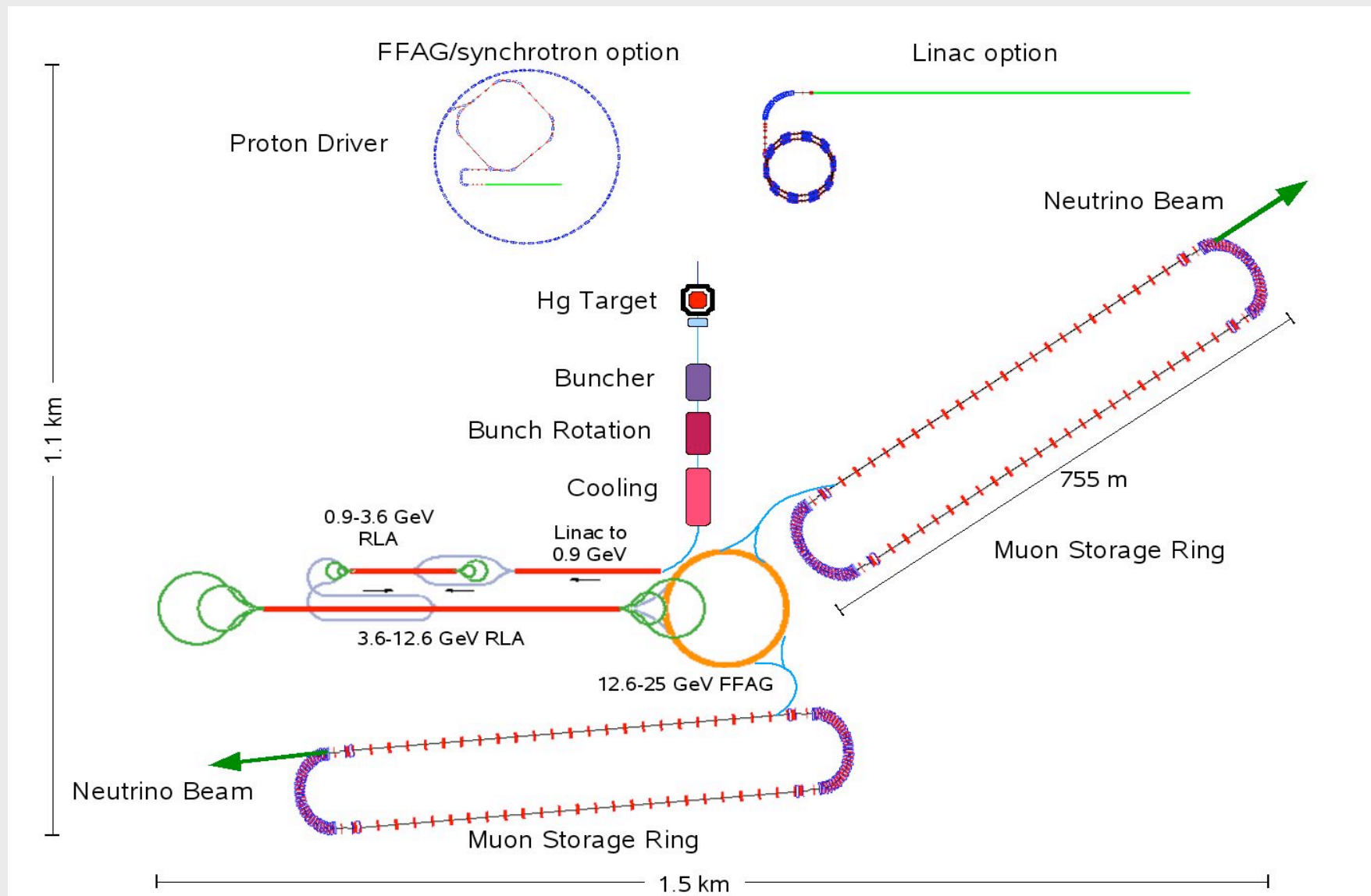




WP3 summary



Overview





Proton driver



R&D for the proton driver is decoupled from IDS as a hosting lab specific solution is assumed,.....but required beam parameters on target have been defined. Within Eurov the proton driver is part of the super beam work package. As the proton driver costing is strongly related to the hosting lab the following conveners have agreed to contribute:

CERN LINAC 4 / SPL : Roland Garoby

Fermilab Project X : Keith Gollwitzer

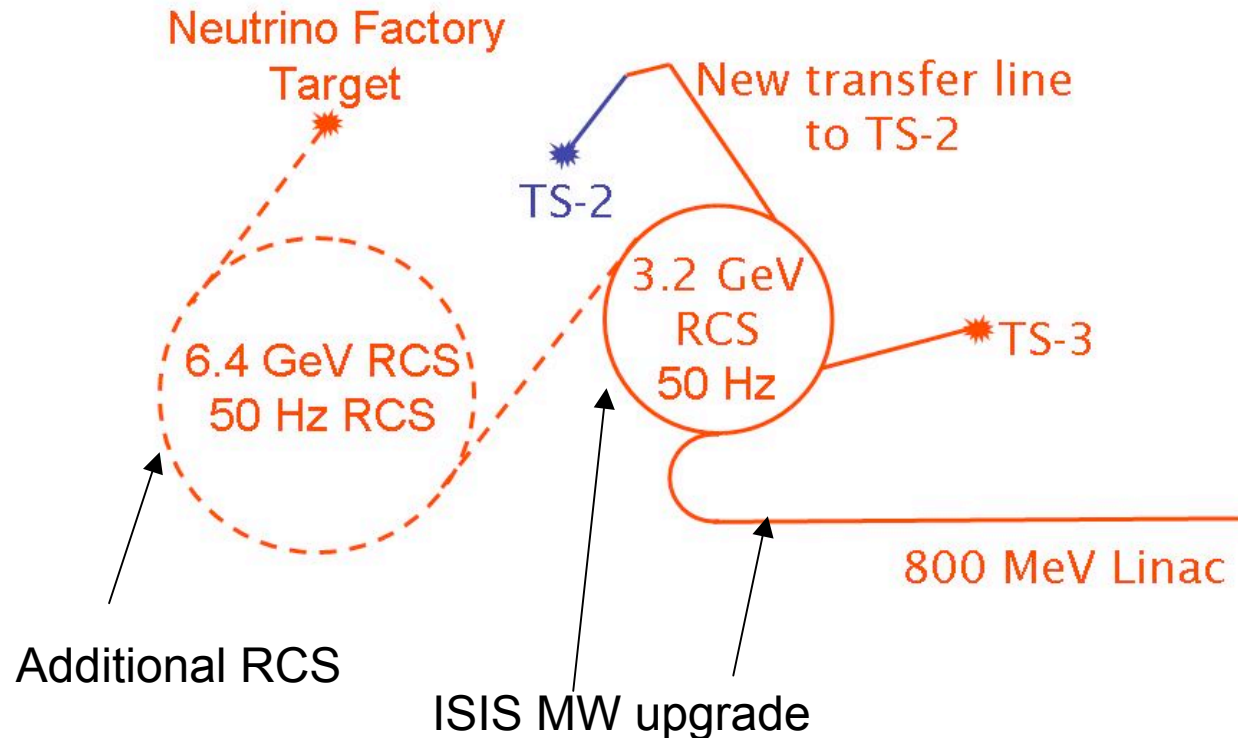
RAL - ISIS upgrade : John Thomason



NF- Proton driver and ISIS upgrade



Jaroslav Pasternak / John Thomason



- Based on MW ISIS upgrade with 0.8 GeV linac and 3.2 GeV RCS.
- Assumes a sharing of the beam power at 3.2 GeV between the two facilities
- Requires additional RCS machine in order to meet the power and energy needs of the Neutrino Factory
- Both facilities can have the same ion source, RFQ, chopper, linac, H⁻ injection, accumulation and acceleration to 3.2 GeV

J. Pozimski, Imperial College @ 2nd annual Eurov meeting 1-4 June 2010, Strasbourg



Proton driver at CERN



High power SPL could be the proton driver for a NF.
Lattice design for accumulator and bunch compressor available.

Detailed particle dynamics studies of accumulator and bunch compressor successfully performed.

3 bunch scenario is able to fulfil NF requirements.

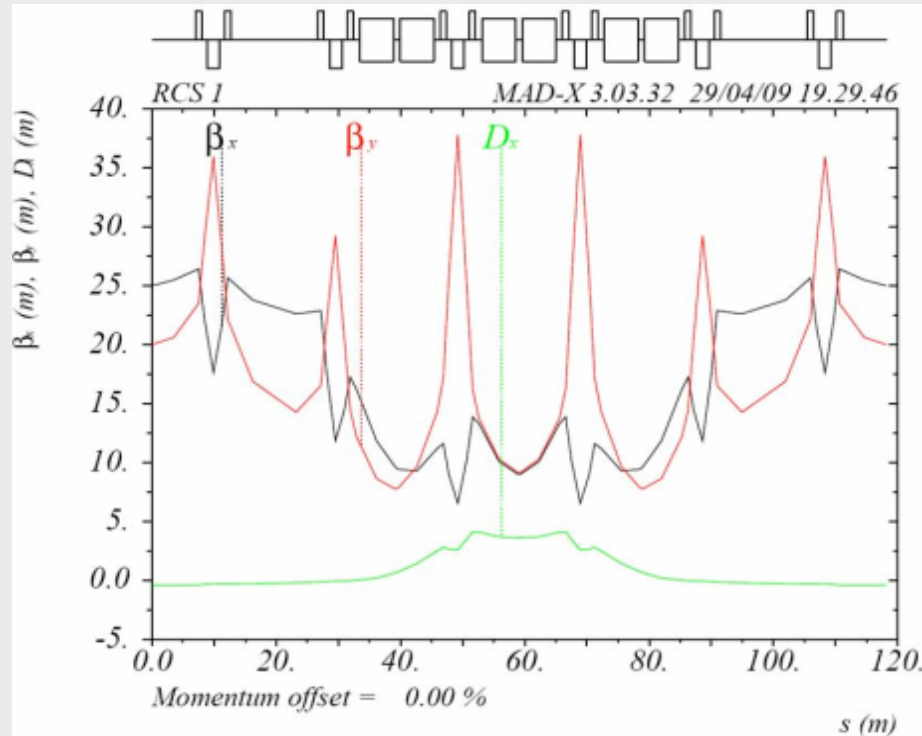
Costing for SPL finished, this has also to be performed for the rings.



Preliminary design of the second RCS



Jaroslav Pasternak



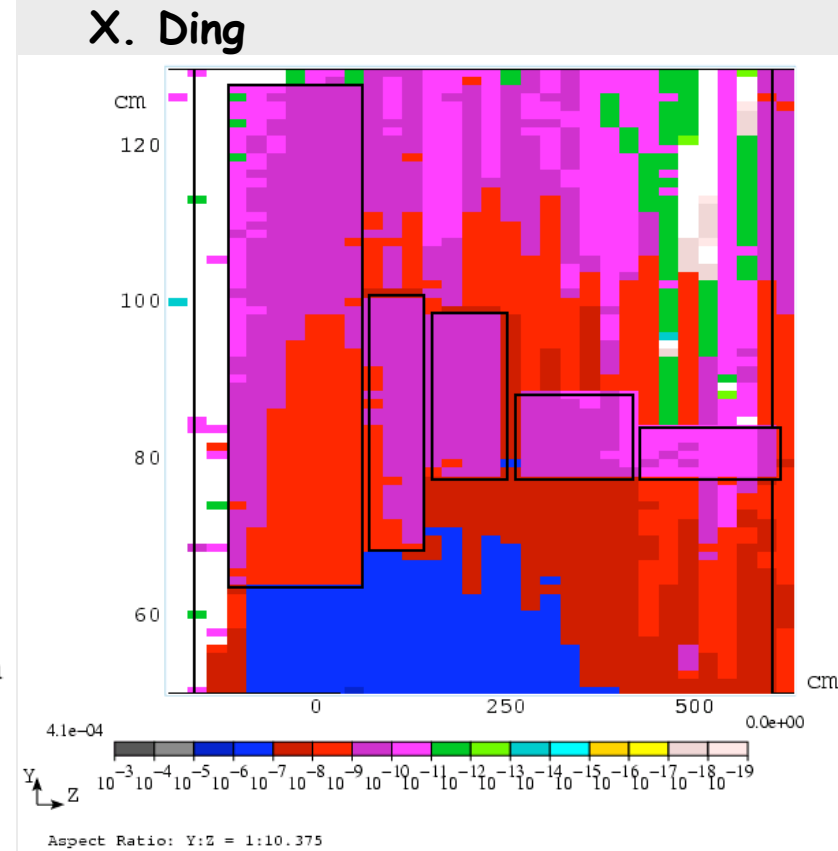
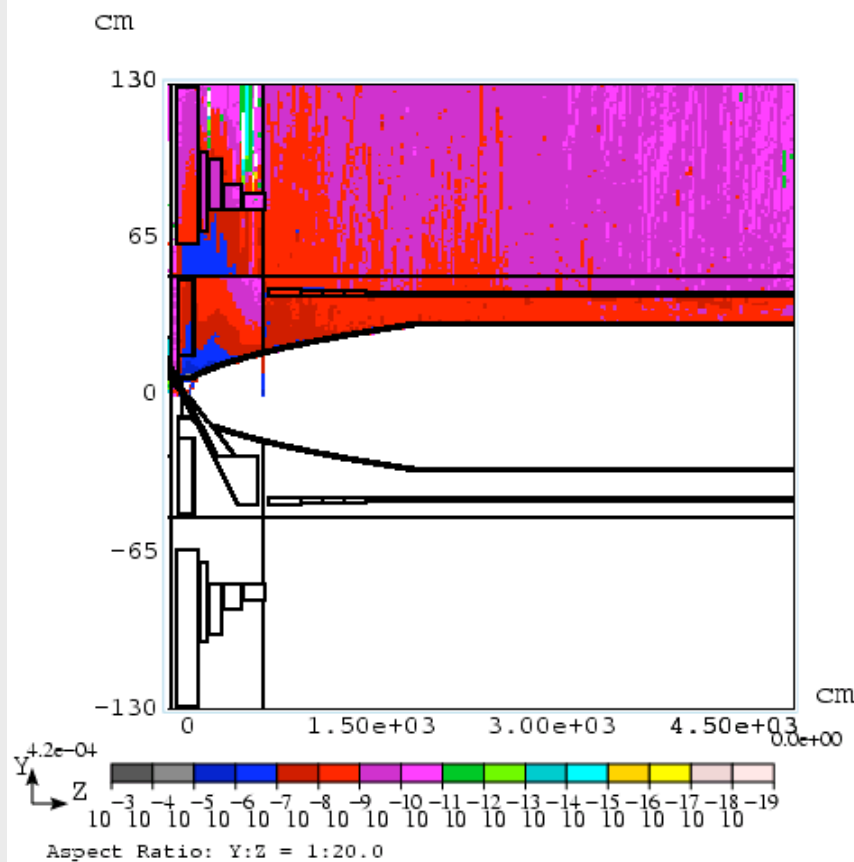
- Lattice may allow for flexibility in gamma transition choice (even with beam).
- Ring is overdesigned in order to allow for 10.3 GeV.
- Optimised solution for 6.4 GeV is in preparation!

