Flat fielding for spectra with AuxTel

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Flat-fielding in spectroscopy



Factorization of the wavelength dependence

Idea:

• Evaluate spatial correlation between flats with different filters

 $ADU(i, j, \lambda) = F_{o.f.}(i, j) imes F_{CCD}(i, j, \lambda) = F_{o.f.}(i, j) imes G_{CCD}(i, j) imes arepsilon_{CCD}(\lambda)$

- See if can factorize out the wavelength dependency on the flats:
 - $F_{o.f.}(i, j)$ = **out of focus artifacts** (vignetting, dust on optical components) = **slow** pixel to pixel variation (real space) or **low** spatial frequency (Fourier space)
 - $F_{CCD}(i, j, \lambda)$ = focused artifacts (dust on the CCD, pixel surface variations) = fast pixel to pixel variation or high spatial frequency
 - \circ We examine the **hypothesis** of $F_{CCD}(i,j,\lambda) = G_{CCD}(i,j) imes arepsilon_{CCD}(\lambda)$
- If ~ true, then we could preliminary use a single flat for spectra reduction
- How can we test this?
 - By examining the **ratio** of flat images **at different wavelengths**

Factorization of the wavelength dependence: pixel correlation



Factorization of the wavelength dependence: pixel correlation



Steps: master bias

- We start with a set of *N* bias images: $B^{(\mu)}(i,j)$ ($\mu=1,2,\ldots,N$)
- We create the **master bias**, *B*, such that the pixel (*i*,*j*) is the **median** of the *N* images at the same pixel, that is,



$$\mid \; \mathcal{B}(i,j) \; = \; median_{\mu} \Big(B^{(\mu)}(i,j) \Big)$$



- We start with a set of *N* flat images with a given filter *b* (FELH0600, SDSSr, SDSSg): $F_b^{(\mu)}(i,j)$ ($\mu = 1, 2, ..., N$), where $F_b^{(\mu)}(i,j) = F_b^{(\mu)}(i,j) B$
- We take the median over all pixels for each image and normalise by it:

$$med^{(\mu)} = median_{(i,j)} \Big(F_b^{(\mu)} \Big) \quad | \quad F_b^{(\mu)}(i,j) \
ightarrow \ F_b^{\prime(\mu)}(i,j) = rac{F_b^{(\mu)}(i,j)}{med^{(\mu)}}$$

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- We take the median over all pixels for each image and normalise by it: $F_{L}^{\prime(\mu)}$ (i,j)



- There are two different components on the signal:
 - Electronics (which we want to keep) + dust on CCD (focused artifacts)
 - Smooth gradients (vignetting) and extended effects (dust and out of focus artifacts)
- To capture the smooth / extended components, we apply a median spatial filter (window of 40x40 pixels)
- We do this process for each register / segment separately
 - To account for gain variations



- We start with a set of *N* flat images with a given filter *b* (FELH0600, SDSSr, SDSSg): $F_b^{(\mu)}(i,j)$ ($\mu = 1, 2, ..., N$), where $F_b^{(\mu)}(i,j) = F_b^{(\mu)}(i,j) \mathcal{B}_b^{(\mu)}(i,j)$
- We take the median over all pixels for each image and normalise by it: F_b^{'(µ)}
 We take the median over the N F_b^{'(µ)} flat images at each pixel (*i,j*): F_b
- We compute the **smooth component**, f_b , by replacing each pixel value by the median in a 40x40 sliding window
- We remove this smooth component:

$${ ilde F}_b o {\mathcal F}_b \hspace{0.1 cm} \mid \hspace{0.1 cm} {\mathcal F}_b(i,j) = rac{{ ilde F}_b(i,j)}{f_b(i,j)}$$

Profile of a column before and after removing the smooth component





Profile of a column before and after removing the smooth component





Master flat / smoothed



Master flat (low + high frequency)



Master flat (high frequency)



Master flats ratio



Master flats ratio



16

Master flats ratio



17



18



le6 FELH0600 le6 SDSSg FELH0600 / SDSSg 1.04 - 3.5 2.0 - 3.0 **Full CCD** 1.02 - 2.5 1.5 DSSDS 1.00 - 2.0 - 1.5 1.0 - 1.0 0.98 - 0.5 0.5 0.96 - 0.0 $\rho = 0.822$ 0.96 0.98 1.00 1.02 1.04 0.0 FELH0600 0.97 1.03 0.96 0.98 0.99 1.00 1.01 1.02 1.04

This seems to be due to an issue on some segments of the CCD



Conclusions and next steps

 We find a good enough spatial correlation (ρ > 0.8) between pixels in flat fields of different colours (FELH0600 and SDSSg) to factor out the λ dependence

 $ADU(i, j, \lambda) = F_{o.f.}(i, j) imes F_{CCD}(i, j, \lambda) = F_{o.f.}(i, j) imes G_{CCD}(i, j) imes arepsilon_{CCD}(\lambda)$

- We need to repeat this with **more filters** (more separated in wavelength)
- If results are consistent, we propose to preliminary **use a single high spatial frequency master flat** (as previously created) for AuxTel spectra deflatening

$$I = rac{D - \mathcal{B}}{\mathcal{F}}$$

Merci

Back-up

Segment numbering convention



λ independence of $F_{o.f.}(i,j)$

- The optical system is supposed to be achromatic
- We check the histograms **before removing** the smooth component:



FELH0600 is bigger, so there is less vignetting coming from its frame



Identical distributions (vignetting tails) for SDSSg (bluer) and SDSSr (redder)

λ independence of $F_{o.f.}(i,j)$

- The optical system is supposed to be **achromatic**
- We check the histograms **before removing** the smooth component
- Equivalent results for SDSSr are found after removing the smooth component



Master flats ratio: spatial correlation of pixel content (SDSSg - SDSSr)



Master flats ratio: spatial correlation of pixel content (SDSSg - SDSSr)



Master flats ratio: spatial correlation of pixel content (SDSSg - SDSSr)

