

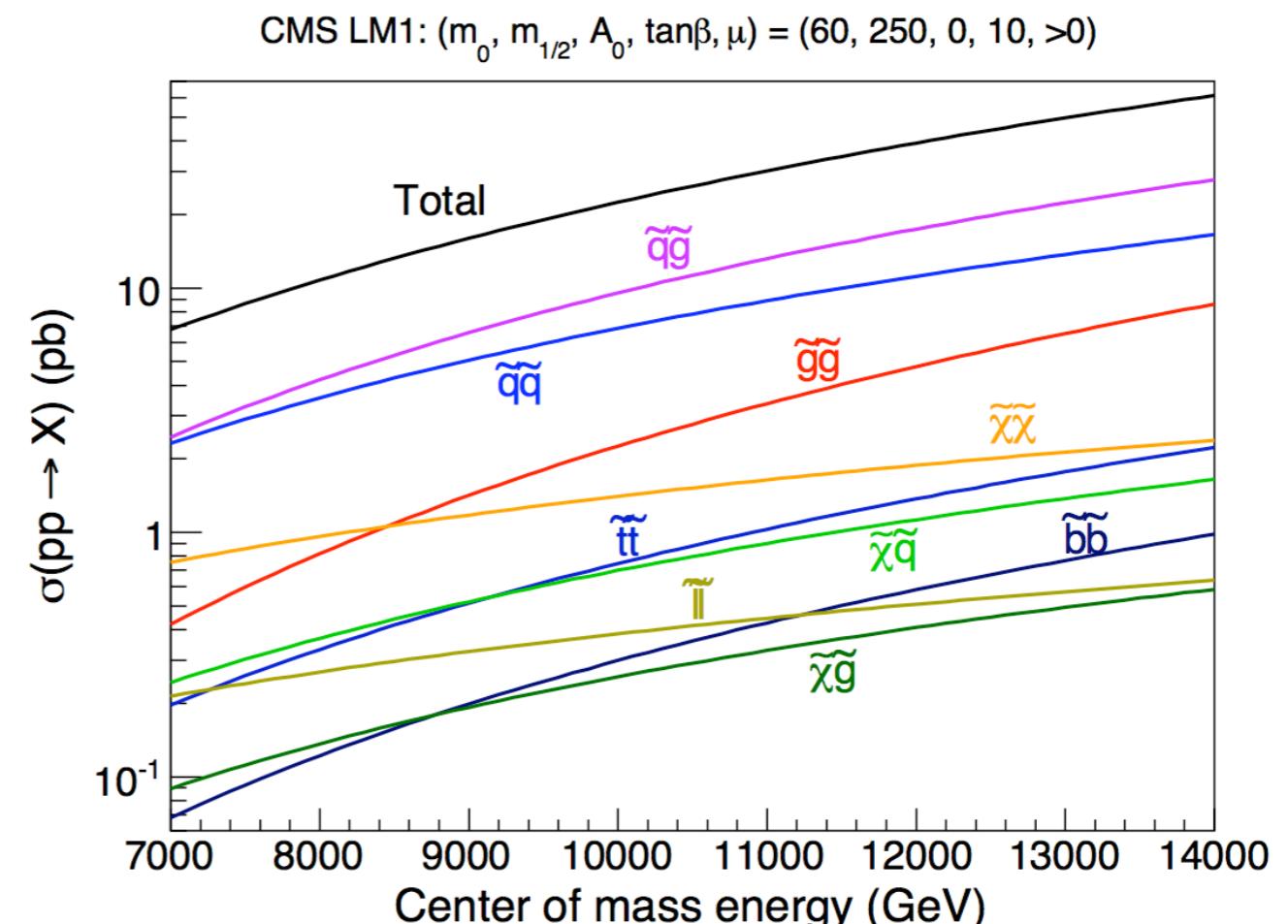
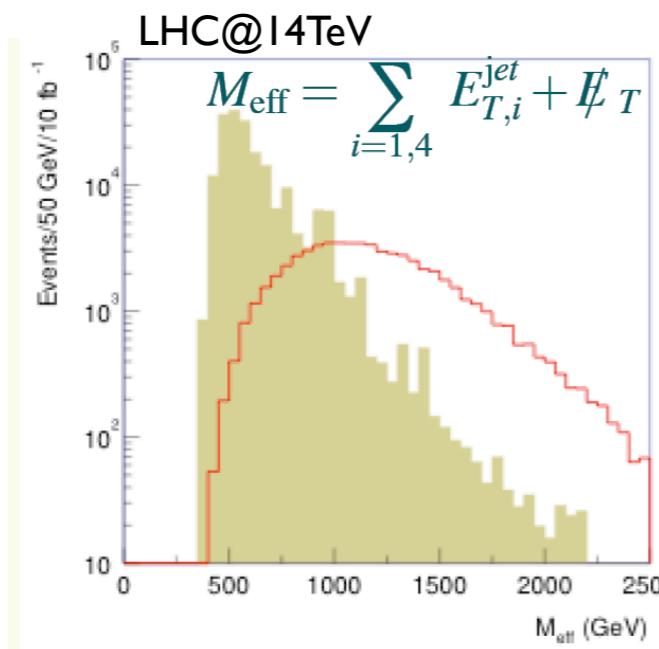
# **Beautiful early SUSY searches**

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Saclay, 29-31 March 2010**

# SUSY at LHC

- General expectation at LHC: large Xsections for squarks and/or gluinos. Strong interaction + the power of phase-space.
- Once produced, squarks/gluinos will decay into lighter sparticles until the LSP\* is reached
  - cascade decays
  - high-p<sub>T</sub> jets
  - large missing E<sub>T</sub>