Number of superperiods	6
Circumference	708.788 m
Harmonic number	6
RF frequency	2.4717-2.5289 MHz
Betatron tunes (Q_H, Q_V)	(7.81, 7.78)
Gamma transition	7.9056
Beam power at 6.4 GeV	4 MW for 2 bunches
Bunch area	1.8 eVs
$\Delta p/p$ at 3.2 GeV	$5.3 \cdot 10^{-3}$
Injection / extraction energy	3.2 / 6.4 [10.3] GeV
Repetition rate	50 Hz
Max B field in dipoles	1.2 T (at 10.3 GeV)
Length of long drift	12 m

Parameters of 6.4 (10.3) GeV RCS



Liquid mercury target -Energy deposition in the target area



Enhanced shield can decrease the power deposition in SC1 coil from 22.1kW to 4.8kW. By replacing the Res Sol by WC shield, the power deposition in SC1 coil can be decreased further to 1.3kW.



Previous Work -Jet Flow



- Free Jet + MHD + EDP

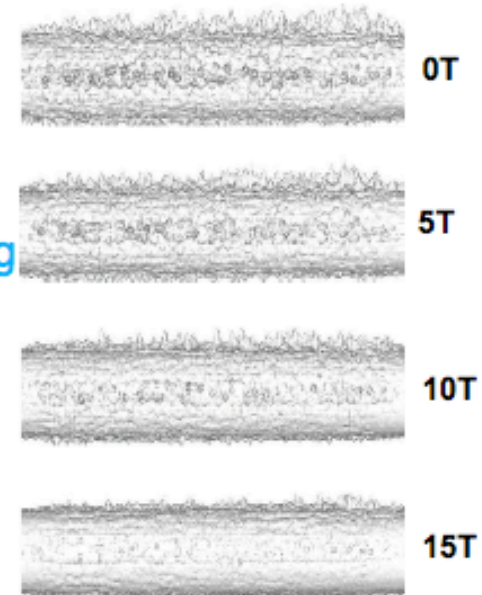
- Numerical Study

- R. Samulyak et al.

FronTier Code based on Front Tracking

- Surface instability
- MHD stabilizing effect
- Filament velocity in the simulations was about **25% smaller** than the experimental value

magnetic field	5T	10T	15T
experiments	54 m/s	50 m/s	35 m/s
simulations	36 m/s	27 m/s	22 m/s

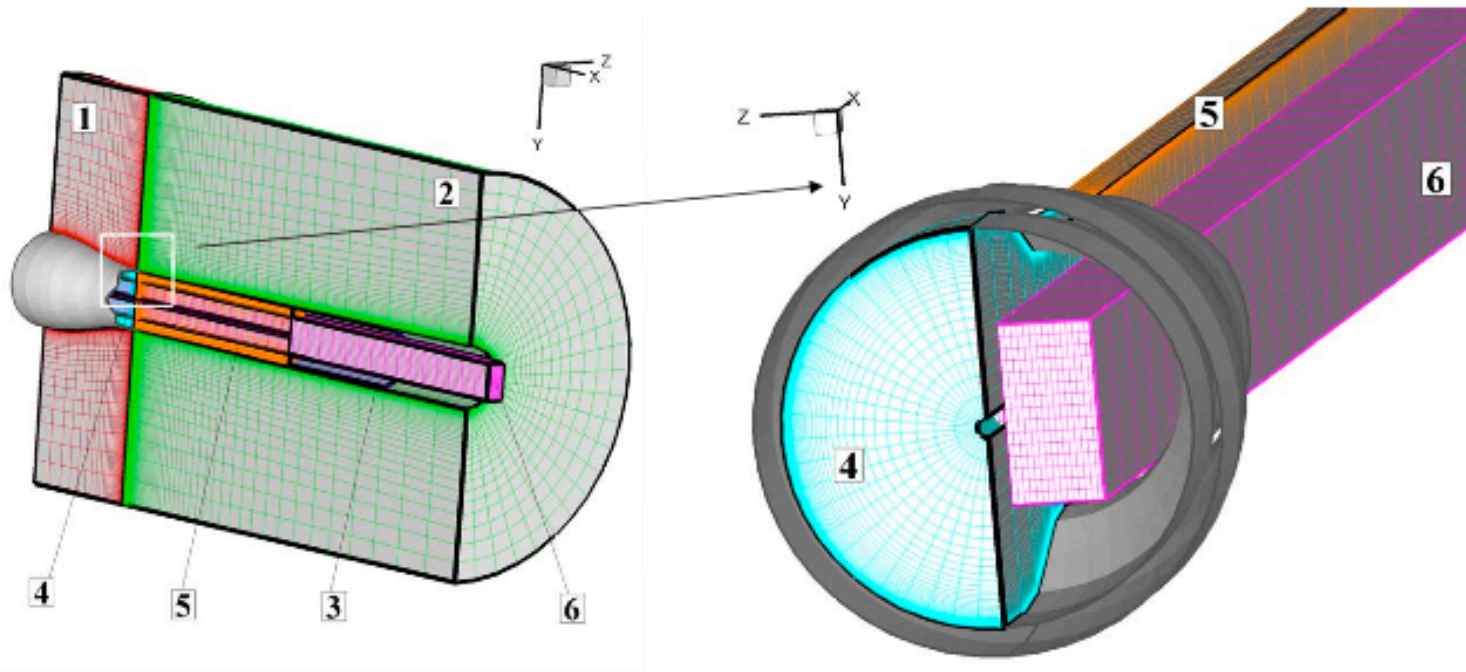


Surface filaments at 150 μs under longitudinal magnetic field



Our Numerical Approach

- CFD-Coupled, Hybrid Nozzle/Plume Calculation
 - Viscous internal flow + plume flow
 - Avoids need to specify jet exit conditions (unknown)

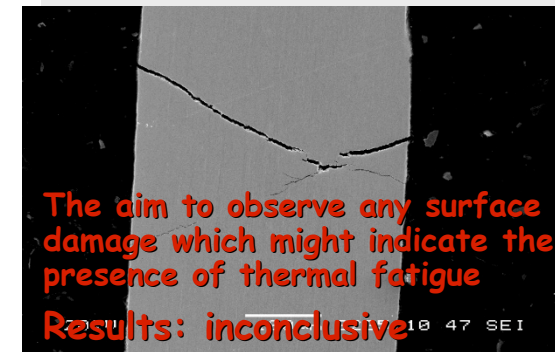
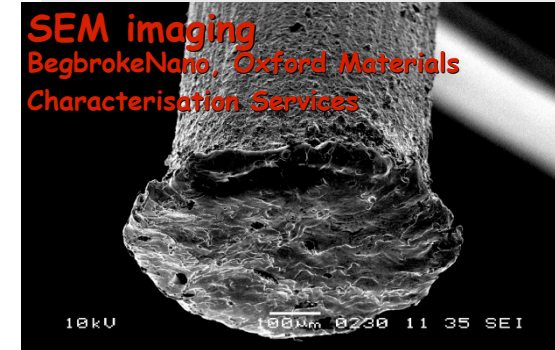
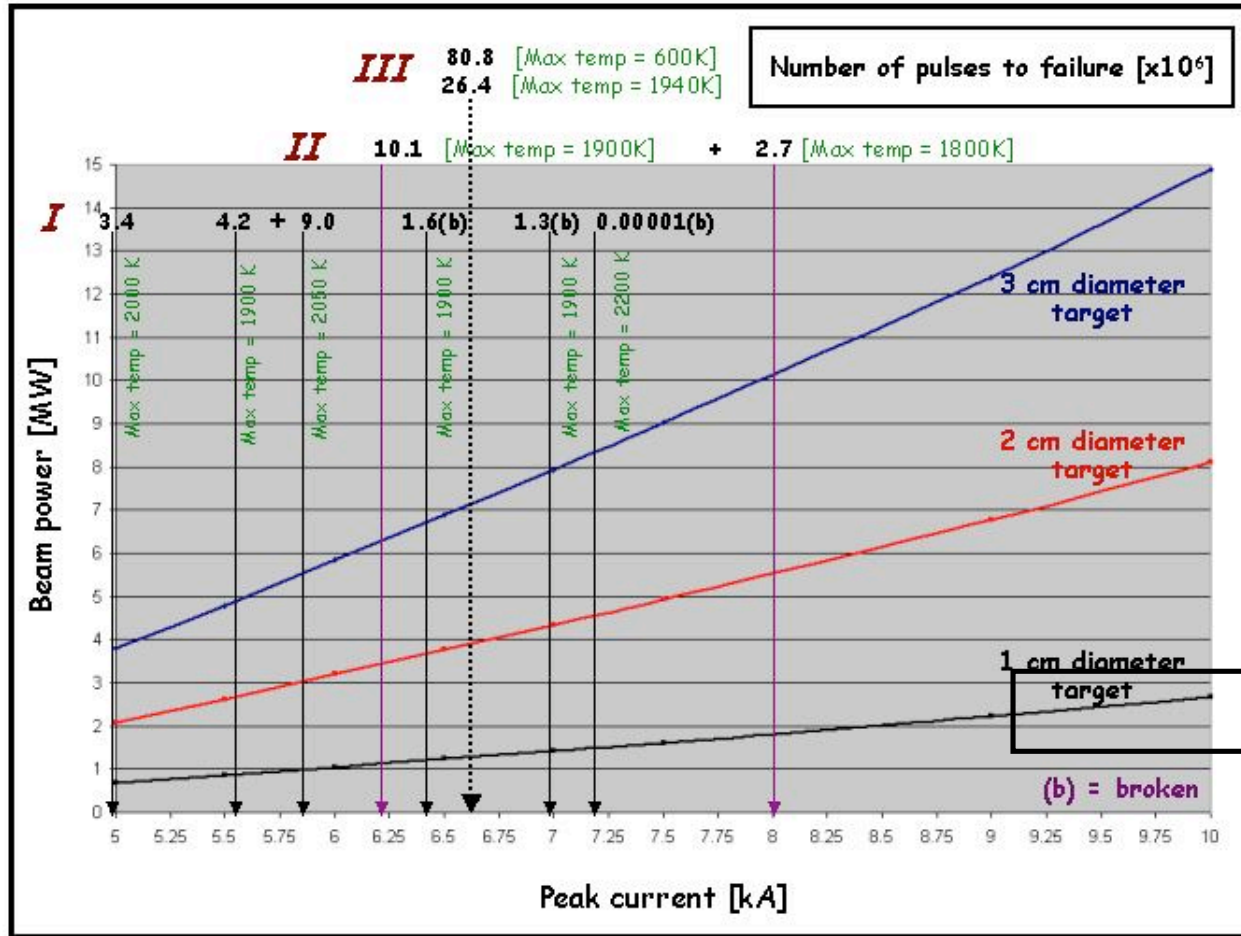




Solid target- lifetime tests



Rob Edgecock



> 10 years for 2cm diameter target
> 20 years for 3cm diameter target

Focus now:
Measure stress;
Confirm modelling.

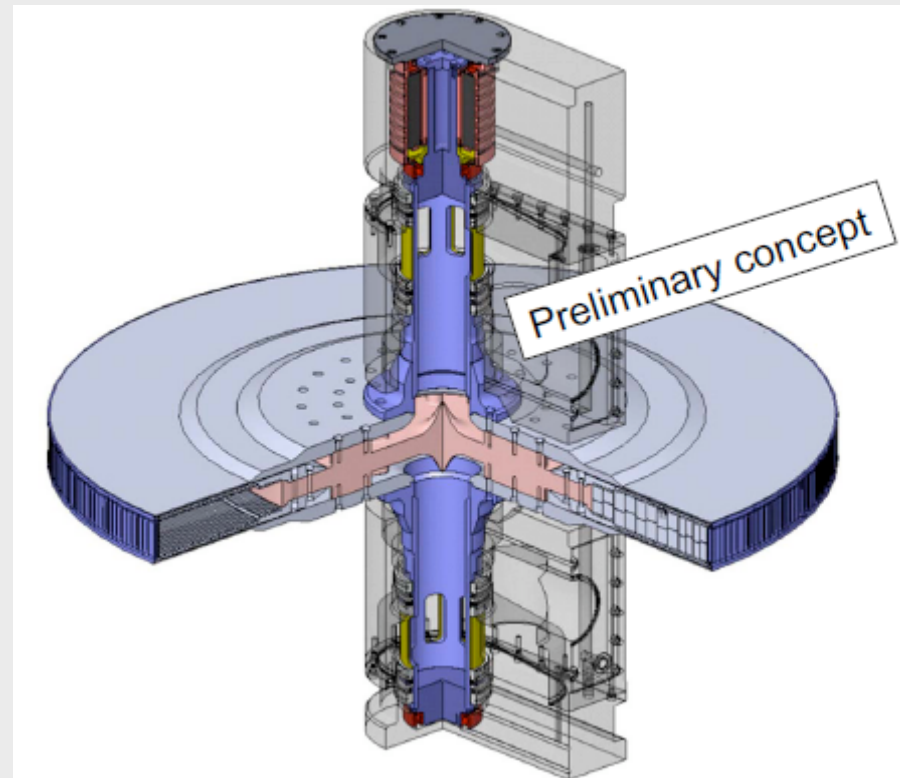


Solid target station



Rob Edgecock

- **Current option: a wheel – being investigated now**
- **Several already used, but most relevant: design study**



Horizontal for compatibility with baseline target station

J. Pozimski, Imperial College @ 2nd annual Eurov meeting 1-4 June 2010, Strasbourg

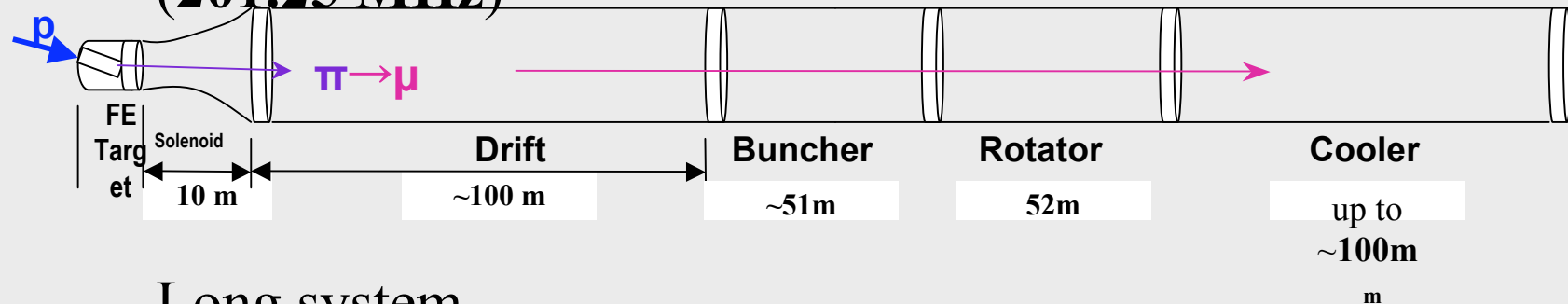


Compact Muon front end



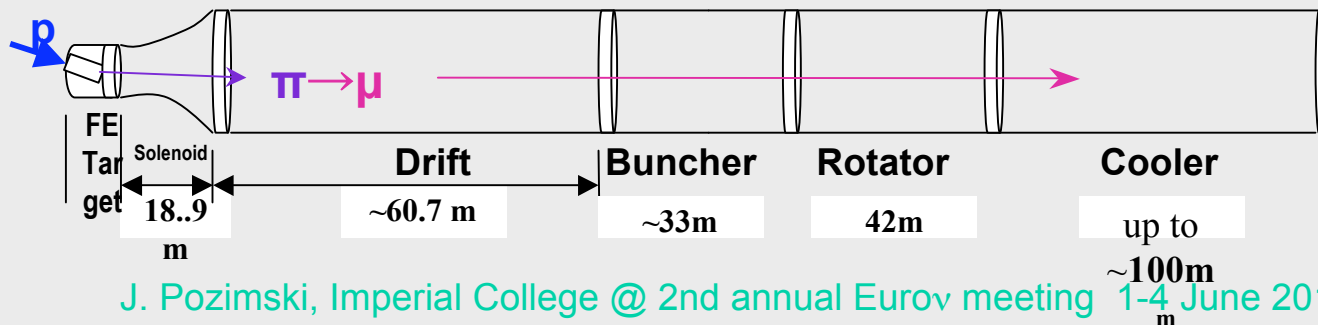
D. Neuffer

- ISS study based on $n_B = 18$ (280 MeV/c to 154 MeV/c)
 - Buncher 0 to 12MV/m; Rotator 12.5MV/m, $B=1.75T$



