Early time gluon fields in relativistic heavy ion collisions

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June 23, 2022

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1. motivation

glasma provides initial conditions for subsequent hydro phase

- 2. structure of the calculation ColourGlassCondensate effective field theory approach
- 3. results:
 - 3.1 isotropization
 - 3.2 azimuthal asymmetries
 - 3.3 angular momentum
 - 3.4 momentum broadening of hard probes
- 4. conclusions

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motivation

goal: describe early time ($au \leq 1$ fm) dynamics of HIC

- evolution of system during this early stage not well understood
- importance: initial conditions for subsequent hydro evolution

more generally: want to understand transition between early-time dynamics \longrightarrow hydro phase

- 1. microscropic theory of non-abelian gauge fields
- = far from equilibrium
- 2. macroscopic effective theory
- based on universal conservation laws
- valid close to equilibrium

MEC, Czajka, Mrówczyński arXiv:2012.03042; 2105.05327; 2112.0681; 2202.00357

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method is based on a separation of scales between

- 1. valence partons with large nucleon momentum fraction (x)
- 2. gluon fields with small x and large occupation numbers
- gluons are in the saturation regime
- distributions are controlled by the saturation scale Q_s

dynamics of gluon fields determined from classical YM equation \rightarrow source provided by the valence partons

L. D. McLerran and R. Venugopalan, Phys. Rev. D 49, 2233 (1994); 49, 3352 (1994); 50, 2225 (1994).

A. Kovner, L. D. McLerran and H. Weigert, Phys. Rev. D 52, 6231 (1995).

J. Jalilian-Marian, A. Kovner, L. D. McLerran and H. Weigert, Phys. Rev. D 55, 5414 (1997).

F. Gelis, E. Iancu, J. Jalilian-Marian and R. Venugopalan, Ann. Rev. Nucl. Part. Sci. 60, 463 (2010). T. Lappi, Int. J. Mod. Phys. E 20, 1 (2011).

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method - notation

light-cone coordinates $x^{\pm} = (t \pm z)/\sqrt{2}$ Milne coordinates $\tau = \sqrt{2x^+x^-} = \sqrt{t^2 - z^2}$ and $\eta = \ln(x^+/x^-)/2 = \ln((t+z)/(t-z))$.

> gauge: sources:

$$\begin{split} A^{\mu}_{\min ne} &= \theta(\tau) \big(0, \, \alpha(\tau, \vec{x}_{\perp}), \vec{\alpha}_{\perp}(\tau, \vec{x}_{\perp}) \big) \\ J^{\mu}(x) &= J^{\mu}_{1}(x) + J^{\mu}_{2}(x) \\ J^{\mu}_{1}(x) &= \delta^{\mu+} g \rho_{1}(x^{-}, \vec{x}_{\perp}) \text{ and } J^{\mu}_{2}(x) = \delta^{\mu-} g \rho_{2}(x^{+}, \vec{x}_{\perp}) \end{split}$$

ansatz:

$$\begin{aligned} A^{+}(x) &= \Theta(x^{-})\Theta(x^{-})x^{+}\alpha(\tau, x_{\perp}) \\ A^{-}(x) &= -\Theta(x^{+})\Theta(x^{-})x^{-}\alpha(\tau, \vec{x}_{\perp}) \\ A^{i}(x) &= \Theta(x^{+})\Theta(x^{-})\alpha_{\perp}^{i}(\tau, \vec{x}_{\perp}) + \Theta(-x^{+})\Theta(x^{-})\beta_{1}^{i}(x^{-}, \vec{x}_{\perp}) + \Theta(x^{+})\Theta(-x^{-})\beta_{2}^{i}(x^{+}, \vec{x}_{\perp}) \end{aligned}$$



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step 1: solve YM equation in the pre-collision region

$$[D_{\mu}, F^{\mu\nu}] = J^{\nu}$$
 with $F_{\mu\nu} = \frac{i}{g}[D_{\mu}, D_{\nu}]$ and $D_{\mu} = \partial_{\mu} - igA_{\mu}$

$$\rho_1(x^+, \vec{x}_\perp) \rightarrow \beta_1^i(x^-, \vec{x}_\perp) \text{ and } \rho_2(x^+, \vec{x}_\perp) \rightarrow \beta_2^i(x^+, \vec{x}_\perp)$$