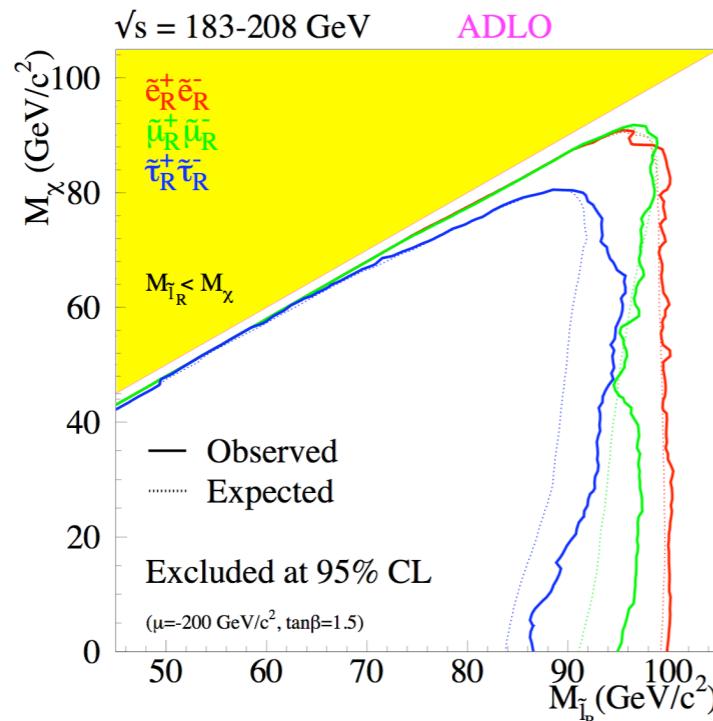
courtesy S. Sekmen

\*) LSP = lightest SUSY particle, stable if R-parity is conserved

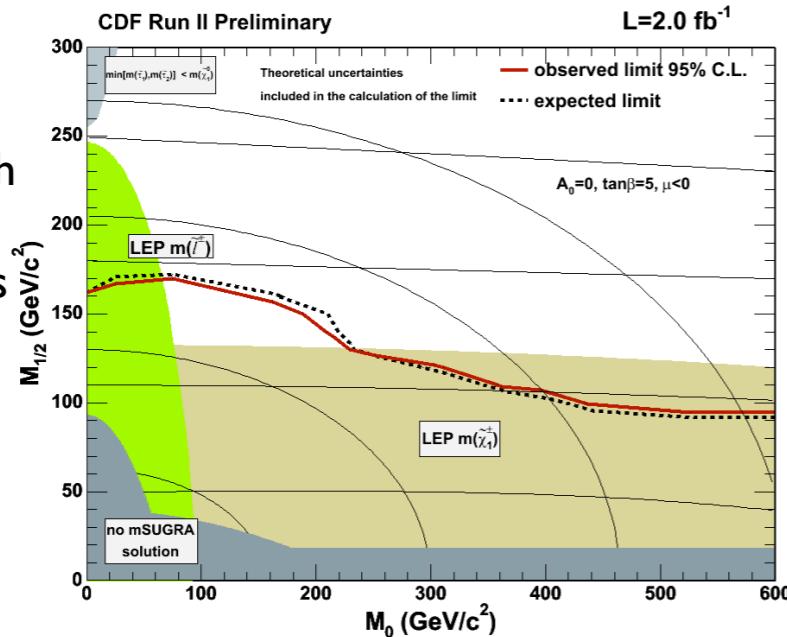
# Limits and constraints

So far only lower mass limits and indirect constraints, e.g.,

LEP: charged  
sparticle  
 $M \gtrsim 100$  GeV

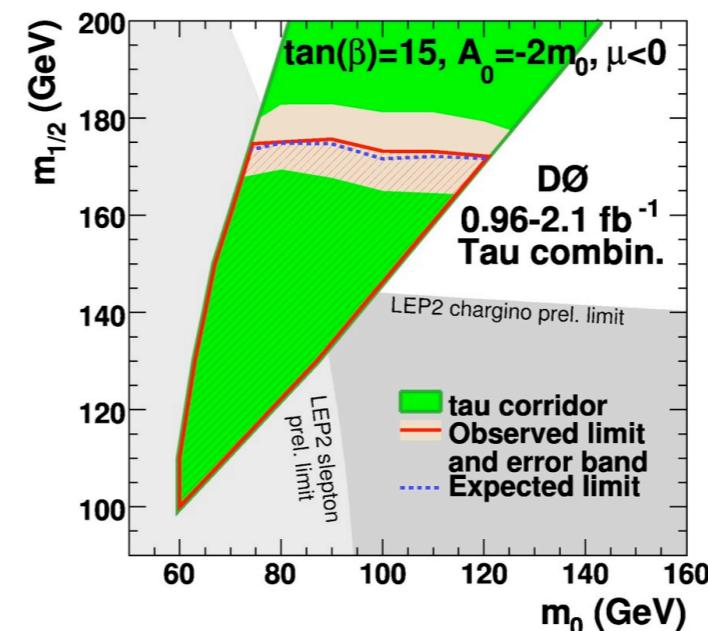


Tevatron:  
begins to reach  
beyond LEP  
but mass limits  
quite model  
dependent



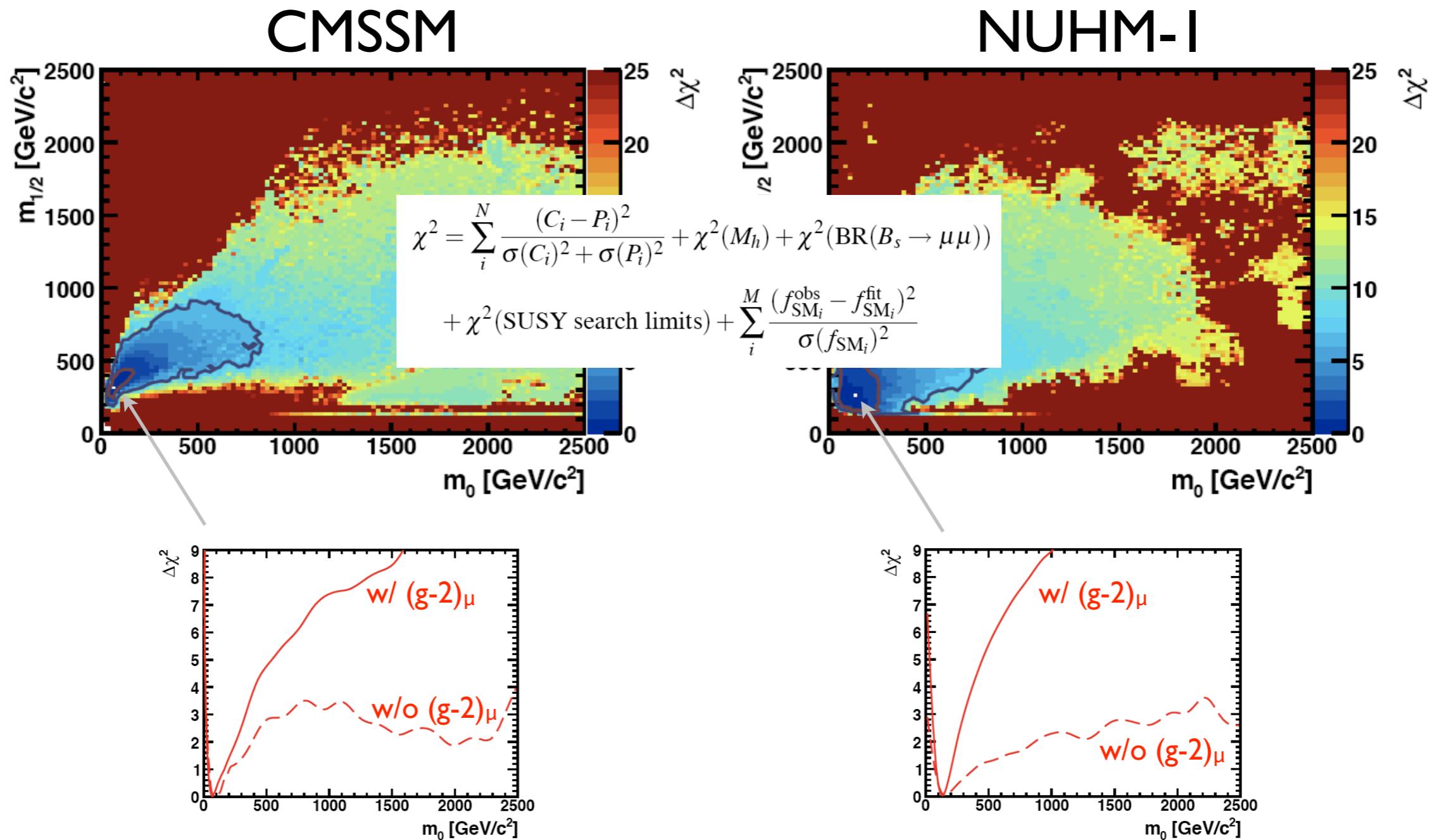
$\text{BR}_{b \rightarrow s\gamma}^{\text{exp}}/\text{BR}_{b \rightarrow s\gamma}^{\text{SM}}$	$1.117 \pm 0.076_{\text{exp}} \pm 0.082_{\text{th(SM)}}$
$\text{BR}(B_s \rightarrow \mu^+ \mu^-)$	$< 4.7 \times 10^{-8}$
$\text{BR}_{B \rightarrow \tau\nu}^{\text{exp}}/\text{BR}_{B \rightarrow \tau\nu}^{\text{SM}}$	$1.25 \pm 0.40_{[\text{exp+th}]}$
$\text{BR}(B_d \rightarrow \mu^+ \mu^-)$	$< 2.3 \times 10^{-8}$
$\text{BR}_{B \rightarrow X_s \ell\ell}^{\text{exp}}/\text{BR}_{B \rightarrow X_s \ell\ell}^{\text{SM}}$	$0.99 \pm 0.32$
$\text{BR}_{K \rightarrow \mu\nu}^{\text{exp}}/\text{BR}_{K \rightarrow \mu\nu}^{\text{SM}}$	$1.008 \pm 0.014_{[\text{exp+th}]}$
$\text{BR}_{K \rightarrow \pi\nu\bar{\nu}}^{\text{exp}}/\text{BR}_{K \rightarrow \pi\nu\bar{\nu}}^{\text{SM}}$	$< 4.5$
$\Delta M_{B_s}^{\text{exp}}/\Delta M_{B_s}^{\text{SM}}$	$0.97 \pm 0.01_{\text{exp}} \pm 0.27_{\text{th(SM)}}$
$(\Delta M_{B_s}^{\text{exp}}/\Delta M_{B_s}^{\text{SM}})/(\Delta M_{B_d}^{\text{exp}}/\Delta M_{B_d}^{\text{SM}})$	$1.00 \pm 0.01_{\text{exp}} \pm 0.13_{\text{th(SM)}}$
$\Delta \epsilon_K^{\text{exp}}/\Delta \epsilon_K^{\text{SM}}$	$1.08 \pm 0.14_{[\text{exp+th}]}$
$a_\mu^{\text{exp}} - a_\mu^{\text{SM}}$	$(30.2 \pm 8.8) \times 10^{-10}$
$M_h$ [GeV]	$> 114.4$ (see text)
$\Omega_{\text{CDM}} h^2$	$0.1099 \pm 0.0062$

B-physics!