– Long system,

- Try shorter version - $n_B = 10$ (233 MeV/c to 154 MeV/c)
 - slightly lower fields (1.5T, 15MV/m)
 - Buncher 0 to 9 MV/m, Rotator 12MV/m
 - Shorter bunch train



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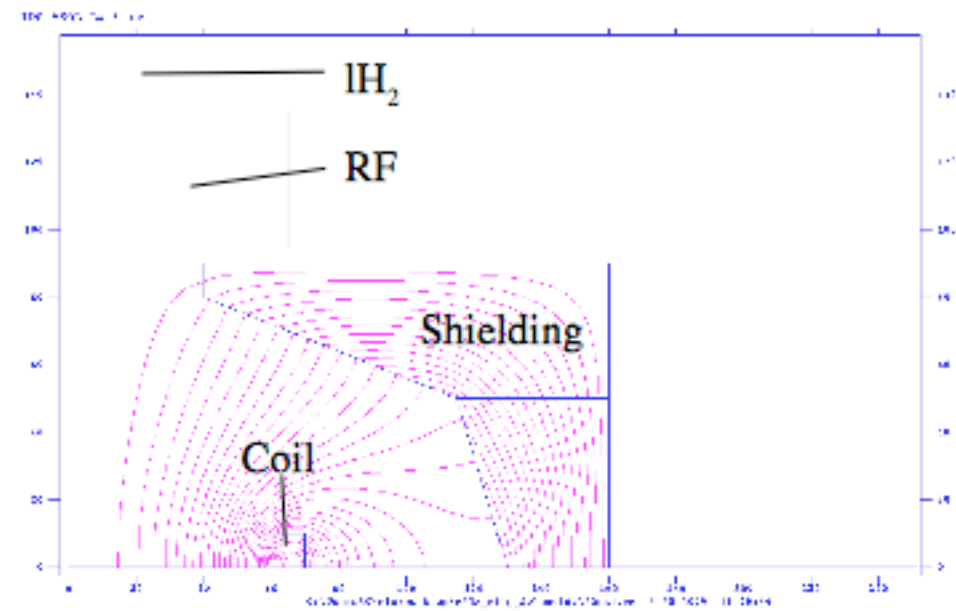
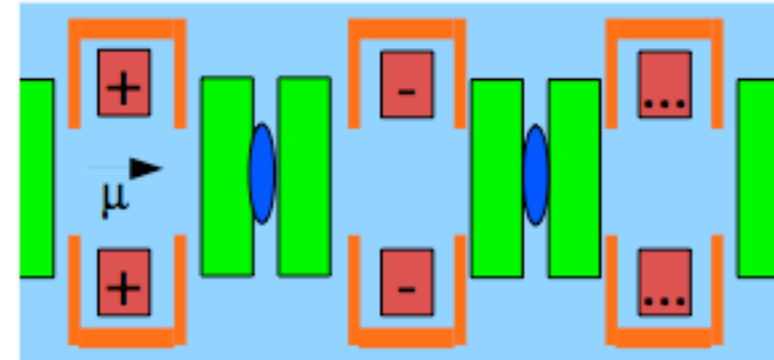


Shielded RF cooling lattice

C. Rogers



- Increase cell length to remove RF from fringe fields
 - Add shielding using iron or bucking coils
- Look at cooling section
 - This is where the RF is most limited
 - This is where optics are most demanding
- How well can we cool in this shielded scenario?
- How well can we optimise the cooling lattice?
- Try to keep RF cavities in < 0.1 T fields
- Liquid Hydrogen absorbers



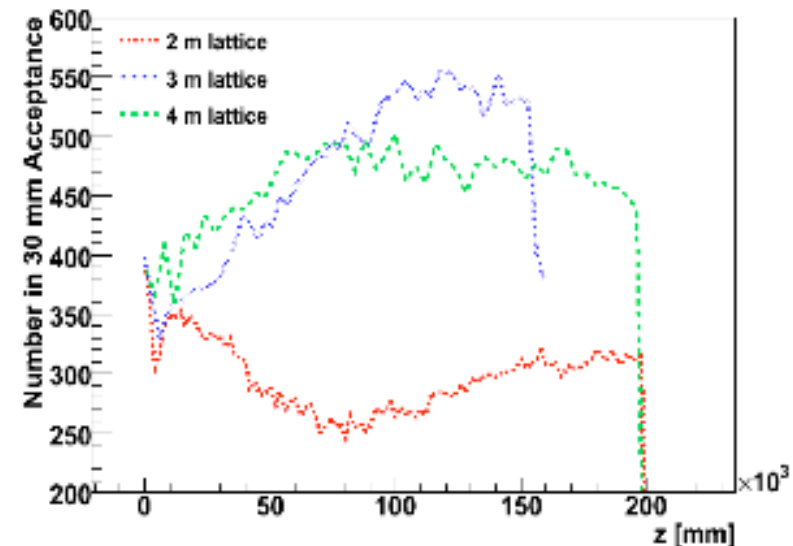
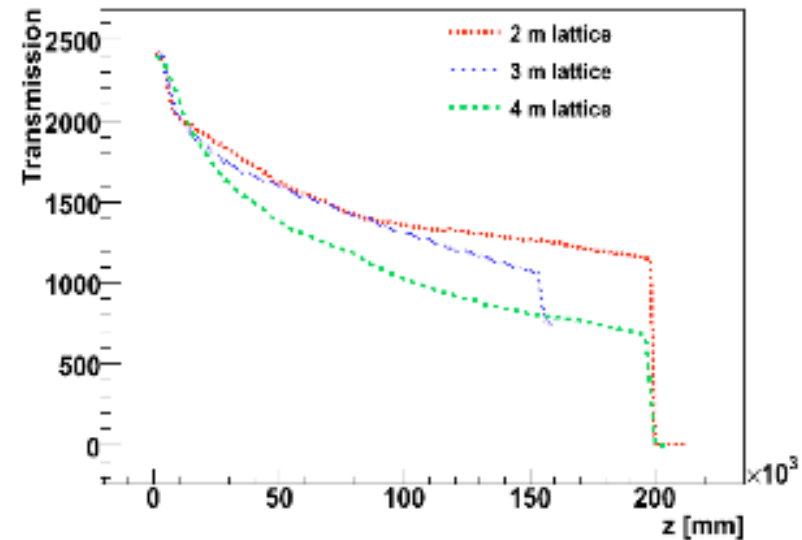


Shielded RF cooling lattice

C. Rogers



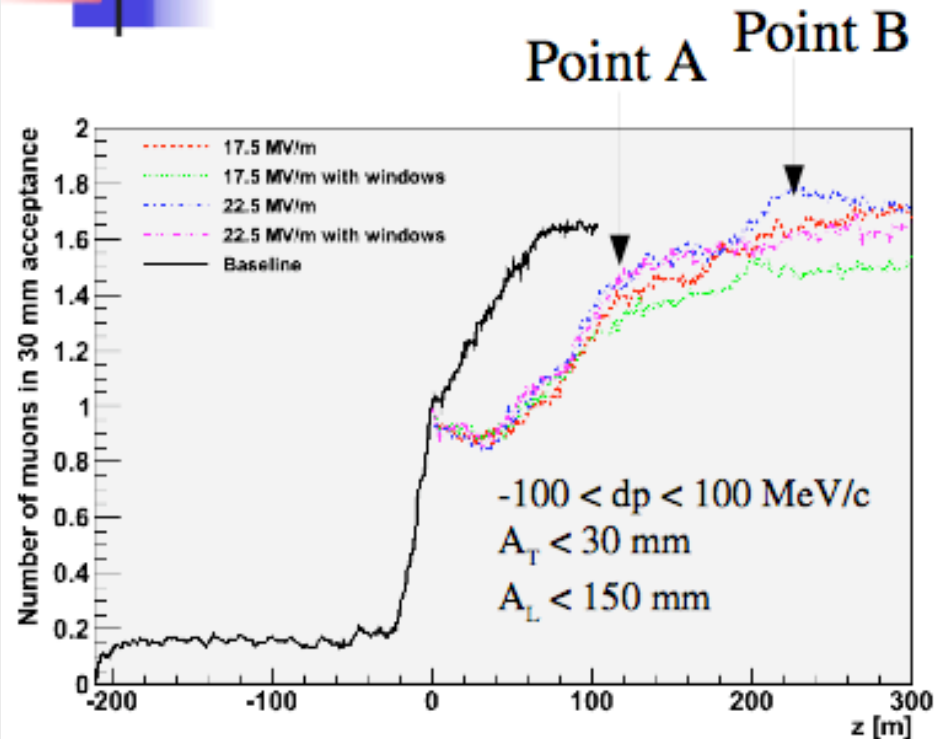
- Cell length optimisation
 - Simulated using long coil option
 - Race between RF packing fraction and β function
 - Higher RF packing \Rightarrow quicker cooling
 - Shorter lattice \Rightarrow lower β function (better equilibrium emittance)
- 3m lattice is optimal
 - Worry about initial beam loss
 - Nb low statistics
 - Get $\sim 40\%$ with long coil (a bit more optimisation is possible)
- Case for beta tapering?





Higher momentum beam

C. Rogers



- Fairly large transmission losses
 - $> \sim 50\%$
- Most of the remaining beam is inside the 30 mm acceptance
- Getting increase in rate of $\sim 70\%$
 - But with more hardware
 - Performance quite similar to baseline

- If I stop at point A - I use roughly the same amount of hardware as the baseline (RF packing fraction $\sim 1/2$ that of the baseline)
 - And lose a few muons
- I can recover baseline performance if I go to Point B
 - But those last few muons are expensive!



Front end work at CERN

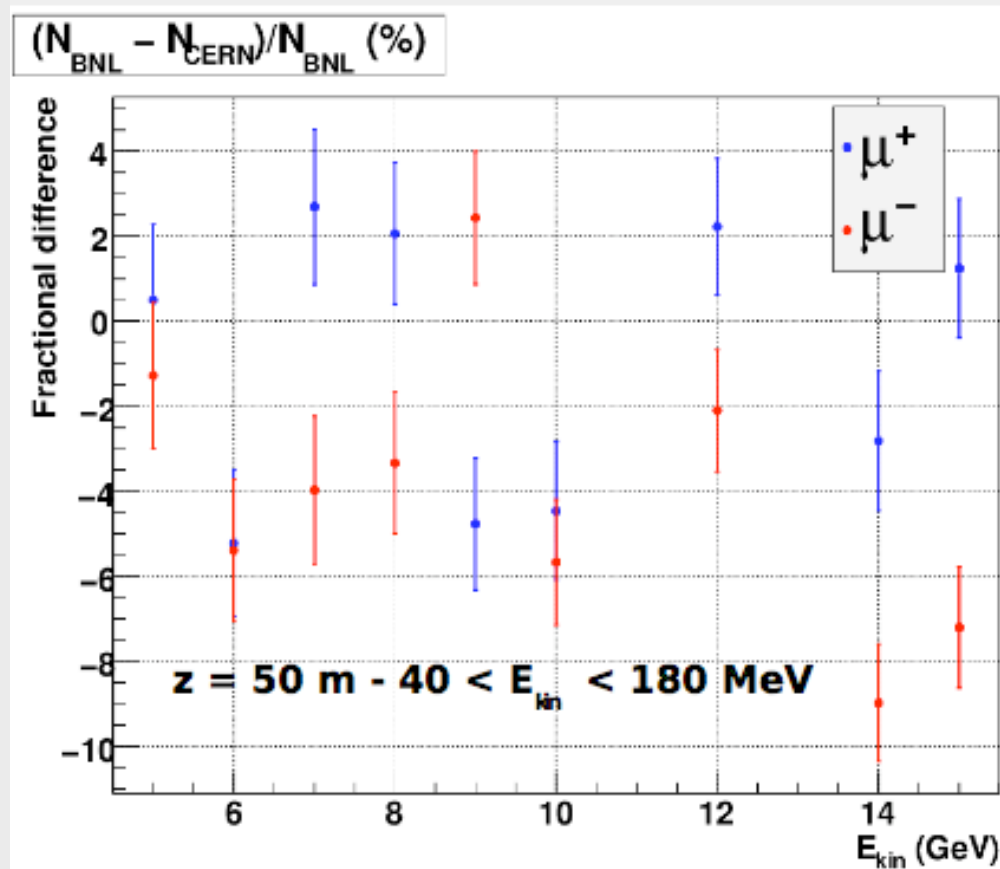
G. Prior



Particles production (coll. with BNL)

Tested the exact same input files on 2 different machines:

- SL4 2.6.9-89.0.11.EL.smp - 32x - little endian (BNL)
- SLC4 2.6.9-89.0.18.EL.cern - 32x - little endian (CERN)



$$\text{Error} = \sqrt{(N \cdot s_2 - s_1^2) / (N - 1)}$$

$$s_2 = \sum w_i^2 \quad s_1 = \sum w_i$$

Difference in the yield HAS to come from different code versions: need to be confirmed by the MARS developers as we don't have access to the source files.