1st ion:
$$\beta_{1}^{i}(x^{-}, \vec{x}_{\perp}) = \frac{i}{g} U_{1}^{\dagger}(x^{-}, \vec{x}_{\perp}) \partial^{i} U_{1}(x^{-}, \vec{x}_{\perp})$$

 $U_{1}(x^{-}, \vec{x}_{\perp}) = \mathcal{P} \exp\left[ig \int_{-\infty}^{x^{-}} dz^{-} \Lambda_{1}(z^{-}, \vec{x}_{\perp})\right]$
 $\Lambda_{1}(x^{-}, \vec{x}_{\perp}) = \frac{1}{2\pi} \int d^{2} z_{\perp} K_{0}(m(\vec{x}_{\perp} - \vec{z}_{\perp})) \rho_{1}(x^{-}, \vec{z}_{\perp})$

 K_0 is a modified Bessel function similar expression for second ion

physics:

 $\overline{1. \rho_1(x^-, \vec{x}_\perp)}$ is independent of the light-cone time x^+

- the static approximation

2. small width across light cone will be taken to 0
J. Jalilian-Marian, A. Kovner, L. D. McLerran and H. Weigert, Phys. Rev. D 55, 5414 (1997).
Y. V. Kovchegov, Phys. Rev. D 60, 034008 (1999).

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step 2: boundary conditions

$$\begin{aligned} \alpha_{\perp}^{i}(0,\vec{x}_{\perp}) &= \alpha_{\perp}^{i(0)}(\vec{x}_{\perp}) = \lim_{w \to 0} \left(\beta_{1}^{i}(x^{-},\vec{x}_{\perp}) + \beta_{2}^{i}(x^{+},\vec{x}_{\perp}) \right) \\ \alpha(0,\vec{x}_{\perp}) &= \alpha^{(0)}(\vec{x}_{\perp}) = -\frac{ig}{2} \lim_{w \to 0} \left[\beta_{1}^{i}(x^{-},\vec{x}_{\perp}), \beta_{2}^{i}(x^{+},\vec{x}_{\perp}) \right] \end{aligned}$$

A. Kovner, L. D. McLerran and H. Weigert, Phys. Rev. D 52, 6231 (1995). MEC, Czajka, Mrówczyński, arXiv:2012.03042.

step 3: glasma fields (at early times) with proper time expansion

$$\alpha(\tau, \vec{x}_{\perp}) = \alpha(\mathbf{0}, \vec{x}_{\perp}) + \tau \alpha^{(1)}(\vec{x}_{\perp}) + \tau^2 \alpha^{(2)}(\vec{x}_{\perp}) + \cdots$$

and similarly for $\alpha_{\perp}^{i}(\tau, \vec{x}_{\perp})$... (dimensionless small parameter is $\tilde{\tau} = \tau Q_{s}$) coefs of expansion: require vector potential satisfies sourceless YM eqn

$$[D_{\mu},F^{\mu
u}]=0$$
 with $F_{\mu
u}=rac{\imath}{g}[D_{\mu},D_{
u}]$ and $D_{\mu}=\partial_{\mu}-igA_{\mu}$

 $\rightarrow \alpha^{(n)}(\vec{x}_{\perp})$ and $\vec{\alpha}_{\perp}^{(n)}(\vec{x}_{\perp})$ in terms of $\alpha(0, \vec{x}_{\perp})$ and $\vec{\alpha}_{\perp}(0, \vec{x}_{\perp})$

R. J. Fries, J. I. Kapusta and Y. Li, Nucl. Phys. A 774, 861 (2006).

- K. Fukushima, Phys. Rev. C 76, 021902 (2007).
- H. Fujii, K. Fukushima and Y. Hidaka, Phys. Rev. C 79, 024909 (2009).
- G. Chen, R. J. Fries, J. I. Kapusta and Y. Li, Phys. Rev. C 92, 064912 (2015). 👝 🧃 🖌 🖉 🕨 🧸 🚍 🕨

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summary of method:



next: colour charge distributions are not known

- assume Gaussian distribution of colour charges in each nucleus - a product of sources is replace by its average over this distro an average over a Gaussian distribution of independent random variables \rightarrow sum over the averages of all possible pairs (*Wick's theorem*) idea of CGC: local fluctuations \propto surface colour charge density μ

$$\langle
ho_1(x^-, \vec{x}_\perp)
ho_1(y^-, \vec{y}_\perp) \rangle \propto g^2 \, \mu_1(\vec{x}_\perp) \delta(x^- - y^-) \delta^2(\vec{x}_\perp - \vec{y}_\perp)$$

analogous expressions for the second ion

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glasma graph approximation \longrightarrow