→ severe constraints  
on parameter space

# Fits to available data (frequentist)

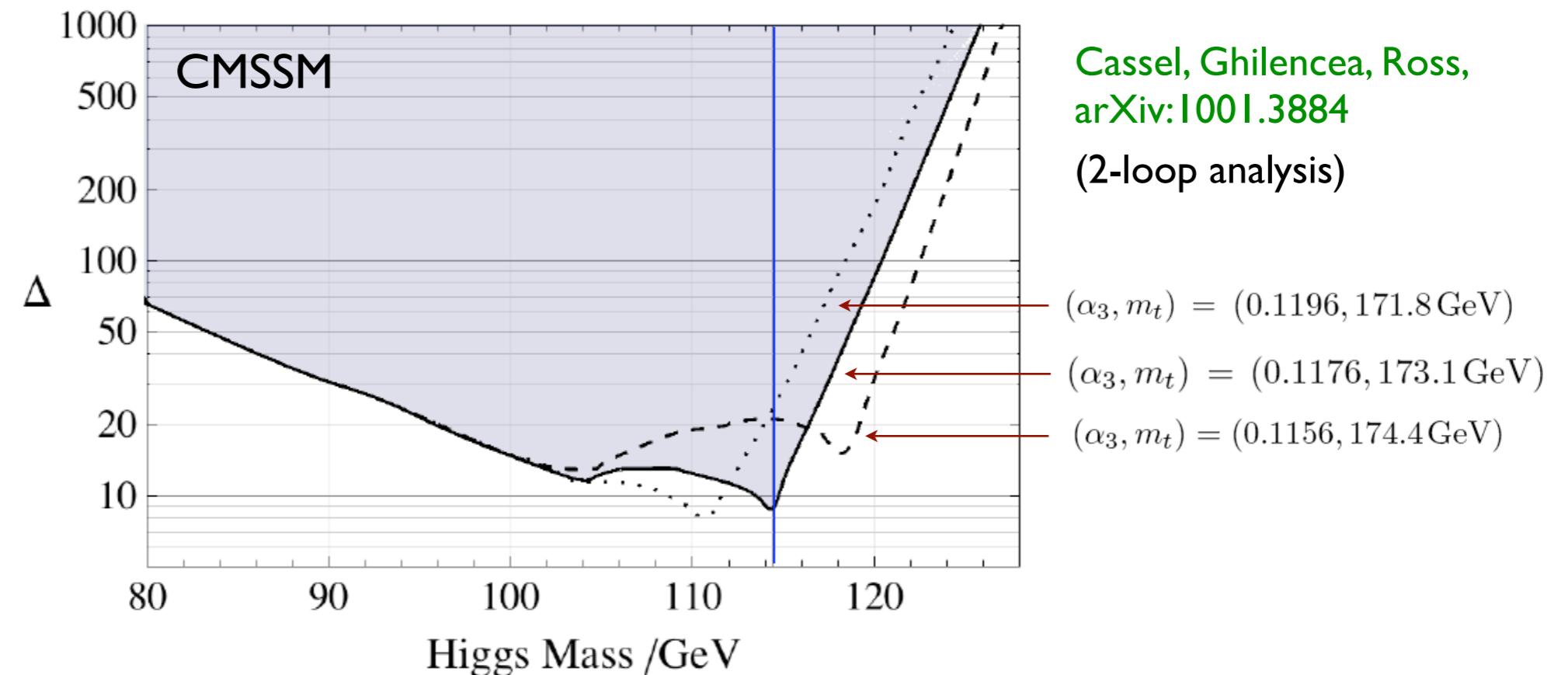


Best fit at  $m_{\tilde{g}} \approx 750/600 \text{ GeV}$ ,  $\tan \beta \approx 11$ .

# The finetuning price

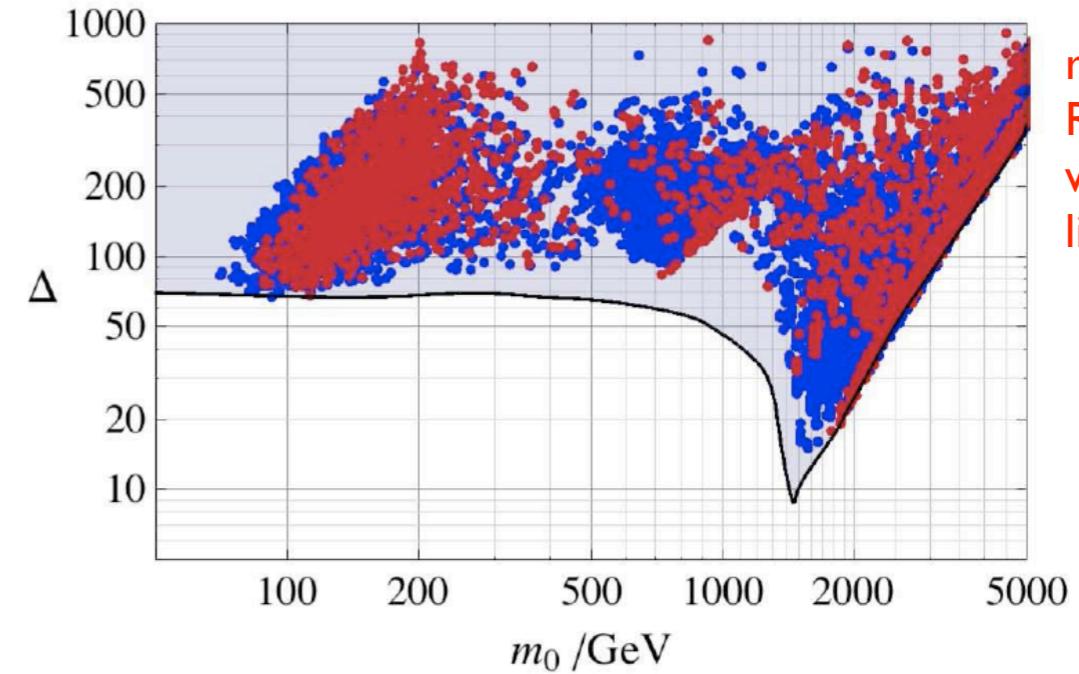
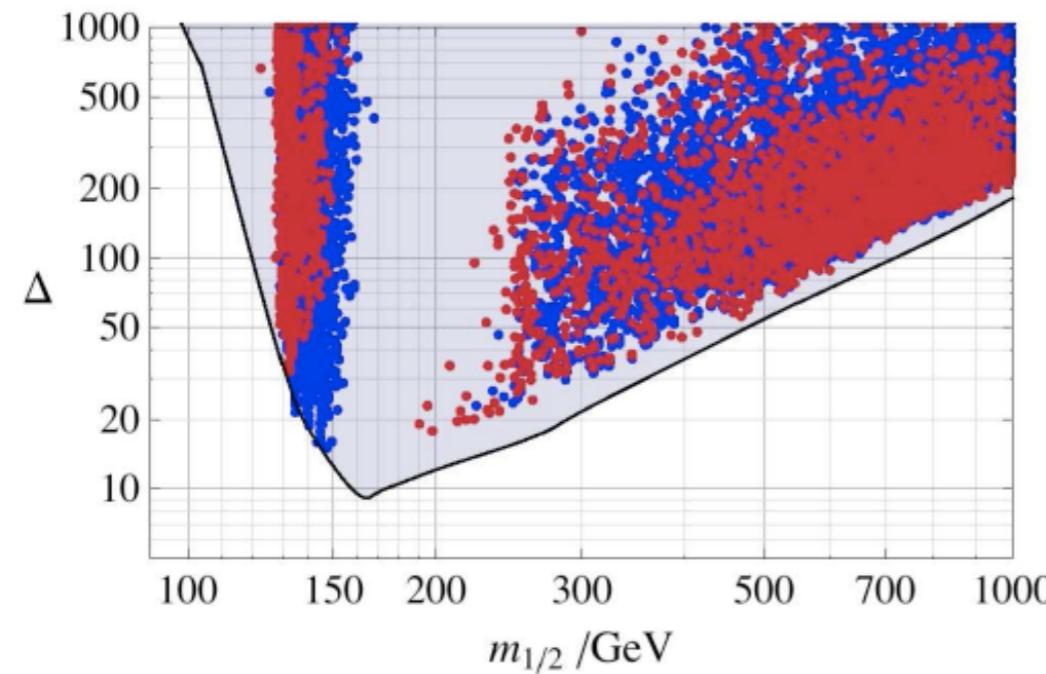
- Finetuning = sensitivity of EW scale to input parameters