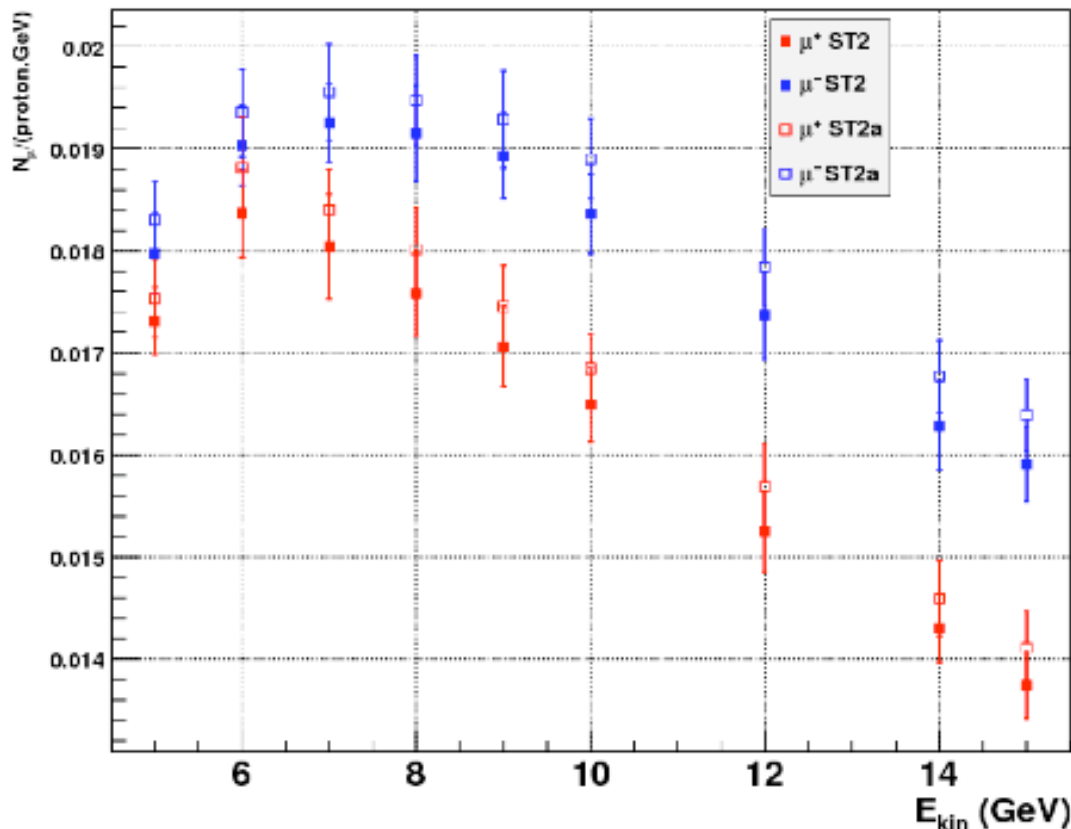
Front end work at CERN

G. Prior



ST2/ST2a comparison

Muon yield – $40 < E_{kin} < 180 \text{ MeV/c} - z = 50 \text{ m}$



Yield = MEAN

Error = STD

ST2a slightly better than ST2 (within the errors bars).

Advantage of a higher field at end of taper but ST2 magnet configuration in a more mature design.

Optimum at 7-8 GeV and yield difference between muon signs is likely to be model dependent.

June 3rd 2010

16

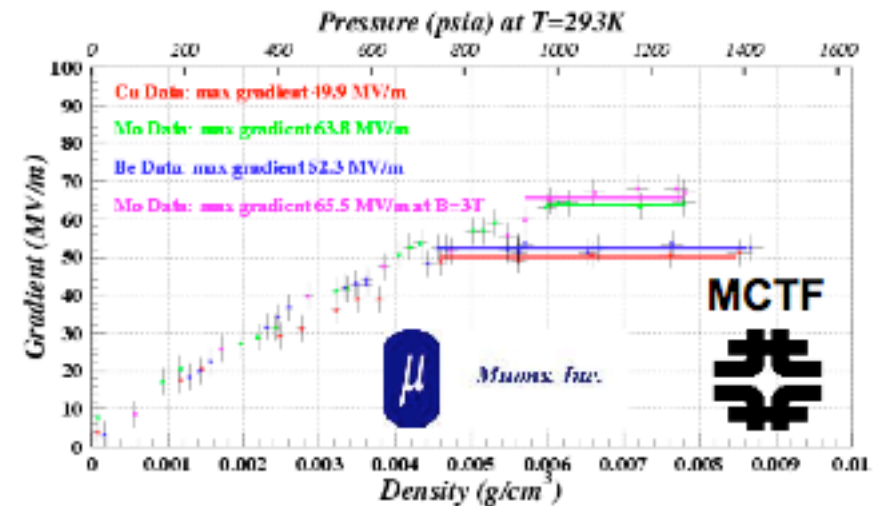
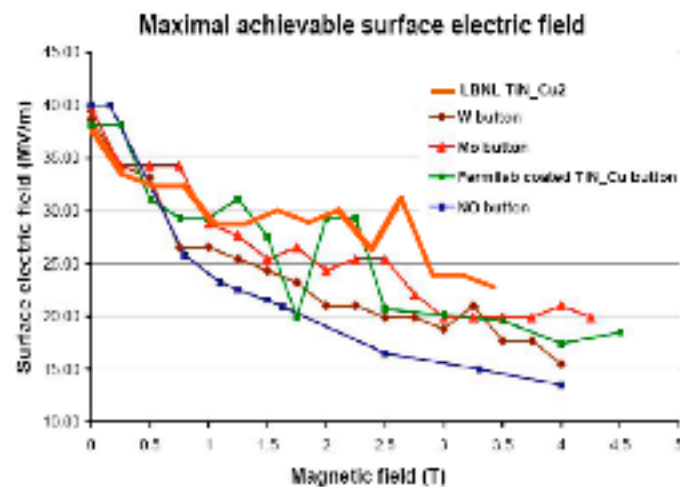


High pressure RF



M. Zisman

- We have evidence that vacuum RF cavity gradient performance degrades in a strong magnetic field
 - alternative approach of HPRF does not
 - though it has other potential issues
- It seems prudent to begin investigating the technical aspects of implementing HPRF in a linear cooling channel
 - minimizes changes in cooling channel layout and hardware

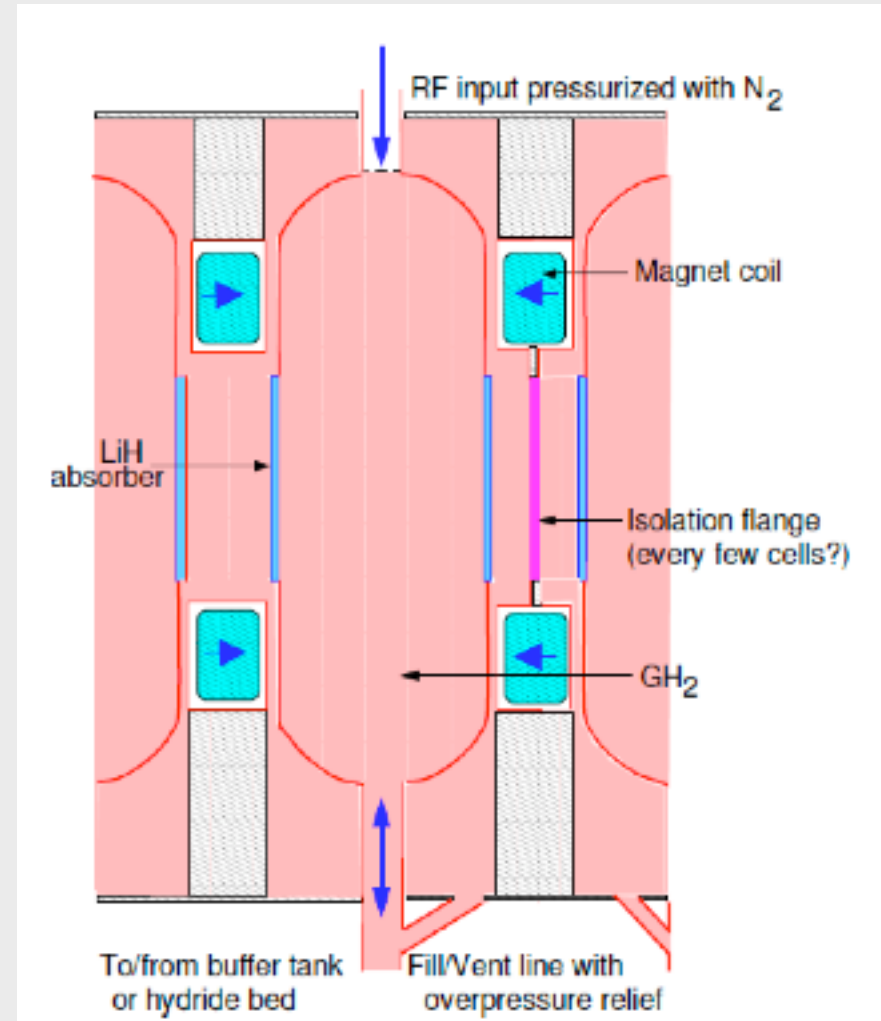
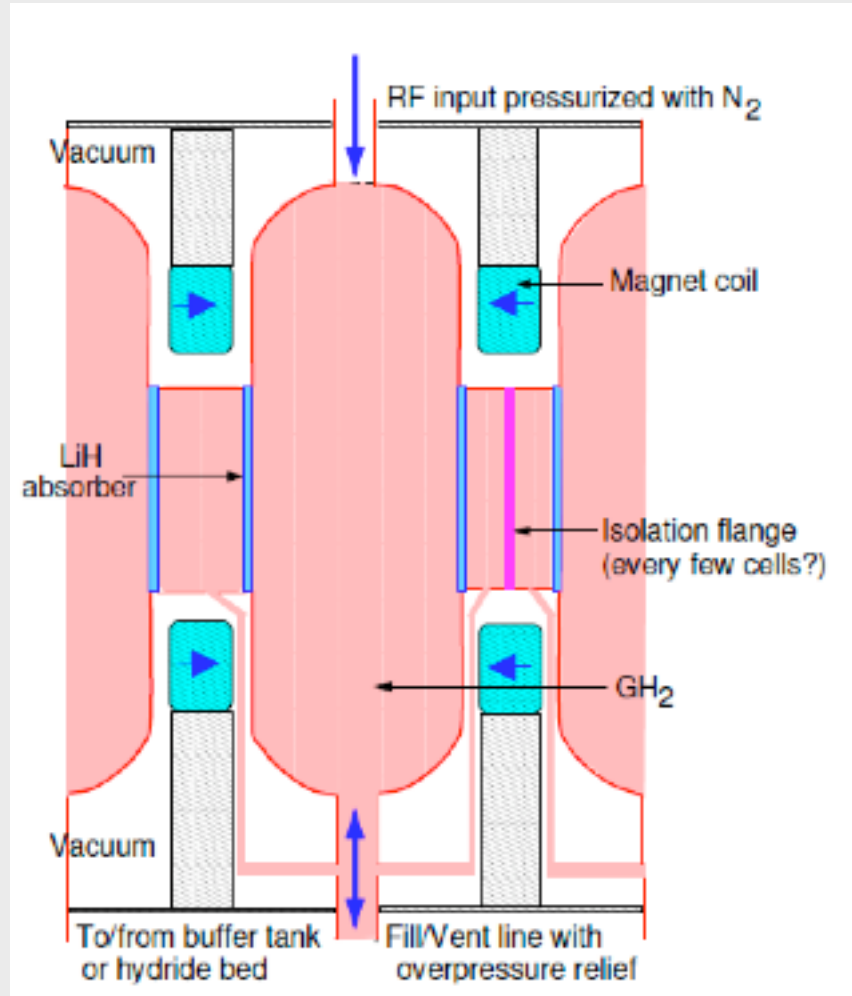




High pressure RF implementation



M. Zisman



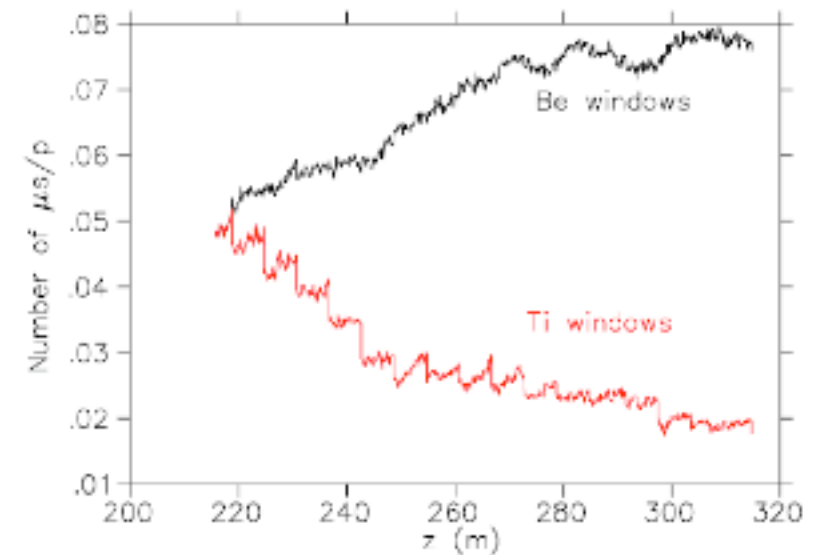
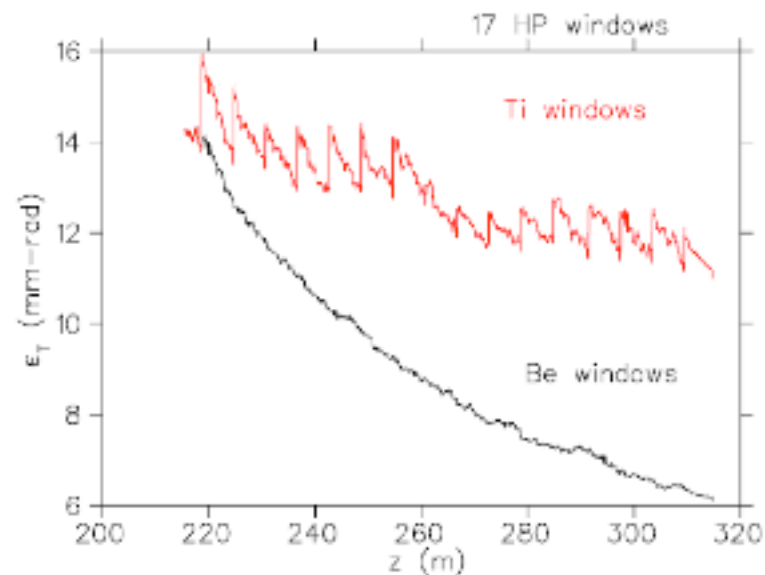


High pressure RF - window material



M. Zisman

- To make sure we were not fooling ourselves, ran cases with both Be and Ti
 - the difference is obvious
- Will next look at Al and AlBeMet windows as time permits
 - Al is okay in terms of hydrogen embrittlement; not yet sure about other materials