J. P. Blaizot, F. Gelis and R. Venugopalan, Nucl. Phys. A 743, 57 (2004).

- K. Fukushima and Y. Hidaka, JHEP 06, 040 (2007).
- F. Fillion-Gourdeau and S. Jeon, Phys. Rev. C 79, 025204 (2009).
- T. Lappi, B. Schenke, S. Schlichting and R. Venugopalan, JHEP 01, 061 (2016).
- J. L. Albacete, P. Guerrero-Rodríguez and C. Marquet, JHEP 01, 073 (2019).

result for correlator of 2 potentials: $(\vec{R} = \frac{1}{2}(\vec{x}_{\perp} + \vec{y}_{\perp}), \vec{r} = \vec{x}_{\perp} - \vec{y}_{\perp})$

$$\begin{split} \delta_{ab} B^{ij}(\vec{x}_{\perp}, \vec{y}_{\perp}) &\equiv \lim_{\mathbf{w} \to 0} \langle \beta_a^i(\mathbf{x}^-, \vec{x}_{\perp}) \beta_b^j(\mathbf{y}^-, \vec{y}_{\perp}) \rangle \\ \lim_{r \to 0} B^{ij}(\vec{x}_{\perp}, \vec{y}_{\perp}) &= \delta^{ij} g^2 \frac{\mu(\vec{R})}{8\pi} \left(\ln \left(\frac{Q_s^2}{m^2} + 1 \right) - \frac{Q_s^2}{Q_s^2 + m^2} \right) + \cdots \end{split}$$

infra-red regulator $m \sim \Lambda_{\rm QCD} \sim 0.2 \text{ GeV}$ ultra-violet regulator = saturation scale = $Q_s = 2 \text{ GeV}$

dots indicate we have kept terms to 2nd order in grad expansion of $\mu(\vec{R})$

J. Jalilian-Marian, A. Kovner, L. McLerran, H. Weigert, Phys. Rev. D 55, 5414 (1997).

H. Fujii, K. Fukushima, Y. Hidaka, Phys. Rev. C 79, 024909 (2009).

G. Chen, R. Fries, J. Kapusta, Y. Li, Phys. Rev. C 92, 064912 (2015).

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surface charge density μ

must specify the form of the surface colour charge density $\mu(\vec{x}_{\perp})$ - 2-dimensional projection of a Woods-Saxon potential

$$\mu(\vec{x}_{\perp}) = \left(\frac{A}{207}\right)^{1/3} \frac{\bar{\mu}}{2a\log(1+e^{R_A/a})} \int_{-\infty}^{\infty} dz \frac{1}{1+\exp\left[(\sqrt{(\vec{x}_{\perp})^2+z^2}-R_A)/a\right]}$$

 R_A and a = radius and skin thickness of nucleus mass number A $r_0 = 1.25$ fm, a = 0.5 fm \rightarrow when A = 207 gives $R_A = r_0 A^{1/3} = 7.4$ fm

normalization: $\mu(\vec{0}) = \bar{\mu} = Q_s^2/g^4$ lead nucleus $g^2 \sqrt{\bar{\mu}} =$ McLerran-Venugopalan (MV) scale

proportional to Q_s - exact relation not determined with CGC approach

E. lancu and R. Venugopalan, in *Quark–Gluon Plasma 3*, eds. R.C. Hwa and X.-N. Wang (World-Scientific, Singapore, 2004), p. 249.

T. Lappi, Eur. Phys. J. C 55, 285-292 (2008).

** numerical results for $\mathcal{E}\dots$ are order of magnitude estimates ratios of different elements of the energy momentum tensor

ightarrow will have much weaker dependence on the MV scale.