$$\frac{M_Z^2}{2} \approx -m_{H_2}^2 - |\mu|^2$$



$$\Delta \equiv \max \left| \Delta_p \right|_{p=\{\mu_0^2, m_0^2, m_{1/2}^2, A_0^2, B_0^2\}}, \quad \Delta_p \equiv \frac{\partial \ln v^2}{\partial \ln p}$$

# CMSSM low finetuning



NB: points with lowest finetuning lie in the focus point region

- gaugino-higgsino mixing
- light gluino
- Xsections a few pb at 7TeV LHC

$BR(\tilde{g} \rightarrow \tilde{\chi}_i^0 g) \sim 10 - 20\%$

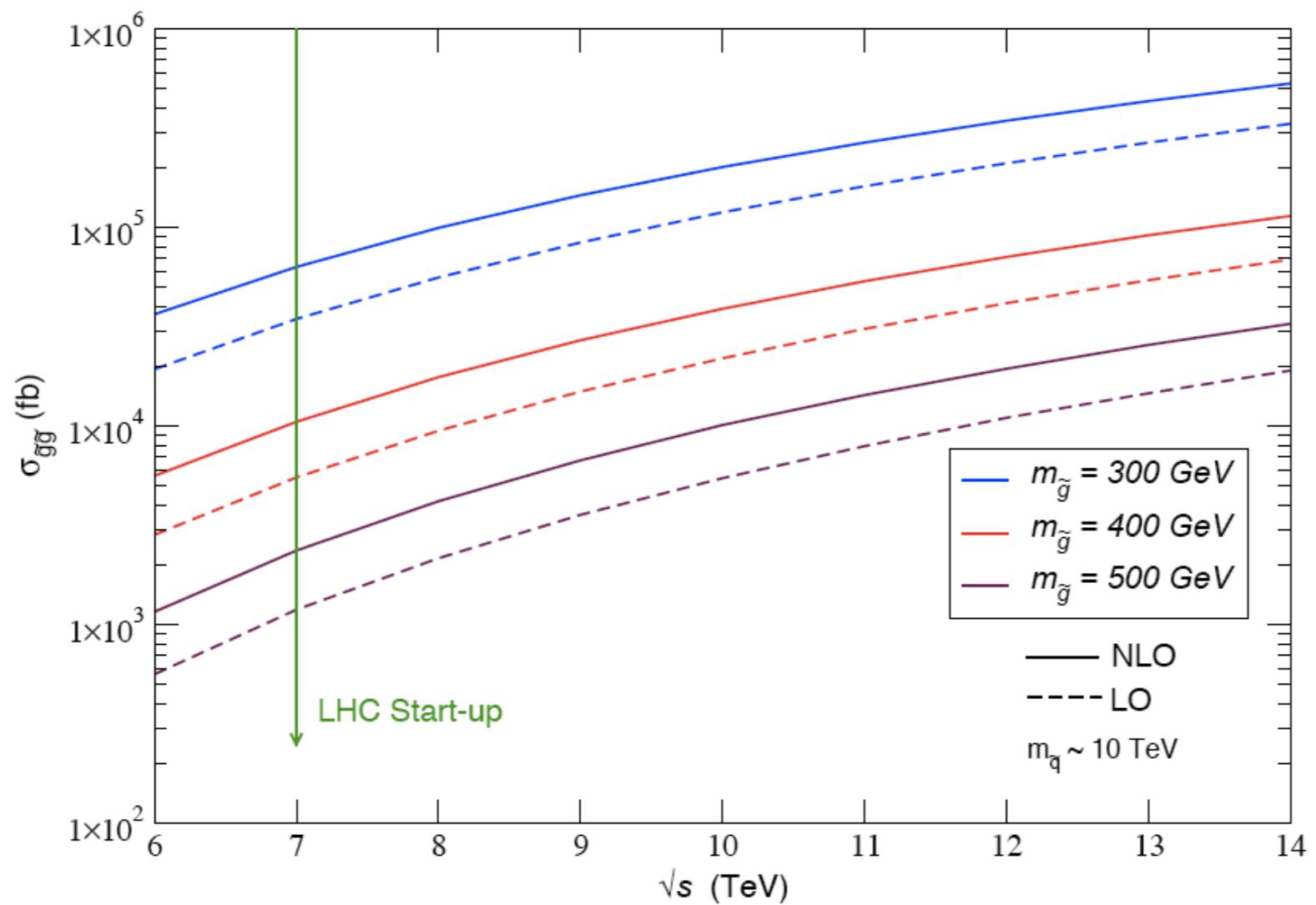
$BR(\tilde{g} \rightarrow \tilde{\chi}_i^0 b\bar{b}, \tilde{\chi}_i^\pm tb) \sim 20\%$

$h^0$	114.5	$\tilde{\chi}_1^0$	79	$\tilde{b}_1$	1147	$\tilde{u}_L$	1444
$H^0$	1264	$\tilde{\chi}_2^0$	142	$\tilde{b}_2$	1369	$\tilde{u}_R$	1446
$H^\pm$	1267	$\tilde{\chi}_3^0$	255	$\tilde{\tau}_1$	1328	$\tilde{d}_L$	1448
$A^0$	1264	$\tilde{\chi}_4^0$	280	$\tilde{\tau}_2$	1368	$\tilde{d}_R$	1446
$\tilde{g}$	549	$\tilde{\chi}_1^\pm$	142	$\tilde{\mu}_L$	1406	$\tilde{s}_L$	1448
$\tilde{\nu}_\tau$	1366	$\tilde{\chi}_2^\pm$	280	$\tilde{\mu}_R$	1406	$\tilde{s}_R$	1446
$\tilde{\nu}_\mu$	1404	$\tilde{t}_1$	873	$\tilde{e}_L$	1406	$\tilde{c}_L$	1444
$\tilde{\nu}_e$	1404	$\tilde{t}_2$	1158	$\tilde{e}_R$	1406	$\tilde{c}_R$	1446