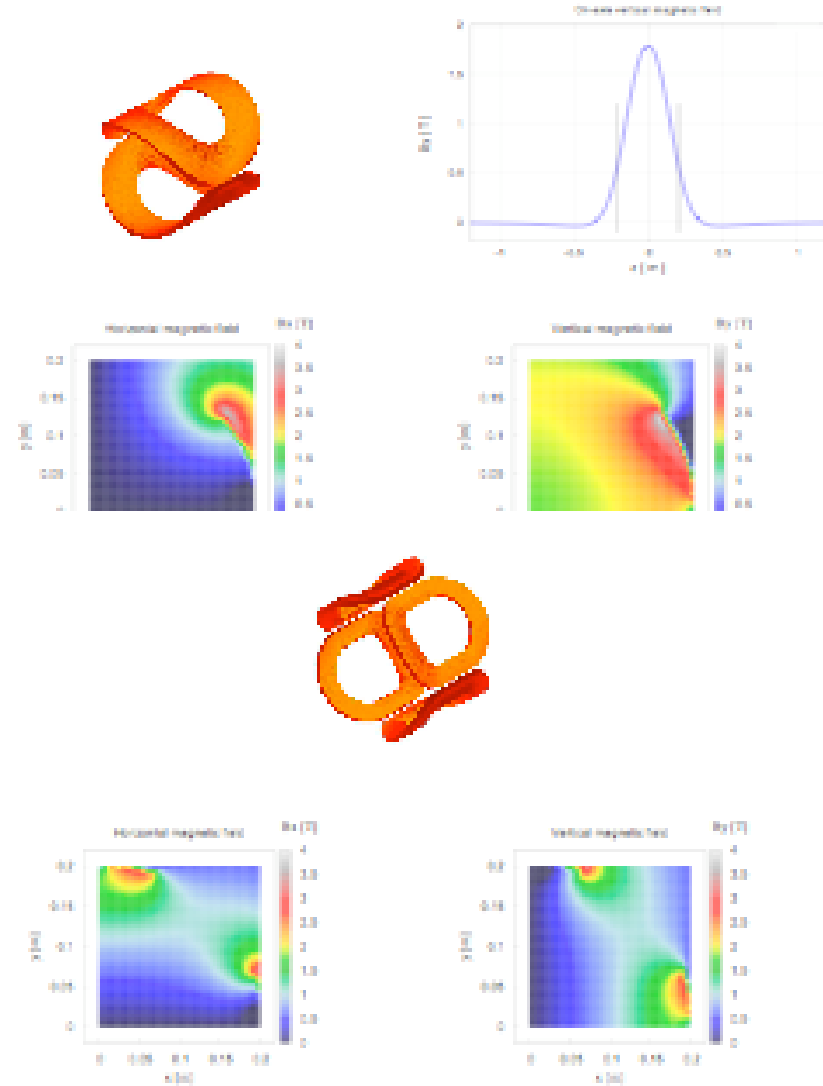
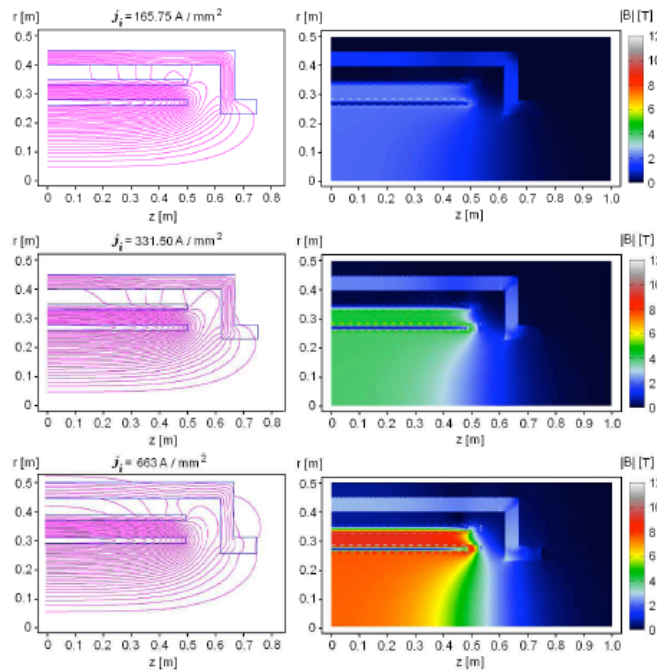
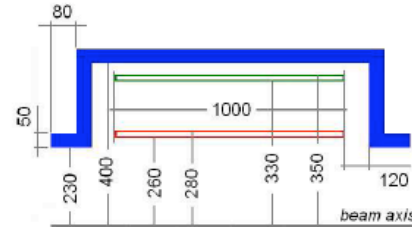
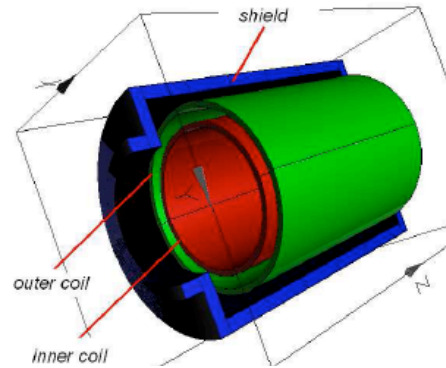




Hardware design magnets



C. Bontoiu



J. Pozimski, Imperial College @ 2nd annual EuroV meeting 1-4 June 2010, Strasbourg

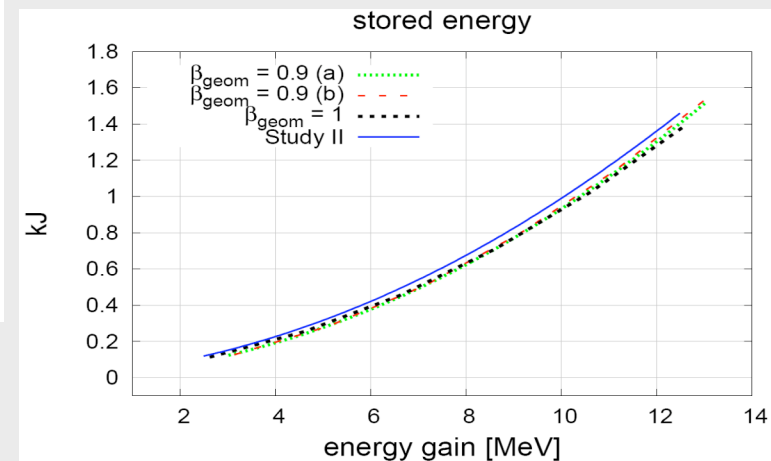
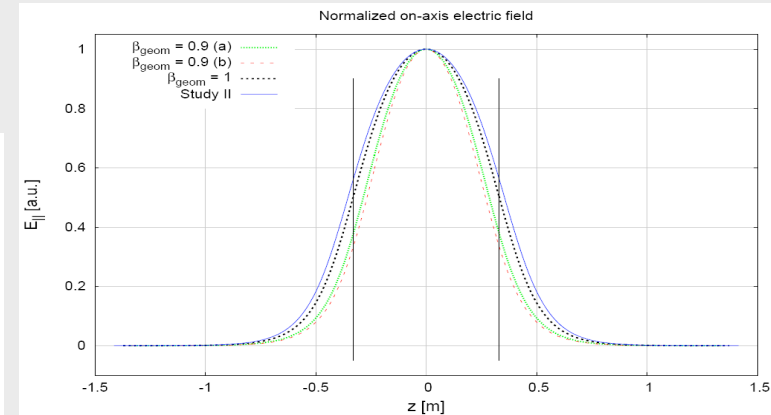


Hardware design RF cavities



C. Bontoiu / J. Pozimski

Parameter	$\beta_{geom} = 1$	$\beta_{geom} = 0.9$ (a)	$\beta_{geom} = 0.9$ (b)	Study II
l_{cav} [m]	0.7448	0.67034	0.67034	0.8282
r [m]	0.6854	0.7042	0.6804	0.6641
f_0 [MHz]	201.247	201.251	201.255	198.575
Q [10^9]	24.67	19.6	18.8	26.7
T	0.650	0.716	0.726	0.591
\hat{E} [MV/m]	26.17	27.19	27.83	26.38
\bar{E} [MV/m]	20.62	20.81	20.53	20.42
$ E _{surface}^{max}$ [MV/m]	21.70	24.87	29.45	19.75
$ H _{surface}^{max}$ [kA/m]	48.06	58.53	61.92	45.00
U [J]	712	772	797	747
$\int_{-\infty}^{+\infty} E(0,z)\cos[\omega t(z)]dz$	8.6142	9.0081	9.1336	8.8466
$\int_{-l_{cav}/2}^{+l_{cav}/2} E(0,z)\cos[\omega t(z)]dz$	10.0000	10.0000	9.9999	10.0000
$\int_{+l_{cav}/2}^{+\infty} E(0,z)\cos[\omega t(z)]dz$	-0.69204	-0.49594	-0.43320	-0.75676
correction [%]	-13.841	-9.9188	-8.6639	-15.135





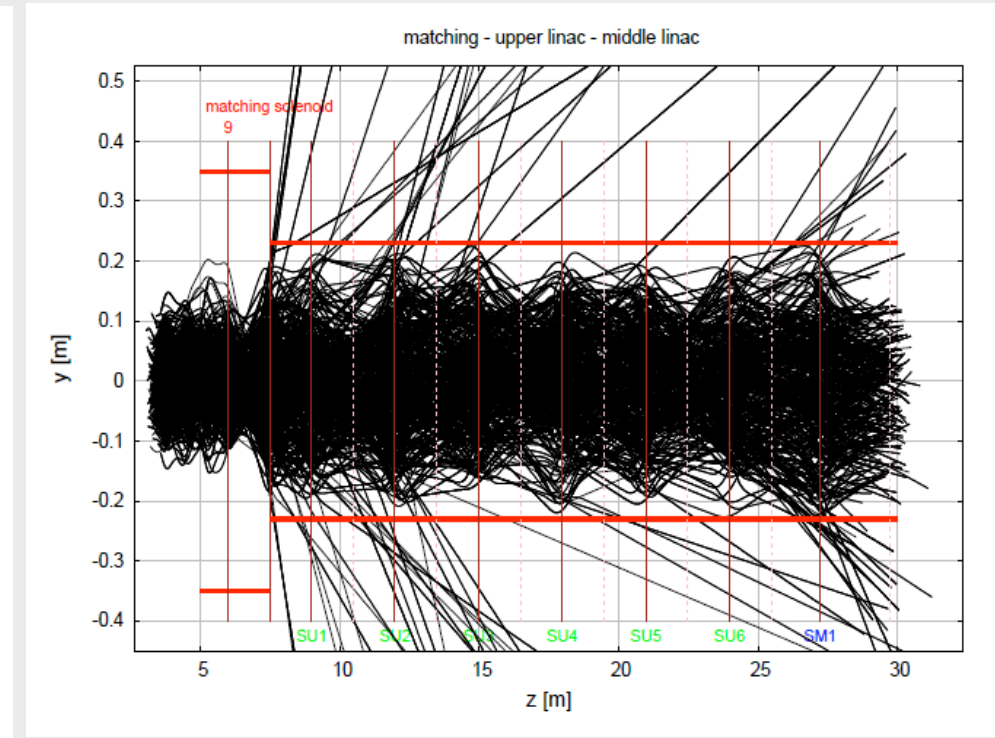
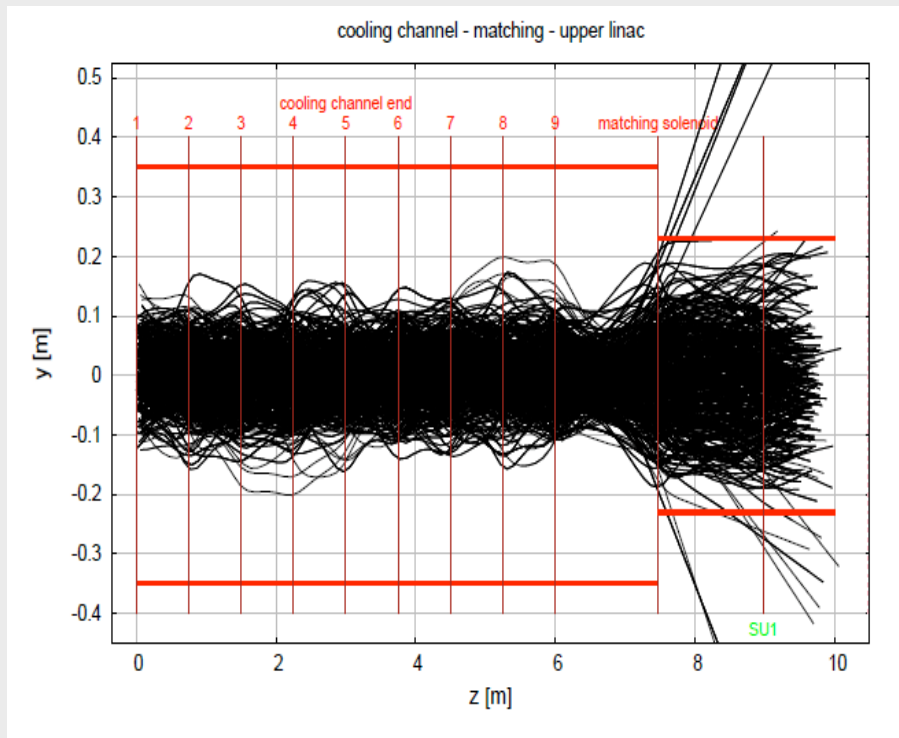
Matching from cooling to linac



