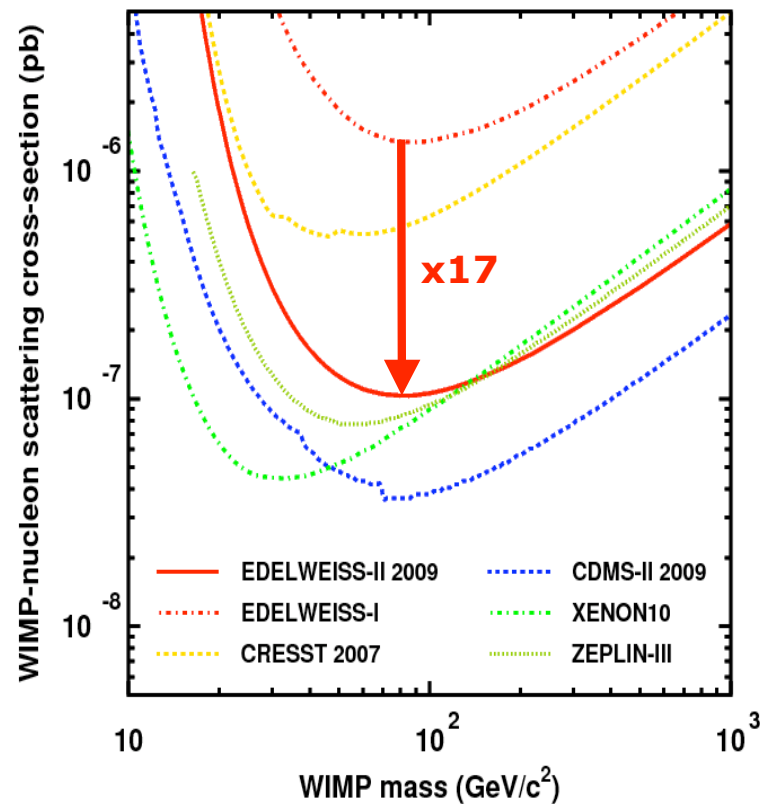
- gamma  $< 0.01$  evt (99.99% rejection)
- beta  $\sim 0.06$  evt (from ID201 calibration+obs. surf. evts)
- neutrons from  $^{238}\text{U}$  in lead  $< 0.1$  evt
- neutrons from  $^{238}\text{U} + (\alpha, n)$  in rock  $\sim 0.03$  evt
- neutrons from muons  $< 0.04$  evt

$< 0.23$  events



# WIMP cross-section limit

[arXiv:0912.0805](https://arxiv.org/abs/0912.0805)

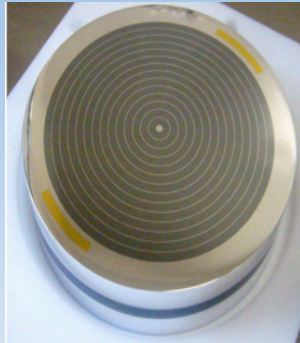


- Limit:  $\sigma_{\text{SI}} \sim 1.0 \times 10^{-7}$  pb  
 $m_{\chi} = 80 \text{ GeV}/c^2$   
90%CL
- $\times 17$  improvement wrt EDW I thanks to active surface-event rejection with ID detectors
- Further improvement in sensitivity expected.

# EDW II - Summary

- Results of 160 kg.d of EDELWEISS II in the direct detection of galactic WIMPs
- New ID detectors
  - Background rejection required for  $\sigma_{\text{SI}} \sim 10^{-8}$  pb
  - Proven to be reliable in real WIMP-search conditions
- WIMP candidate at  $E_r=21$  keV, better background estimation in progress
- Limit:  $\sigma_{\text{SI}} \sim 1.0 \times 10^{-7}$  pb for  $m_\chi = 80$  GeV/c<sup>2</sup>  
To be improved...

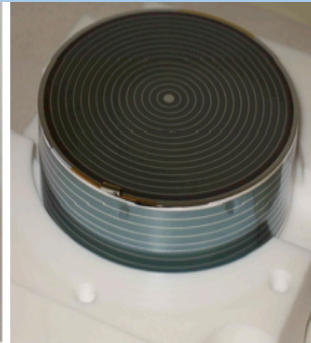
# EDW III - FID detectors



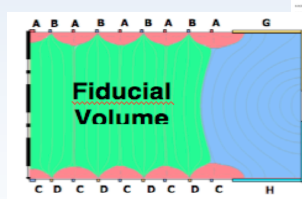
ID401 to 405:  
Φ 70mm, H 20mm, 410g



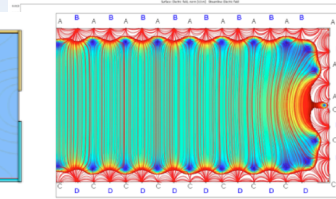
ID2 to ID5:  
Φ 70mm, H 20mm, 410g



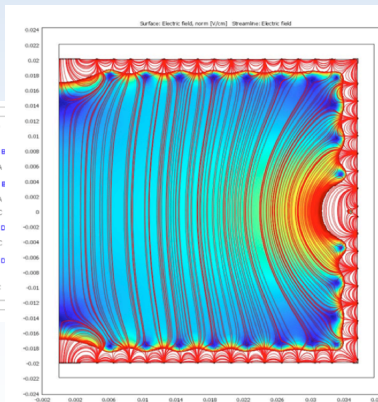
FID401 and FID402:  
Φ 70mm, H 20mm, 410g



ID400 160g



FID400 >300g



FID800 >600g

- Next phase: improve both total mass and fiducial volume/each detector
  - 400g → 800g
- ID → Full ID (FID)
  - the coverage of ring electrodes extends to the sides
- FID  $\beta$  rejection @ LSM  
4/68000 for  $E_r > 25\text{keV}$   
(~similar to ID)



## European Underground Rare Event Calorimeter Array

- EURECA: beyond  $10^{-9}$  pb
  - $\gg 100$  kg cryogenic experiment, multi-target
  - Major efforts in background control and detector development
- Joint effort: EDELWEISS, CRESST, ROSEBUD + others...
- Part of ILIAS/ASPERA European Roadmap
- Preferred site:  $60000 \text{ m}^3$  extension of present LSM, to be dug in 2011-2012