upper limits  
for  $\Delta < 100$

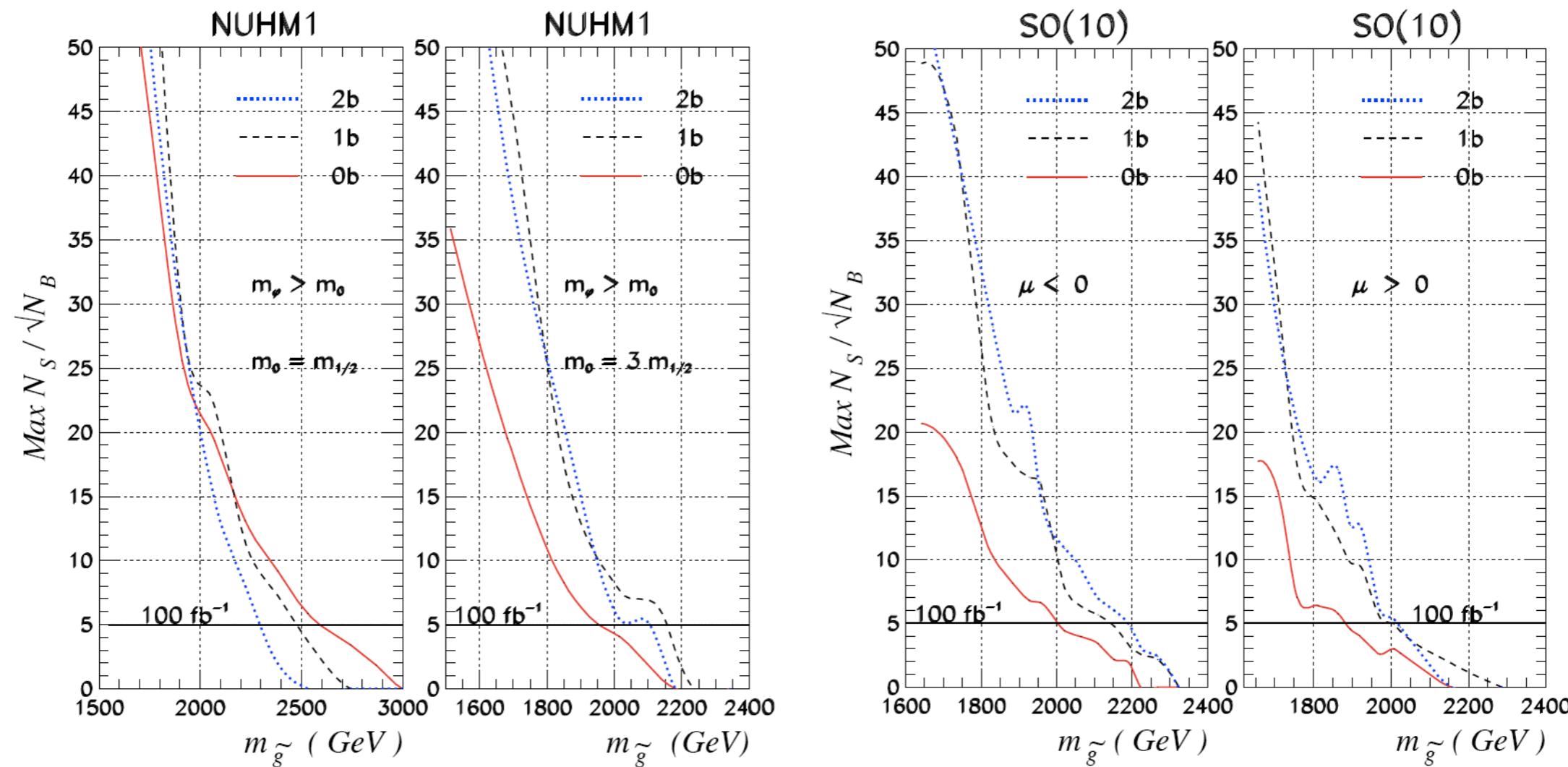
$\tilde{g}$	$\chi_1^0$	$\chi_2^0$	$\chi_3^0$	$\chi_4^0$	$\chi_1^\pm$	$\chi_2^\pm$	$\tilde{t}_1$	$\tilde{t}_2$	$\tilde{b}_1$	$\tilde{b}_2$
1720	305	550	660	665	550	670	2080	2660	2660	3140

# Light gluino: promising for LHC at 7 TeV



# Importance of b-tagging

- Requiring 1, 2, or more b-jets can significantly enhance the signal/bg in certain scenarios, e.g., 15-20% in the CMSSM focus point region.

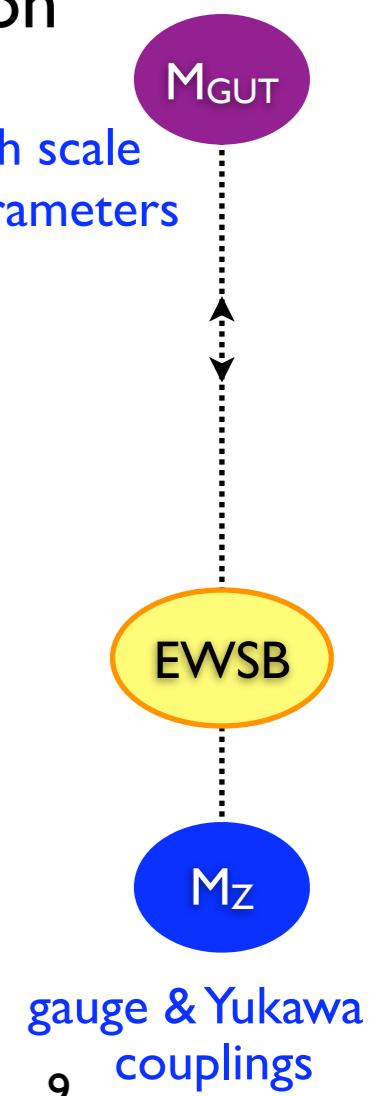


- Typical if 3rd generation is lighter than 1st/2nd gen. and  $m_{\tilde{g}} \ll m_{\tilde{q}}$  ; enhances gluino decays into t or b via on- or off-shell stop/sbottom

# Yukawa-unified SUSY

- SUSY GUTs based on SO(10) are particularly compelling
  - unify all matter of one generation in a 16-plet (incl. r.h. neutrino!)
  - automatic anomaly cancellation
- In the simplest realization the Higgs doublets reside in a 10-plet. This then requires t-b-tau Yukawa coupling unification in addition to gauge coupling unification at  $M_{\text{GUT}}$ .
- Parameter space:
  - common gaugino mass  $m_{1/2}$
  - common sfermion mass parameter  $m_{16}$
  - common Higgs mass parameter  $m_{10}$
  - common trilinear coupling  $A_0$
  - $\tan\beta$  and  $\text{sign}(\mu)$
  - D-term contribution  $M_D^2$  from SO(10) breaking

$$m_{H_{u,d}}^2 = m_{10}^2 \mp M_D^2$$



# Conditions for Yukawa unification (YU)

★ For  $\mu > 0$ , as preferred by  $b \rightarrow s\gamma$ , Yukawa unification (YU) can only be realized for very particular parameter relations

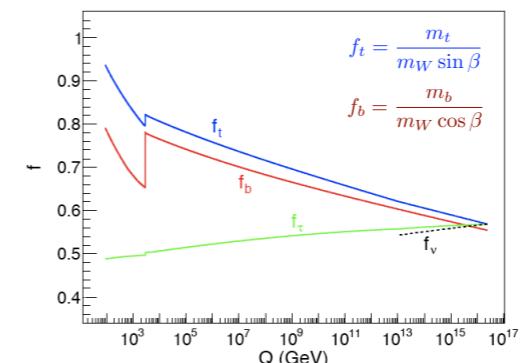
- $m_{16} \sim 5 - 15$  TeV,
- $A_0^2 \simeq 2m_{10}^2 \simeq 4m_{16}^2$ ,  $(A_0 < 0)$
- $m_{1/2} \ll m_{16}$ ,
- $\tan \beta \sim 50$ .

★ D-term splitting

$$\begin{aligned} m_Q^2 = m_E^2 = m_U^2 &= m_{16}^2 + M_D^2 \\ m_D^2 = m_L^2 &= m_{16}^2 - 3M_D^2 \\ m_{\tilde{\nu}_R}^2 &= m_{16}^2 + 5M_D^2 \\ m_{H_{u,d}}^2 &= m_{10}^2 \mp 2M_D^2. \end{aligned}$$

“just-so” Higgs splitting (HS) case

NB: we need  $m_{H_u}^2 < m_{H_d}^2$  at  $M_{\text{GUT}}$ , so  $M_D^2 > 0$ .