C. Bontoiu

cooling → upper
linac

upper → middle linac



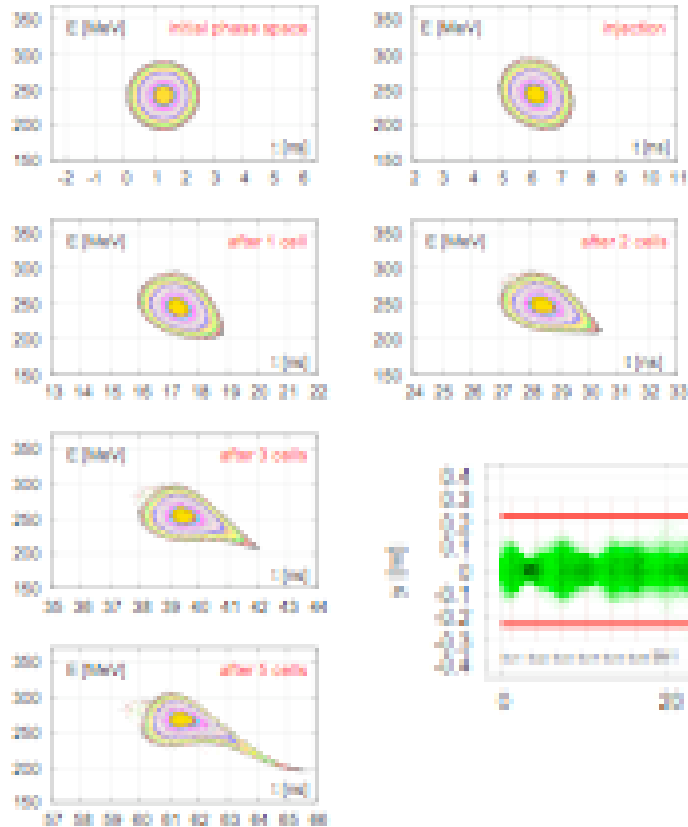


Particle tracking - Linac

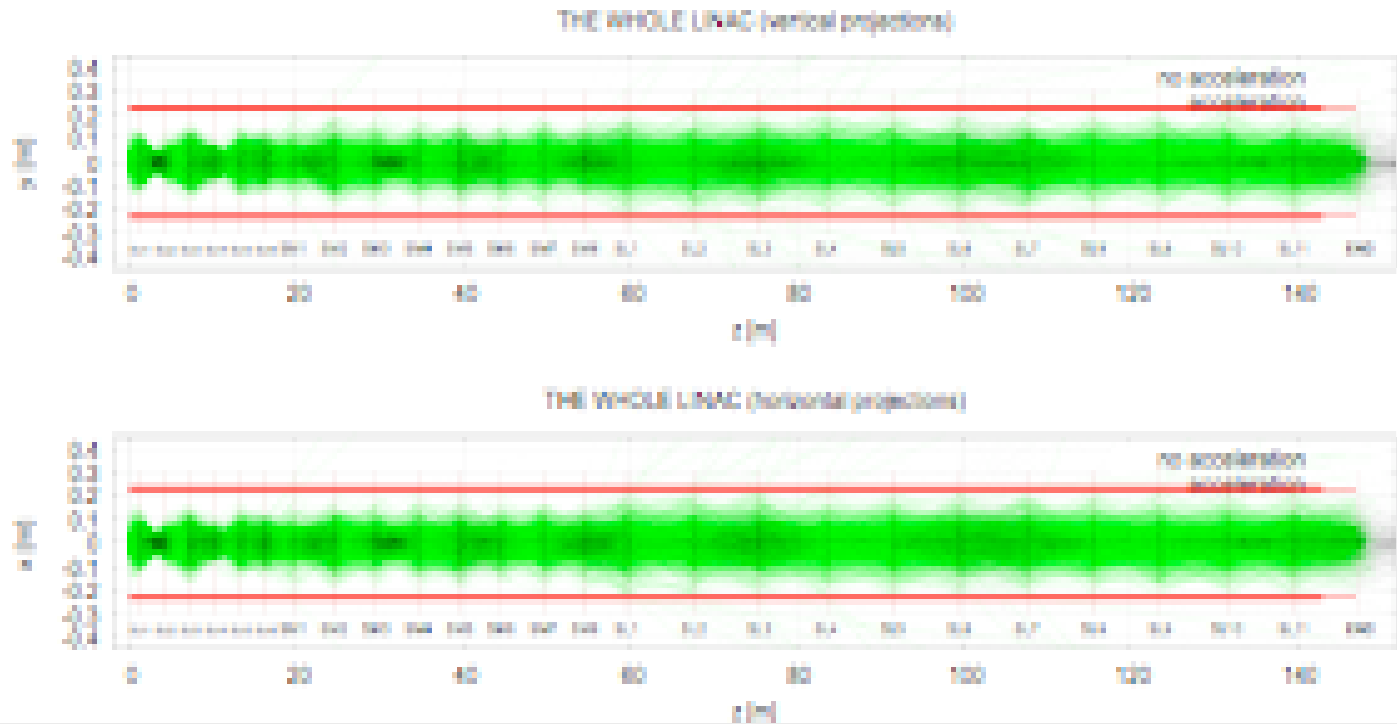


C. Bontoiu

Longitudinal phase space



Transversal projections



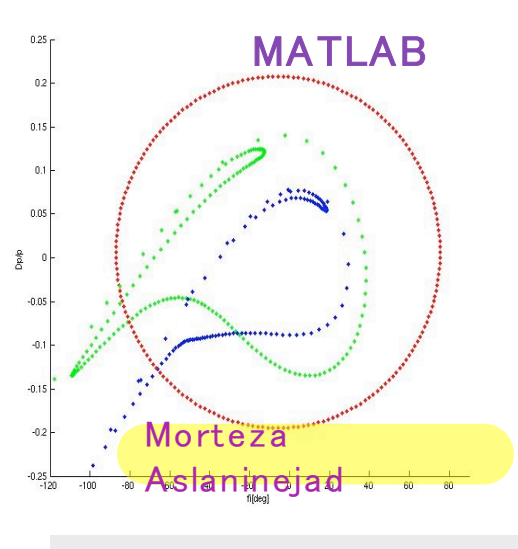
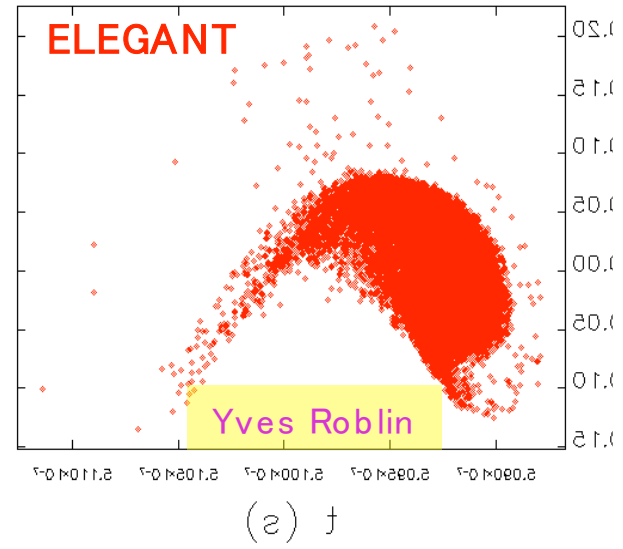
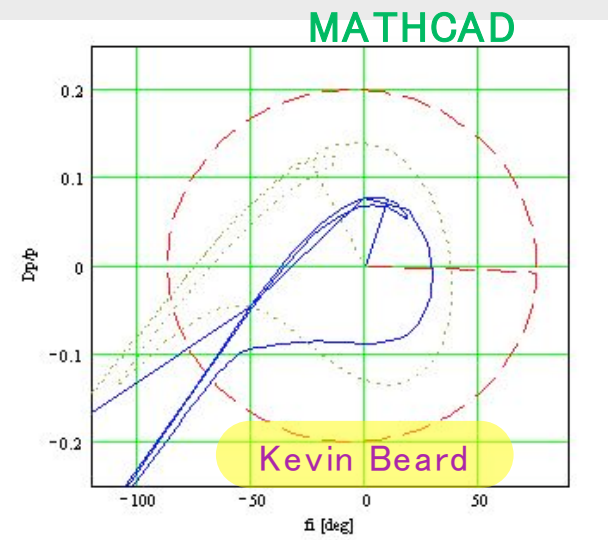
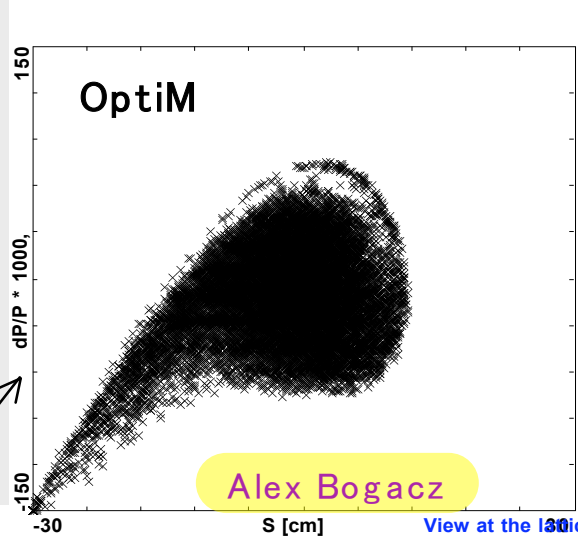
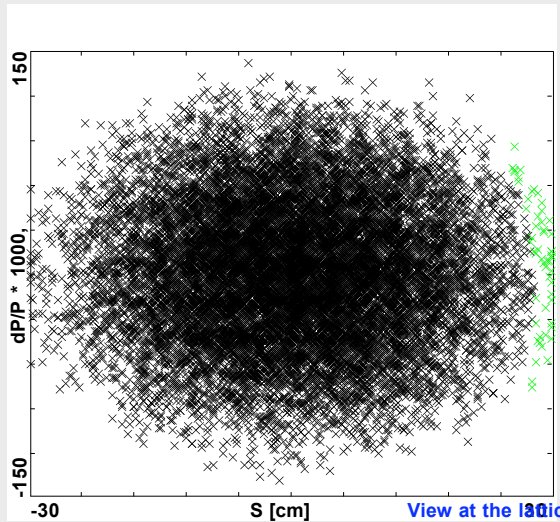


Longitudinal phase space

A. Bogacz



Initial distribution



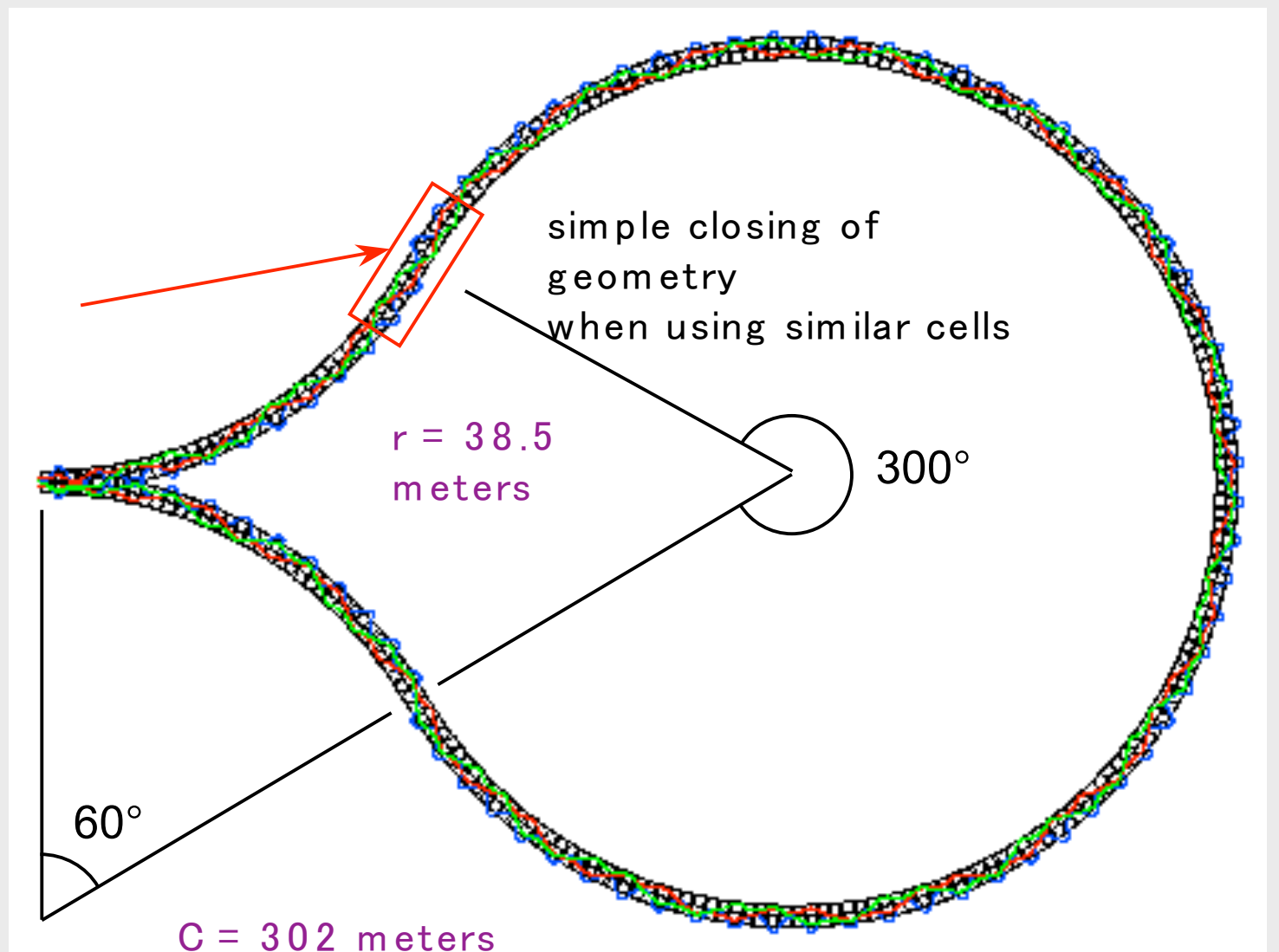


Multipass FFAG arc

A. Bogacz



Vasiliy Morozov





Scaling FFAG alternative for RLA 2



T. Planche

Parameters of a 3.6 to 12.6 GeV muon ring

Lattice type	FDF triplet
Injection/extraction energy	3.6/12.6 GeV
RF frequency	200 MHz
Number of turns	6
RF peak voltage (per turn)	1.8 GV
Synchronous energy	8.04 GeV
Mean radius	~ 160.9 m
B_{max} (@ 12.6 GeV)	3.9 T
Field index k	1390
Total orbit excursion	14.3 cm
Harmonic number h	675
Number of cells	225
Long drift length	~ 1.5 m
Horiz. phase adv. per cell	85.86 deg.
Vert. phase adv. per cell	33.81 deg.

Table 1 - Example of 3.6 to 12.6 GeV muon scaling FFAG ring parameters.

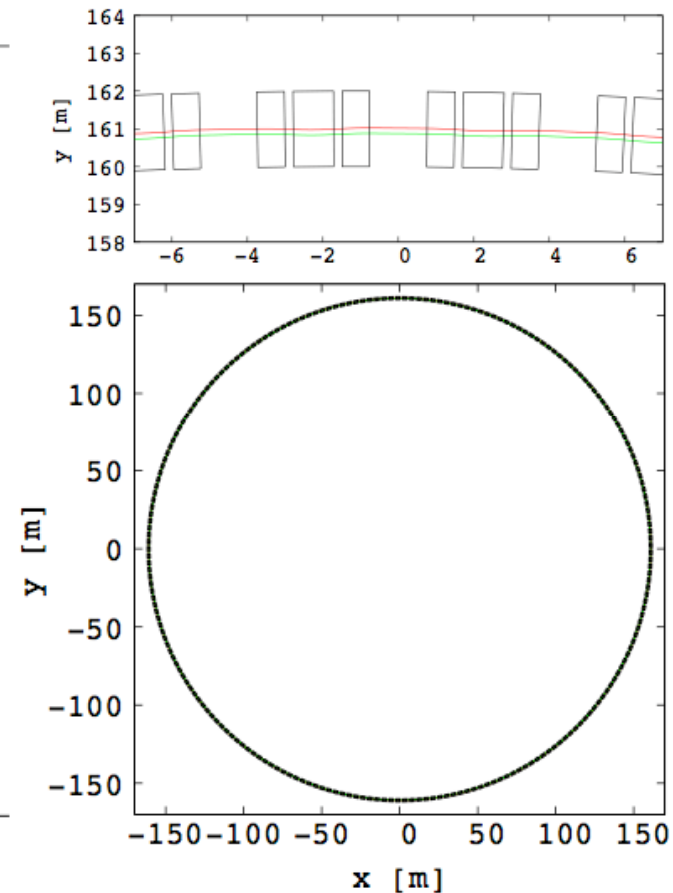


Figure 2 - Ring layout.