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gradient expansion



derivatives are appreciable only in a very small region at the edges non-zero impact parameter (non-central collisions) expand $\mu_{1/2}(\vec{z}_{\perp})$ around ave coord $\vec{R} \mp \vec{b}/2$

summary of method:

YM eqn with average over gaussian distributed valence sources

- \rightarrow correlators of pre-collision fields
- \rightarrow glasma field correlators (b. conds, sourceless YM eqn, τ exp)
- ightarrow correlators of glasma chromodynamic $ec{E}$ and $ec{B}$ fields
- $\Rightarrow \mathsf{observables}$

we work to order $\tau^{\rm 6}$ and study

- 1. isotropization of transverse/longitudinal pressures
- 2. azimuthal momentum distribution and spatial eccentricity
- 3. angular momentum
- 4. momentum broadening of hard probes

comment: many numerical approaches to study initial dynamics our method is fully analytic

- allows control over different approximations and sources of errors
- can be systematically extended
- it has limitations (classical / no fluctuations of positions of nucleons)

at $\tau=\mathrm{0^+}$ the energy-momentum tensor has the diagonal form

$$T(au=0)=\left(egin{array}{cccc} \mathcal{E}_0 & 0 & 0 & 0 \ 0 & -\mathcal{E}_0 & 0 & 0 \ 0 & 0 & \mathcal{E}_0 & 0 \ 0 & 0 & 0 & \mathcal{E}_0 \end{array}
ight)$$

- \rightarrow the longitudinal pressure is large and negative
- system is far from equilibrium
- if the system approaches equilibrium as it evolves:
- the longitudinal pressure must grow
- transverse pressure must decrease ($T_{\mu\nu}$ is traceless)

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to compare longitudinal and transverse pressures ($\tilde{\tau} \equiv \tau Q_s$)

$$A_{TL} \equiv \frac{3(p_T - p_L)}{2p_T + p_L}$$

J. Jankowski, S. Kamata, M. Martinez and M. Spaliński, Phys. Rev. D 104, 074012 (2021). in equilibrium $(p_L = p_T = \mathcal{E}/3) \longrightarrow A_{TL} = 0$



R = 5 fm, $\eta = 0$ and b = 0.

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faster isotropization with smaller impact parameter \rightarrow increased region of overlap



reaction plane defined by collision axis and impact parameter $\phi = 0$ is in reaction plane $\phi = \pi/2$ is perpendicular to reaction plane

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 A_{TL} at order τ^6 $\tau = 0.04$ fm (left panel) $\tau = 0.045$ fm (centre panel) $\tau = 0.05$ fm (right panel)

the axes show R_x and R_y in fm

- the correlator $\langle \beta_a^i(x^-, \vec{x}_\perp) \beta_b^j(y^-, \vec{y}_\perp) \rangle$
- depends on two regulators: m (infra-red) and Q_s (ultra-violet)
- physically related to confinement / saturation scales
- \rightarrow constraints on how to choose them

we used: m = 0.2 GeV and $Q_s = 2.0$ Gev - standard choices

- want results \approx independent of these numbers
- especially since the two scales are pretty close together

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 A_{TL} at order τ^6 as a function of time 3 different values of Q_s with m = 0.2 GeV (left) 3 different values of m with $Q_s = 2.0$ GeV (right) R = 5 fm, b = 0 and $\eta = 0$ at order τ^6 \Rightarrow dependence on these scales is weak



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Radial Flow

transverse momentum flow vector = T_{i0} (trans. Poynting vector) radial flow of the expanding glasma = radial projection $P \equiv \hat{R}_i T_{i0}$ $\phi = \pi/2$ is perpendicular to the reaction plane



- at lowest order P increases linearly with time
- including higher order contros ightarrow P slows as system expands
- order τ^5 shows flattening up to about $\tilde{\tau}=0.5$
- \rightarrow indicates the limit of validity of the τ expansion

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Azimuthal asymmetry

in a non-central collision - initial spatial asymmetry relativistic collision \rightarrow spatial asymmetries rapidly decrease \rightarrow anisotropic momentum flow can develop only in the first fm/c

- sensitive to system properties very early in its evolution
- provides direct information about the early stages of the system

$$\begin{split} \varphi(\vec{x}_{\perp}) &= \tan^{-1} \left(\frac{T^{0y}(\vec{x}_{\perp})}{T^{0x}(\vec{x}_{\perp})} \right) \\ W(\vec{x}_{\perp}) &\equiv \sqrt{\left(T^{0x}(\vec{x}_{\perp}) \right)^2 + \left(T^{0y}(\vec{x}_{\perp}) \right)^2} \\ P(\phi) &\equiv \frac{1}{\Omega} \int d^2 \vec{x}_{\perp} \, \delta(\phi - \varphi(\vec{x}_{\perp})) \, W(\vec{x}_{\perp}), \quad \Omega \equiv \int d^2 \vec{x}_{\perp} \, W(\vec{x}_{\perp}) \\ P(\phi) &= \frac{1}{2\pi} \left(1 + 2 \sum_{n=1}^{\infty} v_n \cos(n\phi) \right) \\ v_n &= \int_0^{2\pi} d\phi \, \cos(n\phi) \, P(\phi) \end{split}$$