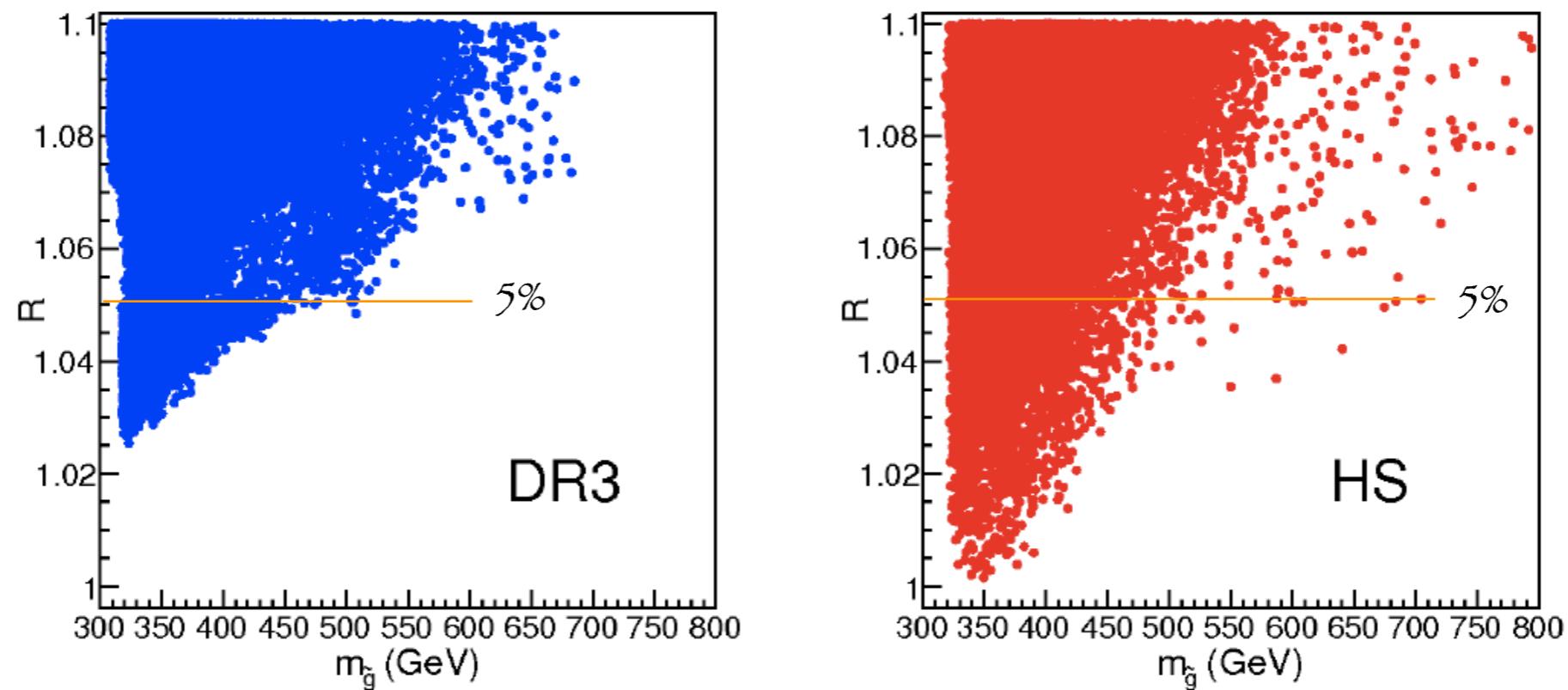
$$R = \frac{\max(f_t, f_b, f_\tau)}{\min(f_t, f_b, f_\tau)}$$

- D-term splitting w/o RHN gives  $R \sim 1.08$  (i.e. 8% unification)
- Splitting of only  $m_H$ 's (“just-so HS”) allows for  $R \sim 1.01$
- D-term splitting with RHN gives  $R \sim 1.04, \dots$
- ... but if we allow in addition small non-degeneracy of 3rd vs. 1st/2nd generation, we get  $R \sim 1.02$

# YU: Typical mass spectra

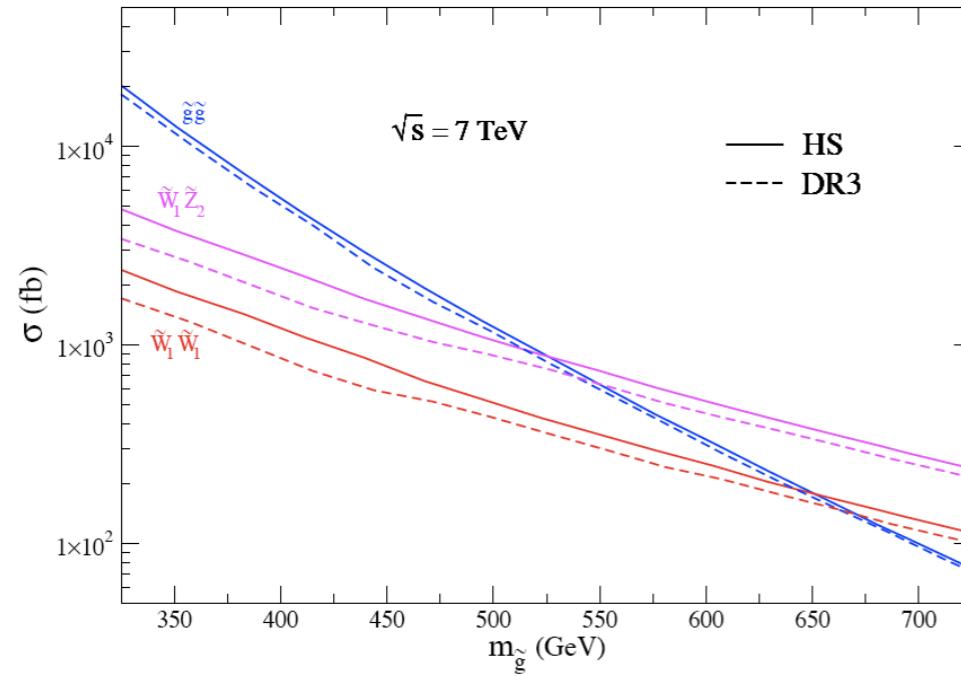
- 1st/2nd generation scalars in the multi-TeV range (5-15 TeV)
- 3rd gen. scalars, heavy Higgses and higgsinos in the 1-3 TeV range
- light gauginos: LSP  $\sim$  50-80 GeV, gluino  $\sim$  300-500 GeV
- c.f “effective SUSY” by Cohen, Kaplan, Nelson ’1996

$$R = \frac{\max(f_t, f_b, f_\tau)}{\min(f_t, f_b, f_\tau)}$$



Points from a MCMC scan for small R

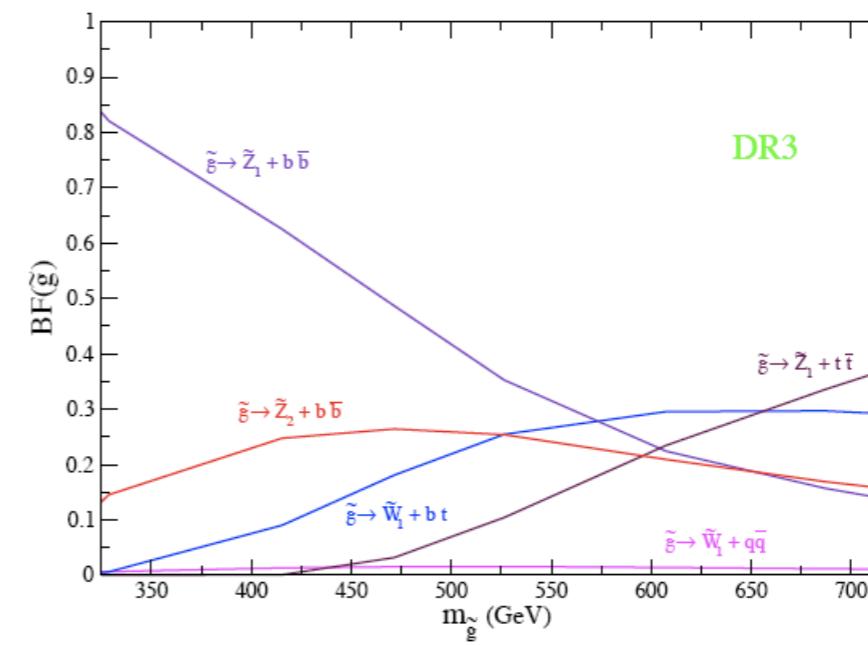
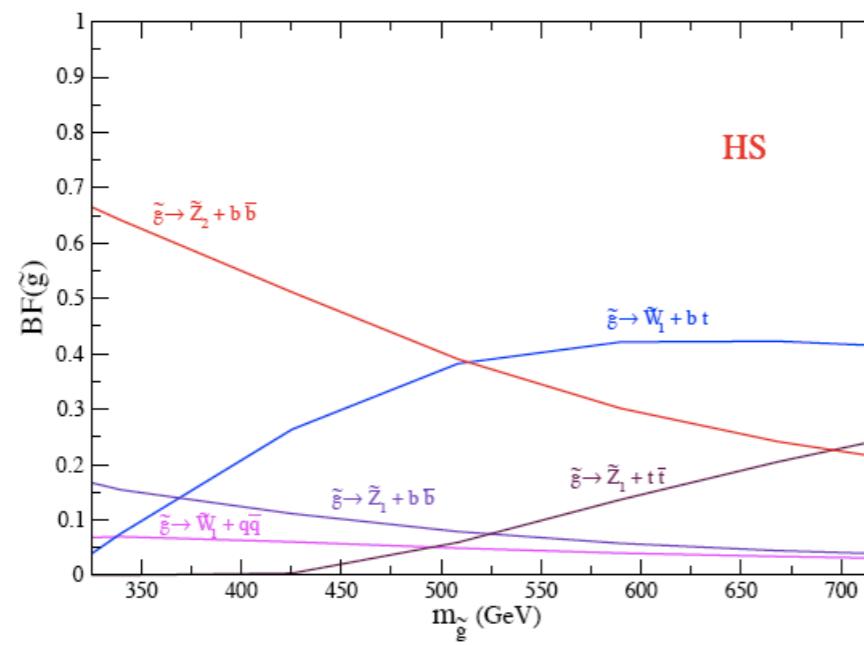
# LHC reach at 7 TeV