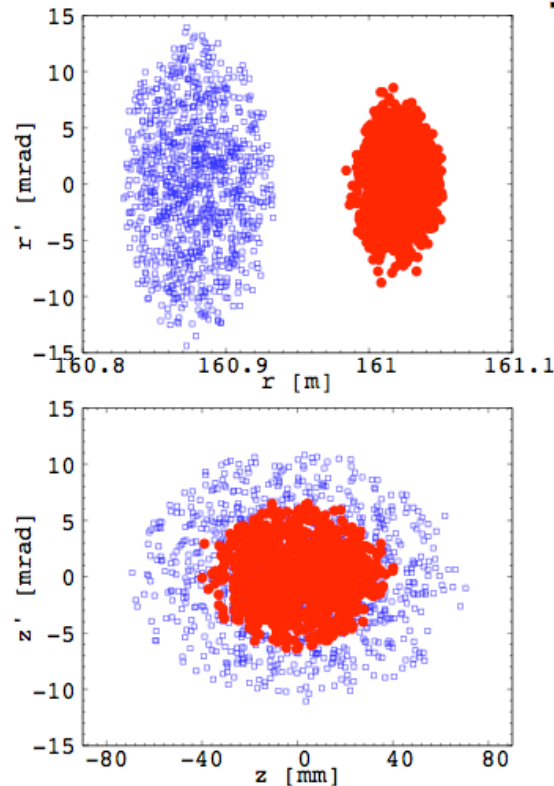
Scaling FFAG alternative for RLA 2



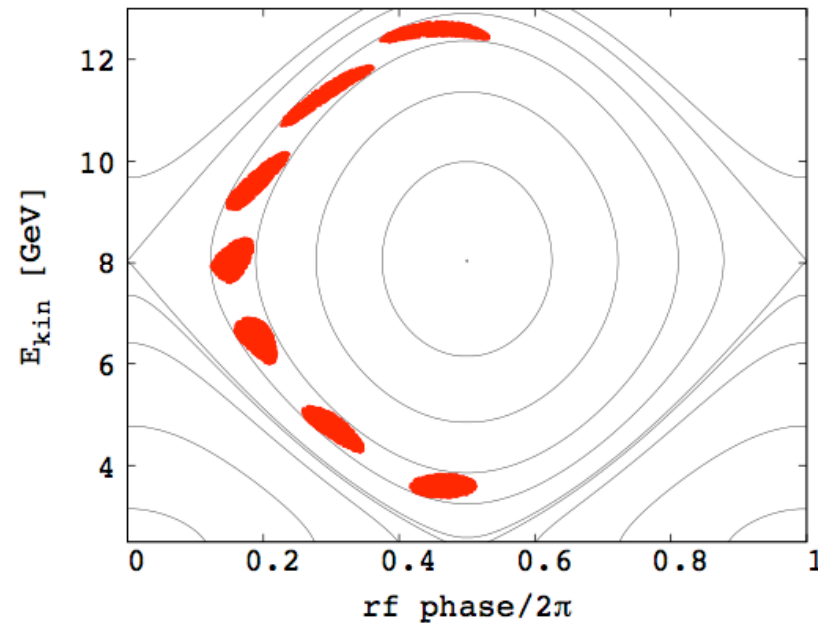
T. Planche

Full acceleration cycle - 6D tracking

- Tracking results -



Figures 8 - Initial (blue) and final (red) particles distribution in the horizontal (top), and vertical (bottom) phase space.



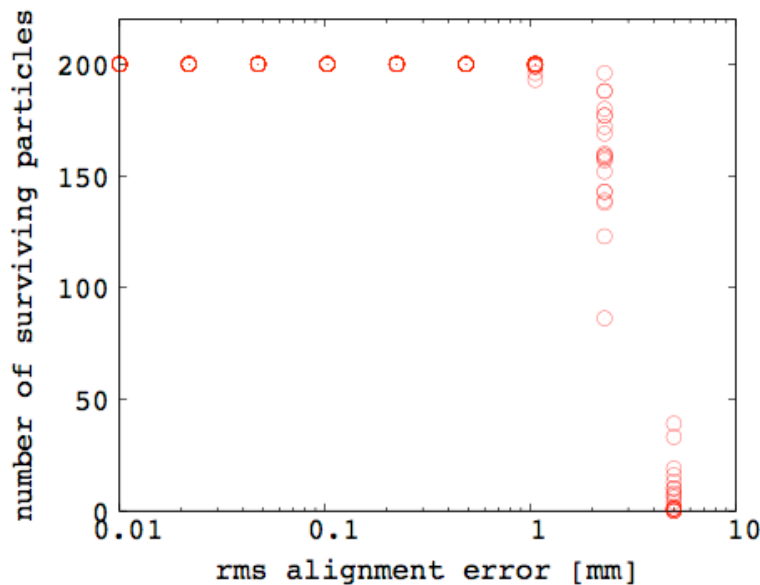
Figures 9 - longitudinal phase space plot showing a 6-turn acceleration cycle. Hamiltonian contours are superimposed.



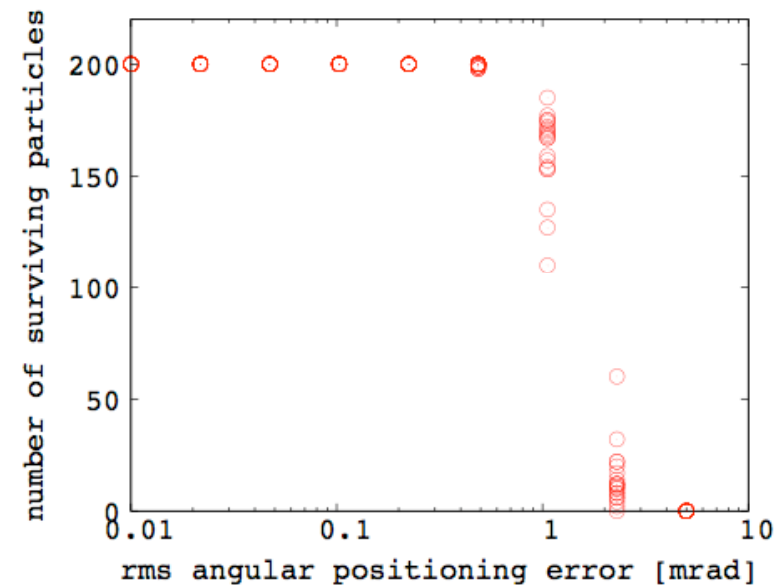
T. Planche

Study with errors

Errors in translation



Errors in rotation



Figures 10 - Number of surviving particles depending on the rms error. 20 different lattices have been generated and tested for each value of rms error.

More tolerant to errors than the NS-FFAG!



FFAG Lattice & EMMA



S. Berg / R. Edgecock / F. Meot

Most promising due to longer drifts for RF and kickers : FDFCC

Minimum E 12.6 GeV

Maximum E 25 GeV

Cells 64

Circumference 546 m

Cavity cells 88

RF Voltage 1119 MV

Decay 5.6%

Turns 11.8

D Radius 11.5 cm

F Radius 15.3 cm

D Field 6.5 T

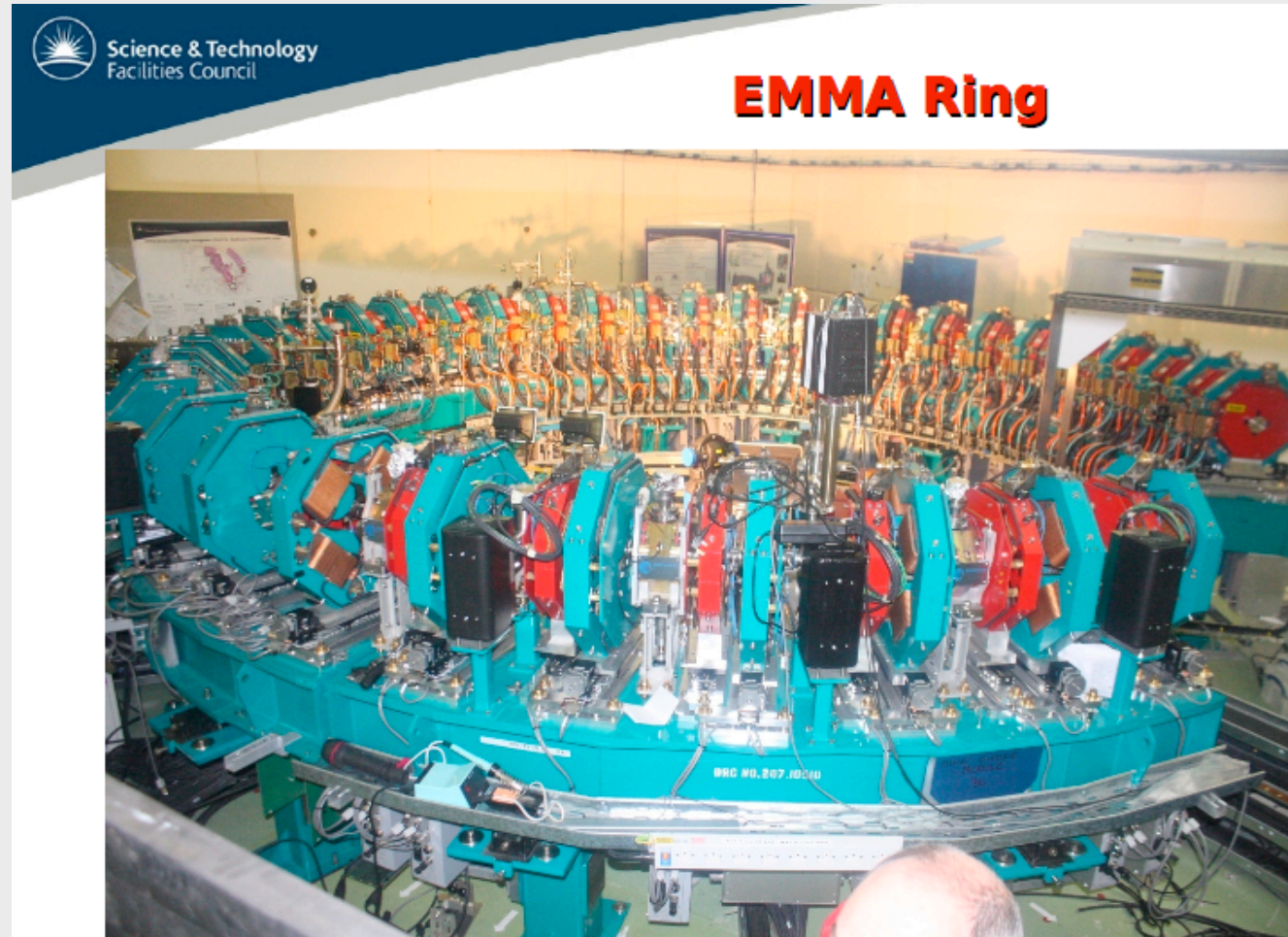
F Field 3.6 T

D Length 1.96 m

F Length 1.29 m

Long drift 3 m

Short drift 0.5 m



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NS-FFAG chromaticity correction



S. Machida

- Reduces time of flight dependence on transverse amplitude
- Downsides
 - Reduction in dynamic aperture
 - Increase in magnet apertures (cost)
 - Modification of time of flight vs. energy
- Still studying optimal choice for this
 - Some modest correction likely included



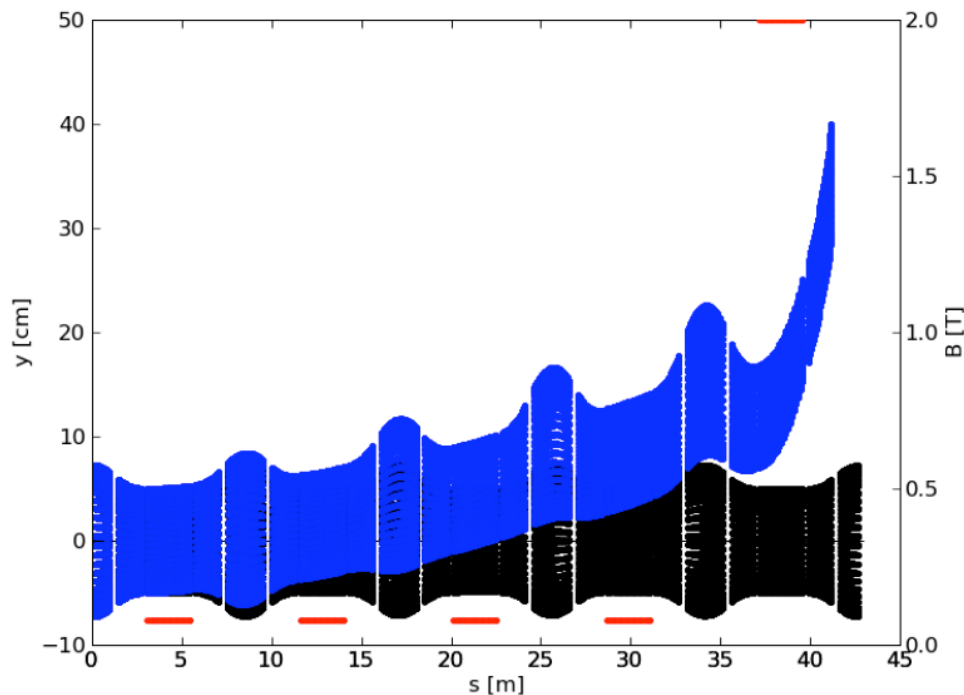
Extraction particle dynamics



J. Pasternak / D. Kelliher

FDFCC - Vertical extraction

- 4 kickers at 0.078 T in consecutive drifts
- Septum at 2T
- Several magnets require large aperture
- Extracted beam 17cm from magnet axis of F after septum



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Injection & Extraction studies

J. Pasternak / D. Kelliher



Summary of parameters for injection/extraction (updated lattices)

Scheme	Triplet 1		FODO		Triplet 2	
	FDFC Injection	FDFC Extraction	FCDC Injection	FCDC Extraction	FDCC Injection	FDCC Extraction
Plane	Horizontal	Vertical	Vertical	Vertical	Horizontal	Vertical
No. Kickers	6	6	6	6	3	4
Kicker top field [T]	0.103	0.103	0.113	0.101	0.089	0.078
Septum top field [T]	>3	>4	>3	>4	2	4
Kicker/Septum length [m]	1.4	1.4	1.4	1.4	2.4	2.4
Cells needed	8	8	5	5	5	6

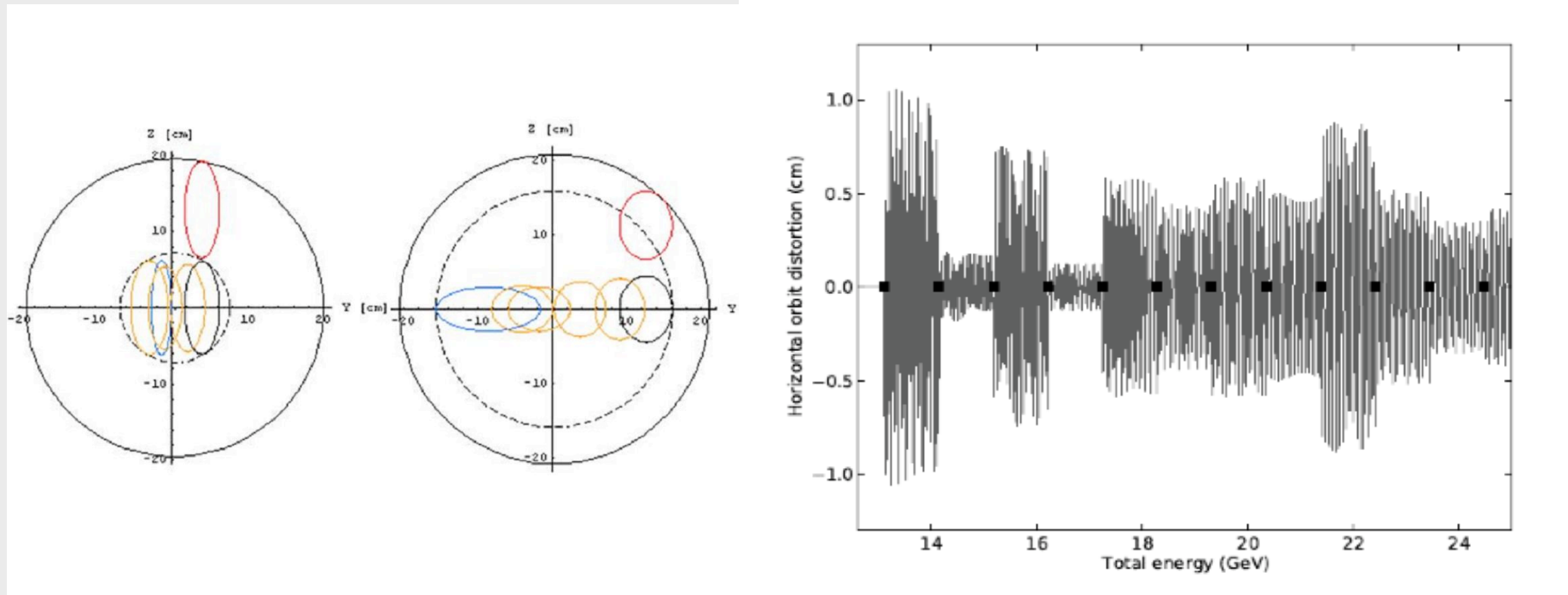


Extraction particle dynamics



J. Pasternak / / D. Kelliher

Effects of special magnets



- Beams close to septum push the magnet aperture.
- Special magnets with higher aperture is needed in injection/extraction regions.
- Those magnets introduce the ring lattice symmetry breaking, which can cause accelerated orbit distortion.
- Current studies show that the effect is not dramatic, but more simulations are required.