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first 3 fourier coefficients at $\tau = 0.04$ fm left panel: b = 2 fm; right panel $\eta = 0.01$



 v_2 and v_3 are \sim experimental values v_1 is bigger than expected

note: usually assumed anisotropy develops mostly during hydro evolution \rightarrow our results for all three Fourier coefficients are large

comment:

experimentally: impact parameter / reaction plane not precisely known our calculation does not correspond exactly to what is measured

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• spatial deviations from azimuthal symmetry

$$\varepsilon_n = -\frac{\int d^2 \vec{R} |\vec{R}| \cos(n\phi) \mathcal{E}(\vec{R})}{\int d^2 \vec{R} |\vec{R}| \mathcal{E}(\vec{R})} \quad \text{with} \quad \phi = \tan^{-1}(R_y/R_x)$$

where $\mathcal{E}(\vec{R})$ denotes the energy density



 $\tau = 0.04 \text{ fm and } \eta = 0 \text{ [normalized to 1 at } b = 0.5 \text{ fm]}$

 \rightarrow correlation btwn spatial asymmetry introduced by the initial geometry and anisotropy of azimuthal momentum distribution

these correlations \sim characteristic of onset of hydrodynamic behaviour

angular momentum

define tensor $M^{\mu\nu\lambda} = T^{\mu\nu}R^{\lambda} - T^{\mu\lambda}R^{\nu}$ with $R^{\mu} = (\tau, \eta, \vec{R})$

 $abla_{\mu} M^{\mu\nu\lambda} = 0
ightarrow$ conserved charges $J^{\nu\lambda} = \int_{\Sigma} d^3 y \sqrt{|\gamma|} n_{\mu} M^{\mu\nu\lambda}$

- n^{μ} is a unit vector perpendicular to the hypersurface Σ
- γ is the induced metric on this hypersurface
- d^3y is the corresponding volume element
- $n^{\mu}=(1,0,0,0)$ in Milne coordinates $ightarrow J^{
 u\lambda}$ defined on a hypersurface of constant au

Pauli-Lubanski vector:
$$L_{\mu}=-rac{1}{2}\epsilon_{\mulphaeta\gamma}J^{lphaeta}u^{\gamma}$$

result: angular momentum per unit rapidity (symmetric collision)

$$\frac{dL^{y}}{d\eta} = -\tau^2 \int d^2 \vec{R} \, R^{\times} T^{0z}$$

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ions moving in +/-z dirns displaced in +/-x dirns $\rightarrow L_y$ is negative

warning: dominant contro to \vec{L} from regions farthest from collision centre = regions gradient expansion is least trusted \rightarrow error bars large

comparison:

 $L_y\sim 10^5$ at RHIC energies for initial system of colliding ions J. H. Gao, S. W. Chen, W. t. Deng, Z. T. Liang, Q. Wang and X. N. Wang



- even larger at LHC energies F. Becattini, F. Piccinini and J. Rizzo, Phys. Rev. C 77, 024906 (2008). idea: initial rapid rotation of glasma

 \rightarrow could be observed via polarization of final state hadrons

- large \vec{L} & spin-orbit coupling ightarrow alignment of spins with \vec{L}

many experimental searches for this polarization

- effect of a few percent observed at RHIC
- at LHC result consistent with zero
- difficult to measure ...
- F. Becattini, M.A. Lisa, Ann. Rev. Nucl. Part. Sci. 70, 395 (2020).

these results are consistent our calculation: glasma carries only tiny imprint of the \vec{L} of the intial state

ightarrow majority of the angular momentum is carried by valence quarks

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idea:

hard probes produced via hard interactions at earliest phase of HIC

- propagate through the evolving medium
- suppression of high- p_T probes (jet quenching)
- \Rightarrow signal of formation of QGP
- deconfined state of matter = significant braking of hard partons

EL and MB of hard probes / equilibrium plasma studied extensively • contro from pre-equilibrium phases has been largely ignored however, see for example: Ruggieri, Das et al, Phys. Rev. D 98, 094024 (2018).

Boguslavski, Kurkela, Lappi, Peuron, JHEP 09, 077 (2020).