Gluino-pair prod. dominated by gg fusion.  
Much less enhancement from heavy squarks.  
 $\sigma(\text{LO}) \sim 1 \text{ pb}$  at  $m(\text{gluino}) \sim 525 \text{ GeV}$

We consider model lines for HS and DR3 cases as function of  $m(\text{gluino})$  up to 700 GeV.

Gluinos decays are again dominated by heavy flavours:  $\tilde{g} \rightarrow \tilde{\chi}_{1,2}^0 b\bar{b}$ ,  $\tilde{\chi}_1^\pm tb$



# LHC reach at 7 TeV

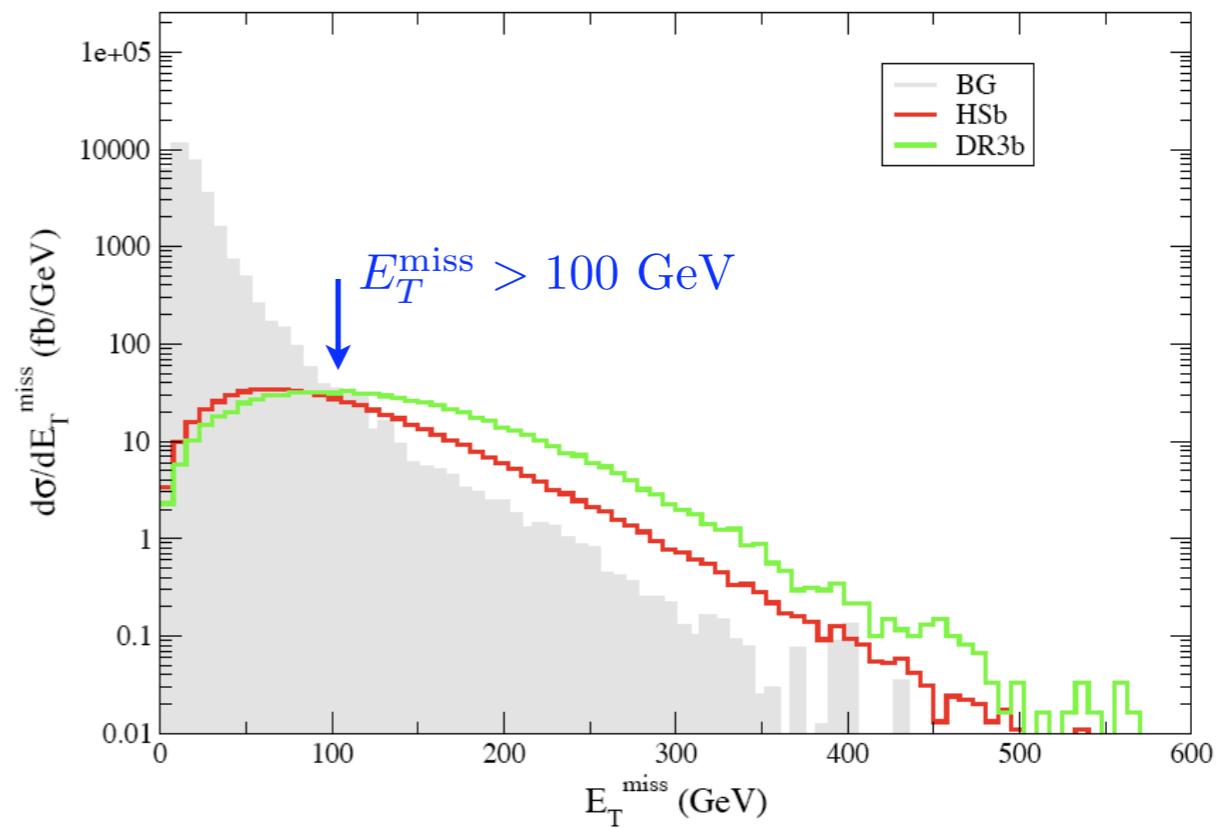
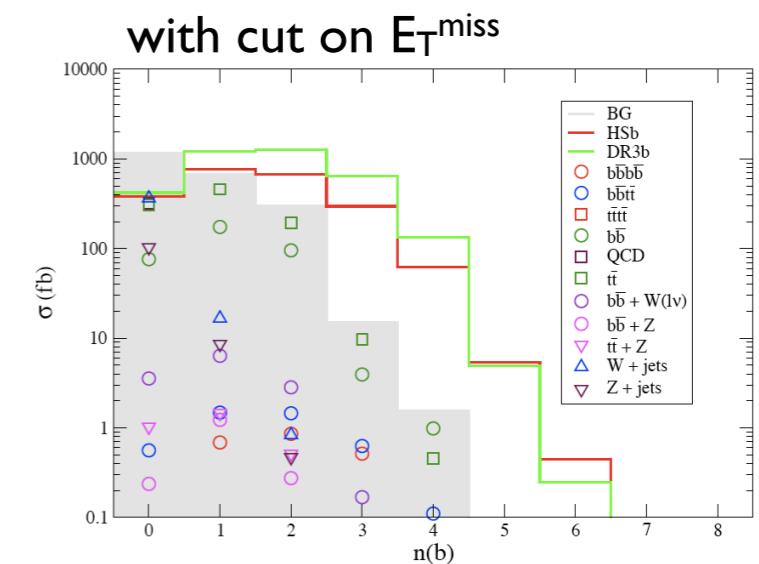
## Event simulation:

- Isajet 7.79 for the signal
- QCD, 2- and 3-bdy BGs with Alpgen
- 4t, 4b, 2t2b BGs with Madgraph
- Phythia for showering and hadronization
- Generic toy detector simulation