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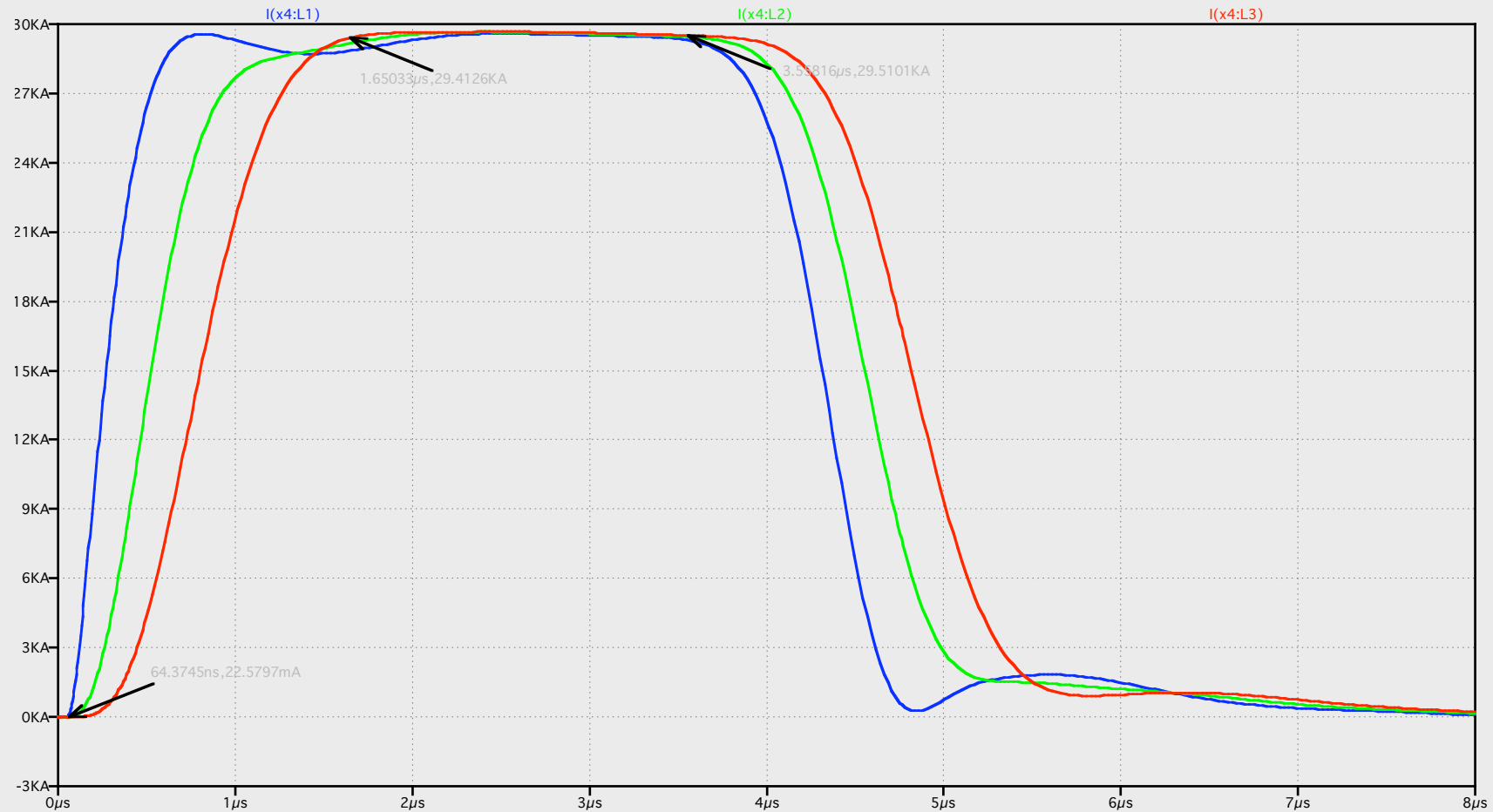


Kicker electronics simulations



J. Pasternak

Current pulses in 3 kicker sections – „travelling wave” using PSPice



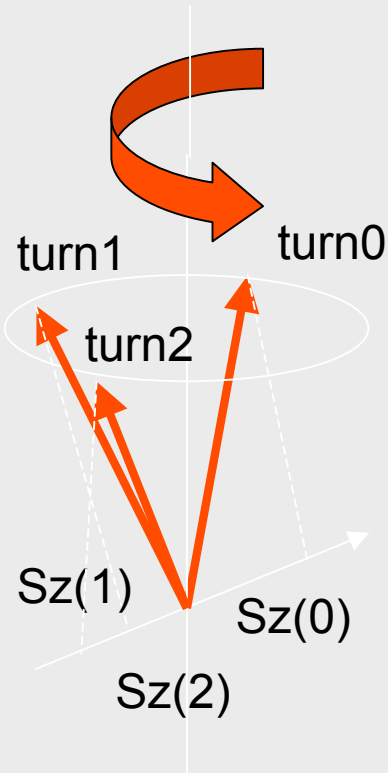
J. Pozimski, Imperial College @ 2nd annual Eurov meeting 1-4 June 2010, Strasbourg



Ring instrumentation - Energy



M. Apollonio



- Energy can be measured using the *Polarisation of the Muon Beam* [Raja-Tollestrup – FERMILAB-Pub-97 / 402] IF some \mathcal{P} is saved after all the message in the machines ...

- Spin precesses in a ring due to coupling with magnetic fields (bending magnets).

- At every turn spin precession is determined by the *SPIN TUNE*:

$$\omega = 2 \pi \gamma a$$

$$a = 1.16E-3$$

This determines a modulation in \mathcal{P}

- NB: if $\Delta E/E = 0 \rightarrow \gamma$ same for all muons $\rightarrow \mathcal{P}$ keeps oscillating
- if $\Delta E/E \neq 0 \rightarrow \mathcal{P}$ goes to 0 after n turns

e^+ spectrum from μ -decay is a function of \mathcal{P} :

$$d^2N/dx d\cos\theta = N_0[(3-2x)x^2 - \mathcal{P}(1-x)x^2 \cos\theta] \quad (\text{CM})$$

- I have modelled the behaviour of a beam ($> 1E5$ muons) all with their spin and energy ($\Delta E/E = [0.01-0.05]$)

- Lorentz Boost

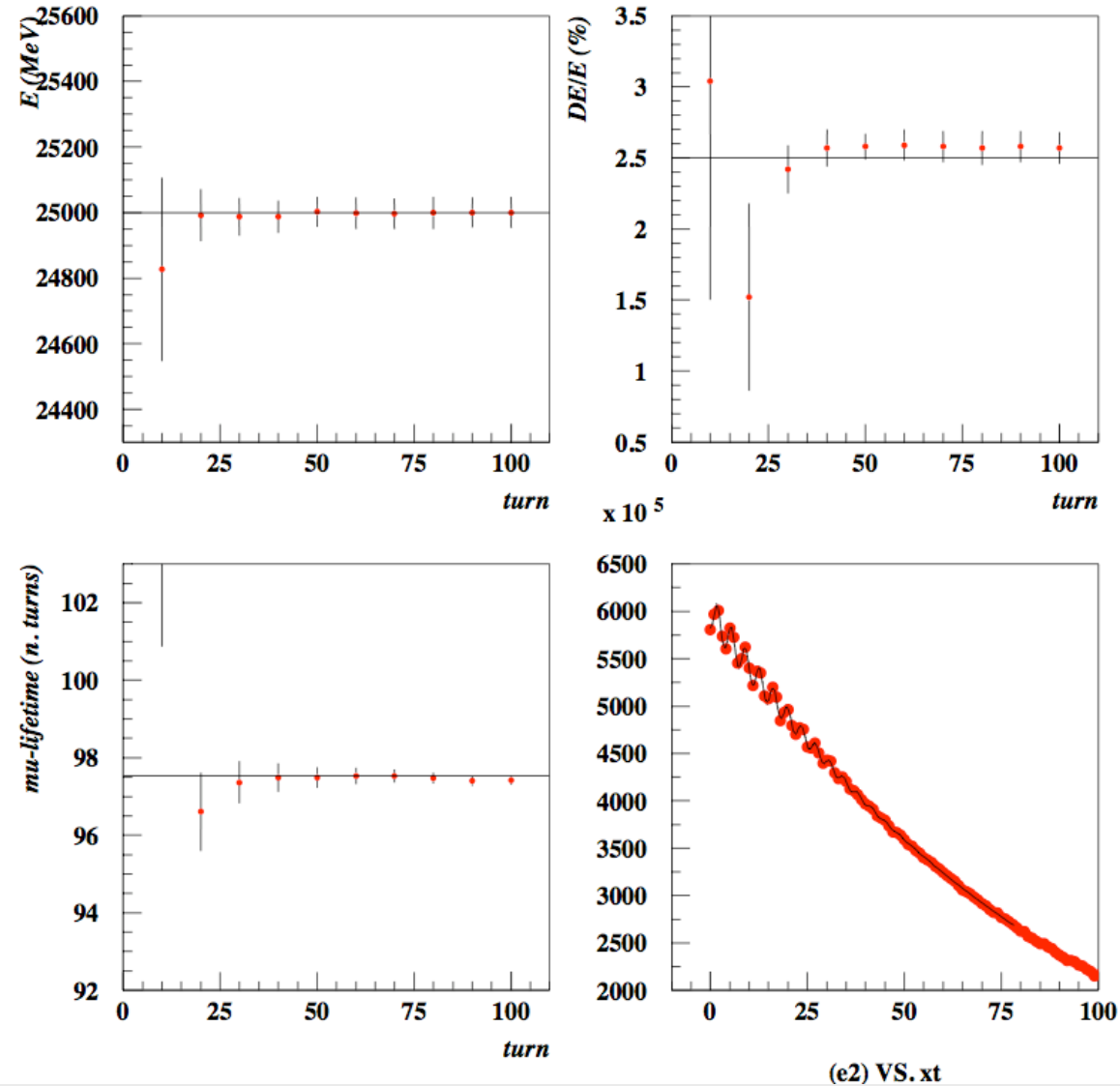
- Modulation in \mathcal{P} produces a modulation in $E(e^+)$



Simulation results



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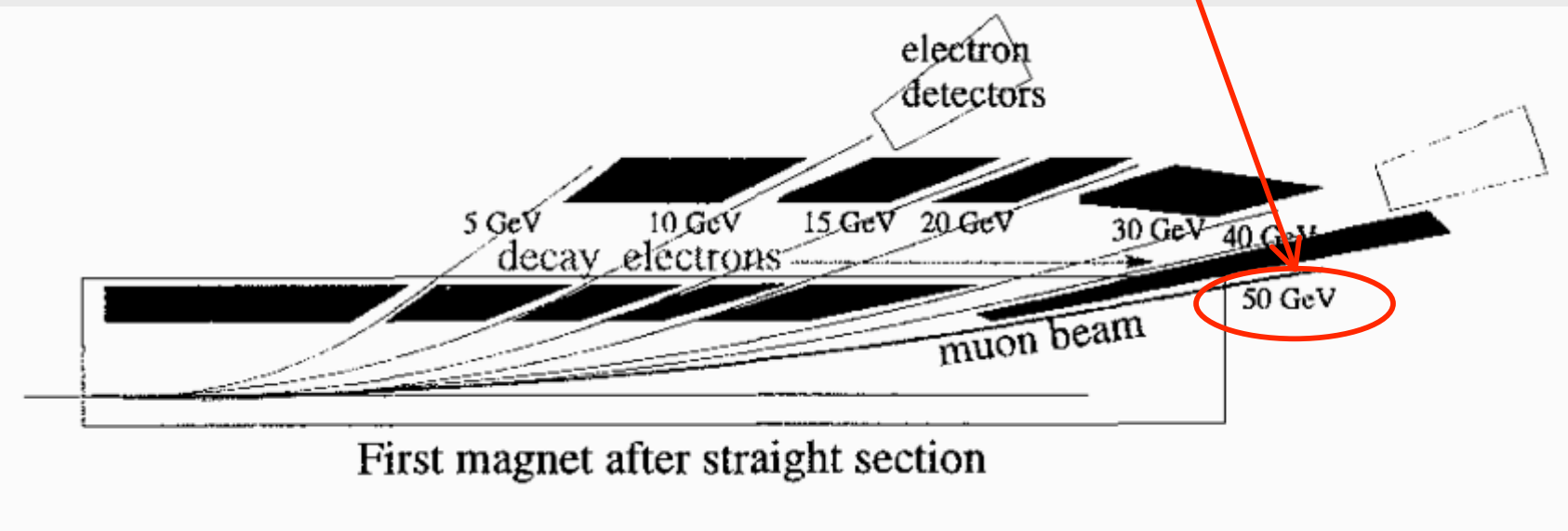
Implementation



M. Apollonio

It has been suggested [Blondel – ECFA 99-197(1999)] to use the *first bending magnet* after the decay straight section to SELECT electron energy bins: what does that mean today with a realistic lattice (25 GeV)?

In fact electron is emitted \sim parallel to μ (due to the high γ)



The spectral power of the 1st magnet depends on its *FIELD* and *LENGTH*

A *G4Beamline simulation* can tell us where electrons impinge after decaying somewhere along the orbit



IDR preparations



- Sub groups presented detailed plans for convergence.
- How do we organize the decision making process ?
- How do we handle alternatives and fall back options in the IDS ?
- IDR chapters should be prepared to be ready for submission to editors after RAL plenary (1st October)
- For IDS “mixed” costing according to readiness of the hardware design (global to detailed)