Ipp, Müller, Schuh, Phys. Lett. B 810, 135810 (2020).

D. Pablos, M. Singh, S. Jeon and C. Gale, arXiv:2202.03414.

Siggi Hauksson, talk presented at SEWM_2022, 21/06.

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physics: frequent small \vec{p} exchanges btwn probe and glasma fields \rightarrow transport equation in Fokker-Planck form

- describes interactions of hard probe interacting with glasma fields see talk by Stanisław Mrówczński, 21/06 14:00

$$\left(\mathcal{D}-\nabla_{p}^{\alpha}X^{\alpha\beta}(\vec{v})\nabla_{p}^{\beta}-\nabla_{p}^{\alpha}Y^{\alpha}(\vec{v})\right)n(t,\vec{x},\vec{p})=0$$

notation: $\alpha \in (1, 2, 3)$

$$\begin{split} n(t, \vec{x}, \vec{p}) &= \text{distribution function of heavy quarks} \\ \vec{v} &= \vec{p}/E_{\vec{p}} = \vec{p}/\sqrt{p^2 + m_Q^2} = \text{velocity of heavy quark} \\ \mathcal{D} &\equiv \frac{\partial}{\partial t} + \vec{v} \cdot \vec{\nabla} = \text{material derivative (drift term)} \end{split}$$

 Y^{lpha} related to collisional energy loss

 $X^{lphaeta}$ related to momentum broadening

St. Mrówczyński, Eur. Phys. J. A 54, 43 (2018).

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$$\begin{aligned} \hat{q} &= \frac{1}{v} \Big(\delta^{\alpha\beta} - \frac{v^{\alpha}v^{\beta}}{v^{2}} \Big) \frac{\langle \Delta p^{\alpha} \Delta p^{\beta} \rangle}{\Delta t} \\ &= \frac{2}{v} \Big(\delta^{\alpha\beta} - \frac{v^{\alpha}v^{\beta}}{v^{2}} \Big) X^{\alpha\beta}(\vec{v}) \\ X^{\alpha\beta}(\vec{v}) &\equiv \frac{1}{2N_{c}} \int_{0}^{t} dt' \operatorname{Tr} \big[\langle \mathcal{F}^{\alpha}(t,\vec{x}) \mathcal{F}^{\beta}(t-t',\vec{x}-\vec{v}t') \rangle \big] \end{aligned}$$

colour Lorentz force: $\vec{\mathcal{F}}(t,\vec{x})\equiv gig(\vec{E}(t,\vec{x})+\vec{v} imes \vec{B}(t,\vec{x})ig)$

$$\begin{split} X^{\alpha\beta}(\mathbf{v}) &= \frac{g^2}{2N_c} \int_0^t dt' \Big[\langle E_a^{\alpha}(t,\vec{x}) E_a^{\beta}(t-t',\vec{y}) \rangle + \epsilon^{\beta\gamma\gamma'} v^{\gamma} \langle E_a^{\alpha}(t,\vec{x}) B_a^{\gamma'}(t-t',\vec{y}) \rangle \\ &+ \epsilon^{\alpha\gamma\gamma'} v^{\gamma} \langle B_a^{\gamma'}(t,\vec{x}) E_a^{\beta}(t-t',\vec{y}) \rangle + \epsilon^{\alpha\gamma\gamma'} \epsilon^{\beta\delta\delta'} v^{\gamma} v^{\delta} \langle B_a^{\gamma'}(t,\mathbf{x}) B_a^{\delta'}(t-t',\vec{y}) \rangle \Big] \end{split}$$

where $\vec{y} = \vec{x} - \vec{v}t'$.

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note: combination of two approaches

1. medium that the hard probe interacts with is a glasma

 \rightarrow described with CGC effective theory with proper time expansion ** description is valid only at very early times

2. FP eqn describes interactions of hard probe with glasma fields ** valid at times long enough that collision terms saturate

 \Rightarrow conflict btwn assumptions that set these two time scales also:

- FP description requires gradient expansion type approximations
- our CGC approach assumes boost invariance
- ** can all these conditions can be satisfied simultaneously?