## Basic Cuts “C0”:

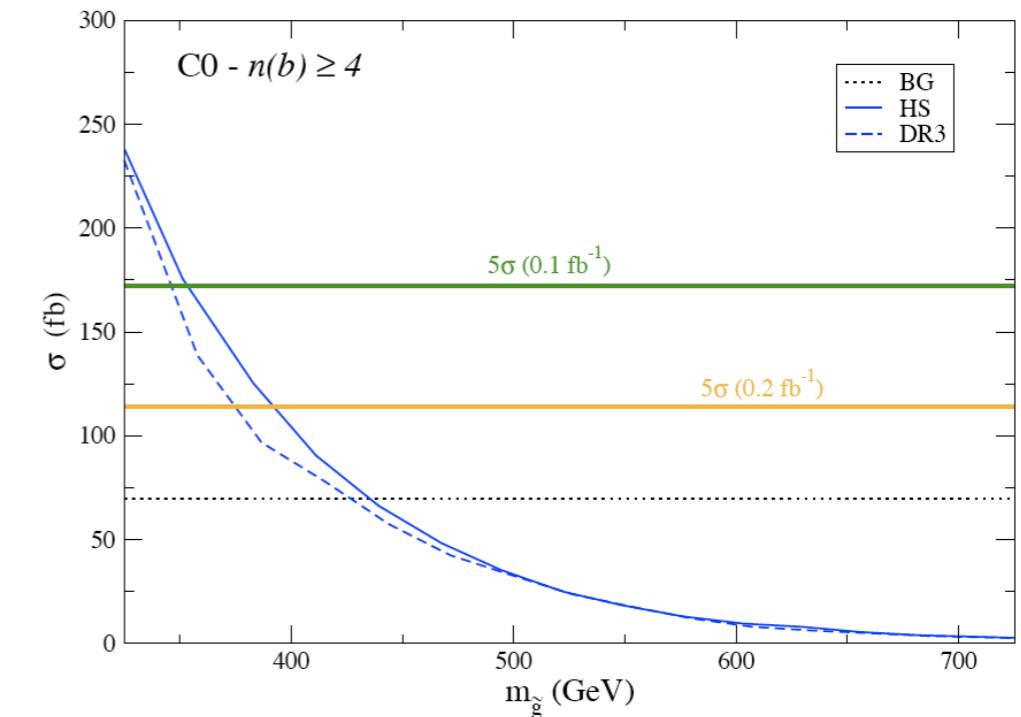
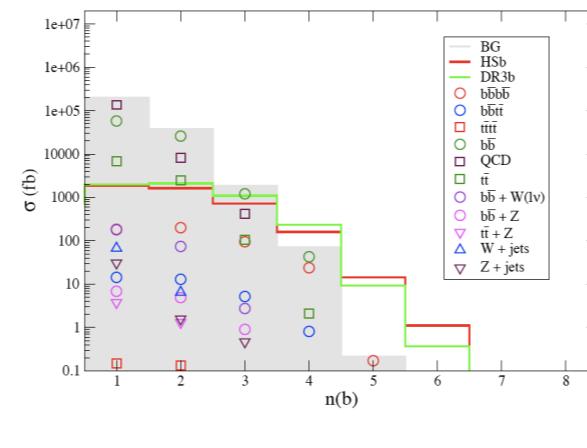
- $n(\text{jets}) \geq 4$  with  $p_T > 50\text{GeV}$
- hardest jet  $p_T > 100\text{ GeV}$
- $S_T \geq 0.2$  (transv sphericity)
- $n(b) \geq 1$  (b-eff. 60%)

Results after C1-based selection			
	$\sigma(n(b) \geq 3)$	$\sigma(n(b) \geq 4)$	$\sigma(\text{OS})$
HSb	364 fb	68 fb	81 fb
DR3b	782 fb	139 fb	23 fb
BG	16 fb	2 fb	9 fb

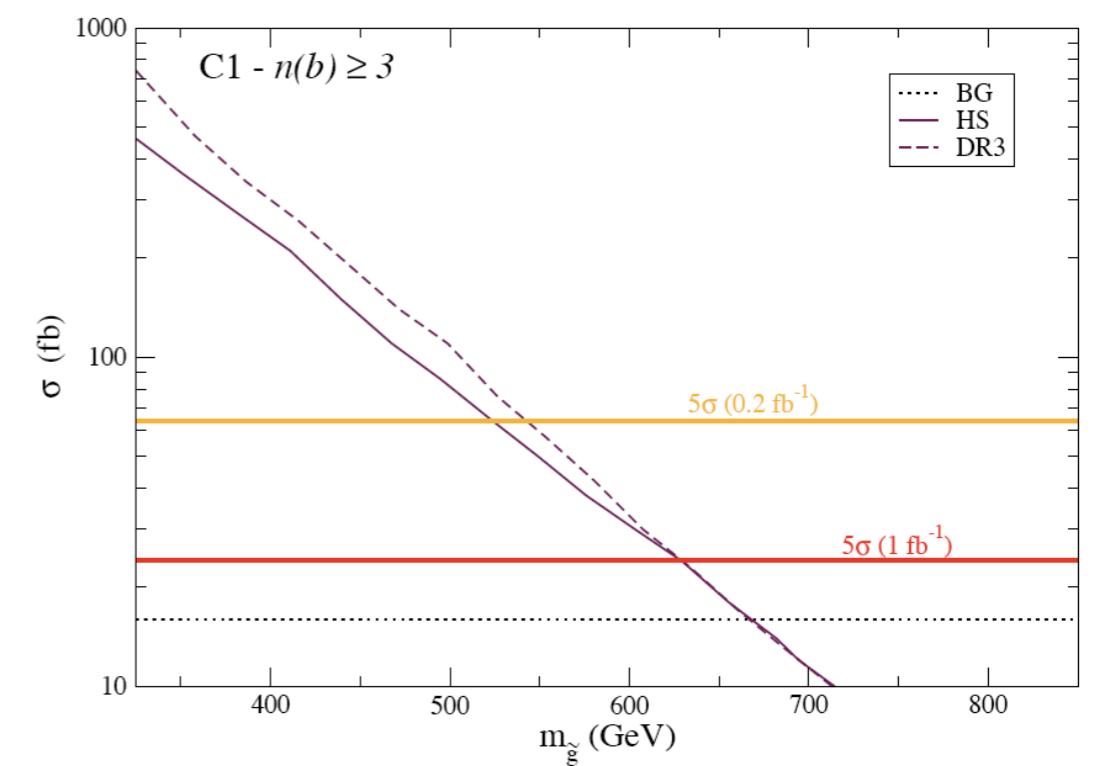
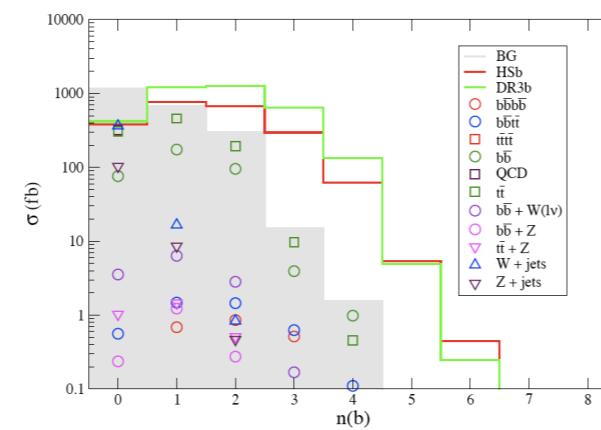


# LHC reach at 7 TeV

Without missing energy measurement:  
up to  $m(\text{gluino})=400$  GeV with  $0.2 \text{ fb}^{-1}$  of data  
requiring 4 b-jets

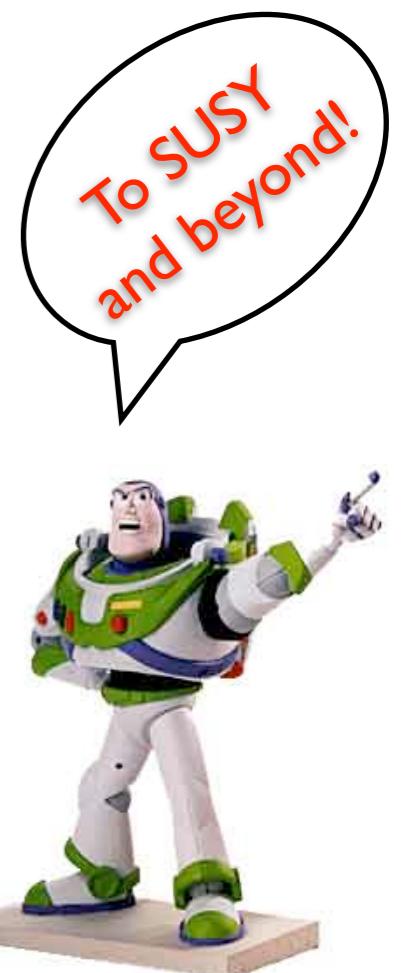


With reliable missing energy measurement:  
reach up to  $m(\text{gluino})=540-630$  GeV  
with  $0.2-1 \text{ fb}^{-1}$  of data,  
 $n(b) \geq 3$



# Conclusions

- Many well-motivated SUSY scenarios feature light gluinos, often in combination with heavy scalars, e.g.,
  - Focus point SUSY
  - Low finetuning scenarios
  - Yukawa-unified SUSY GUTs based on SO(10)
  - Effective SUSY
- Promising potential for LHC @ 7 TeV
- Gluinos often decay into heavy flavours ....  
Search in multi-b channels may essential for early discovery.



There are exciting times ahead of us