• (1) • (2) • (3) • (3) • (3)

result: \hat{q} as a function of τ at different orders in the expansion



key: saturation regime appears before τ expansion breaks down caution:

figure above obtained for $v_\perp = v$ calculation works less well when $v_\parallel \neq 0$

reason: at very early times glasma fields represented as longitudinal flux tubes



 λ_{\perp} can be inferred the 2-point correlator

 \hat{q} built up during time probe is in domain of correlated fields at zeroth order this time is determined by

- transverse correlation length
- orientation and magnitude of the probe's velocity
- \rightarrow saturation is faster if $v_{\parallel}=0$

note: probe's velocity also enters through the Lorentz force



fifth and fourth order results for increasing v_{\parallel}

- saturation is less pronounced as v_{\parallel} increases



 Q_s between 1.9 (bottom) and 2.1 (top) GeV with ratio Q_s/m fixed - one sees that the dependence is fairly weak

boost and translation invariance

calculation of \hat{q} is formulated in Minkowski space

- assumes at least approximate translation invariance

but used correlators of electric and magnetic fields obtained from an boost invariant ansatz for the vector potential

check of consistency: previous result was $z = \eta = 0$ (red line) \hat{q} as a function of τ for three values of η



radiative Eloss/length of probe traversing medium of length L \propto total accumulated transverse momentum broadening

$$\Delta p_T^2 = \int_0^L dt \hat{q}(t)$$

our calculation gives $\hat{q}_{\mathrm{max}} = 6~\mathrm{GeV}^2/\mathrm{fm}$

compare with equilibrium values: hard quark of $p_T > 40 \text{ GeV} \longrightarrow 2 < \hat{q}/T^3 < 4$ - inferred from experimental data *S. Cao et al. [JETSCAPE], Phys. Rev. C* 104, 024905 (2021).

 $\hat{q} = 3T^3$ and $450 > T > 150 \text{ MeV} \rightarrow (0.05 < \hat{q} < 1.0) \text{ GeV}^2/\text{fm}$ \Rightarrow equilibrium value of \hat{q} is much smaller

<u>but</u> τ_{life} of pre-equilibrium phase < 1 fm/c \rightarrow contro of pre-equilibrium phase to jet quenching usually ignored

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schematic representation of the time dependence of \hat{q}



- 1. rapid growth to $\hat{q}_{\mathrm{max}} pprox$ 6 $\mathrm{GeV}^2/\mathrm{fm}$ at $t_{\mathrm{max}} pprox$ 0.06 fm
- this is a rough description of our result
- no saturation region because of time scales
- 2. decrease from $t_{\rm max} \rightarrow t_0$ not captured by our calculation
- is reproduced by the simulations

A. Ipp, D. I. Müller and D. Schuh, Phys. Lett. B 810, 135810 (2020)

$$ightarrow \Delta p_T^2 \Big|^{
m non-eq} = \int_0^{t_0} dt \, \hat{q}(t) = rac{1}{2} \hat{q}_{
m max} t_0 + rac{1}{2} \hat{q}_0(t_0 - t_{
m max})$$

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3. assume hydro evolution from t_0

- using 1d boost invariant hydrodynamics

$$\hat{q}=3T^3$$
 with $T=T_0 \Big(rac{t_0}{t}\Big)^{1/3}$ and

$$\Delta p_T^2 \Big|^{
m eq} = \int_{t_0}^L dt \, \hat{q}(t) = 3T_0^3 t_0 \, \ln rac{L}{t_0}$$

values:

 $t_0 = 0.6 \text{ fm}, \ T_0 {=} 0.45 \text{ GeV}$

C. Shen, U. Heinz, P. Huovinen and H. Song, Phys. Rev. C 84, 044903 (2011) $L=10~{
m fm}$

result:
$$\frac{\Delta p_T^2 [\text{non-equib}]}{\Delta p_T^2 [\text{equib}]} \approx 0.93$$

rough estimate that depends on values of parameters chosen - but result is not very sensitive to values of shape of peak

 \Rightarrow glasma plays an important role in jet quenching _ , , , , ,

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conclusions

- 1. developed an efficient method to calculate correlators of electric and magnetic fields using a CGC approach and a proper time expansion
- 2. 6th order au expansion can be trusted to about au = 0.05 fm
- 3. correlation btwn elliptic flow coef v_2 / spatial eccentricity
 - spatial asymmetry introduced by initial geometry is effectively transmitted to azimuthal distribution of gluon momentum field

 \rightsquigarrow indication of the onset of hydrodynamics.

- 4. most of the angular momentum of the intial system not transmitted to glasma
 - contradicts picture of a rapidly rotating initial glasma state
- 5. glasma plays an important role in jet quenching

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